

The Avalanche Engine

By Barthélemy Paléologue

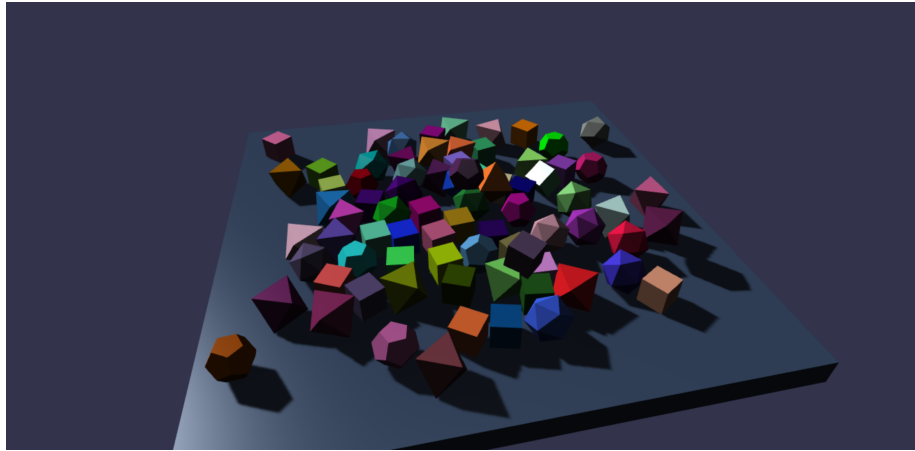


Figure 1: Avalanche Engine

Introduction

For the past month, I have been working on rigidbody physics for my final project. More precisely, I worked on non-penetration constraints, which are used to prevent rigidbodies from intersecting. My goal was to create a fast and accurate physics engine, which could then be used in my other projects. One of the constraints I added myself is working with both convex and concave shapes such as terrain, because I want to use the engine for procedural planets which are made of concave terrain.

The Source Code

The project has been built using WebGL, therefore online demos are available:

The official repo: <https://github.com/BarthPaleologue/Avalanche>

Simple demo: <https://barthpaleologue.github.io/Avalanche/dist/>

Demo terrain: <https://barthpaleologue.github.io/Avalanche/dist/nofriction.html>

Dynamic geometry: <https://barthpaleologue.github.io/Avalanche/dist/carpet.html>

Small planet: <https://barthpaleologue.github.io/Avalanche/dist/planeta.html>

There is a small UI that allows to display bounding boxes, contact points, the hash grid, resting bodies and wireframes. You can click on bodies to apply an

impulse to them.

To run the code on your machine you must have Node.js installed. Then, clone the repo and run the following commands:

```
npm install
npm run serve
```

To run the production speed without overhead, run:

```
npm run serve:prod
```

If you do any changes to the code, you can run it on the production server by running:

```
npm run build
npm run serve:prod
```

Finally, there is a video of the engine in action: [Video](#)

The code of the engine is in the `src/ts` folder. You will find the important files `engine.ts` and `rigidbody.ts`. The other files are either called by them or are just used for the demos. The intersection code is in `src/ts/utils/intersection.ts`.

The Physics Engine

Rigidbody motion

The motions of the bodies are based on what we have done in TP with Euler integration. I use quaternions for rotations.

I added Force fields to alter the motion of the bodies in a continuous way (gravity for example). There is a uniformly directional force field, and a point force field to emulate what we experience on the ground and in space.

Collision detection: Broad phase

AABBs The collision system described in the paper is in two phases. First we want to quickly and broadly find which bodies may be colliding. As the paper describes how AABBs work in great length, I chose to implement this phase using AABBs. One thing I also like about AABBs is that the intersection between two AABBs is another AABB, which will be useful later.

Each body is given 2 AABBs, one for the current state, and one for the next state that is being computed. I add an epsilon-offset to the AABBs to catch more collisions early on. It allows for faster moving objects without interpenetration.

Infinite Spatial Hash Grid To avoid testing all AABBs against each other, I chose a data-structure that allows me to reduce the number of comparisons. I could have used a regular grid, but we would be confined to a finite space. On

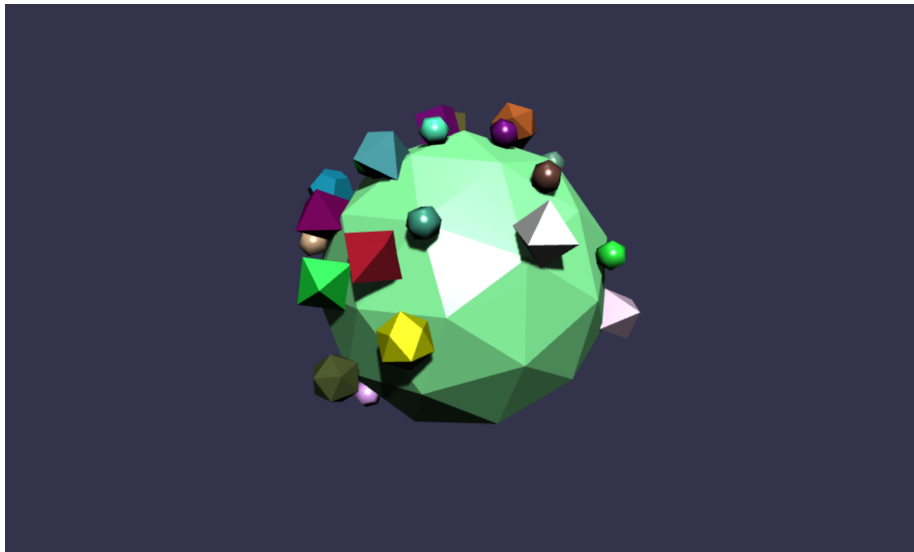


Figure 2: Planetary gravity

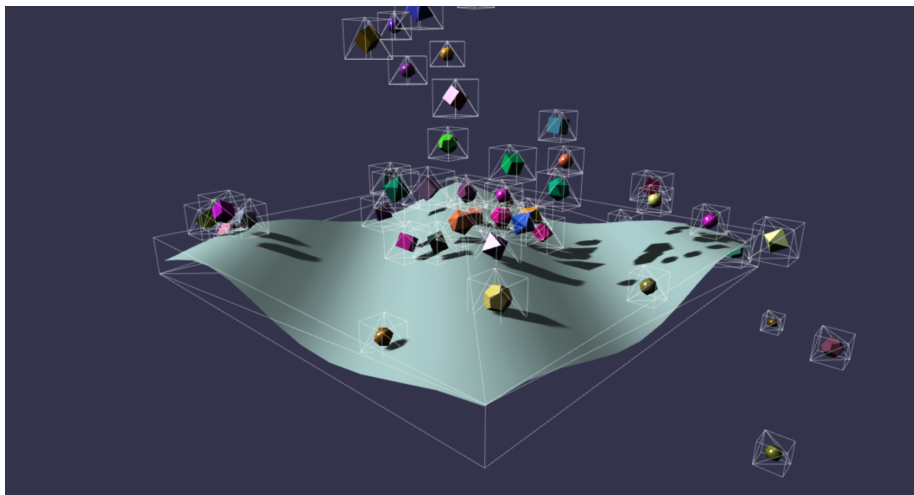


Figure 3: AABBs in action

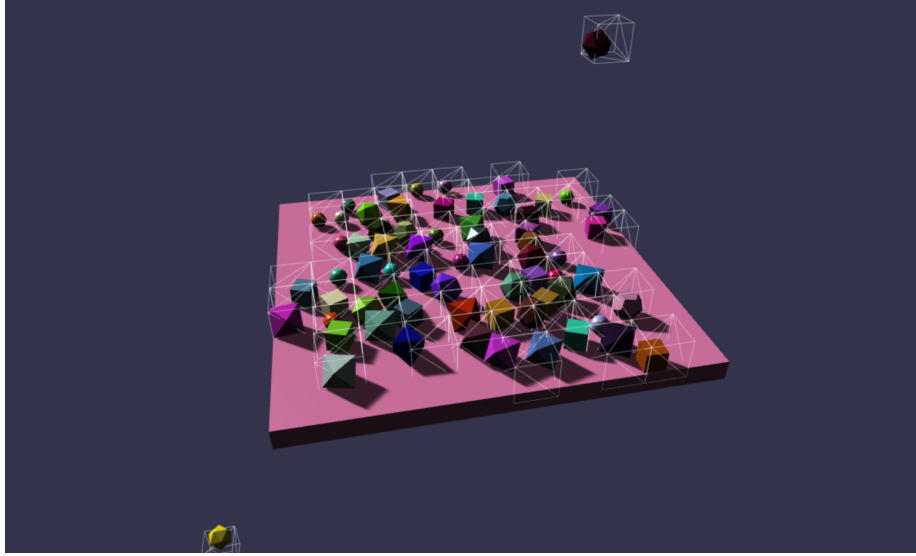


Figure 4: Infinite Spatial Hash Grid in action

the other hand, octrees give dynamic resolution and are interesting but still we are confined to finite space. Yes I hate finite space. That is why I chose to use an infinite spatial hash grid to overcome this limitation. Yet I have to deal with a static resolution.

The grid is rebuilt at each update and each cell contains a list of all the bodies inside of it. During the broad phase I then just have to check the cell the body is in and the adjacent cells. This is a very fast operation, and it allows me to quickly find which bodies may be colliding.

Rest Detection Improving the complexity of the broad phase is not everything, the goal is to eliminate as many meaningless contacts as possible. Maybe two bodies have their AABBs intersecting, but they are resting on the ground and not colliding. It would be faster if we could detect this and avoid computing the expensive narrow phase in that case.

I use different methods to achieve this. First, a body at rest must have minimal linear speed and angular speed. For this we can use a simple threshold. But it is not enough, because a body could be colliding in a way that it negate all its speed before slowly falling on the ground.

It makes sense that the velocity of the body should then be decreasing in the last frames for it to be resting. I use a vector queue to store the last velocities of the body. I then compute the variance of the velocities: if it is low, then the body is either resting or freefalling. If the variance is high, then the body has a changing velocity, hinting at collisions.

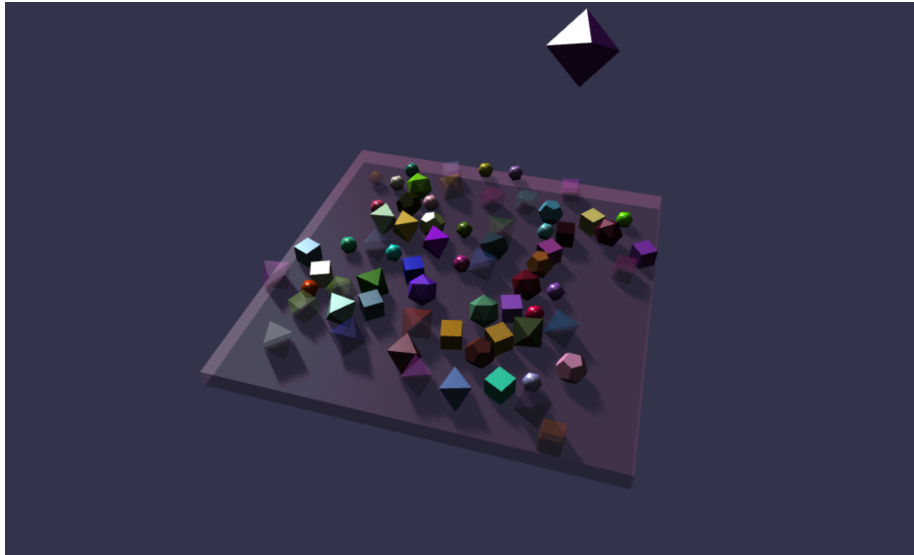


Figure 5: Transparent Bodies are optimized as resting

To prevent considering freefalling objects as resting, I also check that the body is intersecting with at least one other body. If it is not, then it is freefalling and not resting.

The last thing to consider are the neighboring bodies. If the neighbors are not resting, then the body might be colliding with them. I use the same method as above to check if the neighbors are resting. This creates some instability, but the speedup is worth it.

Using the spatial hash grid and the rest detection, I can reach 100 bodies at 60fps on my potatoe laptop.

Collision detection: Narrow phase

Now that we have a list of bodies that may be colliding, we have to check if they are actually colliding. This is the narrow phase. If I was only dealing with convex shapes, I could use the GJK algorithm. But I also want to deal with concave shapes, so I have to use a different algorithm, more general. (See the terrain demo where the terrain is concave)

Vertex to Face What I use is vertex to face collision detection. I consider the ray coming from the center of mass of one body going to its vertex and I consider the intersection with the triangles of the other body:

By doing this, I can detect collisions between convex and concave shapes. It gives us a nice way to compute the penetration distance.

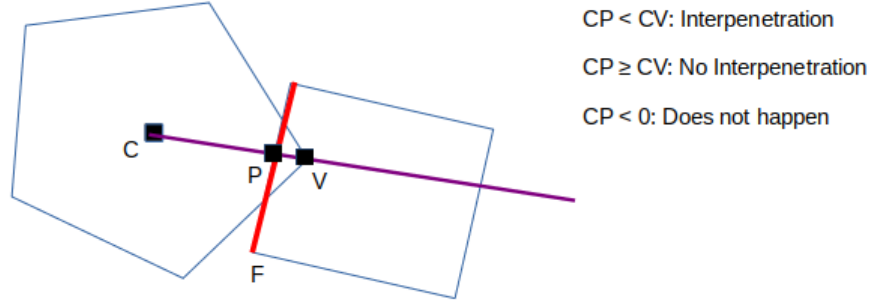


Figure 6: Ray/Face intersection

Then according to the paper, we use bisection to go back to the instant of collision by recomputing the next step for different time steps. It is quite fast and limiting the number of iterations of the bisection to 7 gives a good accuracy.

If bisection cannot converge, I push the bodies apart using a hard translation that prevents any interpenetration. This is not ideal, but it is better than having bodies stuck together.

Then we can use the previous method to compute the contact point and the normal. We can then try to resolve the collision.

Edge to Edge I started to work on edge to edge collisions but it was really slow and imprecise. In the end I decided to not use it as I would have needed more time.

Collision resolution

We use the impulse method to resolve the collision using the huge formula derived in the paper. We can handle static bodies by setting the inverse mass and inverse inertia tensor to zero. The impulse are then applied to the next step and applied to the bodies. Each body comes with a restitution coefficient that determines how much energy is lost in the collision.

Instead of using the resting contacts of the paper, I chose to use friction impulses with my rest detection mechanism I described earlier.

Friction impulses are based on the Coulomb friction model, which is a good approximation for most materials and allows finer control using a per-body friction coefficient.

This approach is very fast, contrary to what was proposed in the paper. The main limitation is that using only vertex to face collisions, the bodies can experience

$$j_t = \frac{-v_r \cdot t}{m_1^{-1} + m_2^{-1} + (I_1^{-1}(r_1 \times t) \times r_1 + I_2^{-1}(r_2 \times t) \times r_2) \cdot t}$$

Figure 7: friction impulse

instability and oscillate. The rest detection can mitigate this issue a little, but shapes with many polygons such as the Stanford bunny will still experience instability.

Another solution would be to use colliders independent from the mesh geometry. This would allow to have more stable and faster collisions, but it would be more complex to implement.

References

Non-penetration constraints: <https://www.cs.cmu.edu/~baraff/sigcourse/notesd2.pdf>

GJK algorithm: <https://cse442-17f.github.io/Gilbert-Johnson-Keerthi-Distance-Algorithm/>

GJK algorithm: <https://www.youtube.com/watch?v=MDusDn8oTSE&t=322s>

Friction Impulse: https://gafferongames.com/post/collision_response_and_coulomb_friction/

Finding inertia tensor of a 3d triangle: https://web.archive.org/web/20161229044620/https://en.wikipedia.org/wiki/Inertia_tensor_of_triangle

Finding inertia tensors for arbitrary polyhedra: <https://citeseerx.ist.psu.edu/viewdoc/download?doi=10.1.1.56.127&rep=rep1&type=pdf>

Spatial Hash Grids: <https://www.youtube.com/watch?v=sx4IIQL0x7c>