Exercise: Due by class time Wednesday, Feb 20, 2013 (no class Feb 18)

# Journal file: YourName\_ex4.pro (in your ~/ASTR2600\_assignments/assignment4/ directory)

(Create an assignment4 subdirectory under your ASTR2600\_assignments directory if necessary.)

Exercises from Chapter 8 and Chapter 9 involve journal files. This is the last assignment with long journal files as exercises. It's probably a good idea to break this up into multiple journal files as described in Assignment 3.

Exercise 8.0: Experimenting with various plot keywords

Exercise 8.1: Title & linestyle keywords, using xyouts for legends

Exercise 8.2: A little zooming Exercise 8.3: 2D plotting Exercise 8.4: Multiple plots

Exercise 9.0: Playing with color definitions.

Exercise 9.1 : Like Exercise 8.1 but with colors

### Whuduzitdo? All Whuduzitdo's from Chapters 8 and 9

Turn in via github.

**Journal files** of the exercises from Chapters 8 & 9.

## Graded Homework: Due by Wednesday, February 20th, 11:59:59 PM

(Readability is 10% of your grade)

All homeworks in this assignment & 8.3 can and should be done by writing script files instead of interactively and creating a journal file.

Homework 8.2: Script file name: YourName\_hw8.2.pro

Postscript file name: YourName\_multiPlanck.ps

Homework 8.3: Script file name: YourName\_hw8.3.pro

# Homework 8.4 (Not in book): Script file, YourName\_hw8.4.pro that does the following:

Read in the lif2b data files used in assignment 3. (Read them from the shared assignment3 directory using the full path. Do not copy them into yours.)

Draw the following plots and save them into JPEG files as described. The plots should all have appropriate  $x \, \mathcal{E} y$  axes labels and an appropriate title.

If the yaxis label appears clipped off the left edge of the window, use the xmargin keyword to remedy this. (Read about it in the book or online documentation.)

a) Plot the flux vs. wavelength. (Déjà vu. You did this in Homework 6.4)

Save an image of this plot as YourName\_fullSpectrum.jpg

b) The plot has a double peak near 990 angstroms.

Edit your script to use the xrange keyword (rather than the where function used in Homework 7.1) to plot about a 5-angstrom range roughly centered on this double peak so that we see it clearly. (This is like "zooming in" on that part of the spectrum.)

Save an image of this plot as YourName\_doublePeak.jpg

c) The plot has a large, thin peak near 1026 angstroms.

Edit the script to use the range keywords to plot about a 1-angstrom range roughly centered on this double peak. This plot should also have a flux range of about 1e-12 flux units so that we can clearly see the structure at this peak. (This means setting a yrange as well.)

Save an image of this plot as YourName\_tallPeak.jpg.

### Homework 9.2 (python; Not in book):

Create a script file YourName\_hw9.2.py that does the following:

Read a binary 500 x 500 array of floats from the file:

/home/shared/astr2600/data/temperature\_iy45.dat The format is simply an array of floats.

There is no first-number "header" with the array size. I'm telling you the size. It's 500 x 500. In order to read data from a binary file in python, you need to use np.fromfile. In order to figure this function out, look at help(np.fromfile). The first line,

```
fromfile(file, dtype='float', count=-1, sep='')
```

tells you the order of parameters. file should be the filename, and dtype is the data type.

HINT: Check to make sure the data look sane! The default numpy data type float is a float64. There is also a float32 and float128. You should figure out which of these is correct. These are numpy data types, so they have to be referred to by np.float32 etc, or you can pass them in with quotes as in the above example ('float'), or the letter f followed by the number of bytes, e.g. 'f4', 'f8', 'f16'.

fromfile will read in the data as a 1D array. Check this using the array.shape method (i.e., if you call your array x, print x.shape). I've told you the correct shape already, so you'll need to reshape the array. Here's an example where I reshape a length-100 array into a  $10 \times 10$  array:

```
x = np. arange (100)

xr = x. reshape ([10, 10])
```

Display that array as a monochrome image with imshow. Monochrome means one color, so you should use one of the black-and-white colormaps. Save the figure with savefig("YourName\_solarTemp\_monochrome.png").

Open another python plotting window using the figure function. Plot the data as a filled color contour map with 10 contours. The contourf command fills in the colors (contour just shows lines). The help for contourf is confusing, but the last unnamed keyword argument is the number of levels, e.g. contour(image,number\_of\_levels). Save a figure using a reasonable colormap (for the sun's surface) using savefig: savefig("YourName\_solarTemp\_auto.png")

Now, in a new figure, define a range of colors where the largest contour values are the brightest. In contour, if you want to specify your own levels, you use a list of levels, e.g. contourf(image,levels=[0,1,2,3]). Look at the min and max values of your array to figure out appropriate level values.

Specify these colors in the contourf command to match the levels. You can specify colors in any way shown in the lecture (color vector, hexadecimal, name), but you need to use 10 different colors. Use multiple hues, e.g. blue to red to yellow. The 10 colors must correspond to 10 levels that you can specify 'by hand' or with a convenience function (e.g., linspace).

The title of the plot should be "Solar Temperature". You don't need to label the axes. Save this contour plot as a JPG file with the name YourName\_solarTemp.jpg

Run this script with %run YourName\_solarTemp.py. Make sure it runs flawlessly!

Turn in via github: Your IDL & python script files:

YourName\_hw8.2.pro
YourName\_hw8.3.pro
YourName\_hw8.4.pro (for one of the plots)
YourName\_hw9.2.py
Your ps, jpg & png files:
YourName\_multiPlanck.ps
YourName\_fullSpectrum.jpg
YourName\_doublePeak.jpg
YourName\_tallPeak.jpg
YourName\_solarTemp\_monochrome.png
YourName\_solarTemp\_auto.png

YourName\_solarTemp.jpg