

Quantum Field Theory in Statistical Physics

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山无数，乱红如雨，不记来时路。

——相对论吧纪念墙

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Preliminaries

1.1 Three Pictures in Quantum Mechanics

Generally, there are merely two kinds of quantities one can directly measure in microscopic systems: the eigenvalue of observables and the corresponding probabilities. To understand the experimental results essentially, people brought in two abstract things in quantum mechanics: observables (unbounded self-adjoint operators) and bras (vectors in rigged Hilbert Space), but which of them undertook the task of time-evolution became the subsequently urgent problem. Depending on the different perspectives on time evolution, people then developed three so-called *pictures*.

Anyway, whatever standpoints we adopt, they must be physically equivalent, i.e., the calculated quantities related to observables must be equal to each other, or

$$\langle \psi^S | \Omega^S | \phi^S \rangle = \langle \psi^H | \Omega^H | \phi^H \rangle = \langle \psi^I | \Omega | \phi^I \rangle. \quad (1.1.1)$$

To discuss the distinction of these three pictures, we set three pictures coincide at initial time for convenience

$$|\psi^S(0)\rangle = |\psi^H(0)\rangle = |\psi^I(0)\rangle.$$

1.1.1 Schrödinger Picture

In the most familiar Schrödinger picture, vectors undertake all the time-evolution brought by Hamiltonian $H(t)$.

The formal solution of Schrödinger equation

$$i\hbar \frac{\partial}{\partial t} |\psi^S(t)\rangle = H(t) |\psi^S(t)\rangle \quad (1.1.2)$$

obviously takes the form of

$$|\psi^S(t)\rangle = U^S(t, t_0) |\psi^S(0)\rangle, \quad (1.1.3)$$

where we have introduced the unitary operator

$$\begin{aligned} U^S(t, t_0) &= \lim_{\Delta t \rightarrow 0} e^{-iH(t)\Delta t/\hbar} e^{-iH(t-\Delta t)\Delta t/\hbar} \dots e^{-iH(t_0)\Delta t/\hbar} \\ &:= T \exp \left(-\frac{i}{\hbar} \int_{t_0}^t d\tau H(\tau) \right). \end{aligned}$$

It is valuable to mention that not all observables in Schrödinger picture are independent of time. Actually, the Hamiltonian, for example, which measures the evolution of system, may be explicit function of time.

1.1.2 Heisenberg Picture

In Heisenberg picture, however, it's the dynamic operators rather than vectors that undertake all the time-evolution brought by Hamiltonian $H(t)$.

Note that Schrödinger equation, as the third principle of axiomatic quantum mechanics, still must work here, so the demands that states never evolve with time equally means an inverse evolution on the Schrödinger picture that:

$$|\psi^H(t)\rangle = U^S(t, t_0)^{-1} |\psi^S(t)\rangle \equiv |\psi(t_0)\rangle. \quad (1.1.4)$$

And to guarantee the conservation of probability amplitude, (1.1.1) gives

$$\Omega^H(t) = U^S(t, t_0)^{-1} \Omega^S U^S(t, t_0). \quad (1.1.5)$$

Differentiate (1.1.4) and (1.1.5), we get

$$\begin{cases} \frac{d}{dt} |\psi^H(t)\rangle = 0, \\ \frac{d}{dt} \Omega^H(t) = \frac{\partial}{\partial t} \Omega^H(t) + \frac{1}{i\hbar} [\Omega^H(t), H^H(t)], \end{cases} \quad (1.1.6)$$

where $\frac{\partial \Omega^H(t)}{\partial t} = U^S(t, t_0)^{-1} \frac{\partial \Omega^S}{\partial t} U^S(t, t_0)$ and $H^H(t) = U^S(t, t_0)^{-1} H(t) U^S(t, t_0)$

1.1.3 Interaction Picture

In perturbation theory, we will divide the Hamiltonian into free and interactive parts $H = H_0 + H_i$. They respectively determine the kinetic and dynamic kinds of evolution. We define the state at time t in interaction picture to be the state in Schrödinger picture that inversely evolves with the *free parts* of Hamiltonian:

$$|\psi^I(t)\rangle := U_0^S(t, t_0)^{-1} |\psi^S(t)\rangle, \quad (1.1.7)$$

where U_0^S denotes the free part evolution, i.e., $U_0^S(t, t_0) = T \exp \left(-\frac{i}{\hbar} \int_{t_0}^t d\tau H_0(\tau) \right)$. Similarly, to ensure the conservation of probability amplitude, we must have

$$|\psi^I(t)\rangle = U^I(t, t_0) |\psi(t_0)\rangle, \quad U^I(t, t_0) \equiv U_0^S(t, t_0)^{-1} U^S(t, t_0) \quad (1.1.8)$$

$$\Omega^I(t) = U_0^S(t, t_0)^{-1} \Omega^S(t) U_0^S(t, t_0). \quad (1.1.9)$$

In quantum fields theory, people would rather call the unitary evolving operator in interactive picture $U^I(t, t_0)$ the *S-matrix*, denoting it by $S(t; t_0)$ instead, i.e.,

$$S(t, t_0) \equiv U_0^{-1}(t, t_0) U(t, t_0), \quad (1.1.10)$$

where we suppress the superscript of picture (because we only use the notation of evolving operation introduced in Schrödinger Picture).

The evolution equation can also be easily obtained by differentiating (1.1.8) and substituting (1.1.9), but it is a little departure of our topic that we will never use in the future, so I just skip it.

1.2 Schrödinger Field and the Second Quantization

Though probability is thoroughly a quantity in quantum mechanics, in the paradigm of quantum fields theory, we regard ψ as just the variable of *classical fields*.

Introducing the *free* lagrangian¹

定义 1(Schrödinger Field)

$$\mathcal{L} = i\hbar\psi^*(\mathbf{r}, t)\dot{\psi}(\mathbf{r}, t) - \frac{\hbar^2}{2m}\nabla\psi^* \cdot \nabla\psi - V\psi^*\psi, \quad (1.2.1)$$

where the complex fields ψ and ψ^* are independent with each other, and applying the celebrated Euler-Lagrangian equation on ψ^* , i.e.,

$$\partial_\mu \frac{\partial \mathcal{L}}{\partial(\partial_\mu \psi^*)} - \frac{\partial \mathcal{L}}{\partial \psi^*} = 0,$$

we get the familiar Schrödinger equation of ψ

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + V\psi.$$

Now that we are in the formalism of classical field theory, the corresponding canonical momentum of field quantity is

$$\pi(\mathbf{r}, t) = \frac{\partial \mathcal{L}}{\partial \dot{\psi}(\mathbf{r}, t)} = i\hbar\psi^*. \quad (1.2.2)$$

And then the Hamiltonian density can be written as

$$\mathcal{H} \equiv \pi\dot{\psi} - \mathcal{L} = \frac{\hbar^2}{2m} \nabla\psi^* \cdot \nabla\psi + V\psi^*\psi. \quad (1.2.3)$$

¹I have to emphasize here that although V is widely interpreted as “interaction” in QM literature, since what we derive from it is the fundamental Schrödinger equation, just as the Klein-Gordon equation derived from non-interactive scalar field Lagrangian and the Dirac equation derived from non-interactive spinor field Lagrangian. So in this sense, we interpret the lagrangian with V here as “free”.

According to the definition of Poisson bracket (at equal time)² that

$$\{f[\psi, \pi], g[\psi, \pi]\} \equiv \int d\mathbf{r}' \left\{ \frac{\delta f[\psi, \pi]}{\delta \psi(\mathbf{r}', t)} \frac{\delta g[\psi, \pi]}{\delta \pi(\mathbf{r}', t)} - \frac{\delta f[\psi, \pi]}{\delta \pi(\mathbf{r}', t)} \frac{\delta g[\psi, \pi]}{\delta \psi(\mathbf{r}', t)} \right\}$$

we get

$$\{\psi(\mathbf{r}, t), \pi(\mathbf{r}', t)\} = \delta(\mathbf{r} - \mathbf{r}'), \quad (1.2.4)$$

$$\{\psi(\mathbf{r}, t), \psi(\mathbf{r}', t)\} = \{\pi(\mathbf{r}, t), \pi(\mathbf{r}', t)\} = 0. \quad (1.2.5)$$

Fortunately, without the annoying gauge invariance in electromagnetic fields, there is no *constraint* in our Schrödinger field, so according to the standard methods for canonically quantizing a constraint system developed by Dirac³, the procedure of second quantization is just simple replacement of “classical field” ψ by the field operators $\hat{\psi}$ and Poisson bracket $\{, \}$ by Dirac bracket⁴ $[\cdot, \cdot]_{\pm}$ with a coefficient $1/i\hbar$, respectively, i.e.,

$$\{A, B\} \mapsto \frac{1}{i\hbar} [\hat{A}, \hat{B}]_{\pm}.$$

So

公理 1(Second Quantization of Schrödinger Field) The canonically quantized variables of ψ and $\pi \equiv i\hbar\psi^*$, denoted as $\hat{\psi}$ and $\hat{\psi}^\dagger$ respectively, in Heisenberg picture, have the commutation relation that

$$[\hat{\psi}(\mathbf{r}, t), \hat{\psi}^\dagger(\mathbf{r}', t)]_{\pm} = i\hbar\delta(\mathbf{r} - \mathbf{r}'), \quad (1.2.6)$$

$$[\hat{\psi}(\mathbf{r}, t), \hat{\psi}^\dagger(\mathbf{r}', t)]_{\pm} = [\hat{\psi}^\dagger(\mathbf{r}, t), \hat{\psi}^\dagger(\mathbf{r}', t)] = 0. \quad (1.2.7)$$

注 1 One may wonder why we recklessly quantize the Schrödinger in *both* the commutative and anti-commutative rule, which is aberrant to what we have learned from QFT that one kind of field corresponds with just one kind of particles (or more precisely, excited by). That is because the energy scale in condensed matter fields is so low (less than 100 eV) that we naturally regard the speed of light not be finite but infinite, and hence in CMP we only consider the *non-relativistic* QFT. So **the confinement of casualties on commutators**⁵ **disappears.**

²cf. chapter 7 in Weinberg I.

³This temporarily works in 3 + 1 dimensional fields theory and 4-dimensional Euclidean fields theory. But things become esoteric for low-dimensional case, which we will study in the future part.

⁴Here positive sign corresponds to Bosons while minus sign corresponds to Fermions

⁵Some reminder on casualty in Relativistic QFT: **Lorentz invariance of S-matrix can be directly deduced from definition of S-matrix, which in turn imposes confinement that the interactive density of Hamiltonian must commute for space-like interval. By definition of fields as well as properties of creation and annihilation operators, any fields operators in QFT can be expanded as multiples of creation and annihilation operations with some coefficients, then the only way to construct the fields operator is linear combination of creation and annihilation fields operators, i.e., the familiar so-called casualty. disappear and we are free to quantize the Schrödinger field haphazardly.** cf. chapters 2-5 in Weinberg I.

1.3 Occupation Number Representation

Given a complete set of function $\{\psi_k(\mathbf{r})\}$ such that

$$\left(-\frac{\hbar^2}{2m}\Delta + V(\mathbf{r})\right)\psi_k = E_k\psi_k$$

then by Sturm-Liouville theorem, this family of functions are orthogonally normalized and complete, i.e., $\langle\psi_k|\psi_l\rangle = \delta_{kl}$ and $\sum_k |\psi_k\rangle\langle\psi_k| = \mathbb{1}$.

So we can expand the field operator by combining the operator a and a^\dagger bearing the non-commutative properties of operators with the coefficients of complete function basis above

断语 1(Expansion of Fields Operators) Generally, fields operators in Schrödinger Picture take the form of (here we write in coordinate representation and add an extra spin label on the complete function basis)

$$\hat{\psi}(\mathbf{r}) = \sum_{k\sigma} a_{k\sigma}\psi_{k\sigma}(\mathbf{r}), \quad \hat{\psi}^\dagger(\mathbf{r}) = \sum_{k\sigma} a_{k\sigma}^\dagger\psi_{k\sigma}^*(\mathbf{r}), \quad (1.3.1)$$

with

$$[a_{k\sigma}, a_{k'\sigma'}^\dagger]_\pm = \delta_{kk'}\delta_{\sigma\sigma'}, \quad [a_{k\sigma}, a_{k'\sigma'}]_\pm = [a_{k\sigma}^\dagger, a_{k'\sigma'}^\dagger]_\pm = 0. \quad (1.3.2)$$

证明 Multiplying a ket $\langle\psi_{k\sigma}|$ on (1.3.1) without choosing of representation, one immediately obtains $a_{k\sigma} = \langle\psi_{k\sigma}|\hat{\psi}\rangle$ and $a_{k\sigma}^\dagger = \langle\hat{\psi}|\psi_{k\sigma}\rangle$, or in coordinate representation

$$a_{k\sigma} = \int d\mathbf{r} \psi_{k\sigma}(\mathbf{r})\hat{\psi}(\mathbf{r}), \quad a_{k\sigma}^\dagger = \int d\mathbf{r} \psi_{k\sigma}^*(\mathbf{r})\hat{\psi}^\dagger(\mathbf{r}).$$

Then (in natural units)

$$\begin{aligned} [a_{k\sigma}, a_{k'\sigma'}^\dagger]_\pm &= \int d\mathbf{r} d\mathbf{r}' [\hat{\psi}(\mathbf{r}, t), \hat{\psi}^\dagger(\mathbf{r}', t)]_\pm \psi_{k\sigma}(\mathbf{r}')\psi_{k'\sigma'}^*(\mathbf{r}) \\ &= \int d\mathbf{r} d\mathbf{r}' \delta(\mathbf{r} - \mathbf{r}') \psi_{k\sigma}(\mathbf{r}')\psi_{k'\sigma'}^*(\mathbf{r}) \\ &= \int d\mathbf{r} \psi_{k'\sigma'}(\mathbf{r})\psi_k^*(\mathbf{r}) = \sum_k \langle\psi_{k'\sigma'}|\psi_{k\sigma}\rangle = \delta_{kk'}\delta_{\sigma\sigma'}, \end{aligned}$$

where in the second line we utilize the commutation relation (1.2.6) in Schrödinger Picture, which is easy to show. \square

例 1(Non-interactive Fermions)

In solid state physics, we already knew that **Bloch electrons**, in flawless periodic lattice, is described by the “box-normalized” plane wave:

$$\psi_{\mathbf{p}}(\mathbf{r}) = \frac{1}{\sqrt{V}} e^{i\mathbf{p}\cdot\mathbf{r}},$$

where V is the volume of the box. In this common case, (1.3.1) takes the form

$$\hat{\psi}(\mathbf{r}, t) = \frac{1}{\sqrt{V}} \sum_{\mathbf{p}} a_{\mathbf{p}} e^{i\mathbf{p} \cdot \mathbf{r}}. \quad (1.3.3)$$

命题 1 The Hamiltonian $\hat{H} \equiv \int d\mathbf{r} \mathcal{H}$ always commutes with the k -mode particle number operator $\hat{N}_k := a_k^\dagger a_k$ if it is not explicit function of time, i.e.,

$$[\hat{H}, \hat{N}_k] = 0, \quad \forall k. \quad (1.3.4)$$

So the number of k -mode quasi-particles is independent of time for a stable system (at equilibrium).

证明 Choosing energy eigen-functions as complete basis, then \hat{H} can be explicit written by creation and annihilation operators:

$$\begin{aligned} \hat{H} &= \int d\mathbf{r} \hat{\psi}^\dagger(\mathbf{r}) \left(-\frac{\hbar^2}{2m} \nabla^2 + V \right) \hat{\psi}(\mathbf{r}) \\ &= \sum_{k, k'} a_{k'}^\dagger a_k E_k \int d\mathbf{r} \psi_{k'}^*(\mathbf{r}, t) \psi_k(\mathbf{r}, t) = \sum_k E_k a_k^\dagger a_k, \end{aligned}$$

where the first line comes from the rule of the second quantization. Obviously, for each k , Hamiltonian commutes with the k -mode quasi-particle number operator. \square

Since each \hat{N}_k also commute with each other, $\{\hat{H}, \hat{N}_k\}$ forms one *complete set of commuting observables*, so they can be simultaneously diagonalized under the *simultaneous eigenstates*, temporarily denoted as $|\psi_{N_k}\rangle$, which must belong to the multi-particle space

$$|\psi_{N_k}\rangle \in \mathcal{H}_{N_k} = \mathcal{H}_1^{(1)} \otimes \cdots \otimes \mathcal{H}_1^{(N_k)} \equiv \bigotimes_{i=1}^{N_k} \mathcal{H}_1^{(i)},$$

where the subscript 1 indicates the single-particle Hilbert space, while the superscript (i) indicates the label of particles.

1.4 Identical Multi-particle State

As is shown above, now we are working in the multi-particle states. If all the k -mode particles are of the same type, then *identical principle* will certainly come into effects and bring into complexity.

Let $|\varphi_\alpha^{(i)}\rangle$ be the α -th eigenstate of the Hamiltonian $H_k^{(i)}$ of the i -th k -mode particle, i.e., in the Hilbert space $\mathcal{H}_1^{(i)}$. Then an arbitrary N -particle state must be written as

$$|\varphi_{N_k}\rangle \equiv |\varphi_{\alpha_1}^{(1)}\rangle \otimes \cdots \otimes |\varphi_{\alpha_N}^{(N)}\rangle = |\varphi_{\alpha_1}^{(1)} \cdots \varphi_{\alpha_N}^{(N)}\rangle. \quad (1.4.1)$$

This basis of \mathcal{H}_N must be complete because each components of it is the complete basis of single-particle state⁶.

According to **identical principle**⁷, **some unitary phases are allowed exits after interchanging two identical particles**

$$|\psi_{N_k}(\cdots, i, \cdots, j, \cdots)\rangle = e^{i\theta(\cdots)} |\psi_{N_k}(\cdots, j, \cdots, i, \cdots)\rangle,$$

where $\theta(\cdots)$ is the **one-dimensional irreducible unitary representation of the group of interchanging particles**. In our case of *three-dimensional* configuration space, this group is the **permutation group**⁸, whose **one-dimensional unitary representation** $\rho : \text{Perm} \rightarrow \mathbb{C}$ is **characterized by either a positive or minus sign**. More explicitly, denote the operator of transposition as P_{ij} such that

$$P_{ij}|\cdots\phi_{\alpha_i}^{(i)}\cdots\phi_{\alpha_j}^{(j)}\cdots\rangle := |\cdots\phi_{\alpha_i}^{(j)}\cdots\phi_{\alpha_j}^{(i)}\cdots\rangle,$$

then

$$P_{ij}|N_k\rangle = \varepsilon|N_k\rangle, \quad \varepsilon = \pm 1, \quad (1.4.2)$$

for the multi-particle state $|N_k\rangle$ including the effects of transposition (compared with the original one in (1.4.1)), with positive for bosons while minus for fermions.

The question is, **since extra structure is included in our theory, the prototype (1.4.1) is reasonable to be invalid**. But fortunately, a mathematical trick succeed in realizing this without discarding this:

定义 1(Symmetrized and Anti-symmetrized Multi-particle State) By introducing a appropriate normalization factor $1/N!$, the prescription⁹

$$|N_k\rangle \equiv |\psi_{N_k}^{(\varepsilon)}\rangle := \frac{1}{N_k!} \hat{S}_\varepsilon |\psi_{\alpha_1}^{(1)} \cdots \psi_{\alpha_{N_k}}^{(N_k)}\rangle, \quad (1.4.3)$$

with $\hat{S}_\varepsilon \equiv \sum_{\mathcal{P}} \varepsilon^{\mathcal{P}} \mathcal{P}$ known as *symmetrization operator*, where p is the number of transpositions constructing the permutation \mathcal{P} through p times of group multiplications, ε is the sign under each transposition, and the sum runs over *all* the permutation for the N -tube $(1, 2, \cdots, N)$.

For $\varepsilon = \pm 1$ we call $|N_k\rangle$ either *symmetrized or anti-symmetrized multi-particle state*. Obviously the space spanned by them are both smaller than the former one. We denote them as $\mathcal{H}_{N_k}^{(\varepsilon)}$, or $\mathcal{SH}_{N_k} \subset \mathcal{H}_{N_k}$ and $\mathcal{AH}_{N_k} \subset \mathcal{H}_{N_k}$, respectively.

⁶As for the reason why these eigenstates of the self-adjoint operator $\hat{\varphi}$ form the complete basis of single-particle Hilbert spaces, one can refer to the celebrated *spectrum decomposition theorem of unbounded self-adjoint operators* proved by Von Neumann. cf. GTM267.

⁷cf. chapter 4 in Weinberg I.

⁸Things alter for two-dimensional case, where the group describing transposition of two particles become the so-called *braid group*, since the process of interchanging two particles involves descriptions on the topology of their worldlines.

⁹Since (1.4.3) is an appropriate mathematical realization of (1.4.2), we naturally inherit the notation $|N_k\rangle$ for multi-particle states with interchanging effects.

To see that (1.4.3) is well-defined, we only need to check realization of (1.4.2). In fact, symmetrized (antisymmetrized) states are nothing but summation of usual ones, which always give out an extra sign ε under another transposition. That is,

$$P_{ij}\hat{S}_\varepsilon = \sum_{\mathcal{P}} \varepsilon^{\mathcal{P}} P_{ij} \mathcal{P} = \sum_{\mathcal{P}} \varepsilon^{\mathcal{P}+1} \mathcal{P} = \varepsilon \hat{S}_\varepsilon.$$

Inner products of symmetrized states will be used in the near future, let us briefly list here in advance. We begin with investigation on properties of the permutation operator:

性质 1

- a) $\mathcal{P} = \mathcal{P}^\dagger$;
- b) $[\mathcal{P}, A] = 0$;
- c) $\mathcal{P}^2 = \mathbb{1}$.

证明 The first property is easy to see. For an arbitrary permutation constructed by multiplication of transposition,

$$\langle \psi_N^{(\varepsilon)} | \mathcal{P}^\dagger | \phi_N^{(\varepsilon)} \rangle \equiv (\langle \phi_N^{(\varepsilon)} | \mathcal{P} | \psi_N^{(\varepsilon)} \rangle)^* = \varepsilon^{\mathcal{P}} (\langle \phi_N^{(\varepsilon)} | \psi_N^{(\varepsilon)} \rangle)^* = \varepsilon^{\mathcal{P}} \langle \psi_N^{(\varepsilon)} | \phi_N^{(\varepsilon)} \rangle = \langle \psi_N^{(\varepsilon)} | \mathcal{P} | \phi_N^{(\varepsilon)} \rangle.$$

To reveal the commutation relation, we invoke the *identical principle* mentioned above that

$$\begin{aligned} \langle \dots \psi_{\alpha_i}^{(i)} \dots \psi_{\alpha_j}^{(j)} \dots | A | \dots \phi_{\alpha_i}^{(i)} \dots \phi_{\alpha_j}^{(j)} \dots \rangle &= \langle \dots \psi_{\alpha_i}^{(j)} \dots \psi_{\alpha_j}^{(i)} \dots | e^{i\theta(\dots)} A e^{-i\theta(\dots)} | \dots \psi_{\alpha_i}^{(j)} \dots \psi_{\alpha_j}^{(i)} \dots \rangle \\ &= \langle \dots \psi_{\alpha_i}^{(j)} \dots \psi_{\alpha_j}^{(i)} \dots | A | \dots \phi_{\alpha_i}^{(j)} \dots \phi_{\alpha_j}^{(i)} \dots \rangle \\ &\equiv \langle \dots \psi_{\alpha_i}^{(i)} \dots \psi_{\alpha_j}^{(j)} \dots | P_{ij}^\dagger A P_{ij} | \dots \phi_{\alpha_i}^{(i)} \dots \phi_{\alpha_j}^{(j)} \dots \rangle, \end{aligned}$$

implying $A = P_{ij}^\dagger A P_{ij}$ or $[P_{ij}, A] = 0$ with the unitarity proven in property one. And since \mathcal{P} is group multiplication of transposition P_{ij} , the second property holds.

Property c) means that \mathcal{P} is *nilpotent*, which is obvious from the nilpotency of transposition operator P_{ij} . \square

引理 1 (Inner Products of Symmetrized (Anti-symmetrized) States)

$$\langle \psi_N^{(\varepsilon)} | A | \phi_N^{(\varepsilon)} \rangle = \langle \psi_{\alpha_1} \dots \psi_{\alpha_N} | A | \phi_N^{(\varepsilon)} \rangle. \quad (1.4.4)$$

证明 The average under symmetrized (anti-symmetrized) states can be shown in using of all properties above:

$$\begin{aligned} \langle \psi_N^{(\varepsilon)} | A | \phi_N^{(\varepsilon)} \rangle &= \frac{1}{N!} \sum_{\mathcal{P}} \varepsilon^{\mathcal{P}} \langle \psi_{\alpha_1} \dots \psi_{\alpha_N} | \mathcal{P}^\dagger A | \phi_N^{(\varepsilon)} \rangle \\ &= \frac{1}{N!} \sum_{\mathcal{P}} \varepsilon^{\mathcal{P}} \langle \psi_{\alpha_1} \dots \psi_{\alpha_N} | A \mathcal{P} | \phi_N^{(\varepsilon)} \rangle \\ &= \frac{1}{N!} \sum_{\mathcal{P}} \varepsilon^{2\mathcal{P}} \langle \psi_{\alpha_1} \dots \psi_{\alpha_N} | A | \phi_N^{(\varepsilon)} \rangle = \langle \psi_{\alpha_1} \dots \psi_{\alpha_N} | A | \phi_N^{(\varepsilon)} \rangle, \end{aligned}$$

where in the second line we make use of both the unitarity and commutative relation of \mathcal{P} and validity of definition on $|\psi_{N_k}^{(\varepsilon)}\rangle$, and in the last line $\varepsilon^{2p} = 1$ whatever the sign of our system is and the sum exactly contains $N!$ terms (that's why we introduce the normalized factor in definition of $|\psi_{N_k}^{(\varepsilon)}\rangle$). \square

Now let us get go on our previous discussion where **different energy level E_k corresponds to a different number of particles N_k .**

One might notice that what we defined before on creation and annihilation operators a_k and a_k^\dagger , as the linear operator on k -mode Hilbert spaces, never mention the essential domain spaces or image spaces of them, as well as the effects they act on states. So here we are to complete the definition:

定义 2(Creation Operators) The creation operator $a_\beta^\dagger : \mathcal{H}_{N_k}^{(\varepsilon)} \rightarrow \mathcal{H}_{N_k+1}^{(\varepsilon)}$ is defined as

$$a_\beta^\dagger |N_k\rangle \equiv a_\beta^\dagger |\varphi_{\alpha_1} \cdots \varphi_{\alpha_{N_k}}\rangle^{(\varepsilon)} := \sqrt{N_k+1} |\varphi_\beta \varphi_{\alpha_1} \cdots \varphi_{\alpha_{N_k}}\rangle^{(\varepsilon)} \equiv \sqrt{N_k+1} |N_k+1\rangle. \quad (1.4.5)$$

Likewise, as the adjoint operator of a_β^\dagger , a_β must also map from $\mathcal{H}_{N_k}^{(\varepsilon)}$ to $\mathcal{H}_{N_k-1}^{(\varepsilon)}$ with some coefficients waiting to be determined, which can be obtained by the commutation relation and the fact that our symmetrized or anti-symmetrized state $|N_k\rangle$ is always the eigenstate of $\hat{N}_k \equiv a_k^\dagger a_k$ such that

$$\hat{N}_k |\varphi_{\alpha_1} \cdots \varphi_{\alpha_{N_k}}\rangle^{(\varepsilon)} = N_k |\varphi_{\alpha_1} \cdots \varphi_{\alpha_{N_k}}\rangle^{(\varepsilon)}. \quad (1.4.6)$$

Hence we have

推论 1(Annihilation Operators)

$$a_{\alpha_i} |N_k\rangle \equiv a_{\alpha_i} |\varphi_{\alpha_1} \cdots \varphi_{\alpha_i} \cdots \varphi_{\alpha_{N_k}}\rangle^{(\varepsilon)} = \sqrt{N_k} |\varphi_{\alpha_1} \cdots \varphi_{\alpha_{i-1}} \varphi_{\alpha_{i+1}} \cdots \varphi_{\alpha_{N_k}}\rangle^{(\varepsilon)} = \sqrt{N_k} |N_k-1\rangle. \quad (1.4.7)$$

So we can now safely conclude that

命题 1 An arbitrary symmetrized or anti-symmetrized states *for a specific energy level* lives in the **single energy level Fork space** $F\mathcal{H}_{N_k}$ such that

$$|N_k\rangle \in \begin{cases} \bigoplus_{N_k=0}^{\infty} \mathcal{S}\mathcal{H}_{N_k} \equiv \bigoplus_{N_k=0}^{\infty} \mathcal{S} \left(\bigotimes_{i=0}^{N_k} \mathcal{H}_1^{(i)} \right), & \text{bosons,} \\ \bigoplus_{N_k=0}^{\infty} \mathcal{A}\mathcal{H}_{N_k} \equiv \bigoplus_{N_k=0}^{\infty} \mathcal{A} \left(\bigotimes_{i=0}^{N_k} \mathcal{H}_1^{(i)} \right), & \text{fermions.} \end{cases}$$

Applying (1.4.7) N_k times and taking the conjugate transpose, one easily gets

命题 2 The symmetrized or anti-symmetrized multi-particle states *for a specific energy level* can be constructed by applying creation operator N_k times on vacuum state

$$|N_k\rangle \equiv |\varphi_{\alpha_1} \cdots \varphi_{\alpha_{N_k}}\rangle^{(\varepsilon)} = \frac{1}{\sqrt{N_k!}} a_{\alpha_1}^\dagger \cdots a_{\alpha_{N_k}}^\dagger |0\rangle. \quad (1.4.8)$$

Although to define the action of creation and annihilation operator we manually pick out one specific energy level with particle number N_k , these two definitions are apparently independent of which energy level we choose. And since creation or annihilation a particle at one energy level has nothing to do with other levels, we naturally demand the *complete* (anti)-symmetrized multi-particle states as simple tensor products of $|N_k\rangle$:

断语 1(Multi-level Symmetrized (Anti-symmetrized) States) Whatever the sign of system we adopt, the complete state always takes the form of

$$|N\rangle = |N_1\rangle \otimes \cdots \otimes |N_k\rangle \otimes \cdots,$$

where $|N\rangle$ denotes the eigenstate of $\hat{N} = \sum_k \hat{N}_k$ with $\hat{N} = \sum_k \hat{N}_k$. Namely, the multi-level symmetrized (or anti-symmetrized) states lives in the tensor products of Fork space of each level

$$|N\rangle \in \bigotimes_{k=1}^m \left(\bigoplus_{N_k=0}^{\infty} \mathcal{H}_{N_k}^{(\varepsilon)} \right),$$

where k runs over m energy level indices.

推论 2(Eigenstates of Bosons \hat{N}) The state of all bosons is

$$|N\rangle_{\text{bosons}} = \frac{1}{\sqrt{N_1! \cdots N_k! \cdots}} (a_{\alpha}^{\dagger})^{N_1} \cdots (a_{\beta}^{\dagger})^{N_k} \cdots |0\rangle, \quad (1.4.9)$$

As for Fermions case, general formula (1.4.9) still holds, but one should note that **the number of particles of k -mode Fermions can only be either zero or one, as is required by *Pauli exclusion principle*:**

引理 2(Pauli Exclusion Principle) For each k , the eigenvalue of \hat{N}_k can only be one or zero.

证明 We show this by proving that at this time \hat{N}_k is the projective operator¹⁰

$$\hat{N}_k^2 = c_k^{\dagger} c_k c_k^{\dagger} c_k = c_k^{\dagger} (1 - c_k^{\dagger} c_k) c_k = \hat{N}_k,$$

where we substitute the anti-commute relation for the creation and annihilation operators for fermions. \square

Since both the factorials of one and zero are one, for fermions we have

推论 3(Eigenstates of Fermions \hat{N}) The state of all fermions is

$$|N\rangle_{\text{fermions}} = c_{\alpha}^{\dagger} \cdots c_{\beta}^{\dagger} \cdots |0\rangle. \quad (1.4.10)$$

1.5 Multi-body Operators

In this section, we will express the operators acting on the Hilbert space we constructed above with creation and annihilation operators.

¹⁰We here use c and c^{\dagger} to make the distinction with boson case.

定义 1(Single-body and Two-body Operator) For an arbitrary observable acting on symmetrized (or anti-symmetrized) Hilbert space of N_i particles at a specific energy level $A \in \text{End}(\mathcal{H}_{N_i}^{(\varepsilon)}) = \text{End}\left(\mathcal{S} \bigotimes_{i=1}^{N_i} \mathcal{H}_1^{(i)}\right)$, the action of A must also be symmetrized (or anti-symmetrized) as the Hilbert space does, i.e.,¹¹ of the form¹²

$$\begin{aligned} A &= A_1 + A_2 + \cdots \\ &= \left(A_1^{(1)} \otimes \mathbb{1} \otimes \cdots + \mathbb{1} \otimes A_1^{(2)} \otimes \cdots + \mathbb{1} \otimes \cdots \otimes A_1^{(N)} \right) \\ &\quad + \left(A_2^{(1)} \otimes A_2^{(2)} \otimes \mathbb{1} \otimes \cdots + A_2^{(1)} \otimes \mathbb{1} \otimes A_2^{(3)} \otimes \mathbb{1} \otimes \cdots + \cdots \right) + \cdots. \end{aligned} \quad (1.5.1)$$

We call the first part of the summation *single-body operator* and the second part *two-body operator*.

例 1(Coulumb Interaction) The Hamiltonian of electrons with Coulumb interaction is simply (without considering influence of lattice)

$$H = H_1 + H_2 = \sum_i^N \frac{p_i^2}{2m} + \frac{1}{2} \sum_{\substack{i,j \\ i \neq j}} \frac{e^2}{4\pi |\mathbf{r}_i - \mathbf{r}_j|}.$$

Although we usual encounter with simple expression as above, **equation (1.5.1) is actually an expansion of interaction including high-order couplings.**

命题 1(Single-body Operator) The single-body operator of A can be expressed as

$$A_1 = \sum_{\alpha, \beta} \langle \varphi_\alpha | A_1^{(i)} | \varphi_\beta \rangle a_\alpha^\dagger a_\beta, \quad (1.5.2)$$

where the superscript i indicates a contribution from the i th single-body¹³ In the continuous form

$$A_1 = \iint d\alpha d\beta \langle \varphi_\alpha | A_1^{(i)} | \varphi_\beta \rangle a_\alpha^\dagger a_\beta.$$

证明 We prove the discrete form, and there is no differences on the continuous one. Utilize the completeness of symmetrized multi-body states $|N\rangle$, we have (focussing on a

¹¹The subscript indicates the number of particles the operator acting on.

¹²Another *wrong* form of multi-body state usually seen in physics literature takes a neat form of

$$A \equiv A_1 + A_2 + \cdots = \sum_{i=1}^N A_1^{(i)} + \frac{1}{2} \sum_{\substack{i,j \\ i \neq j}} A_2^{(i,j)} + \cdots.$$

¹³Since all the k -mode particles of the same energy level are supposed to be identical, contribution from the i th particle has no difference with that of others'.

specific k -mode of N particles)

$$A_1 = \sum_{\substack{\alpha_1 \cdots \alpha_N \\ \beta_1 \cdots \beta_N}} |\varphi_{\alpha_1} \cdots \rangle^{(\varepsilon)(\varepsilon)} \langle \varphi_{\alpha_1} \cdots | A_1 | \varphi_{\beta_1} \cdots \rangle^{(\varepsilon)(\varepsilon)} \langle \varphi_{\beta_1} \cdots | \quad (1.5.3)$$

$$= \frac{1}{N!} \sum_{\substack{\alpha_1 \cdots \alpha_N \\ \beta_1 \cdots \beta_N}} a_{\alpha_1}^\dagger \cdots a_{\alpha_N}^\dagger |0\rangle^{(\varepsilon)} \langle \varphi_{\alpha_1} \cdots | A_1 | \varphi_{\beta_1} \cdots \rangle^{(\varepsilon)} \langle 0 | a_{\beta_N} \cdots a_{\beta_1}. \quad (1.5.4)$$

Here we encounter with an inner product between symmetrized states, which has been shown in (1.4.4) before. Expanding the right hand side of (1.4.4) with the definition of symmetrized (anti-symmetrized) state, one has

$$\begin{aligned} \langle \varphi_{\alpha_1} \cdots | A_1 | \varphi_{\beta_1} \cdots \rangle^{(\varepsilon)} &= \frac{1}{N!} \sum_{\mathcal{P}_\beta} \varepsilon^{p_\beta} \mathcal{P}_\beta \left[\langle \varphi_{\alpha_1}^{(1)} | A_1^{(1)} | \varphi_{\beta_1}^{(1)} \rangle \langle \varphi_{\alpha_2}^{(2)} | \varphi_{\beta_2}^{(2)} \rangle \cdots \langle \varphi_{\alpha_N}^{(N)} | \varphi_{\beta_N}^{(N)} \rangle \right. \\ &\quad \left. + \cdots + \langle \varphi_{\alpha_1}^{(1)} | \varphi_{\beta_1}^{(1)} \rangle \langle \varphi_{\alpha_2}^{(2)} | \varphi_{\beta_2}^{(2)} \rangle \cdots \langle \varphi_{\alpha_N}^{(N)} | A_N^{(N)} | \varphi_{\beta_N}^{(N)} \rangle \right], \end{aligned} \quad (1.5.5)$$

where we make use of the definition of usual multi-body states (as simple tensor products of single-body states) and the fact that single-body operator is actually sum of operators acting on merely one-particle state

$$A_1 = A_1^{(1)} \otimes \mathbb{1} \otimes \cdots + \mathbb{1} \otimes A_1^{(2)} \otimes \cdots + \cdots \otimes \mathbb{1} \otimes A_1^{(N)}.$$

And because both α and β are dummy index under addition in (1.5.5), each term in (1.5.5) should give the same contribution, which can be easily seen by re-labeling them¹⁴. However, because of the existence of permutation operator \mathcal{P}_β acting on all labels β , each transposition (exchange of name of two dummy variables) will contribute an extra factor ε and therefore each term of (1.5.5) will contribute an extra ε^{p_β} overall. Including this extra factor, the coefficient ε^{2p_β} will always be +1 whatever the kind of identical particles we are handling with. So

$$\begin{aligned} A_1 &= \frac{N}{N!} \sum_{\substack{\alpha_1 \cdots \alpha_N \\ \beta_1 \cdots \beta_N}} a_{\alpha_1}^\dagger \cdots a_{\alpha_N}^\dagger |0\rangle \left[\langle \psi_{\alpha_1}^{(1)} | A_1^{(1)} | \psi_{\beta_1}^{(1)} \rangle \delta_{\alpha_2, \beta_2} \cdots \delta_{\alpha_N, \beta_N} \right] \langle 0 | a_{\beta_N} \cdots a_{\beta_1} \\ &= \sum_{\alpha_1, \beta_1} \langle \psi_{\alpha_1}^{(1)} | A_1^{(1)} | \psi_{\beta_1}^{(1)} \rangle a_{\alpha_1}^\dagger a_{\beta_1} \left[\frac{1}{(N-1)!} \sum_{\substack{\alpha_2 \cdots \alpha_N \\ \beta_2 \cdots \beta_N}} a_{\alpha_2}^\dagger \cdots a_{\alpha_N}^\dagger |0\rangle \langle 0 | a_{\alpha_2} \cdots a_{\alpha_N} \right] \\ &= \sum_{\alpha, \beta} \langle \varphi_\alpha | A_1^{(1)} | \varphi_\beta \rangle a_\alpha^\dagger a_\beta^\dagger = \sum_{\alpha, \beta} \langle \varphi_\alpha | A_1^{(i)} | \varphi_\beta \rangle a_\alpha^\dagger a_\beta^\dagger, \end{aligned}$$

where the square bracket in the second line is exactly the completeness relation of $\mathcal{H}_{N-1}^{(\varepsilon)}$. \square

¹⁴ The operation of renaming variables will certainly also affect the order of creation and annihilation operators outside the inner product of symmetrized (anti-symmetrized) states, but by commutation relation there has no contribution.

命题 2 The two-body operator of A can be expressed as¹⁵

$$A_2 \equiv \frac{1}{2} \sum_{\substack{i,j \\ i \neq j}} A_2^{(i,j)} = \frac{1}{2} \sum_{\alpha, \beta, \gamma, \delta} \langle \varphi_\alpha \varphi_\beta | A_2^{(i,j)} | \varphi_\gamma \varphi_\delta \rangle a_\alpha^\dagger a_\beta^\dagger a_\delta a_\gamma. \quad (1.5.6)$$

where (i, j) is an arbitrary pair and $|\varphi_\alpha \varphi_\beta\rangle$ is the usual two-body state $|\varphi_\alpha\rangle \otimes |\varphi_\beta\rangle$, or in the continuous form

$$A_2 = \frac{1}{2} \iiint d\alpha d\beta d\gamma d\delta \langle \varphi_\alpha \varphi_\beta | A_2^{(i,j)} | \varphi_\gamma \varphi_\delta \rangle a_\alpha^\dagger a_\beta^\dagger a_\delta a_\gamma.$$

证明 Similarly. Leave as an exercise. \square

Both (1.5.2) and (1.5.6) involve in merely abstract states in $\mathcal{H}_N^{(\varepsilon)}$, but we always choose some specific representations in real calculation.

推论 1 In the ξ representation (for example, coordinate representation or momentum representation), we can also write the matrix element in (1.5.2) and (1.5.6) as¹⁶

$$A_1 = \sum_{\alpha\beta} \int d\xi \varphi_\alpha^*(\xi) A_1^{(i)}(\xi) \varphi_\beta(\xi) a_\alpha^\dagger a_\beta, \quad (1.5.7)$$

$$A_2 = \frac{1}{2} \sum_{\alpha\beta\gamma\delta} \iint d\xi_1 d\xi_2 \varphi_\alpha^*(\xi_1) \varphi_\beta^*(\xi_2) A_2^{(i,j)}(\xi_1, \xi_2) \varphi_\gamma(\xi_2) \varphi_\delta(\xi_1) a_\alpha^\dagger a_\beta^\dagger a_\gamma a_\delta, \quad (1.5.8)$$

where $A_1^{(i)}(\xi) \equiv \langle \xi | A_1^{(i)} | \xi \rangle$ and $A_2^{(i,j)}(\xi_1, \xi_2) \equiv \langle \xi_1, \xi_2 | A_2^{(i,j)} | \xi_2, \xi_1 \rangle$

证明 We still work in discrete form in consistency with what we have done above, that is, we have a discrete spectrum decomposition of the identity operator in the i th single-body Hilbert space $\mathcal{H}_1^{(i)}$

$$\sum_{\xi} |\xi^{(i)}\rangle \langle \xi^{(i)}| = \mathbb{1}.$$

We first prove the single-body case. One should note that only when $A_1^{(i)}$ acts on states in $\mathcal{H}_1^{(i)}$ that it gives out a non-vanishing value, i.e.,

$$A_1^{(i)} |\xi^{(j)}\rangle = \delta_{ij} A_1^{(i)} |\xi^{(i)}\rangle. \quad (1.5.9)$$

So

$$A_1 = \sum_{\xi^{(i)}} \sum_{\xi^{(j)}} \langle \varphi_\alpha | \xi^{(i)} \rangle \langle \xi^{(i)} | A_1^{(i)} | \xi^{(j)} \rangle \langle \xi^{(j)} | \varphi_\beta \rangle a_\alpha^\dagger a_\beta$$

¹⁵Here we abuse the *wrong notation* for simplicity, but one should always be aware of the correct form (1.5.1).

¹⁶In usual literature we neglect the superscript in the inner product since they all have the same contribution form each particle for single-body operators or each pair for two-body operators. But in order not to cause ambiguities, I maintain them in our expressions.

$$= \sum_{\xi^{(i)}} \varphi_{\alpha}^*(\xi^{(i)}) A_1^{(i)}(\xi^{(i)}) \varphi_{\beta}(\xi^{(i)}) a_{\alpha}^{\dagger} a_{\beta}.$$

There is no difference in the proof of two-body case. \square

例 2(General Interactive Hamiltonian for Schrödinger Field) Consider a general¹⁷ multi-body system with high order couplings¹⁸

$$H = H_1 + H_2 + \cdots = \sum_j \frac{1}{2m} \hat{\mathbf{p}}_i^2 + \frac{1}{2} \sum_{ij}' U(\hat{\mathbf{r}}_i, \hat{\mathbf{r}}_j) + \cdots.$$

Then by rules we deduced before,

$$\begin{aligned} H_1 &= \sum_{\alpha\beta} \frac{1}{2m} \langle \varphi_{\alpha} | \hat{\mathbf{p}}_i^2 | \varphi_{\beta} \rangle a_{\alpha}^{\dagger} a_{\beta} = -\frac{\hbar^2}{2m} \sum_{\alpha\beta} \int d\mathbf{r} \varphi_{\alpha}(\mathbf{r}) \nabla^2 \varphi_{\beta}(\mathbf{r}) a_{\alpha}^{\dagger} a_{\beta} \\ &= -\frac{\hbar^2}{2m} \int d\mathbf{r} \nabla \psi^{\dagger}(\mathbf{r}) \cdot \nabla \psi(\mathbf{r}), \end{aligned}$$

where in the second line we integrate by parts and substitute the definition of fields operators.

Likewise, for the two-body part,

$$\begin{aligned} H_2 &= \frac{1}{2} \sum_{\alpha\beta\mu\nu} \langle \varphi_{\alpha}^{(j)} \varphi_{\beta}^{(i)} | U(\hat{\mathbf{r}}_i, \hat{\mathbf{r}}_j) | \varphi_{\mu}^{(i)} \varphi_{\nu}^{(j)} \rangle a_{\alpha}^{\dagger} a_{\beta}^{\dagger} a_{\mu} a_{\nu} \\ &= \sum_{\alpha\beta\mu\nu} \int d\mathbf{r} d\mathbf{r}' \varphi_{\alpha}^*(\mathbf{r}) \varphi_{\beta}^*(\mathbf{r}') U(\mathbf{r}, \mathbf{r}') \varphi_{\mu}(\mathbf{r}') \varphi_{\nu}(\mathbf{r}) a_{\alpha}^{\dagger} a_{\beta}^{\dagger} a_{\mu} a_{\nu} \\ &= \frac{1}{2} \iint d\mathbf{r} d\mathbf{r}' \psi^{\dagger}(\mathbf{r}) \psi^{\dagger}(\mathbf{r}') U(\mathbf{r}, \mathbf{r}') \psi(\mathbf{r}') \psi(\mathbf{r}). \end{aligned}$$

¹⁷For simplicity we ignore the couplings between spins labels and extra fields here.

¹⁸Again we abuse the wrong notation here.

2

Methods on $T = 0$ Quantum Field Theory

2.1 Grand Canonical Ensemble

In CMP, we would like to handle with the systems with a given chemistry potential rather than a stable number of particles to cater for most of the experiments. Thus we replace our original Hamiltonian with

$$H' = H - \mu N,$$

and modify our definition of fields operators in Heisenberg pictures that

$$\psi_{\alpha}^H(t, \mathbf{r}) = e^{iH't} \psi_{\alpha}^S e^{-iH't}. \quad (2.1.1)$$

We need to check that after applying this change we will obtaining a similar form of (1.3.3). In fact, starting with the “box-normalized” non-interactive field operator in Schödingre picture, we can instantly prove it in usage of (2.1.1)¹:

$$\begin{aligned} \psi^H(\mathbf{r}, t) &= \frac{1}{\sqrt{V}} \left(\sum_{\mathbf{p}} e^{i \sum_{\mathbf{p}'} (\varepsilon_{\mathbf{p}'} - \mu) \hat{n}_{\mathbf{p}'}} \cdot a_{\mathbf{p}} e^{i\mathbf{p} \cdot \mathbf{r}} \cdot e^{i \sum_{\mathbf{p}''} (\varepsilon_{\mathbf{p}''} - \mu) \hat{n}_{\mathbf{p}''}} \right) e^{-i\varepsilon_{\mathbf{p}} t} \\ &= \frac{1}{\sqrt{V}} \sum_{\mathbf{p}} a_{\mathbf{p}} e^{i[\mathbf{p} \cdot \mathbf{r} - (\varepsilon_{\mathbf{p}} - \mu)t]}, \end{aligned} \quad (2.1.2)$$

where in the second line we use the B-C-H formula.

By the probability interpretation of QM, only statistical quantities make sense. But how to calculate the *ensemble average* of the system may bewilder you. Now that

¹Recall that in free theory, the Hamiltonian commutes with the Number operator so both can be diagonalized under $\hat{n}_{\mathbf{p}} \equiv a_{\mathbf{p}}^{\dagger} a_{\mathbf{p}}$.

we allow changes of number of particles in our system, it is natural to combine the theory of *grand canonical ensemble* of quantum statistical physics in our theory:

定义 1 (Density Matrix) The density matrix of the grand canonical ensemble is

$$\rho := \frac{1}{\Xi} e^{-\beta H'}, \quad (2.1.3)$$

where the *grand partition function* is $\Xi \equiv \text{tr}(e^{-\beta H'}) = \text{tr}(e^{-\beta(H-\mu N)})$. And then the average of any operator A is defined by

$$\langle A \rangle := \text{tr}(\rho A).$$

注 1 In the theory of quantum statistical physics, the trace runs all the possible micro-state of our system, i.e.

$$\text{tr}(\rho A) \equiv \sum_s \langle s | \rho A | s \rangle \quad (2.1.4)$$

for any arbitrary set of complete basis of our multi-particle Hilbert space.

2.2 Green Function

Before I put forward the definition of Green function, let us pause for some thoughts on experiments in CMP. Anything encountered for the first time for an pristine human being during the expedition are dangerous black boxes, and to probe the properties of such unknown matters, a natural gesture occurring in his mind may be touching them and feeling their shape and texture. Similarly, for physicists who believe in objective things such as electrons and protons much harder than their bodies and brains, what they first do on the unknown matter is also to **impose an simple influence² on it and observe what the matter will reflect to it**. For example, in solid state physics, we already knew that physicists use the X-ray diffraction to determine the lattice structure of crystals and infrared absorption to study the optic branch of vibration of solids.

In a word, any experiments in CMP, as long as involving in test particles or responding particles, can be categorized to *probing measure*, which is apparently divided into two kinds: **one is to inject a test particle in the sample and measure the responding particles after interactions, or in the language of QFT, propagation³, another one is to take away a particle at first⁴, and then observing the propagation of the “hole”.**

Recall the fields operator and their dual operators have exactly the same effects on multi-particle states, creation and annihilation (otherwise we would not waste time abruptly

²By “simple”, we mean something physicists believe they are familiar with, for example, the linear response of current under external electric fields.

³In some cases, electromagnetic fields will be released from solids and captured by experimental apparatus, but in most cases, the responding particles are collective behavior and quasi-ones.

⁴For example, take away some electrons from the valence band through doping a semiconductor, and then observing the propagation of hole.

swerving our topic to experiments 2333). So it's natural for us to propose the definition of Green function in consideration of these two ways of experiments:

定义 1(Green Function) The single-particle *Green function* is defined as⁵

$$iG_{\alpha\beta}(x, x') = \langle T(\psi_{\alpha}^H(x) \psi_{\beta}^{\dagger H}(x')) \rangle, \quad (2.2.1)$$

where the average takes the meaning of quantum statistical ensembles, the subscript indicate the components of spins of quasi-particles, and T is the time-order operator such that⁶

$$T\psi(x)\psi(y) = \begin{cases} \psi(x)\psi(y), & x^0 > y^0, \\ \varepsilon\psi(y)\psi(x), & y^0 > x^0, \end{cases}$$

where $\varepsilon = \pm 1$ for bosons and fermions.

注 1 The coefficient “ i ” is just a convention in order to make it concord with the Schrödinger equation of no interaction case, as is shown latter, because **in the sense of mathematics, “Green function” should subject to the same differential equation as general solution does, whereas this is true only when we are handling with the free systems in CMP.** We will see this in the near future.

注 2(Subtleties in spins of quasi-particles) In condensed matter physics, the structure of symmetries of space-time, Lorentz group, is removed, and so is the representation of the universal cover of it. In the construction of relativistic quantum fields theory, spins is defined by the the highest weights of the representation of the little group of such universal cover group. Although the structure of group is altered while little authors regard this problem as importance, I believe⁷ **spins can also be obtained form the representation of Galileo group, the central extension of the Lie algebra of which gives out a similar distribution of weights as Lorent group does.**

However, a much more serious question is that **the spin-statistical theorem also breaks in CMP**, hence even if we endow quasi-particles with spins from the representaion of Galileo group, we acturally still **cannot determine what kind of commuation relation their corresponding fields operators subject to**, let alone the complicated anyons. As far as I am concerned, almost all physicists frivolously acquiesce (or never notice) the trueness of spin-statistical thoerem in CMP. In order not to swerve from our mainline, I still have to **suppose the correctness of spin-statistical theorem.** I point out the logical flaws here just to emphasize the inconspicuous ambiguity of the construction of non-relativistic QFT.

In fact, both of the two questions holds only because we take the view of *extreme anti-reductionism*, or *emergence* as Wen does. If one acknowledges that particles in high energy

⁵I will not add the hat on fields operators unless they bring in ambiguity.

⁶Note that **the function of time-order operator is just rearranging the order of operators but never brings in a minus sign in the expression**, which throughly result from the classification of particles.

⁷Lifshiz and Pitaevskii seem to interpret the problem of spin by degeneracy of energy level, but I disagree (and do not understand 2333).

physics are indeed “basic particles”, and collectively exited particles in CMP are composition of them, then all the hinderance above will be cleared. In a word, **whether spins of quasi-particle is a problem is determined by which philosophy—emergence or reductionism—one believes.**

Up to now we have not utilize the condition of *zero-temperature* in the caption of this chapter at all, and let’s see what will it bring to our form of Green functions:

命题 1($T = 0$ Green Function) At $T = 0$ k, the Green function can be computed under the average of *ground state*⁸, i.e.,

$$iG_{\alpha\beta}(x, x') = \langle \Omega | T \psi_{\alpha}^H(x) \psi_{\beta}^{\dagger H}(x') | \Omega \rangle. \quad (2.2.2)$$

证明 Decompose the identity operator of our system by some arbitrary complete set of states (not have to be the symmetrized or anti-symmetrized multiparticle states we introduced in chapter one)

$$\mathbb{1} = \sum_t |t\rangle \langle t|,$$

then by definition of statistical average, we have

$$\begin{aligned} \langle T \psi(x) \psi^{\dagger}(x') \rangle &= \text{tr} (\rho T \psi(x) \psi^{\dagger}(x')) \\ &= \frac{\sum_{s,t} \langle s | e^{-\beta(H-\mu N)} | t \rangle \langle t | T \psi(x) \psi^{\dagger}(x') | s \rangle}{\sum_s \langle s | e^{-\beta(H-\mu N)} | s \rangle} \\ &= \sum_n e^{-\beta(E_s - \mu N_s)} \langle s | T \psi(x) \psi^{\dagger}(x') | s \rangle \bigg/ \sum_s e^{-\beta(E_s - \mu N_s)}, \end{aligned}$$

where in the third line we use the notation E_s and N_s to denote the energy and number of particles of the state $|s\rangle$ and the orthogonal relation of states $|s\rangle$ and $|t\rangle$.

Although we have no idea of the real number E_s and N_s due to the arbitrary choice of complete set of states, we can pick out the unique state $|\Omega\rangle$ with lowest energy by the fundamental **non-degeneracy hypothesis**⁹. Denoting its energy $E'_{\Omega} \equiv E_{\Omega} - \mu N_{\Omega}$, we then have

$$E'_s|_{s \neq \Omega} > E'_{\Omega}.$$

⁸I use $|\Omega\rangle$ just to make it in concord with the notation in high energy physics.

⁹The non-degeneracy hypothesis says that **the ground state of any energy spectrum is unique**. A similar statement in Wightman’s axiomatic construction of quantum field theory also says that **the vacuum state is always non-degenerate**.

So we get

$$iG(t_1, t_2) = \lim_{\beta \rightarrow \infty} \frac{\langle \Omega | T\psi(x)\psi^\dagger(x') | \Omega \rangle + \sum_{s \neq \Omega} e^{-\beta(E'_s - E'_\Omega)} \langle s | T\psi(x)\psi^\dagger(x') | s \rangle + \dots}{1 + \sum_{s \neq \Omega} e^{-\beta(E'_s - E'_\Omega)} + \dots}$$

$$= \langle \Omega | T\psi(x)\psi^\dagger(x') | \Omega \rangle,$$

where all the other terms are *exponentially suppressed* because of the inequality we displayed above. \square

注 3 The method used here is similar to the powerful *zero-mode* in the *heat kernel theory* in mathematical physics, which is tightly related to the *Atiyah-Singer index theorem*.

注 4 Some additional comparison of the ground state $|\Omega\rangle$ in CMP and QED is necessary here:

In CMP, owing to the effects of statistical physics¹⁰, electrons in non-interacting fermion systems have the distribution of *fermi ball* in the momentum space, which is disparate with situations in high energy physics. Thus the terminology “vacuum state” has entirely different meaning in these two paradigms. This difference is apparent under action of creation and annihilation operators:

$$\begin{cases} a_{\mathbf{p}}|\Omega\rangle = 0, & (\varepsilon > \varepsilon_F) \\ a_{\mathbf{p}}^\dagger|\Omega\rangle = 0. & (\varepsilon < \varepsilon_F) \end{cases}$$

Clearly if our system is not ferro-magnetic or put in magnetic fields, Green function must be independent of spins, i.e.,

$$G_{\alpha\beta}(x, x') = \delta_{\alpha\beta} G(x, x').$$

Furthermore, since our physical world for condensed matters physics is *homogeneous* on time and space¹¹, the Green function can only depends on differences between time and spaces

断语 1 Generally, we have

$$G(t_1, \mathbf{r}_1; t_2, \mathbf{r}_2) = G(t; \mathbf{r}), \quad (2.2.3)$$

where $t \equiv t_1 - t_2$, $\mathbf{r} = \mathbf{r}_1 - \mathbf{r}_2$.

But, when it comes to the *isotropy* of time and space, since most of detective approaches in CMP, for example, the X-ray diffraction, which involves in *chronological* scattering proceeds,

¹⁰One should notice that although QED has more structure than CMP in the sense of Lorentz invariance, the theoretical framework of the latter one still own some original features because **CMP is actually the combination of (perturbative) field theory and the theory of Gibbs statistics**. And it is the additional characteristic that makes it valuable to learn.

¹¹For simplicity we start with studying the systems without disorders and defects.

clearly time has not to be invertible. On the other hand, compared with gases and liquids, common systems like solid crystals are sure not of isotropic, so we can only claim that **only in systems with such so-called *microscopic homogeneity*¹² the Green functions satisfy $G(t, \mathbf{r}) = G(t, -\mathbf{r})$. But we can *never* friviously deem that $G(t, \mathbf{r}) = G(-t, \mathbf{t})$.** From this aspect, we must attach importance on the conventional order of the time difference in Green function $t \equiv t_1 - t_2$.

2.3 Free Green Function

2.3.1 Quantum Statistics of Grand Canonical Ensemble

In this section, we review the derivation of distribution function for both bosons and fermions in the frame of multi-particle states and quantum statistical physics.

引理 1(Partition Function) The partition function of the grand canonical ensemble takes the form of

$$\Xi_{\mp} = \prod_{\ell} (1 \mp e^{-\beta(\varepsilon_{\ell} - \mu)})^{\mp 1}, \quad (2.3.1)$$

where minus sign for bosons and possitive sign for fermions.

证明 Recall that the symmetrized or anti-symmetrized multi-particle state for many energy levels lives in tensor product of Fock spaces:

$$|N\rangle \in \bigotimes_k \bigoplus_{n_k} \mathcal{H}_{n_k}^{(\varepsilon)},$$

so does the many-body operators. Hence

$$\begin{aligned} \Xi \equiv \text{tr}(e^{-\beta H'}) &= \text{tr}\left(e^{-\beta \sum_{\ell} H'_{\ell}}\right) = \text{tr}\left(\bigotimes_{\ell} e^{-\beta(\varepsilon_{\ell} - \mu) a_{\ell}^{\dagger} a_{\ell}}\right) \\ &= \prod_{\ell} \text{tr}_{\ell}\left(e^{-\beta(\varepsilon_{\ell} - \mu) a_{\ell}^{\dagger} a_{\ell}}\right) = \prod_{\ell} \sum_{n_{\ell}} \langle n_{\ell} | e^{-\beta(\varepsilon_{\ell} - \mu) a_{\ell}^{\dagger} a_{\ell}} | n_{\ell} \rangle \\ &= \begin{cases} \prod_{\ell} \sum_{n_{\ell}=0}^{\infty} e^{-\beta(\varepsilon_{\ell} - \mu) n_{\ell}}, & \text{Bosons,} \\ \prod_{\ell} \sum_{n_{\ell}=0}^1 e^{-\beta(\varepsilon_{\ell} - \mu) n_{\ell}}, & \text{Fermions.} \end{cases} \end{aligned}$$

and we obtain the familiar form in (2.3.1). □

Now we are able to get the distribution for both kinds:

¹²By *microscopic homogeneity*, we mean not only the system has a homogeneous average density macroscopically but it also gives particles in it a homogeneous distribution of probabilities on different spacial positions.

命题 1(Distribution Function) The distribution function for bosons is

$$n_B(\mathbf{k}) = \frac{1}{e^{\beta(\varepsilon_{\mathbf{k}} - \mu)} - 1}, \quad (2.3.2)$$

while for fermions is

$$n_F(\mathbf{k}) = \frac{1}{e^{\beta(\varepsilon_{\mathbf{k}} - \mu)} + 1}. \quad (2.3.3)$$

证明

$$\begin{aligned} \langle a_k^\dagger a_k \rangle &\equiv \text{tr}(\rho a_k^\dagger a_k) = \text{tr} \left(\frac{1}{\Xi} \bigotimes_{\ell} e^{-\beta(\varepsilon_{\ell} - \mu) a_{\ell}^\dagger a_{\ell}} \cdot a_k^\dagger a_k \right) \\ &= \frac{1}{\Xi} \text{tr}_k \left(e^{-\beta(\varepsilon_k - \mu) a_k^\dagger a_k} a_k^\dagger a_k \right) \text{tr}_{\ell \neq k} \left(\bigotimes_{\ell \neq k} e^{-\beta(\varepsilon_{\ell} - \mu) a_{\ell}^\dagger a_{\ell}} \right) \\ &= \text{tr}_k \left(e^{-\beta(\varepsilon_k - \mu) a_k^\dagger a_k} a_k^\dagger a_k \right) \cdot \left(\prod_{\ell \neq k} \text{tr}_{\ell} \left(e^{-\beta(\varepsilon_{\ell} - \mu) a_{\ell}^\dagger a_{\ell}} \right) \right) / \prod_{\ell} \text{tr}_{\ell} \left(e^{-\beta(\varepsilon_{\ell} - \mu) a_{\ell}^\dagger a_{\ell}} \right) \\ &= -\text{tr}_k \left(\frac{\partial}{\partial(\beta(\varepsilon_k - \mu))} e^{-\beta(\varepsilon_k - \mu) a_k^\dagger a_k} \right) / \text{tr}_k \left(e^{-\beta(\varepsilon_k - \mu) a_k^\dagger a_k} \right) \\ &= -\frac{\partial}{\partial(\beta(\varepsilon_k - \mu))} \text{tr}_k \left(e^{-\beta(\varepsilon_k - \mu) a_k^\dagger a_k} \right) \cdot \text{tr}_k \left(e^{-\beta(\varepsilon_k - \mu) a_k^\dagger a_k} \right) \\ &= -\frac{\partial}{\partial(\beta(\varepsilon_k - \mu))} \ln \left(\text{tr}_k \left(e^{-\beta(\varepsilon_k - \mu) a_k^\dagger a_k} \right) \right) \\ &= -\frac{\partial}{\partial(\beta(\varepsilon_k - \mu))} \ln \left((1 \mp e^{-\beta(\varepsilon_k - \mu) a_k^\dagger a_k}) \right) = \frac{1}{e^{\beta(\varepsilon_k - \mu)} \mp 1} \end{aligned}$$

□

2.3.2 Fermion Fields

At the low temperature limit $\beta \rightarrow \infty$, distribution function reduce to the simple form

$$n_F(\varepsilon_{\mathbf{k}}) = \theta(\mu - \varepsilon_{\mathbf{k}}), \quad (2.3.4)$$

as is illustrated in figure 2.1:

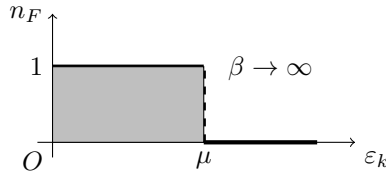


Figure 2.1: Strong-Degenerate free fermion gases

Before putting forward the specific results of Green functions, it is worthy to introduce one generalize function before:

引理 2(Sokhotsky-Weiestrass Formula) Given an integrable function $\varphi(x)$ (with a compact support), we can introduce a linear functional on K space¹³, by

$$\left\langle \varphi(x), \lim_{\delta \rightarrow 0} \frac{1}{x \pm i\delta} \right\rangle \equiv \lim_{\delta \rightarrow \infty} \int dx \varphi(x) \frac{1}{x \pm i\delta} = \mathcal{P} \int dx \frac{\varphi(x)}{x} \mp i\pi\varphi(0), \quad (2.3.5)$$

i.e.,

$$\frac{1}{x \pm i\delta} = \mathcal{P} \frac{1}{x} \mp i\pi\delta(x),$$

where \mathcal{P} is the *Cauchy principle value*.

证明 Out of the main line of this note, so we just skip it. \square

Sometimes we'd like to write $\frac{1}{x \pm i0} \equiv \lim_{\delta \rightarrow 0} \frac{1}{x \pm i\delta}$, then a more widely used form is to re-express such functional by an integral of an exponential function with a slight real translation (this shift is necessary to make it converge)

$$\lim_{\delta \rightarrow +0} \int_0^\infty e^{ist - \delta t} dt = \frac{e^{ist - \delta t}}{s - \delta} \Big|_{t=0}^{t=\infty} = \lim_{\delta \rightarrow +0} \frac{i}{s + i\delta}. \quad (2.3.6)$$

注 1 Note that (2.3.6) holds for all real number s , no matter s is possitive or negative.

Free fields for electrons has be obtained in (2.1.3), so now let us calculate the free Green function

命题 2 The Green funtion of non-interactive fermion gases $G^{(0)}$ in mommentum space is

$$G^{(0)}(\omega, \mathbf{p}) = \frac{1}{\omega + \mu - \varepsilon_{\mathbf{p}} + i\delta \text{sgn}(|\mathbf{p} - \mathbf{p}_F|)}, \quad (2.3.7)$$

where the *Fermi momentum* $p_F \equiv \sqrt{2m\mu}$.

证明 Directly substitute the free fermion fields operator, we have (note that the ground state for non-interacting fermions is the *fermi ball*, where for $k < k_F$ there is no space for more electrons to reside and for $k > k_F$ there is no electron to be annihilated) 2

$$\begin{aligned} iG^{(0)}(\mathbf{r}; t) &= \theta(t) \langle |\Omega| \psi^H \psi^{\dagger H} |\Omega\rangle - \theta(-t) \langle \Omega | \psi^{\dagger H} \psi^H |\Omega\rangle \\ &= \frac{1}{V} \sum_{\mathbf{k}} \left(\theta(t) \langle \Omega | (1 - a_{\mathbf{k}}^\dagger a_{\mathbf{k}}) |\Omega\rangle - \theta(-t) \langle \Omega | a_{\mathbf{k}}^\dagger a_{\mathbf{k}} |\Omega\rangle \right) e^{i[\mathbf{k} \cdot \mathbf{x} - (\varepsilon_{\mathbf{k}} - \mu)t]} \\ &= \frac{1}{V} \sum_{\mathbf{k}} \left(\theta(t) (1 - \theta(\mu - \varepsilon_{\mathbf{k}})) - \theta(-t) \theta(\mu - \varepsilon_{\mathbf{k}}) \right) e^{i[\mathbf{k} \cdot \mathbf{x} - (\varepsilon_{\mathbf{k}} - \mu)t]}, \end{aligned}$$

¹³Some Addition on definition of *basic function space* K . We say the linear combinations of complex functions with bounded support and infinit derivatives forms a space of K if the limit (topology) is defined as follows:

- 1) given a sequence $\{\varphi_m\} \subset K$, there exists a bounded set A such that $\text{supp}(\varphi_m) \subset A$.
- 2) all order of derivatives of $\{\varphi_m\}$, denoted as $\{D^p \varphi_m\}$ are uniformly convergent on $D^p \varphi$, written as $\varphi_m \xrightarrow{K} \varphi$.

where in the second line we have utilized the anticommutative relation for fermionic operators and substituted the distribution of fermions at $T = 0$ K. Next, carefully performing the Fourier transformation, we get

$$\begin{aligned} iG^{(0)}(\mathbf{p}, \omega) &\equiv \int G^{(0)}(\mathbf{r}, t) e^{-i(\mathbf{p} \cdot \mathbf{r} - \omega t)} dt d\mathbf{r} \\ &= \begin{cases} \mu - \varepsilon_{\mathbf{p}} > 0, & \frac{-1}{i} \int \theta(-t) e^{i(\omega - \varepsilon_{\mathbf{p}} + \mu)t} dt = \frac{-1}{i} \int_{-\infty}^0 e^{i(\omega + \mu - \varepsilon_{\mathbf{p}})t} dt, \\ \mu - \varepsilon_{\mathbf{p}} < 0, & \frac{1}{i} \int \theta(t) e^{i(\omega - \varepsilon_{\mathbf{p}} + \mu)t} dt = \frac{1}{i} \int_0^{\infty} e^{i(\omega + \mu - \varepsilon_{\mathbf{p}})t} dt. \end{cases} \end{aligned}$$

One can see that both of the cases can be summarized to an integrand $\int e^{ikt} dt$, which is certainly *divergent*. However, **what outweighs is not the concrete form of the Green function, but the specific indication of the contour you choose in convention when multiplying it on some physically integrable functions in the computation of some feynmann diagrams**, as is seen in the future. In other words, in CMP, similar to the Dirac delta function, Green function is also defined in the sense of integrals and thus a *generalized functions*.

Thus we intentionally add a slight real translation on each term and use (2.3.6), getting¹⁴

$$\begin{aligned} G^{(0)}(\mathbf{p}, \omega) &= \frac{1}{i} \lim_{\delta \rightarrow +0} \int_0^{\infty} dt \left[-\theta(\mu - \varepsilon_{\mathbf{p}}) e^{-i(\omega + \mu - \varepsilon_{\mathbf{p}})t + \delta t} + \theta(\varepsilon_{\mathbf{p}} - \mu) e^{i(\omega + \mu - \varepsilon_{\mathbf{p}})t - \delta t} \right] \\ &= \frac{\theta(|\mathbf{p} - \mathbf{p}_F|)}{\omega + \mu - \varepsilon_{\mathbf{k}} + i0} + \frac{\theta(\mathbf{p}_F - |\mathbf{p}|)}{\omega + \mu - \varepsilon_{\mathbf{k}} - i0}, \end{aligned} \quad (2.3.8)$$

where we have rewritten the theta function $\theta(\mu - \varepsilon_{\mathbf{p}})$ by momentum, and it's easy to check that the above results is equivalent to (2.3.7). \square

Recall that we have mentioned above that the non-interactive Green function in physics concord with that in the sense of mathematics, that is,

推论 1 $G^{(0)}(\mathbf{r}, t)$ we obtained here satisfies the Green equation in mathematical sense that

$$LG(x, s) = \delta(s - x),$$

where the linear differential operator L is exactly the one in Schrödinger equation of fields operators without both external interaction and high order self-interaction terms, i.e.,

$$\left(i \frac{\partial}{\partial t} + \frac{\Delta}{2m} + \mu \right) G^{(0)}(t, \mathbf{r}) = \delta(t) \delta(\mathbf{r}). \quad (2.3.9)$$

¹⁴Because (2.4.2) holds for all real s , so for the $\mu > \varepsilon_{\mathbf{p}}$ term, one just need to perform the variable transformation $t \rightarrow -u$ and absorb the extra minus sign in the bracket, i.e.,

$$\frac{-1}{i} \int_{-\infty}^0 e^{i(\omega + \mu - \varepsilon_{\mathbf{p}})t} dt = \frac{-1}{i} \int_0^{+\infty} e^{i(\varepsilon_{\mathbf{p}} - \omega - \mu)u} du = \lim_{u \rightarrow 0^+} \frac{-i}{\varepsilon_{\mathbf{p}} - \omega - \mu + i\delta}.$$

证明 Rewrite the definition of Green function (2.2.1) in an explicit form

$$\frac{\partial}{\partial t} G_{\alpha\beta} = -i \frac{\partial}{\partial t} \left[\langle \psi_\alpha(x_1) \psi_\beta^\dagger(x_2) \theta(t) - \langle \psi_\beta^\dagger(x_2) \psi_\alpha(x_1) \rangle \theta(-t) \right],$$

and utilize the anti-commutation relation for fermions and the formula $\partial_t \theta(t) \equiv \delta(t)$, we get

$$\frac{\partial}{\partial t} G_{\alpha\beta} = -i \left\langle T \frac{\partial \psi_\alpha(x_1)}{\partial t_1} \psi_\beta^\dagger(x_2) \right\rangle - i \delta_{\alpha\beta} \delta(\mathbf{r}) \delta(t).$$

In addition, the fields operators in Heisenberg picture for our free system satisfies

$$i \frac{\partial \psi_\alpha}{\partial t} = -\frac{1}{2m} \Delta \psi_\alpha - \mu \psi_\alpha, \quad (2.3.10)$$

which can be derived from our Hamiltonian (??) without both external and self interaction.

Substitute this in the above expression, we find the equation of free Green functions:

$$\left(i \frac{\partial}{\partial t} + \frac{\Delta}{2m} + \mu \right) G^{(0)}(t, \mathbf{r}) = \delta(t) \delta(\mathbf{r}).$$

Impose Fourier transformation on it giving

$$\left(\omega - \frac{\mathbf{p}^2}{2m} + \mu \right) G^{(0)}(\omega, \mathbf{p}) = 1, \quad (2.3.11)$$

or the former explicit form

$$G^{(0)}(\omega, \mathbf{p}) = \frac{1}{\omega - \varepsilon_{\mathbf{p}} + \mu + i0 \text{sgn} \omega},$$

where $\varepsilon_{\mathbf{p}} = \mathbf{p}^2/2m$ and we insert the important infinitesimally imaginary term to make it consistent with our former discussion about the choices of integral contour when counting the feynman diagrams (the general analytic properties of Green function will impose such a demand on the imaginary part). \square

2.3.3 Phonon Fields

We start with introducing what is the phonon fields:

Different from the fermion fields operators, **it is because the phonon fields describe the real displacement of lattice that it must be real field operators, or Hermitian operator.** So regard the solid as continuum of springs and follow the same procedure of canonical quantization as Peksin, we have the familiar

定义 1(Phonon Field Operator) The *phonon field operator* (in Heisenberg picture) is

$$\varphi^H(\mathbf{r}, t) = \frac{1}{\sqrt{V}} \sum_{\mathbf{k}} \sqrt{\frac{\omega_{\mathbf{p}}}{2}} \left(b_{\mathbf{k}} e^{i(\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t)} + b_{\mathbf{k}}^\dagger e^{-i(\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t)} \right). \quad (2.3.12)$$

Through the similar process, we can obtain

命题 3 The unperturbed Green function for phonons is

$$D^{(0)}(\mathbf{k}, \omega) = \frac{\omega_{\mathbf{k}}^2}{\omega^2 - \omega_{\mathbf{k}}^2 + i0}. \quad (2.3.13)$$

证明

$$\begin{aligned} D^{(0)}(\mathbf{r}, t) &= \frac{1}{i} \left(\langle \Omega | \theta(t) \varphi^H \varphi^{\dagger H} | \Omega \rangle + \langle \Omega | \theta(-t) \varphi^{\dagger H} \varphi^H | \Omega \rangle \right) \\ &= \frac{1}{iV} \sum_{\mathbf{k}, l} \frac{\sqrt{\omega_{\mathbf{k}} \omega_l}}{2} \left[\theta(t) \langle \Omega | \left(b_{\mathbf{k}} + b_{-\mathbf{k}}^{\dagger} \right) e^{i(\mathbf{k} \cdot \mathbf{r}_1 - \omega_{\mathbf{k}} t_1)} \left(b_l + b_{-l}^{\dagger} \right) e^{i(l \cdot \mathbf{r}_2 - \omega_l t_2)} | \Omega \rangle \right. \\ &\quad \left. + \theta(-t) \langle \Omega | \left(b_l + b_{-l}^{\dagger} \right) e^{i(l \cdot \mathbf{r}_1 - \omega_l t_1)} \left(b_{\mathbf{k}} + b_{-\mathbf{k}}^{\dagger} \right) e^{i(\mathbf{k} \cdot \mathbf{r}_2 - \omega_{\mathbf{k}} t_2)} | \Omega \rangle \right] \\ &= \frac{1}{iV} \sum_{\mathbf{k}} \frac{\omega_{\mathbf{k}}}{2} \left(\theta(t) e^{i(\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t)} + \theta(-t) e^{-i(\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t)} \right), \end{aligned}$$

where in the second we utilize the technique introduced in Peskin to shorten the expression

$$\frac{1}{\sqrt{V}} \sum_{\mathbf{k}} \sqrt{\frac{\omega_{\mathbf{p}}}{2}} \left(b_{\mathbf{k}} e^{i(\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t)} + b_{\mathbf{k}}^{\dagger} e^{-i(\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t)} \right) \equiv \frac{1}{\sqrt{V}} \sum_{\mathbf{k}} \sqrt{\frac{\omega_{\mathbf{p}}}{2}} (b_{\mathbf{k}} + b_{-\mathbf{k}}) e^{i(\mathbf{k} \cdot \mathbf{r} - \omega_{\mathbf{k}} t)}$$

and in the third line we use the commutation relation of $b_{\mathbf{k}}$ and find that the only non-vanishing term is

$$\langle \Omega | b_{\mathbf{k}} b_l^{\dagger} | \Omega \rangle = \delta_{kl}.$$

Because **different form fermions, there is no such momentum ball for bosons and the “physical vacuum state” is exactly the vacuum state in the usual meaning.** So we have $b_{\mathbf{k}} | \Omega \rangle \equiv 0$ while $b_{\mathbf{k}}^{\dagger} | \Omega \rangle = | \phi_{\mathbf{k}} \rangle$. Thus

$$\begin{aligned} D^{(0)}(\mathbf{p}, \omega) &\equiv \int D^{(0)}(\mathbf{r}, t) e^{-i(\mathbf{p} \cdot \mathbf{r} - \omega t)} d\mathbf{r} dt \\ &= \frac{\omega_{\mathbf{p}}}{2i} \left(\int_0^{\infty} e^{i(\omega - \omega_{\mathbf{p}})t} dt + \int_{-\infty}^0 e^{i(\omega + \omega_{\mathbf{p}})t} dt \right) \\ &= \frac{\omega_{\mathbf{p}}}{2i} \left(\frac{i}{\omega - \omega_{\mathbf{p}} + i0} + \frac{-i}{\omega + \omega_{\mathbf{p}} - i0} \right) \\ &= \frac{\omega_{\mathbf{p}}^2}{\omega^2 - \omega_{\mathbf{p}}^2 + i0}, \end{aligned}$$

where in the last line we expand the power function

$$(\omega_{\mathbf{p}} - i0)^2 \sim \omega_{\mathbf{p}}^2 + 2i\omega_{\mathbf{p}}0 \sim \omega_{\mathbf{p}}^2 + i0$$

because by definition, the ground state energy is always no less than zero (and thus would not influence our choice of contour in calculation of integrals). \square

2.4 General Analytical Properties

In this section we will discuss the analytic properties of Green functions around the poles on the ω plane and show that the behavior in free cases is not a special property but a common one.

All our discussion bases on the two postulation that

- **The system is microscopically homogeneous.**
- **The multi-particle state is stable, or space and time invariant.**

Thus for some system with defects or unstability, all analytic properties obtained in this section must be questioned or even become entirely invalid.

Before our derivation, we should recall that for a system with conservative energy and momentum, its fields operator can be rewritten as space-time translation

$$\psi^H(\mathbf{r}, t) = e^{-i(\mathbf{P} \cdot \mathbf{r} - H't)} \psi^S(\mathbf{0}) e^{i(\mathbf{P} \cdot \mathbf{r} - H't)}. \quad (2.4.1)$$

So we have

命题 1 (Källèn-Lehmann Spectral Representaion¹⁵) The Green function for fermions (particularly, electrons) can be rewritten in the momentum space that

$$G(\omega, \mathbf{p}) = \frac{(2\pi)^3}{2} \sum_s \left\{ \frac{A_s \delta(\mathbf{p} - \mathbf{p}_s)}{\omega - \varepsilon_s^+ + \mu + i0} \mp \frac{B_s \delta(\mathbf{p} + \mathbf{p}_s)}{\omega - \varepsilon_s^- + \mu - i0} \right\}, \quad (2.4.2)$$

where sign “ $-$ ” is for bosons while “ $+$ ” for fermions¹⁶, $A_s \equiv |\langle \Omega | \psi_\alpha(0) | \psi_s \rangle|^2$, $B_s \equiv |\langle \psi_s | \psi_\alpha(0) | \Omega \rangle|^2$, $\varepsilon_s^+ = E_s(N+1) - E_\Omega(N)$ and $\varepsilon_s^- = E_\Omega(N) - E_s(N-1)$.

证明 For simplicity we just ignore the label of spins components (if you want to consider it you just need to add an factor of one half for Fermion liquids). Inserting the complete set of one arbitrary basis of Hilbert space, say, $I = \sum_s |s\rangle \langle s|$ and substituting the above decomposition under shifting of spaces and time, we have

$$\begin{aligned} iG(t_1, t_2) &= \langle \Omega | \psi^H(x_1) \psi^{H\dagger}(x_2) | \Omega \rangle \theta(t_1 - t_2) \pm \langle \Omega | \psi^{H\dagger}(x_2) \psi^H(x_1) | \Omega \rangle \theta(t_2 - t_1) \\ &= \sum_s \left(\langle \Omega | e^{-i(\mathbf{P} \cdot \mathbf{r} - H't_1)} \psi e^{i(\mathbf{P} \cdot \mathbf{r} - H't_1)} | s \rangle \langle s | e^{i(\mathbf{P} \cdot \mathbf{r} - H't_2)} \psi e^{-i(\mathbf{P} \cdot \mathbf{r} - H't_2)} | \Omega \rangle \theta(t_1 - t_2) \right. \\ &\quad \left. \pm \langle \Omega | e^{i(\mathbf{P} \cdot \mathbf{r} - H't_2)} \psi e^{-i(\mathbf{P} \cdot \mathbf{r} - H't_2)} | s \rangle \langle s | e^{-i(\mathbf{P} \cdot \mathbf{r} - H't_1)} \psi e^{i(\mathbf{P} \cdot \mathbf{r} - H't_1)} | \Omega \rangle \theta(t_2 - t_1) \right) \\ &= \sum_s \left(\theta(t_1 - t_2) e^{i(E_\Omega - \mu N_\Omega)t_1} \langle \Omega | \psi | s \rangle e^{-i((E_s - \mu N_s)t_1 - \mathbf{P} \cdot \mathbf{x}_1)} \right. \end{aligned}$$

¹⁵I give this such name because in QFT we also have a similar non-perturbative result as the implement of unitarity

¹⁶Conventionally I set the upper sign for bosons, and the lower sign for fermions.

$$\begin{aligned}
& \times e^{i((E_s - \mu N_s)t_2 - \mathbf{P} \cdot \mathbf{x}_2)} \langle s | \psi^\dagger | \Omega \rangle e^{-i(E_\Omega - \mu N_\Omega)t_2} \\
& \pm \theta(t_2 - t_1) e^{i(E_\Omega - \mu N_\Omega)t_2} \langle \Omega | \psi^\dagger | s \rangle e^{-i((E_s - \mu N_s)t_2 - \mathbf{P} \cdot \mathbf{x}_2)} \\
& \times e^{i((E_s - \mu N_s)t_1 - \mathbf{P} \cdot \mathbf{x}_1)} \langle s | \psi | \Omega \rangle e^{-i(E_\Omega - \mu N_\Omega)t_1} \Big),
\end{aligned}$$

where in the third line we use the fact that the momentum of ground state is always zero $P_\Omega = 0$.

Note that we share a similar positive factor $|\langle \Omega | \psi | s \rangle|^2$ and $|\langle s | \psi | \Omega \rangle|^2$ in each terms, while annihilation operator has function of decreasing number of particles for a multiparticle state by one, so the only non-zero contribution demands the confinement separately:

$$\begin{cases} N_s = N_\Omega + 1, & t_1 > t_2, \\ N_s = N_\Omega - 1, & t_1 < t_2. \end{cases}$$

Denote N_Ω as N for short, hence for the first half of term, we have its exponential

$$\begin{aligned}
& \exp \left[i(E_\Omega(N) - \mu N)t_1 - i \left((E_s(N+1) - \mu \cdot (N+1))t_1 - \mathbf{P} \cdot \mathbf{x}_1 \right) + \right. \\
& \quad \left. i \left((E_s(N+1) - \mu \cdot (N+1))t_2 - \mathbf{P} \cdot \mathbf{x}_2 \right) - i(E_\Omega(N) - \mu N)t_2 \right] \\
& = \exp \left[i \left(E_\Omega(N) - E_s(N+1) + \mu \right) (t_1 - t_2) + \mathbf{P} \cdot (\mathbf{x}_1 - \mathbf{x}_2) \right].
\end{aligned}$$

And for another half of term, we also have

$$\begin{aligned}
& \exp \left[i(E_\Omega(N) - \mu N)t_2 - i \left((E_s(N-1) - \mu \cdot (N-1))t_2 - \mathbf{P} \cdot \mathbf{x}_2 \right) + \right. \\
& \quad \left. i \left((E_s(N-1) - \mu \cdot (N-1))t_1 - \mathbf{P} \cdot \mathbf{x}_1 \right) - i(E_\Omega(N) - \mu N)t_1 \right] \\
& = \exp \left[i \left(E_\Omega(N) - E_s(N-1) - \mu \right) (t_2 - t_1) + \mathbf{P} \cdot (\mathbf{x}_2 - \mathbf{x}_1) \right].
\end{aligned}$$

Now introducing the notation (called the *excitation energy*) $\varepsilon_s^+ \equiv E_s(N+1) - E_\Omega(N)$, $\varepsilon_s^- \equiv E_\Omega(N) - E_s(N-1)$ and A_s, B_s above, we have

$$iG = \frac{1}{2} \sum_s \left[\theta(t_1 - t_2) A_s e^{i((-\varepsilon_s^+ + \mu)(t_1 - t_2) + \mathbf{P}_s \cdot (\mathbf{x}_1 - \mathbf{x}_2))} \pm \theta(t_2 - t_1) B_s e^{i((\varepsilon_s^- - \mu)(t_2 - t_1) + \mathbf{P}_s \cdot (\mathbf{x}_2 - \mathbf{x}_1))} \right].$$

Perform Fourier transformation on it, and denote $t \equiv t_1 - t_2$ and $\mathbf{x} = \mathbf{x}_1 - \mathbf{x}_2$ for short, we have

$$\begin{aligned}
iG(\omega, \mathbf{p}) &= \frac{1}{2} \sum_s \int d\mathbf{x} \left[\int_0^\infty dt A_s e^{i((-\varepsilon_s^+ + \mu)t + \mathbf{P}_s \cdot \mathbf{x})} e^{i(\omega t - \mathbf{p} \cdot \mathbf{x})} \right. \\
& \quad \left. \pm \int_{-\infty}^0 dt B_s e^{i((\varepsilon_s^- - \mu)(-t) + \mathbf{P}_s \cdot (-\mathbf{x}))} e^{i(\omega t - \mathbf{p} \cdot \mathbf{x})} \right].
\end{aligned}$$

To compish this integral, we need to shift the exponential slightly and substitute the formula (2.3.6). The first term is easy to done that

$$\int_0^\infty dt A_s e^{i((-\varepsilon_s^+ + \mu)t + \mathbf{P}_s \cdot \mathbf{x})} e^{i(\omega t - \mathbf{p} \cdot \mathbf{x})} = \frac{i A_s \delta(\mathbf{p} - \mathbf{P}_s)}{\omega - \varepsilon_s^+ + \mu + i0}.$$

But the sign of the second term is a little perplexing: perform a variable transformation $t \rightarrow -u$ and reverse the direction of intergal, then the second term becomes

$$\begin{aligned} & \pm \int_{-\infty}^0 dt B_s e^{i((\varepsilon_s^- - \mu)(-t) + \mathbf{P}_s \cdot (-\mathbf{x}))} e^{i(\omega t - \mathbf{p} \cdot \mathbf{x})} \\ &= \pm \int_0^\infty du B_s e^{i((\varepsilon_s^- - \mu - \omega)u - (\mathbf{P}_s + \mathbf{p}) \cdot \mathbf{x})} \\ &= \pm \frac{i B_s \delta(\mathbf{P}_s + \mathbf{p})}{\varepsilon_s^- - \mu - \omega + i0} = \mp \frac{i B_s \delta(\mathbf{P}_s + \mathbf{p})}{\omega + \mu - \varepsilon_s^- - i0}. \end{aligned}$$

So we're done. \square

命题 2 We have the inequality for exitation energy:

$$\begin{cases} \varepsilon_s^+ = E_s(N+1) - E_0(N) > \mu, \\ \varepsilon_s^- = E_0(N) - E_s(N-1) < \mu \end{cases} \quad (2.4.3)$$

证明 Recall that for an open system, we have

$$dU = T dS - p dV + \mu dN. \quad (2.4.4)$$

So at the low temperature limit $T \rightarrow 0$, we have $\mu = \partial E / \partial N$, while $E_\Omega(N+1) - E_\Omega(N) \sim \partial E / \partial N$. So

$$\varepsilon_s^+ \sim \left(E_s(N+1) - E_\Omega(N+1) \right) + \mu.$$

But by definition the difference in the bracket is always possitive, so we're done. \square

2.5 Computation of Macrospic Quantaties

In analogous of the definition of particle number operator in chapter one here we also introduce the definition of number density operator in terms of Heisenberg operators:

定义 1(Particle Density Operator) The *particle density operator* of the multi-particle system is defined to be (with the spin subscript added)

$$\rho_{\alpha\beta}(\mathbf{r}) := \frac{1}{N} \langle \psi_\beta^\dagger(t, \mathbf{r}) \psi_\alpha(t, \mathbf{r}) \rangle. \quad (2.5.1)$$

注 1 Note that this density is of the *real* particles, rather than that of quasi-particles.

Faced with such a familiar expression, I dare to say that one cannot reject to relate it with the Green functions we studied above. In fact, **it's enough to describe all the thermal preperities after knowing the Green function of the system.**

命题 1(Relation of Green Function and Macoscopic Quantaties)

1) The average density of real particles can be computed as

$$\frac{N}{V} = -2i\langle G(t=0, \mathbf{r}=0) \rangle. \quad (2.5.2)$$

2) The distribution of particles with momentum can be computed as

$$N(\mathbf{p}) = -i \lim_{t \rightarrow 0^-} \int_{-\infty}^{\infty} G(\omega, \mathbf{p}) e^{-i\omega t} \frac{d\omega}{2\pi}. \quad (2.5.3)$$

证明 By definition we have

$$G_{\alpha\beta}(x_1, x_2) = \begin{cases} -i\langle \psi_{\alpha}(x_1) \psi_{\beta}^{\dagger}(x_2) \rangle & t_1 > t_2, \\ \mp i\langle \psi_{\beta}^{\dagger}(x_2) \psi_{\alpha}(x_1) \rangle & t_1 < t_2. \end{cases}$$

So for the microscopically homogeneous systems, we have

$$n_{\alpha\beta}(\mathbf{r}) = -iG_{\alpha\beta}(t=0, \mathbf{r}=\mathbf{0}),$$

where $t \equiv t_1 - t_2$, $\mathbf{r} = \mathbf{x}_1 - \mathbf{x}_2$ (so the sign of t should be negative), and (2.5.2) can be gotten by taking the trace of spins.

Performing the Fourier transformation of $n_{\alpha\beta}$ and substitute the form of Green function in momentum space, we finally have

$$\begin{aligned} N(\mathbf{p}) &= \int d\mathbf{r} n(\mathbf{r}) e^{-i\mathbf{p}\cdot\mathbf{r}} = -i \int d\mathbf{r} G(t=0, \mathbf{r}) e^{-i\mathbf{p}\cdot\mathbf{r}} \\ &= -i \lim_{t \rightarrow 0^-} \int_{-\infty}^{\infty} G(\omega, \mathbf{p}) e^{-i\omega t} \frac{d\omega}{2\pi}. \end{aligned}$$

□

注 2 From the free cases we can see that **chemical potential is the parameter of Green function**, so (2.5.2) tells the relation between $\frac{N}{V}$ and μ .

推论 1(Relation of Green Function and Thermal Quantaties) For a system with volume unchangable, its *grand potential*¹⁷ can be computed as

$$\Omega = - \int_0^{\mu} d\mu N(\mu), \quad (2.5.4)$$

where $N(\mu)$ can be deduced from (2.5.2) (though might be just implicit functions).

¹⁷Grand potential $\Omega \equiv U - TS - \mu N$ describes system with volume unchanged.

证明 By definition, $d\Omega = -P dV - S dT - N d\mu$, but by the *third principle of thermal dynamics*¹⁸, entropy S will also go to zero as temperature does, and now $dV = 0$, so we have $d\Omega = -N d\mu$, or (2.5.4). \square

注 3 Note that this corollary holds only at $T = 0$ k.

2.6 Inclusion of Interactions

Now we are to introduce interaction in our theory, where $H' = H'_0 + H'_I$ (here we neglect the upperscript prime). Recall some important relation we derived in chapter one:

$$\begin{cases} |\psi^I(t)\rangle = S(t, t_0)|\psi(0)\rangle, \\ \hat{A}^I(t) = U_0^{-1}(t, t_0)\hat{A}^S(t)U_0(t; t_0), \end{cases}$$

and from (1.1.5) and (1.1.9) we can get the relation between Heisenberg operator and interactive operator:

$$\hat{A}^I(t) = S(t, t_0)\hat{A}^H(t)S^\dagger(t; t_0), \quad (2.6.1)$$

with $S(\infty, -\infty) = T \exp\left(-i \int dt H_I(t)\right)$ the S -matrix of our system.

The main potin in this section is to introduce the celebrated *Gell-mann Low* theorem and to utilize it in computation of Green function. But before that, we have to manually set one significant confinement on interactions in construction of perturbation theory that no one can avoid it.

假设 1(Adiabatic Hypothesis) When considering interactions, one can always regard the interaction absent at long before and in the long future, with a smooth exponent that

$$H = H_0 + H'_\varepsilon = H_0 + e^{-\varepsilon|t|} H'. \quad (2.6.2)$$

注 1 In fact, this hypothesis is equivalent to the technique of *Abel regulation*.

定理 1(Gellmann-Low Theorem) Under the adiabatic hypothesis, one can obtain the ground state of the interactive system through the ground state of non-interactive one by

$$\frac{|\psi_0^H\rangle}{\langle\Omega|\psi_0^H\rangle} = \lim_{\varepsilon \rightarrow \infty} \frac{|\psi_\varepsilon^H\rangle}{\langle\Omega|\psi_\varepsilon^H\rangle} = \lim_{\varepsilon \rightarrow \infty} \frac{S_\varepsilon(0, -\infty)|\Omega\rangle}{\langle\Omega|S_\varepsilon(0, -\infty)|\Omega\rangle}, \quad (2.6.3)$$

where S_ε is the S -matrix with an adiabatic interaction, and $|\psi_\varepsilon^H\rangle$ or $|\psi_0^H\rangle$ the ground state of interactive system in Heisenberg picture.

¹⁸*Nernst theorem*, or *third principle of thermal dynamics* tells that the entropy of the condensed system goes to zero as temperature does in the isothermal processes, i.e.,

$$\lim_{T \rightarrow 0} (\Delta S)_T = 0.$$

证明 The proof of this theorem is combusome and out of the main line of this note. For detailed proof, one can refer to 《固体量子场论》 or wikipedia. \square

注 2 When mentioned “adiabatic hypothesis”, many textbooks just take $|\psi_0^H\rangle = S(0, -\infty)|\Omega\rangle$ for granted. Certainly this is the abuse of terminology, but to some extent this pedasgogical trick makes this theorem more smooth and acceptable. So I skip the ver-bose proof here.

Gell-mann Low theorem allow us to express the average value of the interactive ground state to roducts of Heisenberg operator in time order by the average of non-interactive one. That is,

定理 2(Gellmann-Low Formula) The Green function can be compute by perturbation

$$\langle \psi_0^H | T \hat{A}(t_1) \hat{B}(t_2) \cdots \hat{D}(t_n) | \psi_0^H \rangle = \frac{\langle \Omega | T S(\infty, -\infty) \hat{A}(t_1) \hat{B}(t_2) \cdots \hat{D}(t_n) | \Omega \rangle}{\langle \Omega | S(\infty, -\infty) | \Omega \rangle}. \quad (2.6.4)$$

证明 Suppose $t_1 > \cdots > t_n$ and change the picture from Heisenberg one to interactive one, we get

$$\begin{aligned} & \langle \psi_0^H | T \hat{A}(t_1) \hat{B}(t_2) \cdots \hat{D}(t_n) | \psi_0^H \rangle \\ &= \langle \Omega | S^\dagger(0, -\infty) S(0, t_1) \hat{A}^I(t_1) S(t_1, 0) \cdots S(0, t_n) \hat{D}^I(t_n) S(t_n, 0) S(0, -\infty) | \Omega \rangle \\ &= \langle \Omega | S(-\infty, 0) S(\infty, 0)^\dagger S(\infty, 0) S(0, t_1) \times \\ &\quad \times \hat{A}^I(t_1) S(t_1, t_2) \cdots S(t_{n-1}, t_n) \hat{D}^I(t_n) S(t_n, -\infty) | \Omega \rangle \\ &= \left\langle \Omega \left| S(-\infty, \infty) T \left\{ \hat{A}(t_1) \cdots \hat{D}(t_n) S(\infty, -\infty) \right\} \right| \Omega \right\rangle, \end{aligned}$$

where in the third line we insert $\mathbb{1} \equiv S(\infty, 0)^\dagger S(\infty, 0)$. **By the non-degenerate hypothesis of the non-interactive grond state, we must have the two states $|\Omega\rangle$ and $S(\infty, -\infty)|\Omega\rangle$ differ with just a phase coefficient $e^{i\alpha}$.** Thus

$$\langle \Omega | S(-\infty, \infty) = e^{-i\alpha} |\Omega\rangle = \frac{\langle \Omega |}{\langle \Omega | S(-\infty, \infty) | \Omega \rangle}$$

and we're done. \square

推论 1

$$iG_{\alpha\beta}(x_1; x_2) = \frac{\langle \Omega | T S(\infty, -\infty) \psi_\alpha(x_1) \psi_\beta^\dagger(x_2) | \Omega \rangle}{\langle \Omega | S(\infty, -\infty) | \Omega \rangle}. \quad (2.6.5)$$

This formula will play an important role in our computation of Feynman diagrams in the next section.

2.7 Feynman Diagram

Equipped with the Gellmann-Low theorem above, we now try some low order computation to make the formiliasm more clearly.

引理 1(Wick Theorem) And form of the time order of creation and annihilation fields can be divided into two parts: one is *normal order* term, the other are pairing or complete time order pairs. For example,

$$\begin{aligned} T\{\psi_1\psi_2^\dagger \cdot \psi_3\psi_4\psi_4^\dagger\psi_3^\dagger\} &= \psi_2^\dagger\psi_4^\dagger\psi_3^\dagger\psi_1\psi_3\psi_4 + T\{\psi_1\psi_2^\dagger\}T\{\psi_3\psi_4^\dagger\}T\{\psi_4\psi_3^\dagger\} + (-1) \cdot T\{\psi_1\psi_2^\dagger\}T\{\psi_3\psi_3^\dagger\}T\{\psi_4\psi_4^\dagger\} + \\ &+ (-1) \cdot T\{\psi_1\psi_4^\dagger\}T\{\psi_2^\dagger\psi_3\}T\{\psi_4\psi_3^\dagger\} + T\{\psi_1\psi_3^\dagger\}T\{\psi_2^\dagger\psi_3\}T\{\psi_4\psi_4^\dagger\}. \end{aligned}$$

推论 1 When computing the perturbative series, norm ordering term cancel with vacuum state and so we can only consider the full-contracted terms.

2.7.1 Scalar Electron-electron Interaction

2.7.2 Scalar Electron-phonon Interaction

3

Methods on $T > 0$ Quantum Field Theory

Part I

Beyond Landau Fermi Liquid
Theory

4

Quantum Hall Effect

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