

PKU Lectures - 2019-2020

Magnetic Fluctuations and Resonances

Statics

$$\underline{E} \cdot \underline{B} = 0$$

Physics 235
Spring 2019

→ Stochastic Fields

→ Heat Transport

→ Fields, Flows and Particle Transport

→ Marginal Cells / Percolation

Dynamics

$$t_{\text{turb}} \rightarrow \delta$$

→ Turbulence, MHD Turbulence I
Richardson

→ MHD Turbulence II
Richardson.

→ Helicity and Taylor Relaxation

Lecture II

Physics 235

Prologue

What is this course about?

- transport in disordered / random / turbulent system, different regimes
- system evolution due such transport, i.e. relaxation

Topics:

- → Diffusion & beyond:
 - transport in stochastic magnetic fields
 - scattering + collisions → what is origin of irreversibility
 - $k_{\text{B}}T \ll k_{\text{B}}T_e$
- shear dispersion, cellular arrays, random media
- Percolation and $k_{\text{B}}T \gg 1$ transport
- Intermittency: (beyond Fokker-Planck)
- fractals, multi-scaling → scaling, power laws
- Hurst exponent (correlations)
- Fat tails, Levy Flights

- CTRW and Fractional kinetics
- Relaxation → Avalanche.
- Self-organized Criticality (SOC)
- Traffic Flow, Jams
- Models of SOC
- Turbulence preceding and evolanching
- Selected Topics, TBD.

Why?
 ↗
 ↗
 ↗

Turbulent Transport

"How many magnetic field lines in the universe?" → 1.

1.

I) Case Study: Transport in Stochastic Fields
 (Buried bodies or QLT)

A) Review — Basics of Hamiltonian Chaos
 OH, Chapt. 7 (cf. OH, and other supplement-
 ary material)

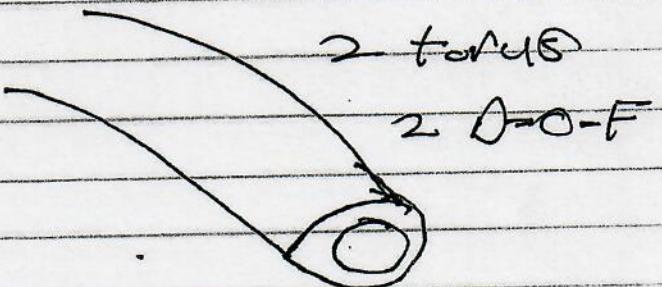
If integrable system, can write:

$$H = H_0(\underline{J})$$

$\underline{J} \equiv$ action variable
 $\underline{\theta} \equiv$ angle variable

$$\text{so } \frac{d\underline{\theta}}{dt} = \frac{\partial H}{\partial \underline{J}} = \underline{\omega}(\underline{J})$$

$$\frac{d\underline{J}}{dt} = \underline{\dot{\theta}}$$



trajectories lie on toroidal surfaces.

For 2-torus, have:

$$\omega_1/\omega_2 = p/q \rightarrow \text{rational number}$$

closed trajectory

$\omega_1/\omega_2 = \text{irrational} \rightarrow \text{ergodic trajectory}$,
 fills surface

recall: Poincaré recurrence....

Surfaces where $\omega_1 / \omega_2 = p/q$ are rational surfaces, and define natural frequencies of system

Now if perturb:

$$H = H_0(\underline{\Omega}) + \epsilon H_1(\underline{\Omega}, \underline{\Omega'})$$

then must implement perturbation theory such that canonical structures maintained, so ΔS (correction to action) needed:
 \rightarrow perturbation of Liouville eqn.

and $\Delta S \sim \epsilon H_1(\underline{\Omega})_m / \underline{\omega \cdot m}$

$$m \cdot \underline{\omega} = 0 \rightarrow \frac{\text{small denominator}}{\text{problem}} \rightarrow \text{resonance.}$$

$\omega - kv$

What happens? \rightarrow central issue in chaos theory

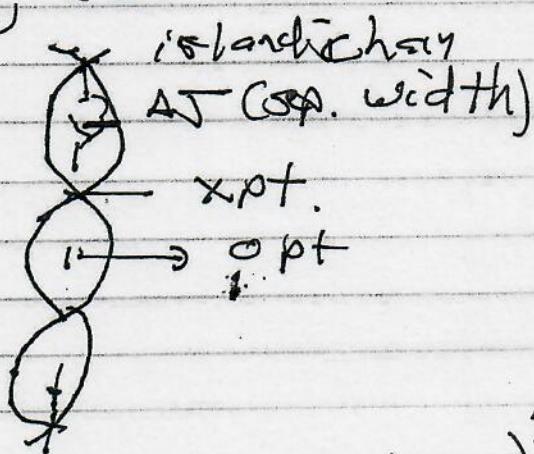
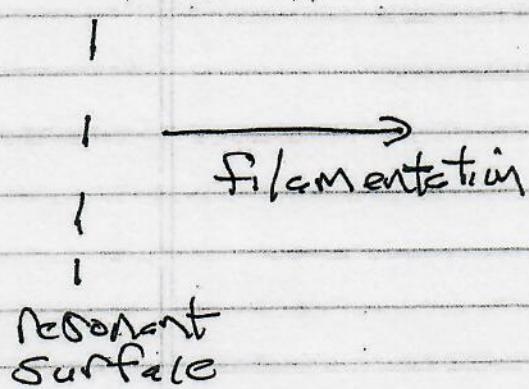
Small denominator problem \leftrightarrow resonance phenomena (n.b. also Landau resonance)

i.e. $m \omega_1 + n \omega_2 = 0$

$$m/n = -\omega_2/\omega_1 = -q/p$$

pitch of trajectory

Now can (for single resonance) resolve small denominator problem by secular perturbation theory (see Supplementary notes), so

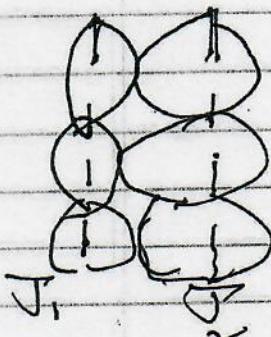


{ lines on
perturbed
surfaces}

$$\Delta T \sim \left(\frac{G H_1}{\Delta W} \right)^{1/2}$$

perturbation strength shear
(diffn/
rotation in
phase $\phi = \omega t$)

Now this fix-up works in the region of a single resonance. But if resonances overlap d.e.



- trajectories:
- wander in radius
- full volume, not surface
- chaos results

4.

Chaos:

- trajectory separation exhibits linear instability, exponentially growing

$$\Delta \sigma = \Delta \sigma_0 e^{\lambda t}$$

↑ ↑

⇒ 1 (at least) Lyapunov exponent > 0

- chaotic motion ⇒ statistical approach for prediction / characterization

⇒ Fokker-Planck Egn.

or
⇒ Hamiltonian dynamics (Liouville Thm)
+ chaos

, " Quasi-linear egn. ($f \rightarrow \text{pdf } f$)

(F-P and QLT equiv. for Hamiltonian)

$$\nabla \cdot V_H = 0$$

N.B.: Approaches limited to low ϵ

- criterium (working) for chaos:

Chirikov overlap:

\rightarrow ~~separation~~ island width

$$\frac{\Delta J_1 + \Delta J_2}{|J_1 - J_2|} > 1 \quad (\text{good working criterion})$$

\hookrightarrow spacing

$$\stackrel{\text{def}}{=} \frac{\Delta w_1 + \Delta w_2}{|B_2 - B_1|} > 1$$

- KAM theory is concerned with ruggedness of irrational surfaces but chaos onset concerned with rational surfaces.

Prime example:

Field Lines in Torus

- magnetic field lines + perturbation

$$\hat{B}_n = \sum_{m,n} B_{mn} e^{(m\theta - n\phi)}$$

- seek D_M \rightarrow diffusivity of field lines in chaotic regime

but who cares about "lines" \Rightarrow seek impact on

- heat, particle, momentum transport and

- \Rightarrow chaotic dynamics always diffusive \Rightarrow

6.

$$\text{def } K_U = \frac{\Delta \omega}{\omega_0} \cdot \frac{\delta B / B}{\Delta r} < 1$$

\downarrow

$$K_U > 1$$

Kubo #: what of $K_U > 1$?

Line Wandering/Diffusion

if $f = f(r, \theta, z) \rightarrow$ line density
i.e. magnetic flux

then, $\underline{B} \cdot \underline{\nabla} f = 0 \quad \text{i.e.} \quad \frac{df}{Br} = \frac{r d\theta}{B\theta} = \frac{B_0 dz}{B_0}$
 z as time

$\text{so if } \underline{B} = B_0 \hat{z} + B_\theta(r) \hat{\theta} + B_r \hat{r} + \tilde{B}_\phi \hat{\phi}$
 $\begin{matrix} \text{totoidal} \\ \text{strong} \end{matrix} \quad \begin{matrix} \text{poloidal} \\ \text{strong} \end{matrix}$

then - Hamiltonian system $\frac{dr}{dz} = \tilde{B}_\phi$
- $f \rightarrow$ phase space density, $r d\theta / dz = \langle B_\theta \rangle + \tilde{B}_\phi$

$$B_0 \partial_z f + \frac{B_\theta(r)}{r} \partial_\theta f + \underline{B} \cdot \underline{\nabla} f = 0$$

$$\partial_z f + \frac{B_\theta(r)}{B_0 r} \partial_\theta f + \frac{\underline{B} \cdot \underline{\nabla} f}{B_0} = 0$$

$$\Rightarrow \partial_z f + \frac{1}{Rg(r)} \partial_\theta f + \frac{\tilde{B}}{B_0} \cdot \underline{\nabla} f = 0$$

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- N.B.: $z \rightarrow$ plays role of time
- periodicity of fast scale perturbations
 - irreversibility of $\langle f \rangle$ evolution

$\Omega \rightarrow$ periodic
 $\text{so, for } \langle f \rangle,$ 2 scales

$$\partial_z \langle f \rangle + \frac{\partial}{\partial r} \left\langle \frac{\tilde{B}_r}{B_0} \tilde{f} \right\rangle = 0$$

$$F_{rB} = \left\langle \frac{\tilde{B}_r}{B_0} \tilde{f} \right\rangle \quad \text{so Fick's Law.}$$

+
flux of line density

How close?

Now characteristics of Liouville Eqn.
 \Rightarrow equations of lines

$$\frac{dr}{B_r} = \frac{d\theta}{\langle B_r(r) \rangle + B_0} = \frac{dz}{B_{z0}}$$

so radial excursion given by:

$$\langle \text{prior} \rangle = \int_{\text{or}}^{\ell^+} \int_{\text{or}}^{\ell} b_{\alpha} b_{\beta} \langle b_{\alpha} b_{\beta} \rangle \delta_{\alpha \beta}$$

$$dr/dz = \tilde{B}_r/B_0$$

$$\therefore dr \approx \int_{\text{or}}^{\ell} (\tilde{B}_r/B_0) dz$$

$$\begin{aligned} d\cdot r &= \int_{\text{or}}^{\ell'} \tilde{B}_r dz \\ dr &= \int_{\text{or}}^{\ell''} b_r dz \end{aligned}$$

Now, into trajectory de-wavers from perturbation for $\ell > \text{lac}$

\rightarrow autocorrelation

length

$-k_{\perp} \ell z$

$$\text{lac} \equiv 1/|\Delta(k_n)|$$

i.e. $\frac{\text{inverse spatial}}$
bandwidth

$$\therefore \left\{ dr \approx \text{lac} \tilde{B}_r/B_0 \right\} \rightarrow \begin{cases} \text{size excursion of} \\ \text{lac} \end{cases}$$

Can identify $\Delta r \equiv$ scatterer radial correlation length (i.e. spatial spectral width) \rightarrow radial coherence length.

$$K_u \equiv \Delta r/\text{Ar} \equiv \frac{\text{lac}}{\Delta r} \tilde{B}_r/B_0 \rightarrow \text{kicks #}$$

and can then post:

$\rightarrow K_u \ll 1 \Rightarrow$ many kicks on coherence length
 \Rightarrow diffusion process

~~Q-~~
9-

$k_{\perp u} \approx 1 \rightarrow$ B.R.K. "natural state" of EM turbulence
 $\{ k_{\perp u} \approx 1 \rightarrow$ critical balance.

$\rightarrow k_{\perp u} > 1 \rightarrow$ more than one Δ_n in $k_{\perp u}$
 \rightarrow strong scattering \leftrightarrow percolation.

QLT

Here $k_{\perp u} \leq 1$, at first. So, proceed via Quasiclassical theory.

$$\Gamma_M = \left\langle \tilde{B}_r \tilde{f} \right\rangle$$

$$= \sum_n \frac{\tilde{B}_{r,n}}{B_0} \tilde{f}_n$$

$$- c \left(k_z - k_{\perp} \frac{B_0}{B_0} \right) \tilde{f}_n = - \tilde{B}_{r,n} \frac{\partial \langle f \rangle}{\partial n}$$

So

$$\Gamma_M = - D_M \frac{\partial \langle f \rangle}{\partial n}$$

$$D_M = \sum_n \left| \frac{\tilde{B}_{r,n}}{B_0} \right|^2 \pi \delta(k_z - k_{\perp} \frac{B_0}{B_0})$$

$$\text{magnetic diffusivity} = \sum_n \left| \frac{\tilde{B}_{r,n}}{B_0} \right|^2 \pi \delta(k_{\perp u}) \quad (\text{RSTZ } \vec{B}_0)$$

$$\approx \left\langle \left(\frac{\tilde{B}_r}{B_0} \right)^2 \right\rangle_{\text{loc}}$$

What is \vec{B}_0 ?

4.

x0.

10.

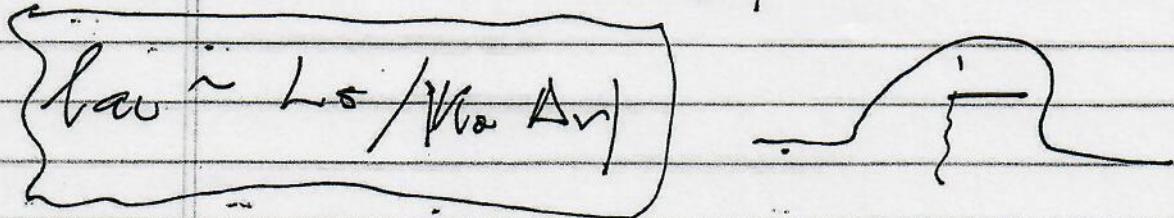
bar

$$\text{N.B. : } \sum_{\underline{n}} = \sum_{m,n} \quad \begin{array}{l} \text{Spatial spread} \\ \text{scatters} \rightarrow l_{sc} \end{array}$$

$$n = \frac{m}{g}, \quad dn = \frac{dm}{g^2} \approx' dx$$

$$\Rightarrow \text{spatial scale of spectral width } (\Delta r)$$

$\text{sets } |k_w| \sim \left| \frac{k_0 \Delta r}{L_0} \right|$



Lined then diffuse sc:

$$\langle dV^2 \rangle \sim D_M z$$

Broaden Recatino \rightarrow Orbit averaging.

N.B. Line Liouville eqn. can be obtained
by reducing/simplifying in OKE

$$\frac{\partial F}{\partial t} + v_{||} \hat{B}_0 \cdot \nabla F + v_{\perp} \cdot \nabla F - \frac{e}{B} \nabla \phi \times \hat{z} \cdot \nabla F$$

$$+ v_{||} \frac{d\hat{B}_1}{B_0} \cdot \nabla F - \frac{e}{m_e} F_{||} \frac{\partial F}{\partial V_{||}} = C(F)$$

11.
S. X-

$$\Rightarrow \nabla_0 \cdot \nabla F + \frac{d}{dr} \frac{\partial B_0}{B_0} \cdot \nabla F = 0 \quad \checkmark$$

Scales

Now, scales:

$B_{ac} \rightarrow (\text{scatters})$

\rightarrow field, line memory length
self-coherence of scattering
field.

$l_0 \rightarrow$ line deconvolution length
(length over which line scattered)

c.e. $\frac{r d\theta}{dz} = \frac{B_0(r)}{B_0}$ from its up to

but $r \sim$ scattered, \Rightarrow

$$\frac{dy}{dz} = B_0(r) + \frac{B_0'(r)}{B_0} dr$$

$$\frac{dy}{dz} \approx \frac{B_0'(r)}{B_0} dr$$

$$\langle dy^2 \rangle = \left\langle \left(\frac{B_0'}{B_0} dr dz \right)^2 \right\rangle$$

~~Ex.~~ $\frac{12}{12}$

\Rightarrow

$$\langle \delta y^2 \rangle \sim \frac{B_0^{1/2}}{B_0} z^2 \langle (\delta r)^2 \rangle$$

$$\sim \frac{B_0^{1/2}}{3B_0} D_M z^3$$

also

$$\langle \delta x^2 \rangle \sim D_M \frac{z^5}{3} \quad \text{on } 10$$

For orbital deacceleration length:

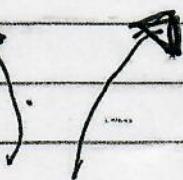
$$k_0^2 \langle \delta y^2 \rangle \sim k_0^2 \frac{B_0^{1/2}}{3B_0} D_M z^3$$

\Rightarrow

$$\frac{L_0}{k} \sim \left(\frac{k_0^2}{3} \frac{B_0^{1/2}}{B_0} D_M \right)^{-1/3}$$

anslagged
to shear
dispersion
 $\left(\frac{k_0^2}{3} \frac{V^2}{B_0} D \right)^{1/3}$

$$\sim \left(\frac{k_0^2}{3} \frac{D_M}{B_0} \right)^{-1/3}$$

Also: 

orbital excretion length
(stretching)
 \rightarrow stretching

show with 2pt. $\langle \delta x(\tau) \delta x(\tau') \rangle$

chaotic \rightarrow stretching