

8.323 Problem Set 2 Solutions

February 21, 2023

Question 1: A Problem with Relativistic Quantum Mechanics (20 points)

The Schrödinger equation for a free non-relativistic particle is:

$$i\partial_t \psi(\vec{x}, t) = -\frac{1}{2m} \nabla^2 \psi(\vec{x}, t)$$

The generalization of the above equation to a free relativistic particle is the so-called Klein-Gordon equation

$$\partial_t^2 \psi(\vec{x}, t) - \nabla^2 \psi(\vec{x}, t) + m^2 \psi(\vec{x}, t) = 0$$

We emphasize that in both these equations, $\psi(\vec{x}, t)$ is interpreted as a wave function for the dynamical variable $\vec{x}(t)$, rather than a dynamical field.

(a) As a reminder, derive from the Schrödinger equation the continuity equation for the probability

$$\partial_t \rho + \nabla \cdot \vec{J} = 0$$

where

$$\rho = |\psi|^2, \quad \vec{J} = -\frac{i}{2m} (\psi^* \nabla \psi - \psi \nabla \psi^*)$$

We compute:

$$\begin{aligned} \partial_t \rho &= \psi \partial_t \psi^* + \psi^* \partial_t \psi = -\frac{i}{2m} \left(\psi \vec{\nabla}^2 \psi^* - \psi^* \vec{\nabla}^2 \psi \right) \\ &= \frac{i}{2m} \vec{\nabla} \cdot \left(\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^* \right) = -\vec{\nabla} \cdot \vec{J} \end{aligned}$$

where we use the Schrödinger equation in the second equality.

(b) Suppose $\psi(\vec{x}, t)$ has the plane wave form, i.e.

$$\psi(\vec{x}, t) \propto e^{i\vec{k} \cdot \vec{x}}$$

for some real vector \vec{k} . Find the solutions to the Klein-Gordon equation above.

We substitute the ansatz $\psi(\vec{x}, t) = e^{i\vec{k} \cdot \vec{x}} \phi(t)$ into the Klein-Gordon equation to get an equation for $\phi(t)$:

$$\partial_t^2 \phi + (\vec{k}^2 + m^2) \phi = 0$$

This has plane-wave solutions of positive and negative frequencies,

$$\phi(t) = A e^{-i\omega_{\vec{k}} t} + B e^{i\omega_{\vec{k}} t}, \quad \omega_{\vec{k}} = \sqrt{\vec{k}^2 + m^2}$$

Hence, the Klein-Gordon equation has solutions

$$\psi(\vec{x}, t) = A e^{i(-\omega_{\mathbf{k}} t + \mathbf{k} \cdot \mathbf{x})} + B e^{i(\omega_{\mathbf{k}} t + \mathbf{k} \cdot \mathbf{x})}$$

(c) Show that the Klein-Gordon equation also leads to a continuity equation, with ρ and \vec{J} now given by

$$\rho = \frac{i}{2m} (\psi^* \partial_t \psi - \psi \partial_t \psi^*). \quad \vec{J} = -\frac{i}{2m} (\psi^* \nabla \psi - \psi \nabla \psi^*)$$

In the same way as in part (a), we compute:

$$\partial_t \rho = \frac{i}{2m} (\psi^* \partial_t^2 \psi - \psi \partial_t^2 \psi^*) = \frac{i}{2m} (\psi^* \nabla^2 \psi - \psi \nabla^2 \psi^*) = -\vec{\nabla} \cdot \vec{J}$$

where we use the Klein-Gordon equation in the second equality.

(d) Argue that this ρ cannot be interpreted as a probability density.

We write

$$\rho = \frac{i}{2m} (\psi^* \partial_t \psi - \psi \partial_t \psi^*) = \frac{1}{m} \text{Im}(\psi \partial_t \psi^*)$$

Any proper probability density must be positive definite, i.e. $\rho \geq 0$. This is not the case here. For instance, for the plane wave solution $A = 0$, B from (b) we compute

$$\rho = \text{Im} \left(B e^{i\omega_{\mathbf{k}} t + i\mathbf{k} \cdot \mathbf{x}} (-i\omega_{\mathbf{k}}) B e^{-i\omega_{\mathbf{k}} t - i\mathbf{k} \cdot \mathbf{x}} \right) = -B^2 \omega_{\mathbf{k}} < 0$$

Since $\omega_{\mathbf{k}} = \sqrt{\mathbf{k}^2 + m^2} > 0$, the existence of negative-frequency solutions means that ρ cannot be positive definite, and cannot be interpreted as a probability density.

Question 2: Commutation relations of creation and annihilation operators (20 points)

For the real scalar field theory discussed in lecture,

$$\mathcal{L} = -\frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{1}{2}m^2\phi^2$$

we showed that the time-evolution of the quantum operator $\phi(\mathbf{x}, t)$ is given by

$$\phi(\mathbf{x}, t) = \int d^3k \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} (a_{\mathbf{k}} u_{\mathbf{k}}(\mathbf{x}, t) + a_{\mathbf{k}}^\dagger u_{\mathbf{k}}^*(\mathbf{x}, t))$$

where

$$\omega_{\mathbf{k}} = \sqrt{\mathbf{k}^2 + m^2}, \quad u_{\mathbf{k}} = e^{-i\omega_{\mathbf{k}}t + i\mathbf{k}\cdot\mathbf{x}}$$

We use $\pi(\mathbf{x}, t)$ to denote the momentum density conjugate to ϕ . The canonical commutation relations among ϕ and π are

$$[\phi(\mathbf{x}, t), \phi(\mathbf{x}', t)] = [\pi(\mathbf{x}, t), \pi(\mathbf{x}', t)] = 0, \quad [\phi(\mathbf{x}, t), \pi(\mathbf{x}', t)] = i\delta^{(3)}(\mathbf{x} - \mathbf{x}')$$

- (a) Show that it is enough to impose the canonical commutation relations at $t = 0$. That is, once we impose them at $t = 0$, then the relations at general t are automatically satisfied.

Note: this statement in fact applies not only to $V(\phi) = \frac{1}{2}m^2\phi^2$, but any potential $V(\phi)$.

In the Heisenberg picture we have:

$$[A(\mathbf{x}, t), B(\mathbf{x}', t)] = [e^{iHt} A(\mathbf{x}, 0) e^{-iHt}, e^{iHt} B(\mathbf{x}', 0) e^{-iHt}] = e^{iHt} [A(\mathbf{x}, 0), B(\mathbf{x}', 0)] e^{-iHt}$$

Now let us impose the canonical commutation relations at $t = 0$. Then, it follows that

$$\begin{aligned} [\phi(\mathbf{x}, t), \phi(\mathbf{x}', t)] &= [\pi(\mathbf{x}, t), \pi(\mathbf{x}', t)] = e^{iHt} 0 e^{-iHt} = 0 \\ [\phi(\mathbf{x}, t), \pi(\mathbf{x}', t)] &= e^{iHt} i\delta^{(3)}(\mathbf{x} - \mathbf{x}') e^{-iHt} = i\delta^{(3)}(\mathbf{x} - \mathbf{x}') \end{aligned}$$

These are again precisely the canonical commutation relations, now at generic t .

- (b) Express $a_{\mathbf{k}}$ and $a_{\mathbf{k}}^\dagger$ in terms of $\phi(\mathbf{k})$ and $\pi(\mathbf{k})$, where $\phi(\mathbf{k})$ and $\pi(\mathbf{k})$ are Fourier transforms of $\phi(\mathbf{x}, t = 0)$ and $\pi(\mathbf{x}, t = 0)$, e.g.

$$\phi(\mathbf{k}) = \int d^3x e^{-i\mathbf{k}\cdot\mathbf{x}} \phi(\mathbf{x}, t = 0)$$

We start with the mode expansions for $\phi(\mathbf{x}, t)$ and $\pi(\mathbf{x}, t)$

$$\begin{aligned} \phi(\mathbf{x}, t) &= \int d^3k \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} (a_{\mathbf{k}} e^{-i\omega_{\mathbf{k}}t + i\mathbf{k}\cdot\mathbf{x}} + a_{\mathbf{k}}^\dagger e^{i\omega_{\mathbf{k}}t - i\mathbf{k}\cdot\mathbf{x}}) \\ \pi(\mathbf{x}, t) &= -i \int d^3k \sqrt{\frac{\omega_{\mathbf{k}}}{2}} (a_{\mathbf{k}} e^{-i\omega_{\mathbf{k}}t + i\mathbf{k}\cdot\mathbf{x}} - a_{\mathbf{k}}^\dagger e^{i\omega_{\mathbf{k}}t - i\mathbf{k}\cdot\mathbf{x}}) \end{aligned}$$

This is almost of the form of a Fourier transform, and by changing variables of one of the terms from $\mathbf{k} \rightarrow -\mathbf{k}$ we have

$$\begin{aligned} \phi(\mathbf{k}, t) &= \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} (a_{\mathbf{k}} e^{-i\omega_{\mathbf{k}}t} + a_{-\mathbf{k}}^\dagger e^{i\omega_{\mathbf{k}}t}) \\ \pi(\mathbf{k}, t) &= -i \sqrt{\frac{\omega_{\mathbf{k}}}{2}} (a_{\mathbf{k}} e^{-i\omega_{\mathbf{k}}t} - a_{-\mathbf{k}}^\dagger e^{i\omega_{\mathbf{k}}t}) \end{aligned}$$

Now, observe that the equations are decoupled in \mathbf{k} . We can take $t = 0$ and solve this as a regular system of equations for $a_{\mathbf{k}}$ and $a_{\mathbf{k}}^\dagger$.

$$\begin{aligned} a_{\mathbf{k}} &= \sqrt{\frac{\omega_{\mathbf{k}}}{2}} \phi(\mathbf{k}) + i \sqrt{\frac{1}{2\omega_{\mathbf{k}}}} \pi(\mathbf{k}) \\ a_{\mathbf{k}}^\dagger &= \sqrt{\frac{\omega_{\mathbf{k}}}{2}} \phi(-\mathbf{k}) - i \sqrt{\frac{1}{2\omega_{\mathbf{k}}}} \pi(-\mathbf{k}) \end{aligned} \quad (1)$$

(c) Using the expressions derived in part (b), deduce the commutation relations

$$[a_{\mathbf{k}}, a_{\mathbf{k}'}], \quad [a_{\mathbf{k}}^\dagger, a_{\mathbf{k}'}^\dagger], \quad [a_{\mathbf{k}}, a_{\mathbf{k}'}^\dagger]$$

from the commutation relations above at $t = 0$.

It is useful to take the Fourier transform \mathcal{F} (from position to momentum space) of the $t = 0$ canonical commutation relations:

$$\begin{aligned} [\phi(\mathbf{k}), \phi(\mathbf{k}')] &= [\pi(\mathbf{k}), \pi(\mathbf{k}')] = \mathcal{F}_{\mathbf{x} \rightarrow \mathbf{k}} \circ \mathcal{F}_{\mathbf{x}' \rightarrow \mathbf{k}'}(0) = 0 \\ [\phi(\mathbf{k}), \pi(\mathbf{k}')] &= \mathcal{F}_{\mathbf{x} \rightarrow \mathbf{k}} \circ \mathcal{F}_{\mathbf{x}' \rightarrow \mathbf{k}'}(i\delta^{(3)}(\mathbf{x} - \mathbf{x}')) = i \int d^3 \mathbf{x} d^3 \mathbf{x}' e^{-i\mathbf{k} \cdot \mathbf{x}} e^{-i\mathbf{k}' \cdot \mathbf{x}'} \delta^{(3)}(\mathbf{x} - \mathbf{x}') \\ &= i \int d^3 \mathbf{x} e^{-i(\mathbf{k} + \mathbf{k}') \cdot \mathbf{x}} = i(2\pi)^3 \delta^{(3)}(\mathbf{k} + \mathbf{k}') \end{aligned}$$

Now we compute commutators of creation and annihilation operators using the results in (b)

$$\begin{aligned} [a_{\mathbf{k}}, a_{\mathbf{k}'}] &= +\frac{i}{2} ([\phi(\mathbf{k}), \pi(\mathbf{k}')] + [\pi(\mathbf{k}), \phi(\mathbf{k}')]) = -\frac{1}{2}(2\pi)^3 (\delta^{(3)}(\mathbf{k} + \mathbf{k}') - \delta^{(3)}(\mathbf{k}' + \mathbf{k})) = 0 \\ [a_{\mathbf{k}}^\dagger, a_{\mathbf{k}'}^\dagger] &= -\frac{i}{2} ([\phi(-\mathbf{k}), \pi(-\mathbf{k}')] + [\pi(-\mathbf{k}), \phi(-\mathbf{k}')]) = -\frac{1}{2}(2\pi)^3 (\delta^{(3)}(-\mathbf{k} - \mathbf{k}') - \delta^{(3)}(-\mathbf{k}' - \mathbf{k})) = 0 \\ [a_{\mathbf{k}}, a_{\mathbf{k}'}^\dagger] &= +\frac{i}{2} (-[\phi(\mathbf{k}), \pi(-\mathbf{k}')] + [\pi(\mathbf{k}), \phi(-\mathbf{k}')]) = -\frac{1}{2}(2\pi)^3 (-\delta^{(3)}(\mathbf{k} - \mathbf{k}') - \delta^{(3)}(\mathbf{k}' - \mathbf{k})) \\ &= (2\pi)^3 \delta^{(3)}(\mathbf{k} - \mathbf{k}') \end{aligned} \quad (2)$$

Question 3: Noether charges in terms of creation and annihilation operators (20 points)

In problem set 1, we obtained the conserved charges associated with spacetime translational symmetries for a complex scalar field theory. The results there can be easily converted to the corresponding expressions for a real scalar field theory.

$$\mathcal{L} = -\frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{1}{2}m^2\phi^2$$

(a) Express the Hamiltonian H of this theory in terms of $a_{\mathbf{k}}$ and $a_{\mathbf{k}}^\dagger$.

From problem set 1, we quote

$$H = \frac{1}{2} \int d^3x (\pi^2 + (\nabla\phi)^2 + m^2\phi^2)$$

It is convenient to first convert this expression into momentum space, before using the decomposition into creation and annihilation operators. We use the identity:

$$\begin{aligned} \int d^3\mathbf{x} f(\mathbf{x})g(\mathbf{x}) &= \int d^3\mathbf{x} d^3\mathbf{k} d^3\mathbf{k}' e^{i(\mathbf{k}+\mathbf{k}')\cdot\mathbf{x}} f(\mathbf{k})g(\mathbf{k}') \\ &= \frac{1}{(2\pi)^3} \int d^3\mathbf{k} d^3\mathbf{k}' \delta^{(3)}(\mathbf{k} + \mathbf{k}') f(\mathbf{k})g(\mathbf{k}') = \int d^3\mathbf{k} f(\mathbf{k})g(-\mathbf{k}) \end{aligned} \quad (3)$$

More generally if there are derivatives acting on f or g , each derivative acting on f drags down a factor of $i\mathbf{k}$, while each derivative acting on g drags down a factor of $i\mathbf{k}'$, which becomes $-i\mathbf{k}$ after performing the $d^3\mathbf{x}$ integral. We further use the shorthand $\mathcal{d}x = dx/2\pi$ (\mathcal{d} is to d as \hbar is to h).

Hence, we can now write

$$\begin{aligned} H &= \frac{1}{2} \int \mathcal{d}^3\mathbf{k} (\pi(\mathbf{k}, t)\pi(-\mathbf{k}, t) + (\mathbf{k}^2 + m^2)\phi(\mathbf{k}, t)\phi(-\mathbf{k}, t)) \\ &= \frac{1}{2} \int \mathcal{d}^3\mathbf{k} (\pi(\mathbf{k}, t)\pi(-\mathbf{k}, t) + \omega_{\mathbf{k}}^2\phi(\mathbf{k}, t)\phi(-\mathbf{k}, t)) \\ &= \frac{1}{2} \int \mathcal{d}^3\mathbf{k} \frac{\omega_{\mathbf{k}}}{2} (a_{\mathbf{k}}(t)a_{\mathbf{k}}(t)^\dagger + a_{\mathbf{k}}(t)^\dagger a_{\mathbf{k}}(t)) = \int \mathcal{d}^3\mathbf{k} \omega_{\mathbf{k}} a_{\mathbf{k}}^\dagger a_{\mathbf{k}} + \frac{1}{2}(2\pi)^3\delta(0) \int \mathcal{d}^3\mathbf{k} \omega_{\mathbf{k}} \end{aligned}$$

In the third equality we use the relations (1) from problem 2(b). In the last equality we use the commutator (2) from problem 2(c), as well as the expressions $a_{\mathbf{k}}(t) = e^{-i\omega_{\mathbf{k}}t}a_{\mathbf{k}}$ and $a_{\mathbf{k}}^\dagger(t) = e^{i\omega_{\mathbf{k}}t}a_{\mathbf{k}}^\dagger$. Note that in the above calculation, we showed that the time-dependence cancels explicitly. We could have also used that H is conserved to remove the time-dependence immediately by evaluating all fields at $t = 0$.

This can be written as

$$H = \int \mathcal{d}^3\mathbf{k} \omega_{\mathbf{k}} N_{\mathbf{k}} + E_0$$

for the number operator $N_{\mathbf{k}} = a_{\mathbf{k}}^\dagger a_{\mathbf{k}}$, and zero-point energy $E_0 = \frac{1}{2}(2\pi)^3\delta(0) \int \mathcal{d}^3\mathbf{k} \omega_{\mathbf{k}}$.

(b) Express the conserved charges P^i for spatial translations, in terms of $a_{\mathbf{k}}$ and $a_{\mathbf{k}}^\dagger$.

Again we quote the charges from problem set 1, and use (3) to write it in momentum space.

$$\begin{aligned} P^i &= \int d^3\mathbf{x} \pi \partial^i \phi = -i \int d^3\mathbf{k} \pi(\mathbf{k}, t) \phi(-\mathbf{k}, t) k^i \\ &= \frac{1}{2} \int \mathcal{d}^3\mathbf{k} k^i (a_{\mathbf{k}}(t) - a_{-\mathbf{k}}(t)^\dagger) (a_{-\mathbf{k}}(t) + a_{\mathbf{k}}(t)^\dagger) \\ &= \frac{1}{2} \int \mathcal{d}^3\mathbf{k} k^i (a_{\mathbf{k}} a_{-\mathbf{k}} e^{-2i\omega_{\mathbf{k}}t} + a_{\mathbf{k}} a_{\mathbf{k}}^\dagger - a_{-\mathbf{k}}^\dagger a_{-\mathbf{k}} - a_{-\mathbf{k}}^\dagger a_{\mathbf{k}}^\dagger e^{2i\omega_{\mathbf{k}}t}) \end{aligned}$$

In the third equality we use the relations (1) from problem 2(b). Observe that due to the k^i factor and the commutation relations (2), the first and fourth terms in our final expression are odd under the change of variables $\mathbf{k} \rightarrow -\mathbf{k}$, so they must vanish. Therefore,

$$P^i = \frac{1}{2} \int d^3\mathbf{k} k^i \left(a_{\mathbf{k}} a_{\mathbf{k}}^\dagger + a_{\mathbf{k}}^\dagger a_{\mathbf{k}} \right) = \int d^3\mathbf{k} k^i a_{\mathbf{k}}^\dagger a_{\mathbf{k}} + \frac{1}{2} \int d^3\mathbf{k} (2\pi)^3 \delta^{(3)}(\mathbf{0}) k^i = \int d^3\mathbf{k} k^i N_{\mathbf{k}}$$

In the second equality we use the commutator (2), and in the third equality we note that the last term vanishes because the integrand is \mathbf{k} -odd.

We may combine this with the expression for H in part (a) to write

$$P^\mu = \int d^3\mathbf{k} k^\mu N_{\mathbf{k}} + \delta^{\mu 0} E_0, \quad k^0 = \omega_{\mathbf{k}} \quad (4)$$

(c) Starting with

$$\phi(0, 0) = \int d^3k \frac{1}{\sqrt{2\omega_{\mathbf{k}}}} \left(a_{\mathbf{k}} + a_{\mathbf{k}}^\dagger \right)$$

show that under the action of translation operators,

$$\phi(\mathbf{x}, t) = e^{iHt - iP^i x^i} \phi(0, 0) e^{-iHt + iP^i x^i}$$

Hint: this problem becomes trivial using the following formula for a harmonic oscillator,

$$e^{i\alpha N} a e^{-i\alpha N} = e^{-i\alpha} a, \quad N = a^\dagger a$$

The formula in the hint follows from the Baker-Campbell-Hausdorff (BCH) formula,

$$e^X Y e^{-X} = Y + [X, Y] + \frac{1}{2!} [X, [X, Y]] + \dots$$

We check:

$$\begin{aligned} [i\alpha N, a] &= i\alpha [a^\dagger a, a] = -i\alpha a \\ e^{i\alpha N} a e^{-i\alpha N} &= a + (-i\alpha) a + \frac{1}{2!} (-i\alpha)^2 a + \dots = e^{-i\alpha} a \end{aligned}$$

and similarly, $e^{-i\alpha N} a^\dagger e^{i\alpha N} = e^{i\alpha} a^\dagger$.

Now we generalize. For $\alpha(\mathbf{k}')$ a real-valued function,

$$\begin{aligned} e^{i \int d^3\mathbf{k}' \alpha(\mathbf{k}') N_{\mathbf{k}'} } a_{\mathbf{k}} e^{-i \int d^3\mathbf{k}' \alpha(\mathbf{k}') N_{\mathbf{k}'} } &= e^{i \int d^3\mathbf{k}' \delta(\mathbf{k}-\mathbf{k}') \alpha(\mathbf{k}') N_{\mathbf{k}'} } a_{\mathbf{k}} e^{-i \int d^3\mathbf{k}' \delta(\mathbf{k}-\mathbf{k}') \alpha(\mathbf{k}') N_{\mathbf{k}'} } \\ &= e^{i\alpha(\mathbf{k}) N_{\mathbf{k}}} a_{\mathbf{k}} e^{-i\alpha(\mathbf{k}) N_{\mathbf{k}}} = e^{-i\alpha(\mathbf{k})} a_{\mathbf{k}} \end{aligned} \quad (5)$$

In the first equality, we use that the 2 sets of operators $\{\alpha_{\mathbf{k}}, \alpha_{\mathbf{k}}^\dagger, N_{\mathbf{k}}\}$ and $\{\alpha_{\mathbf{k}'}, \alpha_{\mathbf{k}'}^\dagger, N_{\mathbf{k}'}\}$ commute with each other for $\mathbf{k} \neq \mathbf{k}'$. This allows us to move all but the $\mathbf{k}' = \mathbf{k}$ exponentials on the left-hand side past the $a_{\mathbf{k}}$ factor, where it cancels out with the exponentials on the right. Note that for this, it is essential that $\alpha(\mathbf{k}')$ is a real-valued function. The last equality follows from the instance of the BCH formula derived above. In the same way, we have that

$$e^{i \int d^3\mathbf{k}' \alpha(\mathbf{k}') N_{\mathbf{k}'} } a_{\mathbf{k}}^\dagger e^{-i \int d^3\mathbf{k}' \alpha(\mathbf{k}') N_{\mathbf{k}'} } = e^{i\alpha(\mathbf{k})} a_{\mathbf{k}}^\dagger \quad (6)$$

Using our expressions for H and P^i in part (a) and (b), identities (5)-(6) allow us to compute

$$\begin{aligned} e^{i(Ht-P^i x^i)} a_{\mathbf{k}} e^{-i(Ht-P^i x^i)} &= e^{i \int d^3 \mathbf{k}' (\omega_{\mathbf{k}'} t - \mathbf{k}' \cdot \mathbf{x}) N_{\mathbf{k}'} + i E_0 t} a_{\mathbf{k}} e^{-i \int d^3 \mathbf{k}' (\omega_{\mathbf{k}'} t - \mathbf{k}' \cdot \mathbf{x}) N_{\mathbf{k}'} - i E_0 t} \\ &= e^{i \int d^3 \mathbf{k}' (\omega_{\mathbf{k}'} t - \mathbf{k}' \cdot \mathbf{x}) N_{\mathbf{k}'}} a_{\mathbf{k}} e^{-i \int d^3 \mathbf{k}' (\omega_{\mathbf{k}'} t - \mathbf{k}' \cdot \mathbf{x}) N_{\mathbf{k}'}} = e^{-i(\omega_{\mathbf{k}} t - \mathbf{k} \cdot \mathbf{x})} a_{\mathbf{k}} \\ e^{i(Ht-P^i x^i)} a_{\mathbf{k}}^\dagger e^{-i(Ht-P^i x^i)} &= e^{i \int d^3 \mathbf{k}' (\omega_{\mathbf{k}'} t - \mathbf{k}' \cdot \mathbf{x}) N_{\mathbf{k}'}} a_{\mathbf{k}}^\dagger e^{-i \int d^3 \mathbf{k}' (\omega_{\mathbf{k}'} t - \mathbf{k}' \cdot \mathbf{x}) N_{\mathbf{k}'}} = e^{i(\omega_{\mathbf{k}} t - \mathbf{k} \cdot \mathbf{x})} a_{\mathbf{k}}^\dagger \end{aligned}$$

Finally, we get

$$\begin{aligned} e^{i(Ht-P^i x^i)} \phi(0,0) e^{-i(Ht-P^i x^i)} &= \int \frac{d^3 \mathbf{k}}{\sqrt{2\omega_{\mathbf{k}}}} e^{i(Ht-P^i x^i)} (a_{\mathbf{k}} + a_{\mathbf{k}}^\dagger) e^{-i(Ht-P^i x^i)} \\ &= \int \frac{d^3 \mathbf{k}}{\sqrt{2\omega_{\mathbf{k}}}} (a_{\mathbf{k}} e^{-i(\omega_{\mathbf{k}} t - \mathbf{k} \cdot \mathbf{x})} + a_{\mathbf{k}}^\dagger e^{i(\omega_{\mathbf{k}} t - \mathbf{k} \cdot \mathbf{x})}) = \phi(\mathbf{x}, t) \end{aligned}$$

Question 4: Noether charges for Lorentz symmetries of a real scalar (20 points + 10 bonus)
In this problem we work out the conserved currents corresponding to the Lorentz symmetries of a real scalar theory,

$$\mathcal{L} = -\frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{1}{2}m^2\phi^2$$

(a) Consider an infinitesimal Lorentz transformation

$$\Lambda_\mu{}^\nu = \delta_\mu{}^\nu + \omega_\mu{}^\nu$$

where $\omega_{\mu\nu} = -\omega_{\nu\mu}$ are infinitesimal numbers. Show that this satisfies

$$\Lambda_\mu{}^\rho\eta_{\rho\lambda}\Lambda_\nu{}^\lambda = \eta_{\mu\nu}$$

to first order in $\omega_{\mu\nu}$, so this does give a Lorentz transformation.

We compute:

$$\begin{aligned}\Lambda_\mu{}^\rho\eta_{\rho\lambda}\Lambda_\nu{}^\lambda &= (\delta_\mu{}^\rho + \omega_\mu{}^\rho)\eta_{\rho\lambda}(\delta_\nu{}^\lambda + \omega_\nu{}^\lambda) = (\eta_{\mu\lambda} + \omega_{\mu\lambda})(\delta_\nu{}^\lambda - \omega_\nu{}^\lambda) \\ &= \eta_{\mu\nu} + \omega_{\mu\nu} - \omega_{\mu\nu} + \omega_{\mu\lambda}\omega_\nu{}^\lambda = \eta_{\mu\nu} + \mathcal{O}(\omega^2)\end{aligned}$$

(b) Write down how ϕ transforms under an infinitesimal Lorentz transformation, and show that the conserved Noether current for this transformation can be written as

$$J^{\mu\lambda\nu} = x^\lambda T^{\mu\nu} - x^\nu T^{\mu\lambda}$$

where $T^{\mu\nu}$ is the conserved energy-momentum tensor derived in problem set 1.

A Lorentz scalar field transforms in a way obeying $\phi'(x') = \phi(x)$. Therefore, under an infinitesimal Lorentz transformation, the scalar field ϕ transforms as

$$\delta\phi = \phi'(x) - \phi(x) = \phi((\Lambda^{-1})^\mu{}_\nu x^\nu) - \phi(x^\nu) = \phi((\delta^\mu{}_\nu - \omega^\mu{}_\nu)x^\nu) - \phi(x^\nu) = -\omega^\mu{}_\nu x^\nu \partial_\mu\phi$$

where in the last equality we Taylor expand $\phi(x^\mu - \omega^\mu{}_\nu x^\nu) = -\omega^\mu{}_\nu x^\nu \partial_\mu\phi$.

Using this, the Lagrangian density transforms as

$$\delta\mathcal{L} = \mathcal{L}[\phi'] - \mathcal{L}[\phi] = \mathcal{L}[\phi - \omega^\lambda{}_\nu x^\nu \partial_\lambda\phi] - \mathcal{L}[\phi] = -\omega^\lambda{}_\nu x^\nu \partial_\lambda\phi \frac{\partial\mathcal{L}}{\partial\phi} = -\omega^\lambda{}_\nu x^\nu \partial_\lambda\mathcal{L} = -\partial_\lambda(\omega^\lambda{}_\nu x^\nu \mathcal{L})$$

We expand only to first order in ω . In the final equality, we use that $\omega_{\mu\nu}$ is antisymmetric, i.e.

$$\partial_\lambda(\omega^\lambda{}_\nu x^\nu f(x)) = \omega^\lambda{}_\nu \delta_\lambda^\nu f(x) + \omega^\lambda{}_\nu x^\nu \partial_\lambda f(x) = \omega^\lambda{}_\lambda f(x) + \omega^\lambda{}_\nu x^\nu \partial_\lambda f(x) = \omega^\lambda{}_\nu x^\nu \partial_\lambda f(x)$$

Hence, $\delta\mathcal{L} = \partial_\mu\mathcal{F}^\mu$, for $\mathcal{F}^\mu = -\omega^\mu{}_\nu x^\nu \mathcal{L}$

The Noether current for the transformation parameterized by $\omega_{\lambda\nu}$ is given by

$$\begin{aligned}j^\mu &= -\frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi)}\delta\phi - \mathcal{F}^\mu = \omega^\lambda{}_\nu x^\nu \partial^\mu\phi \partial_\lambda\phi + \omega^\mu{}_\nu x^\nu \mathcal{L} \\ &= \omega_{\lambda\nu} x^\nu (\partial^\mu\phi \partial^\lambda\phi + \eta^{\mu\lambda} \mathcal{L}) = \omega_{\lambda\nu} x^\nu T^{\mu\lambda}\end{aligned}$$

for the energy-momentum tensor from problem set 1 (now with a real scalar):

$$T^{\mu\nu} = \partial^\mu\phi\partial^\nu\phi + \eta^{\mu\nu}\mathcal{L} = \partial^\mu\phi\partial^\nu\phi - \frac{1}{2}\eta^{\mu\nu}(\partial_\rho\phi\partial^\rho\phi + m^2\phi^2)$$

Note that $\omega_{\lambda\nu}$ is an arbitrary antisymmetric tensor which parameterizes our infinitesimal transformation. In total we have an antisymmetric tensor worth of conserved currents, which we can package in $J^{\mu\lambda\nu}$:

$$J^{\mu\lambda\nu} = x^\lambda T^{\mu\nu} - x^\nu T^{\mu\lambda}, \quad j^\mu = -\frac{1}{2}\omega_{\lambda\nu} J^{\mu\lambda\nu}$$

In writing $J^{\mu\lambda\nu}$ we have made the antisymmetry in λ and ν manifest by explicitly antisymmetrizing over these indices.

(c) Using conservation of the energy-momentum tensor, verify that the current in (b) is conserved, i.e.

$$\partial_\mu J^{\mu\lambda\nu} = 0$$

We compute:

$$\begin{aligned} \partial_\mu J^{\mu\lambda\nu} &= \partial_\mu(x^\lambda T^{\mu\nu} - x^\nu T^{\mu\lambda}) = \delta_\mu^\lambda T^{\mu\nu} + x^\lambda \partial_\mu T^{\mu\nu} - \delta_\mu^\nu T^{\mu\lambda} - x^\nu \partial_\mu T^{\mu\lambda} \\ &= T^{\lambda\nu} - T^{\nu\lambda} = 0 \end{aligned}$$

In the 3rd equality we used the conservation law $\partial_\mu T^{\mu\nu} = 0$, and in the 4th equality we used from problem set 1 that $T^{\mu\nu} = T^{\nu\mu}$ is symmetric.

(d) (**Bonus problem**) Consider the conserved charges associated with $J^{\mu\lambda\nu}$,

$$M^{\lambda\nu} = \int d^3x J^{0\lambda\nu}$$

Express the conserved charges $M^{\mu\nu}$ for the Lorentz symmetries of this theory in terms of $a_{\mathbf{k}}$ and $a_{\mathbf{k}}^\dagger$. From part (b), we have

$$M^{\mu\nu} = \int d^3x J^{0\mu\nu} = \int d^3x (x^\mu T^{0\nu} - x^\nu T^{0\mu})$$

This is antisymmetric in μ and ν , so we need to compute M^{0i} and M^{ij} . To do this, we need to expand $T^{0\mu}$ in terms of creation and annihilation operators:

$$\begin{aligned} T^{0\mu} &= -\pi\partial^\mu\phi - \frac{1}{2}\eta^{0\mu}(-\pi^2 + (\nabla\phi)^2 + m^2\phi^2) \\ &= \left(\frac{1}{2}(\pi^2 + (\nabla\phi)^2 + m^2\phi^2), -\pi\partial^i\phi\right) = (\mathcal{H}(x), \mathcal{P}^i(x)) \end{aligned}$$

Since $M^{\mu\nu}$ are conserved currents, we can compute them at $t = 0$.

First, we need the identity

$$\begin{aligned} \int d^3\mathbf{x} x^i f(\mathbf{x}) g(\mathbf{x}) &= \int d^3\mathbf{x} d^3\mathbf{k} d^3\mathbf{k}' e^{i(\mathbf{k}+\mathbf{k}')\cdot\mathbf{x}} x^i f(\mathbf{k}) g(\mathbf{k}') \\ &= \frac{1}{(2\pi)^3} \int d^3\mathbf{k} d^3\mathbf{k}' (-i\partial_{k_i} \delta^{(3)}(\mathbf{k} + \mathbf{k}')) f(\mathbf{k}) g(\mathbf{k}') = i \int d^3\mathbf{k} \partial_{k_i} f(\mathbf{k}) g(-\mathbf{k}) \end{aligned} \quad (7)$$

where we have used integration by parts in the last equality. More generally if there are derivatives acting on f or g , each derivative acting on f drags down a factor of $i\mathbf{k}$, while each derivative acting on g drags down a factor of $i\mathbf{k}'$, which becomes $-i\mathbf{k}$ after performing the $d^3\mathbf{x}$ integral.

Now we are ready to compute

$$\begin{aligned}
M^{0i} &= \int d^3\mathbf{x} (t\mathcal{P}^i(x) - x^i\mathcal{H}(x))|_{t=0} = -\frac{1}{2} \int d^3\mathbf{x} x^i (\pi^2 + (\nabla\phi)^2 + m^2\phi^2) \\
&= -\frac{i}{2} \int d^3\mathbf{k} (\partial_{k_i}\pi(\mathbf{k})\pi(-\mathbf{k}) + \omega_{\mathbf{k}}^2\partial_{k_i}\phi(\mathbf{k})\phi(-\mathbf{k})) \\
&= -\frac{i}{4} \int d^3\mathbf{k} \left(-\partial_{k_i}(\sqrt{\omega_{\mathbf{k}}}(a_{\mathbf{k}} - a_{-\mathbf{k}}^\dagger)) \cdot (\sqrt{\omega_{\mathbf{k}}}(a_{-\mathbf{k}} - a_{\mathbf{k}}^\dagger)) \right. \\
&\quad \left. + \omega_{\mathbf{k}}^2\partial_{k_i} \left(\frac{1}{\sqrt{\omega_{\mathbf{k}}}}(a_{\mathbf{k}} + a_{-\mathbf{k}}^\dagger) \right) \cdot \left(\frac{1}{\sqrt{\omega_{\mathbf{k}}}}(a_{-\mathbf{k}} + a_{\mathbf{k}}^\dagger) \right) \right) \\
&= +\frac{i}{4} \int d^3\mathbf{k} \left((k^i + \omega_{\mathbf{k}}\partial_{k_i})(a_{\mathbf{k}} - a_{-\mathbf{k}}^\dagger) \cdot (a_{-\mathbf{k}} - a_{\mathbf{k}}^\dagger) + (k^i - \omega_{\mathbf{k}}\partial_{k_i})(a_{\mathbf{k}} + a_{-\mathbf{k}}^\dagger) \cdot (a_{-\mathbf{k}} + a_{\mathbf{k}}^\dagger) \right) \\
&= -\frac{i}{2} \int d^3\mathbf{k} \omega_{\mathbf{k}} \left((\partial_{k_i}a_{\mathbf{k}})a_{\mathbf{k}}^\dagger + (\partial_{k_i}a_{-\mathbf{k}}^\dagger)a_{-\mathbf{k}} \right) = -\frac{i}{2} \int d^3\mathbf{k} \omega_{\mathbf{k}} \left((\partial_{k_i}a_{\mathbf{k}})a_{\mathbf{k}}^\dagger - (\partial_{k_i}a_{\mathbf{k}}^\dagger)a_{\mathbf{k}} \right) \\
&= -\frac{i}{2} \int d^3\mathbf{k} \omega_{\mathbf{k}} \left(a_{\mathbf{k}}^\dagger(\partial_{k_i}a_{\mathbf{k}}) + (2\pi)^3\partial_{k_i}\delta(\mathbf{k} - \mathbf{k}')|_{\mathbf{k}'=\mathbf{k}} - (\partial_{k_i}a_{\mathbf{k}}^\dagger)a_{\mathbf{k}} \right) \\
&= -\frac{i}{2} \int d^3\mathbf{k} \omega_{\mathbf{k}} \left(a_{\mathbf{k}}^\dagger(\partial_{k_i}a_{\mathbf{k}}) - (\partial_{k_i}a_{\mathbf{k}}^\dagger)a_{\mathbf{k}} \right)
\end{aligned}$$

Line 2 follows from identity (7). Line 5 is obtained by noting that all but 2 terms in the integrand of line 4 either cancel out or are odd. In line 6 we use the identity obtained by taking the \mathbf{k} -derivative of $[a_{\mathbf{k}}, a_{\mathbf{k}'}^\dagger] = (2\pi)^3\delta^{(3)}(\mathbf{k} - \mathbf{k}')$, and evaluating at $\mathbf{k}' = \mathbf{k}$. Finally, to reach line 7 we use that the $\delta^{(3)}$ -term in line 6 is odd.

In a similar way, we can also compute for $i \neq j$ (since $M^{ii} = 0$ by asymmetry)

$$\begin{aligned}
M^{ij} &= \int d^3\mathbf{x} (x^i\mathcal{P}^j(x) - x^j\mathcal{P}^i(x)) = - \int d^3\mathbf{x} (x^i\pi\partial^j\phi - x^j\pi\partial^i\phi) \\
&= \frac{1}{2} \int d^3\mathbf{k} (k^j\pi(\mathbf{k})\partial_{k_i}\phi(-\mathbf{k}) - (i \leftrightarrow j)) \\
&= -\frac{i}{4} \int d^3\mathbf{k} \left(k^j\sqrt{\omega_{\mathbf{k}}}(a_{\mathbf{k}} - a_{-\mathbf{k}}^\dagger) \cdot \partial_{k_i} \left(\frac{1}{\sqrt{\omega_{\mathbf{k}}}}(a_{-\mathbf{k}} + a_{\mathbf{k}}^\dagger) \right) - (i \leftrightarrow j) \right) \\
&= -\frac{i}{4} \int d^3\mathbf{k} \left(k^j(a_{\mathbf{k}} - a_{-\mathbf{k}}^\dagger) \cdot \left(-\frac{k^i}{\omega_{\mathbf{k}}} + \partial_{k_i} \right) (a_{-\mathbf{k}} + a_{\mathbf{k}}^\dagger) - (i \leftrightarrow j) \right) \\
&= -\frac{i}{4} \int d^3\mathbf{k} k^j \left(a_{\mathbf{k}}\partial_{k_i}a_{-\mathbf{k}} + a_{\mathbf{k}}\partial_{k_i}a_{\mathbf{k}}^\dagger - a_{-\mathbf{k}}^\dagger\partial_{k_i}a_{-\mathbf{k}} - a_{-\mathbf{k}}^\dagger\partial_{k_i}a_{\mathbf{k}}^\dagger \right) - (i \leftrightarrow j) \\
&= -\frac{i}{4} \int d^3\mathbf{k} k^j \left(a_{\mathbf{k}}\partial_{k_i}a_{\mathbf{k}}^\dagger - a_{-\mathbf{k}}^\dagger\partial_{k_i}a_{-\mathbf{k}} \right) - (i \leftrightarrow j) \\
&= -\frac{i}{4} \int d^3\mathbf{k} k^j \left(-(\partial_{k_i}a_{\mathbf{k}})a_{\mathbf{k}}^\dagger - a_{\mathbf{k}}^\dagger(\partial_{k_i}a_{\mathbf{k}}) \right) - (i \leftrightarrow j) \\
&= -\frac{i}{4} \int d^3\mathbf{k} k^j \left((2\pi)^3\partial_{k_i}\delta(\mathbf{k} - \mathbf{k}')|_{\mathbf{k}'=\mathbf{k}} - 2a_{\mathbf{k}}^\dagger(\partial_{k_i}a_{\mathbf{k}}) \right) - (i \leftrightarrow j) \\
&= -\frac{i}{2} \int d^3\mathbf{k} \left(k^i a_{\mathbf{k}}^\dagger(\partial_{k_j}a_{\mathbf{k}}) - k^j(\partial_{k_i}a_{\mathbf{k}}^\dagger)a_{\mathbf{k}} \right)
\end{aligned}$$

where we make ample use of integration by parts, drop total-derivative terms, and use $i \neq j$ to completely ignore having to deal with $\propto \partial_{k_i}k^j$ terms.

Physically, the conserved quantities are the center of mass velocities M^{0i} and angular momenta M^{ij} .

Altogether, we may write

$$M^{\mu\nu} = -\frac{i}{2} \int d^3\mathbf{k} k^\mu (a_{\mathbf{k}}^\dagger \partial_{k_\nu} a_{\mathbf{k}} - \partial_{k_\nu} a_{\mathbf{k}}^\dagger a_{\mathbf{k}}) - (\mu \rightarrow \nu), \quad k^0 = \omega_{\mathbf{k}}$$

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