

Gammapy – A prototype for the CTA science tools

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Gammapy is a Python package for high-level gamma-ray data analysis built on Numpy, Scipy and Astropy. Starting with event lists and instrument response information, it is possible to analyse gamma-ray data and to create for example sky images, spectra and lightcurves, and to determine the position, morphology and spectra of gamma-ray sources.

So far Gammapy has mostly been used to analyse data from H.E.S.S. and Fermi-LAT, and now it is being used for the simulation and analysis of observations from the Cherenkov Telescope Array (CTA). We have proposed Gammapy as a prototype for the CTA science tools. This contribution will give an overview of the Gammapy package and show an analysis application example with simulated CTA data.

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1. Introduction

The Cherenkov Telescope Array (CTA) will observe the sky in very-high-energy (VHE, $E > 20\text{ GeV}$) gamma-ray light soon. CTA will consist of large telescope arrays at two sites, one in the southern (Chile) and one in the northern (La Palma) hemisphere. It will perform surveys of large parts of the sky, targeted observations on Galactic and extra-galactic sources, and more specialized analyses like a measurement of charged cosmic rays, constraints on the intergalactic medium opacity for gamma-rays and a search for dark matter. Compared to current Cherenkov telescope arrays such as H.E.S.S., VERITAS or MAGIC, CTA will have a much improved detection area, angular and energy resolution, improved signal/background classification and sensitivity. CTA is expected to operate for thirty years, and all astronomers will have access to CTA high-level data, as well as CTA science tools (ST) software. The ST can be used for example to generate sky images and to measure source properties such as morphology, spectra and light curves, using event lists as well as instrument response function (IRF) and auxiliary information as input.

Gammapy is a prototype for the CTA ST, built on the scientific Python stack and Astropy [1], optionally using Sherpa [2, 3, 4] or other packages for modeling and fitting (see Figure 1). Initially, the focus was to implement the “classical TeV analysis”, using 2-dimensional sky images for source detection and morphology fitting, followed by spectral analysis for a given source region. A 3-dimensional analysis with a simultaneous spatial and spectral model of the gamma-ray emission, as well as background (called “cube analysis” in the following) is in development. A first study comparing spectra obtained with the classical 1D analysis and the 3D cube analysis using point source observations with H.E.S.S., to compare the methods and to validate Gammapy, is presented in [5]. Further developments and verification using data from existing Cherenkov telescope arrays such as H.E.S.S. and MAGIC, as well as simulated CTA data is ongoing. Gammapy is starting to be used for scientific studies for existing ground-based gamma-ray telescopes [6, 7], the Fermi-LAT space telescope [8], as well as for CTA [9, 10, 11].

In this writeup we focus on the software and technical aspects of Gammapy. We start with a brief overview of the context in Section 2, followed by a description of the Gammapy package in Section 3, the Gammapy project in Section 4 and finally our conclusions concerning Gammapy as as CTA science tool prototype in Section 5.

2. Context

Prior to Gammapy, in 2011/2012, a collection of Python scripts for the analysis of IACT data was released as a first test of open-source VHE data analysis software: “PyFACT: Python and FITS Analysis for Cherenkov Telescopes” [12]. It has not been updated since, but we mention it to show that the idea to build the CTA science tools as a Python package using Numpy and Scipy started with the CTA project. In 2011, the astronomical Python community came together and created the Astropy project and package [1], which is a key factor making Python the most popular language for astronomical research codes (at least according to this informal analysis [13] and survey [14]). Gammapy is an Astropy affiliated package, which means that where possible it uses the Astropy core package instead of duplicating its functionality, as well as having a certain quality standard such as having automated tests and documentation for the available functionality. In recent



Figure 1: The Gammapy software stack. Required dependencies (Python, Numpy and Astropy) are illustrated with solid arrows, optional dependencies (SciPy and Sherpa) with dashed arrows.

years, several other packages have adopted the same approach, to build on Python, Numpy and Astropy. To name just a few, there is ctapipe (github.com/cta-observatory/ctapipe), the prototype for the low-level CTA data processing pipeline (up to the creation of lists of events and IRFs); Naima for modeling the non-thermal spectral energy distribution of astrophysical sources [15]; PINT, a new software for high-precision pulsar timing (github.com/nanograv/PINT), and Fermipy (github.com/fermipy/fermipy), a Python package that facilitates analysis of data from the Large Area Telescope (LAT) with the Fermi Science Tools and adds some extra functionality.

We note that many other astronomy projects have chosen Python and Astropy as the basis both for their data calibration and reduction pipelines and their science tools. Some prominent examples are the Hubble space telescope (HST) [16], the upcoming James Webb Space Telescope (JWST) [17] and the Chandra X-ray observatory [2, 18]. Even projects like LSST that started their analysis software developments before Astropy existed and are based on C++/SWIG are now actively working towards making their software interoperable with Numpy and Astropy, to avoid duplication of code and development efforts, but also to reduce the learning curve for their science tool software (since many astronomers already are using Python, Numpy and Astropy) [19].

For current ground-based IACTs, data and software are mostly private to the collaborations operating the telescopes. CTA will be, for the first time in VHE gamma-ray astronomy, operated as an open observatory. This implies fundamentally different requirements for the data formats and

software tools. Along this path, the current experiments, H.E.S.S., MAGIC and VERITAS, have started converting their data to FITS format, yet relying on different (some private, some open) analysis tools. To ensure compatibility it became necessary to write down the details of the data format and to a large extent it is now the same, although one has to mention that the development of the CTA high-level data model and formats is work in progress. In particular the methods to link events to IRFs will have to be extended to support multiple event types and the IRF and background model formats will have to be extended to be more precise [20]. Gammapy is participating in and contributing to this effort to prototype and find a good high-level data model and formats for CTA.

3. Gammapy package

Gammapy is a Python package built on Numpy [21], Scipy [22] and Astropy [1]. Optional dependencies are Scipy for integration and interpolation, and Sherpa [2, 3, 4] for modeling and fitting and Matplotlib [23] for plotting. The Gammapy dependency stack is shown in Figure 1.

The functionality is organized into sub-packages, such as for example *gammapy.data*, *gammapy.irf*, *gammapy.spectrum*, The Gammapy features are described in detail in the Gammapy documentation (docs.gammapy.org) and many examples given in the tutorial-style Jupyter notebooks, as well as in [24]. Here, we wanted to focus on Gammapy as a software package and a new approach to build the CTA science tools. So instead of listing the currently available functionality here, we will continue in the next section with a code example explaining how Gammapy works.

Gammapy is written high-level Python code, the data is stored in Numpy arrays or objects such as *astropy.coordinates.SkyCoord* or *astropy.table.Table* that hold Numpy array data members. Almost all functionality needed has been written in C and Python wrappers already. Specifically, *astropy.wcs* is calling into WCSLib ([25]), *astropy.io.fits* uses CFITSIO ([26]) and *astropy.coordinates* as well as *astropy.time* are built on ERFA (github.com/liberfa/erfa), the open-source variant of the IAU Standards of Fundamental Astronomy (SOFA) C library (www.iausofa.org).

An example script that generates a counts image from an event list using Gammapy is shown in Figure 2. The point we want to make here is that it is possible to efficiently work with events and pixels and to implement algorithms from Python, by storing all data in Numpy arrays and processing via calls into existing C extensions in Numpy and Astropy. E.g. here *EventList* stores the RA and DEC columns from the event list as Numpy arrays, and *SkyImage* the pixel data as well, and *image.fill(events)*, and all processing happens in existing C extensions (Numpy histograms and Astropy calls into the CFITSIO and WCSLib C libraries). Should the need arise, that some algorithm can't be efficiently implemented in pure Python, a small C extension can be added to Gammapy, using e.g. Cython [27].

We note that a huge eco-system of scientific Python packages that operate on Numpy arrays is available for advanced users, to implement analysis methods for special use cases (e.g. sky or background models, IRF handling or likelihood or Bayesian analysis methods). To name just a few that we have seen used in conjunction with Gammapy so far (by users, not within the Gammapy package itself): healpy, Scipy, pandas, scikit-image, scikit-learn, iminuit, emcee.

```

1  """Make a counts image with Gammapy."""
2  from gammapy.data import EventList
3  from gammapy.image import SkyImage
4  events = EventList.read('events.fits')
5  image = SkyImage.empty(
6      nxpix=400, nypix=400, binsz=0.02,
7      xref=83.6, yref=22.0,
8      coordsys='CEL', proj='TAN',
9  )
10 image.fill_events(events)
11 image.write('counts.fits')

```

Figure 2: An example script using Gammapy to make a counts image from an event list. This is used in Section 3 to explain how Gammapy achieves efficient processing of event and pixel data from Python: all data is stored in Numpy arrays and passed to existing C extensions in Numpy and Astropy.

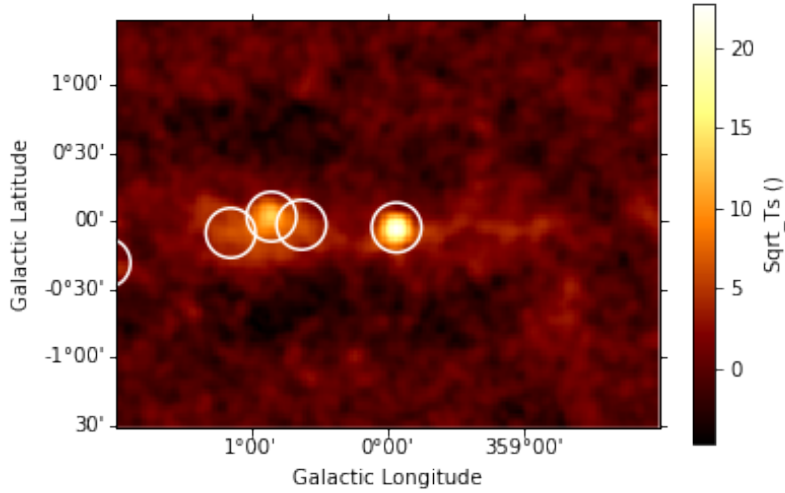


Figure 3: Application example: significance image for the Galactic centre region using 1.5 hours of simulated CTA data. White circles are peaks above 5 sigma.

In this poster we focused on the software and technical aspects of Gammapy. For examples of CTA science studies using Gammapy, we refer you to other posters presented at this conference: Galactic survey [10], PeVatrons [11] and extra-galactic sources [9].

Several other examples using real data from H.E.S.S. and Fermi-LAT, as well as simulated data for CTA can be found via docs.gammapy.org by following the link to “tutorial notebooks”. Figure 3 shows one result of the “CTA data analysis with Gammapy” notebook: a significance sky image of the Galactic center region using 1.5 hours of simulated CTA data. The background was estimated using the ring background estimation technique, and peaks above 5 sigma are shown with white circles.

4. Gammapy project

In this section we describe the current setup of the Gammapy project. We are using the com-

mon tools and services for Python open-source projects for software development, code review, testing, documentation, package distribution and user support.

Gammapy development happens on Github (github.com/gammapy/gammapy). We make extensive use of the pull request system to discuss and review code contributions. For testing we use pytest (pytest.org), for continuous integration Travis-CI (Linux and Mac) as well as Appveyor (Windows). For documentation Sphinx (www.sphinx-doc.org), for tutorial-style documentation Jupyter notebooks (jupyter.org) are used.

Gammapy is distributed and installed in the usual way for Python packages. Each stable release is uploaded to the Python package index (pypi.python.org), and downloaded and installed by users via `pip install gammapy` (pip.pypa.io). Binary packages for conda are available via the conda Astropy channel (anaconda.org/astropy/gammapy) for Linux, Mac and Windows, which conda users can install via `conda install gammapy -c astropy`. Binary packages for the Macports package manager are also available, which users can install via `port install gammapy`. At this time, Gammapy is also available as a Gentoo Linux package (packages.gentoo.org/packages/dev-python/gammapy) and a Debian Linux package is in preparation.

For Gammapy developer team communication we use Slack (gammapy.slack.com). A public mailing list for user support and discussion is available (groups.google.com/forum/#!forum/gammapy). Two face-to-face meetings for Gammapy were organized so far, the first on in June 2016 in Heidelberg as a coding sprint for developers only, the second on in February 2017 in Paris as a workshop for both Gammapy users and developers.

5. Conclusions

In the past two years, we have developed Gammapy as an open-source analysis package for existing gamma-ray telescope and as a prototype for the CTA science tools. Gammapy is a Python package, consisting of functions and classes that can be used as a flexible and extensible toolbox to implement and execute high-level gamma-ray data analyses.

We find that the Gammapy approach, to build on the powerful and well-tested Python packages Numpy and Astropy, brings large benefits: A small codebase that is focused on gamma-ray astronomy in a single high-level language is easy to understand and maintain. It is also easy to modify and extend as new use cases arise, which is important for CTA, since it can be expected that the modeling of the instrument, background and astrophysical emission, as well as the analysis method in general (e.g. likelihood or Bayesian statistical methods) will evolve and improve over the next decade. Last but not least, the Gammapy approach is inherently collaborative, sharing development effort as well as know-how with the larger astronomical community, that to a large degree already has adopted Numpy and Astropy as the basis for astronomical analysis codes in the past 5 years.

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www.cta-observatory.org/consortium_acknowledgments

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