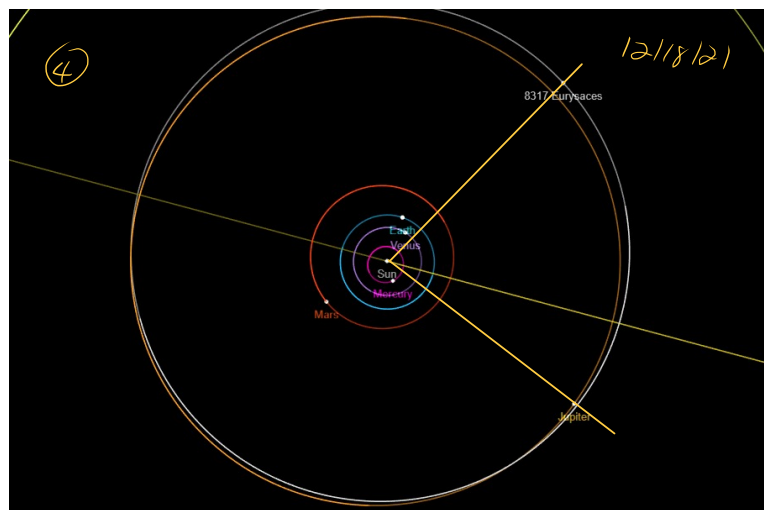
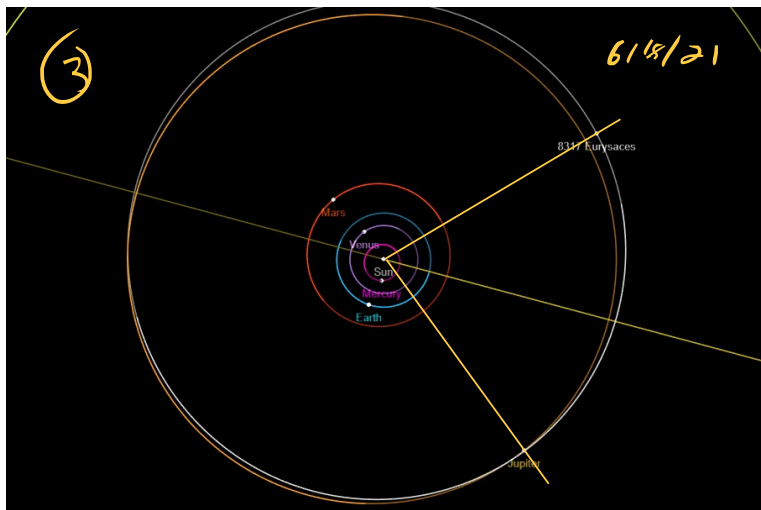
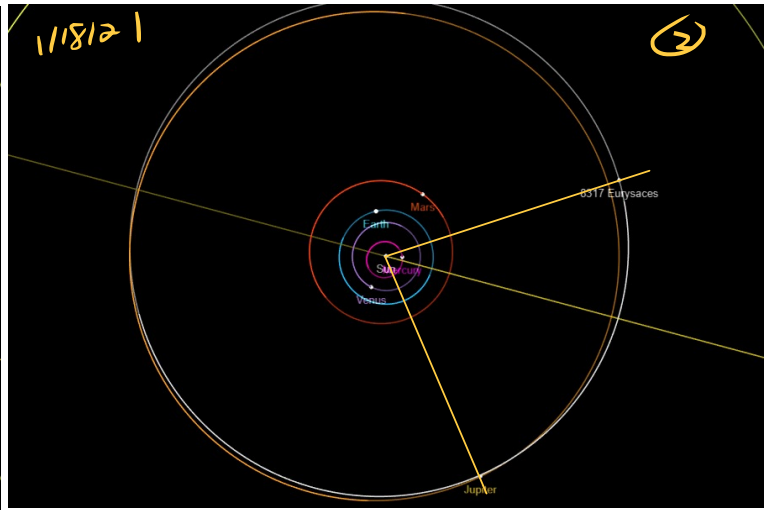
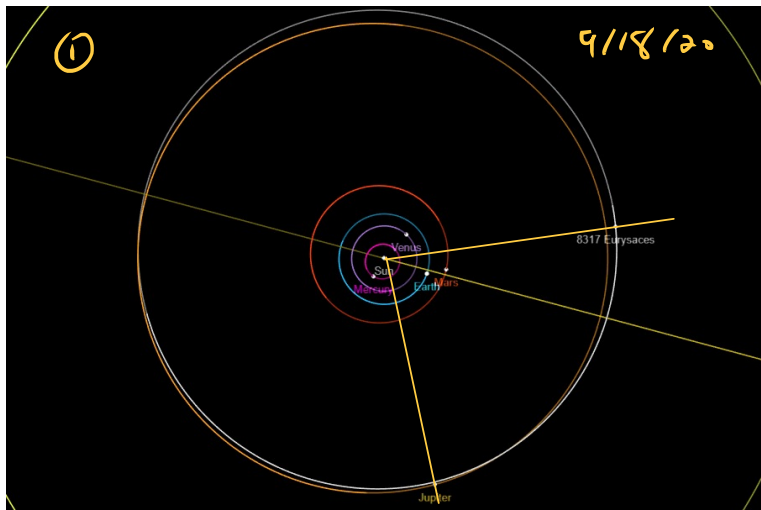


1. a) Euryaces is in the Creek camp as it is in front of the Jupiter.

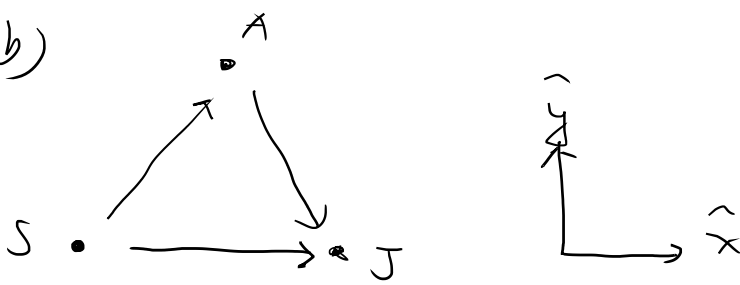
Its orbit has a period of 12.7 years around the Sun, which is slightly longer than that of Jupiter, 11.86 years.



- ① : 86.5°
- ② : 85°
- ③ : 84°
- ④ : 82.5°

The angle decreases approximately 1 degrees every 6 months

1. b)



i)

$$\vec{r}_{SA} + G \underbrace{\frac{(m_S + m_A)}{r_{SA}^3}}_{\text{Dominant}} \vec{r}_{SA} = G m_J \left(\underbrace{\frac{\vec{r}_{AJ}}{r_{AJ}^3}}_{\text{Direct}} - \underbrace{\frac{\vec{r}_{SJ}}{r_{SJ}^3}}_{\text{Indirect}} \right)$$

$$\vec{r}_{SJ} = r_{SJ} \hat{x}, \quad \vec{r}_{SA} = \begin{bmatrix} \cos \frac{\pi}{3} & 0 \\ 0 & \sin \frac{\pi}{3} \end{bmatrix} \vec{r}_{SJ}$$

$$\vec{r}_{AJ} = \vec{r}_{SJ} - \vec{r}_{SA}$$

$$\text{Dominant term: } \| 0.1095 \hat{x} + 0.1897 \hat{y} \| \times 10^{-6} = \underline{2.1910 \times 10^{-7} \text{ km/s}^2}$$

$$\text{Direct term: } \| 0.1046 \hat{x} - 0.1812 \hat{y} \| \times 10^{-9} = \underline{2.0919 \times 10^{-10} \text{ km/s}^2}$$

$$\text{Indirect term: } \| 0.2092 \hat{x} \| \times 10^{-9} = \underline{2.0919 \times 10^{-10} \text{ km/s}^2}$$

$$\text{Net perturbing acceleration: } \| \text{Direct} - \text{Indirect} \| = \underline{2.0919 \times 10^{-10} \text{ km/s}^2}$$

$$\text{total acceleration: } \| \text{Direct} - \text{Indirect} - \text{dominant} \| = \underline{2.1931 \times 10^{-7} \text{ km/s}^2}$$

1. b)

$$ii) \quad \ddot{\vec{r}}_{JA} + \underbrace{\frac{G(M_J + M_A)}{r_{JA}^3}}_{\text{Dominant}} \vec{r}_{JA} = G M_S \left(\underbrace{\frac{\vec{r}_{AS}}{r_{AS}^3}}_{\text{Direct}} - \underbrace{\frac{\vec{r}_{JS}}{r_{JS}^3}}_{\text{Indirect}} \right)$$

$$\text{Dominant: } \|\vec{r}_{JA}\| = \sqrt{(-0.1046 \hat{x} + 0.1812 \hat{y})^2} \times 10^{-9} = \underline{2.0919 \times 10^{-10} \text{ km/s}^2}$$

$$\text{Direct: } \|\vec{r}_{AS}\| = \sqrt{(-0.1095 \hat{x} - 0.1897 \hat{y})^2} \times 10^{-6} = \underline{2.1916 \times 10^{-7} \text{ km/s}^2}$$

$$\text{Indirect: } \|\vec{r}_{JS}\| = \sqrt{(-0.2191 \hat{x})^2} \times 10^{-6} = \underline{2.1910 \times 10^{-7} \text{ km/s}^2}$$

$$\text{Net perturbing acceleration: } \underline{2.1910 \times 10^{-7} \text{ km/s}^2}$$

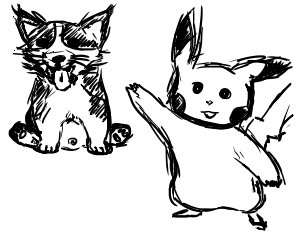
$$\text{Total acceleration: } \underline{2.1931 \times 10^{-7} \text{ km/s}^2}$$

ii) For formulation in part (i), the dominant term is the largest. For formulation in part (ii), the direct & indirect terms, perturbing terms, are equally the largest.

Comparing net perturbations, the Sun has the largest impact. The net accelerations on the asteroid are the same, and they should be the same. It is because the system remains the same and only reference frame has been shifted. The magnitude should be the same but the direction vector will be different.

Both formulations are correct, because this is simply change of reference frame.

We can model the system as a Sun-Asteroid two-body problem, as the acceleration provided by the Jupiter is significantly smaller in the order of 10^{-10} km/s.



***** Changes made to the mission will not be reflected *****
 ***** in the data displayed until the mission is rerun *****

Propagate Command: Propagate1
 Spacecraft : Sat
 Coordinate System: EarthMJ2000Eq

Time System	Gregorian	Modified Julian
UTC Epoch:	19 Sep 2020 16:40:37.256	29112.1948756433
TAI Epoch:	19 Sep 2020 16:41:14.256	29112.1953038840
TT Epoch:	19 Sep 2020 16:41:46.440	29112.1956763840
TDB Epoch:	19 Sep 2020 16:41:46.438	29112.1956763655

Cartesian State	Keplerian State
X = 18030.291531473 km	SMA = 60005.551177493 km
Y = 22.781816125395 km	ECC = 0.6995224279922
Z = 24.319029367737 km	INC = 44.962697100807 deg
VX = -0.0113222174795 km/sec	RAAN = 359.99501442565 deg
VY = 4.3370803782779 km/sec	AOP = 0.1093612816155 deg
VZ = 4.3314356662285 km/sec	TA = 360.00000000000 deg
	MA = 360.00000000000 deg
	EA = 360.00000000000 deg

Spherical State	Other Orbit Data
RMAG = 18030.322324802 km	Mean Motion = 4.295183709e-05 deg/sec
RA = 0.0723949036935 deg	Orbit Energy = -3.3213630545694 km^2/s^2
DEC = 0.0772797148050 deg	C3 = -6.6427261091389 km^2/s^2
VMAG = 6.1295782343427 km/s	Semilatus Rectum = 30642.937174930 km
AZI = 45.037355084042 deg	Angular Momentum = 110518.27128029 km^2/s
VFPA = 90.000000001591 deg	Beta Angle = -1.1272751567757 deg
RAV = 90.149573848317 deg	Periapsis Altitude = 11652.186024802 km
DECV = 44.962592867022 deg	VelPeriapsis = 6.1295782343427 km/s
	VelApoapsis = 1.0837166694310 km/s
	Orbit Period = 146284.43699959 s

Planetodetic Properties

LST	= 0.3378541124879 deg
MHA	= 249.20061968500 deg
Latitude	= 0.1912789599837 deg
Longitude	= 111.13723442749 deg
Altitude	= 11652.186262165 km

Spacecraft Properties

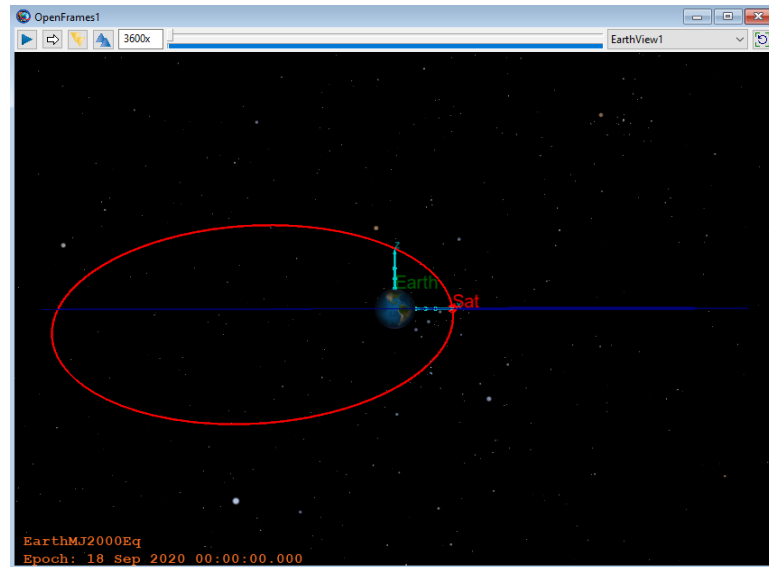
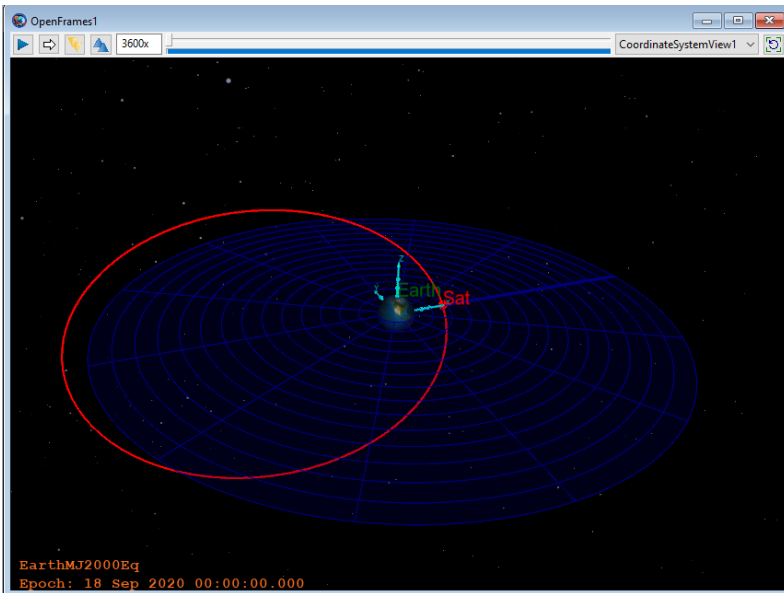
Cd	= 2.200000
Drag area	= 15.00000 m^2
Cr	= 1.800000
Reflective (SRP) area	= 1.000000 m^2
Dry mass	= 850.00000000000 kg
Total mass	= 850.00000000000 kg
SPADDragScaleFactor	= 1.000000
SPADSRPScaleFactor	= 1.000000

- i) Periapsis : 18030.32232 km
- ii) Apoapsis : 101980.6451 km
- iii) Energy : -3.321363055 MJ/kg
- iv) Semi-major axis : 60005.55118 km
- v) semi-latus rectum : 30642.93717 km
- vi) Angular momentum : 110518.2713 km^2/s

$$\begin{array}{lcl}
 v) X : 18000 & \text{km} & V_X = 0 \text{ km/s} \\
 Y : 0 & \text{km} & V_Y = 4.338524937 \text{ km/s} \\
 Z : 0 & \text{km} & V_Z = 4.338524937 \text{ km/s}
 \end{array}$$

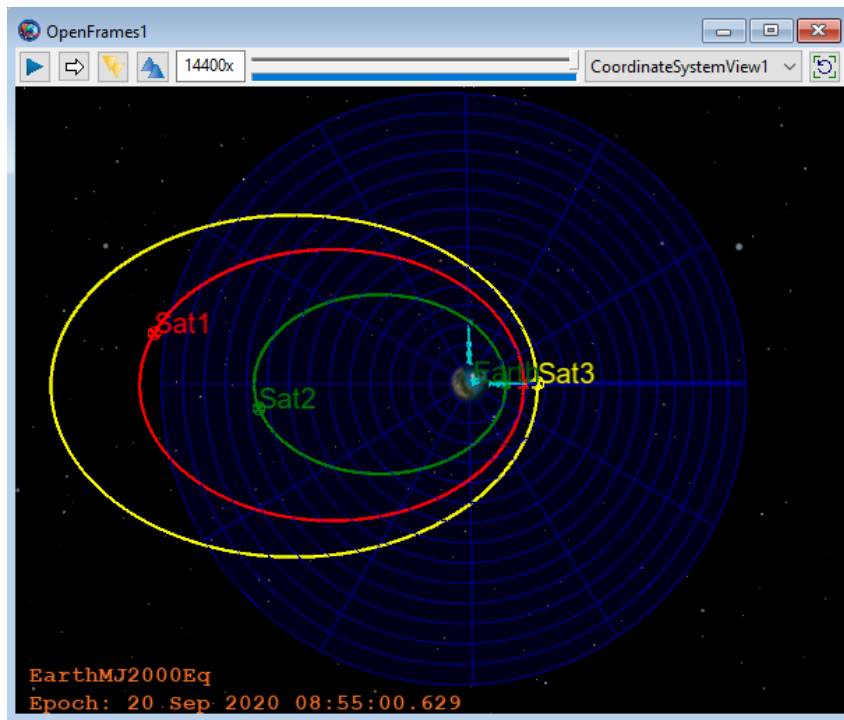
These values are associated with the EarthMJ2000Eq frame

Name	Origin	Axes	Description
EarthMJ2000Eq	Earth	MJ2000Eq	An Earth equator inertial system based on IAU-1976/FK5 theory with 1980 update to nutation.



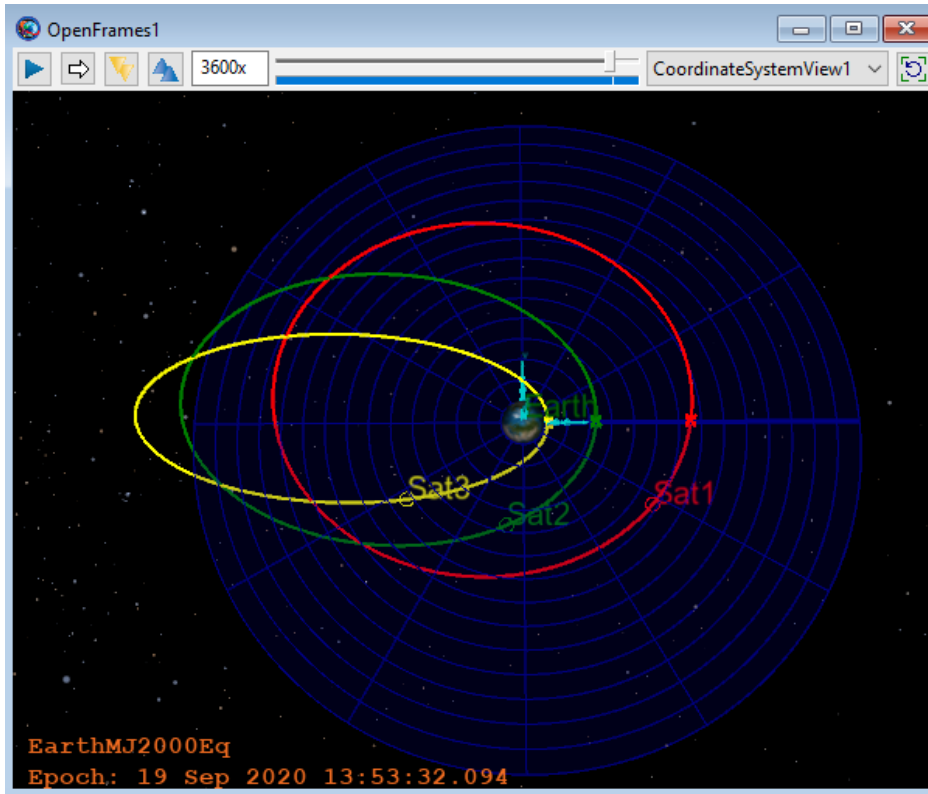
2 b)

case 1: $e=0.7$, $i=0^\circ$, $a=60000, 70000, 75000$

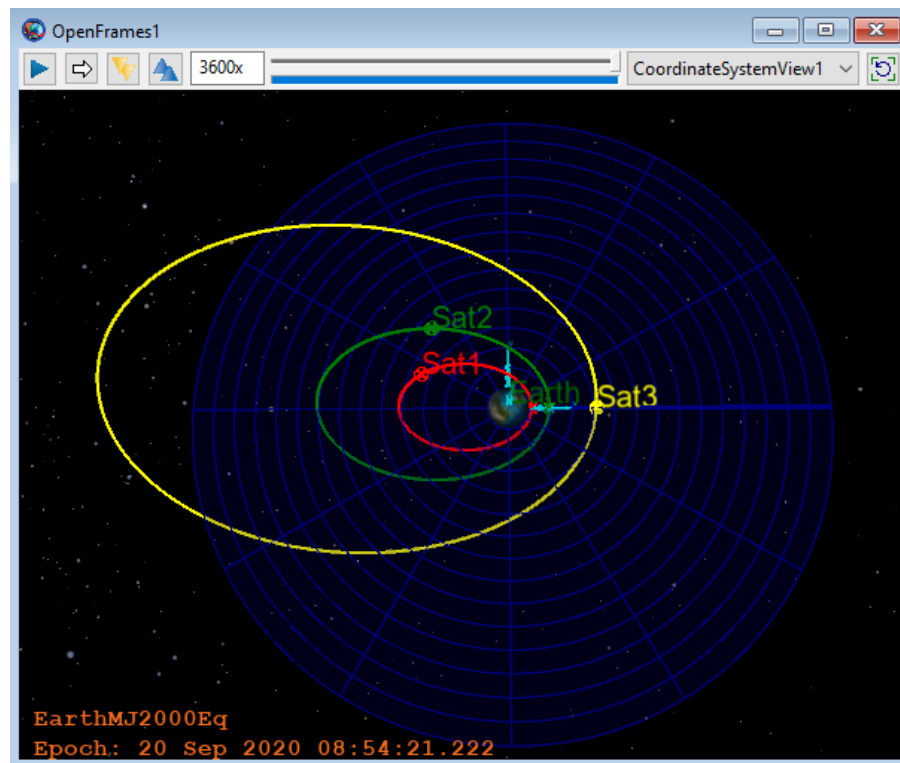


2 b)

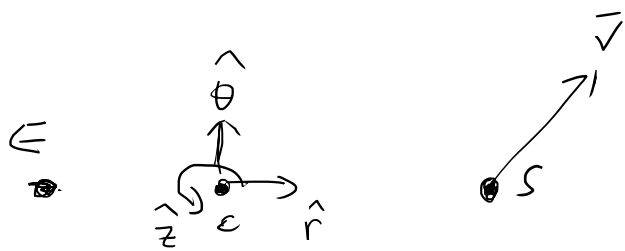
(case 2: $a = 60,000 \text{ km}$, $i = 95^\circ$, $e = 0.2, 0.65, 0.88$)



2 b)
case 3: $e = 0.65$, $i = 45^\circ$, $a = 20000, 35000, 75000 \text{ km}$



3.



$$m_s = 600 \text{ kg}$$

$$h = 8560 \text{ km}$$

$$\vec{v}_{SE} = 2.11 \hat{r} + 4.89 \hat{\theta} \text{ km/s}$$

radius of the Earth: 6357 km.

a) Find \vec{L}_3 , \vec{h} , T , L_4 , $\vec{\zeta}$, \dot{A}

$$\vec{r}_{ES} = 14917 \hat{r} \text{ km.}$$

$$\vec{r}_{EC} \cdot (m_E + m_s) = \vec{r}_{ES} \cdot m_s$$

$$\vec{r}_{EC} = \vec{r}_{ES} \frac{m_s}{m_E + m_s}$$

$$\vec{r}_{EC} = 0.1499 \times 10^{-7} \text{ km.}$$

This value is small enough that we can approximate the center of mass of the system is at the center of the Earth.

$$\text{Let } \vec{r}_{ES} = \vec{r}$$

$$\vec{L}_3 = \frac{m_E m_s}{m_E + m_s} (\vec{r} \times \dot{\vec{r}}) = \underline{4.3766 \times 10^7 \hat{z} \text{ kg km}^2/\text{s}}$$

$$\vec{h} = \vec{r} \times \dot{\vec{r}} = \underline{7.2944 \times 10^4 \hat{z} \text{ km}^2/\text{s}}$$

$$T = \frac{1}{2} \frac{m_E m_s}{m_E + m_s} (\dot{\vec{r}} \cdot \dot{\vec{r}}) = 8.5093 \times 10^3 \text{ kg km}^2/\text{s}^2 \text{ (MJ)}$$

$$U = \frac{G m_E m_S}{|\vec{r}|} = \underline{1.6033 \times 10^4 \text{ kg km}^2/\text{s}^2 \text{ (MJ)}}$$

$$C_\phi = T - U = \underline{-7.5235 \times 10^3 \text{ MJ}}$$

$$\xi = C_\phi \frac{(m_S + m_E)}{m_S \cdot m_E} = \underline{-12.5391 \text{ km}^2/\text{s}^2}$$

$$\dot{A} = \frac{h}{2} = \underline{3.6472 \times 10^4 \text{ km}^2/\text{s}}$$

b) The coefficient is $\frac{m_S + m_E}{m_S \cdot m_E} = 0.0017 \text{ kg}^{-1}$

c) $\mu = G(m_S + m_E) = \underline{3.9860 \times 10^5 \text{ km}^3/\text{s}^2}$

$$p = \frac{h^2}{\mu} = \underline{1.3349 \times 10^4 \text{ km}}$$

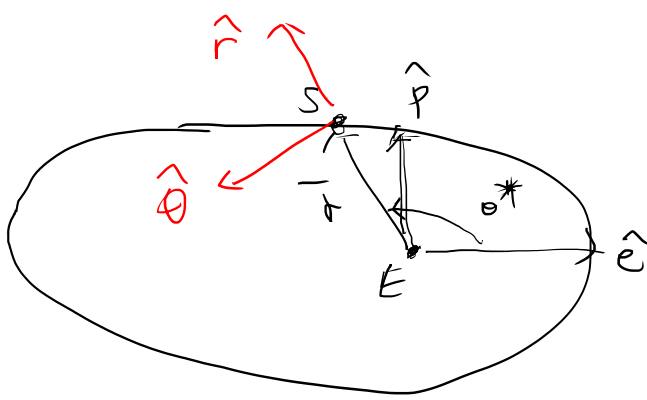
$$e = \sqrt{1 + 2 \frac{\xi h^2}{\mu^2}} = \underline{0.4002}$$

$$a = \frac{p}{(1 - e^2)} = \underline{1.5894 \times 10^4 \text{ km}}$$

$$P = 2\pi \sqrt{\frac{a^3}{\mu}} = \underline{1.9942 \times 10^4 \text{ s}}$$

$$\theta^* = \cos^{-1} \frac{\frac{1}{r} - \frac{\mu}{h^2}}{\sqrt{\frac{\mu^2}{h^4} + 2 \frac{\xi}{h^2}}} = \underline{1.6761 \text{ rad} = 96.0349^\circ}$$

$$\gamma = \tan^{-1} \left(\frac{4.89}{2.11} \right) = \underline{1.4742 \text{ rad} = 84.4642^\circ}$$



$$\hat{r} = \cos\theta^* \hat{e} + \sin\theta^* \hat{p} \quad {}^rR^i = \begin{bmatrix} \cos\theta^* & \sin\theta^* & 0 \\ -\sin\theta^* & \cos\theta^* & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$\hat{\theta} = -\sin\theta^* \hat{e} + \cos\theta^* \hat{p}$$

$${}^i\vec{r}_{SE} = {}^rR^{iT} \cdot {}^r\vec{r}_{SE}$$

$$= (-2.1568 \hat{e} + 1.4834 \hat{p}) \times 10^4 \text{ km}$$

$$d) \quad v_c = \sqrt{\frac{\mu}{r}} = 5.1693 \text{ km/s}$$

$$v = \|\dot{\vec{r}}\| = 5.3258 \text{ km/s}$$

v is bigger than v_c , but not bigger than

$\sqrt{2} v_c$, thus the spacecraft doesn't have the speed to escape the gravity pull from Earth.