

A New Mixed Potential Representation for Unsteady, Incompressible Flow*

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Abstract. We present a new integral representation for the unsteady, incompressible Stokes or Navier–Stokes equations, based on a linear combination of heat and harmonic potentials. For velocity boundary conditions, this leads to a coupled system of integral equations: one for the normal component of velocity and one for the tangential components. Each individual equation is well-conditioned, and we show that using them in predictor-corrector fashion, combined with spectral deferred correction, leads to high-order accuracy solvers. The fundamental unknowns in the mixed potential representation are densities supported on the boundary of the domain. We refer to one as the *vortex source*, the other as the *pressure source*, and to the coupled system as the *combined source integral equation*.

Key words. unsteady Stokes flow, Navier–Stokes equations, boundary integral equations, heat potentials, harmonic potentials, predictor-corrector method, mixed potential formulation, spectral deferred correction method

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1. Introduction. In this paper, we consider the numerical solution of the unsteady, incompressible Navier–Stokes equations,

$$(1.1) \quad \frac{\partial \mathbf{u}}{\partial t} = \nu \Delta \mathbf{u} - \nabla p - (\mathbf{u} \cdot \nabla) \mathbf{u} + \mathbf{f}, \quad \nabla \cdot \mathbf{u} = 0,$$

or their linearization, the unsteady Stokes equations,

$$(1.2) \quad \frac{\partial \mathbf{u}}{\partial t} = \nu \Delta \mathbf{u} - \nabla p + \mathbf{F}, \quad \nabla \cdot \mathbf{u} = 0.$$

The domain, which may be nonstationary, will be denoted at time $t \in [0, T]$ by $D(t)$ with boundary $\Gamma(t)$. For the sake of simplicity, we assume that $\Gamma(t) \in C^2$. The entire space-time domain will be denoted by D_T with boundary Γ_T . Here, $\mathbf{u}(\mathbf{x}, t)$ is the velocity field of interest, ν is the viscosity, and $p(\mathbf{x}, t)$ is the pressure at a point

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$\mathbf{x} \in D(t)$. In (1.1) and (1.2), \mathbf{f} and \mathbf{F} are forcing terms. Initial conditions for the velocity are given by

$$(1.3) \quad \mathbf{u}(\mathbf{x}, 0) = \mathbf{u}_0(\mathbf{x}), \quad \mathbf{x} \in D(0),$$

and we restrict our attention to the case where “velocity” boundary conditions are prescribed:

$$(1.4) \quad \mathbf{u}(\mathbf{x}, t) = \mathbf{g}(\mathbf{x}, t), \quad (\mathbf{x}, t) \in \Gamma_T.$$

We focus on the linearized problem, assuming that \mathbf{F} is known. From a numerical perspective, it already contains the essential difficulty faced by marching schemes for the Navier–Stokes equations, which usually treat the nonlinear, advective term explicitly. That essential difficulty concerns computing the evolution of a diffusing velocity field while maintaining the incompressibility condition

$$(1.5) \quad \nabla \cdot \mathbf{u} = 0$$

through the addition of a pressure gradient.

Beginning with the work of Chorin and Temam [14, 66], one of the most popular approaches for solving this problem is through the use of fractional step “projection” methods. A simple version of such a scheme involves first solving a diffusion equation for the velocity field with an explicit approximation of ∇p and \mathbf{F} , then solving a Poisson equation for the pressure to enforce the incompressibility constraint. Several decisions must be made in such schemes, including the choice of boundary conditions for the diffusion step and the choice of boundary conditions for the pressure correction/projection step. We do not seek to review the literature here and refer the reader to [11, 39, 49] for additional references and a more thorough discussion.

To avoid fractional steps, an alternative is to use a gauge method. Rather than solving the unsteady Stokes equations directly, one solves a system of the form

$$(1.6) \quad \begin{aligned} \frac{\partial \mathbf{m}}{\partial t} &= \nu \Delta \mathbf{m} + \mathbf{F}, \\ \Delta \phi &= \nabla \cdot \mathbf{m}, \end{aligned}$$

from which one obtains \mathbf{u} and p as

$$\begin{aligned} \mathbf{u} &= \mathbf{m} - \nabla \phi, \\ p &= \phi_t - \nu \Delta \phi. \end{aligned}$$

Such schemes require suitable boundary conditions for ϕ and \mathbf{m} , but avoid the fractional step and are more straightforward to discretize with high-order accuracy in time (see, for example, [12, 19, 24, 59, 68]).

One can also obtain an unconstrained formulation by taking the curl of the unsteady Stokes equations, yielding an equation for the evolution of vorticity $\boldsymbol{\omega} = \nabla \times \mathbf{u}$. In three dimensions, we have

$$(1.7) \quad \frac{\partial \boldsymbol{\omega}}{\partial t} = \nu \Delta \boldsymbol{\omega} + \nabla \times \mathbf{F},$$

while in two dimensions,

$$(1.8) \quad \frac{\partial \omega}{\partial t} = \nu \Delta \omega + \left(\frac{\partial F_2}{\partial x_1} - \frac{\partial F_1}{\partial x_2} \right).$$

Here, $\mathbf{u} = (u_1, u_2)$, $\mathbf{F} = (F_1, F_2)$, and vorticity is the scalar $\omega = \frac{\partial u_2}{\partial x_1} - \frac{\partial u_1}{\partial x_2}$. This approach is particularly natural in the two-dimensional setting, where one can introduce a scalar stream function Ψ , with

$$(1.9) \quad \mathbf{u} = \nabla^\perp \Psi = \left(\frac{\partial \Psi}{\partial x_2}, -\frac{\partial \Psi}{\partial x_1} \right),$$

so that the incompressibility constraint is automatically satisfied. It is easy to see that the stream function must satisfy the Poisson equation

$$(1.10) \quad \Delta \Psi = -\omega.$$

A major difficulty with this approach is that the boundary conditions for vorticity are nonlocal [2, 5, 20, 22, 57]. Instead, one can also formulate the two-dimensional Navier–Stokes equations entirely in terms of the stream function [4, 29, 40, 41]:

$$(1.11) \quad \frac{\partial \Delta \Psi}{\partial t} = \nu \Delta^2 \Psi - \left(\frac{\partial F_2}{\partial x_1} - \frac{\partial F_1}{\partial x_2} \right).$$

Since this is a fourth-order partial differential equation, one can directly impose velocity boundary conditions by specifying $\nabla^\perp \Psi$ on $\Gamma(t)$. Unfortunately, the extension of this approach to three dimensions is much more complicated (see, for example, [23]). Other interesting formulations of the Navier–Stokes equations in terms of nonphysical variables include the Clebsch representation [15, 46, 55], streamfunction formulations with auxiliary potentials [58, 61, 60], and methods based on “Eulerian–Lagrangian” potentials [18].

Finally, we should note that there is a Green’s function for the linearized equations (1.2), called the *unsteady Stokeslet*. In [42], integral equation methods were proposed using the corresponding layer potentials. While effective, they are somewhat complicated to implement with existing fast algorithms and high-order accurate quadrature methods. We will return to this point in the concluding section.

Here, we propose a new integral representation for the solution of the unsteady Stokes equations that is divergence-free by construction, involves only the use of harmonic and heat potentials, permits the natural imposition of velocity boundary conditions, and is applicable in either two or three dimensions. Since fast and high-order algorithms have been created for harmonic and heat potentials over the past several decades, powerful numerical machinery can immediately be brought to bear. The heart of our approach is to find a particular solution to the inhomogeneous equation (accounting for the forcing term \mathbf{F}), followed by a solution of the homogeneous, unsteady Stokes problem to enforce the desired boundary conditions. In three dimensions, the latter step involves a representation of the solution of the form

$$\begin{aligned} \mathbf{u}(\mathbf{x}, t) &= \nabla \phi(\mathbf{x}, t) + \nabla \times \mathbf{K}(\mathbf{x}, t), \\ p(\mathbf{x}, t) &= -\phi_t(\mathbf{x}, t), \end{aligned}$$

where ϕ is harmonic and \mathbf{K} satisfies the vector homogeneous heat equation. Both ϕ and \mathbf{K} will be defined in terms of layer potentials on Γ_T , whose source densities will be referred to as the pressure source and the vortex source, respectively. Enforcing velocity boundary conditions will lead to the *combined source integral equation*. An important feature of our approach is that the pressure and vortex sources, restricted to the boundary, are the only unknowns in the method.

The paper is organized as follows. In section 2, we briefly summarize the necessary mathematical background. In section 3, we discuss the mixed potential representation and derive the combined source integral equation. In section 4, we compute the spectrum and condition number of a fully implicit version of the combined source integral equations and in section 5, we present numerical experiments. In section 6, we investigate a kind of predictor-corrector scheme, where we impose the normal and tangential boundary conditions sequentially. In section 7, we show how high-order accuracy can be achieved using a spectral deferred correction scheme. We conclude with an outline of future work.

REMARK 1. *We assume that $\nu = 1$ in the remainder of this paper. This is easily accomplished in the unsteady Stokes equations by rescaling the time variable.*

2. Analytical Preliminaries. For a fixed domain D in \mathbb{R}^d with boundary Γ , we let $L^2(D)$ denote the space of all square integrable functions in D , and we let $L^2(\Gamma)$ denote the space of all square integrable functions on Γ . For the time-varying space-time cylinder $D_T \subset \mathbb{R}^d \times [0, T]$ with boundary Γ_T , we let $L^2(D_T)$ denote the space of all square integrable functions in D_T , and we let $L^2(\Gamma_T)$ denote the space of all square integrable functions on Γ_T . We briefly summarize the necessary aspects of classical potential theory for the Laplace and heat equations in \mathbb{R}^d ($d = 2, 3$). Excellent references for this material include [17, 34, 44, 52, 56].

2.1. Harmonic Potentials. The Green's function for the Laplace equation in free space is given by

$$(2.1) \quad G_L(\mathbf{x}, \mathbf{y}) = \begin{cases} -\frac{1}{2\pi} \ln |\mathbf{x} - \mathbf{y}| & \text{in } \mathbb{R}^2, \\ \frac{1}{4\pi |\mathbf{x} - \mathbf{y}|} & \text{in } \mathbb{R}^3. \end{cases}$$

DEFINITION 2.1. *Let $\mathbf{x} \in \mathbb{R}^d$. The single layer potential \mathcal{S}_L with density $\rho \in L^2(\Gamma)$ is defined by*

$$(2.2) \quad \mathcal{S}_L[\rho](\mathbf{x}) = \int_{\Gamma} G_L(\mathbf{x}, \mathbf{y}) \rho(\mathbf{y}) ds(\mathbf{y}).$$

The volume potential \mathcal{V}_L with density $f \in L^2(D)$ is defined by

$$(2.3) \quad \mathcal{V}_L[f](\mathbf{x}) = \int_D G_L(\mathbf{x}, \mathbf{y}) f(\mathbf{y}) d\mathbf{y}.$$

2.1.1. Jump Relations. For $\mathbf{x}_0 \in \Gamma$, the normal derivative of the single layer potential $\mathcal{S}_L[\rho]$ satisfies the jump relation

$$(2.4) \quad \lim_{\epsilon \rightarrow 0^+} \frac{\partial \mathcal{S}_L[\rho](\mathbf{x}_0 \pm \epsilon \nu(\mathbf{x}_0))}{\partial \nu(\mathbf{x}_0)} = \mp \frac{1}{2} \rho(\mathbf{x}_0) + \mathcal{S}_{L\nu}[\rho](\mathbf{x}_0),$$

where $\nu(\mathbf{x}_0)$ is the unit outward normal vector to Γ at the boundary point \mathbf{x}_0 and

$$(2.5) \quad \mathcal{S}_{L\nu}[\rho](\mathbf{x}_0) = \text{p.v.} \int_{\Gamma} \frac{\partial G_L(\mathbf{x}_0, \mathbf{y})}{\partial \nu(\mathbf{x}_0)} \rho(\mathbf{y}) ds(\mathbf{y}).$$

Since Γ is C^2 , both \mathcal{S}_L and $\mathcal{S}_{L\nu}^*$ are compact operators on $L^2(\Gamma)$ [17, 34, 44].

2.1.2. Tangential Derivatives. In two dimensions, the tangential derivative of the single layer potential is denoted by

$$(2.6) \quad \mathcal{S}_{L\tau}[\rho](\mathbf{x}_0) = \int_{\Gamma} \frac{\partial G_L(\mathbf{x}_0, \mathbf{y})}{\partial \tau(\mathbf{x}_0)} \rho(\mathbf{y}) ds(\mathbf{y}),$$

where $\tau(\mathbf{x}_0)$ is the unit tangential vector at $\mathbf{x}_0 \in \Gamma$. $\mathcal{S}_{L\tau}$ is defined in the Cauchy principal value sense and is a bounded operator on $L^2(\Gamma)$. In three dimensions, the tangential derivatives can be written in the form

$$(2.7) \quad \nu(\mathbf{x}_0) \times \nabla \mathcal{S}_L[\rho](\mathbf{x}_0).$$

This operator is, again, defined in the Cauchy principal value sense and bounded on $L^2(\Gamma)$.

2.2. Heat Potentials. The Green's function for the heat equation, $u_t = \Delta u$, is

$$(2.8) \quad G_H(\mathbf{x}, t) = \frac{1}{(4\pi t)^{d/2}} e^{-\frac{|\mathbf{x}|^2}{4t}}, \quad \mathbf{x} \in \mathbb{R}^d.$$

DEFINITION 2.2. Let $u_0 \in L^2(D(0))$, $f \in L^2(D_T)$, and $\mu \in L^2(\Gamma_T)$. Then the initial heat potential \mathcal{I}_H is defined by

$$(2.9) \quad \mathcal{I}_H[u_0](\mathbf{x}, t) = \int_D G_H(\mathbf{x} - \mathbf{y}, t) u_0(\mathbf{y}) d\mathbf{y},$$

the volume heat potential \mathcal{V}_H is defined by

$$(2.10) \quad \mathcal{V}_H[g](\mathbf{x}, t) = \int_0^t \int_{D(t')} G_H(\mathbf{x} - \mathbf{y}, t - t') f(\mathbf{y}, t') d\mathbf{y} dt',$$

and the single layer heat potential \mathcal{S}_H is defined by

$$(2.11) \quad \mathcal{S}_H[\mu](\mathbf{x}, t) = \int_0^t \int_{\Gamma(t')} G_H(\mathbf{x} - \mathbf{y}, t - t') \mu(\mathbf{y}, t') ds(\mathbf{y}) dt'.$$

2.2.1. Jump Relations. It is well known that the initial heat potential \mathcal{I}_H is a compact operator on $L^2(D(0))$ and that the volume heat potential \mathcal{V}_H is a compact operator on $L^2(D_T)$. As in the harmonic case, the normal derivative of the single layer heat potential $\mathcal{S}_{H\nu}[\mu]$ satisfies the jump relations [56]

$$(2.12) \quad \lim_{\epsilon \rightarrow 0^+} \mathcal{S}_{H\nu}[\mu](\mathbf{x}_0 \pm \epsilon \nu(\mathbf{x}_0), t) = \mp \frac{1}{2} \mu(\mathbf{x}_0, t) + \mathcal{S}_{H\nu}[\mu](\mathbf{x}_0, t), \quad \mathbf{x}_0 \in \Gamma,$$

where $\mathcal{S}_{H\nu}[\mu](\mathbf{x}_0, t)$ is the principal value of

$$\mathcal{S}_{H\nu}[\mu](\mathbf{x}_0, t) = \int_0^t \int_{\Gamma(t')} \frac{\partial G_H(\mathbf{x}_0 - \mathbf{y}, t - t')}{\partial \nu(\mathbf{x}_0)} \mu(\mathbf{y}, t') ds(\mathbf{y}) dt'.$$

In two dimensions, the tangential derivative of the single layer heat potential is denoted by

$$(2.13) \quad \mathcal{S}_{H\tau}[\mu](\mathbf{x}_0, t) = \int_0^t \int_{\Gamma(t')} \frac{\partial G_H(\mathbf{x}_0 - \mathbf{y}, t - t')}{\partial \tau(\mathbf{x}_0)} \mu(\mathbf{y}, t') ds(\mathbf{y}) dt',$$

where $\tau(\mathbf{x}_0)$ is the unit tangential vector at $\mathbf{x}_0 \in \Gamma$. $\mathcal{S}_{H\tau}$ is defined in the Cauchy principal value sense and is a bounded operator on $L^2(\Gamma)$.

2.2.2. The Vector Heat Potential. In three dimensions, we will also make use of a vector heat potential, defined by

$$(2.14) \quad \mathbf{K}_H[\mathbf{J}](\mathbf{x}, t) = \int_0^t \int_{\Gamma(t')} G_H(\mathbf{x} - \mathbf{y}, t - t') \mathbf{J}(\mathbf{y}, t') ds(\mathbf{y}) dt',$$

where \mathbf{J} is a tangential vector field on Γ .

DEFINITION 2.3. *For reasons that will become clear below, we will refer to \mathbf{J} or μ as the vortex source. ($\mathbf{K}_H[\mathbf{J}]$ will play a role analogous to that of the vector potential in electromagnetic theory, where the source is a surface electric current.)*

It is straightforward to verify that, as in the electromagnetic case, the tangential components of $\nabla \times \mathbf{K}_H[\mathbf{J}]$ are given by [16]

$$(2.15) \quad \lim_{\epsilon \rightarrow 0^+} \nu(\mathbf{x}_0) \times (\nabla \times \mathbf{K}_H)[\mathbf{J}](\mathbf{x}_0 \pm \epsilon \nu(\mathbf{x}_0), t) = \pm \frac{1}{2} \mathbf{J}(\mathbf{x}_0, t) + \mathbf{M}_H[\mathbf{J}](\mathbf{x}_0, t),$$

where

$$(2.16) \quad \mathbf{M}_H[\mathbf{J}](\mathbf{x}_0, t) = \nu(\mathbf{x}_0) \times (\nabla \times \mathbf{K}_H)[\mathbf{J}](\mathbf{x}_0, t), \quad \mathbf{x}_0 \in \Gamma,$$

interpreted in the principal value sense. $\mathbf{M}_H[\mathbf{J}]$ is a compact operator when Γ is C^1 [16]. Finally,

$$(2.17) \quad \nu(\mathbf{x}_0) \cdot \nabla \times \mathbf{K}_H[\mathbf{J}](\mathbf{x}_0, t)$$

is defined in the Cauchy principal value sense, continuous across the boundary of $\Gamma \in C^1$, and bounded on $L^2(\Gamma)$.

2.3. The Helmholtz Decomposition of a Vector Field. Let $D \subset \mathbb{R}^d$ be a bounded domain ($d = 2, 3$). It is well known [28] that every vector field $\mathbf{F} \in L^2(D)$ has a decomposition of the form

$$(2.18) \quad \mathbf{F} = \nabla \phi + \mathbf{w},$$

where \mathbf{w} is divergence-free (or *solenoidal*) and $\nabla \phi$ is curl-free (or *irrotational*). We will sometimes write

$$(2.19) \quad \mathbf{F} = \mathbf{F}_G + \mathbf{F}_S$$

instead of (2.18), where \mathbf{F}_G is irrotational and \mathbf{F}_S is solenoidal.

Without boundary conditions on \mathbf{w} or ϕ , the Helmholtz decomposition is not unique. Nevertheless, assuming \mathbf{F} is sufficiently smooth, a simple explicit construction is easily computed.

LEMMA 2.4 ([3]). *Let \mathbf{F} be a twice differentiable vector field in a domain D with boundary Γ in \mathbb{R}^3 , and let*

$$\begin{aligned} \phi(\mathbf{x}) &= - \int_D G_L(\mathbf{x} - \mathbf{y}) (\nabla_{\mathbf{y}} \cdot \mathbf{F}(\mathbf{y})) d\mathbf{y} + \int_{\Gamma} G_L(\mathbf{x} - \mathbf{y}) (\nu(\mathbf{y}) \cdot \mathbf{F}(\mathbf{y})) ds(\mathbf{y}), \\ \mathbf{A}(\mathbf{x}) &= \int_D G_L(\mathbf{x} - \mathbf{y}) (\nabla_{\mathbf{y}} \times \mathbf{F}(\mathbf{y})) d\mathbf{y} - \int_{\Gamma} G_L(\mathbf{x} - \mathbf{y}) (\nu(\mathbf{y}) \times \mathbf{F}(\mathbf{y})) ds(\mathbf{y}). \end{aligned}$$

Then

$$\mathbf{F} = \nabla \times \mathbf{A} + \nabla \phi.$$

In \mathbb{R}^2 , if $\mathbf{F} = (F_1, F_2)$ is twice differentiable in a domain D with boundary Γ , let ϕ be defined as above and let

$$\psi(\mathbf{x}) = \int_D G_L(\mathbf{x} - \mathbf{y}) (\nabla_{\mathbf{y}}^\perp \cdot \mathbf{F}(\mathbf{y})) d\mathbf{y} - \int_\Gamma G_L(\mathbf{x} - \mathbf{y}) (\tau(\mathbf{y}) \cdot \mathbf{F}(\mathbf{y})) ds(\mathbf{y}),$$

where $\nabla_{\mathbf{x}}^\perp \cdot \mathbf{F}(\mathbf{x}) = \frac{\partial F_2}{\partial x_1} - \frac{\partial F_1}{\partial x_2}$ and τ denotes the unit tangent vector along Γ . Then

$$\mathbf{F} = \nabla^\perp \psi + \nabla \phi.$$

Using the notation above, we can write this more compactly as

$$\begin{aligned} \phi(\mathbf{x}) &= -\mathcal{V}_L[\nabla \cdot \mathbf{F}](\mathbf{x}) + \mathcal{S}_L[\nu \cdot \mathbf{F}](\mathbf{x}), \\ \mathbf{A}(\mathbf{x}) &= \mathcal{V}_L[\nabla \times \mathbf{F}](\mathbf{x}) - \mathcal{S}_L[\nu \times \mathbf{F}](\mathbf{x}), \\ \psi(\mathbf{x}) &= \mathcal{V}_L[\nabla^\perp \cdot \mathbf{F}](\mathbf{x}) - \mathcal{S}_L[\tau \cdot \mathbf{F}](\mathbf{x}). \end{aligned} \quad (2.20)$$

Both the harmonic volume potentials and the harmonic single layer potentials can be computed in optimal time, and with high-order accuracy, using the fast multipole method and suitable quadrature rules [9, 13, 25, 30, 32, 36, 47, 51, 64].

REMARK 2. In free space, there is an even simpler construction for the Helmholtz decomposition (assuming sufficiently rapid decay of \mathbf{F}).

LEMMA 2.5 ([45]). If $\mathbf{F} \in L^2(\mathbb{R}^d)$, then

$$\mathbf{F}_G = -\nabla \left(\nabla \cdot \int_{\mathbb{R}^d} G_L(\mathbf{x} - \mathbf{y}) \mathbf{F}(\mathbf{y}) d\mathbf{y} \right), \quad \mathbf{F}_S = \mathbf{F} - \mathbf{F}_G,$$

where G_L is the Green's function for the Laplace equation.

3. Potential Theory for the Unsteady Stokes Equations. Before turning to the full boundary value problem, it is worth stating a fundamental, but rarely used, fact about the unsteady Stokes equations in the absence of physical boundaries.

LEMMA 3.1 ([45, Chapter 4]). Let $\mathbf{F}(\mathbf{x}, t) \in L^2(\mathbb{R}^d)$, where $d = 2, 3$, with the Helmholtz decomposition

$$\mathbf{F}(\mathbf{x}, t) = \mathbf{F}_S(\mathbf{x}, t) + \mathbf{F}_G(\mathbf{x}, t),$$

where \mathbf{F}_S is solenoidal and \mathbf{F}_G is irrotational. Then the solution to (1.2) in \mathbb{R}^d with divergence-free initial data $\mathbf{u}_0(\mathbf{x})$ is given by

$$\begin{aligned} \mathbf{u}^{(F)}(\mathbf{x}, t) &= \mathcal{I}_H[\mathbf{u}_0](\mathbf{x}, t) + \mathcal{V}_H[\mathbf{F}_S](\mathbf{x}, t), \\ \nabla p^{(F)}(\mathbf{x}, t) &= \mathbf{F}_G(\mathbf{x}, t). \end{aligned} \quad (3.1)$$

(The operators \mathcal{I}_H and \mathcal{V}_H are assumed to be defined on \mathbb{R}^d rather than a bounded domain D .)

In short, given the Helmholtz decomposition of the forcing term \mathbf{F} , the unsteady Stokes equations have an explicit solution in free space by quadrature. This turns out to be true in a bounded domain as well.

LEMMA 3.2. Let $\mathbf{F}(\mathbf{x}, t) \in L^2(D)$, where $d = 2, 3$, with the Helmholtz decomposition

$$\mathbf{F}(\mathbf{x}, t) = \mathbf{F}_S(\mathbf{x}, t) + \nabla \phi(\mathbf{x}, t),$$

where $\mathbf{F}_S = \nabla \times \mathbf{A}$ in \mathbb{R}^3 and $\mathbf{F}_S = \nabla^\perp \psi$ in \mathbb{R}^2 . Then a particular solution to (1.2), (1.3) is given by

$$\begin{aligned} \mathbf{u}^{(F)}(\mathbf{x}, t) &= \mathcal{I}_H[\mathbf{u}_0](\mathbf{x}, t) + \nabla \times \mathcal{V}_H[\mathbf{A}](\mathbf{x}, t) && \text{in } \mathbb{R}^3, \\ \nabla p^{(F)}(\mathbf{x}, t) &= \nabla \phi(\mathbf{x}, t), \\ (3.2) \quad \mathbf{u}^{(F)}(\mathbf{x}, t) &= \mathcal{I}_H[\mathbf{u}_0](\mathbf{x}, t) + \nabla^\perp \mathcal{V}_H[\psi](\mathbf{x}, t) && \text{in } \mathbb{R}^2, \\ \nabla p^{(F)}(\mathbf{x}, t) &= \nabla \phi(\mathbf{x}, t), \end{aligned}$$

where the initial and volume heat potentials are given in Definition 2.2.

Proof. It is straightforward to verify that the partial differential equation (1.2) is satisfied. The fact that $\mathbf{u}^{(F)}(\mathbf{x}, t)$ is divergence-free follows immediately from Lemma 3.1 for the term $\mathcal{I}_H[\mathbf{u}_0](\mathbf{x}, t)$ and by construction for the term involving $\mathcal{V}_H[\mathbf{A}](\mathbf{x}, t)$ or $\mathcal{V}_H[\psi](\mathbf{x}, t)$. \square

Thus, from the preceding lemma, we may represent the solution to the full unsteady Stokes equations in the form

$$\mathbf{u} = \mathbf{u}^{(F)} + \mathbf{u}^{(B)}, \quad \nabla p = \nabla p^{(F)} + \nabla p^{(B)},$$

where $(\mathbf{u}^{(B)}, \nabla p^{(B)})$ satisfy the homogeneous, linearized equations

$$\begin{aligned} (3.3) \quad \frac{\partial \mathbf{u}^{(B)}}{\partial t} &= \Delta \mathbf{u}^{(B)} - \nabla p^{(B)}, && (\mathbf{x}, t) \in D_T, \\ \nabla \cdot \mathbf{u}^{(B)} &= 0, && (\mathbf{x}, t) \in D_T, \\ \mathbf{u}^{(B)}(\mathbf{x}, 0) &= \mathbf{0}, && \mathbf{x} \in D(0), \\ \mathbf{u}^{(B)}(\mathbf{x}, t) &= \tilde{\mathbf{g}}(\mathbf{x}, t) := \mathbf{g}(\mathbf{x}, t) - \mathbf{u}^{(F)}(\mathbf{x}, t), && (\mathbf{x}, t) \in \Gamma_T. \end{aligned}$$

There is a significant advantage in solving the homogeneous equations (3.3) rather than (1.2)–(1.4), as we shall now see.

3.1. The Mixed Potential Representation. Let us represent the solution to the homogeneous system,

$$(\mathbf{u}^{(B)}(\mathbf{x}, t), p^{(B)}(\mathbf{x}, t)),$$

in terms of harmonic and heat layer potentials. In three dimensions, we define

$$\begin{aligned} (3.4) \quad \mathbf{u}^{(B)}(\mathbf{x}, t) &= \nabla \mathcal{S}_L[\rho](\mathbf{x}, t) + \nabla \times \mathbf{K}_H[\mathbf{J}](\mathbf{x}, t), \\ p^{(B)}(\mathbf{x}, t) &= -\frac{\partial}{\partial t} \mathcal{S}_L[\rho](\mathbf{x}, t), \end{aligned}$$

while in two dimensions, we define

$$\begin{aligned} (3.5) \quad \mathbf{u}^{(B)}(\mathbf{x}, t) &= \nabla \mathcal{S}_L[\rho](\mathbf{x}, t) + \nabla^\perp \mathcal{S}_H[\mu](\mathbf{x}, t), \\ p^{(B)}(\mathbf{x}, t) &= -\frac{\partial}{\partial t} \mathcal{S}_L[\rho](\mathbf{x}, t). \end{aligned}$$

Here, ρ , \mathbf{J} , and μ are unknown *boundary densities* to be determined. It is straightforward to verify that the representations (3.4) and (3.5) satisfy the first three equations in (3.3).

DEFINITION 3.3. *Because of the preceding relations, we will refer to ρ as the pressure source or pressure source density.*

3.1.1. The Combined Source Integral Equation. If we decompose the velocity field into a sum of normal and tangential components on the boundary, then imposing velocity boundary conditions leads, in two dimensions, to the system of integral equations

$$(3.6) \quad \begin{aligned} \frac{1}{2}\rho(\mathbf{x}, t) + \mathcal{S}_{L\nu}[\rho](\mathbf{x}, t) + \mathcal{S}_{H\tau}[\mu](\mathbf{x}, t) &= \nu \cdot \tilde{\mathbf{g}}(\mathbf{x}, t), \\ \frac{1}{2}\mu(\mathbf{x}, t) + \mathcal{S}_{H\nu}[\mu](\mathbf{x}, t) - \mathcal{S}_{L\tau}[\rho](\mathbf{x}, t) &= -\tau \cdot \tilde{\mathbf{g}}(\mathbf{x}, t) \end{aligned}$$

for the unknowns ρ and μ , where \mathbf{x} is a point on the boundary $\Gamma(t)$.

In three dimensions, we obtain the system of integral equations

$$(3.7) \quad \begin{aligned} \frac{1}{2}\rho(\mathbf{x}, t) + \mathcal{S}_{L\nu}[\rho](\mathbf{x}, t) + \nu(\mathbf{x}, t) \cdot \nabla \times \mathbf{K}_H[\mathbf{J}](\mathbf{x}, t) &= \nu(\mathbf{x}, t) \cdot \tilde{\mathbf{g}}(\mathbf{x}, t), \\ \frac{1}{2}\mathbf{J}(\mathbf{x}, t) + \mathbf{M}_H[\mathbf{J}](\mathbf{x}, t) + \nu(\mathbf{x}, t) \times \nabla \mathcal{S}_L[\rho](\mathbf{x}, t) &= \nu(\mathbf{x}, t) \times \tilde{\mathbf{g}}(\mathbf{x}, t) \end{aligned}$$

for the unknowns ρ and \mathbf{J} , where \mathbf{x} is a point on the boundary $\Gamma(t)$. We will refer to either (3.6) or (3.7) as the *combined source integral equation* (CSIE).

One major advantage of the mixed potential representation is that the unknown densities (the pressure source and the vortex source) correspond to physical quantities of interest. The harmonic potential $\phi(\mathbf{x}, t) = \mathcal{S}_L[\rho](\mathbf{x}, t)$ determines the pressure, according to (3.4) and (3.5), while the heat potentials determine the vorticity. More precisely, in three dimensions,

$$\boldsymbol{\omega}(\mathbf{x}, t) = \nabla \times (\nabla \times \mathbf{K}_H)[\mathbf{J}](\mathbf{x}, t) + \nabla \times \mathbf{u}^{(F)},$$

while in two dimensions,

$$\omega(\mathbf{x}, t) = -\partial_t \mathcal{S}_H[\mu](\mathbf{x}, t) + \frac{\partial u_2^{(F)}}{\partial x_1} - \frac{\partial u_1^{(F)}}{\partial x_2},$$

where $\mathbf{u}^{(F)} = (u_1^{(F)}, u_2^{(F)})$. These relations may be of some direct interest in analysis.

REMARK 3. *It is, perhaps, worth noting that in the mixed potential representation, the boundary conditions for \mathbf{K} and ϕ are local, but they yield exact, nonlocal expressions for the pressure and vorticity through the formulae above.*

3.2. Discretization. For the sake of simplicity, we restrict our attention to the integral equation system (3.6) in two dimensions, and begin by semidiscretization in time (i.e., discretization with respect to the time variable alone). For this, we let

$$\rho_j = \rho(\mathbf{x}, j\Delta t), \quad \mu_j = \mu(\mathbf{x}, j\Delta t),$$

$$\boldsymbol{\rho}_j = [\rho_0, \dots, \rho_j], \quad \boldsymbol{\mu}_j = [\mu_0, \dots, \mu_j].$$

We then write

$$\mathcal{S}_{H\nu}[\boldsymbol{\mu}_j](\mathbf{x}, t) = \mathcal{S}_{H\nu}^{\text{far}}[\boldsymbol{\mu}_{j-1}](\mathbf{x}, t) + \mathcal{S}_{H\nu}^{\text{loc}}[\boldsymbol{\mu}_j](\mathbf{x}, t),$$

$$\mathcal{S}_{H\tau}[\boldsymbol{\mu}_j](\mathbf{x}, t) = \mathcal{S}_{H\tau}^{\text{far}}[\boldsymbol{\mu}_{j-1}](\mathbf{x}, t) + \mathcal{S}_{H\tau}^{\text{loc}}[\boldsymbol{\mu}_j](\mathbf{x}, t),$$

to denote the semidiscrete approximations of $\mathcal{S}_{H\nu}[\mu]$ and $\mathcal{S}_{H\tau}[\mu]$, where

$$\begin{aligned}
 \mathcal{S}_{H\nu}^{\text{far}}[\mu_{j-1}](\mathbf{x}, t) &= \sum_{l=1}^{j-1} \int_{(l-1)\Delta t}^{l\Delta t} \int_{\Gamma(t')} \frac{\partial G_H(\mathbf{x} - \mathbf{y}, t - t')}{\partial \nu(\mathbf{x})} P_k^I[\mu_l](\mathbf{y}, t') ds(\mathbf{y}) dt', \\
 \mathcal{S}_{H\tau}^{\text{far}}[\mu_{j-1}](\mathbf{x}, t) &= \sum_{l=1}^{j-1} \int_{(l-1)\Delta t}^{l\Delta t} \int_{\Gamma(t')} \frac{\partial G_H(\mathbf{x} - \mathbf{y}, t - t')}{\partial \tau(\mathbf{x})} P_k^I[\mu_l](\mathbf{y}, t') ds(\mathbf{y}) dt', \\
 \mathcal{S}_{H\nu}^{\text{loc}}[\mu_j](\mathbf{x}, t) &= \int_{(j-1)\Delta t}^{j\Delta t} \int_{\Gamma(t')} \frac{\partial G_H(\mathbf{x} - \mathbf{y}, t - t')}{\partial \nu(\mathbf{x})} P_k^I[\mu_j](\mathbf{y}, t') ds(\mathbf{y}) dt', \\
 \mathcal{S}_{H\tau}^{\text{loc}}[\mu_j](\mathbf{x}, t) &= \int_{(j-1)\Delta t}^{j\Delta t} \int_{\Gamma(t')} \frac{\partial G_H(\mathbf{x} - \mathbf{y}, t - t')}{\partial \tau(\mathbf{x})} P_k^I[\mu_j](\mathbf{y}, t') ds(\mathbf{y}) dt'.
 \end{aligned}
 \tag{3.8}$$

Here, $P_k^I[\mu_l]$ is the k th-order Lagrange interpolant of the data

$$\{\mu_l, \mu_{l-1}, \dots, \mu_{l-k}\}$$

at the $(k+1)$ uniformly spaced time points $\{l\Delta t, (l-1)\Delta t, \dots, (l-k)\Delta t\}$.

Note that we have separated out the contributions to $\mathcal{S}_{H\nu}$ and $\mathcal{S}_{H\tau}$ from the early time steps ($\mathcal{S}_{H\nu}^{\text{far}}, \mathcal{S}_{H\tau}^{\text{far}}$) from the contributions on the most recent time interval ($\mathcal{S}_{H\nu}^{\text{loc}}, \mathcal{S}_{H\tau}^{\text{loc}}$). The superscript I in the expression $P_k^I[\mu_l]$ indicates that the latest time point $t_l = l\Delta t$ is being used in the polynomial interpolant. Since $\mathcal{S}_{H\nu}^{\text{loc}}$ and $\mathcal{S}_{H\tau}^{\text{loc}}$ are linear operators acting on the densities $\{\mu_j, \mu_{j-1}, \dots, \mu_{j-k}\}$, we will also have occasion to write

$$\begin{aligned}
 \mathcal{S}_{H\nu}^{\text{loc}}[\mu_j](\mathbf{x}, j\Delta t) &= \left[A_{H\nu}^{j,k} \mu_{j-k} + \dots + A_{H\nu}^{j,1} \mu_{j-1} \right] + A_{H\nu}^{j,0} \mu_j = F_{H\nu}^{j,k}[\mu_{j-1}] + A_{H\nu}^{j,0} \mu_j, \\
 \mathcal{S}_{H\tau}^{\text{loc}}[\mu_j](\mathbf{x}, j\Delta t) &= \left[A_{H\tau}^{j,k} \mu_{j-k} + \dots + A_{H\tau}^{j,1} \mu_{j-1} \right] + A_{H\tau}^{j,0} \mu_j = F_{H\tau}^{j,k}[\mu_{j-1}] + A_{H\tau}^{j,0} \mu_j.
 \end{aligned}
 \tag{3.9}$$

This makes explicit the contributions of the densities at the various time steps to the local potentials $\mathcal{S}_{H\nu}^{\text{loc}}$ and $\mathcal{S}_{H\tau}^{\text{loc}}$.

In the present paper, following previous work with the unsteady Stokeslet [42], we interchange the order of integration in space and time, as proposed in [48], and carry out the time integration analytically. For $k=1$, using an implicit interpolation rule, P_1^I , to achieve second-order accuracy in Δt , the kernels of the spatial operators $A_{H\nu}^{j,0}$ and $A_{H\tau}^{j,0}$ are

$$\begin{aligned}
 G_{H\nu}^{\text{loc}}(\mathbf{x}, \mathbf{y}) &= -\frac{(\mathbf{x} - \mathbf{y}) \cdot \nu}{2\pi \|\mathbf{x} - \mathbf{y}\|^2} e^{-\frac{\|\mathbf{x} - \mathbf{y}\|^2}{4\Delta t}} + \frac{(\mathbf{x} - \mathbf{y}) \cdot \nu}{8\pi \Delta t} E_1\left(\frac{\|\mathbf{x} - \mathbf{y}\|^2}{4\Delta t}\right), \\
 G_{H\tau}^{\text{loc}}(\mathbf{x}, \mathbf{y}) &= -\frac{(\mathbf{x} - \mathbf{y}) \cdot \tau}{2\pi \|\mathbf{x} - \mathbf{y}\|^2} e^{-\frac{\|\mathbf{x} - \mathbf{y}\|^2}{4\Delta t}} + \frac{(\mathbf{x} - \mathbf{y}) \cdot \tau}{8\pi \Delta t} E_1\left(\frac{\|\mathbf{x} - \mathbf{y}\|^2}{4\Delta t}\right),
 \end{aligned}
 \tag{3.10}$$

where $E_1(x) = \int_x^\infty \frac{e^{-t}}{t} dt$ is the exponential integral function [54].

We will have occasion to make use of the explicit form of the interpolant as well.

DEFINITION 3.4. $P_k^E[\mu_{j-1}](t)$ is defined to be the k th-order Lagrange extrapolant of the data

$$[\mu_{j-1}, \dots, \mu_{j-(k+1)}]$$

evaluated at $t = j\Delta t$.

As in multistep methods for ordinary differential equations, when the interpolation order $k > 1$, some care is required in initialization—that is, computing the first $k - 1$ time steps with sufficient accuracy. We will ignore this issue for the moment to avoid distractions.

Finally, the spatial integrals in $\mathcal{S}_{H\tau}^{\text{loc}}$, $\mathcal{S}_{H\nu}^{\text{loc}}$, $\mathcal{S}_{L\tau}$, and $\mathcal{S}_{L\nu}$ involve either logarithmic singularities or principal value type integrals. We use the quadrature schemes of [1] to discretize these integrals to 16th-order accuracy on smooth curves.

3.3. History Dependence and Fast Algorithms. From a practical perspective, $\mathcal{S}_{H\tau}^{\text{far}}$ and $\mathcal{S}_{H\nu}^{\text{far}}$ clearly depend on the entire space-time history of the problem at hand. In the absence of suitable algorithms, the cost of their evaluation would be prohibitive. Fortunately, a number of fast algorithms have been developed for precisely this purpose [31, 33, 50, 65] that permit their evaluation in $O(NM \log M)$ work rather than $O(N^2 M^2)$ work, where N denotes the number of time steps and M denotes the number of points in the discretization of the boundary. We refer the reader to those papers for further details.

4. Spectrum of the Fully Implicit Combined Source Integral Equation. For the sake of simplicity, we restrict our attention to the coupled integral equations (3.6) in two dimensions on a stationary boundary. While at first glance this might appear to involve a compact perturbation of the identity, that is not the case. Informally, this can be seen as follows. First, we note that the operators $\mathcal{S}_{L\tau}$ and $\mathcal{S}_{H\tau}$ are compact perturbations of the operator $-\frac{1}{2}H$, where H denotes the Hilbert transform operator on the circle with perimeter L (L here is the length of Γ [44]):

$$H[f](s) = \frac{1}{2\pi} p.v. \int_0^L \cot\left(\frac{\pi(s-s')}{L}\right) f(s') ds'.$$

Thus, the system can be written in the form

$$(4.1) \quad \begin{bmatrix} \frac{1}{2}I & \frac{1}{2}H \\ -\frac{1}{2}H & \frac{1}{2}I \end{bmatrix} \begin{bmatrix} \rho \\ \mu \end{bmatrix} + \mathbf{C} \begin{bmatrix} \rho \\ \mu \end{bmatrix} = \begin{bmatrix} \nu \cdot \tilde{\mathbf{g}}(\mathbf{x}, t) \\ -\tau \cdot \tilde{\mathbf{g}}(\mathbf{x}, t) \end{bmatrix},$$

where \mathbf{C} is compact. Since $H^2 = -I$, the determinant of the leading part of the system vanishes, so that the coupled system fails to be a Fredholm equation of the second kind.

We now study the spectrum of the integral equation in detail when Γ is a circle of radius r . The resulting properties follow qualitatively for any smooth curve. More specifically, let us assume we are using a second-order accurate one-step implicit marching scheme, as described in subsection 3.2. This yields an integral equation at the j th time step of the form

$$(4.2) \quad \begin{bmatrix} \frac{1}{2}I + \mathcal{S}_{L\nu} & A_{H\tau}^{j,0} \\ -\mathcal{S}_{L\tau} & \frac{1}{2}I + A_{H\nu}^{j,0} \end{bmatrix} \begin{bmatrix} \rho_j \\ \mu_j \end{bmatrix} = \begin{bmatrix} \nu \cdot \tilde{\mathbf{g}} - \mathcal{S}_{H\tau}^{\text{far}}[\mu_{j-1}] - F_{H\tau}^{j,1}[\mu_{j-1}] \\ -\tau \cdot \tilde{\mathbf{g}} - \mathcal{S}_{H\nu}^{\text{far}}[\rho_{j-1}] - F_{H\nu}^{j,1}[\rho_{j-1}] \end{bmatrix}$$

using the notation of subsection 3.2.

The kernels of the operators on the left-hand side of (4.2) are given by

$$(4.3) \quad G_{L\nu}(\mathbf{x}, \mathbf{y}) = -\frac{(\mathbf{x} - \mathbf{y}) \cdot \nu}{2\pi \|\mathbf{x} - \mathbf{y}\|^2}, \quad G_{L\tau}(\mathbf{x}, \mathbf{y}) = -\frac{(\mathbf{x} - \mathbf{y}) \cdot \tau}{2\pi \|\mathbf{x} - \mathbf{y}\|^2},$$

and (3.10). On a circle of radius r , we have

$$\mathbf{x} = (r \cos s', r \sin s'), \quad \mathbf{y} = (r \cos s, r \sin s)$$

and

$$(\mathbf{x}-\mathbf{y}) \cdot \nu = r(1 - \cos(s' - s)), \quad (\mathbf{x}-\mathbf{y}) \cdot \tau = r \sin(s' - s), \quad \|\mathbf{x}-\mathbf{y}\|^2 = 2r^2(1 - \cos(s' - s)).$$

We then have

$$(4.4) \quad G_{L\nu}(\mathbf{x}, \mathbf{y}) = -\frac{1}{4\pi r}, \quad G_{L\tau}(\mathbf{x}, \mathbf{y}) = -\frac{1}{4\pi r} \cot\left(\frac{s' - s}{2}\right).$$

That is, the kernel of $S_{L\nu}$ is constant and $S_{L\tau} = -\frac{1}{2}H$, where H is the Hilbert transform on the unit circle. It is easy to verify that all of these operators are diagonalized by the Fourier transform. Thus, we only need to consider the 2×2 block for each Fourier mode e^{iks} with $k \in \mathbb{Z}$.

For $k = 0$, we have

$$(4.5) \quad \begin{aligned} \left(\frac{1}{2}I + S_{L\nu}\right)[1] &= \frac{1}{2} - \frac{1}{4\pi r} \int_0^{2\pi} 1 \cdot r ds = 0, \\ S_{L\tau}[1] &= -\frac{1}{2}H[1] = 0. \end{aligned}$$

$A_{H\tau}^{j,0}[1] = 0$ by symmetry, and

$$(4.6) \quad \left(\frac{1}{2}I + A_{H\nu}^{j,0}\right)[1] = \lambda_0,$$

where

$$(4.7) \quad \lambda_0 = \frac{1}{2} - \int_0^{2\pi} \left[\frac{e^{-r^2(1-\cos s)/(2\Delta t)}}{4\pi} - \frac{(1 - \cos s) r^2 E_1\left(\frac{r^2(1-\cos s)}{2\Delta t}\right)}{8\pi\Delta t} \right] ds.$$

Since $E_1(x) > 0$ for $x > 0$, we have $\lambda_0 > 0$ for any $r > 0$ and $\Delta t > 0$. Thus, the system (4.2) has eigenvalues 0 and λ_0 with eigenvectors $[1 \ 0]^T$ and $[0 \ 1]^T$.

For $k \neq 0$, we have

$$(4.8) \quad \begin{aligned} \left(\frac{1}{2}I + S_{L\nu}\right)[e^{iks}](s') &= \frac{1}{2}e^{iks'}, \quad S_{L\tau}[e^{iks}](s') = -\frac{1}{2}H[e^{iks}](s') = \frac{1}{2}i \operatorname{sgn}(k)e^{iks'}, \\ \left(\frac{1}{2}I + A_{H\nu}^{j,0}\right)[e^{iks}](s') &= a_k e^{iks'}, \quad A_{H\tau}^{j,0}[e^{iks}](s') = b_k e^{iks'}, \end{aligned}$$

with a_k, b_k defined by the formulae

$$(4.9) \quad \begin{aligned} a_k &= \frac{1}{2} - \frac{1}{4\pi} \int_{-\pi}^{\pi} e^{-\frac{r^2 \sin^2(s/2)}{\Delta t}} \cos(ks) ds \\ &\quad + \frac{r^2}{8\pi\Delta t} \int_{-\pi}^{\pi} (1 - \cos s) E_1\left(\frac{r^2 \sin^2(s/2)}{\Delta t}\right) \cos(ks) ds, \\ b_k &= \frac{i}{4\pi} \int_{-\pi}^{\pi} \cot\left(\frac{s}{2}\right) e^{-\frac{r^2 \sin^2(s/2)}{\Delta t}} \sin(ks) ds \\ &\quad - \frac{ir^2}{8\pi\Delta t} \int_{-\pi}^{\pi} \sin(s) E_1\left(\frac{r^2 \sin^2(s/2)}{\Delta t}\right) \sin(ks) ds. \end{aligned}$$

We note that $E_1(x)$ has a series expansion $E_1(x) = \ln x + \gamma + x + \frac{x^2}{4} + \cdots$ and that the spectrum of an integral operator with a smooth kernel decays exponentially fast. Thus, we have

$$(4.10) \quad \begin{aligned} a_k &\approx \frac{1}{2} + \frac{r^2}{8\pi\Delta t} \int_{-\pi}^{\pi} (1 - \cos s) \ln(\sin^2(s/2)) \cos(ks) ds, \\ b_k &\approx \frac{i}{4\pi} \int_{-\pi}^{\pi} \cot\left(\frac{s}{2}\right) \sin(ks) ds - \frac{ir^2}{8\pi\Delta t} \int_{-\pi}^{\pi} \sin(s) \ln(\sin^2(s/2)) \sin(ks) ds. \end{aligned}$$

Using the facts (see, for example, [43]) that

$$\int_{-\pi}^{\pi} \ln(\sin^2(s/2)) \cos(ks) ds = -\frac{2\pi}{|k|}, \quad \frac{1}{2\pi} \int_{-\pi}^{\pi} \cot\left(\frac{s}{2}\right) \sin(ks) ds = \operatorname{sgn}(k),$$

we obtain

$$(4.11) \quad a_k \approx \frac{1}{2} - \frac{r^2}{4\Delta t |k|^3}, \quad b_k \approx \frac{i}{2} \operatorname{sgn}(k) - \frac{ir^2}{4\Delta t k^2} \operatorname{sgn}(k).$$

Combining all the above, we see that for the Fourier mode e^{iks} with k large, the following 2×2 matrix determines its spectral behavior:

$$(4.12) \quad \begin{bmatrix} \frac{1}{2} & \frac{i}{2} \operatorname{sgn}(k) [1 - \frac{r^2}{2\Delta t k^2}] \\ -\frac{i}{2} \operatorname{sgn}(k) & \frac{1}{2} \end{bmatrix}.$$

The above matrix has roughly equal eigenvalues and singular values with $\lambda_{k1} \approx \sigma_{k1} \approx 1 - \frac{r^2}{8\Delta t k^2}$ and $\lambda_{k2} \approx \sigma_{k2} \approx \frac{r^2}{8\Delta t k^2}$ for k large. In summary, the integral equation system (3.6) has eigenvalues (and roughly equal singular values) $0, \lambda_0, \frac{r^2}{8\Delta t k^2}, 1 - \frac{r^2}{8\Delta t k^2}$ for k large. From this, in the complement of the one-dimensional nullspace, the condition number of the linear system can be seen to be of the order $O(\Delta t/h^2)$ with $h = 2\pi/N$ the spatial discretization size, assuming a uniform grid (so that $k \approx N \approx 1/h$). In short, the condition number is $O(N)$ for $\Delta t = O(h)$ and $O(1)$ for $\Delta t = O(h^2)$. For a fixed time step Δt independent of N , the condition number is $O(N^2)$ (see Figure 4.1). These estimates are essentially the same as the conditioning of an implicit finite difference approximation applied to the heat equation with the same Δt and h .

The preceding analysis can be extended, in part, to the case of an arbitrary smooth curve.

LEMMA 4.1. *The nullspace of the system of integral equations (3.6) contains functions of the form $[\rho_0(\mathbf{x})f(t) \ 0]^T$, where $\rho_0(\mathbf{x})$ spans the one-dimensional nullspace of the operator $\frac{1}{2}I + S_{L\nu}$ and $f(t)$ is an arbitrary smooth function on $[0, T]$.*

Proof. It is well known that the operator $\frac{1}{2}I + S_{L\nu}$ has a one-dimensional nullspace (see, for example, [44]). Let us denote a corresponding null vector by ρ_0 . It is easy to see that the function $v = S_L[\rho_0]$ solves the interior Neumann problem for the Laplace equation with zero boundary data. From well-known properties of harmonic functions, this implies that v must be constant in D , so that its tangential derivative must be zero on the boundary. Thus, $S_{L\nu}[\rho_0] = 0$, completing the proof. \square

REMARK 4. *Numerical experiments indicate that the only null vectors of (3.6) are the functions identified in Lemma 4.1. We conjecture that the coupled system of integral equations is exactly rank one deficient in any simply connected domain.*

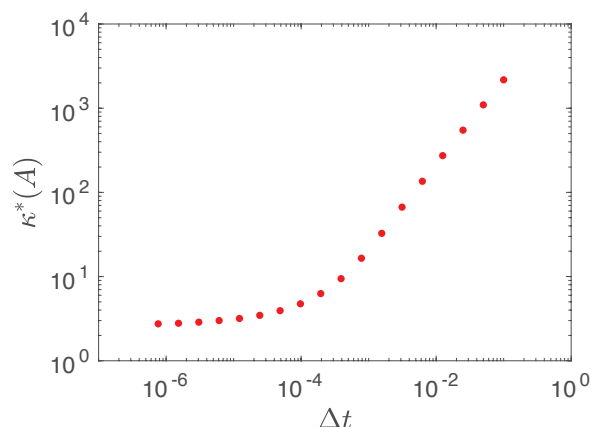


Fig. 4.1 The condition number of the FI-CSIE (in the complement of the one-dimensional nullspace) as a function of the time step for a circle of radius $r = 0.6$, discretized with 128 points.

5. Numerical Results for the Coupled Integral Equation System. Let us first consider the behavior of the implicit, second-order accurate one-step marching scheme described above. We will refer to solving the resulting system of the form (4.2) as the fully implicit combined source integral equation (FI-CSIE). As discussed in subsection 3.2, we use a 16th-order accurate spatial quadrature rule [1] so that the spatial error is negligible and we accelerate the computation of the history part using the Fourier spectral method of [33, 42].

In Figure 5.1, we plot the eigenvalues of the system matrix with $n = 64$ and $n = 128$, respectively, when Γ is a circle of radius $r = 0.6$. Note that the asymptotic analysis is in close agreement with direct discretization. It is worth noting that

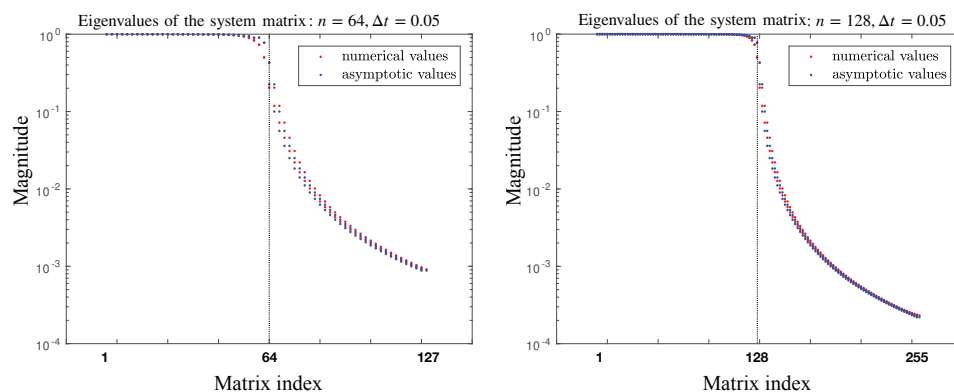


Fig. 5.1 Magnitude of the eigenvalues of the system matrix (of size $2n$) for a circle of radius $r = 0.6$, discretized with n points. The red dots are numerical values and the blue dots are the asymptotic values from section 4. The x -axis corresponds to the eigenvalue index, plotted in decreasing order. The eigenvalues from the asymptotic analysis are ordered in the corresponding fashion. We omit the exact zero eigenvalue from the Fourier analysis, which is manifested by a single eigenvalue of order 10^{-15} in the matrix analysis.

Table 5.1 Numerical results for the ellipse using the second-order accurate fully implicit marching scheme. The matrix size is 192×192 .

| N | Δt | N_{its} | Error | Ratio |
|-----|------------|------------------|---------------------|-------|
| 10 | 1/10 | 93 | $2.0 \cdot 10^{-2}$ | |
| 20 | 1/20 | 93 | $1.2 \cdot 10^{-2}$ | 1.6 |
| 40 | 1/40 | 92 | $3.8 \cdot 10^{-3}$ | 3.8 |
| 80 | 1/80 | 90 | $7.5 \cdot 10^{-4}$ | 4.2 |
| 160 | 1/160 | 86 | $1.8 \cdot 10^{-4}$ | 4.1 |

Table 5.2 Numerical results for the hexagram using the second-order accurate fully implicit marching scheme. The matrix size is 320×320 .

| N | Δt | N_{its} | Error | Ratio |
|-----|------------|------------------|---------------------|-------|
| 10 | 1/10 | 172 | $1.8 \cdot 10^{-2}$ | |
| 20 | 1/20 | 178 | $6.9 \cdot 10^{-3}$ | 2.6 |
| 40 | 1/40 | 183 | $1.5 \cdot 10^{-3}$ | 4.5 |
| 80 | 1/80 | 185 | $2.8 \cdot 10^{-4}$ | 5.5 |
| 160 | 1/160 | 183 | $6.1 \cdot 10^{-5}$ | 4.6 |

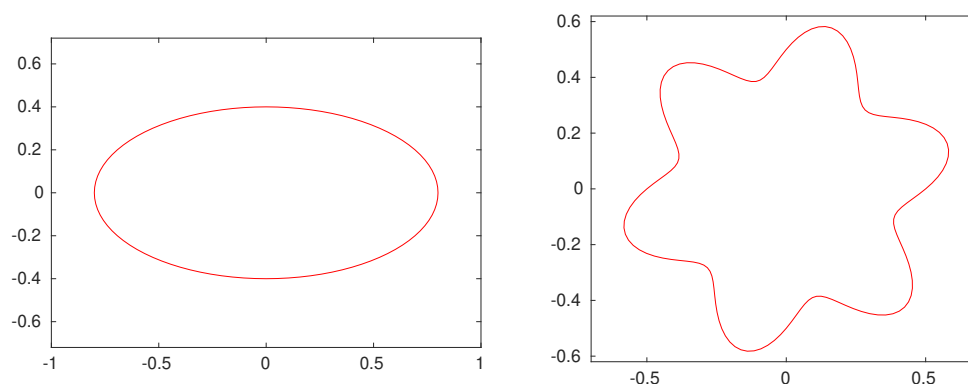


Fig. 5.2 Boundary curves for Tables 5.1 and 5.2, respectively.

a physical constraint on the boundary data \mathbf{g} or $\tilde{\mathbf{g}}$ is that the normal component integrates to zero on Γ . Such data has no projection onto the null vector of the system matrix so that an iterative method such as GMRES can be applied without any modification. With a stopping criterion for the residual set to 10^{-12} , Tables 5.1 and 5.2 show the performance of GMRES and the error obtained when the boundary is either an ellipse with aspect ratio 2 : 1 or a smooth hexagram, respectively (see Figure 5.2). For all the tables presented in this paper, we take as the exact solution the divergence-free velocity field

$$\begin{aligned}
 \mathbf{u}(\mathbf{x}, t) = & \sum_{j=1}^{10} \sum_{k=0}^{\lfloor t/(2h) \rfloor} \frac{(x_2 - x_{2j}, x_{1j} - x_1)}{|\mathbf{x} - \mathbf{x}_j|^2} \left(e^{-\frac{|\mathbf{x} - \mathbf{x}_j|^2}{4(t - (2k+1)h)}} - e^{-\frac{|\mathbf{x} - \mathbf{x}_j|^2}{4(t - (2k+2)h)}} \right) \\
 & + t \cos(313\pi t)(x_1, -x_2) + \frac{t^2}{4} \cos(233\pi t)e^{x_1}(\cos x_2, -\sin x_2) \\
 & + 2t \sin(299\pi t)e^{x_2}(\cos x_1, \sin x_1),
 \end{aligned}
 \tag{5.1}$$

where $h = 0.1$, $\mathbf{x} = (x_1, x_2)$, and the $\{\mathbf{x}_j\}$ are chosen to be equispaced on the unit circle, which encloses both domains of interest. (This is a solution to (3.3), determined uniquely in the interiors of the computational domains by its boundary data.) 96 points are used to discretize the ellipse and 160 points are used to discretize the hexagram. In both tables, the first column lists the total number of time steps N needed to reach $t = 1$; the second column lists the time step; the third column lists the average number of GMRES iterations required to solve the system to the desired tolerance; the fourth column lists the relative l_2 error at 20 random points in the computational domain; the last column lists the ratio of the errors for each doubling of N . Note that the data are consistent with second-order accuracy in time. Note also that the number of iterations is approximately equal to the number of points on the boundary, as expected for a large time step with $\Delta t \approx 1/N$.

6. The Predictor-Corrector Combined Source Integral Equation (PC-CSIE).

The previous section shows that the FI-CSIE yields a somewhat ill-conditioned system of equations for large time steps. We now investigate a method for solving the unsteady Stokes equations using a rule of predictor-corrector type.

DEFINITION 6.1. *The k th-order predictor-corrector scheme for the CSIE, denoted by PC-CSIE(k), is given as follows (using the notation of subsection 3.2):*

Step 1 : Set $\mu_j = P_k^E[\mu_{j-1}](j\Delta t)$.

Step 2 : Solve $\left(\frac{1}{2}I + \mathcal{S}_{L\nu}\right)\rho_j = \nu \cdot \tilde{\mathbf{g}} + \mathcal{S}_{H\tau}^{\text{far}}[\mu_{j-1}] + \mathcal{S}_{H\tau}^{\text{loc}}[\mu_j]$.

Step 3 : Solve $\left(\frac{1}{2}I + A_{H\nu}^{j,0}\right)\mu_j = \tau \cdot \tilde{\mathbf{g}} - \mathcal{S}_{H\nu}^{\text{far}}[\mu_{j-1}] - F_{H\nu}^{j,k}[\mu_{j-1}] - \mathcal{S}_{L\tau}[\rho_j]$.

That is, we first extrapolate μ_j from previous time steps, then solve for ρ_j , and finally solve for μ_j given the newly computed ρ_j . Note that the integral equation in Step 2 of PC-CSIE(k) is a Fredholm equation of the second kind. While it has a one-dimensional nullspace, it is simply the classical equation for the harmonic interior Neumann problem obtained when representing the solution as a single layer potential. Since the right-hand side is compatible, the resulting linear system is easily solved using GMRES with $O(1)$ iterations. The integral equation in Step 3 is a Volterra equation of the second kind. It is always invertible and well-conditioned (at least on smooth curves).

The obvious difference between PC-CSIE(k) and the fully implicit version comes from the extrapolation step. While the order of accuracy can be arbitrarily high, and the individual integral equations are well-conditioned, the stability of the resulting scheme remains to be studied.

Preliminary experiments suggest that the schemes PC-CSIE(2) and PC-CSIE(3) are stable even for moderately large time steps, while PC-CSIE(4) is not, but a thorough analysis remains to be undertaken (see section 8). Using the same geometries as above (see Figure 5.2), we obtain the results in Table 6.1 for the ellipse (top) and the hexagram (bottom). Note that many fewer GMRES iterations are required for each step. The convergence rates estimated by the error ratios in the last column are somewhat erratic, but generally better than the theoretical estimate 2^k .

Table 6.1 Numerical results for the ellipse and hexagram using the PC-CSIE(2) method. The matrix size is 96×96 for the ellipse and 256×256 for the hexagram at each stage.

| Ellipse | | | | |
|---------|------------|------------------|---------------------|-------|
| N | Δt | N_{its} | Error | Ratio |
| 80 | 1/40 | 7 | $4.2 \cdot 10^{-2}$ | |
| 160 | 1/80 | 7 | $1.7 \cdot 10^{-3}$ | 24 |
| 320 | 1/160 | 7 | $2.9 \cdot 10^{-4}$ | 6.0 |
| 640 | 1/320 | 7 | $4.8 \cdot 10^{-5}$ | 6.0 |
| 1280 | 1/640 | 7 | $9.0 \cdot 10^{-6}$ | 5.3 |

| Hexagram | | | | |
|----------|------------|------------------|---------------------|-------|
| N | Δt | N_{its} | Error | Ratio |
| 80 | 1/40 | 12 | $9.6 \cdot 10^{-3}$ | |
| 160 | 1/80 | 12 | $3.6 \cdot 10^{-3}$ | 2.7 |
| 320 | 1/160 | 11 | $5.8 \cdot 10^{-4}$ | 6.1 |
| 640 | 1/320 | 11 | $9.9 \cdot 10^{-5}$ | 5.8 |
| 1280 | 1/640 | 11 | $1.7 \cdot 10^{-5}$ | 5.7 |

7. High-Order Schemes Using Spectral Deferred Corrections. Instead of seeking to develop stable higher-order predictor-corrector type schemes, we now show that more rapid convergence is easily achieved by combining PC-CSIE(2) with spectral deferred correction (SDC) [7, 21, 35, 53].

A very brief introduction to deferred corrections follows: suppose that we seek the solution $v(t)$ of some time-dependent problem starting at $t = 0$, and that an approximate solution can be computed for k steps on $[0, \Delta t]$ using a low-order accurate method with an error of the order $O(\Delta t^m)$ for some $m < k$. We denote the discrete solution at those k points by $\mathbf{v}_k^{[0]}$. One can then interpolate the low-order solution by a polynomial in t of order k , namely, $P_k^I[\mathbf{v}_k^{[0]}]$, defined in subsection 3.2. This allows us to define a continuous error function

$$\delta^{[0]}(t) = v(t) - P_k^I[\mathbf{v}_k^{[0]}](t),$$

which can be substituted into the governing equation for $v(t)$ and solved for $\delta^{[0]}(t)$, using the same low-order scheme. This generates the discrete solution vector $\delta_k^{[0]}$. A corrected approximation is then defined by

$$\mathbf{v}_k^{[1]} = \mathbf{v}_k^{[0]} + \delta_k^{[0]}.$$

It is straightforward to show that the error in $\mathbf{v}_k^{[1]}$ is of the order $O(h^{2m})$, so long as $2m < k$ and all computations involving the known function $P_k^I[\mathbf{v}_k^{[0]}](t)$ are carried out with k th-order accuracy. For further details, see the references above and [6]. The correction procedure is easily iterated until k th-order accuracy is achieved. The process can then be repeated on the next time interval $[\Delta t, 2\Delta t]$, etc. (The phrase *spectral deferred correction* is typically used when the underlying problem has been formulated as an integral equation and the k stages are chosen at nodes corresponding to some high-order spectral discretization, typically of Gauss or Gauss–Radau type.)

In the present context, let us assume that we have divided $[0, T]$ into N equal subintervals $[t_{i-1}, t_i]$ with $t_i = i\Delta t$, $\Delta t = T/N$ for $i = 1, \dots, N$. We restrict our attention to the i th such interval $[t_{i-1}, t_i]$, which we will denote by $[\alpha, \beta]$ when the context is clear. Given a positive integer k , we will denote by $\alpha < \tau_1, \dots, \tau_k = \beta$

the k Gauss–Radau nodes shifted and scaled to the interval $[\alpha, \beta]$ (see, for example, [27]), with $\tau_0 = \alpha$. Let us denote the values of the densities ρ and μ at these nodes by $\boldsymbol{\rho}_k = (\rho_0, \rho_1, \dots, \rho_k)^T$ and $\boldsymbol{\mu}_k = (\mu_0, \mu_1, \dots, \mu_k)^T$, respectively. Since we are discretizing in time only, recall that ρ_i, μ_i are functions of the spatial variable $\mathbf{y} \in \Gamma(t)$.

Following the principle outlined above, the first step of SDC for the mixed potential formulation is to use some low-order scheme to obtain $\boldsymbol{\rho}_k^{[0]}$ and $\boldsymbol{\mu}_k^{[0]}$. We then use these two vectors to obtain interpolating polynomials of degree $k-1$ in time, namely, $P_k^I[\boldsymbol{\rho}_k^{[0]}]$ and $P_k^I[\boldsymbol{\mu}_k^{[0]}]$, and define

$$\delta_\mu^{[0]}(t) = \mu(t) - P_k^I[\boldsymbol{\mu}_k^{[0]}], \quad \delta_\rho^{[0]}(t) = \rho(t) - P_k^I[\boldsymbol{\rho}_k^{[0]}].$$

Inserting this representation into (3.6), we obtain

$$(7.1) \quad \begin{aligned} \frac{1}{2} \delta_\rho^{[0]}(\mathbf{x}, t) + \mathcal{S}_{L\nu}[\delta_\rho^{[0]}](\mathbf{x}, t) + \mathcal{S}_{H\tau}[\delta_\mu^{[0]}](\mathbf{x}, t) &= R_1(\mathbf{x}, t), \\ -\mathcal{S}_{L\tau}[\delta_\rho^{[0]}](\mathbf{x}, t) + \frac{1}{2} \delta_\mu^{[0]}(\mathbf{x}, t) + \mathcal{S}_{H\nu}[\delta_\mu^{[0]}](\mathbf{x}, t) &= R_2(\mathbf{x}, t), \end{aligned}$$

where the residuals R_1 and R_2 are given by

$$\begin{aligned} R_1(\mathbf{x}, t) &= \nu \cdot \tilde{\mathbf{g}}(\mathbf{x}, t) - \frac{1}{2} P_k^I[\boldsymbol{\rho}_k^{[0]}](\mathbf{x}, t) - \mathcal{S}_{L\nu}[P_k^I[\boldsymbol{\rho}_k^{[0]}]](\mathbf{x}, t) - \mathcal{S}_{H\tau}[P_k^I[\boldsymbol{\mu}_k^{[0]}]](\mathbf{x}, t), \\ R_2(\mathbf{x}, t) &= -\tau \cdot \tilde{\mathbf{g}}(\mathbf{x}, t) + \mathcal{S}_{L\tau}[P_k^I[\boldsymbol{\rho}_k^{[0]}]](\mathbf{x}, t) - \frac{1}{2} P_k^I[\boldsymbol{\mu}_k^{[0]}](\mathbf{x}, t) - \mathcal{S}_{H\nu}[P_k^I[\boldsymbol{\mu}_k^{[0]}]](\mathbf{x}, t). \end{aligned}$$

Note that this is exactly the same equation as (3.6), but with a different right-hand side. Since, as noted above, SDC requires that all residuals be computed with high-order accuracy, we have provided some of the integrals needed at the intermediate stages in the appendix.

After solving (7.1) at the same k stages, yielding $[\delta_\rho^{[0]}]_k, [\delta_\mu^{[0]}]_k$, we let

$$(7.2) \quad \boldsymbol{\rho}_k^{[1]} = \boldsymbol{\rho}_k^{[0]} + [\delta_\rho^{[0]}]_k, \quad \boldsymbol{\mu}_k^{[1]} = \boldsymbol{\mu}_k^{[0]} + [\delta_\mu^{[0]}]_k.$$

This procedure may be repeated until the desired order of accuracy is achieved.

We have implemented SDC using the second-order predictor-corrector scheme PC-CSIE(2) described in the previous section. For simplicity, we provide numerical results in Table 7.1 for the case of a circle. In this table, SDC_k^j denotes the scheme with k Gauss–Radau nodes on each subinterval and j iterations of deferred correction. In particular, SDC_k^0 is simply the uncorrected solution obtained with PC-CSIE(2). N is the number of subintervals, Δt is the time step size for each subinterval, E is the relative l^2 error for the method indicated in the subscript, and N_{it} is the average number of iterations for GMRES to reach the requested tolerance 10^{-12} . The total number of time steps is Nk and the expected error reduction should be $2^{2(j+1)}$ for SDC_k^j until $2(j+1) > k$, since we are driving the deferred correction process with a second-order accurate scheme. While the behavior of the SDC schemes as a function of the number of correction sweeps is somewhat erratic, it is more or less consistent with the asymptotic estimates. That is, SDC_5^0 converges approximately like a second-order scheme, while SDC_5^1 converges at a much higher rate. Further sweeps of deferred correction don't increase the convergence rate significantly, since the degree of polynomial approximation is only four, limiting the order of accuracy, as discussed above. Note, however, that these further sweeps have no impact on stability.

Table 7.1 Numerical results for the circle of radius 0.5, using SDC. The number of points in the spatial discretization is 200.

| Nk | $\Delta t/k$ | N_{its} | $E_{SDC_5^0}$ | $E_{SDC_5^1}$ | $E_{SDC_5^2}$ | $E_{SDC_5^3}$ | $E_{SDC_5^4}$ |
|------|--------------|------------------|---------------------|---------------------|---------------------|---------------------|---------------------|
| 40 | 1/20 | 2.7 | $1.5 \cdot 10^{-2}$ | $2.5 \cdot 10^{-2}$ | $2.5 \cdot 10^{-2}$ | $2.5 \cdot 10^{-2}$ | $2.5 \cdot 10^{-2}$ |
| 80 | 1/40 | 2.6 | $3.0 \cdot 10^{-3}$ | $3.5 \cdot 10^{-5}$ | $1.8 \cdot 10^{-4}$ | $1.8 \cdot 10^{-4}$ | $1.8 \cdot 10^{-4}$ |
| 160 | 1/80 | 2.7 | $2.7 \cdot 10^{-3}$ | $3.8 \cdot 10^{-5}$ | $3.7 \cdot 10^{-5}$ | $3.7 \cdot 10^{-5}$ | $3.7 \cdot 10^{-5}$ |
| 320 | 1/160 | 2.6 | $6.9 \cdot 10^{-4}$ | $1.5 \cdot 10^{-7}$ | $1.3 \cdot 10^{-7}$ | $9.8 \cdot 10^{-8}$ | $4.5 \cdot 10^{-8}$ |
| 640 | 1/320 | 2.6 | $1.6 \cdot 10^{-4}$ | $1.1 \cdot 10^{-8}$ | $5.3 \cdot 10^{-9}$ | $4.5 \cdot 10^{-9}$ | $4.9 \cdot 10^{-9}$ |

REMARK 5. A nice feature of the mixed potential representation is the complete separation of the instantaneous pressure source from the vortex source. As a result, even though the exact solution defined in (5.1) has a highly oscillatory pressure field, the numerical results in Tables 5.1, 5.2, 6.1, and 7.1 show that high accuracy is achieved even for large time steps that underresolve the oscillatory behavior of the pressure.

8. Conclusions and Future Work. We have developed a new integral representation for the unsteady Stokes equations which makes use of simple harmonic and heat potentials, leading to the *combined source integral equation* (CSIE). Unlike schemes based on the unsteady Stokeslet [42], this permits the direct application of well-developed fast algorithms (see, for example, [31, 33, 50, 65, 69, 70] and [13, 25, 30, 32, 47, 51]).

While the fully coupled CSIE is not of Fredholm type, we have shown that each individual equation is well-conditioned, and we found that a second-order predictor-corrector type method is effective even for large time steps. Moreover, one can achieve high-order accuracy through the use of spectral deferred correction. Since our primary goal in the present paper is the development of the mathematical representation itself, a more thorough investigation of various predictor-corrector, Runge–Kutta, and implicit-explicit type marching schemes will be carried out at a later date.

It is worth noting that the mixed potential representation and existing gauge methods (see (1.6)) bear some resemblance to one another. The principle differences are (1) that we are working in an integral equation–based framework, (2) that we apply the Helmholtz decomposition to the inhomogeneous data rather than to the auxiliary unknown vector field \mathbf{m} , and (3) that we have as unknowns only the vortex source and pressure source, which are restricted to the boundary and for which imposing velocity boundary conditions is straightforward.

We should also note that the potential theoretic aspects of our method can be extended to piecewise smooth boundaries [26, 44, 67], and that recent advances on understanding corner singularities have made integral equation methods extremely efficient for such problems [8, 10, 37, 38, 62, 63].

A number of open questions remain, including the completeness of the representation for multiply connected domains, a detailed characterization of the nullspace of the coupled system, and the extension of the mixed potential formulation to other boundary/interface conditions. While our representation is formally valid in either fixed or moving geometries, we have not yet investigated its performance in the non-stationary case. Moreover, while fast algorithms for harmonic and heat potentials are available in the three-dimensional setting, suitable quadrature rules and high-order surface representations are still areas of active research, especially for moving geometries.

To solve the equations with a forcing term we also need to couple the solver described here with volume-integral based codes for the Helmholtz decomposition, as discussed in section 2.3. The full Navier–Stokes equations are, of course, nonlinear but most marching schemes (as noted in the introduction) treat the nonlinear term explicitly so that the approach developed here applies. Fully implicit schemes will require the introduction of unknowns in the interior of the domain as well as the development of a solver for nonlinear volume integral equations. We are presently investigating many of these topics and will report on our progress at a later date.

Appendix A. For the SDC method of section 7, using product integration in time as in [42], the residual is required at intermediate stage $\tau_i \in [\alpha, \beta]$. This involves integrals beyond those given by equation (4.9) in [42]. We provide those integrals here, which are sufficient to obtain fourth-order accuracy:

$$(A.1) \quad \int_{\alpha}^{\alpha+\tau_i} e^{-r^2/4(\alpha+\tau_i-\tau)} \frac{(\beta-\tau)^j}{(\alpha+\tau_i-\tau)^2} d\tau = \begin{cases} \frac{4}{r^2} e^{-c}, & j=0, \\ E_1(c) + \frac{b}{a} e^{-c}, & j=1, \\ \left(\tau_i + \frac{b^2}{a}\right) e^{-c} - E_1(c)(a-2b), & j=2, \\ -\frac{e^{-c}}{2} (\tau_i a - x_i^2 - 6\tau_i b - 2b^3/a) + E_1(c)(a^2 - 6ab + 6b^2), & j=3, \\ \frac{e^{-c}}{6} (2\tau_i^3 - \tau_i^2(a-12b) + \tau_i(a-6b)^2 + 6b^4/a) \\ - \frac{E_1(c)}{6} (a^3 - 12a^2b + 36ab^2 - 24b^3), & j=4, \end{cases}$$

where $r = \|\mathbf{x} - \mathbf{y}\|$, $c = \frac{r^2}{4\tau_i}$, $a = \frac{r^2}{4}$, $b = \Delta t - \tau_i$.

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