

## Lecture VII

- MHD Turbulence II
  - review of Kraichnan - Inertial Theory
  - Goldreich - Sridhar and Critical Balance
  - Spectra / Triads
- Taylor Relaxation I
  - Relaxation  $\rightarrow$  Impact of Turbulence (Dynamo) on Mean Field
  - Magnetic Helicity
  - Taylor Relaxation - an Intro.

# Review — MHD Turbulence

①

→  $K41$

similarity,  $S$ , and  $R$

$$l_d < \ell < l_0 , \quad \epsilon \sim \frac{v(\ell)^3}{\ell}$$

$$v(\ell) \sim \epsilon^{1/4} \ell^{1/2}$$

$$l_d \sim \left( v^3 / \epsilon \right)^{1/4}$$

→ MHD



wave packet streams

$$\delta z_\ell \sim T_A \delta r z_\ell$$

$$\sim T_A \frac{z_\ell^2}{\ell} \quad \text{random}$$

$$\Omega \delta r z_\ell \sim \frac{z_\ell^2}{\ell} \quad \leftrightarrow \quad \delta r x \sim \frac{\phi}{V}$$

mean times

$$\langle \delta z_\ell^2 \rangle \sim \tau_{\text{av}} \left( \frac{z_\ell^2}{\ell} \right)^2 +$$

$$T_{\text{fr}} \quad \langle \delta z_\ell^2 \rangle \sim \langle z_\ell^2 \rangle$$

→  $P_m = \pm$  → effect on  $f_d$  ?

6

→ Result. → K-I spectrum.

→ + anisotropy  $\Rightarrow G-S$ .

## Review

6.

### Kraichnan - Inoshikov

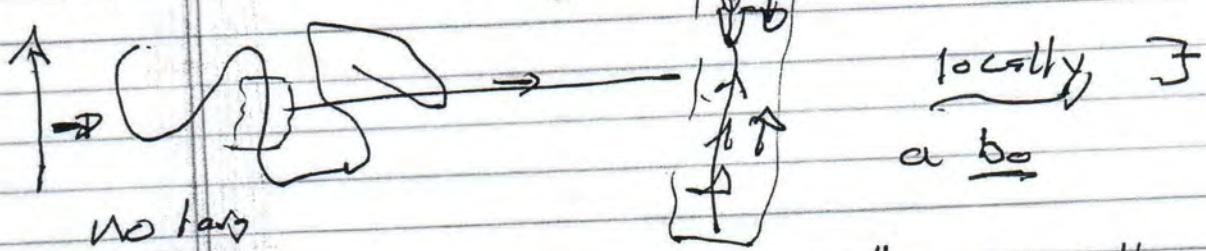
#### so, some Phenomenology:

#### classic

(Inoshikov,  
Kraichnan)  
(64, 65)

→ nonlinear transfer - (2) isotropic turbulence

For weak mean  $\bar{B}$ ,  $b_0 \sim b_{rms} > TB$



→ so, nonlinear scattering from "collisions"  
of counter-propagating packets

$$M \rightarrow \leftarrow M$$

so (modulo 4/5!) - TBD.

[universality  
wave, =  
counter E's  
populations]

$$\rightarrow \left\{ \epsilon \sim \frac{Z^2}{T^2} \right\} \quad l \equiv \text{scale} \quad \text{locality?}$$

→ A linear wave cascade  
 → const. energy thru-set  
 rate

Now A linear transit time on scale  $l$ , in  $b_0$   
is  $\frac{l}{b_0}$

$$\boxed{T_A = l/b_0}$$

For transfer:

- transfer by wave scattering  $\xrightarrow{\text{random walk}}$  of amplitude/alters
- define  $T_{Tr}$  by:

$$\langle \delta Z^2(T_{Tr}) \rangle \sim \langle Z^2 \rangle$$

i.e. randomization of scattering in amplitude

$$\delta Z_{\text{rms}} \sim Z$$

$T_{Tr}$  sufficient # kicks to make

(3)

$$Z_e(t + T_A) = Z_e(t) + T_A \delta Z_p$$

$\downarrow$  diffusion process  $\xrightarrow{t \ll t}$   
kick in  $T_A$  after transit time

random walk.

$$\delta Z_p \approx T_A \frac{Z_e^2}{\ell} \rightarrow \text{kick in } T_A$$

dimin.

$$\text{so } \frac{Z_e^2}{\ell} \rightarrow \text{kick.}$$

$$\Rightarrow \langle \delta Z_e^2 \rangle = \left( T_A \frac{Z_e^2}{\ell} \right)^2 \xrightarrow{\text{accumulated kicks}} D t$$

thus:

$$\langle \delta Z_e^2 \rangle \sim \langle Z_e^2 \rangle$$

$\rightarrow$  defined  $T_{Tr}$ .

$$\Rightarrow \delta Z_p^2 \approx T_A^2 \frac{(Z_e^2)^2}{\ell^2} \frac{T_{Tr}}{T_A}$$

$$\therefore \frac{1}{T_{Tr}} \approx \frac{Z_e^2}{\ell^2} T_A$$

$\rightarrow$  determines transfer rate.

$$\langle \delta Z_e^2 \rangle \sim \langle Z_e^2 \rangle$$

$$T_A + \frac{\langle Z_e^3 \rangle}{\langle Z_e^2 \rangle} \sim \langle Z_e^2 \rangle$$

1.

$$\underline{\underline{\sigma}} \frac{1}{T_{Tr}} \sim \frac{Z_e^2}{\ell^2} \frac{\ell}{b_0}$$

$$\sim \frac{Z_e^2}{\ell b_0}$$

Thus:

$$\underline{\underline{\epsilon}} \sim \frac{Z_e^2}{T_{Tr}} \sim \frac{Z_e^2 Z_e^2}{\ell b_0}$$

const. rate balance

$$\boxed{Z_e^3 \sim \sqrt{b_0 \epsilon} \ell^{1/2}}$$

$$Z_e \sim (b_0 \epsilon)^{1/4} \ell^{1/4}$$

$$\boxed{E(k) \sim \sqrt{b_0} k^{-3/2}}$$

Kraichnan spectrum

N.B. - as work with  $Z$

$$E_M \sim E_K$$

(equal Elsasser regulations)

$$-\frac{1}{T_{Tr}} \sim \frac{Z_e^2}{\ell^2} \frac{\ell}{b_0} \equiv \underbrace{\frac{\gamma_A}{T_{Elly}}}_{T_{Elly}} \frac{1}{\gamma_{Elly}}$$

$$P_{\text{out}} = \sin - \text{out} \\ = \sum_{\omega, \theta} P_{\omega, \theta} e^{-i\omega t} \\ - \frac{1}{2} \int_{-\infty}^{\infty} \sum_{\omega, \theta} P_{\omega, \theta} e^{-i\omega t} \frac{d\omega}{\Delta \omega} \sim$$

$$T_{\text{tr}} = \sum_{\omega} |K + i\hbar/\omega|^2 \delta(E_p - E_i - \hbar\omega)$$

202  
t <  $\tau_{\text{edd}}$

$$\text{compare } K(4) : \frac{1}{T_{\text{tr}}} \sim \frac{1}{\tau_{\text{edd}}}$$

Transfer rate reduced by factor  $T_A / \tau_{\text{edd}}$ !

Also, recall for weak wave turbulence:

$$C_{QNG} = \sum_{k'} |\nabla I|^2 N_k' N_{k'} T_C \rightarrow \begin{matrix} \text{general structure} \\ \text{of collision} \\ \text{integral} \end{matrix}$$

$\downarrow$   
 $C_C$                        $\uparrow$   
 $(\text{tried}) \text{ coherence}$   
 $\text{time} \rightarrow \text{interaction}$

$$\frac{1}{T_{\text{tr}}} \simeq \sum_{k'} |\nabla I|^2 N_k' T_C$$

$$\text{but } N_k' \rightarrow Z_e^2 \text{ (intensity)} \quad (\text{assumes local transfer})$$

$$|\nabla I|^2 \sim |\nabla Z_e| \sim \frac{1}{L^2} \rightarrow \frac{1}{L^2} \quad C_C$$

$$T_C \sim T_A \rightarrow \text{tried coherence}$$

$$\xrightarrow{\uparrow} \text{Atmospheric packet transit time}$$

$$\frac{1}{T_{\text{tr}}} \sim \frac{Z_e^2}{L^2} T_A \hookrightarrow \text{W.T.T.}$$

Fundamentally wave scattering process

- 'triad'  $\rightarrow$  2 A<sub>WS</sub> + self.  
 $\rightarrow$  akin NLLD.
  - Alfvén waves non-dispersive, but  
 coherence time controlled by  
 packet propagation
  - $\Rightarrow$  presents negative spectra, etc.  
 (Coherence times not so large)
  - Strong B<sub>0</sub>. (explicit)
- Some simple observations:
- turbulence clearly anisotropic
  - nonlinear transfer in k<sub>||</sub>.

So, consider weak wave turbulence!

Consider: k<sub>||</sub>  $\gg$  l<sub>perp</sub>

k<sub>||</sub>

$$k_{\perp} \gg k_{||}$$

$$\text{so } E \sim \frac{(Z(l_{\perp})^2)(Z(l_{\parallel})^2)}{l_{\perp}^2 |k_{||} V_A|}$$

i.e. tacitly  $|k_{||} V_A| \sim |k_{\perp} V_A|$

$\Delta k_{||} V_A \rightarrow$  coh. time

as far as stochastic field

11.

$$\exists (\ell_{\perp})^2 \approx (\epsilon k_{\parallel} V_A)^{1/2} \ell_{\perp}$$

$\Rightarrow E(k_{\perp}) \approx (\epsilon k_{\parallel} V_A)^{1/2} / k_{\perp}^2$  short  
 and neglecting for 0.05 in  $k_{\parallel}$ :  $E(k_{\perp}, k_{\parallel}) \sim [\epsilon V_A]^{1/2} / k_{\perp}^2 k_{\parallel}^{1/2}$  int. restricted

- strongly anisotropic

weak wave  
turbulence  
effectively

-  $k_{\parallel}$  frozen.

Now,  $Z(\ell_{\perp}) \sim \partial B(\ell_{\perp}) \sim (\epsilon k_{\parallel} V_A)^{1/4} \ell_{\perp}^{1/2}$

$$\omega \sim k_{\parallel} V_A$$

recall:  $D_{\parallel} = D_z + \frac{\partial B_L}{B_0} \cdot D_{\perp}$

Kubo #

(see 235)  $k_{\parallel} \sim \frac{\partial B_L D_{\perp}}{B_0} / D_z \sim \frac{k_{\perp} \partial B_L / B}{\Delta_{\perp}}$

# kicks in  
coherence  
length

Now,  $\frac{\partial B_L}{\ell_{\perp}} \sim (\epsilon k_{\parallel} V_A)^{1/4} / \ell_{\perp}^{1/2}$

$$k_{\parallel} \sim \frac{1}{(\Delta k_{\parallel})} \frac{(\epsilon k_{\parallel} V_A)^{1/2}}{\ell_{\perp}^{3/2}}$$

B

⇒ For W.T.T., expect:

$k_{\text{tr}} < l_1$  (diffusive picture)

but,

$k_{\text{tr}}$  rises as  $l_1$  drops

i.e.  $k_{\text{tr}} \uparrow$  as progress thru  $l_1$  cascade → ?? what happens

⇒ Begs the question:

How high can  $k_{\text{tr}}$  go and still retain physics of Alfvén wave cascade?

need Park T.R. enter the:  $\langle \dots \rangle$

B86 comment '78

Critical Balance Conjecture

- Goldreich - Sridhar (1995)  
(cf also Kondratenko - Pogutse, 1978)

⇒ MHD (charis) ranges in strong field will sit at  $k_{\text{tr}} \approx l_1$

i.e.  $\delta B_z \cdot \nabla \perp \sim B_0 T_H$

i.e. transit time sets

$$\frac{z(l_1)}{l_1} \sim k_{\text{tr}} V_A$$

based on ext. strength

$$\frac{T_A}{T_B} \rightarrow_k, \frac{T_B}{T_A} \rightarrow_l$$

$k_{\text{tr}}$  vs  $k_L$  rel.

Teddy

13.

but:  $E \sim \frac{(z(l_\perp))^2 (z(l_\parallel))^2}{l_\perp^2} \frac{1}{k_{\text{in}} v_4}$

$$\rightarrow \frac{(z(l_\perp))^2 (z(l_\parallel))^2}{l_\perp^2 \frac{z(l_\perp)}{l_\perp}} \quad \text{Kary}$$
$$= \frac{z(l_\perp)^3}{l_\perp}$$

$$\Rightarrow z(l_\perp) \sim (E l_\perp)^{1/3}$$

$$E(k_\perp k_\parallel) \sim E^{2/3} k_\perp^{-5/3}$$

and anisotropy:

$$k_{\text{in}} \sim, \quad R_\theta \cdot \nabla \sim \vec{B}_\perp \cdot \nabla_\perp$$

$$B_\theta k_\parallel \sim E^{1/3} \frac{l_\perp^{1/3}}{l_\perp}$$

$k_{\text{in}}$  vs  
 $k_\perp$   
relation

$$k_{\text{in}} \sim \frac{E^{1/3}}{B_\theta} k_\perp^{+2/3}$$

$\leftarrow$  e-s spectrum  
back to  $k^{4/3}$   
but diff  
physics  
 $\rightarrow$  fits dists

$\leftarrow$   
 $k_{\text{in}}$  does "cascade"  
through more slowly  
specific "GS cone"  
in  $k$  space

i.e.  $k_{\text{in}} \ll k_\perp$

$\rightarrow$  cascade develops  
preferentially in  
perp.

$$k_{\text{in}} \sim k_\perp^{2/3}$$

14.

- Why Believe

Is there

→ analogues of 4/5 Law (August, Politano)

$$-\frac{4}{3} \epsilon^T f = \left\langle (\partial U \cdot \partial U + \partial B \cdot \partial B) \partial U \right\rangle$$

$\xrightarrow{\text{non local}} \text{form } \tilde{E}_S$

$$-2 \left\langle (\partial U \cdot \partial B) B \right\rangle$$

total energy  
conversion

induction  $\xrightarrow{\text{P}} \text{form } \tilde{E}_M$

~ reflects Flip-Flop in energy between  
channels

or

$$-\frac{4}{3} \epsilon^\pm f = \left\langle (\partial Z^\pm \cdot \partial Z^\mp) \partial Z^\mp \right\rangle$$

no dissipation for 1 stream only ↓.

→ Why? → relate induced E-field.

## MHD Turbulence II

- Anisotropic Cascades and Critical Balance  $\leftrightarrow$  A closer look.
- Extending the 4/5 Law.
- Selective Decay and Relaxation.
- 2D MHD - A Study in Turbulent Relaxation.

(i) Anisotropic Cascades and Critical Balance - A closer look.

Recall : I-K Phenomenology :

$$\epsilon \approx \frac{Z(l)^2}{T_{rr}(l)} \quad (\text{Bo weak, but } B_{\text{rms}} \text{ generated})$$

$$1/T_{rr}(l) \approx \frac{Z(l)^2}{\rho^2} T_A \quad T_A \sim \frac{l}{B_0}$$

$$\Rightarrow Z_l \sim (B_0 \epsilon)^{1/4} l^{1/4} \quad \theta_{\text{rms}}$$

$$E(k) \sim \sqrt{\epsilon k} k^{-3/2}$$

and with strong  $B_0$  :

$$\epsilon \sim \frac{Z(l_\perp)^2 Z(l_\parallel)}{\delta_\perp^2 |k_\parallel V_A|}$$

so W.T.T. Alfvénic cascade:

$$E(k_z, k_{\perp}) \sim (\epsilon V_A)^{1/2} / k_{\perp}^{1/2} k_z^2 \sim " \text{hard}" \text{ in } k_z$$

However, note:

$$z(l_z) \sim \partial B(l_z) \sim (\epsilon k_n V_A l_z)^{1/4} l_z^{1/2}$$

$\Omega$

$$\partial B \cdot \vec{v}_z \sim \frac{1}{l_z^{1/2}} (\epsilon k_n V_A l_z)^{1/4}$$

But recall:

- Alfvén wave:

$$\omega \approx k_n V_A$$

derived from:

$$\partial_t A = B_0 \vec{v}_{\parallel} \phi + \dots$$

$$\partial_t \vec{v}_{\parallel} \phi = B_0 \vec{v}_{\parallel} \vec{J}_{\parallel} + \dots$$

$$\vec{v}_{\parallel} = \underset{\text{Linear}}{\underset{\uparrow}{\partial z}} + \underset{\text{Nonlinear}}{\underset{\uparrow}{\frac{\partial B_z}{B_0} \cdot \vec{v}_{\perp}}}$$

$$\text{Ratio } \frac{\text{Nonlocal}}{\text{Linear}} = k_u = \frac{\partial B_{\perp}}{\partial z} \cdot \frac{D_{\perp}}{B_0}$$

$$\Rightarrow k_u \sim \frac{\partial B}{\partial z} \frac{D_{\perp}}{B_0 \Delta_{\perp}}$$

$\ell_{\text{loc}} \rightarrow \text{parallel auto correlation length}$

see stochastic fields discussion of Phys 235  
2016.

$\Delta_{\perp} \rightarrow \perp \text{ correlation length.}$

$$\text{Point: } B_0 \partial_z C + \partial B_{\perp} \cdot D_{\perp} C = 0$$

$$\partial_z C + \frac{\partial B_{\perp}}{\partial z} \cdot D_{\perp} C = 0$$

$k_u < 1 \rightarrow C \text{ evolves by many kicks in } \Delta_{\perp}$   
 $\rightarrow \text{diffusion}$

$\rightarrow$  in WTT wave interactions are diffusive in character.

$k_u > 1 \rightarrow C \text{ scattered} \rightarrow \Delta_{\perp}$   
in one step

$\rightarrow$  fast transport in random media  $\rightarrow$  percolation

Analogy  $\partial_t C + \nabla \cdot D C = 0$

$$k_u = \frac{\nabla T_{ac}}{\Delta_1}$$

so we have a concern:

- Physics of MHD turbulence understood in terms of Alfvén wave interactions.
- but scalings of WTT spectrum suggest that wave character lost as cascade progresses

c.f.

$$k_u \sim \frac{k_u^{-1} [\epsilon k_u v_A]}{k_\perp^{q/2}}^{1/4}$$

$\uparrow$  as  $k_\perp \downarrow$

i.e. How high can  $k_u$  go and still be consistent with physics of Alfvén Wave Cascade

⇒ Critical Balance Conjecture

(GS 75, H+P'78)

$\Rightarrow$  MHD inertial range in strong field will set at  $k_{\perp} \sim 1$ .

$$\text{c.e. } \rightarrow \partial B_{\perp} \cdot D_{\perp} \sim \frac{Z(l_1)}{\ell_1} \sim B_0 D_{\parallel}$$

$$\Rightarrow \frac{Z(l_1)}{\ell_1} \simeq k_{\perp} V_A$$

$$\rightarrow \frac{T_A}{T_{\text{Eddy}}} \rightarrow \perp \quad T_{\text{Tr}} \rightarrow T_{\text{Eddy}} \sim T_A.$$

- c.e. all timescales equalize  
 $\rightarrow k_{\perp} \sim 1$  is maximum  $k_{\perp}$  and still retain Alfvénic character.  
 $\rightarrow$  Why?

Recall:

$$-\omega \pi \quad T_{\text{c}} \sim T_{\text{ac}}$$

$\left. \begin{array}{l} \text{Triad coherence} \\ \text{set by wave dispersion} \end{array} \right\}$

$$\rightarrow \pi \delta(\omega_{\perp} - \omega_{\perp}^{\text{eff}} - \omega_{\parallel}')$$

- STT - Renormalized Theory

$$T_{\text{c}} \sim T_{\text{c}}$$

$\left. \begin{array}{l} \text{Triad coherence} \\ \text{set by nonlinear} \\ \text{scattering, etc.} \end{array} \right\}$

$$\rightarrow I / (\Delta \omega_{\perp} + \Delta \omega_{\perp}^{\text{eff}} + \Delta \omega_{\parallel}^{\text{eff}})$$

$$= \Theta_{k_1, k_2, k_3}$$

\* 5.

So, renormalized wave interaction theory  $\Rightarrow$

$$\Theta_{k_1 k_2 k''} = \frac{\Delta\omega_{k_1} + \Delta\omega_{k_2}^0 + \Delta\omega_{k''}^0}{(\omega_{k_1} - \omega_{k''} - \omega_{k_2})^2 + (\Delta\omega_{k_1} + \Delta\omega_{k_2}^0 + \Delta\omega_{k''}^0)^2}$$

$\rightarrow$  recovers both limits ✓

Now,  $\Theta_{k_1 k_2 k''}$  clearly sets  $T_{tr}$ .

so, can re-write phenomenological transfer balance as:

$$E \sim \frac{1}{k^2} \frac{Z(l_1)^2 Z(l_2)^2}{T_{tr}(l_1)}$$

$$\frac{1}{k^2} T_{tr}(l_1) = \left[ \frac{(k_1 v_A)^2 + (Z(l_1))^2}{k^2} \right]^{1/2}$$

↑  
comparable at  $k_1 \approx 1$

by analogy with  $\Theta_{k_1 k_2 k''}$ .

7.

$$\textcircled{1} > \textcircled{2} \rightarrow \text{W.T.T.}$$

$$\textcircled{1} \leq \textcircled{2} \rightarrow \text{S.T.T.}$$

$$E \sim \frac{1}{l_1} \frac{Z(l_1)^2 Z(l_1)^2}{Z(l_1)/l_1}$$

-  $Z(l_1)^3/l_1$

and  $\underbrace{Z(l_1) \sim (El_1)^{1/3}}$  - Back to  
K 4).!

Point :  $\langle Z(k)^2 \rangle \sim E^{2/3} k_1^{-5/3}$  - GS spectrum,  
but different physics! - softer than

-  $\frac{Z(l_1)}{l_1}$  vs.  $\ln V_d$

- WTT.  
- Great power  
law on sky !!

$$\left(\frac{El_1}{l_1}\right)^{1/3} \sim \frac{E^{1/3}}{l_1^{2/3}} \rightarrow \text{rate increases}$$

as  $l_1 \downarrow$   
 $\Rightarrow$  constant constant  
 $\{V_{\ln V_d}\}$

$$- \frac{Z(l_1)}{l_1} \sim \frac{E^{1/3}}{l_1^{2/3}} \rightarrow \frac{\delta B}{B_0} \cdot D_1$$

8-

then  $k_{\parallel \perp} \sim \frac{1}{t}$   $\Rightarrow$

$$\frac{\epsilon^{1/3}}{k_1^{2/3}} \sim k_{\parallel \perp}$$

$$\Rightarrow \boxed{k_{\parallel \perp} \sim \epsilon^{1/3} k_1^{2/3}} - GS cone.$$

$\rightarrow$  Critical Balance is a hypothesis.

- Plausible answer to question of "how maintain Alfvénic cascade in state of strong (i.e. non-weak) turbulence?"

- anisotropy of spectrum supported by simulations (cf. Galtier),

But

- hypothesis, only

$\rightarrow$  Computational support semi-quantitative.

$\rightarrow$   $5/3$  vs  $3/2$  etc. still ongoing.

→ A word about tricks.

In wave turbulence cascade,  
must satisfy:

$$\underline{k} = \underline{f} + \underline{\varepsilon}$$

$$\underline{\omega}_k = \underline{\omega}_p \pm \underline{\omega}_{\varepsilon} \quad (\text{WTT})$$

→ resonance  
criterion

Conditions satisfied

by:

$$q_{\perp n} = 0$$

(i.e.  $\underline{\varepsilon}$  is a cell,  
driven by beats)

$$\text{so } k_n = p_n$$

$$\underline{k}_{\perp} = \underline{p}_{\perp} + \underline{\varepsilon}_{\perp}$$

and  $\underline{\omega}_n = \underline{\omega}_p \pm \cancel{\varepsilon}$

→ - deformation of Alfvénic wave  
packet directly related to  
its interaction with 2D part  
of wave packet travelling  
in opposite direction.

- interaction passive w/r  $k_n$ .  
⇒  $\perp$  transfer in long time  
limit.

A.

## Basics of Helicity

→ Magnetic Helicity → constraint on Relaxation

- another conserved quantity in ideal MHD is magnetic helicity  $K$

$$K = \int d^3x \underline{A} \cdot \underline{B}$$

$V$  is taken to be the volume of a 'flux tube'.

- what, yet another invariant!  $|P|$

→  $K$  is different ⇒ has topological interpretation

$$K = \int_V d^3x \underline{A} \cdot \underline{\nabla} \times \underline{A}$$

→  $\underline{x} \rightarrow -\underline{x}$  flips sign of  $K$

→  $K$  is a pseudo-scalar  
i.e. has orientation or "handedness" --

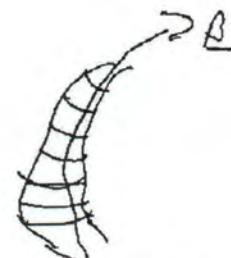
Proceed via:

- show  $K$  conservation
- discuss interpretation of  $K$
- comment on utility ⇒ Taylor Relaxation

N.B.: Important →  $K$  is gauge invariant

i.e. if  $\underline{A} \rightarrow \underline{A} + \underline{\nabla} \chi$

$$\begin{aligned} K &\rightarrow K + \int d^3x \underline{\nabla} \cdot \underline{B} \\ &= K + \int d^3x \underline{\nabla} \cdot (\underline{B} \underline{x}) \\ &= 0, \text{ to surface term. } \left\{ \begin{array}{l} \underline{B} \cdot \hat{n} = 0 \text{ on surface of} \\ \text{tube} \end{array} \right. \end{aligned}$$



Now, consider a blob of MHD fluid in motion



$$\text{can show } \frac{dK}{dt} =$$

$$\begin{aligned} \underline{E} + \frac{\underline{v} \times \underline{B}}{c} &= n \underline{J} \\ \underline{E} &= -\frac{1}{c} \frac{\partial \underline{A}}{\partial t} - \underline{\nabla} \phi \end{aligned} \quad \left\{ \begin{array}{l} \text{Field in} \\ \text{blob} \end{array} \right.$$

$\Rightarrow$

$$\frac{\partial \underline{A}}{\partial t} = \underline{v} \times \underline{\nabla} \times \underline{A} - c \underline{\nabla} \phi - cn \underline{J}$$

$$\frac{\partial \underline{B}}{\partial t} = -\underline{v} \cdot \underline{\nabla} \underline{B} + \underline{B} \cdot \underline{\nabla} \underline{v} - \underline{B} \underline{\nabla} \cdot \underline{v} + n \underline{\nabla}^2 \underline{B}$$

$$\frac{dK}{dt} = \frac{d}{dt} \int d^3x (\underline{A} \cdot \underline{B})$$

$$= \int d^3x \left( \frac{d\underline{A} \cdot \underline{B}}{dt} + \underline{A} \cdot \frac{d\underline{B}}{dt} \right) + \int \frac{\underline{A} \cdot \underline{B}}{dt} d^3x$$

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$$\frac{dK}{dt} = \int d^3x \left( \frac{\partial \underline{A}}{\partial t} \cdot \underline{B} + (\underline{V} \cdot \nabla \underline{A}) \cdot \underline{B} + \underline{A} \cdot \frac{\partial \underline{B}}{\partial t} + \underline{A} \cdot (\underline{V} \cdot \nabla \underline{B}) \right) + \underline{A} \cdot \underline{B} \cdot \nabla \cdot \underline{V}$$

where  $\frac{d}{dt} d^3x = \underline{D} \cdot \underline{V}$

$$\begin{aligned} \text{i.e. } \frac{d}{dt} d^3x &= \frac{d}{dt} d\underline{r} \cdot d\underline{l} + d\underline{r} \cdot \frac{d}{dt} d\underline{l} \\ &= -d\underline{l} \cdot \underline{D} \underline{V} \cdot d\underline{r} + (\underline{L} \cdot \underline{V})(d\underline{r} \cdot d\underline{l}) + d\underline{l} \cdot \underline{D} \underline{V} \cdot d\underline{r} \end{aligned}$$

$$= \underline{D} \cdot \underline{V} \frac{d^3x}{d\underline{r} \cdot d\underline{l}}$$

s.t. and  $\underline{B} \cdot \underline{n}$  on surface of tube.

$$\begin{aligned} \frac{dK}{dt} &= \int d^3x \left[ (\underline{B} \cdot \cancel{\underline{V} \times \underline{B}} - c_1 \cancel{\underline{D} \cdot \underline{B}} - c_2 \cancel{\underline{J} \cdot \underline{B}}) \right. \\ &\quad \left. + \underline{A} \cdot (\underline{D} \times (\underline{V} \times \underline{B})) + \underline{D} \cdot ((\underline{A} \cdot \underline{B}) \underline{V}) + \underline{A} \cdot \underline{n} \underline{D}^2 \underline{B} \right] \end{aligned}$$

$\cancel{\text{dis}}$  flux

where  $\underline{A} \cdot (\underline{V} \cdot \underline{D} \underline{B}) + \underline{B} \cdot (\underline{V} \cdot \underline{D} \underline{A}) + \underline{A} \cdot \underline{B} \cdot \underline{D} \cdot \underline{V} = \underline{D} \cdot (\underline{V} \underline{A} \cdot \underline{B})$

$$\begin{aligned} \frac{dK}{dt} &= \int d^3x \left[ \underline{D} \cdot ((\underline{A} \cdot \underline{B}) \underline{V}) + \underline{D} \cdot ((\underline{V} \times \underline{B}) \times \underline{A}) + (\underline{V} \times \underline{B}) \cdot (\underline{D} \times \underline{A}) \right. \\ &\quad \left. - c_1 \underline{D} \cdot \underline{B} - \eta (\underline{A} \cdot \underline{D} \times \underline{B}) \right] \end{aligned}$$

60.

$$\Rightarrow \frac{d\mathbf{r}}{dt} = \int d^3x \left\{ \underline{\mathbf{v}} \cdot [(\underline{\mathbf{A}} \cdot \underline{\mathbf{B}}) \underline{\mathbf{v}} + (\underline{\mathbf{v}} \times \underline{\mathbf{B}}) \times \underline{\mathbf{A}} + c_1 (\underline{\mathbf{A}} \times \underline{\mathbf{J}})] - c_1 \underline{\mathbf{J}} \cdot \underline{\mathbf{B}} - c_1 \underline{\mathbf{J}} \cdot \underline{\mathbf{B}} \right\}$$

$$= \int d\underline{s} \cdot [(\underline{\mathbf{A}} \cdot \underline{\mathbf{B}}) \underline{\mathbf{v}} + (\underline{\mathbf{v}} \times \underline{\mathbf{B}}) \times \underline{\mathbf{A}} + c_1 \underline{\mathbf{A}} \times \underline{\mathbf{J}}] \\ - 2 \int d^3x [c_1 \underline{\mathbf{J}} \cdot \underline{\mathbf{B}}]$$

$$= \int d\underline{s} \cdot [(\cancel{\underline{\mathbf{A}}} \cdot \cancel{\underline{\mathbf{B}}}) \underline{\mathbf{v}} - (\cancel{\underline{\mathbf{A}}} \cdot \cancel{\underline{\mathbf{B}}}) \underline{\mathbf{v}} + (\cancel{\underline{\mathbf{A}}} \cdot \cancel{\underline{\mathbf{v}}}) \underline{\mathbf{B}}] - c_1 \int d\underline{s} \cdot \underline{\mathbf{J}} \times \underline{\mathbf{A}}$$

$$- 2c_1 \int d^3x (\underline{\mathbf{J}} \cdot \underline{\mathbf{B}})$$

$\cancel{\underline{\mathbf{B}}} \cdot \cancel{\underline{\mathbf{n}}} = 0$ , on tube

$$= -c_1 \int d\underline{s} \cdot [\cancel{\underline{\mathbf{B}}} \cdot \cancel{\underline{\mathbf{A}}} - \cancel{\underline{\mathbf{A}}} \cdot \cancel{\underline{\mathbf{B}}}] - 2c_1 \int d^3x \underline{\mathbf{J}} \cdot \underline{\mathbf{B}}$$

$$= -2c_1 \int d^3x (\underline{\mathbf{J}} \cdot \underline{\mathbf{B}})$$

$\Rightarrow$  have shown:

$$\boxed{\frac{d\mathbf{r}}{dt} = -2c_1 \int d^3x (\underline{\mathbf{J}} \cdot \underline{\mathbf{B}})}$$

\*

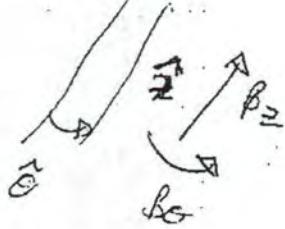
and clearly!  $\frac{d\mathcal{H}}{dt} \rightarrow 0 \text{ as } \alpha \rightarrow 0$   
 (non-singular  $J$ )

∴ Helicity is conserved in ideal MHD.  
 Magnetic, Can assign helicity to each  
 flux tube.

→ Magnetic helicity conserved, but what does it mean?

- helicity is non-trivial  $\Rightarrow$  more than just helical field lines.

interesting to note:  $\mathcal{H}(r) = \frac{r B_z}{r B_\theta(r)} = \frac{1}{R A(r)}$



$$A(r) = \frac{B_\theta(r)}{r B_z} \rightarrow \begin{cases} \text{field line} \\ \text{pitch length scale} \\ \text{length scale over} \\ \text{entire helicity} \\ \text{volume} \end{cases}$$

cylindrical plasma  $\rightarrow B = B(r)$

Now,  $A_\theta = \frac{1}{r} \int_0^r B_z dr$

$A_z = - \int_0^r B_\theta dr$

RF

$$\text{so } \underline{\underline{A}} \cdot \underline{\underline{B}} = \int_0^r B_z dr - B_z \int_0^r B_0 dr \\ = \mu B_z \int_0^r \frac{B_0}{\mu} dr - B_z \int_0^r B_0 dr$$

$$\underline{\underline{A}} \cdot \underline{\underline{B}} = B_z \left[ \mu \int_0^r \frac{B_0}{\mu} dr - \int_0^r B_0 dr \right] \\ = 0 \text{ for constant } \mu$$

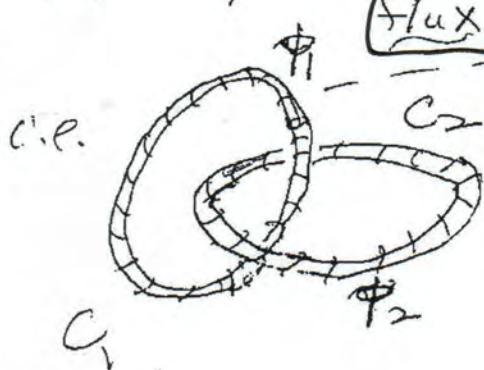
i.e. non-zero helicity requires  $\mu = \mu(r)$

i.e. — pitch varies with radius

$\Rightarrow$  magnetic shear

twist

- physically  $\rightarrow$  helicity means self-linkage of 2



$$\Phi = \int \underline{\underline{A}} \cdot \underline{\underline{B}} = \int_{x-\text{section}}^{\text{area}} \text{const}$$

tube 2:  $\Phi = \Phi_2$

field in loops, only

63

Now, for volume  $V_1$  of tube I

$$k = \int_{V_1} A \cdot B \, d^3x = \oint dl \int_{\text{surf}} A \cdot B$$

$C_1$ :  
 ↓  
 elong 100P  
 X-sect area

$$= \oint_{C_1} A \cdot dl \int_{S_1} B \cdot \hat{n} \, dA$$

$$= \oint_{C_1} \oint_{S_1} A \cdot dl$$

Now, can shrink  $C_1$ , as no field outside loops



→ in x section:



but  $\oint_{C_1} A \cdot dl = \int_{A \text{ enclosed}} B \cdot dS = \oint_{C_2}$

→ dynamics? - how does relaxation occur

→ more in discussion of kinetic  
tearing.

$$so \dots K_1 = \phi_1 \phi_2 \rightarrow \underline{\text{product of fluxes}}$$

similarly

$$K_2 = \phi_2 \phi_1$$

$$\therefore K = 2\phi_1 \phi_2$$

$$\text{if } n \text{ windings} \quad K = K_1 + K_2 = \pm 2n\phi_1 \phi_2$$

$\Rightarrow$  helicity is measure of self-linkage of magnetic configuration.

Scope

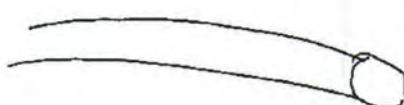
Topological complexity.

Why care  $\rightarrow$  Taylor Conjecture

(1974)  
(J.B. Taylor)

- in magnetic confinement, of great interest to determine how fields, currents self-organize

- RFP



$\sim$  toroid  
 $\sim$  toroidal current

$$\text{well fit by } B_z = B_0 J_0(\alpha r) \\ B_\theta = B_0 J_1(\alpha r)$$

$$\underline{J} \times \underline{B} = 0$$

$\Rightarrow$  why so robust?  
especially since RFP is turbulent

force free

- Taylor conjectured conservation of magnetic helicity constraints relaxation to force-free state.

Key Point - helicity conserved in flux tubes, to if

- toroidal plasma  $\rightarrow$  many small tubes



etc.

- recall Sweet-Parker model:  
magnetic reconnection / resistive dissipation effective on small scales.

$\Rightarrow$  Taylor Conjecture: At finite  $\eta$ , helicity of small tubes dissipated but)  $\underline{\text{global}}$  helicity conserved.

$$\stackrel{\text{c.e.}}{=} \int_{\text{plasma volume}} \underline{A} \cdot \underline{B} d^3x = k_0 \rightarrow \textcircled{a} \text{ conserved.}$$

$\therefore$  Taylor conjectured that optical magnetic configuration could be explained by minimum principle:

$$\delta \left[ \int_V d^3x \frac{B^2}{8\pi} + \lambda \int_V d^3x A \cdot B \right] = 0$$

i.e. minimize magnetic energy subject to constraint of conserved global helicity,

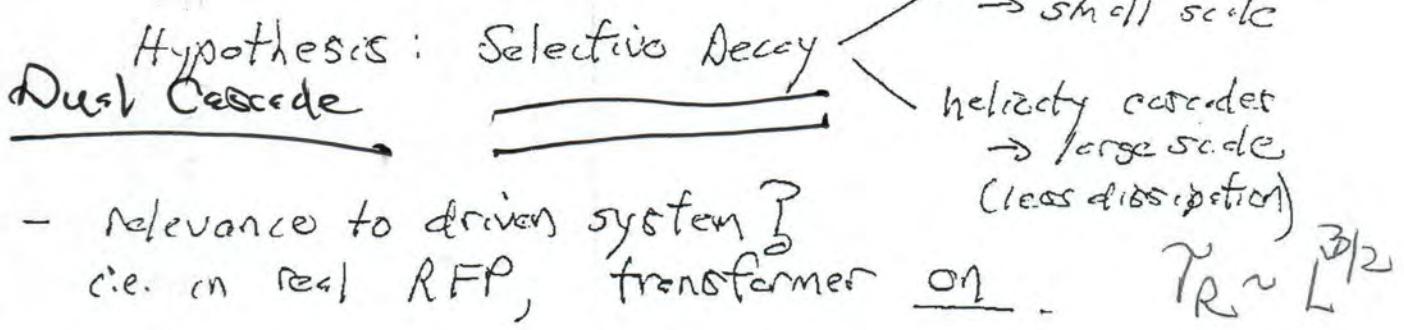
Comments:

→ it works! - indeed amazingly well - for

RFPs, spheromaks, etc. Departures only recently being discovered.

→ inspired idea of helicity injection as way to maintain configurations

→ it is a conjecture → no proof. ||



- Taylor conjectured conservation of magnetic helicity constrains relaxation to force-free state.

Key Point - helicity conserved in flux tubes to  $\eta$

- toroidal plasma  $\rightarrow$  many small tubes
- recall Sweet-Parker model : magnetic reconnection / resistive dissipation effective on small scales.

$\Rightarrow$  Taylor Conjecture : At finite  $\eta$ , helicity of small tubes dissipated but helicity conserved. global

c.e.

$$\int \underline{A} \cdot \underline{\beta} d^3x = k_0 \rightarrow \textcircled{1} \text{ conserved.}$$

Plasma volume

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$$\oint \left[ \int d^3x \frac{B^2}{8\pi} + \lambda \int d^3x A \cdot B \right] = 0$$

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Comments:

- it works! - indeed amazingly well - for RFPs, spheromaks, etc. Departures only recently being discovered
  - inspired idea of helicity injection as way to maintain configurations
  - it is a conjecture → no proof.
- Hypothesis: Selective Decay
- energy cascades
  - small scale
  - helicity cascades
  - large scale (less dissipation)
- Relevance to driven system?  
i.e. in real RFP, transformer on.

60%

→ dynamics? - how does relaxation occur  
 → more in discussion of kinks,  
 tearing.

$$\int \left[ \int d\mathbf{x} \left[ \frac{\delta^2}{\delta H} + \lambda \underline{A} \cdot \underline{B} \right] \right] =$$

$$\frac{\underline{B} \cdot \delta \underline{B}}{4\pi} + \lambda \underline{A} \cdot \delta \underline{B} = 0$$

$$\frac{\delta \times \underline{A}}{4\pi} + \lambda \underline{A} = 0$$

$\underline{U}_X$

$$\underline{I} = \mu \underline{B}$$

$$\underline{\delta \times B} = \mu \underline{B}$$

ans

$$\frac{\underline{I} \cdot \underline{B}}{B^2} = \mu$$

↓  
const

force force

$\delta J_H = 0 \rightarrow$  parallel current  
 homogenized