

# ASTR 119: Session 18

## Astropy Tutorial

# Outline

- 1) Visualization of the Day
- 2) Final projects
- 3) Astropy / FITS Image Tutorial
- 4) Final project organization time
- 5) Save your work to GitHub



# Final Projects

- 1) You will be asked to perform one of four final projects, working in a **group of up to 4 people of your choosing from your SECTION**. If you would like to be assigned to a group, please let me know immediately.
- 2) Please notify me of your group by **8:00am Thursday Nov 19**. Anyone who does not respond by **8:00am Thursday** letting me know their group or their desire to be assigned to a group will work on their own final project.
- 3) You will choose from one of the following 4 project topics, described in the following slides:
  - 1) Damped, driven pendulum
  - 2) Logistic map and chaos
  - 3) Astronomical source detection
  - 4) Monte Carlo Integration
- 4) Groups must select a final project topic by **Tues Nov 24**. Detailed instructions for each project will be distributed Thursday Nov 19.
- 5) Groups should be organized through GitHub, invite your group members and us (TAs) as collaborators on your project. Tag us (TAs) when your final project is ready to grade.
- 6) Each python module should indicate which student authored which part of the code.
- 7) Final projects are due (tags must happen by) **Tuesday, December 15, 2020 at 3pm**.



The Astropy Project is a community effort to develop a **common core package** for Astronomy in Python and foster an ecosystem of **interoperable astronomy packages**.

Please remember to **acknowledge and cite** the use of Astropy!

## What's new in **Astropy 3.0?**

Current Version: 3.0.5

# File Download

Download the session\_18.tar file from the class Canvas site.

Open a terminal and use “tar” to expand the tar file into the separate files:

```
3. brant@eduroam-169-233-156-109.ucsc.edu: /Users/brant/Desktop/classes/astr_119/Session 18 (bash)
[17:55:57][brant@eduroam-169-233-156-109:~/Desktop/classes/astr_119/Session 18]$ tar -xvf session_18.tar
x ./_HorseHead.fits
x HorseHead.fits
x ./_M13_blue_0001.fits
x M13_blue_0001.fits
x ./_M13_blue_0002.fits
x M13_blue_0002.fits
x ./_M13_blue_0003.fits
x M13_blue_0003.fits
x ./_M13_blue_0004.fits
x M13_blue_0004.fits
x ./_M13_blue_0005.fits
x M13_blue_0005.fits
[17:56:07][brant@eduroam-169-233-156-109:~/Desktop/classes/astr_119/Session 18]$
```

|                              | Date Created | Date Modified | Modified By     | Size   |   |
|------------------------------|--------------|---------------|-----------------|--------|---|
| astr_119_final_project_1.pdf | Nov 15, 2018 | Nov 15, 2018  | Brant Robertson | 317 KB | ✓ |
| astr_119_final_project_2.pdf | Nov 20, 2018 | Nov 20, 2018  | Brant Robertson | 270 KB | ✓ |
| astr_119_final_project_3.pdf | Nov 15, 2018 | Nov 15, 2018  | Brant Robertson | 85 KB  | ✓ |
| astr_119_final_project_4.pdf | Nov 15, 2018 | Nov 15, 2018  | Brant Robertson | 125 KB | ✓ |
| astr_119.1.pdf               | Sep 27, 2018 | Sep 27, 2018  | Brant Robertson | 3.5 MB | ✓ |
| astr_119.2.pdf               | Oct 2, 2018  | Oct 2, 2018   | Brant Robertson | 6.5 MB | ✓ |
| astr_119.3.pdf               | Oct 4, 2018  | Oct 4, 2018   | Brant Robertson | 2.8 MB | ✓ |



# Astropy Tutorial

## Use Astropy to analyze FITS images

Based on a tutorial by Lia Corrales

```
In [1]: %matplotlib inline
import numpy as np
import matplotlib.pyplot as plt
from astropy.io import fits
```

### Open the FITS file

```
In [3]: fname = "HorseHead.fits"
hdu_list = fits.open(fname)
hdu_list.info()
```

Filename: HorseHead.fits

| No. | Name    | Ver | Type       | Cards | Dimensions | Format                   |
|-----|---------|-----|------------|-------|------------|--------------------------|
| 0   | PRIMARY | 1   | PrimaryHDU | 161   | (891, 893) | int16                    |
| 1   | er.mask | 1   | TableHDU   | 25    | 1600R x 4C | [F6.2, F6.2, F6.2, F6.2] |

# Astropy Tutorial

Generally, the image information is located in the PRIMARY block. The blocks are numbered and can be accessed by indexing `hdu_list`.

```
In [4]: image_data = hdu_list[0].data
```

Our data is now stored as a 2-D numpy array. But how do we know the dimensions of the image? We can simply look at the shape of the array.

```
In [6]: print(type(image_data))  
print(image_data.shape)  
  
<class 'numpy.ndarray'>  
(893, 891)
```

At this point, we can close the FITS file because we've stored everything we wanted to a variable.

```
In [7]: hdu_list.close()
```



# Astropy Tutorial

**Shortcut: use "getdata()" to just read in the image data and close the file.**

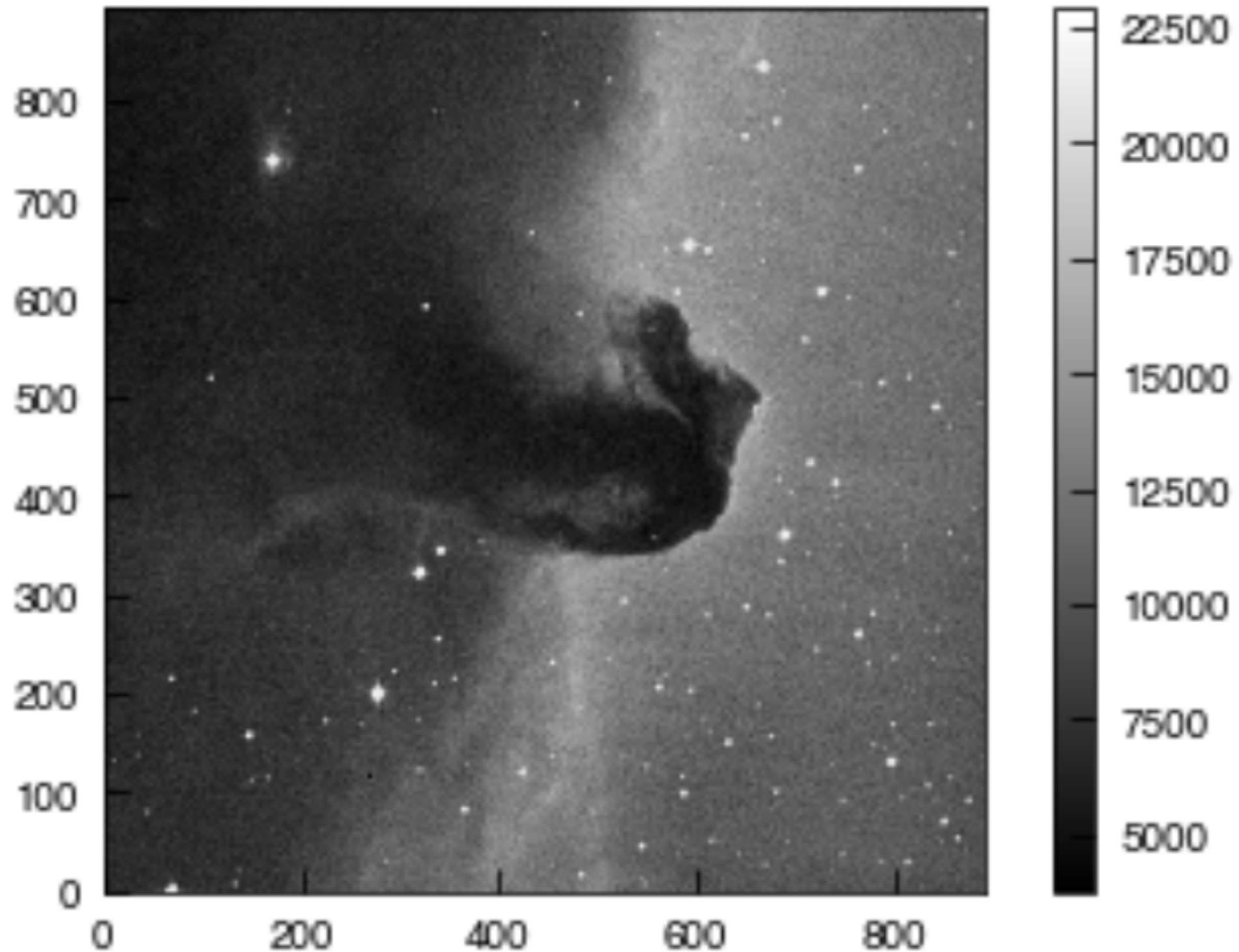
```
In [8]: image_data = fits.getdata(fname)
print(type(image_data))
print(image_data.shape)
```

```
<class 'numpy.ndarray'>
(893, 891)
```

**Let's show the data**

```
In [9]: plt.imshow(image_data, cmap='gray')
plt.colorbar()
```

# Astropy Tutorial



# Astropy Tutorial

**Let's get some basic statistics about our image:**

```
In [10]: print('Min:', np.min(image_data))  
         print('Max:', np.max(image_data))  
         print('Mean:', np.mean(image_data))  
         print('Stdev:', np.std(image_data))
```

```
Min: 3759  
Max: 22918  
Mean: 9831.48167629  
Stdev: 3032.3927542
```

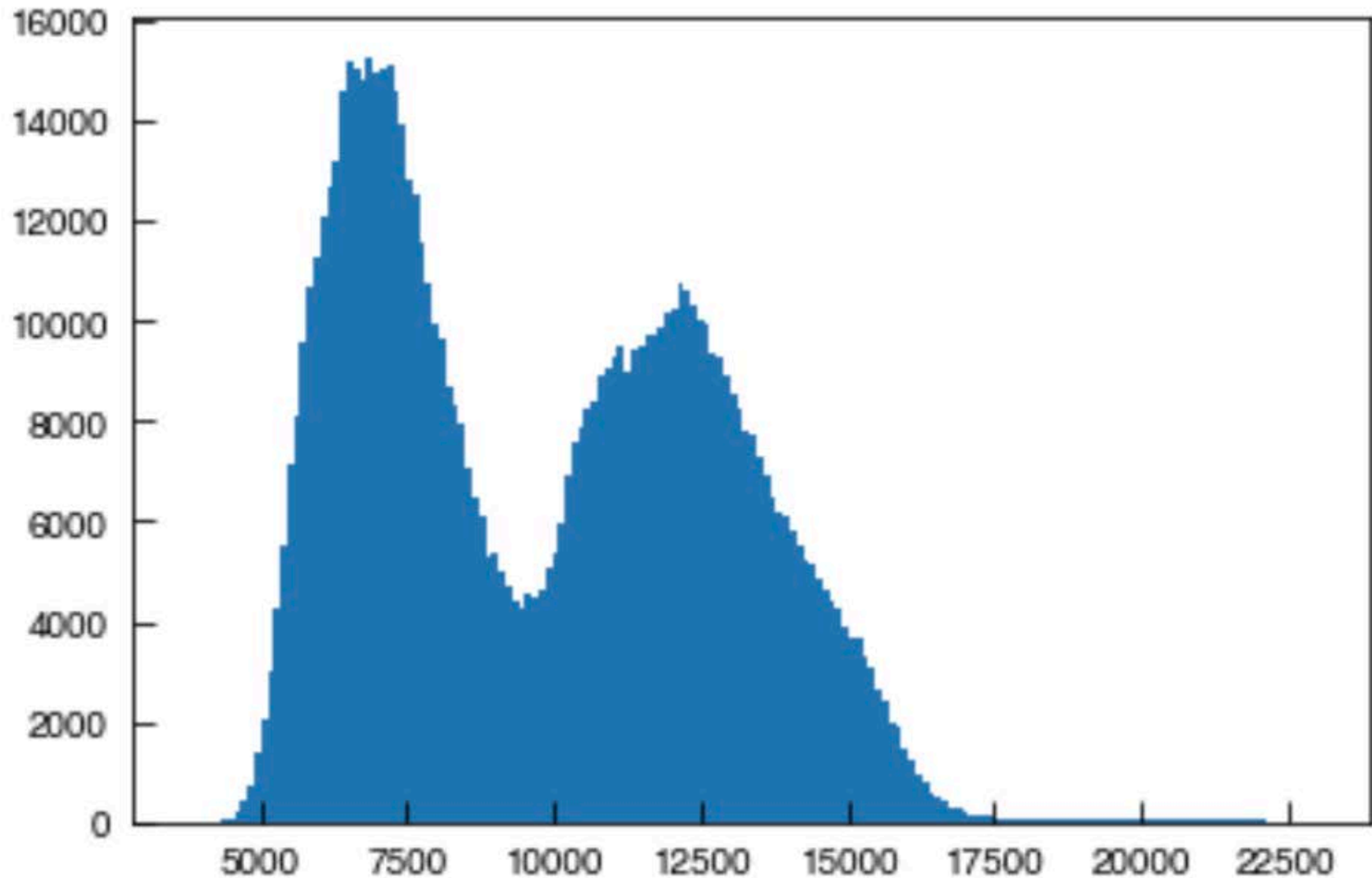
## Plotting a histogram

To make a histogram with `matplotlib.pyplot.hist()`, we'll need to cast the data from a 2-D array to something one dimensional.

In this case, let's use the `ndarray.flatten()` to return a 1-D numpy array.

```
In [11]: histogram = plt.hist(image_data.flatten(), bins='auto')
```

# Astropy Tutorial



# Astropy Tutorial

## Displaying the image with a logarithmic scale

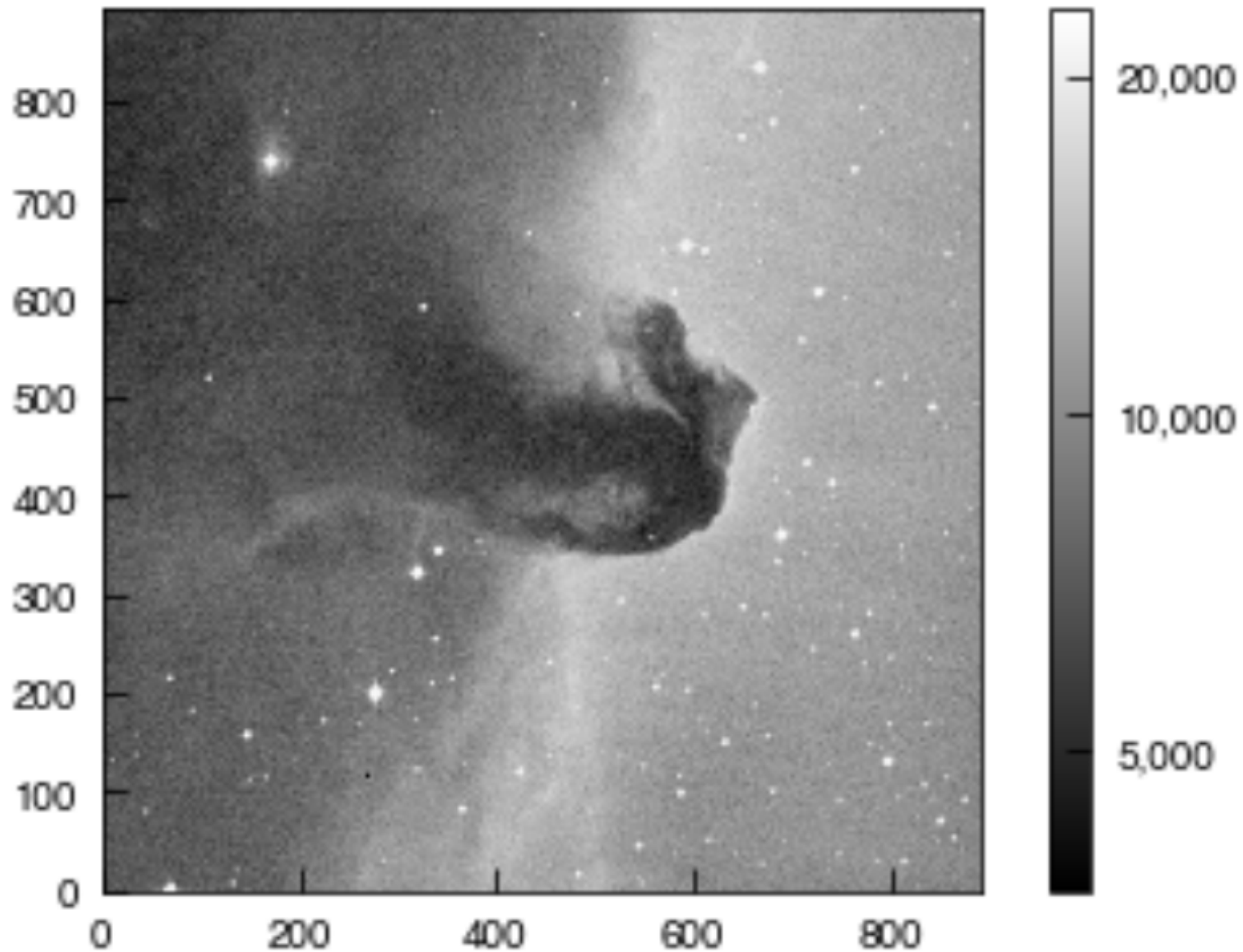
What if we want to use a logarithmic color scale? To do so, we can load the LogNorm object from matplotlib.

```
In [12]: from matplotlib.colors import LogNorm
```

```
In [13]: plt.imshow(image_data, cmap='gray', norm=LogNorm())  
  
# Choose the tick marks based on the histogram above  
cbar = plt.colorbar(ticks=[5.e3, 1.e4, 2.e4])  
cbar.ax.set_yticklabels(['5,000', '10,000', '20,000'])
```

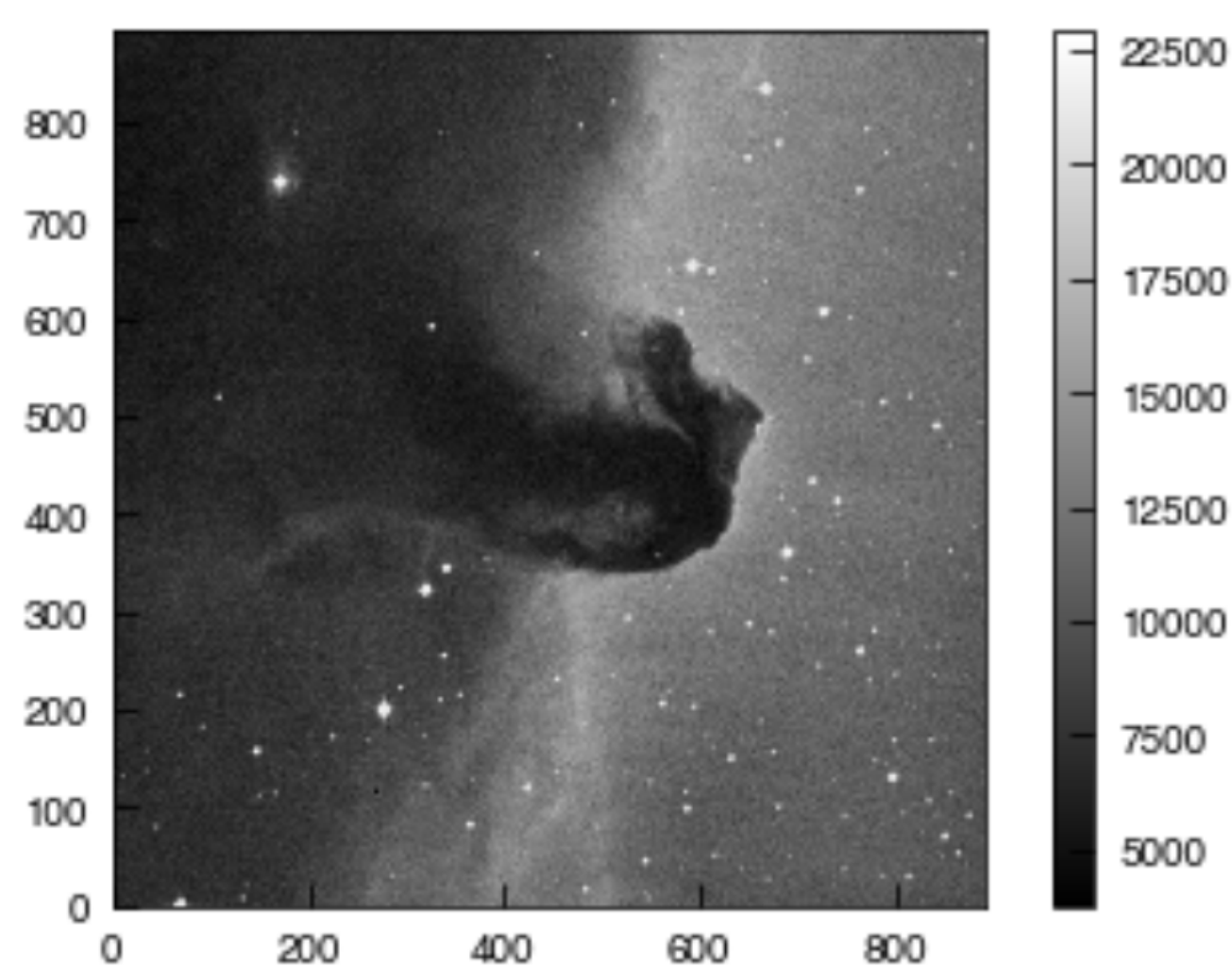


# Astropy Tutorial

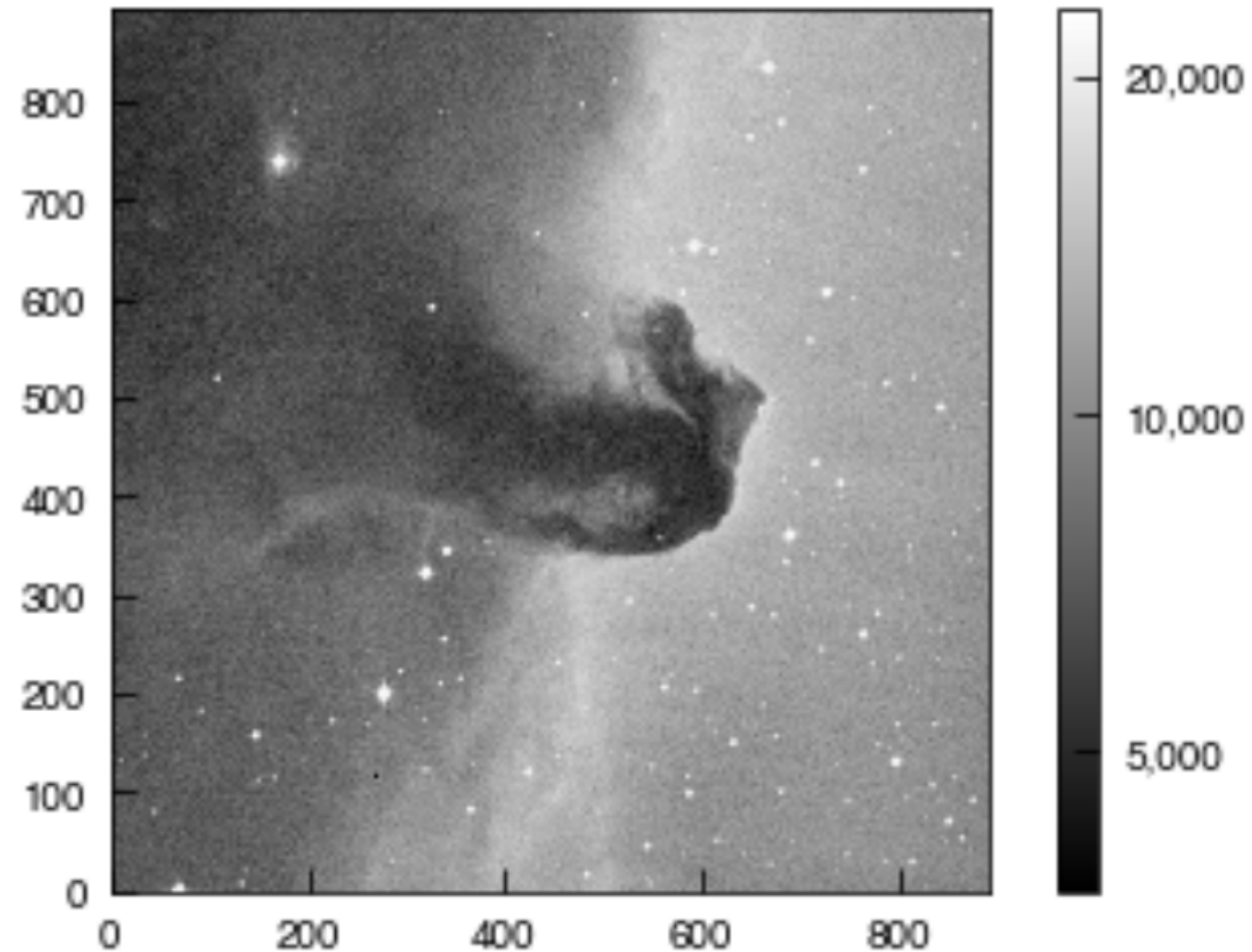




# Astropy Tutorial



Linear



Logarithmic

# Astropy Tutorial

## Stacking Images

Since the noise in an image results from a random process, we use stacking of separate images to improve the signal to noise ratio of objects we observe. Here we are going to stack 5 images of M13 taken with a 10 inch telescope.

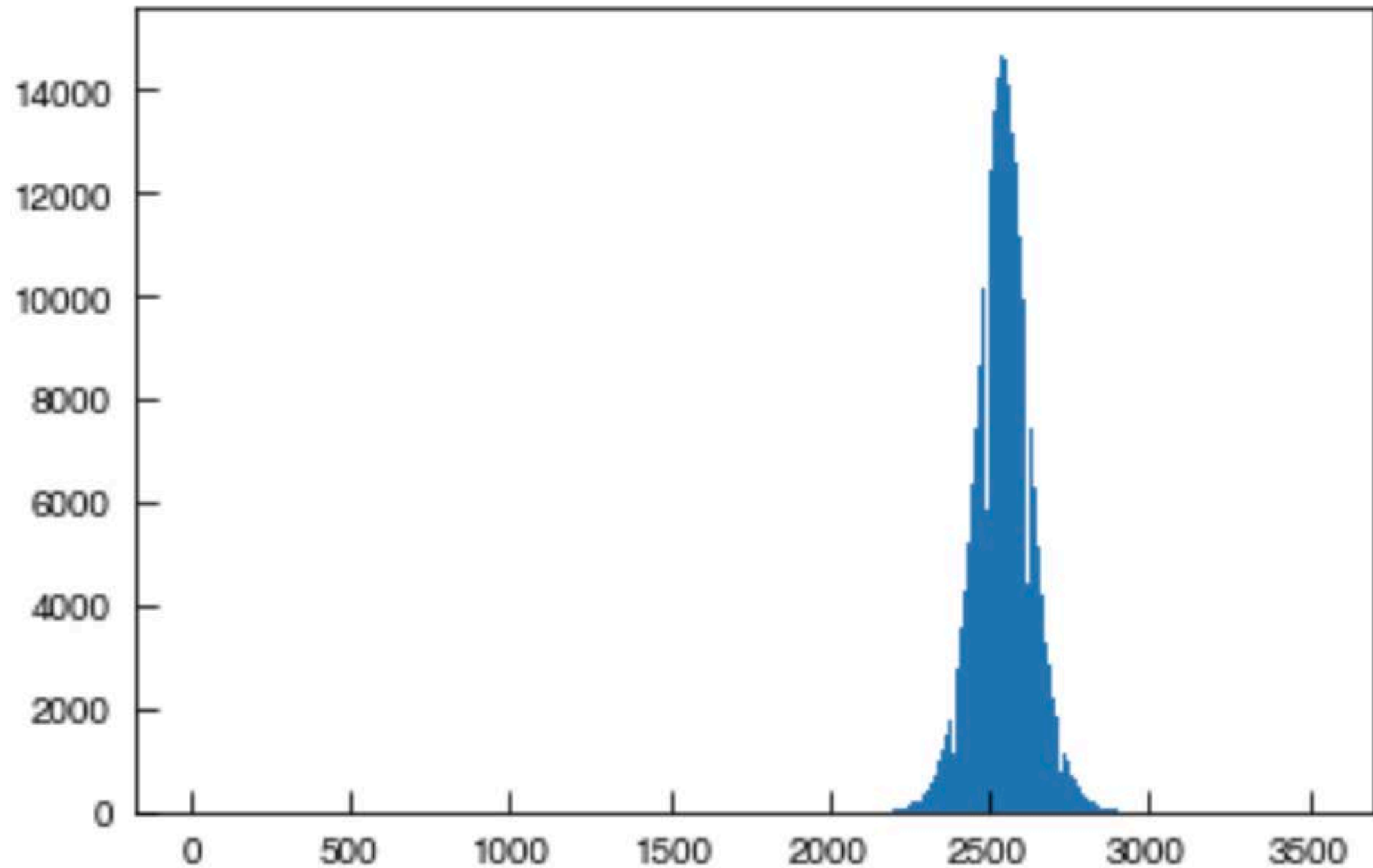
```
]# make a list of filenames
image_list = ['M13_blue_0001.fits', 'M13_blue_0002.fits', 'M13_blue_0003.fits', \
              'M13_blue_0004.fits', 'M13_blue_0005.fits']
```

```
]#make an array of images from the list of images
image_concat = [fits.getdata(image) for image in image_list]
```

```
]#sum the images together
final_image = np.sum(image_concat, axis=0)
```

```
]#plot a histogram of the image pixel values
image_hist = plt.hist(final_image.flatten(), bins='auto')
```

# Astropy Tutorial

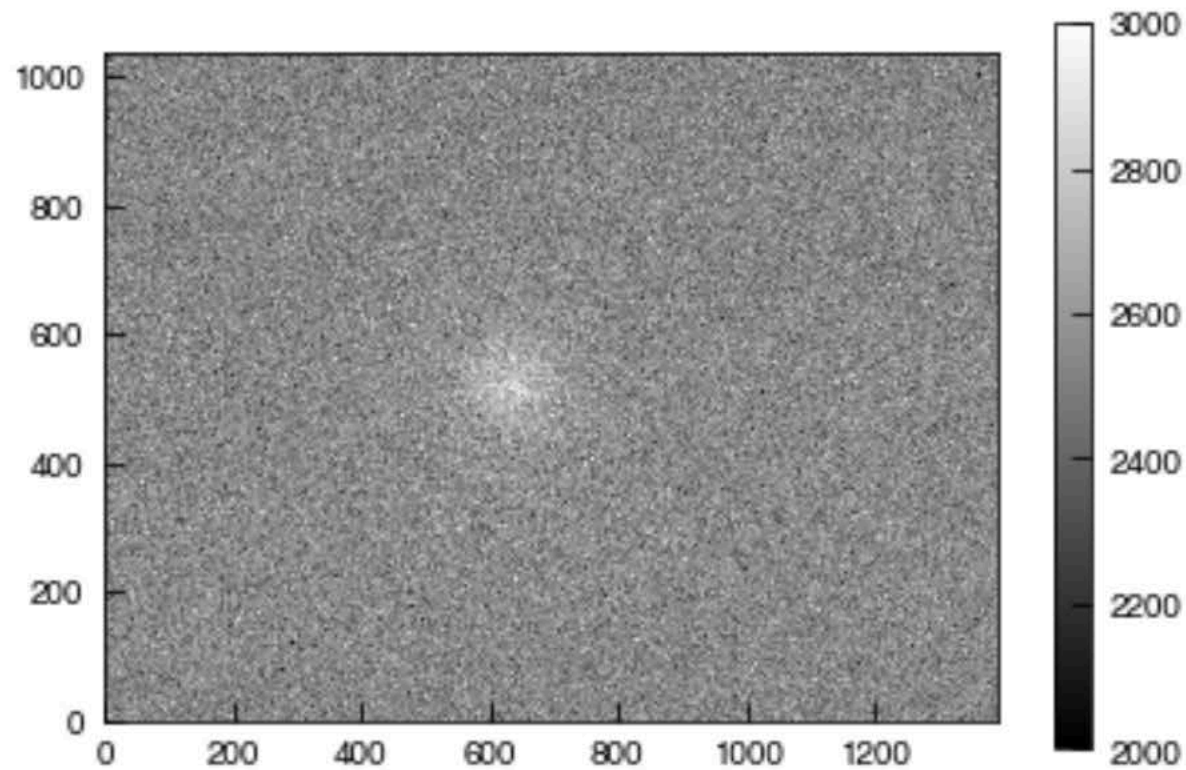


# Astropy Tutorial

We'll use the keywords `vmin` and `vmax` to set limits on the color scaling for `imshow`.

```
In [18]: plt.imshow(final_image, cmap='gray', vmin=2E3, vmax=3E3)  
plt.colorbar()
```

```
Out[18]: <matplotlib.colorbar.Colorbar at 0x118685978>
```





# Astropy Tutorial

## Writing a new FITS file

We can easily do this with the `writeto()` method.

Warning: you'll receive an error if the file you are trying to write already exists. That's why we've set `clobber=True`.

```
outfile = 'stacked_M13_blue.fits'  
hdu = fits.PrimaryHDU(final_image)  
hdu.writeto(outfile, overwrite=True)
```

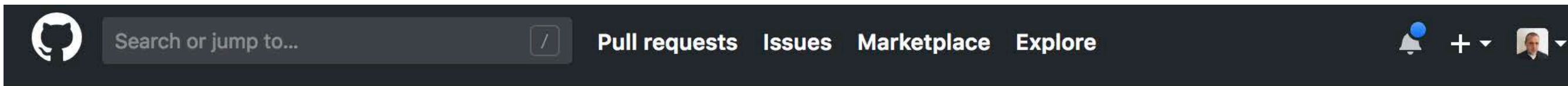
# Final Projects

- 1) You have chosen from one of the following 4 project topics:
  - 1) Damped, driven pendulum
  - 2) Logistic map and chaos
  - 3) Astronomical source detection
  - 4) Monte Carlo Integration
- 2) Groups should be organized through GitHub, invite your group members and us (TA + prof) as collaborators on your project. Tag us (TA+prof when your final project is ready to grade).
- 3) Each python module should indicate which student authored which part of the code.
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# Save Your Work

Make a GitHub project “astr-119-session-17”, and commit the programs you made today.



## Create a new repository

A repository contains all the files for your project, including the revision history.

Owner

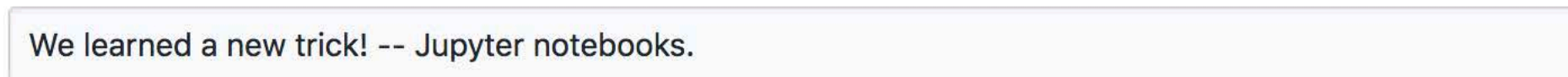


Repository name



Great repository names are short and memorable. Need inspiration? How about **fantastic-spork**.

Description (optional)



☒  **Public**

Anyone can see this repository. You choose who can commit.

☐  **Private**

You choose who can see and commit to this repository.