

PKU - HUST Lecture Series

Lecture VII

= More on Avalanches and Turbulence & reading

①

= As time (and energy) in short supply :

⇒ Rather a general discussion

⇒ Conceptual rather than detail - intensive.

⇒ Plenty of references ...

Good area for projects ---- can discuss

②

Avalanches - Review

BTW

SOC



Ph I / S

Dynamical
of SOC
States

⇒ "Noise propagates thru the scaling clusters by means of a domino effect upsetting the minimally stable states"

⇒ Space - time propagation.

percolative
transport

⇒ Trends :

$$P \sim \frac{1}{\Delta}$$

Zipf's Law

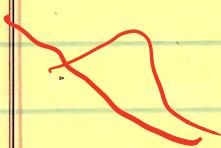
$$\Im(\omega) \sim 1/\omega$$



→ We want an equation ...

$$\text{J R Symmetry} \rightarrow x \rightarrow -x \\ \delta p \rightarrow -\delta p$$

Ginsburg -
Leaden



$\Rightarrow \Gamma^v(\mathcal{O}_P)$ invariant

$$\partial_t \delta p + \partial_x \tilde{F}(\delta p) - D_a \partial_x^2 \delta p = S$$

⇒ Burgers!

, long wavelength.

$$\partial_t \delta p + \cancel{\partial_x \delta p^2} - D \tilde{\partial_x} \delta p = \tilde{S}$$

(long wavelength limit)

$$\partial_x v + u \partial_x v - r \partial_x^2 v = 0$$

Crank → N.O. corrections?

RE

Coupled word pile

So:

→ shock as model for Austracnehe

→ Pulse velocity

$$V(\delta p) \sim \sqrt{\delta p}$$

[big pulses go faster]

noisy

→ $\delta x \sim c_f$

("Ballistic")

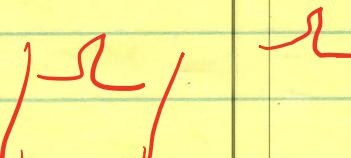
Burgers Solution:

($\tilde{S} \rightarrow 0$)

Whitham

$$\partial_t \delta p + \partial_x [\kappa \delta p^2 - D \partial_x \delta p] = 0$$

$$\partial_t \int dx \delta p = 0$$



area of pulse profile
invariant in time, even with

$D \approx 1$.

$\rightarrow 50$

$$Re = \frac{1}{\rho} \int_{-\infty}^{\infty} \rho dx$$

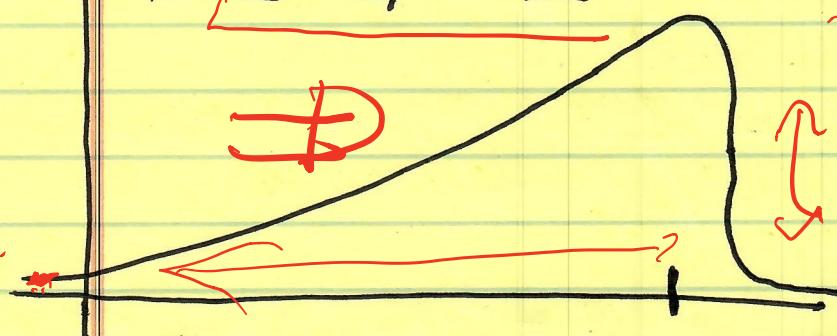
$$\Rightarrow \int_{-\infty}^{\infty} \rho dx \rightarrow \text{const}$$

why?

so (see Whitham), pulse shape:

$$= \pm \left[v \cdot (t/2A)^{1/2} \right]$$

i.e. $F = C_0 + A \delta(x)$
impulse



A cons.

$$AL \rightarrow AL \rightarrow$$

$$R^{-1} * (2At)^{-1/2} \xrightarrow{A}$$

i.e.

$$v \sim \frac{x}{t}$$

$$(0 < x < 2\sqrt{At})$$

TTP

details noise matter

but

Where's the Physics?

Direction

③ → Can we learn from REAL avalanches?

What's a real avalanche?



→ Flood Wave



→ Flood Wave \rightarrow Kinematic Waves

"Kinematic" waves:

(Lighthill + Whitham)

$$\partial_t \rho + \cancel{(\nu(\rho) \rho)} - \cancel{\nu \frac{\partial^2}{\partial x^2} \rho} = 0$$

$\nu(\rho)$ specified

\rightarrow regularization

$$\partial_t \rho + \partial_x Q(\rho) = 0$$

$$Q(\rho) = \nu(\rho) \rho$$

④ prop speed:

(cogn r)

$$\partial_t \rho + \partial_x Q(\rho) = 0$$

r

$$\partial_t \rho + \frac{f Q}{\partial \rho} \partial_x \rho = 0$$

$$C(\rho) = \frac{\partial Q}{\partial \rho}$$

$C(\rho)$

$$c(\rho) = v(\rho) + \rho \frac{\partial v(\rho)}{\partial \rho}$$

$v(\rho)$

$c(\rho) \uparrow$

of course - $c(\rho)$ increasing with ρ
 decreasing \Rightarrow shock structure.

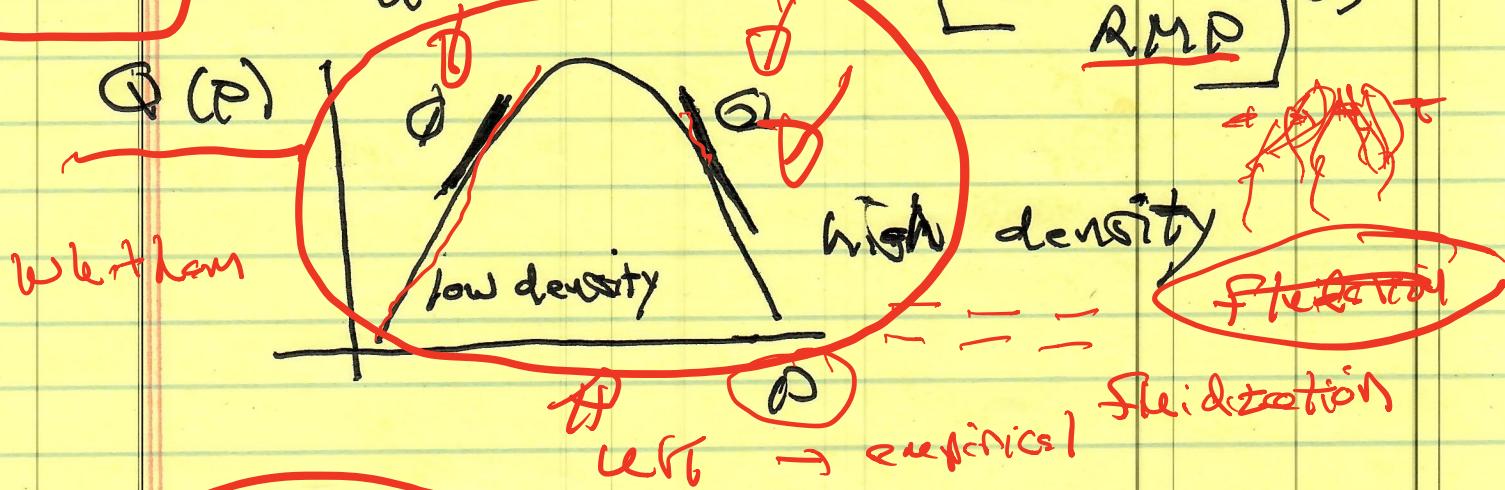
classic: Traffic Flow

$\rho =$ car density

$v(\rho)$

$v(\rho) :=$

[see Helbing, RMP]



① $\delta Q / \delta \rho > 0 \rightarrow$ usual

\Rightarrow overtaking, forward shock

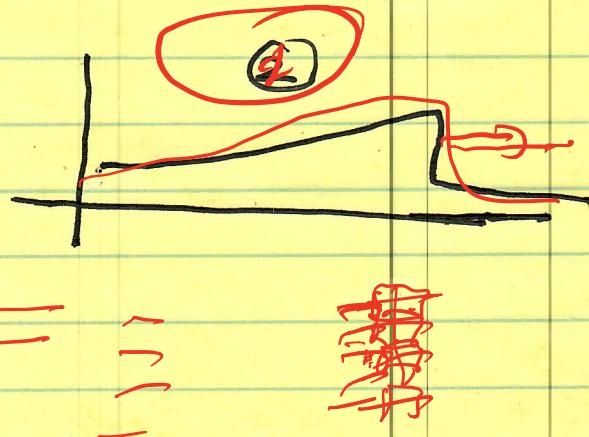
② $\delta Q / \delta \rho < 0 \rightarrow$ backward shock
 high density goes down

$v(c)$?



not

① as if coming upon toll booth



→ Beyond kinematic:

$$\partial_t \rho + \partial_x Q = 0$$

$$Q = Q(c) - v \partial_x \rho$$

$$v = v(c) - v \rho \partial_x$$

Now:

$$\partial_t \rho + \partial_x (\rho v) = 0$$

Dyn. shock \rightarrow discontinuity

whereas

gas diff

P

Cross

8.

$$\frac{\partial_t V}{T} + V \frac{\partial_x V}{T} = -\frac{1}{T} [V - \bar{V}(P) + \frac{V}{P} P_x]$$

i.e. V relaxes to $\bar{V}(P)$

(distribution) ?

Important:

$1/\tau$

→ rate of relaxation

τ = driver reaction time

$$\tau = \gamma \left[\frac{1}{\text{none}} \right] \quad \begin{cases} \text{none} \rightarrow \text{short} \\ \text{lots} \rightarrow \text{long} \end{cases}$$

Note: Another appearance of the mysterious time

delay (recall LCO). !

Now, if linearized:

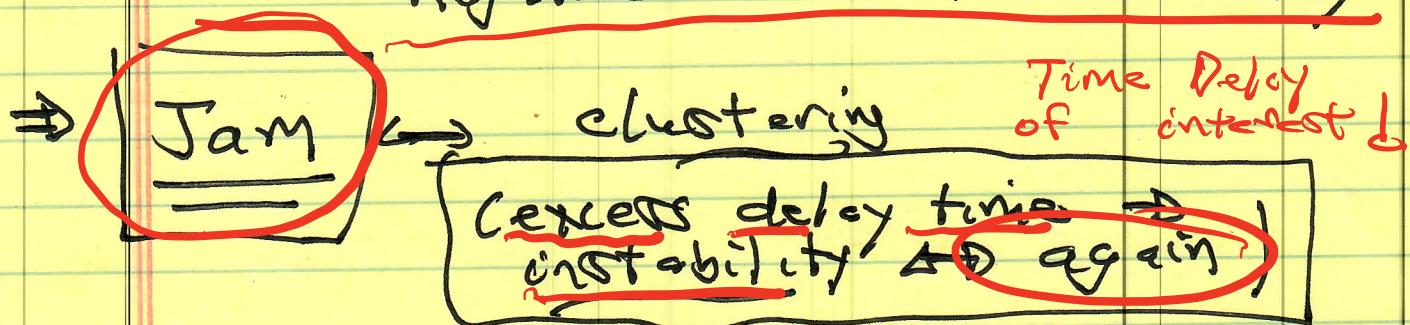


$$r \rightarrow r \rightarrow (v_0 - c_0)^2 T$$

$$\partial_t^2 r + c_0 \frac{\partial r}{\partial x} = \left[\check{v} - (v_0 - c_0)^2 T \right] \frac{\partial^2 r}{\partial x^2}$$

Long $T \rightarrow r - (v_0 - c_0)^2 T < 0 \rightarrow$

negative diffusion instability



N.B. In avalanches (heat) :

$$\bar{V} \Rightarrow \langle Q \rangle - \text{mean heat flux}$$

$$V \Rightarrow Q - \text{instantaneous heat flux}$$

Heat Flux Jams \Rightarrow



T Jamtons

Layers, staircase ; etc.

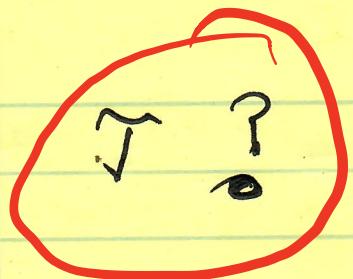
Kosyagin P.G. Gurcan

PRL (2012)
+ POP

$$1/\gamma \rightarrow 1/k_L$$

10.

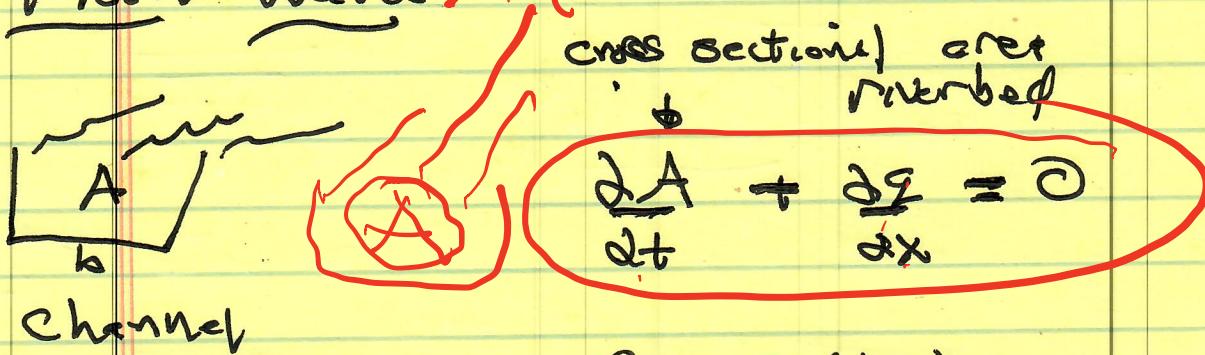
Key:



→ Oct by time dynamics
of heat flux response,
(C.f. Yan Qinghai, PD
ongoing)

T → larger near meander, etc.

→ Flood Waves:



$$\underline{Q = Q(A, x)}$$

$$\underline{\frac{\partial A}{\partial t} + \frac{\partial Q}{\partial A} \frac{\partial A}{\partial x} = - \frac{\partial Q}{\partial x}}$$

const. flux

$$\underline{C = \frac{\partial Q}{\partial A}} = \frac{1}{b} \frac{\partial Q}{\partial h}$$

$$\underline{V = Q/A}$$

if neglect R.H.S. :

Flood Wave

δ
avalanche

$$\partial_t A + CCA|A_x = 0$$

~ kinematics.

hydrodynamics

shallow water theory



→ Beyond kinematics:

h

$$\partial_t h + \partial_x (hv) = 0$$

mass continuity

sigma stress

$$\partial_t (hv) + \partial_x \left(hv^2 + \frac{1}{2} g' h^2 \right) = g' h S - C_F v^2$$

gravity

S

$\frac{\partial}{\partial t}$

friction

C_F

hydraulics



→ downhill avalanche of fluid ... Whitham



Whitham

$$\rightarrow \partial_t h + \partial_x (hv) = 0$$

and RHS balance

$$v \approx \left(\frac{g' S}{C_F} h \right)^{1/2}$$

$$\rightarrow V_0$$

shocks

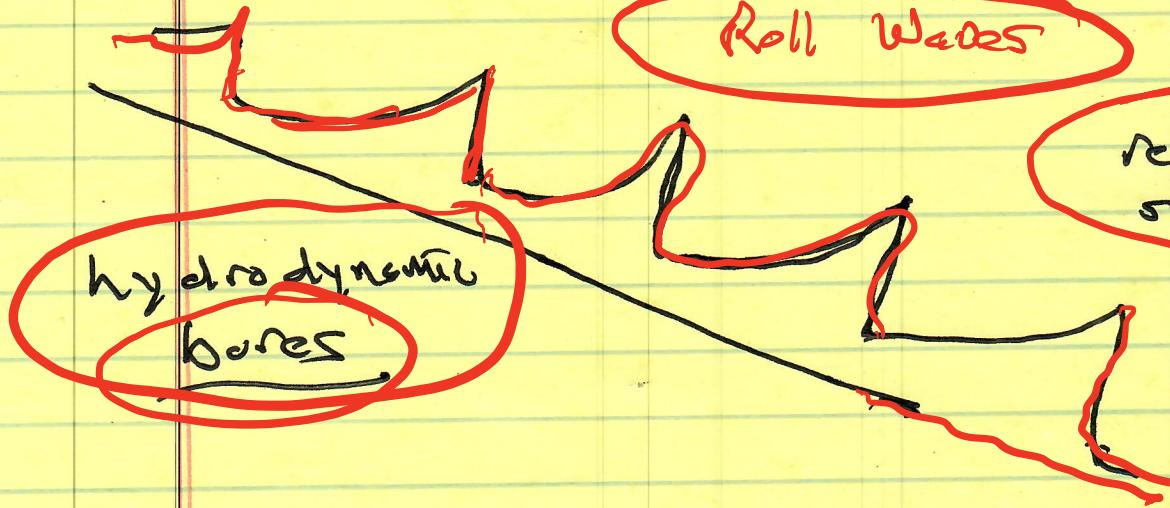
⇒ roll waves.

→ type of kinematic wave

Roll Waves

represented
shock train

hydrodynamic
bores



Can go unstable and jam for

$$C_0 < C_- \quad , \quad C_0 > C_+$$

$$C_+ = V_0 + \sqrt{g h_0}$$

$$C_0 = \frac{3V_0}{2}$$

$$C_- = V_0 - \sqrt{g h_0}$$

→ $V_0 \sim 2\sqrt{g h_0}$

⇒ $\zeta < 4 C_F$

Model =
here + turb. vibs

- Promising model. ...
- See Balmforth, et al.

2004

Incorporates turbulent mixing of
dense rare

shocks: []



- Interesting problem :

Noisy flood wave model of
Avalanche

$\sum \epsilon > 0$

Model avalanche



2018

Turbulence Spreading

Jal Flor. Phys.
soc.

See Hahn, P. D.
review + refs
therein

- Why ?

- ubiquity of turbulence, apart

ITB, ETB

- Penetration of stable zones



- Old idea = $K - \epsilon$ models

$K \equiv$ turbulence energy

Spreading

\propto ρ

$$K(t) \approx \langle \delta z^2 \rangle$$

$$\frac{\partial K}{\partial t} + \langle u \cdot \nabla K \rangle = - \frac{1}{\tau} \sqrt{K} \frac{f}{l} \cdot \nabla K = - \frac{K^{3/2}}{l} \frac{f}{\tau} t \dots$$

$\frac{P}{Re}$
Reynolds

$E \rightarrow E(t)$

"Spreading" = spreading of turbulence
by self-scattering

→ "Toward" → entrainment
street wall shear flow

→ Obviously,

$$\Gamma = -\sqrt{k} l \quad D \text{ } \textcircled{H}$$

\uparrow related

$$\Gamma = -\times \textcircled{\delta p^2}$$

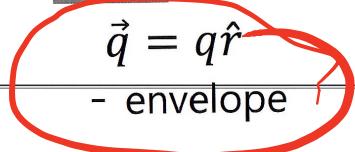
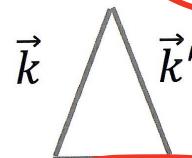
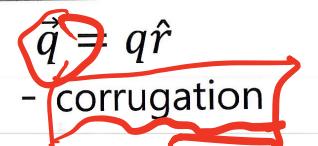
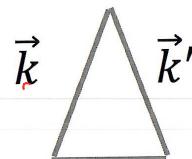
overturning,
pulse propagation

\Rightarrow turbulence
propagation.

Q: What is the relation / difference
between turbulence spreading
and avalanching?

Turbulence Spreading vs Avalanching

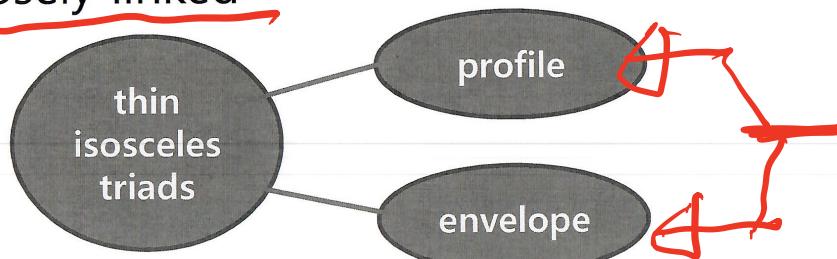
- Both: (non-Brownian) radial propagation of excitation
- **Avalanching:**
 - via overturning and mixing of neighboring cells
 - Coupling via $\nabla \langle P \rangle$
 - $\partial_t \delta P \sim \partial_x (\alpha \delta P^2)$
- Turbulence spreading (t.s. by T.S.)
 - via spatial scattering due nonlinear coupling
 - Couple via turbulence intensity field
 - Usually $\partial_t I \sim \partial_x (D_0 I \partial_x I)$



I

• Bottom Line:

- Very closely linked



- ~ impossible to have one without other
- t.s. can persist in strong driven, non-marginal regimes
- Which effect more dramatic is variable → specifics?
- Controversy sociological (or sociopathic)...

Physics + NLD \rightarrow MFE



\rightarrow The NL Proto-type : Non linear
Diffusion

Insert disked

J. D. Murray

$$\frac{\partial n}{\partial t} = - \nabla \cdot J$$

Mathematics
Biology

$$J = -D(n) \nabla n$$

$$\frac{\partial D}{\partial n} > 0$$

\approx 1D :

$$\frac{\partial J}{\partial t} = \frac{\partial D \partial P \partial T}{\partial x}$$

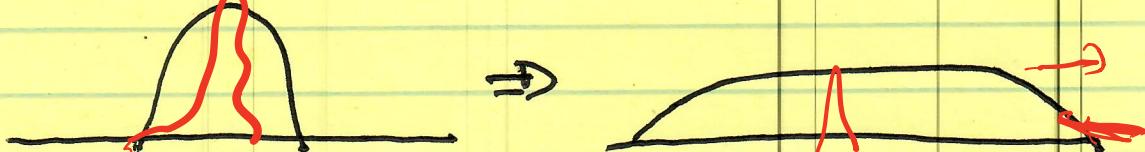
$$\frac{\partial n}{\partial t} = D_0 \frac{\partial}{\partial x} \left[\left(\frac{n}{n_0} \right)^m \frac{\partial n}{\partial x} \right]$$

NL Heat Conduction

Point : $n \rightarrow 0 \Rightarrow D \rightarrow 0$

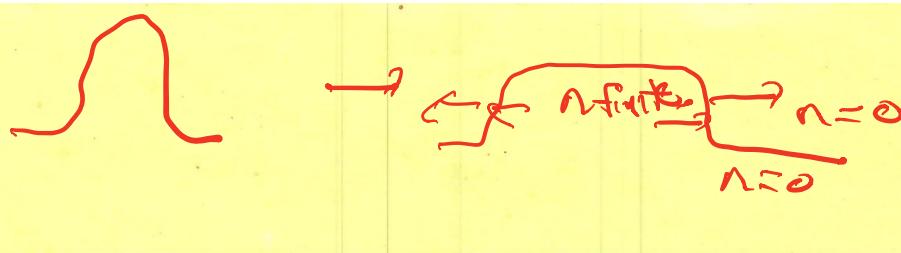
δC

\rightarrow Front :



$$\frac{D_0 t}{r_0^2} = \frac{x^2}{Dt}$$

NL \rightarrow Front
not Ballistic



Can solve exactly:



$$n(x, t) = n_0 [\lambda(t)]^{-1} \left[1 - \left(\frac{x}{r_0 \lambda(t)} \right)^2 \right]$$

$|x| < r_0 \lambda(t)$

$$= 0 \quad , \quad |x| > r_0 \lambda(t)$$

$$\lambda(t) = (t/t_0)^{1/2+m}$$

$$t_0 = R^2 M / 2 D_m (m+2)$$

$$r_0 = \sqrt{\frac{R^2 (m+3/2)}{[\pi^{1/2} \Gamma(m+1)]}}$$

Front: $x \sim r_0 \lambda(t)$

$$\sim r_0 (t/t_0)^{1/2+m}$$

(blower then diffusion)

$$x \sim \sqrt{D_m t}$$

$$\partial_t \mathbf{r} - \partial_x \mathbf{b}_0 \cdot \partial_x \mathbf{u} = \mathbf{r}, \dots$$

→ Other issues:

- with growth → back to Fisher!

See Gurcan, et.al. 2005

- Subcritical? → See Heinonen, P.D. 2018.
 ↗ ?

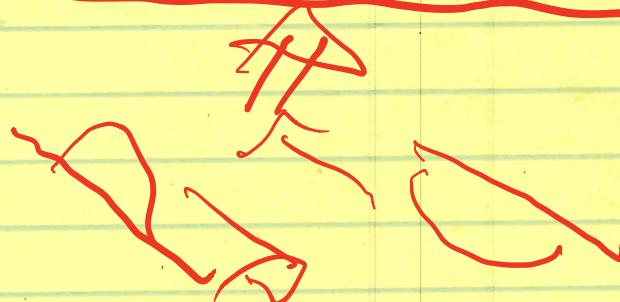
- local or non-local → See Yen, P.D.
in preparation ...
 ↗ ?

→ Spreading of mysterious, oft-invoked god Matter !!

a "deus ex Machina"

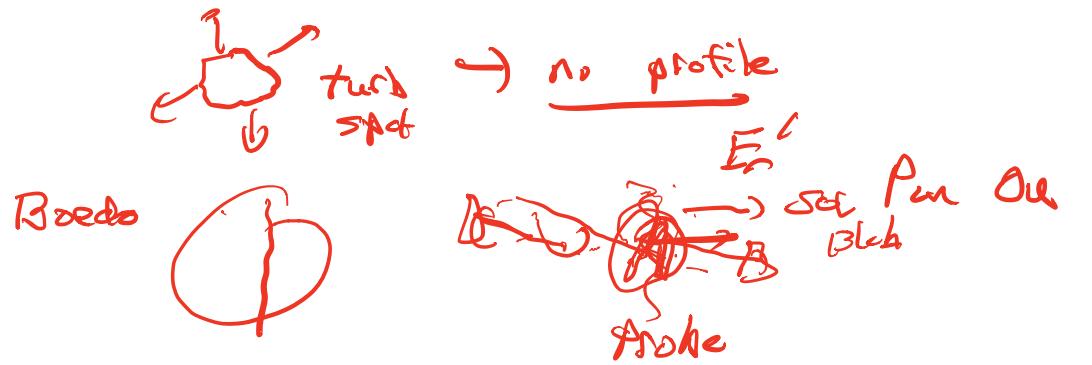
→ Not easy to understand by simulation

→ Should use theory to clarify !



$E_r \rightarrow$ Idemur

Avalanching \rightarrow profile retracted



Mo Choi JPS Y. Xu NF
2018-19

Driver \rightarrow Avalanches observed δ_{in} exper.
ex. $T < \frac{L_0}{D}$
SOC (Sandpile) example syst \rightarrow avalanches

Model

\hookrightarrow Avalanches more general than SOC

$$P \sim 2/\Delta$$



\rightarrow Physical system Avalanche

- JPS + Burgers



- Flood wave physics

turbulent

Turbulent waves

$$\partial_t \phi + \partial_x (u \phi) / \varepsilon$$

— — Bare

0(0)