

PKU Lectures - Lecture IV.

→ PV Dynamics, Charney - DeZily,
Modulation

- A closer look at PV.
- Momentum Theorem for Flows
 \Leftrightarrow PV and WMD.
- Modulations/ Interaction and
Predator - Prey Model.

TH

$$q = \rho v = \underbrace{\nabla^2 \phi}_{\substack{\text{Fluid} \\ \text{local}}} + \beta y$$

↓
potential
surface

↓
planetary

→ 2D conservative advection, according
Hamiltonian system

$$\frac{dq}{dt} = -\nabla \phi \times \hat{e}^1$$

→ M.B.:

$$\frac{dq}{dt} = 0 +$$

$$q = \begin{cases} \nabla^2 \phi & \rightarrow 2D \\ \nabla^2 \phi + \beta y & \rightarrow QG \\ \ln \frac{p_1}{p_0} + \phi - \nabla^2 \phi & \rightarrow HM \end{cases}$$

+
⋮

↔ obvious analogy to:

$$\frac{df}{dt} = 0, \quad \partial_t f + v \partial_x f + \sum_m E_m v f = 0$$

$$= C(f)$$

with collisions

$$H = \frac{mv^2}{2} + q\phi$$

$$f = \langle f \rangle + \delta f$$

i.e. A few observations:

H-W system \rightarrow

$$(0 \cdot \nabla = 0)$$

$$\frac{d}{dt} \nabla^2 \phi = + D_{11} \nabla_1^2 (\phi - n) + r \nabla^2 \nabla^2 \phi$$

$$\frac{d}{dt} n = + D_{11} \nabla_1^2 (n - \phi) + D_0 \nabla^2 n$$

includes
 $\frac{\partial}{\partial n}$

n.b. \tilde{T}_i conserves

i.e.

$$\frac{d}{dt} (n - \nabla^2 \phi) = 0$$

$$PV = \mathcal{E} = \mathcal{E}_{\text{out}} + \mathcal{E}_{\text{pol.}}$$

$n \rightarrow$ electron/GC density

$\nabla^2 \phi \rightarrow$ polarization charge (i.e. QC)
(\sim ion drift \rightarrow pol. drift off surface)

$$PV = n - \nabla^2 \phi$$

\rightarrow conserved along orbits

Reynolds force \rightarrow flux polarization charge.

N.B.

Trade offs possible! \rightarrow coupling

to flows,

$$\frac{dz}{dt} = 0$$

$$z = \bar{n} \bar{v} \bar{\phi}$$

$$\Rightarrow \langle \tilde{v}(z) \rangle = \underbrace{\langle \tilde{v} \tilde{n} \rangle}_{\text{Can transport PV via particle flux}} - \underbrace{\langle \tilde{v} \bar{\sigma} \bar{\phi} \rangle}_{\text{"bad" - transport and/or via Viscosity Flux}} \quad \text{ZF}$$

Can transport PV via particle flux
 ↳ "bad" - transport and/or
 via Viscosity Flux \Leftrightarrow "good" - ZF

→ A Closer Look at ρV and
What it Means

" What is EFD [Low Freq. Dyn. Mag. Plasma] "

- The Fluid Dynamics of Potential
Vorticity "

- Rick Salmon (P.D.-insert)

Recall $\frac{\omega}{\rho}$ frozen-in :

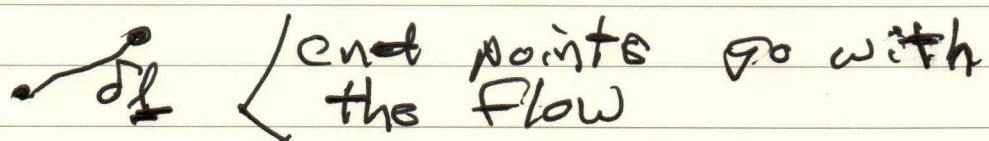
$$\frac{d}{dt} \left(\frac{\omega + 2\Omega}{\rho} \right) = \left(\frac{\omega + 2\Omega}{\rho} \right) \cdot \nabla V$$

$\frac{\omega}{\rho}$ force $\nabla \cdot V = 0$

$\frac{\omega}{\rho}$ for $\nabla \cdot V \neq 0$

$\frac{\omega + 2\Omega}{\rho}$ with rotation,

Why "frozen in"



"Frozen in"

$$\underline{\omega} \quad \text{[redacted]} \quad \frac{d}{dt} \underline{\delta f} = \underline{\omega} \cdot \underline{\nabla} V$$

$$\Rightarrow \frac{d}{dt} \left(\frac{\underline{\omega} + 2\underline{\Omega}}{\rho} \right) = \left(\frac{\underline{\omega} + 2\underline{\Omega}}{\rho} \right) \cdot \underline{\nabla} V$$

$$\frac{d}{dt} \underline{\delta f} = \underline{\delta f} \cdot \underline{\nabla} V$$

→ same equation $\Leftrightarrow \frac{\underline{\omega} + 2\underline{\Omega}}{\rho}$ frozen in

Also observe, for:

ψ = positive scalar field

$$\frac{d\psi}{dt} = 0$$

$$\frac{d}{dt} (\psi_1 - \psi_2) = 0$$

$$\int \underline{\nabla} \underline{\delta f}^2$$

$$\underline{\delta \psi} = \underline{\nabla} \psi - \underline{\delta f}$$

and

$$\frac{d}{dt} (\underline{\nabla} \psi \cdot \underline{\delta f}) = 0$$

$$\delta \ell \nrightarrow \frac{\underline{\omega} + 2\underline{\Omega}}{\rho}$$

as
satisfy some

80

$$\frac{d}{dt} \left(\frac{(\underline{\omega} + 2\underline{\Omega}) \cdot \nabla \Psi}{\rho} \right) = 0$$

latitude
 θ

$$\frac{q}{\rho} = \frac{(\underline{\omega} + 2\underline{\Omega}) \cdot \nabla \Psi}{\rho} \equiv \text{partial defn of } PV$$

\uparrow
PV
charge
 \uparrow
depth

$$Q = \int d^3x \sum q \quad \text{gt total PV}$$

charge

$$\frac{dQ}{dt} = 0 \rightarrow \text{PV conservation}$$

Symmetry \leftrightarrow Particle re-labeling symmetry

Particular can be re-labeled without changing the thermodynamic state

11

How do we get something useful out of this? (from ion)

→ Specialize to Plasma (ions)

$$\frac{d}{dt} \left[\frac{(\underline{\omega} + \underline{\Omega}) \cdot \nabla \psi}{\rho} \right] = 0$$

For plasma, $\rightarrow \underline{B}_0$

$$\underline{\Omega} = \underline{\omega}_z \hat{z}$$

$$\rho = n_0(r) + \tilde{n}$$

$$\nabla \psi = \hat{z} \quad (\text{const. vector})$$

$$\frac{d}{dt} \left[\frac{\omega_z + \Omega_0}{n_0(r) + \tilde{n}} \right] = 0$$

$$\Rightarrow \rightarrow \text{includes } \nabla \cdot \nabla n_0$$

$$\frac{1}{n_0} \frac{d}{dt} \tilde{\omega}_z - \frac{\Omega_0}{n_0^2} \frac{d}{dt} \tilde{n}_0 = 0$$

$$\tilde{\omega}_z = \frac{c}{B_0} \nabla^2 \psi$$

$$\frac{\tilde{n}_0}{n_0} = \frac{\tilde{n}_0}{n_0} = k \frac{1}{B} \phi$$

Date

50

$$\frac{d}{dt} \frac{c}{B_0} \vec{V} \cdot \vec{\phi} - \frac{N_0}{B_0} \left(\frac{\partial}{\partial t} \frac{c \vec{V} \cdot \vec{\phi}}{B_0} + \frac{\vec{V}_E}{N_0} \frac{\partial N_0}{\partial r} \right) = 0$$

$$= 0$$

\Rightarrow

$$\boxed{\frac{d}{dt} \left(\frac{c \vec{V} \cdot \vec{\phi}}{T} - \vec{A} \cdot \vec{V} \frac{c \vec{V} \cdot \vec{\phi}}{T} \right) + V_{\perp} \frac{d}{dr} \frac{c \vec{V} \cdot \vec{\phi}}{T} = 0}$$

$$\rightarrow H-M \text{ eqn. 6'}$$

\rightsquigarrow Can develop PV much further.
 See 216 (UCSD) notes, refs.

→ A key result: Taylor Identity

$$\begin{aligned}
 \langle v_y \nabla^2 \phi \rangle_x &= \langle \partial_x \phi (\partial_x^2 \phi + \partial_y^2 \phi) \rangle_x \\
 &= \langle \partial_x (\cancel{\partial_x(\partial_x \phi)^2}) \rangle_x \\
 &\quad + \langle \partial_x \phi \partial_y^2 \phi \rangle_x \\
 &= 0 + \partial_y \langle \partial_x \phi \partial_y \phi \rangle - \langle \partial_{xy} \phi \partial_y \phi \rangle \\
 &= \partial_y \langle \partial_x \phi \partial_y \phi \rangle_x - \langle \partial_x (\cancel{\partial_y \phi})^2 \rangle_x \\
 &\quad \uparrow \\
 &\quad R.S \\
 \rightarrow &= - \partial_y \langle \tilde{v}_y \tilde{v}_x \rangle
 \end{aligned}$$

Vorticity (+ PV) Flux = Reynolds Force.

N.B. Need 1 direction of symmetry, to eliminate odd moments.

→ Key result: Reynolds Force = vorticity Flux.

→ Links Reynolds force to polarization charge flux

Related:

Often useful to think of flow via vorticity.

Vorticity \leftrightarrow Polarity Charge.

Thus for zonal flow:

$$\vec{D} \cdot \vec{J} = 0$$

$$\nabla_{\perp} \cdot \vec{J}_{\perp} + \nabla_{\parallel} \vec{J}_{\parallel} + \vec{D}_{\perp} \cdot \vec{J}_{PS} = 0$$

Polaroidal symmetry / $m=0$:

$$\nabla_r \langle D_{r\perp} \rangle + \nabla_r \left\langle \frac{\tilde{B}_r}{\tilde{B}_\theta} \tilde{J}_{\theta\perp} \right\rangle = 0$$

Show from
GR

$$\Rightarrow \partial_t \langle D_r^2 \phi \rangle = - \partial_r \langle \tilde{V}_r D_r^2 \phi \rangle + \partial_r \langle \tilde{b}_r \tilde{J}_{\theta\parallel} \rangle$$

51

$$\partial_t \langle Q \rangle = - \partial_r \langle \tilde{V}_r \tilde{Q} \rangle + \partial_r \langle \tilde{b}_r \tilde{J}_{\theta\parallel} \rangle$$

\downarrow
total charge density

charge flux

inhomogeneous
field line tilting

$$\text{Dn } \langle \tilde{V}_r \tilde{J}_z \rangle \rightarrow \textcirclearrowleft \rightarrow \textcirclearrowleft \rightarrow \textcirclearrowleft \rightarrow$$

$$\text{Dn } \langle \tilde{B}_r \tilde{J}_{\parallel} \rangle \rightarrow$$


$$\text{and } \text{as } \bar{T}_w = -\bar{\sigma}^2 A_{\parallel\parallel}$$

$$\text{so: } \langle \tilde{B}_r \tilde{J}_{\parallel} \rangle = \text{Dn } \langle \tilde{B}_r \tilde{B}_{\phi} \rangle$$

Magnetic stress.
(show).

\Rightarrow

$$\partial_r \langle \varepsilon \rangle = -\bar{\sigma}^2 [\langle \tilde{V}_r \tilde{B}_{\phi} \rangle - \langle \tilde{B}_r \tilde{B}_{\phi} \rangle]$$

key point/summary for zonal structure generated by turbulence:

Zonal Flow/Jet results from:

1) inhomogeneous PV mixing
+

2) 1 direction of symmetry

N.B. { Symmetry need not be equilibrium symmetry
a priori }

→ Taylor Identity enables:

Krey Quasilinear Flow Theorem
→ Charney - DeZeeuw Thm.

Now, PV conservation \Rightarrow

$$\partial_t \tilde{z} + \nabla \cdot \nabla \tilde{z} = -\nabla^2 \tilde{z} - u \tilde{z}$$

↑ scale independent damping

Then, for mean:

$\underbrace{\langle \tilde{z} \rangle}_{\text{natural quantity of interest}}$

$$\partial_t \langle \tilde{z} \rangle = -\partial_r \langle \tilde{v}_y \tilde{z} \rangle + \nabla \partial_y^2 \langle \tilde{z} \rangle$$

$\langle \tilde{z} \rangle$ slow.

$-u \langle \tilde{z} \rangle$

↑
friction.

$\langle \tilde{v}_y \tilde{z} \rangle$ is PV Flux \rightarrow key quantity of interest.

→ Now, could proceed as QLT, with linear response \tilde{z} , etc.

on:

→ examining $\langle \tilde{\zeta}^2 \rangle \rightarrow$ potential enstrophy balance

$$\text{A.b. } -\frac{dE}{dt} = 0$$

not
enstroph.
 $\langle \tilde{\zeta}^2 \rangle$ conserved

Q.E.

- energy conserved

$$\partial_t \tilde{\zeta} + \Omega \cdot \nabla \tilde{\zeta} = - \tilde{U}_r \frac{d \langle \tilde{\zeta}^2 \rangle}{dr}$$

$$+ r \nabla^2 \tilde{\zeta} \quad \cancel{- u \tilde{\zeta}}$$

Q.E.D.

$$\partial_t \langle \tilde{\zeta}^2 \rangle + \cancel{\Omega \cdot \nabla} \langle \tilde{U}_r \tilde{\zeta}^2 \rangle = - \langle \tilde{U}_y \tilde{\zeta} \rangle \frac{d \langle \tilde{\zeta}^2 \rangle}{dy} - \cancel{r \langle (\partial \tilde{\zeta})^2 \rangle}$$

but have:

$$\partial_t \langle U_x \rangle + u \langle U_x \rangle = - \partial_y \langle \tilde{U}_y \tilde{U}_x \rangle$$

\uparrow
zonal flow

Now, Taylor Identity:

$$-\partial_y \langle \tilde{U}_y \tilde{U}_x \rangle = \langle \tilde{U}_y \Omega^2 \tilde{\phi} \rangle = \langle \tilde{U}_y \tilde{\zeta} \rangle$$

$$as \quad \bar{q} = (\bar{v}_y + D^2\phi) \\ \bar{\zeta} = D^2\bar{\phi}$$

so

$$\partial_t \langle v_x \rangle + u \langle v_x \rangle = \langle \bar{v}_y \bar{\zeta} \rangle$$

$$\Rightarrow \partial_t \left\langle \frac{\bar{\zeta}^2}{2} \right\rangle + D \cdot \left\langle \nabla \frac{\bar{\zeta}^2}{2} \right\rangle = - \left[- \partial_y \langle \bar{v}_y \bar{v}_x \rangle \right] \frac{d \langle \bar{v}_x \rangle}{dy}$$

$$= - \left[\partial_t \langle v_x \rangle + u \langle v_x \rangle \right] \frac{d \langle \bar{v}_x \rangle}{dy}$$

Then for $\partial \langle \bar{v}_x \rangle / \partial y$ slowly varying in space, time:

$$\left\{ \partial_t \left\{ \frac{\langle \bar{\zeta}^2 \rangle}{2} + \langle v_x \rangle \right\} + D_y \frac{\langle \nabla \bar{\zeta}^2 \rangle}{d \langle \bar{v}_x \rangle / dy} \right\} \\ = -u \langle v_x \rangle$$

→ Charney - Derazin Thm.

→ Conservation of balance and Taylor identity.

→ Momentum Thm ($\sim Q \text{Lt}$) for Rossby - Zonal interaction.

what does it mean?

$$\text{① } \frac{\partial}{\partial y} \left\{ \frac{\langle \tilde{\beta}^2 \rangle}{2 d\langle \tilde{\epsilon} \rangle / dy} + \langle u_x \rangle \right\} + \partial_y \frac{\langle \tilde{v}_y \tilde{\beta}^2 / 2 \rangle}{d\langle \tilde{\epsilon} \rangle / dy}$$

$$\text{④} = -u \langle u_x \rangle$$

$$\text{① } \frac{\langle \tilde{\beta}^2 \rangle}{2 d\langle \tilde{\epsilon} \rangle / dy} ?$$

$$(P_0 = I)$$

$$Q = \nabla^2 \phi + \beta y$$

$$\tilde{Q} = \nabla^2 \tilde{\phi}$$

$$\text{①} = \sum_k \frac{(k_x^2)^2 |\phi_{kl}|^2}{2 d\langle \tilde{\epsilon} \rangle / dy}$$

$$= - \sum_k - \frac{k_x^2 (k_x l)}{2 \frac{k_x d\langle \tilde{\epsilon} \rangle / dy}{k_x^2}} k_x$$

wave action density

$$\textcircled{1} = - \sum_k k_x N_k$$

\rightarrow Quasi-particle density

$$\omega = - \frac{k_x \sqrt{\rho}}{k^2}$$

$$N_k = \frac{E_{kx}}{\omega_k}$$

Wave energy density

Wave action density

Thus

$$\textcircled{1} = - WMD \equiv - \rho \text{ pseudomomentum (PM)} \\ A \qquad \rightarrow \text{wave action density}$$

$$\lim_{\rho \rightarrow \infty} \text{Pseudomomentum} = WMD$$

and:

$$\textcircled{3} = \left\langle \tilde{v}_y \frac{\tilde{g}^2}{2 \frac{d \tilde{v}_x}{dx}} \right\rangle$$

$$= + \left\langle \tilde{v}_x (-\tilde{p}/\tilde{M}) \right\rangle$$

\Rightarrow obviously
related to
turbulence
spreading \rightarrow
beyond QLT.

$$\partial_t \left\{ \langle v_x \rangle - k_x N \right\} - \partial_y \langle \tilde{\mathcal{D}}_y (\tilde{W}_{MD}) \rangle \\ = - u \langle v_x \rangle$$

So:

\rightarrow up to ^{drag} spreading, flow locked to wave momentum density, i.e.

$$\partial_t \langle v_x \rangle = \partial_t \left(\sum_n k_x N_n \right)$$

charge wave amplitude \leftrightarrow charge in fluid

\rightarrow no-slip theorem for fluid relative to inter-penetrating fluid of quasi-particles,

\rightarrow spreading term \rightarrow linked to fluctuation envelope scale \rightarrow

can break local, but not global no-slip. Phase?

$H W \rightarrow$ generalize for $D, r = 0$
 Hasagawa - Wakatani system

Theoretical Observations:

- ① → theorem is general, with no truncation, though leaves enstrophy flux $\langle \nabla \vec{g}^2 \rangle$ un-calculated
 - ② → no linear response closure like QL used
 - ③ → WMD defined beyond level of small amplitude perturbation
- ① → ③ enabled by { PV conservation \rightarrow
 enstrophy balance
 Taylor identity }

contrast to:

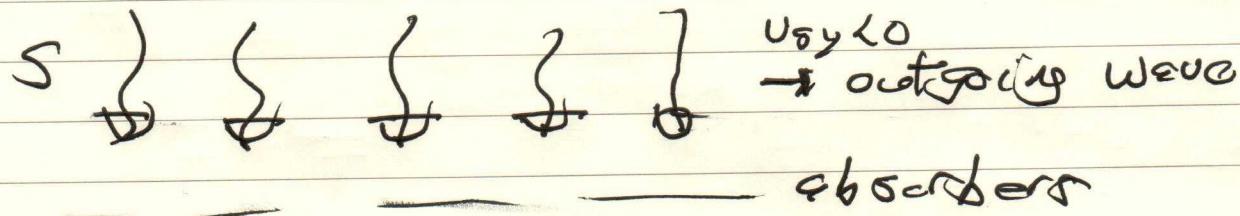
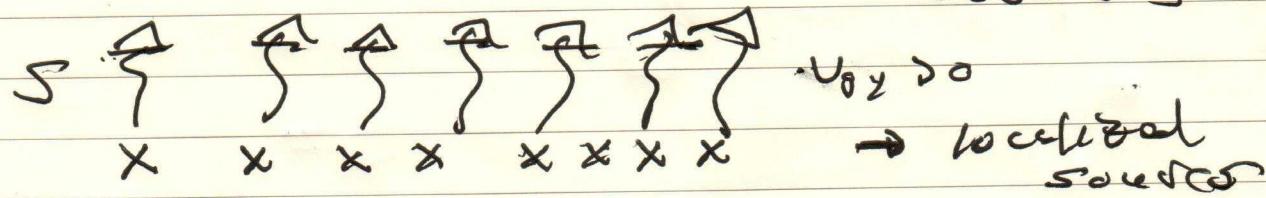
$$2\zeta \{ R P M D + W M D \} = 0$$

Aside \rightarrow Physical Picture.

889

Date

\rightarrow ZF's and "Negative Viscosity" absorbers



Circularity \Rightarrow selects sign v_{0y} for outgoing wave

$$\omega = -\frac{\beta k_x}{k^2}$$

n.b. irreversibility
 \Rightarrow outgoing wave

$$v_{0y} = \frac{2\beta [k_x k_y]}{(k^2)^2}$$

\rightarrow prop velocity
sgn ($\beta k_x k_y$) \rightarrow
sgn v_{0y}

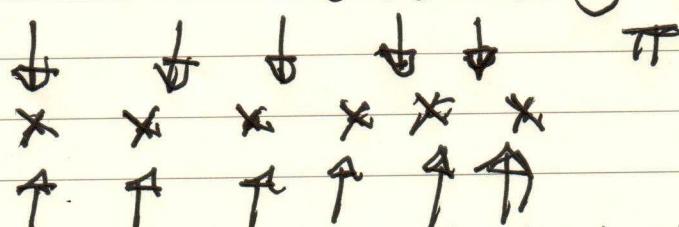
but

$$\langle v_y v_x \rangle = [-] \sum_h [k_x k_y] |k_{\perp h}|^2$$

energy flux

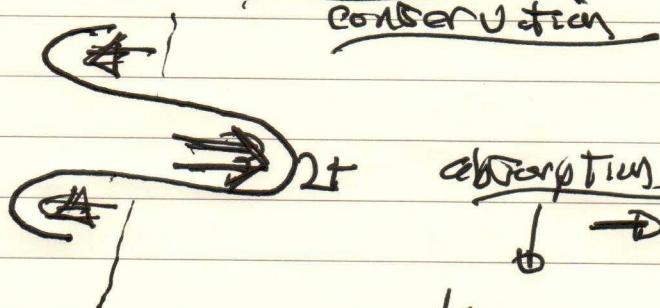
$\beta > 0 \rightarrow$ outgoing ($v_{0y} > 0$) $\wedge \Rightarrow$

incoming momentum flux.



10

Momentum
conservation



at absorption



builds till

flow modifier

dispersion

sufficiently

on

by
change

momentum
transported up
gradient \rightarrow not
a simple Ficks
Law.

Prob-
Abey

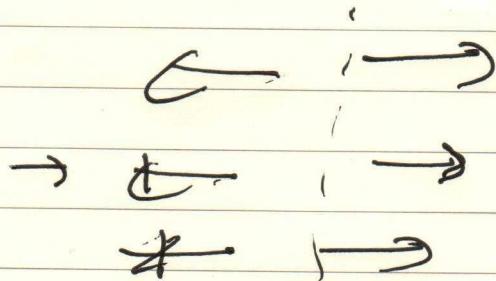
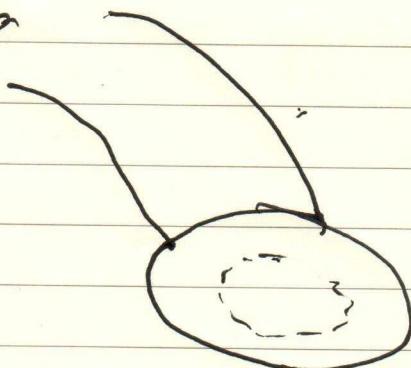
feed blocks on
source
(instability)

"...the central result that a rapidly rotating flow, when stirred in a localized region will convert angular momentum into that region."

- Isaac Held.

\rightarrow applies to Plasma with oblique transmutation

i.e.



etc. $\Rightarrow V_{gr}$

$\Rightarrow V'_E$ modif' V_{gr}

→ Whole we're here :-

Drift Wave - Zonal Flow Interaction:

Modulational Instability and the Predator-Prey Model

- Have established connection between zonal flow and pseudo-momentum / WMD.
 - Of course, coupling of WMD + flow must extract energy from turbulence \Rightarrow feedback loop for flow on turbulence.
 - Structure closely analogous to theory of weak Langmuir turbulence
- | | | |
|----------------------|-------------------|--|
| Plasma Wave | \leftrightarrow | Drift Wave |
| \updownarrow | | \updownarrow |
| Ion Acoustic
Mode | | Zonal Flow
$(\rightarrow \text{frequency})$ |
| (Finite frequency) | | |

Recall:

Laguerre wave envelope

$$i\omega \partial_t \Sigma = \frac{c_p^2}{2} \frac{\partial N}{N_0} \Sigma - \propto V_{\text{The}}^2 \nabla^2 \Sigma$$

↑
 refraction
 in density
 perturbation

↑
 diffraction

$$\nabla^2 \frac{\partial N}{N_0} - c_s^2 V^2 \frac{\partial N}{N_0} = \nabla^2 \left(\frac{i \sigma l^2}{8 \pi n m c} \right)$$

↑
 radiation
 pressure

treating plasma waves
 at level of eikonal theory
 \Rightarrow rays, can write:

$$\frac{\partial N}{\partial t} + k_p \cdot \nabla N - \frac{\partial(\omega)}{\partial x} \cdot \frac{\partial N}{\partial k} = 0$$

i.e.

$$\frac{dN}{dt} = 0$$

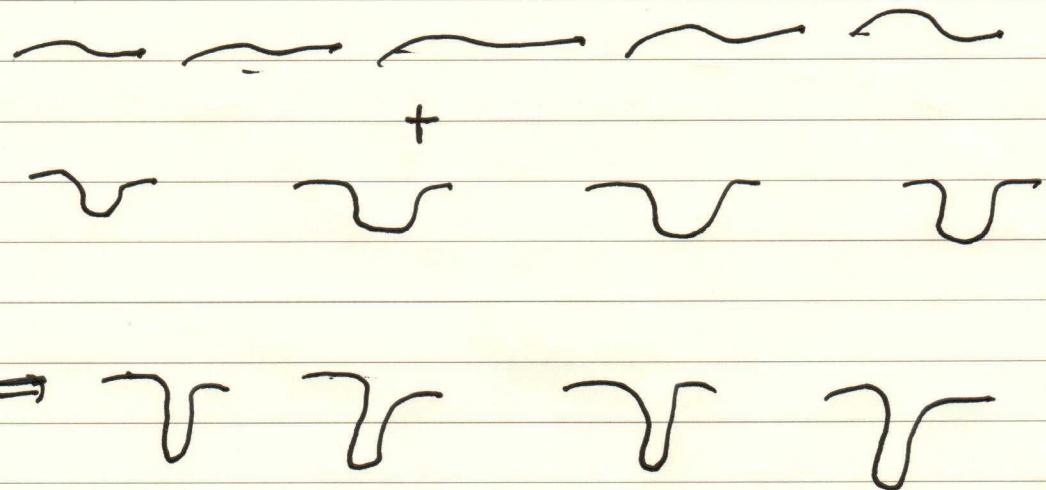
\rightarrow Statement of
 conservation of
 wave action density

Conservation Action \Rightarrow Phase Symmetry

$$N = \sum \omega_n \Rightarrow \text{population density in } x, k$$

\rightarrow phase space density

Modulation:



i.e. plasma wave goes unstable to
~~non~~ reflecting density perturbations.

→ perturbations grow

→ plasma wave ensemble depleted.

To calculate:

- I A w growth rate
 → field energy density

$$\frac{|\vec{E}|^2}{8\pi} = \left(\frac{1}{2}\right) \omega_n^2$$

$$= \int \left(\frac{1}{2}\right) \omega_n^2 N_n \, dk \quad \text{modulated wave}$$

then

$$\frac{\partial^2}{\partial t^2} \frac{\partial n}{n_0} - C_s^2 \nabla^2 \frac{\partial n}{n} = D^2 \left(\frac{1}{M_e} \int dk \omega_n^2 \tilde{N}_A \right)$$

Date

Where:

 $\frac{k}{w_n} \rightarrow$ Plasma wave

 \tilde{N}_{ξ_2}
 $\frac{2}{\Sigma} \rightarrow$ Modulation field

modulated WAD

Straightforward linearization:

$$\tilde{N}_{\xi_2} = -\frac{2}{\Omega - 2 \cdot \underline{v}_{gn}} \frac{\pi}{2n_0} \xi_2 \cdot \frac{\partial \ln N}{\partial \underline{v}}$$

where

$$\frac{\underline{v}}{\Omega - 2 \cdot \underline{v}_{gn}} = \frac{P}{\Omega - \underline{v}_{gn}} - i\pi C(\Omega - \underline{v}_{gn})$$

Q

$$\vec{\nabla}_x \cdot \vec{B}_n - c_s^2 \nabla^2 \frac{\vec{B}_n}{n_0} = \nabla^2 \left(\frac{1}{m_e} \int dk \omega_n + \right.$$

$$\left. - \frac{\sum w_{p0}}{\Omega - \sum V_p n^{2/3}} \frac{\vec{B}_n}{n_0} \cdot \frac{\partial \langle N \rangle}{\partial \vec{k}} \right)$$

#

$$\Omega^2 = q^2 c_s^2$$

$$+ \frac{q^2 w_{p0}^2}{2 m_e} \int dk \text{ if } (\Omega - \sum V_p) \frac{\vec{B}_n}{n_0} \cdot \frac{\partial \langle N \rangle}{\partial \vec{k}}$$

instability of IAW if

$$\frac{\partial \langle N \rangle}{\partial \vec{k}} \cdot \vec{B}_n > 0 \quad \text{at:}$$

$$\Omega = \sum V_p$$

i.e. population inversion.

→ energy decays from plasma waves
to IAW.

→ $\Omega_{\text{real}} \approx c_s$ as well.

What of Plasma Waves?

$\langle N \rangle$ evolves!

$$\frac{\partial \langle N \rangle}{\partial t} = \frac{\partial}{\partial \underline{N}} \cdot \left\langle \frac{\partial \tilde{\omega}}{\partial \underline{x}} \hat{N} \right\rangle$$


 -> An approximation of
in PL

⇒

$$\frac{\partial \langle N \rangle}{\partial t} = \frac{\partial}{\partial \underline{N}} \cdot \underline{D}_n \cdot \frac{\partial \langle N \rangle}{\partial \underline{N}}$$

- Quasi-particle diffusion equation

- "Induced Diffusion"

For energy

$$E = \int d\underline{N} \tilde{\omega} \langle N \rangle$$

$$\frac{d'E}{dt} = - \int \frac{\partial \tilde{\omega}}{\partial \underline{N}} \cdot \underline{D}_n \cdot \frac{\partial \langle N \rangle}{\partial \underline{N}}$$


 sign v_{gr} enters!

Sign $\frac{\partial \omega}{\partial k}$ $\frac{\partial \omega_{NS}}{\partial k}$ sets ref

gain / loss energy from plasmons.

$$\boxed{D_k = e \left(\frac{\omega_p^2}{2m} \right) \sum_{q, \Omega} |\tilde{U}_{kq, \Omega}|^2 \frac{c}{\Omega - \epsilon_q v_g}}$$

- physics is stochastic refraction

- requires: Ray chaos

i.e. overlap of phase speed \leftrightarrow
group speed resonances ($Q-P$)

- energetic Plasmon $Q-P$ energy
vs IAW energy.

- structure var const.

For plasma wave $\frac{\partial \omega}{\partial k} > 0$, so

$\frac{\partial \omega}{\partial k} >$ for decay.

- how related to wave interaction,

$$\underline{\omega}_{\text{sum}} = \underline{\omega}_{\text{H}'} - \underline{\omega}_{\text{y}}$$

$$|\underline{\omega}| < |\underline{\omega}'|$$

$$= \frac{i}{\underline{\omega}_{\text{H}'} + \underline{n} \cdot \frac{\partial \omega}{\partial \underline{\omega}} - \underline{\omega}_{\text{H}'} - \underline{\omega}_{\text{y}}}$$

$$= \frac{i}{\underline{n} \cdot v_{\text{g}}(\underline{\omega}') - \underline{\omega}_{\text{y}}} \quad \checkmark$$

To Zonal Flow + DW !

- key points:

$$\textcircled{1} \quad \frac{\partial \omega}{\partial \underline{n}} < 0 \quad \text{so} \quad \frac{\partial N}{\partial \underline{n}} < 0 \quad \text{for depletion DW.}$$

$$\textcircled{2} \quad N = \frac{(1 + k_z^2 \beta^2) |\phi_{\text{H}'}|^2}{k_0 V_* / (1 + k_z^2 \beta^2)} = \frac{\Sigma}{\omega_{\text{y}}}$$

$$N = \frac{\Sigma}{k_0 V_*}$$

but $\frac{dy}{dt} = -\frac{\partial}{\partial x} (\underline{k}_x \cdot \underline{v} + \psi)$

$$= -\frac{\partial}{\partial x} (h_0 V_0(r))$$

$$\rightarrow \frac{dh_0}{dt} = -\frac{\partial}{\partial r} (h_0 V_0(r))$$

i.e. h_0 constant

Potential

Action Conservation \leftrightarrow Enstrophy

Conservation in Shearing Field