

## Experimental and Theoretical Investigation of HC<sub>5</sub>N Adsorption on Amorphous Ice Surface: Simulation of the Interstellar Chemistry

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HC<sub>5</sub>N adsorbed on amorphous water ice at 10 K presents an interaction with the ice surface and induces the restructuring of the ice amorphous bulk. Warming up the sample induces the HC<sub>5</sub>N desorption from the H<sub>2</sub>O ice film, between 120 and 160 K, and the associated desorption energy is 90 kJ/mol. This value is in good agreement with that calculated  $E_d$  (80 kJ/mol) and gives evidence that the amorphous ice surface is essentially dynamic. From theoretical calculations, it is shown that the HC<sub>5</sub>N moiety presents a curvature and is no more linear and stabilized by two strong N···H bonds (2.09 and 2.29 Å) and one H···O bond (1.84 Å).

### Introduction

At the present time, around 130 molecular species, neglecting isotopic variants, have been detected in interstellar clouds and circumstellar regions. Most molecules are detected in the opaque circumstellar envelopes of stars, among these molecules long carbon-chain molecules and cyanopolyynes (HC<sub>*n*</sub>N with *n* = 3–11) have also been observed in these areas.<sup>1</sup>

The rich chemistry of this medium results in the variety of physical conditions of the various interstellar objects. The emission of stellar UV photons drives a quickly evolving photochemistry. To understand the active chemistry that takes place on grain mantles (constituted by a silicate and/or carbonaceous core surrounded by an ice mantle, the latter essentially constituted of water), numerous laboratory models have been developed.

Furthermore, Titan, the largest Saturnian satellite, has a dense atmosphere similar to the terrestrial one, essentially composed of N<sub>2</sub> (≥90%) and CH<sub>4</sub> (≈2%), which allows nitrile formation under UV photon irradiations.<sup>2</sup> Up to this day, four gas-phase nitriles have already been identified in Titan's atmosphere: HCN, CH<sub>3</sub>CN, HC<sub>3</sub>N, and C<sub>2</sub>N<sub>2</sub>. The fifth one, C<sub>4</sub>N<sub>2</sub>, has been detected as a solid.<sup>3</sup> These compounds are essential for the synthesis of amino acids and are particularly helpful for understanding Titan's chemistry and probably also that of the early Earth. Since the detection of H<sub>2</sub>O in the Titan's atmosphere,<sup>4</sup> we can suppose that water ice can play a role in the chemistry of species adsorbed or trapped in its lattice. Ever since its discovery in space,<sup>5</sup> the cyanodiacetylene (2,4-pentadiynenitrile and 1-cyano-1,3-butadiyne) has drawn the attention of spectroscopists and astrochemists. It is also a potential constituent of the atmosphere of Titan and has been found in laboratory simulations with Titan-like gas mixtures.<sup>6</sup>

In previous papers,<sup>7,8</sup> the results of the structure and energy of 1:1 complexes formed between cyanopolyynes (HC<sub>5</sub>N and HC<sub>3</sub>N) and water in argon matrix have been presented. Two kinds of complexes were observed, with water playing the role of either acceptor or donor of a proton.

The purpose of the present work is to assess the chemical stability of HC<sub>5</sub>N adsorbed on water ice surface and to describe

the nature of the adsorption sites of HC<sub>5</sub>N on amorphous water ice. We wanted to answer the question “does HC<sub>5</sub>N interact with water ice like HC<sub>3</sub>N or C<sub>4</sub>N<sub>2</sub>?”

### Experimental Section

Pure cyanobutadiyne was synthesized using the method described by Trolez and Guillemin.<sup>9</sup> The sample was distilled before each deposition. Moreover, the first fraction of HC<sub>5</sub>N was always eliminated. H<sub>2</sub>O was doubly distilled before use.

Low-density amorphous ice films were obtained from a water/argon (1/50) gaseous mixture deposited on a highly polished Au-plated copper cube held at 80 K. The deposition was made under pumping to 10<sup>−7</sup> mbar to outgas Ar and obtain a porous solid.<sup>10</sup> Then, the ice film was recooled to 10 K to deposit pure HC<sub>5</sub>N (ca. 5 μmol, corresponding approximatively to a monolayer) and to study the adsorption process. Adsorption and desorption of HC<sub>5</sub>N on the amorphous ice film were monitored by infrared spectroscopy using a Nicolet series II Magma system 750 in the 4000 to 650 cm<sup>−1</sup> wavenumber range at a resolution of 1 cm<sup>−1</sup>.

Infrared spectra of amorphous ice have been extensively described in the literature.<sup>11</sup> Above 80 K, during a temperature increase, an irreversible modification of the amorphous ice toward crystalline form was observed near 145 K. Ice remains crystalline, until the sample sublimation at 180 K. This form is characterized by shoulders<sup>12</sup> at 3340 and 3150 cm<sup>−1</sup> on the large profile of the ν<sub>OH</sub> mode at 3260 cm<sup>−1</sup>.

Several temperature-programmed desorption (TPD) experiments, adapted to Fourier transform infrared (FTIR) spectroscopy as previously described,<sup>13</sup> were conducted using different β heating rates. For each β heating rate value, we could evaluate the fractional surface coverage from the normalized integrated absorbance of the HC<sub>5</sub>N vs *T*.

### Computational Details

To represent the ice surface, two general strategies are available. The molecular cluster method was carried on in our last studies;<sup>14</sup> the ice surface was represented by a *closed* cluster of hydrogen-bonded water molecules, up to 10 units generally. This option enables the use of sophisticated methods and is very

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TABLE 1: Experimental and Calculated Frequencies of HC<sub>5</sub>N<sup>a</sup>

mode	theory				experiment					
	CCSD(T) anharmonic and harmonic ( <i>italic</i> ) <sup>19</sup>		B3LYP/aug-cc-pVTZ harmonic, scaled <sup>20</sup>		Ar matrix <sup>22</sup>		solid		adsorbed on amorphous ice	
	$\nu$ (cm <sup>-1</sup> )	intensity (km/mol)	$\nu$ (cm <sup>-1</sup> )	intensity (km/mol)	$\nu$ (cm <sup>-1</sup> )	relative intensity*	$\nu$ (cm <sup>-1</sup> )	relative intensity	$\nu$ (cm <sup>-1</sup> )	relative intensity
$\nu_3 + \nu_5$	3334	39.0	3345	n/a	3320.3	15	3227	74	not found	
$\nu_1$	3382	n/a								
	3349	49.0	3321	121	3304.0	60	3209	92	not found	
$2\nu_5$	3455	91								
	2262	29.6	2306	n/a	2334.3	12	not found		not found	
$\nu_2$	2314	n/a								
	2319	33.6	2256	98	2248.5	45	2251	100	2250	100
$\nu_3$	2318	66								
	2198	5.5	2192	2	2186.4	12	2187	18	2187	15
n/a	2225	3.9								
							2175	11	2177	14
$\nu_4$	2067	0.08	2061	0	2056.2	1	2049	20	2046	18
	2090	0.03								

<sup>a</sup> \*Percent intensities (measured with the accuracy of ca. 20%) relative to the strongest band  $\nu_7$  (643.6 cm<sup>-1</sup>).

useful for spectroscopy applications. The less satisfactory aspect is that the solid three-dimensional (3D) boundary conditions cannot be included and also that the very long computation times restrict the system to a very small number of constituents.

Another quantum method can be chosen in the domain of solid-state physics. In this case, the periodic boundary conditions are fully taken into account. The periodic 3D calculations were performed within the framework of the gradient-corrected density functional theory using the *PWscf* computer code from the Quantum-Espresso package.<sup>15</sup> Exchange as well as the correlation functionals are Perdew–Burke–Ernzerhof (PBE). A plane wave basis set used was associated to Vanderbilt ultrasoft pseudopotentials<sup>16</sup> and an energy cutoff of 32 Rydberg (435 eV). The same method was carried out during our study on carbon monoxide adsorption on ice<sup>17</sup> and proved efficient even if the CO–ice interactions are very small. Of course, the plane wave basis set does not generate any problem of BSSE on the binding energy (BE).

The BE of a moiety *M* (HC<sub>5</sub>N, H<sub>2</sub>O, etc.) to ice surface is evaluated as

$$\Delta E_{\text{BE}} = E(M + \text{H}_2\text{O}) - E(M) - E(\text{H}_2\text{O}_{\text{surface}}) \quad (1)$$

This value is also referred in the text as “adsorption energy” or “interaction energy”; the desorption energy in this framework is simply  $-\Delta E_{\text{BE}}$ .

As before, we adopted the ice surface unit cell optimized by the Turin group<sup>18</sup> known as proton ordered “ $\pi$ -ice” model. The supercell is made of a two layer slab containing 32 water molecules.

## Results and Discussion

### Adsorption and Desorption of HC<sub>5</sub>N on Amorphous Ice.

The spectrum of solid HC<sub>5</sub>N, in the 650–4000 cm<sup>-1</sup> range, shows four stretching fundamental modes labeled  $\nu_1$  to  $\nu_4$  (Table 1). These bands observed at 3209, 2251, 2187, and 2049 cm<sup>-1</sup> have been assigned, respectively, to CH, CN, and two CC stretching modes.<sup>19–23</sup> We observe another band at 3227 cm<sup>-1</sup> attributed to the combination mode  $\nu_3 + \nu_5$  with regard to our previous study.<sup>23</sup> In solid phase, the  $2\nu_7$  combination mode expected around 1270 cm<sup>-1</sup> is not observed, while another intense band located at 2175 cm<sup>-1</sup> is obtained but not yet assigned.

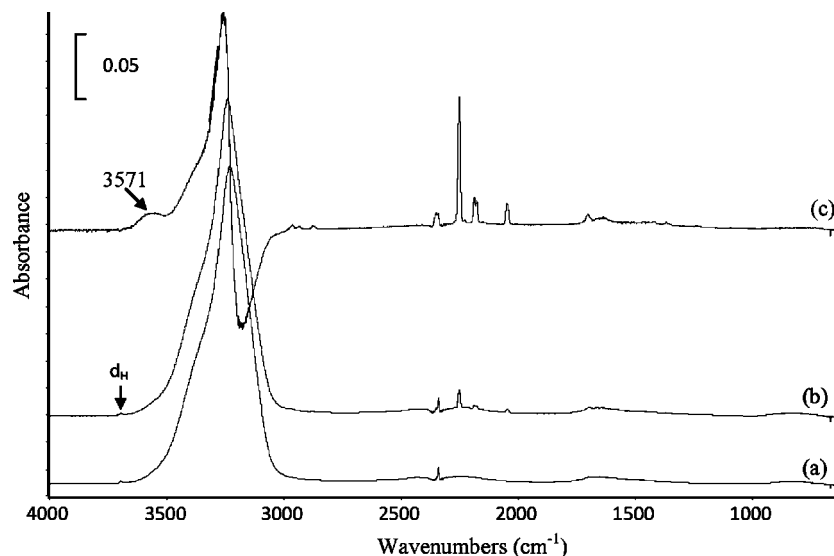
Amorphous ice displays a weak feature at 3695 cm<sup>-1</sup>, which is relative to the dangling H mode ( $d_{\text{H}}$ ).<sup>24</sup> This latter is generally used to follow the adsorption of molecules on ice surfaces. Two other vibrational surface modes exist at 3549 and 3503 cm<sup>-1</sup> attributed to the  $d_{\text{O}}$  and  $S_4$  sites, respectively.<sup>25</sup> The HC<sub>5</sub>N adsorption at 10 K on a bare amorphous ice surface induces no change in the  $d_{\text{H}}$  position, indicating that no major interaction between  $d_{\text{H}}$  and HC<sub>5</sub>N exists as with HC<sub>3</sub>N.<sup>8</sup>

However, we observe an important modification in the ice spectrum (Figure 1). The position of bulk OH stretching modes at 3230 cm<sup>-1</sup> is shifted to higher wavenumbers by 10 cm<sup>-1</sup>. This shift just as in the case of C<sub>4</sub>N<sub>2</sub><sup>26</sup> indicates a strong interaction between the HC<sub>5</sub>N and the ice film.

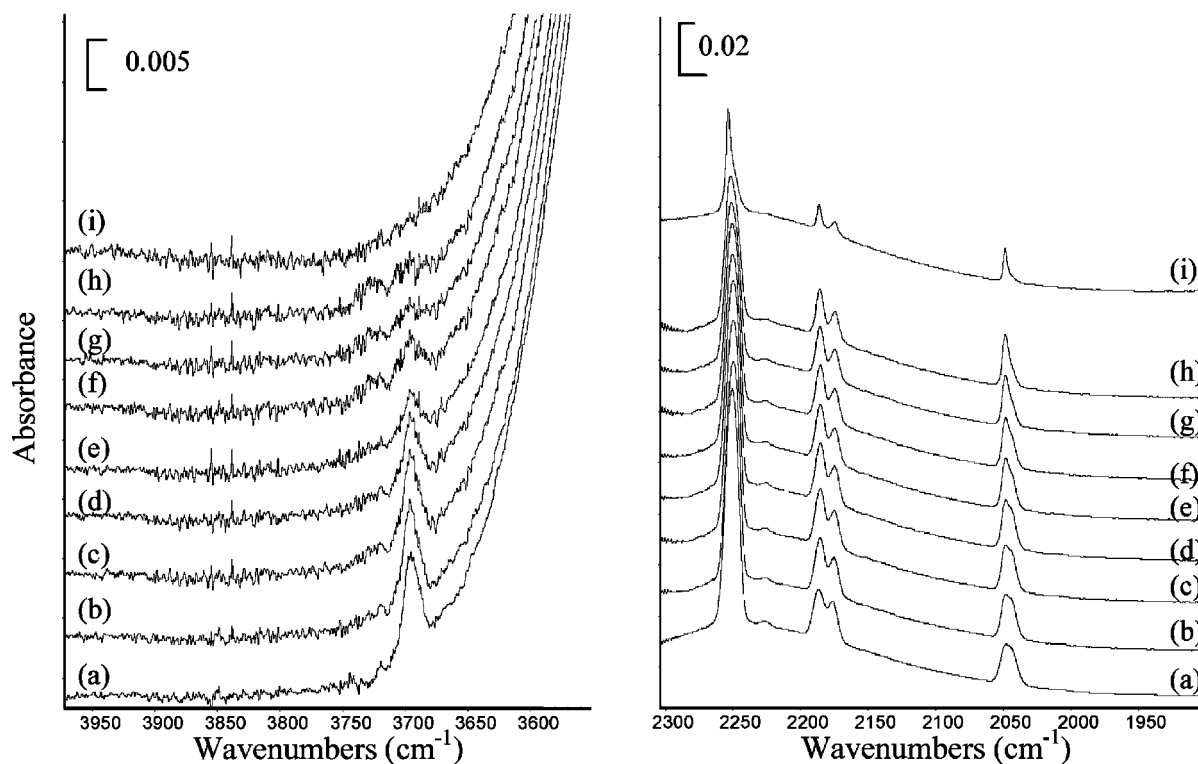
During this adsorption, a large band at 3571 cm<sup>-1</sup> grows. This band is located in the  $d_{\text{O}}$  and  $S_4$  sites regions of bare ice. Recent works<sup>25</sup> have shown that CO, CH<sub>4</sub>, and rare gas atoms (Ar and Kr) adsorbed on  $d_{\text{O}}$  and  $S_4$  sites led to a decrease in the infrared responses of these sites. However, in the case of the C<sub>4</sub>N<sub>2</sub> adsorption, an opposite effect was observed. The present results show that HC<sub>5</sub>N (just like C<sub>4</sub>N<sub>2</sub>) interacts mostly with amorphous ice sites other than  $d_{\text{H}}$ .

The infrared band shapes and positions of HC<sub>5</sub>N modes are almost the same as those observed for the solid HC<sub>5</sub>N (Table 1). Contrary to our studies in the argon matrix and those for the bulk HC<sub>5</sub>N, we could not detect the  $\nu_1$  band of HC<sub>5</sub>N deposited on ice, as it was probably masked by the strong OH stretching band of bulk ice.

Figure 2 shows selected spectra recorded during the sample warm-up from 10 K to the HC<sub>5</sub>N sublimation temperature with a heating rate  $\beta$  of 0.5 K min<sup>-1</sup>. Starting from 15 K, the band located at 3571 cm<sup>-1</sup> decreased in intensity, to disappear at 80 K. Above 85 K, the  $d_{\text{H}}$  signal began to decrease in intensity and disappeared at 130 K (Figure 2). The latter phenomenon is due to the collapse of amorphous ice pores during their evolution toward crystalline ice. The desorption process of HC<sub>5</sub>N can be followed in Figure 3 showing the variations of the  $\nu_2$  band intensity. Between 80 and 120 K, the intensity of this band remains unchanged, which is indicative of a stable adsorption state. At 140 K, we observe an abrupt decrease that results from the sublimation of adsorbed HC<sub>5</sub>N. It is remarkable that (i) this sublimation occurs at the beginning of the ice crystallization



**Figure 1.** Infrared spectra resulting from a HC<sub>5</sub>N exposure on an amorphous ice film maintained at 10 K: (a) Bare amorphous ice, (b) 5  $\mu$ mol of HC<sub>5</sub>N deposited on this ice, (c) subtraction spectrum b – a, and spectrum c multiplied by 5.

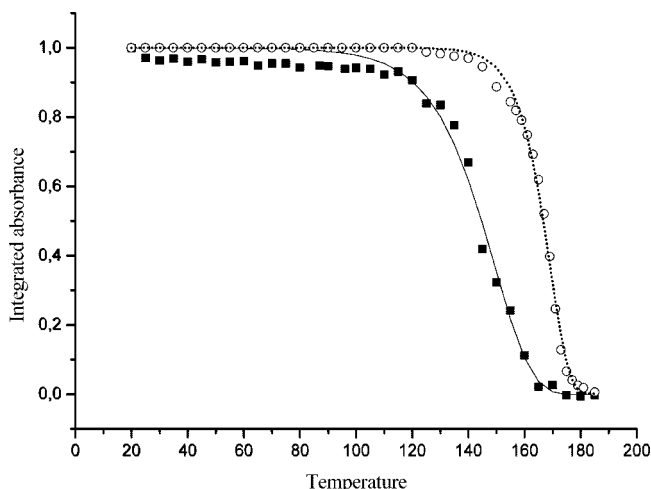


**Figure 2.** Selected infrared spectra during the TPD (heating rate  $\beta = 0.5$  K min<sup>-1</sup>) of HC<sub>5</sub>N from amorphous ice surface in the OH stretching region of ice and in the  $\nu_2$ ,  $\nu_3$ , and  $\nu_4$  mode region of HC<sub>5</sub>N: (a) 10, (b) 90, (c) 100, (d) 110, (e) 120, (f) 125, (g) 135, (h) 140, and (i) 155 K.

process near the phase transition temperature and (ii) it occurs at a lower temperature than for bulk HC<sub>5</sub>N ( $\approx 170$  K, Figure 3). Recent studies<sup>8,26</sup> have shown that the sublimation of C<sub>4</sub>N<sub>2</sub> adsorbed on ice occurs similarly at approximately 150 K, while that of HC<sub>3</sub>N takes place at a lower temperature of 135 K.

Just as for other polyyynes, the HC<sub>5</sub>N desorption follows a first-order kinetic model. To characterize the interaction energy between HC<sub>5</sub>N and amorphous ice surface, we have performed a TPD study. Table 2 lists experimental results obtained using four  $\beta$  heating rate values between 0.3 and 0.9 K min<sup>-1</sup>.  $T_p$  (temperature for which the desorption is maximal) is evaluated for each heating rate. From the linear plot of  $\ln \beta / RT_p^2$  vs  $1/T_p$  (Figure 4), we have derived the activation energy of desorption,

$E_d = 90 \pm 20$  kJ/mol, and the pre-exponential factor  $A_d = 10^{13}$  s<sup>-1</sup>. This  $E_d$  value is higher than the one evaluated with TPD for pure solid HC<sub>5</sub>N sublimation and different from that of previous reports regarding HC<sub>3</sub>N ( $39 \pm 8$  kJ/mol)<sup>8</sup> and C<sub>4</sub>N<sub>2</sub> ( $42 \pm 5$  kJ/mol).<sup>26</sup> At last, the activation desorption energy ( $90 \pm 20$  kJ/mol) measured from the HC<sub>5</sub>N desorption of the ice surface is consistent with numerous interactions between the ice surface and the HC<sub>5</sub>N moiety. Notwithstanding that the two first members of the cyanopolyynes family present some similarity during the complexation observed in cryogenic matrix, the activation desorption energy is very different. Indeed, the HC<sub>5</sub>N is the double of HC<sub>3</sub>N desorption's energy. This is probably due to the interaction of ice and the  $\Pi$  system, which must be more important in the first case.



**Figure 3.** Evolution of integrated normalized absorbance of the  $\nu_2$  mode of HC<sub>5</sub>N adsorbed on ice surface (square) and for pure HC<sub>5</sub>N (ball) with temperature (heating rate  $\beta = 0.5$  K min<sup>-1</sup>).

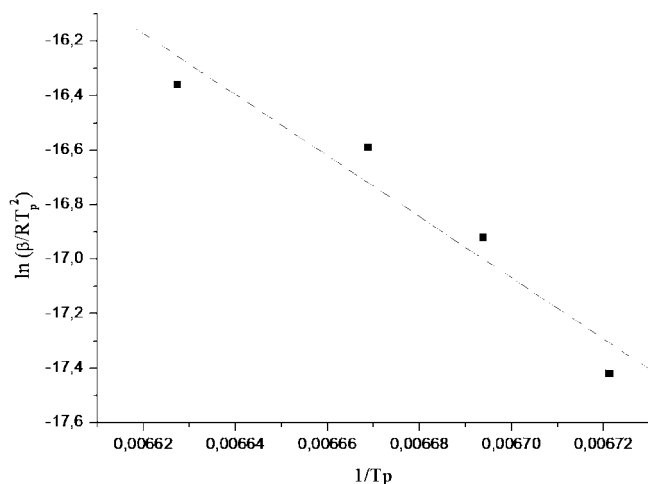
**TABLE 2: Desorption Peak Maximum Temperatures ( $T_p$ ) for Various Heating Rates ( $\beta$ )**

heating rate $\beta$ (K min <sup>-1</sup> )	$T_p$ (K)
0.3	148.8
0.5	149.4
0.7	150.0
0.9	150.9

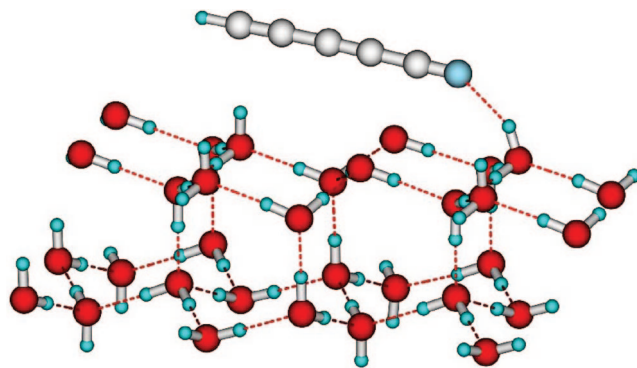
When crystalline ice is submitted to a progressive HC<sub>5</sub>N exposure, no change in the spectral line shape is observed. The infrared spectrum is similar to that of the HC<sub>5</sub>N solid. During the temperature increase, the HC<sub>5</sub>N deposited on crystalline ice sublimates around 155 K. No bands are detected in the 3500 cm<sup>-1</sup> area corresponding to a stable adsorption site.

**Quantum Study of HC<sub>5</sub>N Adsorption on Water Ice Surface.** Our experimental results indicate that HC<sub>5</sub>N adsorption does not proceed through on the dangling OH sites of the amorphous ice surface. Theoretical calculations were carried out to model the comportment of HC<sub>5</sub>N adsorbed on the amorphous ice surface.

As previously evoked, we adopted the “ $\pi$ -ice” model optimized by the Turin group<sup>18</sup> known as proton ordered. In ref 17 as well as in ref 27, this model of surface proved to be



**Figure 4.** Desorption kinetics of HC<sub>5</sub>N from an amorphous ice surface obtained for four TPD experiments. The line is derived from a linear regression analysis of the data points.



**Figure 5.** Adsorption of the HC<sub>5</sub>N molecule on bare  $\pi$ -ice surface.

realistic enough to represent accurately the ice–carbon monoxide and acetic acid interaction, although the experiments were performed on amorphous ice surfaces. Nevertheless, these molecules are very small in dimension, and within very short distance,  $\pi$ -ice is a good model of the physical environment. The situation is lightly different with HC<sub>5</sub>N, which is more than 7 Å long. Therefore, the optimized structure for adsorption on the bare  $\pi$ -ice surface looks like Figure 5.

The starting structure for geometry optimization featured the HC<sub>5</sub>N molecule roughly parallel to the ice surface; the final geometry is displayed in Figure 5. By analogy with the structure of HC<sub>5</sub>N:H<sub>2</sub>O complexes formed in solid argon,<sup>7</sup> the HC<sub>5</sub>N molecule was placed with its terminal hydrogen interacting with a free dangling O of the ice surface, while the terminal nitrogen was in interaction with a free dangling H.

The molecule is bonded to the substrate by only one weak H-bond between the nitrogen atom and the one dangling H. Consequently, the BE is small, −34 kJ/mol, and the interaction with the ice surface is weak on the contrary with our experimental finding.

Because the ice surface is essentially dynamic, water molecules can desorb and readsorb again, thus hydrating the adsorbed HC<sub>5</sub>N. Therefore, the idea was to add in sequence extra H<sub>2</sub>O on the surface around the molecule. The problem is that the molecule exhibits two hydrophilic zones at its ends, whereas the carbon atoms are less sensitive to water interaction; then, we concentrated our efforts in a model hydration of the N and H atoms. Six water molecules were enough to hydrate N, and  $\Delta E_{BE}$  dropped down to −70 kJ/mol. Two more waters yield the structure displayed in Figure 6, and the adsorption energy on this perturbed ice surface is now 80 kJ/mol. We can hardly pretend that this structure is really the physical one, but we do pretend that it is a good model for adsorption on a disordered ice surface. The structure of the adsorbed moiety is very slightly affected: The molecule undergoes a curvature and is no more linear, but the bond lengths are not modified. The system stability is ensured by two strong N...H bonds (2.09 and 2.29 Å) and a H...O bond (1.84 Å). Calculating the phonon structure of such a large crystal cell (127 atoms) is hardly possible; therefore, we have tried that on the cluster corresponding to one unit cell. The whole system is calculated using Gaussian 03<sup>28</sup> and the ONIOM model<sup>29</sup> with B3LYP/6-31 g (d,p) for the high-level cluster and AM1 for the low-level cluster.

It is worth-mentioning that the presently described interactions between the HC<sub>5</sub>N and the ice surface are very similar to those observed in the argon matrix between HC<sub>5</sub>N and water molecules. At last, it is important to note that when HC<sub>5</sub>N is adsorbed upright a d<sub>0</sub> site, the optimization leads the HC<sub>5</sub>N to be parallel to the ice surface.



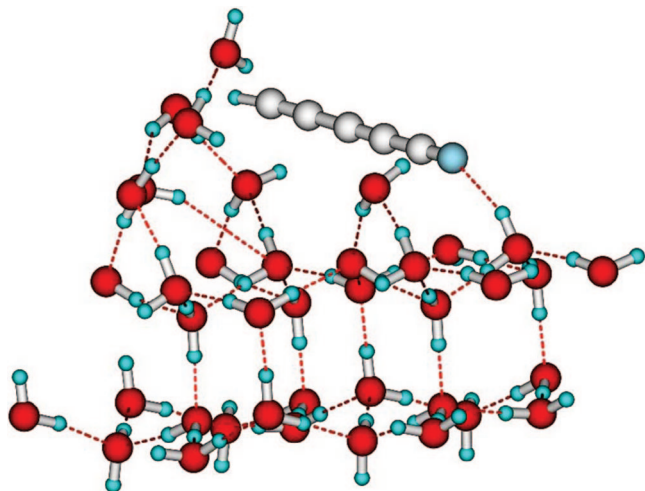


Figure 6. Adsorption and hydration of the HC<sub>5</sub>N molecule on  $\pi$ -ice.

TABLE 3: Optimized Structural Parameters for the HC<sub>5</sub>N Adsorbed on an Amorphous Ice Surface Obtained by Theoretical Calculations<sup>a</sup>

parameters	free HC <sub>5</sub> N	dH sites	HC <sub>5</sub> N adsorbed
r N <sub>1</sub> C <sub>2</sub>	1.17		1.17
r C <sub>2</sub> C <sub>3</sub>	1.35		1.35
r C <sub>3</sub> C <sub>4</sub>	1.22		1.23
r C <sub>4</sub> C <sub>5</sub>	1.34		1.34
r C <sub>5</sub> C <sub>6</sub>	1.22		1.22
r C <sub>6</sub> H <sub>7</sub>	1.10		1.11
$\angle$ N <sub>1</sub> C <sub>2</sub> C <sub>3</sub>	180.0		177.3
$\angle$ C <sub>2</sub> C <sub>3</sub> C <sub>4</sub>	180.0		176.2
$\angle$ C <sub>3</sub> C <sub>4</sub> C <sub>5</sub>	180.0		177.5
$\angle$ C <sub>4</sub> C <sub>5</sub> C <sub>6</sub>	180.0		178.0
$\angle$ C <sub>5</sub> C <sub>6</sub> H <sub>7</sub>	180.0		173.8
rOH		0.97	0.97

<sup>a</sup> Bond lengths are given in Å, and angles are given in degrees.

Table 3 gives the optimized geometrical parameters obtained for HC<sub>5</sub>N and for the dangling H sites involved in the hydrogen bonding. These data indicate that the bond lengths of HC<sub>5</sub>N adsorbed on water ice are nearly unchanged with respect to the monomer structure. Only the bond angles are modified, deviating from linearity by 2–6°. Consequently, only slight shifts of the infrared absorption bands were detected. It should be recalled that we do not detect any HC<sub>5</sub>N bending modes in the present experimental conditions. The molecule undergoes a bending and is no more linear.

Moreover, only the OH bond length of the d<sub>H</sub> site in interaction with HC<sub>5</sub>N is slightly perturbed by the adsorption (Table 3). This bond is longer than in the bare ice by only 0.001 Å. Considering the clusters that support this calculation, only the d<sub>H</sub> previously evoked is considered; in the other site, this mode is assumed to be unaffected.

We can also note that the adsorption of HC<sub>5</sub>N on the amorphous water ice leads to dramatic modifications of the ice structure after optimization (Figure 6). This result is consistent with the presently detected infrared absorption frequency shifts for the bulk OH stretching mode. The same phenomenon has already been observed for C<sub>4</sub>N<sub>2</sub> adsorbed on the ice surface.<sup>26</sup>

The predicted interaction energy of HC<sub>5</sub>N with water ice surface is in the same order of magnitude as the experimental one obtained by the TPD. This lack of accuracy is understandable, given the qualitative character of the applied theoretical model.

## Conclusion

According to our measurements of infrared absorption spectra, HC<sub>5</sub>N is stable on the ice surface up to the temperature of 150 K, which makes this adsorption interesting for the interstellar medium chemistry. HC<sub>5</sub>N is slightly curved on the ice surface and stabilized by a strong hydrogen bond (1.84 Å long) and two strong N...H bonds (2.09 and 2.29 Å). The adsorption of HC<sub>5</sub>N on amorphous ice surface induces a strong reorganization of the surface structure. The observation of a new strong band at 3571 cm<sup>-1</sup> is indicative of a major restructuring of the ice bulk coming from the fact that the amorphous ice is dynamic. This vibrational absorption comes from water molecules, which embed the HC<sub>5</sub>N moiety. However, the HC<sub>5</sub>N adsorption drives to structural modifications of the ice surface as for C<sub>4</sub>N<sub>2</sub>. No similar modifications have been reported for the HC<sub>3</sub>N adsorption. We can assume that this great ice modification comes from the  $\Pi$  system, which is similar between HC<sub>5</sub>N and C<sub>4</sub>N<sub>2</sub>. In the case of C<sub>4</sub>N<sub>2</sub> (like observed for C<sub>3</sub>O<sub>2</sub><sup>30</sup>), the absorption induces the growing of a band at 3649 cm<sup>-1</sup>, which is due to the interaction between the d<sub>H</sub> and the nitrogen atoms of C<sub>4</sub>N<sub>2</sub>. The calculations show that the C<sub>4</sub>N<sub>2</sub> or C<sub>3</sub>O<sub>2</sub> molecules were flattened on the ice surface, the two terminal atoms interacting with two free OH bonds of the ice surface. No similar interaction has been observed in our experiment.

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## References and Notes

- (1) Millar, T. J.; Rawlings, J. M. C.; Bennett, A.; Brown, P. D.; Charnley, S. B. *Astron. Astrophys., Suppl. Ser.* **1991**, 87, 585.
- (2) (a) Okabe, H. *J. Chem. Phys.* **1981**, 75, 2772. (b) Thaddeus, P.; McCarthy, M. C.; Travers, M. J.; Gottlieb, A.; Chen, W. *Faraday Discuss.* **1998**, 109, 121.
- (3) Coustenis, A.; Schmitt, B.; Khanna, R. K.; Trotta, F. *Planet Space Sci.* **1999**, 47, 1305.
- (4) Coustenis, A.; Salama, A.; Lellouch, E.; Encrenaz, Th.; Bjoraker, G. L.; Samuelson, R. E.; de Graauw, Th.; Feuchtgruber, H.; Kessler, M. F. *Astron. Astrophys.* **1998**, 336, L85.
- (5) Avery, L. W.; Broten, N. W.; MacLeod, J. M.; Oka, T.; and Kroto, H. W. *Ap. J. (Lett.)* **1976**, 205, L173.
- (6) Coll, P.; Coscia, D.; Smith, N.; Gazeau, M. C.; Ramírez, S. I.; Cernogora, G.; Israël, G.; Raulin, F. *Planet Sp. Sci.* **1999**, 47, 1331.
- (7) Coupeaud, A.; Piétri, N.; Aycard, J. P.; Couturier-Tamburelli, I. *Phys. Chem. Chem. Phys.* **2007**, 9, 3985.
- (8) Borget, F.; Chiavassa, T.; Allouche, A.; Marinelli, F.; Aycard, J. P. *J. Am. Chem. Soc.* **2001**, 123, 10668.
- (9) Trolez, Y.; Guillemin, J. C. *Angew. Chem. Int.* **2005**, 44, 7224.
- (10) (a) Givan, A.; Loewenschuss, A.; Nielsen, C. *J. Phys. Chem. B* **1997**, 101, 8696. (b) Givan, A.; Loewenschuss, A.; Nielsen, C. *J. Chem. Phys. Lett.* **1997**, 275, 98.
- (11) Rowland, D.; Devlin, J. P. *J. Chem. Phys.* **1991**, 94, 812.
- (12) Graham, J. D.; Roberts, J. T. *Geophys. Res. Lett.* **1995**, 22, 251.
- (13) Couturier-Tamburelli, I.; Chiavassa, T.; Pourcin, J. J. *Phys. Chem. B* **1999**, 103, 3677.
- (14) (a) Allouche, A. *J. Phys. Chem. A* **1999**, 103, 9150. (b) Ferro, Y.; Allouche, A. *J. Chem. Phys.* **2003**, 118, 10461. (c) Ferro, Y.; Allouche, A.; Kempter, V. *J. Chem. Phys.* **2004**, 120, 8683.
- (15) Scandolo, S.; Giannozzi, P.; Cavazzoni, C.; de Gironcoli, S.; Pasquarello, A.; Baroni, S. Z. *Kristallogr.* **2005**, 220, 574; Quantum Espresso website <http://www.quantum-espresso.org>.
- (16) Laasonen, K.; Pasquarello, A.; Lee, C.; Car, R.; Vanderbilt, D. *Phys. Rev. B* **1993**, 47, 10142.
- (17) Manca, C.; Martin, C.; Allouche, A.; Roubin, P. *J. Phys. Chem. B* **2001**, 105, 12861.

- (18) (a) Pisani, C.; Casassa, S.; Ugliengo, P. *Chem. Phys. Lett.* **1996**, 253, 201. (b) Casassa, S.; Ugliengo, P.; Pisani, C. *J. Chem. Phys.* **1997**, 106, 8030.
- (19) Botschwina, P.; Oswald, M. *Spectrochim. Acta* **1997**, A53, 1097.
- (20) Gronowski, M.; Kołos, R. *Chem. Phys. Lett.* **2006**, 428, 245.
- (21) Benilan, Y.; Ferraday, T.; Fray, N.; Jolly, A.; Raulin, F.; Guillemin, J. C. *Bull. Am. Astron. Soc.* **2005**, 37, 717.
- (22) Coupeaud, A.; Kołos, R.; Couturier-Tamburelli, I.; Aycard, J. P.; Piétri, N. *J. Phys. Chem. A* **2006**, 110, 2371.
- (23) Coupeaud, A.; Turowski, M.; Gronowski, M.; Piétri, N.; Couturier-Tamburelli, I.; Kołos, R.; Aycard, J. P. *J. Chem. Phys.* **2007**, 126, 164301.
- (24) Mayer, E.; Pletzer, R. *Nature* **1986**, 319, 298.
- (25) (a) Manca, C.; Martin, C.; Roubin, P. *Chem. Phys.* **2004**, 300, 53. (b) Manca, C. Thesis of Université de Provence, 2002. (c) Manca, C.; Martin, C.; Allouche, A.; Roubin, P. *J. Phys. Chem. B* **2001**, 105, 12861.
- (26) (a) Guennoun, Z. Thesis of Université de Provence, 2004. (b) Guennoun, Z.; Couturier-Tamburelli, I.; Piétri, N.; Aycard, J. P. *J. Phys. Chem. B* **2005**, 109, 3437.
- (27) Allouche, A.; Bahr, S. *J. Phys. Chem. B* **2006**, 110, 8640.
- (28) Frisch, M. J.; Trucks, G. W.; Schlegel, H. B.; Scuseria, G. E.; Robb, M. A.; Cheeseman, J. R.; Montgomery, J. A., Jr.; Vreven, T.; Kudin, K. N.; Burant, J. C.; Millam, J. M.; Iyengar, S. S.; Tomasi, J.; Barone, V.; Mennucci, B.; Cossi, M.; Scalmani, G.; Rega, N.; Petersson, G. A.; Nakatsuji, H.; Hada, M.; Ehara, M.; Toyota, K.; Fukuda, R.; Hasegawa, J.; Ishida, M.; Nakajima, T.; Honda, Y.; Kitao, O.; Nakai, H.; Klene, M.; Li, X.; Knox, J. E.; Hratchian, H. P.; Cross, J. B.; Bakken, V.; Adamo, C.; Jaramillo, J.; Gomperts, R.; Stratmann, R. E.; Yazyev, O.; Austin, A. J.; Cammi, R.; Pomelli, C.; Ochterski, J. W.; Ayala, P. Y.; Morokuma, K.; Voth, G. A.; Salvador, P.; Dannenberg, J. J.; Zakrzewski, V. G.; Dapprich, S.; Daniels, A. D.; Strain, M. C.; Farkas, O.; Malick, D. K.; Rabuck, A. D.; Raghavachari, K.; Foresman, J. B.; Ortiz, J. V.; Cui, Q.; Baboul, A. G.; Clifford, S.; Cioslowski, J.; Stefanov, B. B.; Liu, G.; Liashenko, A.; Piskorz, P.; Komaromi, I.; Martin, R. L.; Fox, D. J.; Keith, T.; Al-Laham, M. A.; Peng, C. Y.; Nanayakkara, A.; Challacombe, M.; Gill, P. M. W.; Johnson, B.; Chen, W.; Wong, M. W.; Gonzalez, C.; Pople, J. A. Gaussian 03, revision B.04; Gaussian, Inc.: Wallingford, CT, 2004.
- (29) Dapprich, S.; Komaromi, I.; Byun, K. S.; Morokuma, K.; Frisch, M. J. *J. Mol. Struct. Theochem.* **1999**, 461, 1.
- (30) Allouche, A.; Couturier-Tamburelli, I.; Chiavassa, T. *J. Phys. Chem. B* **2000**, 104, 1497.

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