# Meteo 515 - Assignment 2 - Parametric Distributions

Part 2 – Examining Marshall–Palmer fits to a rain drop size distribution (DSD)

The data are from a 2-D video disdrometer in the western Pacific on 02-Jan-15

```
In [1]:
```

```
from __future__ import division, print_function
#import datetime as dt

import matplotlib.pyplot as plt
from netCDF4 import Dataset, num2date
import numpy as np
#import pandas as pd
from scipy.optimize import curve_fit
#import scipy.stats as ss
#import statsmodels.api as sm
```

In [2]:

```
plt.style.use('seaborn-darkgrid') %matplotlib notebook
```

## Load the data

using netcdf4-python from Unidata

```
In [3]:
```

```
fname = './data/twpvdisC3.b1.20150102.000000.cdf'
d = Dataset(fname)

base_time = d['base_time']  # seconds since 1970-1-1 0:00:00 0:00
time_offset = d['time_offset']  # same as time...
time = d['time']  # seconds since base time
assert( np.all(time[:] == time_offset[:]) )
t_dt = num2date(time[:], time.units)  # create datetimes

drop_diameter = d['drop_diameter']
num_density = d['num_density']

intercept_param = d['intercept_parameter']
slope_param = d['slope_parameter']
```

# a) DSD at chosen minute

Here, we choose the time at which the sum of the larger bins is maximized.

The simpler Marshall-Palmer fit that we will use for our binned drop number density vs. drop diameter data has the form

$$n(D) = n_0 e^{-\lambda D}$$

where n(D) is the number density of raindrops of size D (technically in the size range [D, D + dD]),  $n_0$  is called the intercept parameter, and  $\lambda$  is called the slope parameter. n(D) is also called the drop size distribution (function).

## b) More complex fit

The more complex fit we will use has the form

$$n(D) = n_0 D^{\mu} e^{-\lambda D}$$

which has an added 3rd parameter,  $\mu$ . Note that the k-th moment of the drop size distribution n(D) is given by  $M_k = \int_0^\infty n(D) \, D^k \, dD$ 

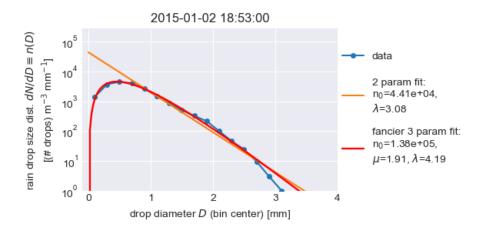
\*\*Note:\*\* in the assignment N is used for number concentration, but I prefer n for number concentration and N for a total particle count.

#### In [4]:

```
def ND3(D, N_0, mu, lamb):
    """More complex Marshall Palmer number density -- drop size relation"""
    return N_0 * D**mu * np.exp(-lamb*D)
```

#### In [5]:

```
i_t choice = np.argmax(num_density[:,10:].sum(axis=1))
t choice = t dt[i t choice]
Dplot = np.linspace(0, 10, 400)
f1, a = plt.subplots(figsize=(6, 3.0), num='dsd')
a.semilogy(drop_diameter[:], num_density[i_t_choice], 'o-', ms=4, label='data') # the data
a.semilogy(Dplot, intercept_param[i_t_choice]*np.exp(-slope_param[i_t_choice]*Dplot),
          label='2 param fit:\nn$_0$=%.3g,\n$\lambda$=%.3g' % (intercept_param[i_t_choice], slope_param[i
_t_choice])) # the exp fit
popt, pcov = curve fit(ND3, drop_diameter[:], num_density[i_t_choice],
                         p0=(1e5, 1, 1), xtol=1e-10)
a.plot(Dplot, ND3(Dplot, *popt), 'r-', lw=1.7,
       label='fancier 3 param fit:\nn$_0$=%.3g,\n$\mu$=%.3g, $\lambda$=%.3g' % tuple(popt))
a.set_xlabel('drop diameter $D$ (bin center) [{:s}]'.format(drop diameter.units))
a.set_xlim((-0.1, 4))
a.set_ylabel('rain drop size dist. dN/dD \neq (0) \ln[{:s}]'.format('(# drops) m^{-3}$ mm^{-1}$'))
#num density.units))
a.set_ylim(ymin=num_density[i_t_choice][num_density[i_t_choice] > 0].min()) # must be a cleaner way to do
this...
a.set title(str(t choice)[:19]) # don't care about sub-second
a.legend(loc='center left', bbox_to_anchor=(0.98, 0.5), labelspacing=1.5)
f1.tight layout();
```



### **Discussion:**

We can see (easily with the log y-axis) that the simpler fit is not able to capture the distribution behavior at the edges, and in fact is unphysical (has a nonzero value at D = 0). The 3 parameter fit captures the shape much better, although does not quite capture the sharpness of the changes at higher D's.

# Save figures locally

Probably a silly thing for a Jupyter Notebook to do but whatever