

Time independent perturbation theory:

→ Non-degenerate Case:

Suppose $H = H_0 + \lambda V$
 ↑
 unperturbed
 Hamiltonian.

Assume we know $H_0 |n^{(0)}\rangle = E_n^{(0)} |n^{(0)}\rangle$

Now want to solve: $(H_0 + \lambda V) |n\rangle = E_n |n\rangle$

assume λ is small; Now find $E_n, |n\rangle$ in series of λ .

Ex: 2-level system.

$$H_0 = \begin{pmatrix} E_1^{(0)} & 0 \\ 0 & E_2^{(0)} \end{pmatrix} \quad \text{with } |1^{(0)}\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$$

$$|2^{(0)}\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

$$\lambda V = \begin{pmatrix} 0 & \lambda V_{12} \\ \lambda V_{21} & 0 \end{pmatrix} \quad \text{with } V_{12} = \langle 1^{(0)} | V | 2^{(0)} \rangle$$

$$\sqrt{V_{21}} = V_{12}$$

For simplicity, $V_{11} = V_{22} = 0$.

$$H = \begin{pmatrix} E_1^{(0)} & \lambda V_{12} \\ \lambda V_{21} & E_2^{(0)} \end{pmatrix}$$

First diagonalize Hamiltonian.

$$(E_1^{(0)} - E)(E_2^{(0)} - E) - \lambda^2 |V_{12}|^2 = 0$$

$$E = \frac{E_1^{(0)} + E_2^{(0)}}{2} \pm \sqrt{\left(\frac{E_1^{(0)} - E_2^{(0)}}{2}\right)^2 + \lambda^2 |V_{12}|^2}$$

Expansion requires

$$\left(\frac{2\lambda V_{12}}{E_1^{(0)} - E_2^{(0)}}\right)^2 < 1 \quad \approx \quad \frac{E_1^{(0)} + E_2^{(0)}}{2} \pm \frac{E_1^{(0)} - E_2^{(0)}}{2} \sqrt{1 + \left(\frac{2\lambda V_{12}}{E_1^{(0)} - E_2^{(0)}}\right)^2}$$

$$\text{then } E_1 = E_1^{(0)} + \lambda^2 \frac{|V_{12}|^2}{E_1^{(0)} - E_2^{(0)}} + \dots \quad \left. \begin{array}{l} \text{require } \frac{2\lambda V_{12}}{E_1^{(0)} - E_2^{(0)}} \ll 1 \\ \text{so } \lambda V_{12} \ll \frac{E_1^{(0)} - E_2^{(0)}}{2} \end{array} \right\}$$

Formal Theory of Perturbation Expansion:

$$H_0 |n^{(0)}\rangle = E_n^{(0)} |n^{(0)}\rangle$$

assume $\langle n | n \rangle \neq 1$
but $\langle n | n^{(0)} \rangle = 1$

Suppose $\Delta_n = E_n - E_n^{(0)}$

λ is small parameter.
expect $|n\rangle = |n^{(0)}\rangle + \lambda |n^{(1)}\rangle + \lambda^2 |n^{(2)}\rangle + \dots$

$$\Delta_n = \lambda \Delta_n^{(1)} + \lambda^2 \Delta_n^{(2)} + \lambda^3 \Delta_n^{(3)} \dots$$

Formal Development of Perturbation Expansion:

Suppose we know the unperturbed energy eigenket and eigen energy:

$$H_0 |n^{(0)}\rangle = E_n^{(0)} |n^{(0)}\rangle$$

Here $|n^{(0)}\rangle$ is complete such that $\sum_n |n^{(0)}\rangle \langle n^{(0)}| = 1$.

Assume energy spectrum is non-degenerate:

Now we add perturbation and introduce parameter λ where λ goes from $[0, 1]$. We set $\lambda = 1$ in the end.

$$(H_0 + \lambda V) |n\rangle = E_n^{(\lambda)} |n\rangle$$

Now define energy shift from unperturbed state as:

$$\Delta_n = E_n - E_n^{(0)}$$

Now we expect $|n\rangle$ to expand in λ :

$$|n\rangle = |n^{(0)}\rangle + \lambda |n^{(1)}\rangle + \lambda^2 |n^{(2)}\rangle + \dots$$

assume $|n\rangle$ is not normalized, $\langle n | n \rangle \neq 1$
but $\langle n | n^{(0)} \rangle = 1$

$$\Delta_n = \lambda \Delta_n^{(1)} + \lambda^2 \Delta_n^{(2)} + \dots$$

and

Then we need to solve for

$$(H_0 + \lambda V)|n\rangle_\lambda = E_n^{(0)}|n\rangle_\lambda$$

$$\hookrightarrow (H_0 + \lambda V)|n\rangle_\lambda = (\Delta_n + E_n^{(0)})|n\rangle_\lambda$$

$$\hookrightarrow \boxed{(E_n^{(0)} - H_0)|n\rangle_\lambda = (\lambda V - \Delta_n)|n\rangle_\lambda}$$

Note: We cannot do $|n\rangle_\lambda = \frac{(\lambda V - \Delta_n)}{E_n^{(0)} - H_0}|n\rangle_\lambda$

Since it can be non-invertible since when $|n\rangle$ contains $|n^{(0)}\rangle$
which give $\frac{1}{E_n^{(0)} - E_n^{(0)}} = \text{bad.}$

Now multiply $\langle n^{(0)} |$ on both sides.

$$\underbrace{\langle n^{(0)} | E_n^{(0)} - H_0 | n \rangle_\lambda}_{=0} = \langle n^{(0)} | \lambda V - \Delta_n | n \rangle_\lambda$$

so $\boxed{\langle n^{(0)} | \lambda V - \Delta_n | n \rangle_\lambda = 0} *$

* ϕ_n gets rid of $|n^{(0)}\rangle$ term.
Now let's introduce: $\phi_n = 1 - |n^{(0)}\rangle \langle n^{(0)}| = \sum_{k \neq n} |k^{(0)}\rangle \langle k^{(0)}|$

It has property $\begin{cases} \phi_n |n^{(0)}\rangle = \langle n^{(0)}| \phi_n = 0 \\ \phi_n^2 = \phi_n \end{cases}$

also note

here we used:
 $\langle n^{(0)} | \lambda V - \Delta_n | n \rangle_\lambda = 0$

$$\phi_n (\lambda V - \Delta_n) |n\rangle = (1 - |n^{(0)}\rangle \langle n^{(0)}|)(\lambda V - \Delta_n) |n\rangle = (\lambda V - \Delta_n) |n\rangle$$

Note that $\frac{1}{E_n^{(0)} - H_0}$ is defined when multiplied to ϕ_n since there is no $|n^{(0)}\rangle \langle n^{(0)}|$ term.

$$\Rightarrow \frac{1}{E_n^{(0)} - H_0} \phi_n = \sum_{k \neq n} \frac{1}{E_n^{(0)} - E_k^{(0)}} |k^{(0)}\rangle \langle k^{(0)}|$$

Now you might be tempted to write

$$(E_n^{(0)} - H_0)|n\rangle = (\lambda V - \Delta_n)|n\rangle \Rightarrow |n\rangle = \frac{1}{(E_n^{(0)} - H_0)} \phi_n (\lambda V - \Delta_n) |n\rangle$$

However this cannot be write since we must have

$|n\rangle \rightarrow |n^{(0)}\rangle$ as $\lambda \rightarrow 0$ and $\Delta_n \rightarrow 0$. So let's add a homogeneous solution $\underline{c_n(\lambda) |n^{(0)}\rangle}$

$$|n\rangle = \cancel{c_n(\lambda)} |n^{(0)}\rangle + \frac{1}{E_n^{(0)} - H_0} \phi_n (\lambda V - \Delta_n) |n\rangle$$

$$\text{where } \lim_{\lambda \rightarrow 0} c_n(\lambda) = \lim_{\lambda \rightarrow 0} \langle n^{(0)} | n \rangle = 1$$

Right we set normalization $\langle n^{(0)} | n \rangle = c_n(\lambda) = 1$.

$$\therefore |n\rangle = |n^{(0)}\rangle + \frac{\phi_n}{E_n^{(0)} - H_0} (\lambda V - \Delta_n) |n\rangle$$

Previously we saw $\langle n^{(0)} | \lambda V - \Delta_n | n \rangle = 0$

Rearrange: $\lambda \langle n^{(0)} | V | n \rangle = \Delta_n \underbrace{\langle n^{(0)} | n \rangle}_{\text{some } \#} = 1 \text{ by normalization}$

so $\Delta_n = \lambda \langle n^{(0)} | V | n \rangle$

now with $|n\rangle = |n^{(0)}\rangle + \lambda |n^{(1)}\rangle + \lambda^2 |n^{(2)}\rangle + \dots$

$$\Delta_n = \lambda \Delta_n^{(1)} + \lambda^2 \Delta_n^{(2)} + \lambda^3 \Delta_n^{(3)} + \dots$$

$\hookrightarrow \Delta_n = \lambda \langle n^{(0)} | V | n \rangle$

$\hookrightarrow \lambda \Delta_n^{(1)} + \lambda^2 \Delta_n^{(2)} + \lambda^3 \Delta_n^{(3)} = \lambda \langle n^{(0)} | V (|n^{(0)}\rangle + \lambda |n^{(1)}\rangle + \lambda^2 |n^{(2)}\rangle + \dots)$

By matching λ terms:

$$\mathcal{O}(\lambda^1): \Delta_n^{(1)} = \langle n^{(0)} | V | n^{(0)} \rangle$$

$$\mathcal{O}(\lambda^2): \Delta_n^{(2)} = \langle n^{(0)} | V | n^{(1)} \rangle$$

$$\mathcal{O}(\lambda^3): \Delta_n^{(3)} = \langle n^{(0)} | V | n^{(2)} \rangle$$

$$\vdots \quad \vdots \quad \vdots$$

$$\mathcal{O}(\lambda^N): \Delta_n^{(N)} = \langle n^{(0)} | V | n^{(N-1)} \rangle$$

Note we need energy correction of previous terms, in order to proceed

but we still need to know terms $|n^{(\neq 0)}\rangle$.

Now let's find $|n^{(\neq 0)}\rangle$ in order to evaluate Δ_n :

we know $|n\rangle = |n^{(0)}\rangle + \frac{\phi_n}{E_n^{(0)} - H_0} (\lambda V - \Delta_n) |n\rangle$

expand $|n\rangle$ and Δ_n in terms of λ :

$$|n^{(0)}\rangle + \lambda |n^{(1)}\rangle + \dots = |n^{(0)}\rangle + \frac{\phi_n}{E_n^{(0)} - H_0} (\lambda V - \lambda \Delta_n^{(1)} - \lambda^2 \Delta_n^{(2)} - \dots) (|n^{(0)}\rangle + \lambda |n^{(1)}\rangle + \dots)$$

Now collect terms in λ :

$$\mathcal{O}(\lambda^0) : |n^{(0)}\rangle = |n^{(0)}\rangle$$

$$\mathcal{O}(\lambda) : |n^{(1)}\rangle = \frac{\phi_n}{E_n^{(0)} - H_0} (V - \Delta_n^{(1)}) |n^{(0)}\rangle = \frac{\phi_n}{E_n^{(0)} - H_0} V |n^{(0)}\rangle$$

then with

$$\Delta_n^{(1)} = \langle n^{(0)} | V | n^{(0)} \rangle$$

$$\Rightarrow \boxed{\Delta_n^{(1)} = \langle n^{(0)} | V \frac{\phi_n}{E_n^{(0)} - H_0} V | n^{(0)} \rangle}$$

Now let's rewrite $\frac{\phi_n}{E_n^{(0)} - H_0} = \sum_{k \neq n} \frac{1}{E_n^{(0)} - E_k^{(0)}} |k\rangle \langle k|$

then $|n^{(1)}\rangle = \frac{\phi_n}{E_n^{(0)} - H_0} V |n^{(0)}\rangle$

$$= \sum_{k \neq n} \frac{1}{E_n^{(0)} - E_k^{(0)}} |k\rangle \langle k| V |n^{(0)}\rangle$$

$$\boxed{|n^{(1)}\rangle = \sum_{k \neq n} \frac{V_{kn}}{E_n^{(0)} - E_k^{(0)}} |k\rangle}$$

for $V_{kn} = \langle k | V | n^{(0)} \rangle$

Here note that
 $\phi_n \Delta_n^{(1)} |n^{(0)}\rangle = 0$
 $\hookrightarrow \Delta_n^{(1)} (1 - |n^{(0)}\rangle \langle n^{(0)}|) |n^{(0)}\rangle = 0$

$$\text{Similarly } \Delta_n^{(2)} = \langle n^{(0)} | V \frac{\phi}{E_n^{(0)} - E_0} V | n^{(0)} \rangle$$

$$= \langle n^{(0)} | V \sum_{k \neq n} \frac{| k^{(0)} \rangle \langle k^{(0)} |}{E_n^{(0)} - E_k} V | n^{(0)} \rangle$$

$$\boxed{\Delta_n^{(2)} = \sum_{k \neq n} \frac{|V_{kn}|^2}{E_n^{(0)} - E_k}} \quad \leftarrow \text{2nd order energy correction.}$$

Remarks:

- 1st order energy correction only requires eigenket of the unperturbed state.
- 2nd order energy correction:

→ suppose $n=0$, ground state, we see $\frac{|V_{k0}|^2}{E_0^{(0)} - E_k}$ always give negative value, so the overall energy goes down.

→ the first excited level, $n=1$, we see $\frac{|V_{k1}|^2}{E_1^{(0)} - E_k}$ give slightly less negative compared to ground level $n=0$, so their overall energy goes in opposite direction, which is the no-level-crossing theorem.

⇒ Applicability of perturbation theory: $\lambda |V_{kn}|^2 \ll E_n^{(0)} - E_k^{(0)}$

* Summary: $|n\rangle = |n^{(0)}\rangle + \lambda \sum_{k \neq n} \frac{|V_{kn}|}{E_n^{(0)} - E_k^{(0)}} |k^{(0)}\rangle$

$$+ \lambda^2 \left(\sum_{k \neq n} \sum_{l \neq n} \frac{|k^{(0)}\rangle \langle k_l| V_{kl} V_{ln}}{(E_n^{(0)} - E_k^{(0)})(E_n^{(0)} - E_l^{(0)})} - \sum_{k \neq n} \frac{|k^{(0)}\rangle \langle V_{nn} V_{kn}|}{(E_n^{(0)} - E_k^{(0)})^2} \right) + \dots$$

$$E_n = E_n^{(0)} + \lambda V_{nn} + \lambda^2 \sum_{k \neq n} \frac{|V_{nk}|^2}{E_n^{(0)} - E_k^{(0)}}$$

$$+ \lambda^3 \sum_{k \neq n} \sum_{l \neq n} \frac{|V_{nk} V_{lk} V_{ln}|}{(E_n^{(0)} - E_k^{(0)})(E_n^{(0)} - E_l^{(0)})} - V_{nn} \sum_k \frac{|V_{nk}|^2}{(E_n^{(0)} - E_k^{(0)})^2} + \dots$$

Note that $|n\rangle = |n^{(0)}\rangle + \lambda|n^{(1)}\rangle + \lambda^2|n^{(2)}\rangle + \dots$
is not normalized.

To normalize:

$$N \langle n|n \rangle_N = z_n \langle n|n \rangle = 1$$

so $z_n^{-1} = \langle n|n \rangle$

$$\begin{aligned} &= 1 + \lambda^2 \langle n^{(1)}|n^{(0)}\rangle + \mathcal{O}(\lambda^3) \\ &= 1 + \lambda^2 \sum_k' \frac{|V_{kn}|^2}{(E_n^{(0)} - E_k^{(0)})^2} \end{aligned}$$

then

$$z_n = 1 - \lambda^2 \sum_k \frac{|V_{kn}|^2}{(E_n^{(0)} - E_k^{(0)})^2}$$

let $|n\rangle_N = \sqrt{z_n} |n\rangle$

$$\langle n^{(0)}|n\rangle_N = \sqrt{z_n} \underbrace{\langle n^{(0)}|n\rangle}_{=1} \text{ by previous definition of } |n\rangle$$

$$z_N = |\langle n^{(0)}|n\rangle_N|^2$$

$$z_N = 1 - \lambda^2 \sum_{k \neq n} \frac{|V_{kn}|^2}{(E_n^{(0)} - E_k^{(0)})^2}$$

← side note: $z_N = \frac{\partial E_n}{\partial E_n^{(0)}}$

probability for leakage to states other than $|n^{(0)}\rangle$.

Ex: $H = \frac{p^2}{2m} + \frac{1}{2}m\omega^2 x^2 + \frac{1}{2}\epsilon m\omega_0^2 x^2$

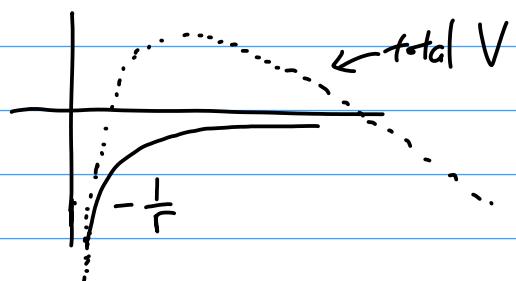
Two ways: 1) exact solution and expand in ϵ .
2) perturbation theory in ϵ

Quadratic Stark Effect:

$$H = \frac{p^2}{2m} + V_0(r) + -e\vec{E} \cdot \vec{z}$$

$V_0 \propto \frac{1}{r}$

Imagine spinless electron:



Find shift in ground state.

$$\begin{aligned} \Delta_K &= \Delta_K^{(1)} + \Delta_K^{(2)} \\ &= e|\vec{E}| \cancel{Z_{KK}} + e^2 |E|^2 \sum_j \frac{|Z_{Kj}|^2}{E_K^{(0)} - E_j^{(0)}} = \frac{1}{2} \alpha |E|^2 \end{aligned}$$

but $\langle K^{(0)} | Z | K^{(0)} \rangle = 0$ here $|K^{(0)}\rangle$ is non-degenerate so it is a parity eigenket.

$$\underbrace{\langle K^{(0)} |}_{\langle K^{(0)} | \epsilon_K} \cancel{\pi^\dagger \pi z \pi^\dagger \pi} |K^{(0)}\rangle \underbrace{- z \epsilon_K |K^{(0)}\rangle}_{\langle K^{(0)} | Z | K^{(0)} \rangle} \rightarrow \text{never happen, so } \langle K^{(0)} | Z | K^{(0)} \rangle = 0.$$

$$\hookrightarrow \langle K^{(0)} | Z | K^{(0)} \rangle = -\langle K^{(0)} | Z | K^{(0)} \rangle \epsilon_K^2 \quad \text{parity-selection rule.}$$

$$\langle n' l' m' | Z | n l m \rangle = 0 \quad , \text{ here } z = T_{q=0}^{k=1}$$

unless $|j-k| \leq j' \leq j+k \Rightarrow j' = \pm 1$, $m' = q+m \rightarrow m$.

and parity: $\langle n' l' m' | Z | n l m \rangle \Rightarrow (-1)^{l'} (-1)^l = -1$, then $l'-l = \underline{\text{odd}}$.

Now consider ground state $|1, 0, 0\rangle$

$$\boxed{\alpha = -2e^2 \sum_{K=0}^{\infty} \frac{|\langle h^{(0)} | Z | 1, 0, 0 \rangle|^2}{E_0^{(0)} - E_K^{(0)}}}$$

dipole moment.

$$\vec{p} = \frac{\partial \epsilon}{\partial \vec{E}} = \alpha \vec{E}$$

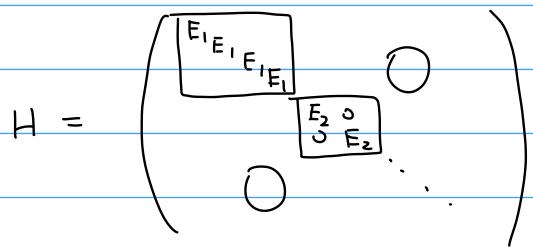
$\epsilon \sim$ wave of hydrogen

$$\text{applicability: } |V_{kn}| \ll |E_n^{(0)} - E_k^{(0)}|$$

Degenerate Perturbation Theory:

Since $|n^{(0)}\rangle \sim \sum_{k \neq n} \frac{V_{nk}}{E_n^{(0)} - E_k^{(0)}}$

Degenerate Spectrum:



→ If $V_{nk} \neq 0$, and $E_n^{(0)} = E_k^{(0)}$, i.e. degenerate, then we have trouble.

* Idea: choose zeroth order solution basis such that V has no off diagonal element.

i.e. let $V_{nk} = 0$ for $k \neq n$.

→ Suppose there is a g-fold degeneracy, before perturbation V.

Then there are g different eigenkets with eigenenergy $E_D^{(0)}$.

→ Let $|m^{(0)}\rangle$, $m=1, 2, 3, \dots, g$ to denote those eigenket of degenerate levels.

→ Define

$$|\ell^{(0)}\rangle = \sum_{m \in D} \langle m^{(0)} | \ell^{(0)} \rangle |m^{(0)}\rangle$$

↑ where $|\ell^{(0)}\rangle$ diagonalizes V

With result:

$$\Delta_1^{(1)} = \langle \ell^{(0)} | V | \ell^{(0)} \rangle = \langle m^{(0)} | V | m^{(0)} \rangle \left| \langle m^{(0)} | \ell^{(0)} \rangle \right|^2$$

$$\Delta_1^{(2)} = \sum_{k \notin D} \frac{|V_{ke}|^2}{E_D^{(0)} - E_k^{(0)}}$$

$$\Rightarrow \det[V - (E - E_D^{(0)})] = 0, \text{ i.e. } \boxed{\Delta_1^{(1)} = \text{eigenvalue of } \langle m^{(0)} | V | m^{(0)} \rangle}$$

Linear Stark Effect: $|n, l, m\rangle$

hydrogen atom: $a_0 = \frac{\pi^2}{m_e e^2}$: Bohr radius:

$$E = -\frac{e^2}{a_0} \frac{1}{2n^2}$$

For $|1, 0, 0\rangle \rightarrow E = -\frac{e^2}{2a_0}$

$\left. \begin{array}{l} |2, 0, 0\rangle \rightarrow E = -\frac{e^2}{8a_0} \\ |2, 1, \pm 1\rangle \rightarrow E = -\frac{e^2}{8a_0} \end{array} \right\}$ unperturbed state.

$$H_0 = \frac{p^2}{2m} - \frac{e^2}{r}, \text{ let } V = -ez|\vec{E}|$$

$V_0 \leftarrow z$ has non-vanishing element only between $l=0$ and $l=1$
also require m to be the same sign

$$z \text{ is } V_0^! \cdot \begin{matrix} 2S & 2p, m=0 & 2p, m=1 & 2p, m=-1 \end{matrix}$$

since z is odd, we contribution from $|l\rangle$ together to be odd.

so need $l=0$ and $l=1$

$$V = \begin{pmatrix} 0 & \langle 2S | V | 2p, m=0 \rangle & 0 & 0 \\ \langle 2p, m=0 | V | 2S \rangle & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$

$$\langle 2S | V | 2p, m=0 \rangle = 3ea_0 |\vec{E}|$$

$$V = 3ea_0 |\vec{E}| \left(\underbrace{\begin{pmatrix} \delta_x & 0 \\ 0 & 0 \end{pmatrix}}_{\text{matrix}} \right)$$

$$\lambda = +1, -1, 0, 0$$

$$\text{and } \Delta^{(1)} = \langle l^{(1)} | V | l^{(0)} \rangle = \text{eigenvalue } \{ \langle m^{(0)} | V | m^{(0)} \rangle \}$$

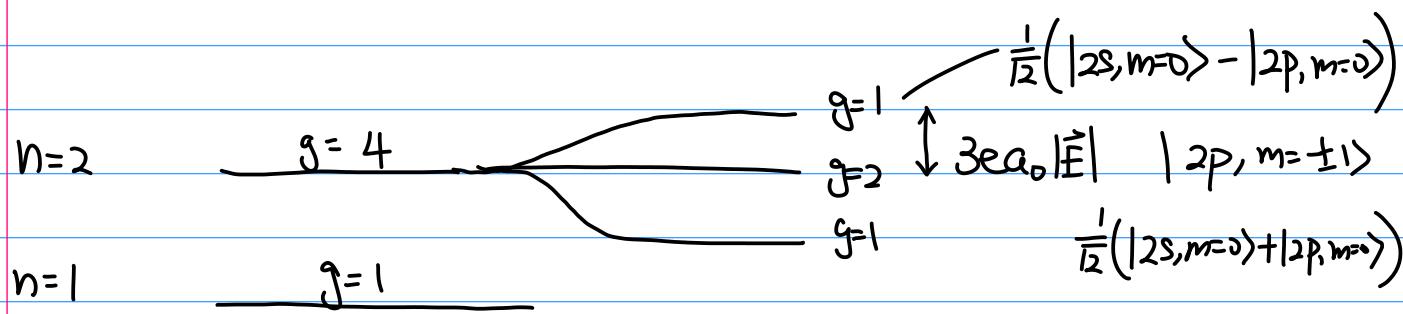
then

(1)

$$\Delta_{\pm} = \pm 3e |\vec{E}| a_0 \quad \leftarrow \text{linear stark effect.}$$

$$\Delta_0^{(1)} = 0$$

with $| \pm \rangle = \frac{1}{\sqrt{2}} (| 2s, m=0 \rangle \pm | 2p, m=0 \rangle)$



Fine Structure of atomic terms, Hydrogen atom:

$$H_0 = \frac{p^2}{2m} + V(r) , \text{ let } V(r) = -\frac{ze^2}{r}$$

$V(r)$ - central potential, rotational symmetry.

$$\left[-\frac{\hbar^2}{2m} \left(\frac{1}{r^2} \partial_r r^2 \partial_r - \frac{l^2}{r^2} \right) + V(r) \right] \psi = E \psi$$

$$\vec{l} = \frac{\vec{L}}{\hbar}$$

$$\text{and } l^2 = -\left(\frac{1}{\sin\theta} \partial_\theta \sin\theta \partial_\theta + \frac{1}{\sin^2\theta} \partial_\phi^2 \right)$$

$$\psi = R_{nl}(r) Y_l^m(\theta, \phi)$$

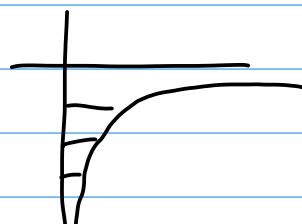
$$\hookrightarrow \left[-\frac{\hbar^2}{2m} \left(\frac{1}{r^2} \partial_r r^2 \partial_r - \frac{l(l+1)}{r^2} \right) + V(r) \right] R_{nl}(r) = E_{nl} R_{nl}(r)$$

$$V(r) = -\frac{ze^2}{r}, \text{ Hydrogen atom: } z = 1$$

$$\begin{aligned} \text{Typical kinetic energy: } & \frac{p^2}{2m} \sim \frac{1}{2m} \left(\frac{\hbar}{r_0} \right)^2 \\ \text{Typical potential energy: } & \sim -\frac{ze^2}{r_0} \end{aligned} \quad \left. \right\} r_0 \sim \frac{\hbar^2}{mze^2}$$

$$\text{then } E \sim -\frac{1}{2} \frac{me^4}{\hbar^2} z^2$$

$$\text{then For } z=1: E_{z=1} = -\frac{1}{2} \frac{me^4}{\hbar^2} \sim -13.5 \text{ eV}$$



$$r_0 = \frac{\hbar^2}{me^2} = a_0 \sim 0.529 \times 10^{-8} \text{ cm}$$

$$\text{with } z \neq 1, \Rightarrow e \rightarrow \sqrt{z} e \quad m \rightarrow m' = \frac{mm'}{m+z}$$

Now rewrite $E = -\frac{Z^2}{2n^2} \frac{me^4}{\hbar^2} = -\frac{Z^2}{2n^2} \frac{\hbar^2}{ma_0^2}$

$$r \rightarrow p = \frac{2Zr}{na_0}$$

then $\left(\frac{1}{p^2} \partial_p p^2 \partial_p - \frac{l(l+1)}{p^2} + \frac{n}{p} \right) R_{nl}(p) = \frac{1}{4} R_{nl}(p)$

as $p \rightarrow \infty$: $\partial_p^2 R_{nl} = \frac{1}{4} R_{nl} \Rightarrow R_{nl} \sim e^{-\frac{p}{2}}$

Here assume $R_{nl} \sim e^{-p}$

as $p \rightarrow 0$: $R_{nl} \sim \alpha(\alpha+1)p^{\alpha-2} - l(l+1)p^{\alpha-2} + np^{\alpha-1} = \frac{1}{4}p^{\alpha-2}$

assume $R_{nl} \sim p^\alpha \quad \downarrow \quad \underline{\alpha=1}, \quad -l \cancel{< 1}$
 as $p \rightarrow 0$

let $R_{nl} = p^l e^{-\frac{p}{2}} f_{nl}(p)$

$$\hookrightarrow [p \partial_p^2 + (\underbrace{2l+2-p}_{\beta}) \partial_p - \underbrace{(l+1-n)}_{2} f_{nl} = 0$$

then $F(\alpha, \beta; p)$ - confluent hypergeometric series.

when n : integer, it doesn't diverge anywhere.

so require n to be integer.

then $f_{nl} = L_{nl}^{2l+1}(p) = F(l+1-n, 2l+2, p)$ for n : integer
 $n \geq l+1$

↑ generalized Laguerre Polynomial.

if $n=3$:

$$\begin{array}{ll} \ell=2 & \rightarrow 5 \\ \ell=1 & \rightarrow 3 \\ \ell=0 & \rightarrow 1 \end{array} \quad \underbrace{\qquad}_{9}$$

so for n , we have n^2 degeneracy.

$$L_n^m(\rho) = (-1)^m \frac{n!}{(n-m)!} e^\rho \rho^{-m} \partial_\rho^{n-m} (e^{-\rho} \rho^n) \leftarrow \text{generalized Laguerre.}$$

$$L_n^m = L_n(\rho) = e^\rho \partial_\rho^n (e^{-\rho} \rho^n) \leftarrow \text{Laguerre Polynomial.}$$

Normalization:

$$\int e^{-\rho} \rho^{2\ell} \left[L_{n+\ell}^{2\ell+1}(\rho) \right]^2 \rho^2 d\rho = \frac{2n [(n+\ell)!]^3}{(n-\ell-1)!}$$

then

$$R_{nl}(\rho) = - \left[\left(\frac{2z}{na_0} \right)^3 - \frac{(n-\ell-1)!}{2n[(n+\ell)!]^3} \right]^{1/2} e^{-\frac{\rho}{2}} \rho^\ell L_{n+\ell}^{2\ell+1}(\rho)$$

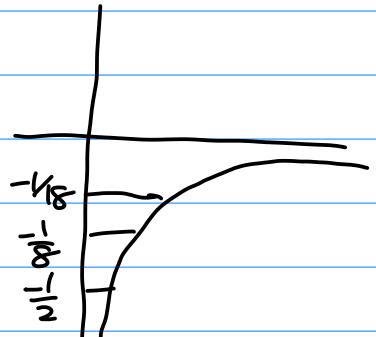
and $\psi_{nlm} = R_{nl}(\rho) Y_l^m(\theta, \phi) \rightarrow \text{require } n \geq \ell+1 \text{ and integer.}$

$$E_n = -\frac{1}{2n^2} \frac{z^2 e^2}{a_0} \quad \rho = \frac{2z}{na_0} r, \quad a_0 = \frac{\hbar^2}{me^2}$$

$$\text{Ex: } R_{10} = \left(\frac{z}{a_0} \right)^{3/2} 2 e^{-\frac{zr}{a_0}}$$

$$R_{20} = \left(\frac{z}{2a_0} \right)^{3/2} \left(2 - \frac{zr}{a_0} \right) e^{-\frac{zr}{2a_0}}$$

$$R_{21} = \left(\frac{z}{2a_0} \right)^{3/2} \frac{zr}{\sqrt{3}a_0} e^{-\frac{zr}{2a_0}}$$



<u>n</u>	<u>ℓ</u>	Name	<u>g</u>	without spin
1	0	1s	1	
2	0	2s	1	$\left. \begin{array}{l} 4 \text{ for energy} \\ 3 < m = \pm 1, 0 \end{array} \right\}$
	1	2p		
3	0	3s	1	$\left. \begin{array}{l} 9 \text{ for energy} \\ 5 \end{array} \right\}$
	1	3p	3	
	2	3d	5	

$$\langle r^k \rangle = \int_0^\infty dr r^2 R_m(r)^2 r^k \int \underbrace{\frac{dl}{l!} Y_l^m(\theta, \phi)}_{=1}$$

Sakurai S.3

$$\left. \begin{array}{l} \langle r \rangle = \int_0^\infty dr r^2 R_m(r)^2 r = \frac{a_0}{2Z} (3n^2 - \ell(\ell+1)) \\ \langle r^2 \rangle = \int_0^\infty dr r^2 R_m(r)^2 r^2 = \frac{a_0 n^2}{2Z^2} [5n^2 + 1 - 3\ell(\ell+1)] \\ \langle \frac{1}{r} \rangle = \frac{Z}{n^2 a_0} \\ \langle \frac{1}{r^2} \rangle = \frac{Z}{n^3 a_0^2 (\ell + \frac{1}{2})} \end{array} \right\}$$

Now consider spin: then we have $2n^2$ degeneracy.

Fine structure corrections:

fine structure

- 1) Electron move at $\sim \frac{c}{137}$, need relativistic corrections $\sim Z\alpha^2$
- 2) Radiative Correction: $\sim \text{lamb shift} \sim (Z\alpha)^2 \alpha \ln \frac{1}{z}$ small effect
- 3) external field.
- 4) Electron Interactions. (hard problem)
- 5) Nuclear spin \rightarrow cause magnetic field. \rightarrow Hyperfine Structure. $(Z\alpha)^2 \frac{m_e}{M_{\text{nuc}}}$

Focus on Relativistic Correction:

$$H_{\text{Dirac}} = c\vec{\alpha}(\vec{p} - \frac{e}{c}\vec{A}) + \beta mc^2 + e\phi$$

$\frac{1}{c}$ expansion $\xrightarrow{[Foldy-Wouthuysen \text{ Transformation}]}$

spectrum of H_{Dirac} .

$$E = \pm \sqrt{(mc^2)^2 + p^2 c^2}$$

$$\xrightarrow{} H = mc^2 + \frac{1}{2m} (\vec{p} - \frac{e}{c}\vec{A})^2 - \frac{p^4}{8m^3 c^2} + \theta$$

+ $e\phi - \frac{e\hbar}{2mc} \vec{\sigma} \cdot \vec{B}$ $- i \frac{e\hbar^2}{8m^2 c^2} \vec{\sigma} \cdot (\vec{\nabla} \times \vec{E}) - \frac{e\hbar}{4m^2 c^2} \vec{\sigma} \cdot (\vec{E} \times \vec{p}) - \frac{e\hbar^2}{8m^3 c^2} (\vec{\nabla} \cdot \vec{E})$

$\underbrace{\text{Zeeman term}}_{\text{ignore.}}$ $\underbrace{\text{Spin-orbital interaction}}_{\text{ignore.}}$ $\underbrace{\text{Darwin Term.}}_{\text{(ignore)}}$

Consider spin - orbit interaction:

$$V_{LS} = -\frac{e\hbar}{4m^2c^2} \vec{\sigma} \cdot (\vec{E} \times \vec{p}) = -\frac{e\hbar}{4m^2c^2} \left(-\frac{dV}{dr} \right) \frac{1}{r} \vec{\sigma} \cdot (\underbrace{\vec{r} \times \vec{p}}_{\vec{L}})$$

$$\vec{E} = -\vec{\nabla} V \frac{1}{e}$$

$$= -\frac{dV}{dr} \hat{r} \frac{1}{e}$$

$$= -\frac{dV}{dr} \frac{1}{r} \frac{1}{e} \hat{r}$$

$$\vec{E} = -\vec{\nabla} \phi$$

$$\vec{S} = \frac{\hbar}{2} \vec{\sigma}$$

$$\hookrightarrow \boxed{V_{LS} = \frac{1}{2m^2c^2} \frac{1}{r} \frac{dV}{dr} \vec{L} \cdot \vec{S}}$$

thomas precession.

Spin - Orbital Interaction:

$$H = \frac{\vec{p}^2}{2m} + V(r) \rightarrow |n, l, m_l\rangle$$

with spin $\rightarrow |n, l, m_l, s, m_s\rangle$

$$[H_{LS}, \vec{L}^i] \neq 0$$

$$\text{but } [H_{LS}, \underbrace{\vec{L} + \vec{S}}_{\vec{J}}] = 0 \quad \text{and} \quad \vec{L} \cdot \vec{S} = \frac{1}{2} [(l+s)^2 - l^2 - s^2] = \frac{1}{2} (j^2 - l^2 - s^2)$$

$$= \frac{1}{2} (j(j+1) - l(l+1) - s(s+1))$$

where $j = |l-s|, \dots, l+s$

$$\text{then } \vec{L} \cdot \vec{S} = \begin{cases} \frac{\hbar^2}{2} l & j = l + \frac{1}{2} \\ -\frac{\hbar^2}{2}(l+1) & j = l - \frac{1}{2} \end{cases}$$

So instead of $|n, l, m_l, s, m_s\rangle \rightarrow |n, l, s, j, m\rangle$

Since we're in basis that diagonalizes V_{LS} , we just use first order perturbation theory.

$$\begin{aligned}
 \text{then } E^{(1)} &= \langle n, l, s, j, m | V_{LS} | n, l, s, j, m \rangle \\
 &= \langle n, l, s, j, m | \frac{1}{2m^2c^2} \frac{1}{r} \frac{dV}{dr} [\vec{L} \cdot \vec{s}] | n, l, s, j, m \rangle \\
 &= \underbrace{\frac{1}{2m^2c^2} \int_0^\infty dr r^2 R_{nl} \frac{1}{r} \frac{dV}{dr}}_{\propto \frac{A_{nl}}{n^2}} \underbrace{\langle \vec{L} \cdot \vec{s} \rangle}_{S=\frac{1}{2}} \\
 &\quad j = l \pm \frac{1}{2} \\
 &= \frac{A_{nl}}{2} [j(j+1) - ((l+1) - s(s+1))]
 \end{aligned}$$

Let's find the difference in split.

$$\begin{aligned}
 E_{nj}^{(1)} - E_{n(l-j-1)}^{(1)} &= A_{nl} \frac{1}{2} [j(j+1) - (j-1)j] \\
 \boxed{E_{nj}^{(1)} - E_{n(l-j-1)}^{(1)} = A_{nl} j} &\quad \leftarrow \text{splitting energy.}
 \end{aligned}$$

$$\begin{aligned}
 A_{nl} &= \left\langle \frac{\hbar^2}{2m^2c^2} \frac{1}{r} \frac{dV}{dr} \right\rangle \sim \frac{e^2}{a_0^3} \frac{\hbar^3}{mc^2}, \quad a_0 = \frac{\hbar^2}{me^2} \\
 &\sim \frac{e^2}{a_0} \frac{\hbar^2}{ma_0^2} \frac{1}{mc^2} \sim \text{small.} \\
 \lambda = \frac{e}{\hbar c} &\quad \hookrightarrow A_{nl} \sim \frac{e^2}{a_0} \frac{e^4}{\hbar^2 c^2} \\
 &\quad \frac{1}{\lambda^2} = \left(\frac{e^2}{\hbar c}\right)^2 \sim \left(\frac{1}{137}\right)^2
 \end{aligned}$$

Scales:

energy: $E \sim \frac{e^2}{a_0^3} \frac{\hbar^2}{m_e c^2}$ or $\frac{e^2}{a_0} \alpha^2$

Length: $\underbrace{\frac{e^2}{m_e c^2}}_{\text{classical radius}}$ or $\underbrace{\frac{\hbar}{m_e c}}_{\text{compton radius}}$ or $\underbrace{a_0 = \frac{\hbar^2}{m_e c^2}}_{\text{bohr radius}} = 1 : \frac{1}{\alpha} : \frac{1}{\alpha^2} = 1 : 137 : 137^2$

estimate $\frac{p^4}{8m_e^3 c^2} \sim \alpha^2 \frac{e^2}{a_0}$

Zeeman Effect (magnetic field)

→ Consider at uniform magnetic field:

$$\vec{p} \rightarrow \vec{p} - \frac{e}{c} \vec{A} \quad , \quad \vec{\nabla} \cdot \vec{A} = \vec{B}$$

Now choose gauge: $\vec{A} = \frac{1}{2} (\vec{B} \times \vec{r})$
↑ radial gauge.

$$\Rightarrow H = \frac{\vec{p}^2}{2m} + V_c(r) - \frac{e}{2mc} (\vec{p} \cdot \vec{A} + \vec{A} \cdot \vec{p}) + \frac{e^2 \vec{B}^2}{2mc^2}$$

$$\text{choose } \vec{B} = B_0 \hat{z} \quad \text{then } \vec{A} = -\frac{1}{2} B_0 (-y \hat{x} - x \hat{y})$$

generally: $\vec{p} \cdot \vec{A} = \vec{A} \cdot \vec{p} - i\hbar (\vec{\nabla} \cdot \vec{A})$
 $\vec{\nabla} \cdot \vec{A} = 0$ in the this gauge.

$$\text{then } \vec{A} \cdot \vec{p} = -\frac{1}{2} B_0 (-y P_x - x P_y) = \frac{1}{2} B_0 L_z$$

$$\text{and } A^2 = \frac{1}{4} B_0^2 (x^2 + y^2)$$

Finally: $H = \frac{\vec{p}^2}{2m} + V_c(r) - \frac{e}{2mc} B_0 L_z + \underbrace{\frac{e^2}{8mc^2} B_0^2 (x^2 + y^2)}_{\propto B^2, \text{ drop}} - \vec{\mu} \cdot \vec{B}$

added spin
 magnetic-moment
 interaction

After omitting B^2 term: $\frac{-e}{mc} \vec{S} \cdot \vec{B} = \frac{-e}{mc} B S_z$

$$\text{let } H_0 = \frac{\vec{p}^2}{2m} + V_c(r)$$

$$H_{LS} = \frac{1}{2m_e c^2} \frac{1}{r} \frac{dV}{dr} \vec{L} \cdot \vec{S} \sim \frac{e^2}{a_0} \left(\frac{1}{137} \right)^2$$

$$H_B = \frac{-eB}{2mc} (L_z + 2S_z) \sim \frac{eB}{2mc}$$

In weak Magnetic field limit: if $H_B \ll H_{LS}$ (Lande Limit)
 strong Magnetic field limit: if $H_B \gg H_{LS}$ (Paschen-Bach)

	Dominant Interaction	Almost good (good for specific case)	No good at all	All case Quantum #s always good
$ j, l, s, m\rangle$	weak B (Lande)	H_{LS}	$J^2, \vec{L} \cdot \vec{S}$	L_z, S_z
$ \ell, s, m_l, m_s\rangle$	strong B \uparrow Paschen-Bach	H_B	L_z, S_z	$J^2, \vec{L} \cdot \vec{S}$

Consider Weak Magnetic limit ($H_B \ll H_{LS}$) (Lande Limit):

Good Quantum #: L^2, S^2, J_z }
 Almost good: J^2 } $|j, l, s, m_z\rangle$

$$\Delta E = -\frac{eB}{2m_ec} \langle J_z + S_z \rangle = -\frac{eB}{2m_ec} (tm + \langle S_z \rangle_{j=1 \pm \frac{1}{2}})$$

$$|j=1 \pm \frac{1}{2}, m\rangle = \pm \sqrt{\frac{1+m+\frac{1}{2}}{2l+1}} |m_l=m-\frac{1}{2}, m_s=\frac{1}{2}\rangle$$

$$C_+ \quad + \sqrt{\frac{1+m+\frac{1}{2}}{2l+1}} |m_l=m+\frac{1}{2}, m_s=-\frac{1}{2}\rangle$$

$$\begin{aligned} \langle S_z \rangle &= \frac{\hbar}{2} (|C_+|^2 - |C_-|^2) \\ &= \frac{\hbar}{2} \frac{1}{2l+1} [(1 \pm m + \frac{1}{2}) - (1 \mp m + \frac{1}{2})] = \pm \frac{m\hbar}{2l+1} \end{aligned}$$

$$\boxed{\Delta E_B = -\frac{e\hbar B}{2m_ec} m \left[1 \pm \frac{1}{2l+1} \right]}$$

← energy shift. (Lande formula)

Another Derivation:

assume $\vec{S} = \alpha \vec{J}$
 ↘ constant

then $S_z = \alpha S_z$

then $\underbrace{\vec{S} \cdot \vec{J}}_{\text{same constant}} = \alpha J^2$

Find $\vec{J} \cdot \vec{J}$:

$$(\vec{J} - \vec{S})^2 = \vec{L}^2 \quad \hookrightarrow -2\vec{J} \cdot \vec{S} + J(J+1) + S(S+1) = l(l+1)$$

$$\hookrightarrow -2\vec{J} \cdot \vec{S} + J(J+1) + S(S+1) = l(l+1)$$

$$\hookrightarrow \vec{J} \cdot \vec{S} = \frac{j(j+1) - l(l+1) + s(s+1)}{2} = \alpha j(j+1)$$

$$\hookrightarrow \alpha = \frac{j(j+1) - l(l+1) + s(s+1)}{2 J(j+1)}$$

so $S_z = \alpha \hbar m$

introduce: Lande factor : $g = 1 + \frac{j(j+1) - l(l+1) + s(s+1)}{2 j(j+1)}$

if $s = \frac{1}{2}, j = l \pm \frac{1}{2} \rightarrow g = 1 \pm \frac{1}{2l+1}$

then

$$\boxed{\Delta E_B = \frac{-e\hbar B}{2m_e c} \left[1 \pm \frac{1}{2l+1} \right]}$$

↙ same answer as before.

Time - Dependent Potential:

Consider Hamiltonian, H such that it can be split into two parts

$$H = H_0 + V(t)$$

part that doesn't contain time variation
contain time variation

if only H_0 : $|\alpha(t)\rangle = e^{-\frac{iH_0 t}{\hbar}} |\alpha(t=0)\rangle$

Suppose at $t=0$, we have $|\alpha\rangle = \sum_n c_n(0) |n\rangle$

We wish to find at later times.

$$|\alpha, t_0=0; t\rangle_S = \sum_n c_n(t) e^{-\frac{iE_n t}{\hbar}} |n\rangle$$

Interaction picture gets rid of this

Interaction Picture.

Define

$$|\alpha, t_0; t\rangle_I = e^{iH_0 t/\hbar} |\alpha, t_0; t\rangle_S = e^{\frac{iH_0 t}{\hbar} - \frac{iH t}{\hbar}} |\alpha, t_0; t\rangle_H$$

↑
Interaction Picture. ↑ Schrödinger Picture ↑ Heisenberg Picture.

For operators (observables):

$$A_I = e^{\frac{iH_0 t}{\hbar}} A_S e^{-\frac{iH_0 t}{\hbar}}$$

In particular: $V_I = e^{\frac{iH_0 t}{\hbar}} V e^{-\frac{iH_0 t}{\hbar}}$

↑ time-dependent potential in
Schrödinger picture.

Recall Heisenberg: $|\alpha\rangle_H = e^{\frac{iH t}{\hbar}} |\alpha, t_0=0; t\rangle_S$

$$A_H = e^{\frac{iH t}{\hbar}} A_S e^{-\frac{iH t}{\hbar}}$$

Now let's derive the fundamental differential equation that characterizes time evolution of state ket in interaction picture:

$$\begin{aligned}
 i\hbar \frac{dt}{i\hbar} |\alpha, t_0; t\rangle_I &= i\hbar \frac{dt}{i\hbar} (e^{iH_0 t/\hbar} |\alpha, t_0; t\rangle_S) \\
 &= -H_0 e^{\frac{iH_0 t}{\hbar}} |\alpha, t_0; t\rangle_S + e^{\frac{iH_0 t}{\hbar}} \frac{dt}{i\hbar} |\alpha, t_0; t\rangle_S \\
 &= (H_0 + V) |\alpha, t_0; t\rangle_S \\
 &= e^{\frac{iH_0 t}{\hbar}} V |\alpha, t_0; t\rangle_S
 \end{aligned}$$

$$i\hbar \frac{dt}{i\hbar} |\alpha, t_0; t\rangle_I = e^{\frac{iH_0 t}{\hbar}} V e^{-\frac{iH_0 t}{\hbar}} |\alpha, t_0; t\rangle_I$$

$$\therefore i\hbar \frac{dt}{i\hbar} |\alpha, t_0; t\rangle_I = V_I |\alpha, t_0; t\rangle_I$$

Therefore, $|\alpha, t_0; t\rangle_I$ is constant if $V_I = 0$

We can also show for observable A, (that doesn't contain time explicitly in Schrodinger) that:

$$\frac{dA_I}{dt} = \frac{1}{i\hbar} [A_I, H_0]$$

\leftarrow Heisenberg-like, but replaced $H \rightarrow H_0$

Table Summary:

	Heisenberg	Interaction	Schrodinger
State ket	No change	determined by V_I	determined by H
Observable	determined by H	determined by H_0	No change.

$$\text{Let } |\alpha; t_0; t\rangle_I = \sum_n C_n(t) |n\rangle$$

$$\text{then } i\hbar \frac{d}{dt} \underbrace{\langle n | \alpha, t_0; t \rangle}_I = \sum_m \underbrace{\langle n | V_I | m \rangle}_{C_m(t)} \underbrace{\langle m | \alpha, t_0; t \rangle}_I$$

$$= C_n(t) \langle n | e^{\frac{iH_0 t}{\hbar}} V e^{-\frac{iH_0 t}{\hbar}} | m \rangle$$

System of ODEs
to solve

$$\hookrightarrow = V_{nm}(t) e^{\frac{i(E_n - E_m)}{\hbar} t}$$

**

$$i\hbar \frac{d}{dt} C_n(t) = \sum_m V_{nm} e^{i\omega_{nm} t} C_m(t)$$

$$\text{where } \omega_{nm} = \frac{E_n - E_m}{\hbar}$$

$$\hookrightarrow i\hbar \begin{pmatrix} \dot{c}_1 \\ c_2 \\ \vdots \end{pmatrix} = \begin{pmatrix} V_{11} & V_{12} e^{i\omega_{12} t} & \dots & \dots \\ V_{21} e^{i\omega_{21} t} & V_{22} & V_{32} & \dots \\ \vdots & & \ddots & \ddots \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \\ \vdots \end{pmatrix}$$

Time-Dependent Two-state problem:

Problem with exact solution: sinusoidal oscillating potential.

$$H_0 = E_1 |1\rangle\langle 1| + E_2 |2\rangle\langle 2| \quad (E_2 > E_1)$$

$$V(t) = \gamma e^{i\omega t} |1\rangle\langle 2| + \gamma e^{-i\omega t} |2\rangle\langle 1|$$

then $H = \begin{pmatrix} E_1 & \gamma e^{i\omega t} \\ \gamma e^{-i\omega t} & E_2 \end{pmatrix}$ with initial condition $C_1(0)=1$
 $C_2(0)=0$

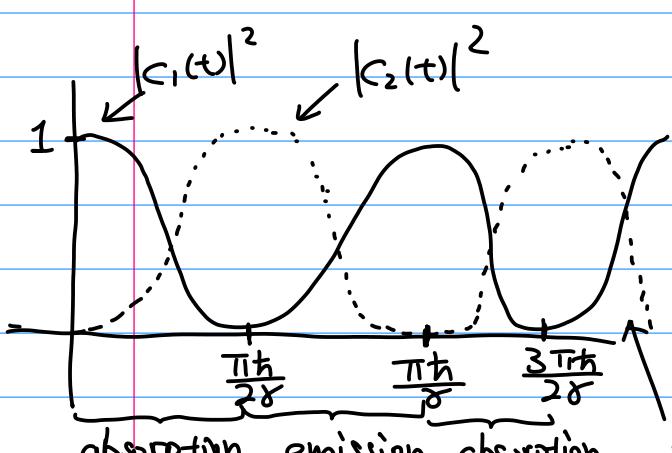
then we have:

$$i\hbar \begin{pmatrix} \dot{C}_1 \\ \dot{C}_2 \end{pmatrix} = \begin{pmatrix} 0 & \gamma e^{i\omega t} \\ \gamma e^{-i\omega t} & 0 \end{pmatrix} \begin{pmatrix} C_1 \\ C_2 \end{pmatrix}$$

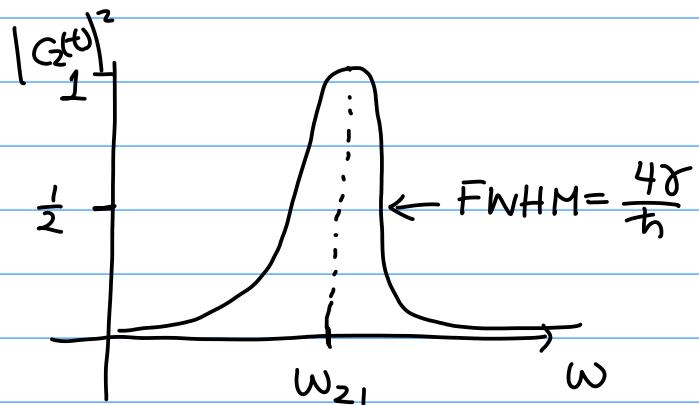
This has solution:

probability of finding upper state $\Rightarrow |C_2(t)|^2 = \frac{\gamma^2/\hbar^2}{\gamma^2/\hbar^2 + (\omega - \omega_{21})^2/4} \sin^2 \left\{ \sqrt{\frac{\gamma^2}{\hbar^2} + \frac{(\omega - \omega_{21})^2}{4}} t \right\}$

probability of finding lower state $\Rightarrow |C_1(t)|^2 = 1 - |C_2(t)|^2$ resonance when $\omega = \omega_{21}$



(higher $|C_2(t)|^2$) (lower $|C_1(t)|^2$)
 so absorb to upper state emits photon to lower state
 when $\omega = \omega_{21}$ otherwise $|C_2(t)|^2_{\text{max}} \neq 1$



Time-Dependent Perturbation Theory:

Dyson Series:

Previously we derived: $i\hbar \partial_t |\alpha\rangle_I = V_I |\alpha\rangle_I$

$$\text{where } V_I = e^{\frac{i}{\hbar} H_0 t} V_0 e^{-\frac{i}{\hbar} H_0 t}$$

$$|\alpha, t\rangle_I = U_I(t, t_0) |\alpha, t=t_0\rangle$$

↑
time evolution operator in interaction picture.

$$\text{then } i\hbar \partial_t U_I(t, t_0) |\alpha, t=t_0\rangle = V_I U_I(t, t_0) |\alpha, t=t_0\rangle$$

$$\hookrightarrow ① \quad i\hbar \partial_t U_I(t, t_0) = V_I(t) U_I(t, t_0)$$

$$\text{since } |\alpha, t=t_0\rangle_I = U_I(t=t_0, t_0) |\alpha, t=t_0\rangle_I$$

$$\hookrightarrow ② \text{ require } U_I(t, t_0) \Big|_{t=t_0} = 1 \text{ as initial condition.}$$

$$\text{using } ① : \quad dt \left[\underbrace{\frac{d}{dt} U_I(t', t_0)}_{U_I(t, t_0)} \right] \Bigg|_{t_0}^t = \int_{t_0}^t -\frac{i}{\hbar} V_I(t') U_I(t', t_0) dt'$$

$$\hookrightarrow U_I(t, t_0) - 1 = \int_{t_0}^t -\frac{i}{\hbar} V_I(t') U_I(t', t_0) dt'$$

$$\hookrightarrow \boxed{U_I(t, t_0) = 1 - \frac{i}{\hbar} \int_{t_0}^t V_I(t') U_I(t', t_0) dt'}$$

Solve above equation via iteration:

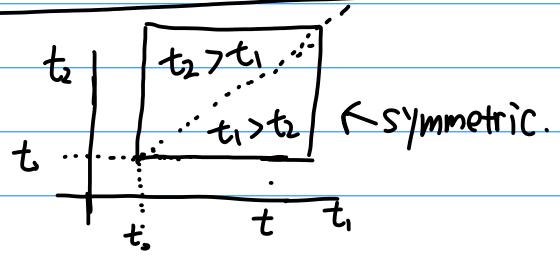
~~Dyson Series~~

$$U_I(t, t_0) = 1 - \frac{i}{\hbar} \int_{t_0}^t dt' V_I(t') + \left(\frac{-i}{\hbar}\right)^2 \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' V_I(t') V_I(t'')$$

$$+ \dots + \left(\frac{-i}{\hbar}\right)^n \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' \dots \int_{t_0}^{t^{(n-1)}} dt^{(n)} V_I(t') V_I(t'') \dots V_I(t^{(n)})$$

with requirement: $t > t' > t'' > \dots > t^{(n-1)} > t^{(n)} > t_0$

Rewriting in a compact form



then $U_I(t, t_0) = \sum_{n=0}^{\infty} \hat{T}_t \frac{1}{n!} \left(\int_{t_0}^t dt_1 V_I(t_1) \right)^n$

↓ time ordering operator that orders
time into $t > t_1 > t_2 \dots$

$$= \hat{T}_t \exp \left\{ \int_{t_0}^t dt_1 V_I(t_1) \right\}$$

Transition Probability

Suppose we're given

$$|i, t_0=0; t\rangle_I = U_I(t, 0) |i, t_0=0; t=t_0\rangle \quad \downarrow \text{initial state}$$

$$= \sum_n |n\rangle \underbrace{\langle n | U_I(t, 0)}_{C_n(t)} |i\rangle$$

$$|i, t_0=0; t\rangle_I = \sum_n C_n(t) |n\rangle, \quad C_n(t) = \langle n | U_I(t, 0) | i \rangle$$

Now explore connection between $U(t, t_0)$ and $U_I(t, t_0)$

$$|\alpha, t_0; t\rangle_I = e^{\frac{iH_0t}{\hbar}} |\alpha, t_0; t\rangle_S$$

$$= e^{\frac{iH_0t}{\hbar}} U(t, t_0) |\alpha, t_0; t=t_0\rangle_S$$

$$|\alpha, t_0; t\rangle_I = e^{\frac{iH_0t}{\hbar}} \underbrace{U_S(t, t_0) e^{-\frac{iH_0t}{\hbar}}}_{U_I(t, t_0)} |\alpha, t_0; t=t_0\rangle_I$$

So
$$U_I(t, t_0) = e^{\frac{iH_0t}{\hbar}} U_S(t, t_0) e^{-\frac{iH_0t}{\hbar}}$$

So $\langle n | U_I(t, t_0) | i \rangle = \langle n | e^{\frac{iH_0t}{\hbar}} U_S(t, t_0) e^{-\frac{iH_0t}{\hbar}} | i \rangle$

$$\langle n | U_I(t, t_0) | i \rangle = e^{\frac{i(E_i t - E_i t_0)}{\hbar}} \underbrace{\langle n | U_S(t, t_0) | i \rangle}_{\text{transition amplitude.}}$$

so we see that $\langle n | V_I(t, t_0) | i \rangle \neq \langle n | V_S(t, t_0) | i \rangle$

However if $|n\rangle$ and $|i\rangle$ are both energy eigenstates:

then $|\langle n | V_I(t, t_0) | i \rangle|^2 = |\langle n | V_S(t, t_0) | i \rangle|^2$

but if we use $|a'\rangle$ and $|b'\rangle$, which are eigenstates of operator A, B, but $[H, A] \neq 0$ $[H, B] \neq 0$, i.e. not simultaneous eigenket of H, then in general,

$$|\langle b' | V_I(t, t_0) | a' \rangle|^2 \neq |\langle b' | V_S(t, t_0) | a' \rangle|^2$$

*

Determine Transition Probability: $P(i \rightarrow n) = \sum_{n+i} |C_n^{(1)}(t) + C_n^{(2)}(t) + \dots|^2$

Determining: $C_n(t) = \langle n | V_I(t, t_0) | i \rangle$

$$C_n(t) = C_n^{(0)}(t) + C_n^{(1)}(t) + C_n^{(2)}(t) + \dots$$

$$= \langle n | V_I(t, t_0) | i \rangle$$

$$= \langle n | 1 - \frac{i}{\hbar} \int_{t_0}^t dt' V_I(t') + \left(\frac{-i}{\hbar}\right)^2 \int_{t_0}^t dt' \int_{t'}^t dt'' V_I(t') V_I(t'') + \dots | i \rangle$$

$$= \langle n | i \rangle - \frac{i}{\hbar} \int_{t_0}^t \langle n | V_I(t') | i \rangle dt' \quad \text{note: } V_I = e^{\frac{iH_0 t}{\hbar}} V_S e^{\frac{-iH_0 t}{\hbar}}$$

$$+ \left(\frac{-i}{\hbar}\right)^2 \int_{t_0}^t dt' \int_{t'}^t dt'' \langle n | V_I(t') V_I(t'') | i \rangle$$

$$C_n(t) = \delta_{ni} - \frac{i}{\hbar} \int_{t_0}^t e^{iW_{ni} t'} V_{ni}(t') dt'$$

$$+ \left(\frac{-i}{\hbar}\right)^2 \sum_m \int_{t_0}^t dt' \int_{t'}^t dt'' e^{iW_{nm} t'} V_{nm}(t') e^{iW_{mi} t''} V_{mi}(t'') + \dots$$

assume
 n, i are
eigenkets
of H

Ex: Constant Perturbation:

$$V(t) = \begin{cases} 0 & \text{for } t < 0 \\ V & \text{for } t \geq 0 \end{cases}$$

assume no time dependence, but made up of operators like \hat{x} , \hat{p} and \hat{s} .

SUPPOSE at $t_0 = 0$, we're in eigenket $|i\rangle$.

Then $C_n^{(0)} = \delta_{ni}$

$$C_n^{(1)} = -\frac{i}{\hbar} \int_{t_0}^t e^{i\omega_n t'} V_{ni}(t') dt'$$

but $V_{ni}(t') = V_{ni}$

$$\begin{aligned} &= -\frac{i}{\hbar} V_{ni} \int_0^t e^{i\omega_n t'} dt' \\ &= -\frac{V_{ni}}{\hbar \omega_n} e^{i\omega_n t'} \Big|_0^t \end{aligned}$$

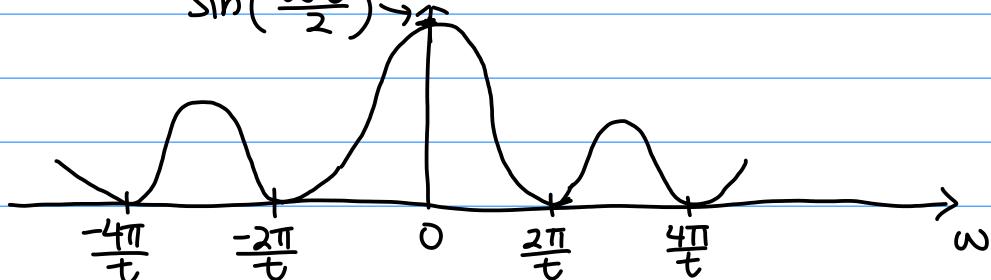
$$C_n^{(1)} = \frac{V_{ni}}{E_n - E_i} \left(1 - e^{i\omega_n t} \right)$$

Probability to go to other state is:

$$P_{i \rightarrow n} \approx |C_n^{(1)}|^2 = \frac{|V_{ni}|^2}{|E_n - E_i|^2} (2 - 2 \cos \omega_n t)$$

$$P_{i \rightarrow n} = \frac{4 |V_{ni}|^2}{|E_n - E_i|^2} \sin^2 \left[\frac{(E_n - E_i)t}{2\hbar} \right]$$

$$\sin^2 \left(\frac{\omega t}{2} \right) \rightarrow$$



As $E_n \rightarrow E_i$:

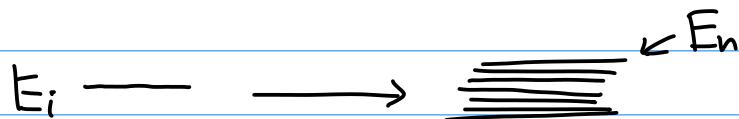
$$P_{i \rightarrow n} = \frac{4|V_{ni}|^2}{|E_n - E_i|^2} \sin^2\left(\frac{(E_n - E_i)t}{2\hbar}\right)$$

$$= \frac{4|V_{ni}|^2}{(\hbar\omega_{ni})^2} \sin^2\left(\frac{\omega_{ni}t}{2}\right)$$

take limit $E_n \rightarrow E_i$, $\omega_{ni} \rightarrow 0$, so $\sin^2\left(\frac{\omega t}{2}\right) \sim \left(\frac{\omega}{2}t\right)^2$

then $P_{i \rightarrow n} = |C_n^{(i)}|^2 \approx \frac{1}{\hbar^2} |V_{ni}|^2 t^2$

Since $E_n \sim E_i$, the final states form a continuous energy spectrum near E_i :



⇒ If there are many of those states, we want to calculate the total probability, that is, the transition probability summed over final states with $E_n \approx E_i$.

or $P_{i \rightarrow n, E_n \approx E_i} = \sum_{n, E_n \approx E_i} |C_n^{(i)}|^2$

Σ sum over all states with their energy $\approx E_i$

Introduce density of final states with energy interval, $(E, E + dE)$:

$$\rightarrow \rho(E) dE$$

$$\hookrightarrow \sum_{n, E_n \approx E_i} |C_n^{(1)}|^2 \Rightarrow \int dE_n \rho(E_n) |C_n^{(1)}|^2 \\ = 4 \int \sin^2 \left[\frac{(E_n - E_i)t}{2\hbar} \right] \frac{|\psi_{ni}|^2}{|E_n - E_i|^2} \rho(E_n) dE_n$$

as $t \rightarrow \infty$, using $\lim_{x \rightarrow \infty} \frac{1}{\pi} \frac{\sin^2 x}{x^2} = \delta(x)$

$$\lim_{t \rightarrow \infty} \frac{1}{|E_n - E_i|^2} \sin^2 \left(\frac{(E_n - E_i)t}{2\hbar} \right) = \frac{\pi t}{2\hbar} \delta(E_n - E_i)$$

$\xrightarrow{t \rightarrow \infty}$ take average of $|\psi_{ni}|^2$

$$4 \int \rho(E_n) \frac{\pi t}{2\hbar} \delta(E_n - E_i) |\psi_{ni}|^2 dE_n$$

$$\lim_{t \rightarrow \infty} \sum_{n, E_n \approx E_i} |C_n^{(1)}|^2 = \frac{2\pi t}{\hbar} |\psi_{ni}|^2 \rho(E_n) \Big|_{E_n \approx E_i}$$

Now consider

$$\text{Transition rate: } W_{i \rightarrow [n]} = \frac{d}{dt} P_{i \rightarrow [n]} = \frac{d}{dt} \sum_n |C_n^{(i)}|^2$$

$[n]$ means a group of states

using

$$\lim_{t \rightarrow \infty} \sum_{n, E_n \approx E_i} |C_n^{(i)}|^2 = \frac{2\pi t}{\hbar} |\langle V_{ni} \rangle|^2 f(E_n) \Big|_{E_n \approx E_i}$$

then $W_{i \rightarrow [n]} = \frac{d}{dt} \sum_n |C_n^{(i)}|^2 = \frac{2\pi}{\hbar} |\langle V_{ni} \rangle|^2 f(E_n) \Big|_{E_n \approx E_i}$

constant rate, instead of oscillating

$$\hookrightarrow W_{i \rightarrow n} = \frac{2\pi}{\hbar} |\langle V_{ni} \rangle|^2 \delta(E_n - E_i)$$

switched to a particular state, n .

← Fermi's Golden Rule

and note $W_{i \rightarrow [n]} = \int dE_n f(E_n) W_{i \rightarrow n}$

Now consider 2nd order:

$$C_n^{(2)}(t) = \left(\frac{-i}{\hbar}\right)^2 \sum_m V_{nm} V_{mi} \int_0^t dt' e^{i\omega_{nm} t'} \int_0^{t'} dt'' e^{i\omega_{ni} t''}$$

$$= \frac{i}{\hbar} \sum_m \frac{V_{nm} V_{mi}}{E_m - E_i} \int_0^t (e^{i\omega_{ni} t'} - e^{i\omega_{nm} t'}) dt$$

Same as $C_n^{(1)}$
 and $\sim t$ as $t \rightarrow \infty$

 \uparrow
 \uparrow
 gives rise to
 rapid oscillation
 as $t \rightarrow \infty$, so it
 doesn't grow in t ,
 then don't care.

Now with $C_n^{(1)}(t)$ and $C_n^{(2)}(t)$:

$$\omega_{i \rightarrow [n]} = \frac{2\pi}{\hbar} \overline{\left| V_{ni} + \sum_m \frac{V_{nm} V_{mi}}{E_i - E_m} \right|^2} \rho(E_n) \Big|_{E_n \approx E_i}$$

Harmonic Perturbation:

$$V(t) = V e^{i\omega t} + V^+ e^{-i\omega t} \quad t \geq 0$$

Assume 1 eigenstate of H_0 is populated initially,
assume perturbation is on at $t=0$.

then

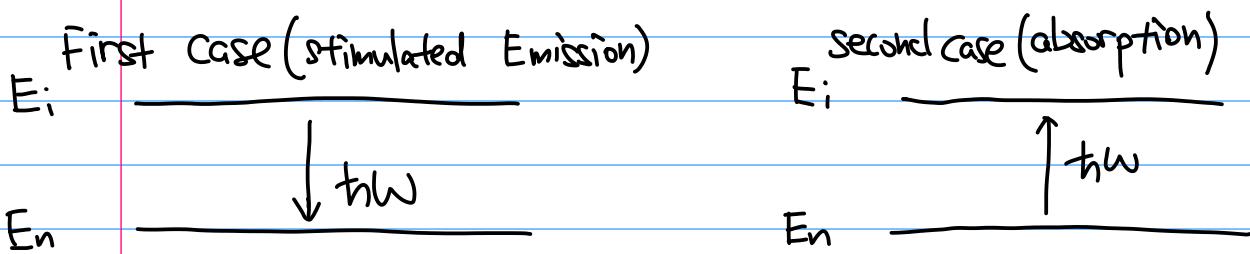
$$\begin{aligned} C_n^{(1)}(t) &= \frac{-i}{\hbar} \int_0^t (V_{ni} e^{i\omega t} + V_{ni}^+ e^{-i\omega t}) e^{i\omega_n t'} dt' \\ &= \frac{1}{\hbar} \left[\frac{1 - e^{i(\omega_n - \omega)t}}{\omega + \omega_n} V_{ni} + \frac{1 - e^{i(\omega_n - \omega)t'}}{-\omega + \omega_n} V_{ni}^+ \right] \end{aligned}$$

We see as $t \rightarrow \infty$: $|C_n^{(1)}|^2$ valid when:

For first term: $\omega_n + \omega \approx 0$ or $E_n \approx E_i - \hbar\omega$

For second term: $\omega_n - \omega \approx 0$ or $E_n \approx E_i + \hbar\omega$

When first term is important, second term is not. Vice-versa.



then the transition rates are:

$$W_{i \rightarrow [n]} = \frac{2\pi}{\hbar} \overline{|V_{ni}|^2} f(E_n) \Big|_{E_n \approx E_i - \hbar\omega}$$

(stimulated Emission)
First case.

$$W_{i \rightarrow [n]} = \frac{2\pi}{\hbar} \overline{|V_{ni}^+|^2} f(E_n) \Big|_{E_n \approx E_i + \hbar\omega}$$

(Absorption)
Second case.

More commonly:

$$W_{i \rightarrow n} = \frac{2\pi}{\hbar} \left\{ \frac{|V_{ni}|^2}{|V_{ni+}^+|^2} \right\} \delta(E_n - E_i \pm \hbar\omega)$$

Note since $|V_{ni}|^2 = |V_{ni+}^+|^2$

$$\frac{W_{i \rightarrow [n]}}{P(E_n)} = \frac{W_{n \rightarrow [i]}}{P(E_i)} \quad \leftarrow \text{detailed balance}$$

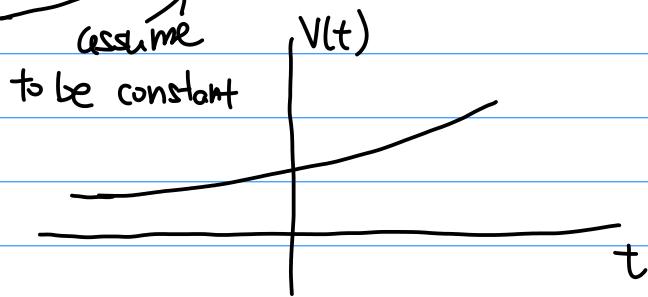
Energy Shift and Decay Width:

Question: What happens to $C_i(t)$ itself, i.e. $i \rightarrow n=i$.

To avoid effect of sudden change in Hamiltonian,
let's increase perturbation very slowly.

$$V(t) = e^{\eta t} V, \quad \text{so } V(t) = 0 \text{ as } t=t_0 = -\infty$$

let $\eta \rightarrow 0$
in the end
to have
constant potential.



let's first calculate $i \rightarrow n \neq i$:

$$\begin{aligned} C_n^{(0)}(t) &= 0 \\ C_n^{(1)}(t) &= \frac{-i}{\hbar} V_{ni} \lim_{t_0 \rightarrow -\infty} \int_{t_0}^t e^{\eta t'} e^{i\omega_{ni} t'} dt' \\ &= \frac{-i}{\hbar} V_{ni} \frac{e^{\eta t + i\omega_{ni} t}}{\eta + i\omega_{ni}} \end{aligned}$$

then: $|C_n^{(1)}(t)|^2 \simeq \frac{|V_{ni}|^2}{\hbar^2} \frac{e^{2\eta t}}{\eta^2 + \omega_{ni}^2}$

then the transition rate:

$$\gamma_{i \rightarrow n} = \frac{d}{dt} |C_n^{(1)}(t)|^2 \simeq \frac{2|V_{ni}|^2}{\hbar^2} \left(\frac{\eta e^{2\eta t}}{\eta^2 + \omega_{ni}^2} \right)$$

Now let $\eta \rightarrow 0^+$,

$$\text{but note: } \lim_{\eta \rightarrow 0^+} \frac{\eta}{\eta^2 + \omega_{ni}^2} = \pi \delta(\omega_{ni}) = \pi \delta\left(\frac{E_n - E_i}{\hbar}\right) = \pi \hbar \delta(E_n - E_i)$$

then

$$W_{i \rightarrow n} \underset{\text{again, the golden rule.}}{\approx} \left(\frac{2\pi}{\hbar}\right) |V_{ni}|^2 \delta(E_n - E_i)$$

Now lets consider case $n = i$:

$$C_i^{(0)} = 1$$

$$C_i^{(1)} = \frac{-i}{\hbar} V_{ii} \lim_{t_0 \rightarrow -\infty} \int_{t_0}^t e^{\eta t'} dt' = \frac{-i}{\eta \hbar} V_{ii} e^{\eta t}$$

$$\begin{aligned} C_i^{(2)} &= \left(\frac{-i}{\hbar}\right)^2 \sum_m |V_{mi}|^2 \lim_{t_0 \rightarrow -\infty} \int_{t_0}^t dt' e^{i\omega_{mi}t' + \eta t'} \frac{e^{i\omega_{mi}t' + \eta t}}{i(\omega_{mi} - i\eta)} \\ &\stackrel{\substack{\text{separate} \\ \text{to } m=i \\ \text{and } m \neq i}}{=} \left(\frac{-i}{\hbar}\right)^2 |V_{ii}|^2 \frac{e^{2\eta t}}{2\eta^2} + \left(\frac{-i}{\hbar}\right) \sum_{m \neq i} \frac{|V_{mi}|^2 e^{2\eta t}}{2\eta(E_i - E_m + i\hbar\eta)} \end{aligned}$$

Now evaluate up to 2nd order:

$$\begin{aligned} \frac{\dot{C}_i}{C_i} &\approx \frac{-i}{\hbar} V_{ii} + \left(\frac{-i}{\hbar}\right)^2 \frac{|V_{ii}|^2}{\eta} + \left(\frac{-i}{\hbar}\right) \sum_{m \neq i} \frac{|V_{mi}|^2}{(E_i - E_m + i\hbar\eta)} \\ &\quad \downarrow \frac{1 - \frac{i}{\hbar} \frac{V_{ii}}{\eta}}{} \\ &\approx \frac{-i}{\hbar} V_{ii} + \left(\frac{-i}{\hbar}\right) \sum_{m \neq i} \frac{|V_{mi}|^2}{E_i - E_m + i\hbar\eta} \end{aligned}$$

Now try ansatz: $C_i(t) = e^{\frac{-i}{\hbar} \Delta_i t}$ and $\frac{\dot{C}_i(t)}{C_i(t)} = \frac{-i}{\hbar} \Delta_i$

Now in if in interaction picture:

$$C_i(t)_{\text{I}} = e^{\frac{-i}{\hbar} \Delta_i t}$$

then in Schrödinger picture we have:

$$C_i(t)_S = e^{\frac{-i}{\hbar} \Delta_i t} e^{\frac{-i}{\hbar} E_i t} = e^{\frac{-i}{\hbar} (E_i + \Delta_i) t} \quad \leftarrow \text{we see it's shifting by } \Delta_i$$

$$\text{or } E_i \rightarrow E_i + \Delta_i$$

If we let $\Delta_i = \Delta_i^{(1)} + \Delta_i^{(2)} + \dots$

then $\Delta_i^{(1)} = V_{ii}$ \leftarrow exactly time-independent perturbation theory.

To find $\Delta_i^{(2)}$, first note: $\lim_{\epsilon \rightarrow 0} \frac{1}{x+i\epsilon} = \text{Pr}\{\frac{1}{x}\} - i\pi\delta(x)$

Thus

$$\Delta_i^{(2)} = \underbrace{\text{Pr}\left\{\sum_{m \neq i} \frac{|V_{mi}|^2}{E_i - E_m}\right\}}_{\text{exactly 2nd order time-independent perturbation theory.}} - i\pi \underbrace{\sum_{m \neq i} |V_{mi}|^2}_{\frac{\hbar}{2\pi} \sum_{m \neq i} W_{i \rightarrow m}} \delta(E_i - E_m)$$

Note $\sum_{m \neq i} W_{i \rightarrow m} = \frac{2\pi}{\hbar} \sum_{m \neq i} |V_{mi}|^2 \delta(E_i - E_m) = -\frac{2}{\hbar} \text{Im}(\Delta_i^{(2)})$

In the end we have:

$$C_i(t) = e^{\underbrace{-\frac{i}{\hbar} \text{Re}(\Delta_i)t}_{\text{shift in energy}}} e^{\underbrace{\frac{-i}{\hbar} \text{Im}(\Delta_i)t}_{\text{decay}}} \\ E_i \rightarrow E_i + \text{Re}(\Delta_i)$$

if we define $\frac{\Gamma_i}{\hbar} = -\frac{2}{\hbar} \text{Im}(\Delta_i) < \text{life time}$.

$$\text{then } |C_i(t)|^2 = e^{\frac{2 \text{Im}(\Delta_i)t}{\hbar}} = e^{-\frac{\Gamma_i t}{\hbar}}$$

define decay width: $\Gamma_i = \frac{\hbar}{\Gamma_i}$

Also note probability conservation: for small t

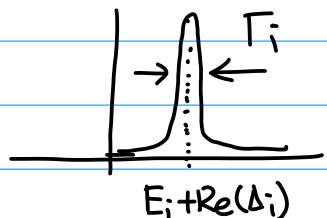
$$|C_i|^2 + \sum_{m \neq i} |C_m|^2 = \left(1 - \frac{\Gamma_i t}{\hbar}\right) + \underbrace{\sum_{m \neq i} W_{i \rightarrow m} t}_{= \frac{\Gamma_i}{\hbar} t} = 1$$

To see why Γ_i is the width: take Fourier decomposition:

$$\int f(E) e^{-\frac{iE}{\hbar}t} dE = e^{\frac{i}{\hbar}(E_i + \text{Re}(\Delta_i)t)} e^{-\frac{\Gamma_i t}{2\hbar}}$$

then find $f(E)$ by inverse transform:

$$|f(E)|^2 \propto \frac{1}{\{E - [E_i + \text{Re}(\Delta_i)]\}^2 + \frac{\Gamma_i^2}{4}}$$



Ex: staying in ground state:

- Particle in ground state of $U = -\alpha \delta(x)$ for $t < 0$.
- A weak uniform field $V(x,t) = -x F_0 \sin \omega t$ applied for $t > 0$

Find $W_0(t)$, probability that particle stays in ground state at t ?

First write out solution of the bound state:

$$\psi_{\text{bound}}(x, t=0) = \frac{\sqrt{m\alpha}}{\hbar} e^{-\frac{m\alpha}{\hbar^2} |x|} \quad \text{with } E_0 = -\frac{m\alpha^2}{2\hbar^2} = -\frac{k^2}{2} \alpha$$

The upper energy states are free particle states since $E \gg E_0$.

$$\psi_k(x) \sim e^{ikx}$$

To normalize, a very large box of length $[0, L]$, then

$$\int |\psi_k(x)|^2 dx = L \rightarrow \text{so } \psi_k(x) = \frac{1}{\sqrt{L}} e^{ikx}$$

To use Fermi's Golden Rule, rewrite the potential in the form:

$$V(t) = V_0 e^{i\omega t} + V_0^+ e^{-i\omega t}$$

for $V(t) = -x F_0 \left(\frac{e^{i\omega t} - e^{-i\omega t}}{2i} \right)$

$$= \frac{-xF_0}{2i} (e^{i\omega t} - e^{-i\omega t})$$

or $V = \frac{-xF_0}{2i}$

using Fermi's Golden Rule:

$$\omega_{0 \rightarrow k} = \frac{2\pi}{\hbar} |\psi_{k_0}|^2 \int \delta(E_k - E_0 - \hbar\omega) f(E) dE$$

$$|\langle \psi_k | -\frac{\nabla F_0}{2i} | \psi_{k_0} \rangle|^2 \quad E_k = \frac{\hbar^2 k^2}{2m}$$

If we imagine free particle in box of length L , then we have

$$\psi(x) = \sqrt{\frac{1}{L}} \sin(kx) \text{ and } \psi(x=L)=0 \Rightarrow kL = 2\pi n$$

$$\text{then we require } k = \frac{2\pi}{L} n$$

$$\hookrightarrow f(k) = \frac{dn}{dk} dk = \frac{L}{2\pi} dk$$

$$\text{so } \omega_{0 \rightarrow k} = \frac{L}{\hbar} \int |\psi_{k_0}|^2 \underbrace{\delta\left(\frac{\hbar^2 k^2}{2m} - E_0 - \hbar\omega_0\right)}_{f(k)} dk$$

$$\text{for } \delta(f(k)) = \sum_i \frac{\delta(k-k_i)}{|f'(k=k_i)|} \text{ of } f(k)$$

k_i is the i th root

$$\text{we know: } k_i = \pm \sqrt{\frac{2m(E_0 + \hbar\omega_0)}{\hbar^2}}$$

$$= \frac{L}{\hbar} \int |\psi_{k_0}|^2 \underbrace{\delta\left(k + \sqrt{\frac{2m(E_0 + \hbar\omega_0)}{\hbar^2}}\right) + \delta\left(k - \sqrt{\frac{2m(E_0 + \hbar\omega_0)}{\hbar^2}}\right)}_{\frac{\hbar^2}{m} |k|} dk$$

Now evaluate ψ_{k0} :

$$\begin{aligned}\psi_{k0} &= -\frac{F_0}{2i} \int_{-\infty}^{\infty} dx \underbrace{\frac{1}{L} e^{-ikx}}_{\gamma_k^*} \times \frac{\sqrt{m\alpha}}{\hbar} e^{-\frac{m\alpha}{\hbar^2}|x|} \\ &= -\frac{F_0}{2i} \frac{\sqrt{m\alpha}}{\hbar} \frac{1}{L} \int_{-\infty}^{\infty} dx e^{-ikx - \frac{m\alpha}{\hbar^2}|x|} \\ &= -\frac{F_0}{i} \frac{1}{L} \frac{\left(\frac{m\alpha}{\hbar^2}\right)^{3/2}}{k^2 + \left(\frac{m\alpha}{\hbar^2}\right)^2} \xrightarrow{\text{L}} \frac{2 \frac{m\alpha}{\hbar^2}}{k^2 + \left(\frac{m\alpha}{\hbar^2}\right)^2}\end{aligned}$$

$$|\psi_{k0}|^2 = \frac{F_0^2}{L} \frac{\left(\frac{m\alpha}{\hbar^2}\right)^3}{\left[k^2 + \left(\frac{m\alpha}{\hbar^2}\right)^2\right]^2}$$

$$\begin{aligned}&\stackrel{\text{SD}}{=} \omega_{0 \rightarrow k} = \frac{k}{\hbar} \int \frac{F_0^2}{L} \frac{\left(\frac{m\alpha}{\hbar^2}\right)^3}{\left[k^2 + \left(\frac{m\alpha}{\hbar^2}\right)^2\right]^2} \delta\left(k + \sqrt{\frac{2m(F_0 + \hbar\omega_0)}{\hbar^2}}\right) + \delta\left(k - \sqrt{\frac{2m(F_0 + \hbar\omega_0)}{\hbar^2}}\right) dk \\ &= \frac{m F_0^2}{\hbar^3} \frac{\left(\frac{m\alpha}{\hbar^2}\right)^3}{\left[\frac{2m(F_0 + \hbar\omega_0)}{\hbar^2} + \left(\frac{m\alpha}{\hbar^2}\right)^2\right]^2} \frac{2}{\sqrt{\frac{2m}{\hbar^2}} \sqrt{F_0 + \hbar\omega_0}} \\ &\text{set } K = \frac{m\alpha}{\hbar^2} \\ &= \frac{F_0^2}{\hbar^2} \sqrt{\frac{2m}{\hbar^2}} \frac{K^3}{\left(\frac{2m(-\hbar^2 K^2)}{\hbar^2} + \frac{2m\hbar\omega_0}{\hbar^2} + K^2\right)^2} \frac{1}{\sqrt{F_0 + \hbar\omega_0}} \\ &= \frac{F_0^2 \hbar^2}{(2m)^{3/2}} \frac{-K^3}{(\hbar\omega_0)^2} \frac{1}{\sqrt{F_0 + \hbar\omega_0}} \\ &\stackrel{!}{=} -\frac{F_0^2 \hbar^2 K^3}{(2m)^{3/2} (\hbar\omega_0)^2} \frac{1}{\sqrt{F_0 + \hbar\omega_0}}\end{aligned}$$

Since F_0 is negative
 $E_0 = -\frac{\hbar^2 K^2}{2m}$

Now

$$\frac{d}{dt} |C_n|^2 = \underbrace{\sum_i w_{i \rightarrow n} |C_i|^2}_{\text{other states going to state } n} - \underbrace{w_{n \rightarrow i} |C_n|^2}_{\text{state } n \text{ leaving}}$$

assume $w_{i \rightarrow n}$ or $w_{k \rightarrow 0} = 0$, then

$$\frac{d}{dt} |C_0|^2 = -w_{0 \rightarrow k} |C_0|^2 \rightarrow \boxed{|C_0(t)|^2 \stackrel{\substack{\text{probability to} \\ \text{stay in ground} \\ \text{state}}}{=} e^{-w_{0 \rightarrow k} t}}$$

(semi-classical)

WKB Approximation: $\hbar \rightarrow 0$ (Taylor in \hbar)

$$-\frac{\hbar^2}{2m} \nabla^2 \psi = (E - U(x)) \psi$$

↑
smooth in the scale of deBroglie wavelength

Assume $\psi(x) = \exp\left\{\frac{i}{\hbar} \delta(x)\right\}$

$$\hookrightarrow -\frac{\hbar^2}{2m} \vec{\nabla} \left(\frac{i}{\hbar} \exp\left\{\frac{i}{\hbar} \delta(x)\right\} \vec{\nabla} \delta(x) \right) = (E - U) \exp\left\{\frac{i}{\hbar} \delta(x)\right\}$$

$$\hookrightarrow -\frac{\hbar^2}{2m} \frac{i}{\hbar} \left\{ \nabla^2 \delta(x) + \frac{i}{\hbar} (\vec{\nabla} \delta(x))^2 \right\} = E - U$$

$$\hookrightarrow \boxed{-i\hbar \nabla^2 \delta(x) + (\vec{\nabla} \delta(x))^2 = 2m(E - U(x))} \quad \leftarrow \text{Equation to solve}$$

Let $\delta(x) = \delta^{(0)} + \frac{\hbar}{i} \delta^{(1)} + \left(\frac{\hbar}{i}\right)^2 \delta^{(2)} + \dots$

Now, let's consider 1D:

then we have:

$$(dx \delta(x))^2 - i\hbar dx^2 \delta(x) = P(x)$$

where

$$P(x) = \sqrt{2m(E - U(x))}$$

← Equation that we need to solve for.

First: 0th order solution:

$$(dx \delta^{(0)}(x))^2 - i\hbar dx \delta^{(0)}(x) = P(x)$$

$$(dx \delta^{(0)}(x))^2 = P(x) \Rightarrow \boxed{\delta^{(0)}(x) = \int^x \pm P(x') dx'}$$

zeroth
order
solution

$$\psi_0(x) = C_1 e^{i \int^x P(x') dx'} + C_2 e^{-i \int^x P(x') dx'}$$

$$\text{then } \Psi^{(0)}(x, t) = \psi(x) e^{-i\frac{E}{\hbar}t}$$

$$\Psi^{(0)} \underset{\int}{\approx} \exp \left\{ \pm i \int_{-\infty}^x p(x') dx' - \frac{i}{\hbar} Et \right\}$$

Applicability of WKB:

$$\text{First require } \hbar \partial_x^2 \delta^{(0)}(x) \ll (\partial_x \delta^{(0)}(x))^2$$

$$\text{we saw } \delta^{(0)}(x) \sim \pm \int_{-\infty}^x p(x') dx'$$

$$\text{so } \partial_x \delta^{(0)}(x) = p(x)$$

$$\text{then } \partial_x^2 \delta^{(0)}(x) = p'(x)$$

Hence condition becomes

$$\boxed{\hbar p'(x) \ll p(x)^2}$$

$$\text{Now since } \hbar k(x) = p(x)$$

$$\hookrightarrow k'(x) \ll k(x)^2$$

$$\text{using } k(x) = \frac{2\pi}{\lambda} \hookrightarrow \frac{\partial_x k(x)}{k^2} \ll 1$$

$$\hookrightarrow \left| \frac{d}{dx} \left(\frac{1}{k(x)} \right) \right| \ll 1$$

$$\text{or } * \quad \boxed{\left| \frac{d}{dx} \frac{\lambda}{2\pi} \right| \ll 1}$$

\leftarrow change in wavelength
is small compare
to the wavelength itself.

Another interpretation:

Starting from $\hbar p'(x) \ll p(x)^2$

since $p(x) = \sqrt{2m(E - U(x))}$

$$\text{so } \frac{d}{dx} p(x) = -\frac{m}{p} \frac{dU(x)}{dx} = \frac{mF(x)}{p}$$

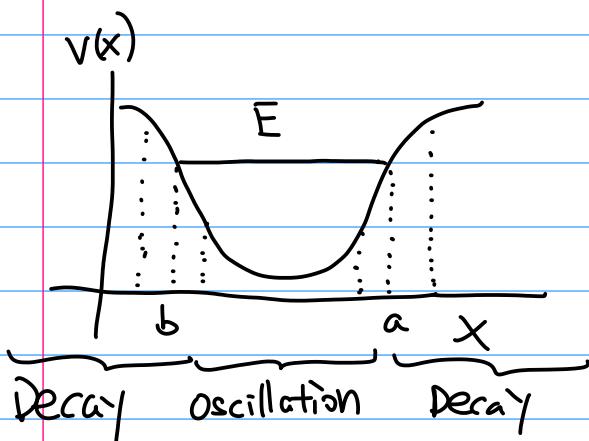
then $\frac{\hbar m}{p} F(x) \ll p^2$

or $\boxed{\frac{\hbar m}{p^3} F(x) \ll 1}$

the overall change
of the local momentum
must be small compared
to the momentum itself

or $\boxed{\lambda F(x) \ll \frac{p^2}{m}}$

← the potential energy
change over a wavelength
is much less than
local kinetic energy.



At turning point, $E \sim U(x)$, so
 $p \approx 0$, so $\frac{\hbar m F(x)}{p^3} \ll 1$

is not valid.

Now expand in higher orders:

$$\sigma(x) = \sigma^{(0)} + \left(\frac{i\hbar}{i}\right) \sigma^{(1)} + \left(\frac{i\hbar}{i}\right)^2 \sigma^{(2)} + \dots$$

$$\Rightarrow \left(\partial_x \sigma(x) \right)^2 - i\hbar \partial_x^2 \sigma(x) = P(x)$$

$$\hookrightarrow \left[\partial_x \left(\sigma^{(0)} + \left(\frac{i\hbar}{i}\right) \sigma^{(1)} + \left(\frac{i\hbar}{i}\right)^2 \sigma^{(2)} \right) \right]^2 - i\hbar \partial_x^2 \left(\sigma^{(0)} + \frac{i\hbar}{i} \sigma^{(1)} \right) = P(x)$$

$$\left(\partial_x \sigma^{(0)} \right)^2 + \frac{2i\hbar}{i} \left(\partial_x \sigma^{(0)} \right) \left(\partial_x \sigma^{(1)} \right) + \left(\frac{i\hbar}{i}\right)^2 \left[\left(\partial_x \sigma^{(1)} \right)^2 + 2 \partial_x \sigma^{(0)} \partial_x \sigma^{(2)} \right] - i\hbar \partial_x^2 \left(\sigma^{(0)} + \frac{i\hbar}{i} \sigma^{(1)} \right) = P(x)$$

Group by $\frac{i\hbar}{i}$:

$$i\hbar^{(0)} : \left[\partial_x \sigma^{(0)} \right]^2 = P(x) \Rightarrow \boxed{\sigma^{(0)} = \pm \int^x p(x') dx'}$$

$$\hookrightarrow \boxed{\psi^{(0)}(x) = C_1 e^{\pm \frac{i}{\hbar} \int^x p(x') dx'} + C_2 e^{\mp \frac{i}{\hbar} \int^x p(x') dx'}}$$

$$i\hbar^{(1)} : 2 \frac{i\hbar}{i} \left[\partial_x \sigma^{(0)} \right] \left[\partial_x \sigma^{(1)} \right] - i\hbar \partial_x^2 \sigma^{(0)} = 0$$

$$\hookrightarrow \partial_x \sigma^{(0)} \partial_x \sigma^{(1)} + \frac{1}{2} \partial_x^2 \sigma^{(0)} = 0$$

know $\partial_x \sigma^{(0)} = P$

$$\hookrightarrow \partial_x \sigma^{(1)} = \frac{\frac{d}{dx} P(x)}{2P(x)}$$

$$\hookrightarrow \boxed{\sigma^{(1)} = -\frac{1}{2} \ln P(x)}$$

then $\psi(x) = \exp \left\{ \frac{i}{\hbar} \left(\pm \int^x p(x') dx' - \frac{1}{2} \ln P(x) \right) \right\}$

$$\boxed{\psi^{(0+1)}(x) = \frac{C_1}{\sqrt{P}} e^{\pm \frac{i}{\hbar} \int^x p(x') dx'} + \frac{C_2}{\sqrt{P}} e^{\mp \frac{i}{\hbar} \int^x p(x') dx'}}$$

$$(5)^2: \left(\frac{t}{i}\right)^2 \left([\partial_x \sigma^{(1)}]^2 + 2 \partial_x \sigma^{(0)} \partial_x \sigma^{(2)} \right) - t^2 \partial_x^2 \sigma^{(1)} = 0$$

$$\hookrightarrow \partial_x \sigma^{(0)} \partial_x \sigma^{(2)} + \frac{1}{2} [\partial_x \sigma^{(1)}]^2 + \frac{1}{2} \partial_x^2 \sigma^{(1)} = 0$$

$$\hookrightarrow \partial_x \sigma^{(2)} = \frac{-\frac{1}{2} \left([\partial_x \sigma^{(1)}]^2 + \partial_x^2 \sigma^{(1)} \right)}{\partial_x \sigma^{(0)}}$$

$$= \frac{-\frac{1}{2} \left[\left(\frac{-1}{2} \frac{P'(x)}{P(x)} \right)^2 + \left(\frac{-1}{2} \right) \frac{P''(x)P(x) - (P'(x))^2}{P(x)^2} \right]}{P(x)}$$

$$\partial_x \sigma^{(2)} = \frac{1}{4} \frac{P''(x)}{P(x)^2} - \frac{3}{8} \frac{(P'(x))^2}{P(x)^3}$$

$$\hookrightarrow \zeta^{(2)}(x) = \int dx' \frac{P''}{4P} - \frac{3}{8} \frac{(P')^2}{P^3}$$

$$= \frac{P'}{4P} + \int \frac{(P')^2}{2P^3} - \frac{3(P')^2}{8P^3} dx'$$

using $F = \frac{PP'}{m}$

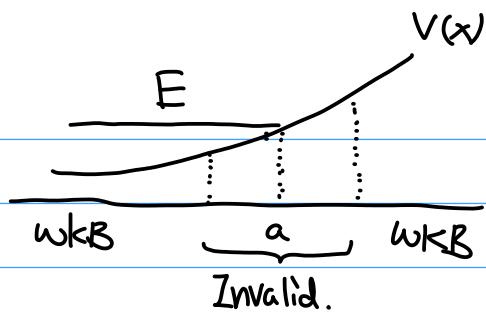
$$= \frac{1}{4} \frac{mF}{P^3} + \frac{m^2}{8} \int^x \frac{F^2}{P^5} dx'$$

then $\hat{\gamma}(x)^{(0+1+2)} = \exp \left\{ \frac{i}{\hbar} \left(\sigma^{(0)} + \frac{t}{i} \sigma^{(1)} + \left(\frac{t}{i} \right)^2 \sigma^{(2)} \right) \right\}$

$$= \left[\frac{C_1}{\sqrt{P}} e^{\int^x P(x) dx'} + \frac{C_2}{\sqrt{P}} e^{-\int^x P(x) dx'} \right] \left(\exp \left\{ \frac{t}{i} \sigma^{(2)} \right\} \right)$$

$$\boxed{\hat{\gamma}(x)^{(0+1+2)} = \left[\frac{C}{\sqrt{P}} e^{i \int^x P(x) dx'} + \frac{C_2}{\sqrt{P}} e^{-i \int^x P(x) dx'} \right] \left(1 - it \sigma^{(2)} \right)}$$

Connection Formula:



wKB is invalid about turning point:

Turning point, a and b, when $p(a) = p(b) = 0$

$$\text{or } E = U(a) = U(b)$$

$$\text{Define } k = \frac{1}{\hbar} \sqrt{2m(E - U(x))} \quad \text{when } E > U(x)$$

$$k = \frac{1}{\hbar} \sqrt{2m(U(x) - E)} \quad \text{when } E < U(x)$$

then wKB with zeroth and first order tell us:

(grow/decay): $\psi_{x \gg a}(x) = \frac{A}{\Gamma k} e^{-\int_a^x k(x') dx'} + \frac{B}{\Gamma k} e^{\int_a^x k(x') dx'}$ *

oscillation: $\psi_{x \ll a}(x) = \frac{C}{\Gamma k} e^{-i \int_a^x k(x') dx'} + \frac{D}{\Gamma k} e^{i \int_a^x k(x') dx'}$

Now the question is how is A,B related with C,D.

To connect, we need to solve Schrodinger more accurately near turning points. by approximating the potential as a linear potential near turning point.

Suppose we want to solve a specific case:

at $x \approx a$ (turning point):

linear potential: $U(x) - E = g(x-a)$ where $g > 0$

then we have $-\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} + g(x-a)\psi = 0$

lets convert $x \rightarrow z = \left(\frac{2mg}{\hbar^2}\right)^{1/3}(x-a)$

$$\Rightarrow \frac{d^2\psi}{dz^2} - z\psi = 0$$

sine $k^2 = \frac{2m}{\hbar^2} (E - U(x))$

$$= -\frac{2mg}{\hbar^2} (x-a) \quad \Rightarrow E - U(x) = -g(x-a)$$

$$k^2 = -\left(\frac{2mg}{\hbar^2}\right)^{2/3} z$$

using applicability condition:

$$\left| \frac{d}{dx} k(x) \right| \ll k(x)^2 \quad \text{or} \quad \left| \frac{d}{dx} \left(\frac{1}{k} \right) \right| \ll 1$$

$$\hookrightarrow \frac{d}{dz} \left(\frac{2mg}{\hbar^2} \right)^{1/3} \left(-\left(\frac{2mg}{\hbar^2}\right)^{-1/3} z^{-1/2} \right) \ll 1$$

$$\hookrightarrow \frac{1}{2} z^{-3/2} \ll 1$$

or $|z|^{3/2} \gg \frac{1}{2}$ ← For WKB-Applicability.

The exact solution to

$$\frac{d^2}{dz^2} \psi - z \psi = 0$$

are Airy Functions: $A_i(z)$ and $B_i(z)$

Since WKB condition is $|z|^{3/2} \gg \frac{1}{2}$

this means we can use WKB for Airy Functions for very large z in both positive and negative.

In general:

$$A_i(z) = \frac{1}{\pi} \int_0^\infty \cos\left(\frac{s^3}{3} + sz\right) ds$$

$$B_i(z) = \frac{1}{\pi} \int_0^\infty \left(\exp\left\{-\frac{s^3}{3} + sz\right\} + \sin\left(\frac{s^3}{3} + sz\right) \right) ds$$

For $z \rightarrow +\infty$:

$$A_i(z) \approx \frac{1}{2\pi} z^{-1/4} e^{-\frac{2}{3}|z|^{3/2}}$$

$$B_i(z) \approx \frac{1}{\pi} z^{-1/4} e^{\frac{2}{3}|z|^{3/2}}$$

For $z \rightarrow -\infty$:

$$A_i(z) \approx \frac{1}{\pi} |z|^{-1/4} \cos\left(\frac{2}{3}|z|^{3/2} - \frac{\pi}{4}\right)$$

$$B_i(z) \approx -\frac{1}{\pi} |z|^{-1/4} \sin\left(\frac{2}{3}|z|^{3/2} - \frac{\pi}{4}\right)$$

Now recognize for linear potential case:

$$\int_x^a K(x) dx = \frac{2}{3} \left(\frac{2mg}{\hbar^2} \right)^{1/2} (a-x)^{3/2} = \frac{2}{3} |z|^{3/2} \quad \text{for } z < 0$$

$$\int_x^a K(x) dx = \frac{2}{3} \left(\frac{2mg}{\hbar^2} \right)^{1/2} (x-a)^{3/2} = \frac{2}{3} |z|^{3/2} \quad \text{for } z > 0$$

We see that the asymptotic solution of the exact solution are the same as WKB-solutions.

Therefore: we use the fact of asymptotic solution and write:

$$\frac{2A}{\sqrt{K(x)}} \cos \left(\int_x^a K(x) dx - \frac{\pi i}{4} \right) - \frac{B}{\sqrt{K(x)}} \sin \left(\int_x^a K(x) dx - \frac{\pi i}{4} \right) \quad \begin{matrix} x \ll a \\ a \end{matrix} \quad (\text{rising slope})$$

$$= \frac{A}{\sqrt{K(x)}} \exp \left(- \int_a^x K(x) dx \right) + \frac{B}{\sqrt{K(x)}} \exp \left(\int_a^x K(x) dx \right) \quad x \gg a$$

Similarly:

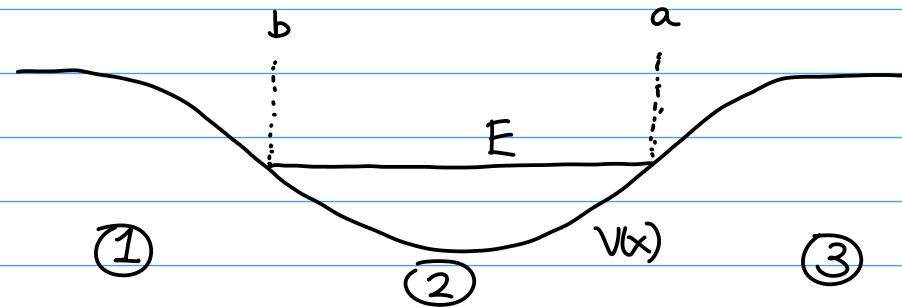
connection formula

$$\frac{A}{\sqrt{K(x)}} \exp \left(- \int_x^b K(x) dx \right) + \frac{B}{\sqrt{K(x)}} \exp \left(\int_x^b K(x) dx \right) \quad \begin{matrix} x \ll b \\ b \end{matrix} \quad (\text{lowering slope})$$

$$= \frac{2A}{\sqrt{K(x)}} \cos \left(\int_b^x K(x) dx - \frac{\pi i}{4} \right) - \frac{B}{\sqrt{K(x)}} \sin \left(\int_b^x K(x) dx - \frac{\pi i}{4} \right) \quad x \gg b$$

- Remarks:
- 1) If turning points are too close, then cannot use WKB, invalid regime spans entirely.
 - 2) If $V(x)$ is singular then $K(x)$ also becomes singular.

Example: Bound state: (1D potential well)



In region ①: since wave function must decay as $x \rightarrow -\infty$,

$$\psi(x) \approx \frac{1}{\sqrt{k}} \exp\left\{-\int_x^b k(x) dx\right\} \quad \text{for } x < b$$

Since for $x < b$, we have no growing term, then $B=0$
we also observe $A=1$ by matching.

so in region ②:

$$\psi(x) = \frac{2}{\sqrt{k}} \cos\left(\int_b^x k(x) dx - \frac{\pi}{4}\right) \quad \text{for } b > x > a$$

$$= \frac{2}{\sqrt{k}} \cos\left(\int_b^a k(x) dx - \int_x^a k(x) dx - \frac{\pi}{4}\right)$$

$$= -\frac{2}{\sqrt{k}} \cos\left(\int_b^a k(x) dx\right) \sin\left(\int_x^a k(x) dx - \frac{\pi}{4}\right)$$

$$+ \frac{2}{\sqrt{k}} \sin\left(\int_b^a k(x) dx\right) \cos\left(\int_x^a k(x) dx - \frac{\pi}{4}\right)$$

using connection formula we see

$$B = \frac{2}{\sqrt{k}} \cos\left(\int_b^a k(x) dx\right) \quad \text{and} \quad A = \frac{1}{\sqrt{k}} \sin\left(\int_b^a k(x) dx\right)$$

Then in region ③:

$$\psi(x) = \underbrace{\frac{1}{\Gamma K} \sin \left(\int_b^a k(x) dx \right)}_{A} \frac{1}{\Gamma K} \exp \left(- \int_a^x k(x) dx \right)$$

$$+ \underbrace{\frac{2}{\Gamma K} \cos \left(\int_b^a k(x) dx \right)}_{B} \frac{1}{\Gamma K} \exp \left(\int_a^x k(x) dx \right)$$

However, we don't want growing terms in ③:
so choose $B=0$

or $\cos \left(\int_b^a k(x) dx \right) = 0$

this means

$$\boxed{\int_b^a k(x) dx = \left(n + \frac{1}{2}\right)\pi \quad \text{for } n=0, 1, 2, \dots}$$

Bohr-Sommerfeld Formula.

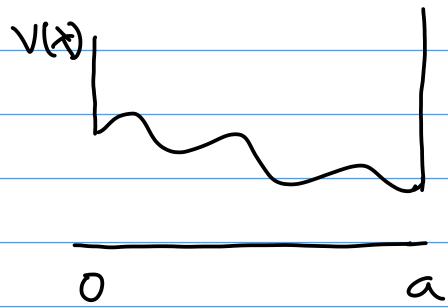
then $\psi(x) = \frac{2}{\Gamma K} \sin \left(\int_b^a k(x) dx \right) \cos \left(\int_x^a k(x) dx - \frac{\pi}{4} \right)$ for $b > x > a$

and $\psi(x) = \frac{1}{\Gamma K} \sin \left(\int_b^a k(x) dx \right) \frac{1}{\Gamma K} \exp \left\{ - \int_a^x k(x) dx \right\}$ for $x > a$

or

$$\boxed{\oint p dx = 2\pi\hbar \left(n + \frac{1}{2}\right)}$$

Bound state: potential well with 2 vertical walls



$$\text{know } \psi(x \leq 0) = 0 \quad \psi(x \geq a) = 0$$

we have oscillatory solution for $x \in [0, a]$

$$\begin{aligned} \psi &= \frac{A}{P} \exp\left\{+i \int_0^x P dx'\right\} + \frac{B}{P} \exp\left\{-i \int_0^x P dx'\right\} \\ \psi(x) &= \frac{A}{P} \cos\left(\int_0^x k dx'\right) + \frac{B}{P} \sin\left(\int_0^x k dx'\right) \end{aligned}$$

$$\Rightarrow \psi(x=0) = \frac{A}{P(0)} \underbrace{\cos(0)}_{=1} = 0$$

$$\text{so } A=0$$

$$\Rightarrow \psi(x=a) = \frac{B}{P(a)} \underbrace{\sin\left(\int_0^a k(x') dx'\right)}_{=\pi n} = 0$$

so

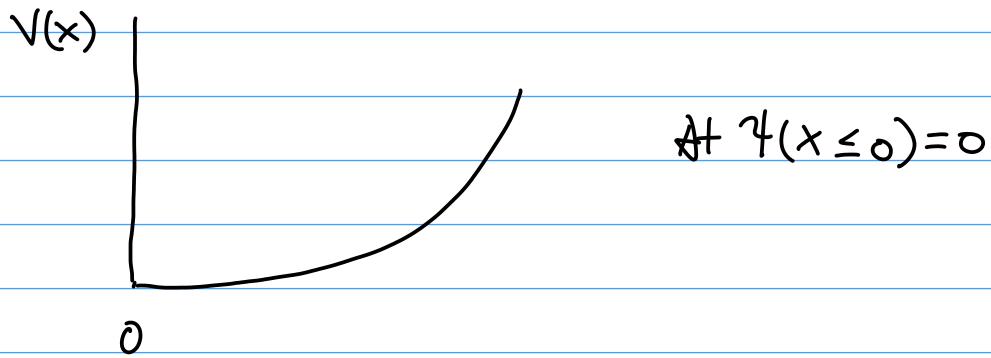
$$\int_0^a k(x') dx' = \pi n$$

or

$$\int_0^a P(x') dx' = n \pi \hbar$$

condition for
bound state when
there are two vertical
walls.

Bound state: $1 \propto$ potential wall:



Using connection formula for rising slope:

$$\begin{aligned} & \frac{2A}{\Gamma K} \cos\left(\int_x^{x_0} k(x') dx' - \frac{\pi}{4}\right) - \frac{B}{\Gamma K} \sin\left(\int_x^{x_0} k(x') dx' - \frac{\pi}{4}\right) \quad \text{for } x \ll x_0 \\ = & \frac{A}{\Gamma K} \exp\left\{-\int_{x_0}^x k(x') dx'\right\} + \frac{B}{\Gamma K} \exp\left\{\int_{x_0}^x k(x') dx'\right\} \quad \text{for } x \gg x_0 \end{aligned}$$

At $x \gg x_0$, we do not expect growing solutions, so $B=0$

So for $x \ll x_0$, we have:

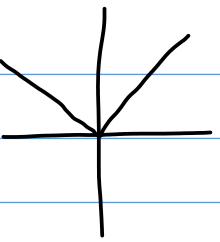
$$\psi(x) = \frac{2A}{\Gamma K} \cos\left(\int_x^{x_0} k(x') dx' - \frac{\pi}{4}\right)$$

$$\psi(x=0) = \frac{2A}{\Gamma K(0)} \cos\left(\int_0^{x_0} k(x') dx' - \frac{\pi}{4}\right) = 0$$

Require $\int_0^{x_0} k(x') dx' - \frac{\pi}{4} = (n + \frac{1}{2})\pi$

or $\boxed{\int_0^{x_0} k(x') dx' = \left(n + \frac{3}{4}\right)\pi}$

example: $-\frac{\hbar^2}{2m} \frac{d^2}{dx^2} \psi + g|x|\psi = E\psi$

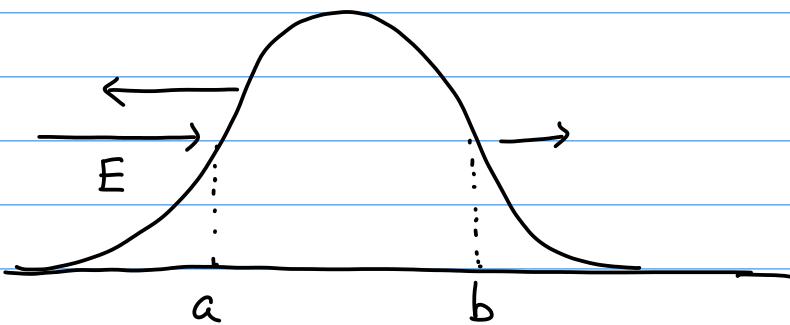


$$2 \int_0^{E/g} \sqrt{\frac{2m}{\hbar^2}} (E - gx) = (n + \frac{1}{2}) \pi$$

$$E_{\text{WKB}}^3 = \frac{g}{32} \pi^2 \left(n + \frac{1}{2}\right)^2 \frac{g^2 \hbar^2}{m}$$

$$E_{\text{WKB}} \approx \left(n + \frac{1}{2}\right)^{2/3}$$

Transmission through a barrier:



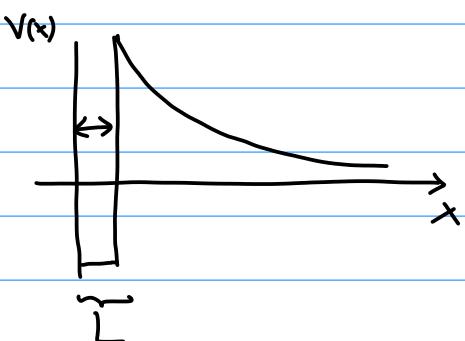
find $T = |t|^2 \approx e^{-2 \int_a^b K(x) dx}$ when WKB is valid

probability of emission

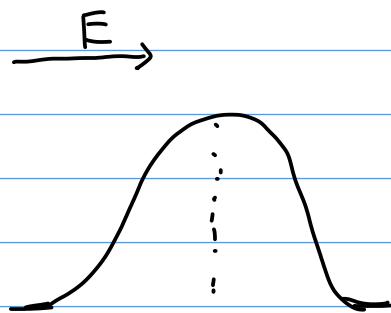
$$R = \frac{1}{T}$$

average life time. $\bar{\tau} = \frac{2L}{\lambda} e^{2\gamma}$ \hookrightarrow frequency of collisions inside the potential.

energy width $\rightarrow \Gamma = \frac{\hbar}{\bar{\tau}} = \hbar \frac{v}{2L} e^{-2\gamma}$



Reflection about Barrier:



if classical, particle slows down as it goes through potential, but all goes through.

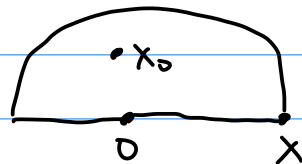
$$\psi = \frac{1}{\sqrt{K(x)}} e^{i \int_0^x K(x) dx}$$

Find turning points: $V(x) = E$

but in case $E > V(x)$, we get complex turning points.

$$Ex: \quad V(x) = -\frac{1}{2}ax^2 = E \rightarrow x = \pm i\sqrt{\frac{2E}{a}}$$

Now solve



then

$$R = e^{-4 \operatorname{Im} \int_0^{x_0} K(x) dx}$$

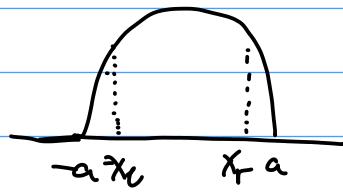
$$\stackrel{!}{=} e^{-2\pi i \epsilon}$$

vs exact $\frac{1}{1 + e^{2\pi i \epsilon}}$

$$\text{where } \epsilon = \frac{E}{\hbar} \sqrt{\frac{m}{a}}$$

Tunneling examples:

1) $U(x) = U_0 \left(1 - \left(\frac{x}{a}\right)^2\right)$



$$T = e^{-2\gamma} = \exp \left\{ -2 \int_{x_L}^{x_R} K(x) dx \right\}$$

$$= \exp \left\{ -2 \int_{x_L}^{x_R} \sqrt{\frac{2m}{\hbar^2} (U_0(1 - (\frac{x}{a})^2) - E)} dx \right\}$$

Find classical turning point:

$$E = U(x_L, x_R) = U_0 \left(1 - \left(\frac{x}{a}\right)^2\right)$$

$$\hookrightarrow x_{L,R} = \pm a \sqrt{1 - \frac{E}{U_0}}$$

$$\text{so } T = \exp \left\{ -2 \int_{-a\sqrt{1-E/U_0}}^{a\sqrt{1-E/U_0}} \sqrt{\frac{2m}{\hbar^2} (U_0 - E - U_0(\frac{x}{a})^2)} dx \right\}$$

$$= \exp \left\{ -2 \sqrt{\frac{2m(U_0 - E)}{\hbar^2}} \int \sqrt{1 - \frac{U_0}{U_0 - E} \left(\frac{x}{a}\right)^2} dx \right\}$$

$$\text{let } \zeta = \sqrt{\frac{U_0}{U_0 - E}} \frac{x}{a} \quad \text{then} \quad dx = \sqrt{\frac{(U_0 - E)}{U_0}} a d\zeta$$

$$= \exp \left\{ -2 \frac{(U_0 - E)}{\hbar} \sqrt{\frac{2m}{U_0}} a \int_{-1}^{+1} \sqrt{1 - \zeta^2} d\zeta \right\}$$

$$T = \exp \left\{ -\frac{\pi a}{\hbar} (U_0 - E) \sqrt{\frac{2m}{U_0}} \right\}$$

Applicability is to require
 $T \ll 1$ or $\frac{\pi a}{\hbar} (U_0 - E) \sqrt{\frac{2m}{U_0}} \gg 1$

Adiabatic Approximation (Born - Oppenheimer)

Hamiltonian changing slowly:

$$i\hbar \partial_t \psi = H(\lambda(t)) \psi$$

assume $H(\lambda(t)) \psi_n = E_n(t) \psi_n$

↖ instantaneous eigenstate.

then $i\hbar \partial_t \psi_n = E_n(t) \psi_n$

$$\psi_n = \boxed{e^{-\frac{i}{\hbar} \int_0^t E_n(t') dt'}} \psi_n(t=0)$$

Dynamical
phase, $i\theta_n(t)$

↖ but we're not
solving for time-dependent
Schrodinger.

But we can do better:

$$\psi(t) = \sum_m c_m(t) \psi_m(t) e^{i\theta_m(t)} \quad \text{and} \quad \langle \psi_n | \psi_m \rangle = \delta_{nm}$$

↪ plug into $i\hbar \partial_t \psi = H \psi$

$$\Rightarrow \dot{c}_m(t) = - \sum_n c_n \langle \psi_m | \dot{\psi}_n \rangle e^{i(\theta_n - \theta_m)}$$

$$\text{then } \underset{m \neq n}{\langle \psi_m | \dot{H} | \psi_n \rangle} = (E_n - E_m) \langle \psi_m | \psi_n \rangle$$

$$\dot{c}_m(t) = - \underbrace{c_m \langle \psi_m | \dot{\psi}_m \rangle}_{\text{when } n=m} - \sum_{m \neq n} c_m \frac{\langle \psi_m | \dot{H} | \psi_m \rangle}{E_n - E_m} e^{\frac{i}{\hbar} \int_0^t [E_n(t') - E_m(t')] dt'}$$

If H changes slowly, i.e. $\dot{H} \ll 1$, and energy do not cross, i.e. $E_n \neq E_m$

then

$$\dot{C}_m(t) = -\underbrace{C_m}_{\text{when } n=m} \langle \psi_m | \dot{\psi}_m \rangle$$

so $C_m(t) = C_m(0) e^{i\gamma_m(t)}$

$$\boxed{\gamma_m(t) = i \int_0^t \langle \psi_m(t') | \frac{\partial}{\partial t} \psi_m(t') \rangle dt'}$$

↑
Berry's phase.

then $\boxed{\psi_n(t) = e^{i\theta_n(t)} e^{i\gamma_m(t)} \psi_m(0)}$

$$H \quad H(t) \rightarrow H(\lambda(t))$$

then $\gamma_m(t) = i \int_0^t \langle \psi_m | \nabla_{\lambda} \psi_m \rangle \frac{d\lambda}{dt} dt$

$$= i \int_C d\lambda \langle \psi_m | \nabla_{\lambda} \psi_m \rangle$$

↳ geometric phase.
or Berry phase.

$$Ex: i\hbar \frac{d}{dt} \vec{\gamma} = -\mu \vec{B}(t) \vec{\jmath} \vec{\gamma}$$

↳ slow

$$H = -\mu \vec{B} \cdot \vec{\sigma} \quad \rightarrow E = \pm \mu |\vec{B}|$$

$$\chi_- \left\{ \begin{array}{l} \sin \frac{\theta}{2} e^{-i\phi} \\ -\cos \frac{\theta}{2} \end{array} \right.$$

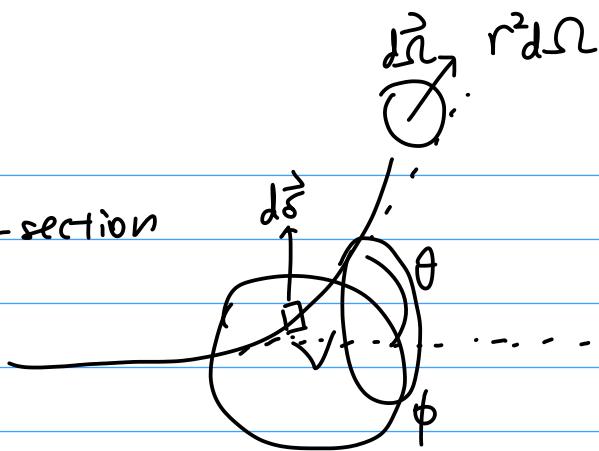
$$\vec{B} = B(t) [\sin \theta \cos \phi \hat{x}, \sin \theta \sin \phi \hat{y}, \cos \theta \hat{z}]$$

$$\text{Find } \gamma_- = \frac{1}{2} \oint_{\gamma_+} d\phi (1 - \cos \theta)$$

↳ initial direction
of magnetic field

Scattering theory

incident flux $\vec{J}_{\text{incident}} \cdot d\vec{\sigma}$



$$dN = r^2 \vec{J}_{\text{scatter}} d\vec{\Omega}$$

$$\hookrightarrow \vec{J}_{\text{incident}} d\vec{\sigma} = r^2 \vec{J}_{\text{scattered}} d\vec{\Omega}$$

$$d\vec{\sigma} = \frac{J_{\text{scatter}}}{J_{\text{incident}}} r^2 d\vec{\Omega}$$

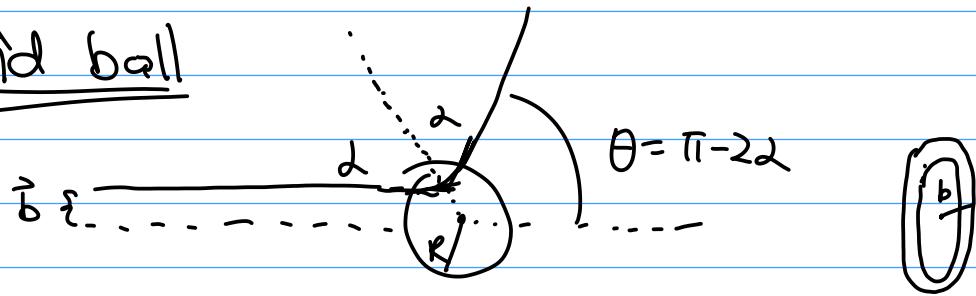
$$\frac{d\vec{\sigma}}{d\vec{\Omega}} = \frac{J_{\text{scatter}}}{J_{\text{incident}}} r^2 \quad (\theta, \phi)$$

assume azimuthal symmetry.

then

$$\Sigma = \int \underbrace{\frac{d\sigma}{d\Omega} d\Omega}_{\text{sum over } \phi} \quad \text{total cross-section.}$$

ex: solid ball



$$b = R \sin \frac{\theta}{2}$$

$$= R \sin \left(\frac{\pi - \theta}{2} \right)$$

$$= R \cos \frac{\theta}{2}$$

$$dS = 2\pi b db$$

$$= \frac{dS}{d\theta} 2\pi \approx \frac{db}{d\theta} d\theta$$

$$= 2\pi R \cos \frac{\theta}{2} dR \cos \frac{\theta}{2}$$

$$= 2\pi R^2 \underbrace{\cos \frac{\theta}{2} \sin \frac{\theta}{2} \frac{1}{2}}_{\frac{1}{4} \sin \theta} d\theta$$

compare with

$$\frac{dS}{d\theta} 2\pi \sin \theta d\theta$$

then find

$$\boxed{\frac{dS}{d\theta} = \frac{R^2}{4}}$$

using classical mechanics.

ex 2:

For $V = \frac{ze^2}{r}$, we find

using solid ball

$$\frac{dS}{d\theta} = \frac{ze^2}{E} \frac{1}{16 \sin^4 \frac{\theta}{2}}$$

In quantum mechanics -

$$e^{ikx} \text{ with } E = \frac{\hbar^2 k^2}{2m}$$



$$\frac{e^{i\vec{k} \cdot \vec{r}}}{r} f(\theta, \phi)$$

$$\phi_{\text{incident}} \sim e^{ikx} \quad \leftarrow \text{Cartesian.}$$

$$\phi_{\text{scatter}} \sim \frac{e^{ikr}}{r} f(\theta, \phi) \quad \rightarrow \text{spherical coordinates}$$

$$\text{so } |\phi_{\text{scatter}}|^2 \sim J_{\text{scattered}}$$

$$\text{so } d\sigma = |f(\theta, \phi)|^2 d\Omega$$



determine this for scattering wave func
solve perturbatively using Bohr-Approximation