

# Downscaling of global solar irradiation in R

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## Abstract

A methodology for downscaling solar irradiation from satellite-derived databases is described using R software. Different packages such as `raster`, `parallel`, `solarR`, `gstat`, `sp` and `rasterVis` are considered in this study for improving solar resource estimation in areas with complex topography, in which downscaling is a very useful tool for reducing inherent deviations in satellite-derived irradiation databases, which lack of high global spatial resolution. A topographical analysis of horizon blocking and sky-view is developed with a digital elevation model to determine what fraction of hourly solar irradiation reaches the Earth's surface. Eventually, kriging with external drift is applied for a better estimation of solar irradiation throughout the region analyzed. This methodology has been implemented as an example within the region of La Rioja in northern Spain, and the mean absolute error found is a striking 25.5% lower than with the original database.

**Keywords:** Solar irradiation, R, `raster`, `solarR`, digital elevation model, shade analysis, downscaling

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## 1. Introduction

During the last few years the development of photovoltaic energy has flourished in developing countries with both multi-megawatt power plants and micro installations. However, the scarcity of long-term, reliable solar irradiation data from pyranometers in many of these countries makes it necessary to estimate solar irradiation from other meteorological variables or satellite photographs (Schulz et al., 2009). In such cases, models need to be validated via nearby pyranometer records, since they lack spatial generalization. Thus, in some regions in which there are no pyranometers nearby these models are ruled out as an option and irradiation data must be obtained from satellite estimates. Although satellite-derived irradiation databases such as NASA's Surface meteorology and Solar Energy (SSE)<sup>1</sup>, the National Renewable Energy Laboratory (NREL)<sup>2</sup>, INPE<sup>3</sup>, SODA<sup>4</sup> and the Climate Monitoring Satellite Application Facility (CM SAF)<sup>5</sup> provide wide spatial coverage, only NASA and some CM SAF climate data sets give global coverage, albeit at a reduced spatial resolution (Table 1). The spatial resolutions of

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<sup>1</sup><http://maps.nrel.gov/SWERA>

<sup>2</sup><http://www.nrel.gov/gis/solar.html>

<sup>3</sup><http://www.inpe.br>

<sup>4</sup><http://www.soda-is.com/eng/index.html>

<sup>5</sup><http://www.cmsaf.eu>

satellite estimates are generally in the range of kilometers: they tend to average solar irradiation and omit the impact of topography within each cell. As a result, intra-cell variations can be very significant in areas with local micro-climatic characteristics and in areas with complex topography (which are often one and the same). In this case, the irradiation data might not be accurate enough to enable a photovoltaic installation to be designed. [Perez et al. \(1994\)](#) analyze the spatial behavior of solar irradiation and conclude that the break-even distance from satellite estimates to pyranometers is in the order of 7 km and that variations are hard to estimate for distances greater than 40 km. [Antonanzas-Torres et al. \(2013\)](#) reject ordinary kriging as a spatial interpolation method for solar irradiation in Spain with stations more than 50 km apart in mountainous regions, as a result of the high spatial variability in such areas. The NASA-SSE and CM SAF SIS Climate Data Sets (GHI) provide global coverage with resolutions of  $1 \times 1^\circ$  and  $0.25 \times 0.25^\circ$  (Table 1), which in most latitudes implies a grosser resolution than the previously mentioned 40-50 km.

One of the alternatives for obtaining higher spatial resolution of solar irradiation is the downscaling of satellite estimates. Irradiation downscaling can be based on interpolation kriging techniques when pyranometer records are available, with the implementation of continuous irradiation-related variables such as elevation, sky-view-factor and other meteorological variables as external drifts ([Alsamamra et al., 2009](#); [Batlles et al., 2008](#)). Downscaling is generally based on digital elevation models (DEM) with satellite-derived irradiation data to account for the effect of complex topography. It has previously been applied in mountainous areas such as the Mont Blanc Massif (France) ([Corripio, 2003](#)) and Sierra Nevada (Spain) ([Bosch et al., 2010](#); [Ruiz-Arias et al., 2010](#)) with image resolutions of  $3.5 \times 3.5$  km. However, the NASA-SSE and CM SAF SIS Climate Data Sets are based on much lower resolutions and are the only irradiation datasets in numerous countries where there has been recent interest in solar energy. In this paper, a downscaling methodology of global solar irradiation is explained by means of R software and studied in the region of La Rioja (a very mountainous region in northern Spain). Data from the CM SAF with  $0.03 \times 0.03^\circ$  resolution is considered and then downscaled to a higher resolution ( $200 \times 200$  m). In a second step, *kriging with external drift*, also referred to as *universal kriging*, is applied to interpolate data from 6 on-ground pyranometers in the region, and this downscaled CM SAF data is considered as an explanatory variable. Finally, a downscaled map of annual global solar radiation throughout this region is obtained.

## 2. Data

The CM SAF was funded in 1992 as a joint venture of several European meteorological institutes, with the collaboration of the European Organization for the Exploitation of Meteorological Satellites (EUMETSAT) to retrieve, archive and distribute climate data to be used for climate monitoring and climate analysis ([Posselt et al., 2012](#)). Two categories are provided: operational products and climate data. Operational products are built on data validated with on-ground stations and provided in near-to-present time and climate data are long-term series for evaluating inter-annual variability. This study is built on hourly surface incoming solar radiation and direct irradiation climate data, denoted as SIS and SID by CM SAF respectively, for the year 2005. These climate data are derived from Meteosat first generation satellites (Meteosat 2 to 7, 1982-2005) and validated using on-ground records from the Baseline Surface Radiation Network (BSRN) as a reference. The target accuracy of SIS and SID in hourly means is  $15 \text{ W/m}^2$  ([Posselt et al., 2011](#)), providing a maximum spatial resolution of  $0.03 \times 0.03^\circ$ . In the study, SIS and SID data are selected with spatial resolution of  $0.03 \times 0.03^\circ$ . Data is freely accessible via FTP through

the CM SAF website. Hourly GHI records from SOS Rioja<sup>6</sup>, taken from 6 meteorological stations (shown in Figure 1 and Table 2) in 2005 serve as complementary measurements for downscaling within the region studied. These stations have First Class pyranometers (according to ISO 9060) with uncertainty levels of 5% in daily totals. These data are filtered from spurious, assuming when relevant the average between the previous and following hourly measurements. The digital elevation model (DEM) is also freely obtained from product MDT-200 by the ©Spanish Institute of Geography<sup>7</sup> with a spatial resolution of 200x200 m.

### 3. Method

This section describes the methodology proposed. Figure 2 displays the method diagram using red ellipses and lines for data sources, blue ellipses and lines for derived rasters (results), and black rectangles and lines for operations.

#### 3.1. Irradiation decomposition

Initially, diffuse horizontal irradiation ( $DHI$ ) is obtained from the difference between global horizontal irradiation ( $GHI$ ) and beam horizontal irradiation ( $BHI$ ) rasters, previously obtained from CM SAF.  $DHI$  and  $BHI$  are firstly disaggregated from the original gross resolution ( $0.03 \times 0.03^\circ$ ) into the DEM resolution (200x200 m), leading to similar values remaining in disaggregated pixels to the original gross resolution pixel. In a second step,  $DHI$  can be divided in two components: isotropic diffuse irradiation ( $DHI_{iso}$ ), and anisotropic diffuse irradiation ( $DHI_{ani}$ ) as per the model by Hay & McKay (Hay and McKay, 1985) (Equation 1). This model is based on the anisotropy index ( $k_1$ ), defined as the ratio of the beam irradiance ( $B(0)$ ) to the extra-terrestrial irradiance ( $B_0(0)$ ), as shown in Equation 2. High  $k_1$  values are typical in clear sky atmospheres, while low  $k_1$  values are frequent in overcast atmospheres and those with a high aerosol density.

$$DHI = DHI_{iso} + DHI_{ani} \quad (1)$$

$$k_1 = \frac{B(0)}{B_0(0)} \quad (2)$$

The  $DHI_{iso}$  accounts for the incoming diffuse irradiation portion from an isotropic sky, and is more significant on very cloudy days (Equation 3).

$$DHI_{iso} = DHI \cdot (1 - k_1) \quad (3)$$

$DHI_{ani}$ , also denoted as circumsolar diffuse irradiation, considers the incoming portion from the circumsolar disk and can be analyzed as beam irradiation (Perpiñán-Lamigueiro, 2013) (Equation 4).

$$DHI_{ani} = DHI \cdot k_1 \quad (4)$$

<sup>6</sup><http://www.larioja.org/npRioja/default/defaultpage.jsp?idtab=442821>

<sup>7</sup><http://www.ign.es>

### 3.2. Sky view factor and horizon blocking

Topographical analysis is performed accounting for the visible sky sphere (sky view) and horizon blocking. The  $DHI_{iso}$  is directly dependent on the sky-view factor (SVF), which computes the proportion of visible sky related to a flat horizon. The sky-view factor is considered in earlier irradiation assessments (Corripio, 2003; Ruiz-Arias et al., 2010). It is calculated in each DEM pixel by considering 72 vectors (separated by  $5^\circ$  each) and evaluating the maximum horizon angle ( $\theta_{hor}$ ) over 20 km in each vector (Equation 5). The  $\theta_{hor}$  stands for the maximum angle between the altitude of a location and the elevation of the group of points along each vector, related to a horizontal plane on the location. Locations without horizon blocking have SVFs close to 1, which means a whole visible semi-sphere of sky.

$$SVF = 1 - \int_0^{2\pi} \sin^2 \theta_{hor} d\theta \quad (5)$$

Eventually, the downscaled  $DHI_{iso}$  ( $DHI_{iso,down}$ ) is computed with Equation 6.

$$DHI_{iso,down} = DHI_{iso} \cdot SVF \quad (6)$$

Horizon blocking is analyzed by evaluating the solar geometry in 15 minute samples, particularly the solar elevation ( $\gamma_s$ ) and the solar azimuth ( $\psi_s$ ). Secondly, the mean hourly  $\gamma_s$  and  $\psi_s$  (from those 15 minute rasters) are calculated and then disaggregated as explained above for  $DHI$  and  $BHI$  rasters. The decision to solve the solar geometry with low resolution rasters enables computation time to be reduced significantly without penalizing the results. The  $\theta_{hor}$  corresponding to each  $\psi_s$  is compared with the  $\theta_{zs}$ . As a consequence, if the  $\theta_{zs}$  is greater than the  $\theta_{hor}$ , then there is horizon blocking on the surface analyzed and therefore,  $BHI$  and  $DHI_{ani}$  are blocked. Finally, the sum of  $DHI_{ani,down}$ ,  $DHI_{iso,down}$  and  $BHI_{iso,down}$  constitutes the downscaled global horizontal irradiation  $GHI_{down}$ .

### 3.3. Post-processing: kriging with external drift

The fact that this downscaling accounts for the irradiation loss due to horizon blocking and the sky-view factor leads us to introduce a trend in estimates (lowering them) compared to the original data (gross resolution data). However, satellite-derived irradiation data implicitly considers shade, as a consequence of the lower albedo recorded in these zones, although it is later averaged over the pixel.  $GHI_{down}$  can be considered as a useful bias of the behavior of solar irradiation within the region studied. *Universal kriging* or *kriging with external drift* (KED) includes information from exhaustively-sampled explanatory variables in the interpolation. As a result,  $GHI_{down}$  is considered as the explanatory variable for interpolating measured irradiation data from on-ground calibrated pyranometers, which is denoted as *post-processing*.  $GHI_{down}$  is correlated with the DEM as a consequence of the major influence of horizon blocking with topography, estimations can be derived by separating the deterministic ( $\hat{m}(\mathbf{s}_\theta)$ ) and stochastic components ( $\hat{\epsilon}(\mathbf{s}_\theta)$ ) (Equations 7 and 8).

$$\hat{z}(\mathbf{s}_\theta) = \hat{m}(\mathbf{s}_\theta) + \hat{\epsilon}(\mathbf{s}_\theta) \quad (7)$$

$$\hat{z}(\mathbf{s}_\theta) = \sum_{k=0}^p \hat{\beta}_k q_k(\mathbf{s}_\theta) + \sum_{i=1}^n \lambda_i \epsilon(\mathbf{s}_i) \quad (8)$$

where  $\hat{\beta}_k$  are the estimated coefficients of the deterministic model,  $q_k(\mathbf{s}_\theta)$  are the auxiliary predictors obtained from the fitted values of the explanatory variable at the new location,  $\lambda_i$  are

the kriging weights determined by the spatial dependence structure of the residual, and  $\epsilon(\mathbf{s}_i)$  are the residual at location  $\mathbf{s}_i$  (?).

The semivariogram is a function defined as Equation 9 based on a constant variance of  $\epsilon$  and also on the assumption that spatial correlation of  $\epsilon$  depends on the distance amongst instances ( $\mathbf{h}$ ) rather than on their position (Pebesma, 2004).

$$\gamma(\mathbf{h}) = \frac{1}{2}E(\epsilon(\mathbf{s}) - \epsilon(\mathbf{s} + \mathbf{h}))^2 \quad (9)$$

Given that the above sample variogram only collates estimates from observed points, a fitting model of this variogram is generally considered to extrapolate the spatial behavior of observed points to the area studied. In the literature different variogram functions are commonly defined such as the exponential, Gaussian or spherical models. Along these lines, different parameters such as the sill, range and nugget of the model must be adjusted to best fit the sample variogram (Hengl, 2009). The nugget effect, generally associated with intrinsic micro-variability and measurement error, models the discontinuity of the variogram at the source. It must be highlighted that when the nugget effect is recorded, kriging differs from a regular interpolation and as a result estimates are different from measured values. The variogram model of solar horizontal irradiation is evaluated in Spain, and the conclusion reached is that a pure nugget fitting behaves best, which implies no spatial auto-correlation on residuals (Antonanzas-Torres et al., 2013).

## 4. Implementation

The method proposed is applied in the region of La Rioja (northern Spain). Figure 3 shows the corresponding annual global horizontal irradiation from CM SAF with resolution  $0.03 \times 0.03^\circ$ .

### 4.1. Packages

The downscaling described in this paper has been implemented using the free software environment R (R Development Core Team, 2013) and various contributed packages:

- raster (Hijmans and van Etten, 2013) for spatial data manipulation and analysis.
- solaR (Perpiñán-Lamigueiro, 2012) for solar geometry.
- gstat (Pebesma and Graeler, 2013) and sp (Pebesma et al., 2013) for geostatistical analysis.
- parallel for multi-core parallelization.
- rasterVis (Perpiñán-Lamigueiro and Hijmans, 2013) for spatial data visualization methods.

```
R> library(sp)
R> library(raster)
R> rasterOptions(todisk=FALSE)
R> rasterOptions(chunksize = 1e+06, maxmemory = 1e+07)
R> library(maptools)
R> library(gstat)
R> library(lattice)
R> library(latticeExtra)
R> library(rasterVis)
R> library(solaR)
R> library(parallel)
```

## 180 4.2. Data

181 Satellite data can be freely downloaded after registration from CM SAF<sup>8</sup> by going to the data  
182 access area, selecting *web user interface* and *climate data sets* and then choosing the hourly climate  
183 data sets named *SIS* (Global Horizontal Irradiation) and *SID* (Beam Horizontal Irradiation) for  
184 2005. Both rasters are projected to the UTM projection for compatibility with the DEM.

```
185 R> projUTM <- CRS('+proj=utm+zone=30')
186 R> projLonLat <- CRS('+proj=longlat+ellps=WGS84')
187
188 R> listFich <- dir(pattern='SIShm2005')
189 R> stackSIS <- stack(listFich)
190 R> stackSIS <- projectRaster(stackSIS, crs=projUTM)
191
192 R> listFich <- dir(pattern='SIDhm2005')
193 R> stackSID <- stack(listFich)
194 R> stackSID <- projectRaster(stackSID, crs=projUTM)
```

195 We compute the annual global irradiation, which will be used as a reference for subsequent  
196 steps.

```
197 R> SISa2005 <- calc(stackSIS, sum, na.rm=TRUE)
```

198 The Spanish Digital Elevation Model can be obtained after registration from the ©Spanish  
199 Institute of Geography<sup>9</sup> by going to the *free download of digital geographic information for non-*  
200 *commercial use* area, and then cropping to the region analyzed (La Rioja). As stated above, this  
201 DEM uses the UTM projection.

```
202 R> elevSpain <- raster('elevSpain.grd')
203 R> elev <- crop(elevSpain, extent(479600, 616200, 4639600, 4728400))
204 R> names(elev) <- 'elev'
```

## 205 4.3. Sun geometry

206 The first step is to compute the sun angles (height and azimuth) and the extraterrestrial so-  
207 lar irradiation for each cell of the CM SAF rasters. The function `calcSol` from the `solar` pack-  
208 age calculates the daily and intradaily sun geometry. By means of this function and `overlay`  
209 from the `raster` package, three multilayer raster objects are generated with the sun geometry  
210 needed for the next steps. For the sake of brevity we show only the procedure for extraterres-  
211 trial solar irradiation. The sun geometry is calculated with the resolution of CM SAF. First, it  
212 is defined a function to extract the hour for aggregation, choose the annual irradiation raster as  
213 reference, and define a raster with longitude and latitude coordinates.

```
214 R> hour <- function(tt) as.POSIXct(trunc(tt, 'hours'))
215
216 R> r <- SISa2005
217
218 R> latlon <- stack(init(r, v='y'), init(r, v='x'))
219 R> names(latlon) <- c('lat', 'lon')
```

220 The extraterrestrial irradiation is calculated with 5-min samples. Each point is a column of  
221 the data frame `locs`. Its columns are traversed with `lapply`, so for each point of the raster

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<sup>8</sup>[www.cmsaf.eu](http://www.cmsaf.eu)

<sup>9</sup><http://www.ign.es>

object a time series of extraterrestrial solar irradiation is computed. The result, B05min, is a RasterBrick object with a layer for each element of the time index BTi, which is aggregated to an hourly raster with zApply and transformed to the UTM projection.

```

225 R> BTi <- seq(as.POSIXct('2005-01-01_00:00:00'),
226 + as.POSIXct('2005-12-31_23:55:00'), by='5_min')
227
228 R> B05min <- overlay(latlon, fun=function(lat, lon){
229 + locs <- as.data.frame(rbind(lat, lon))
230 + b <- lapply(locs, function(p){
231 +
232 + hh <- local2Solar(BTi, p[2])
233 + sol <- calcSol(p[1], BTi=hh)
234 + Bo0 <- as.data.frameI(sol)$Bo0
235 + Bo0 })
236 + res <- do.call(rbind, b)})
237
238 R> B05min <- setZ(B05min, BTi)
239 R> names(B05min) <- as.character(BTi)
240
241 R> B0h <- zApply(B05min, by=hour, fun=mean)
242 R> projectRaster(B0h, crs=projUTM)

```

#### 4.4. Irradiation components

The CMSAF rasters must be transformed to the higher resolution of the DEM (UTM 200x200 m). Because of the differences in pixel geometry between DEM (square) and irradiation rasters (rectangle) the process is performed in two steps.

The first step increases the spatial resolution of the irradiation rasters to a similar and also larger pixel size than the DEM with disaggregated data, where *sf* is the scale factor. The second step post-processes the previous step by means of a bilinear interpolation which resamples the raster layer and achieves the same DEM resolution (*resample*). This two-step disaggregation prevents the loss of the original values of the gross resolution raster that would be directly interpolated with the one-step disaggregation.

```

253 R> sf <- res(stackSID)/res(elev)
254
255 R> SIDd <- disaggregate(stackSID, sf)
256 R> SIDdr <- resample(SIDd, elev)
257
258 R> SISd <- disaggregate(stackSIS, sf)
259 R> SISdr <- resample(SISd, elev)

```

On the other hand, the diffuse irradiation is obtained from the global and beam irradiation rasters. The two components of the diffuse irradiation, isotropic and anisotropic, can be separated with the anisotropy index, computed as the ratio between beam and extraterrestrial irradiation.

```

264 R> Difdr <- SISdr - SIDdr
265
266 R> B0hd <- disaggregate(B0h, sf)
267 R> B0hdr <- resample(B0hd, elev)
268
269 R> k1 <- SIDdr/B0hdr

```

```

270
271 R> Difiso <- (1-k1) * Difdr
272 R> Difani <- k1 * Difdr

```

#### 273 4.5. Sky view factor and horizon blocking

##### 274 4.5.1. Horizon angle

275 The maximum horizon angle required for the horizon blocking analysis and to derive the  
 276 SVF is obtained with the next code. The alpha vector is visited with `mclapply` (using parallel  
 277 computing). For each direction angle (elements of this vector) the maximum horizon angle is  
 278 calculated for a set of points across that direction from each of the locations defined in `xyelev`  
 279 (derived from the DEM raster and transformed in the matrix `locs` visited by rows).

```

280 R> xyelev <- stack(init(elev, v='x'),
281 + init(elev, v='y'),
282 + elev)
283 R> names(xyelev) <- c('x', 'y', 'elev')
284
285
286 R> inc <- pi/36
287 R> alfa <- seq(-0.5*pi, (1.5*pi-inc), inc)
288
289 R> locs <- as.matrix(xyelev)

```

290 Separations between the source locations and points along each direction are defined by  
 291 `resD`, the maximum resolution of the DEM, `d`, maximum distance to visit, and consequently in  
 292 the vector `seps`.

```

293 R> resD <- max(res(elev))
294
295 R> d <- 20000
296 R> seps <- seq(resD, d, by=resD)

```

297 The elevation (`z1`) of each point in `xyelev` is converted into the horizon angle: the largest of  
 298 these angles is the horizon angle for that direction. The result of each apply step is a matrix,  
 299 which is used to fill in a `RasterLayer` (`r`). The result of `mclapply` is a list, `hor`, of `RasterLayer`  
 300 which can be converted into a `RasterStack` with `stack`. Each layer of this `RasterStack` corre-  
 301 sponds to a different direction.

```

302 R> hor <- mclapply(alfa, function(ang){
303 + h <- apply(locs, 1, function(p){
304 + x1 <- p[1]+cos(ang)*seps
305 + y1 <- p[2]+sin(ang)*seps
306 + p1 <- cbind(x1,y1)
307 + z1 <- elevSpain[cellFromXY(elevSpain,p1)]
308 + hor <- r2d(atan2(z1-p[3], seps))
309 + maxHor <- max(hor[which.max(hor)], 0)
310 + })
311 + r <- raster(elev)
312 + r[] <- matrix(h, nrow=nrow(r), byrow=TRUE)
313 + r}, mc.cores=8)
314 R> horizon <- stack(hor)

```

315 This operation is very time-consuming as it is necessary to work with high resolution files.  
 316 Computation time can be decreased by increasing the sampling space (200 m) or the sectoral  
 317 angle ( $5^\circ$ ) or by reducing the maximum distance (20 km).



#### 318 4.5.2. Horizon blocking

319 Horizon blocking is analyzed by evaluating the solar geometry in 15 minute samples, partic-  
320 ularly the solar elevation and azimuth angles from the original irradiation raster. Secondly, the  
321 hourly averages are calculated, disaggregated and post-processed as explained above for the  
322 irradiation rasters. The decision to solve the solar geometry with low resolution rasters enables  
323 a significant reduction to be obtained in computation time without penalizing the results.

324 First, the azimuth raster is cut into different classes according to the alpha vector (direc-  
325 tions). The values of the horizon raster corresponding to each angle class are extracted using  
326 `stackSelect`.

```
327 R> idxAngle <- cut(AzShr, breaks=r2d(alfa))  
328 R> AngAlt <- stackSelect(horizon, idxAngle)
```

329 The number of layers of `AngAlt` is the same as `idxAngle` and can therefore be used for  
330 comparison with the solar height angle, `AlShr`. If `AngAlt` is greater, there is horizon blocking  
331 (`dilogical=0`).

```
332 R> dilogical <- ((AngAlt-AlShr) < 0)
```

333 With this binary raster, beam irradiation and diffuse anisotropic irradiation can be corrected  
334 with horizon blocking.

```
335 R> Dirh <- SIDdr * dilogical  
336 R> Difani <- Difani * dilogical
```

#### 337 4.5.3. Sky view factor

338 The sky-view factor can be easily computed from the horizon object with the equation pro-  
339 posed above. This factor corrects the isotropic component of the diffuse irradiation.

```
340 R> SVFRuizArias <- calc(horizon, function(x) sin(d2r(x))^2)  
341 R> SVF <- 1 - mean(SVFRuizArias)  
342  
343 R> Difiso <- Difiso * SVF
```

344 Finally, the global irradiation is the sum of the three corrected components, beam and anisotropic  
345 diffuse irradiation including horizon blocking, and isotropic diffuse irradiation with the sky  
346 view factor.

```
347 R> GHIh <- Difanis + Difiso + Dirh  
348 R> GHI2005a <- calc(GHIh, fun=sum)
```

#### 349 4.6. Kriging with external drift

350 The downscaled irradiation rasters can be improved by using kriging with external drift.  
351 Irradiation data from on-ground meteorological stations is interpolated with the downscaled  
352 irradiation raster as the explanatory variable. To define the variogram here we use the results  
353 previously published in ([Antonanzas-Torres et al., 2013](#)).

```
354 R> load('Stations.RData')  
355 R> UTM <- SpatialPointsDataFrame(Stations[,c(2,3)], Stations[, -c(2,3)],  
356 + proj4string=CRS('+proj=utm+zone=30+ellps=WGS84'))  
357  
358  
359 R> vgmCMSAF <- variogram(GHImed ~ GHICMSAF, UTM)  
360 R> fitvgmCMSAF <- fit.variogram(vgmCMSAF, vgm(model='Nug'))
```

```

361
362 R> gModel <- gstat(NULL, id='G0yKrig',
363 + formula= GHImed ~ GHICmsaf,
364 + locations=UTM, model=fitvgmCMSAF)
365
366 R> names(GHI2005a) <- 'GHICmsaf'
367 R> G0yKrig <- interpolate(GHI2005a, gModel, xyOnly=FALSE)

```

#### 368 4.7. Analysis of the results

369 Figure 3 shows the annual GHI as per CM SAF with the gross resolution analyzed (0.03x0.03°)  
370 and Figures 4 and 5 show the downscaled maps (200x200 m) without and with the KED.

##### 371 4.7.1. Model performance

372 In order to evaluate the performance of the method proposed, relative differences evalu-  
373 ated with station measurements are shown in 6. As can be deduced from this Figure, relative  
374 differences are smaller in *downscaling with KED* than in CM SAF or *downscaling without KED*,  
375 at  $\pm 15\%$ . The mean absolute error (MAE) and root mean square error (RMSE), described in  
376 Equations 10 and 11, are used as indicators of model performance.

$$MAE = \frac{\sum_{i=1}^n |x_{est} - x_{meas}|}{n} \quad (10)$$

$$RMSE = \sqrt{\frac{\sum_{i=1}^n (x_{est} - x_{meas})^2}{n}} \quad (11)$$

377 where  $n$  is number of stations and  $x_{est}$  and  $x_{meas}$  the annual estimated and measured irradi-  
378 ation, respectively.

379 Table 3 shows the MAE and RMSE obtained with CM SAF and with the methodology pro-  
380 posed before and after the KED. The KED leads to a significant improvement in estimates: the  
381 MAE is down by 25.5% and the RMSE by 27.4% compared to CM SAF.

382 The higher MAE recorded in station locations in CM SAF and *downscaling without KED* is  
383 also explained in the irradiation maps shown in Figures 3 and 4. The  $GHI_{annual}$  is lowered too  
384 far in certain regions of the area studied with *downscaling without KED* compared to  $GHI_{down,ked}$ ,  
385 which is also shown in Figure 6.

##### 386 4.7.2. Zonal variability

387 The intrapixel variability due to the downscaling procedure is indicative of the importance  
388 of the topography as an attenuator of solar irradiation. As a result, this zonal variability is  
389 higher in pixels with complex topographies and downscaling is more useful. Figure 7 shows  
390 the relative difference between downscaling with KED and CM SAF. As might be deduced, CM  
391 SAF over-estimates GHI in this region by between 11 and 22%. Figures 8 and 9 display the  
392 standard deviations of the downscaled maps within each cell of the original CM SAF raster  
393 (0.03x0.03°). The zonal function from the raster library permits this calculation, explaining  
394 the intrinsic variability of solar radiation within gross resolution pixels. Consequently, in those  
395 pixels with higher standard deviations there will be greater variability. Figure 9 shows how  
396 the KED method smooths the deviation within pixels and also the range of solar irradiation in  
397 the region (Figures 4 and 5).

## 398 5. Concluding comments

399 A methodology for downscaling solar irradiation is described and presented using R soft-  
400 ware. This methodology is useful for increasing the accuracy and spatial resolution of gross  
401 resolution satellite-estimates of solar irradiation.

402 It has been proved that areas whose topography is complex show greater differences with  
403 the original gross resolution data as a consequence of horizon blocking and lower sky-view  
404 factors, so downscaling is highly recommended in these areas.

405 *Kriging with external drift* with the *gstat* package has proved very useful in downscaling so-  
406 lar irradiation when on-ground registers are available and an explanatory variable is provided.

407 This methodology is implemented as an example in the region of La Rioja in northern Spain,  
408 and striking reductions of 25.5% and 27.4% in MAE and RMSE are obtained compared to the  
409 original gross resolution database. The high repeatability of this methodology and the reduction  
410 in errors obtained might be also very useful in the downscaling of meteorological variables  
411 other than solar irradiation.

## 412 Session information

413 The results discussed in this paper were obtained in a session with these characteristics:

- 414 • R version 2.15.2 (2012-10-26), x86\_64-apple-darwin9.8.0
- 415 • Locale: es\_ES.UTF-8/es\_ES.UTF-8/es\_ES.UTF-8/C/es\_ES.UTF-8/es\_ES.UTF-8
- 416 • Base packages: base, datasets, graphics, grDevices, grid, methods, parallel,  
417 stats,utils
- 418 • Other packages: foreign 0.8-51, gstat 1.0-16, hexbin 1.26.0, lattice 0.20-15,  
419 latticeExtra 0.6-19, maptools 0.8-14, raster 2.1-16, rasterVis 0.20-01,  
420 RColorBrewer 1.0-5, rgdal 0.8-01, solar 0.33, sp 1.0-8, zoo 1.7-9
- 421 • Loaded via a namespace (and not attached): intervals 0.14.0, spacetime 1.0-4,  
422 tools 2.15.2, xts 0.9-3

## 423 Acknowledgements

424 We are indebted to the University of La Rioja (fellowship FPI2012) and the Research Institute  
425 of La Rioja (IER) for funding parts of this research.

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Database	Product	Spatial coverage		Spatial resolution	Temporal coverage	Temporal resolution
CM SAF	SIS Climate Data Set (GHI)	Global		0.25x0.25°	1982-2009	Daily means
CM SAF	SIS Climate Data Set (GHI)	70S-70N,	70W-70E	0.03x0.03°	1983-2005	Hourly means
CM SAF	SID Climate Data Set (BHI)	70S-70N,	70W-70E	0.03x0.03°	1983-2005	Hourly means
SODA	Helioclim 3 v2 and v3 (GHI)	66S-66N,66W-66E		5km	2005	15 minutes
SODA	Helioclim 3 v2 and v3 (GHI)	66S-66N,66W-66E		5km	2005	15 minutes
NREL	GHI Moderate resolution	Central and South America, Africa, India, East Asia		40x40km	1985-1991	Monthly means of daily GHI
NASA	SSE	Global		1x1°	1983-2005	Daily means

Table 1: Summary of solar irradiation databases

#	Name	Net.	Lat.(°)	Long.(°)	Alt.	$GHI_a$
1	Ezcaray	SOS	42.33	-3.00	1000	1479
2	Logroño	SOS	42.45	-2.74	408	1504
3	Moncalvillo	SOS	42.32	-2.61	1495	1329
4	San Roman	SOS	42.23	-2.45	1094	1504
5	Ventrosa	SOS	42.17	-2.84	1565	1277
6	Yerga	SOS	42.14	-1.97	1235	1448

Table 2: Summary of the meteorological stations selected.

	CM SAF	without KED	with KED
MAE	101.35	175.63	75.54
RMSE	118.65	196.53	86.18

Table 3: Summary of errors obtained in  $kWh/m^2$ .

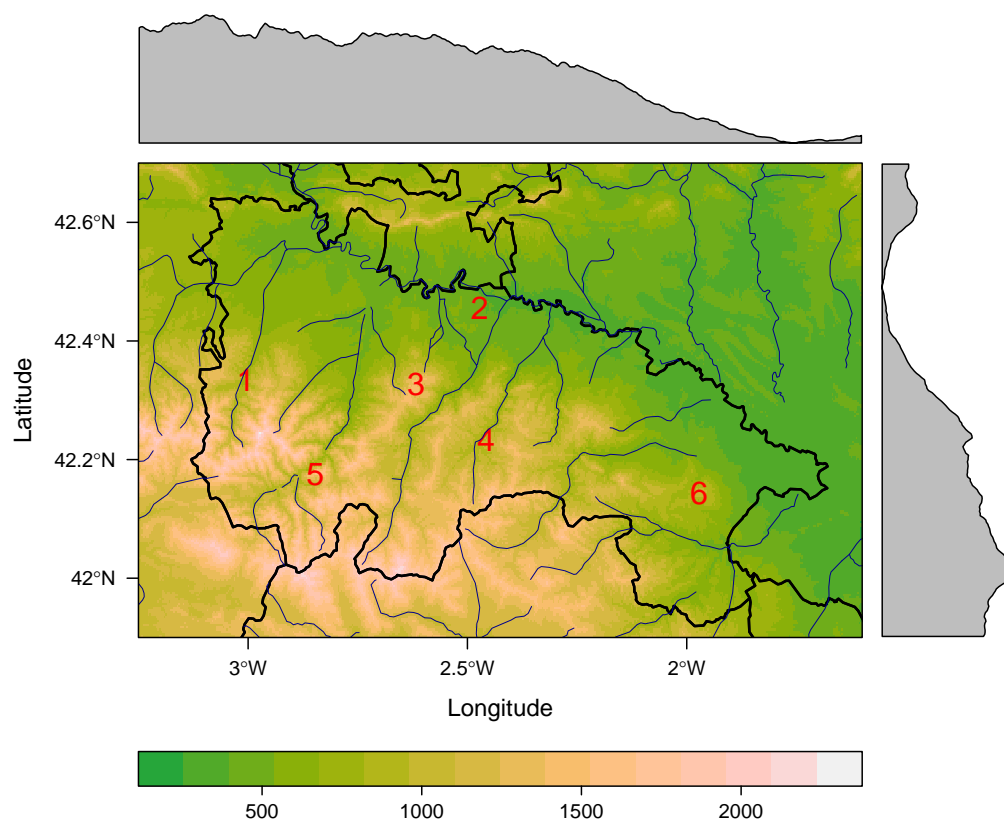


Figure 1: Region analyzed and meteorological stations considered

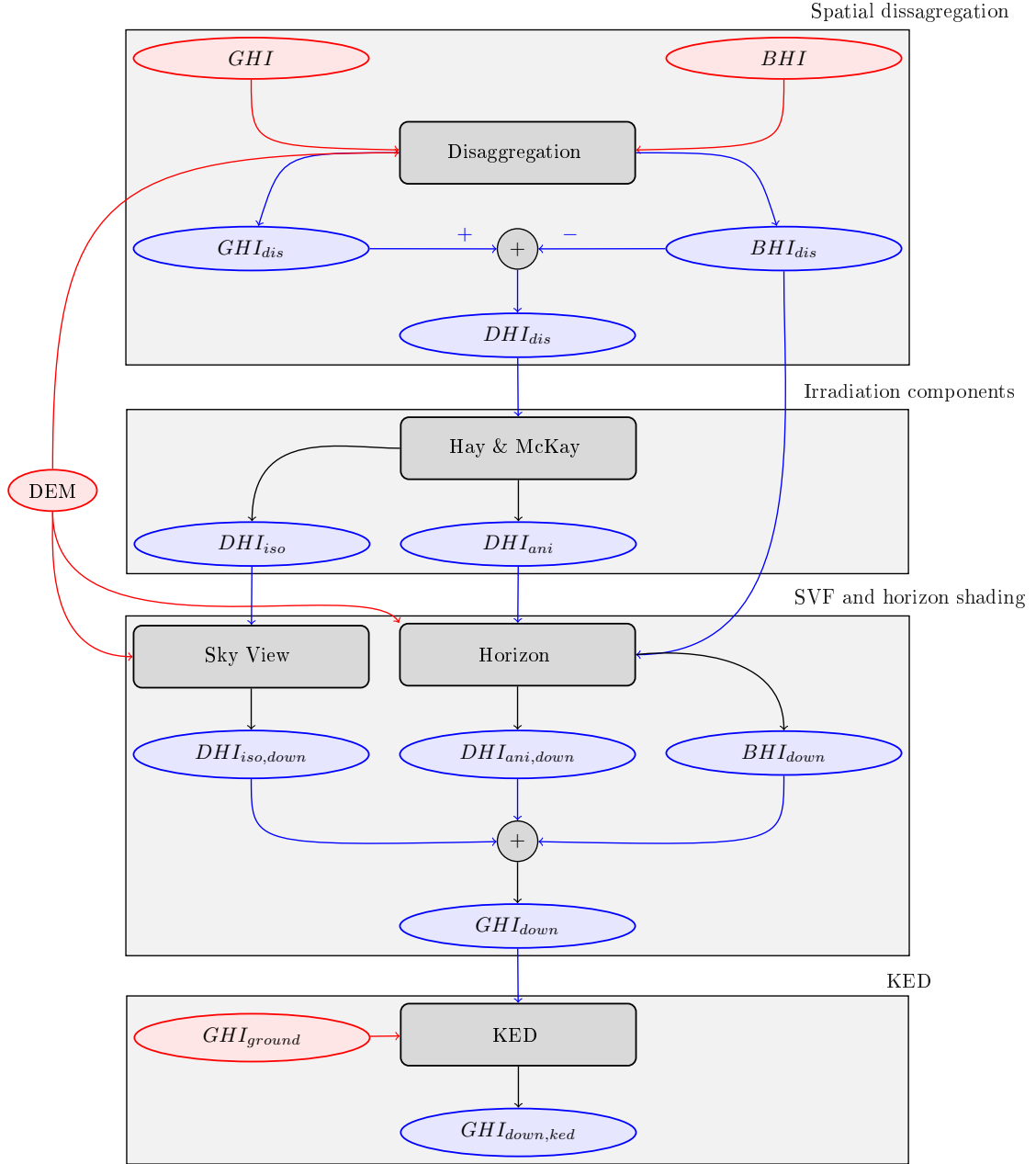


Figure 2: Methodology of downscaling: this figure uses red ellipses and lines for data sources, blue ellipses and lines for derived rasters (results), and black rectangles and lines for operations.



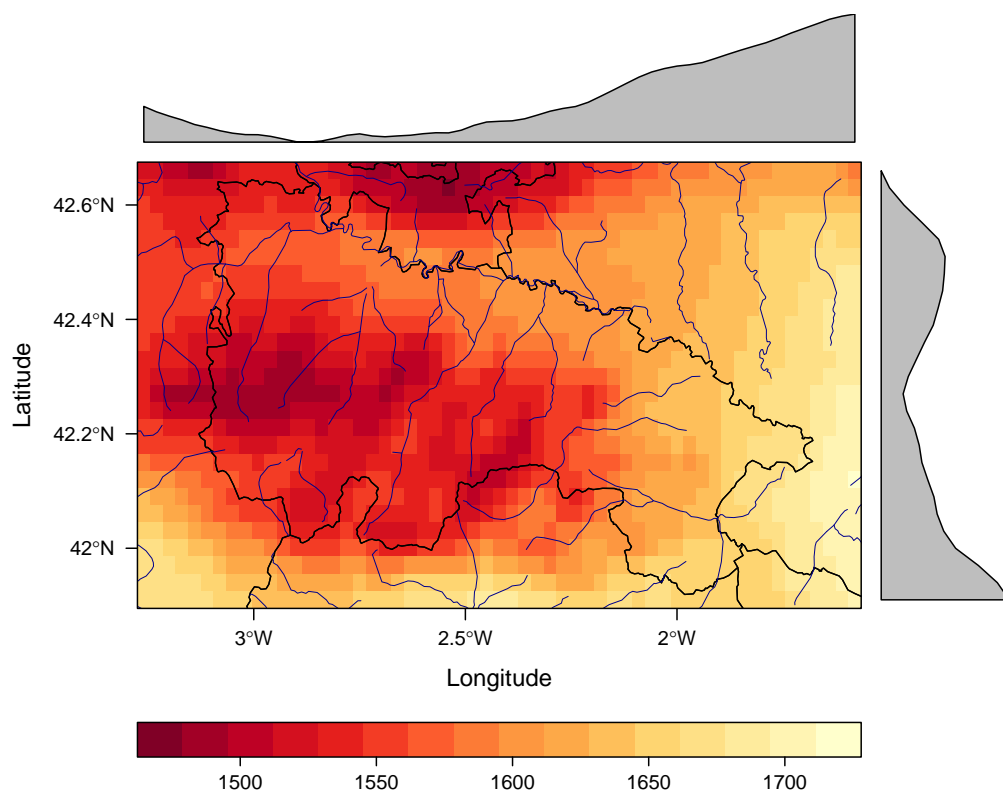


Figure 3: Annual GHI of 2005 from CM SAF estimates (0.03x0.03°) in La Rioja

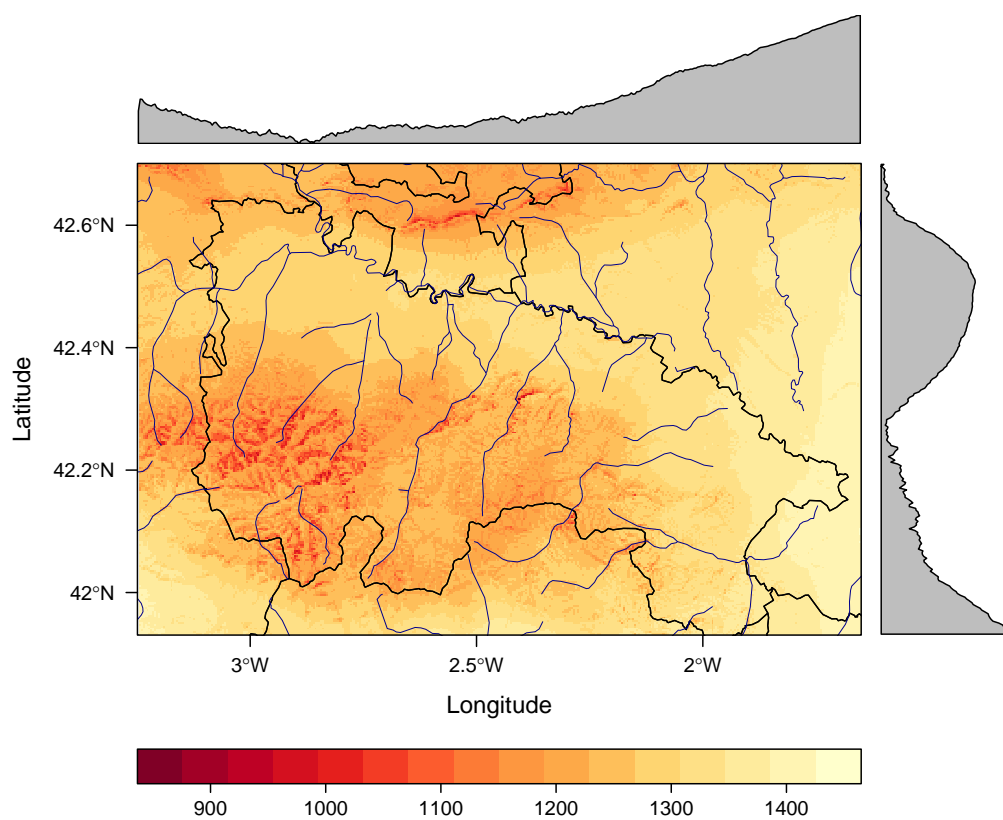


Figure 4: Annual GHI of 2005 downscaled without KED ( $0.03 \times 0.03^\circ$ ) in La Rioja

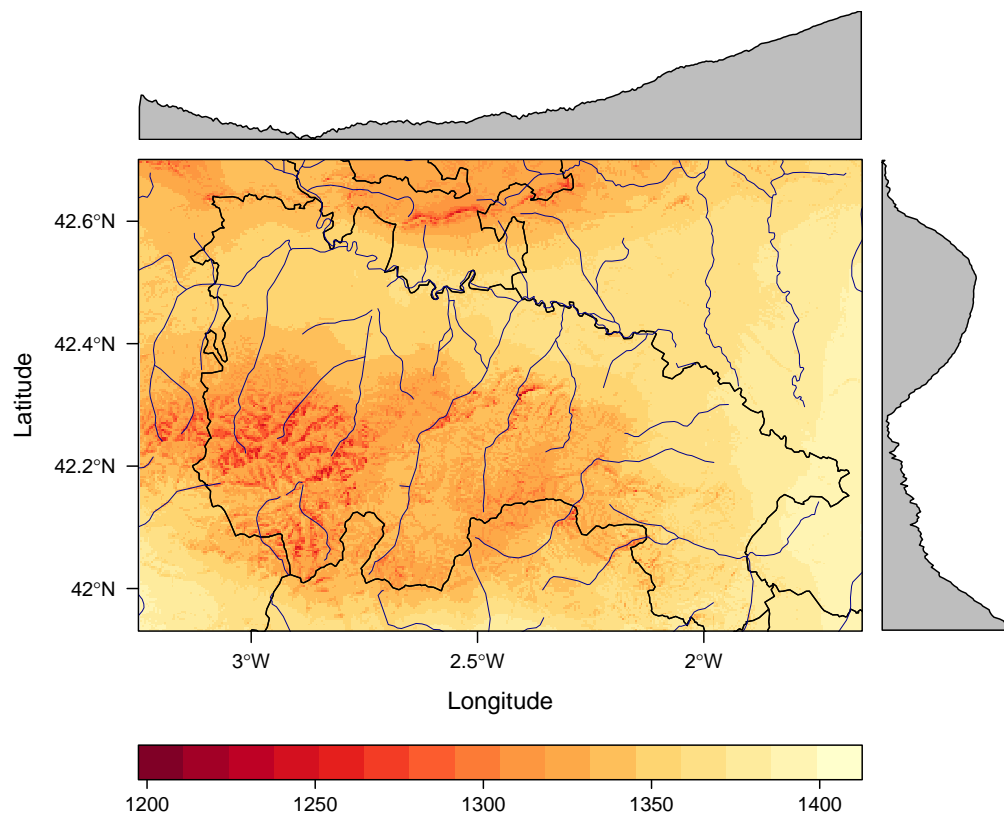


Figure 5: Annual GHI of 2005 downscaled with KED ( $0.03 \times 0.03^\circ$ ) in La Rioja

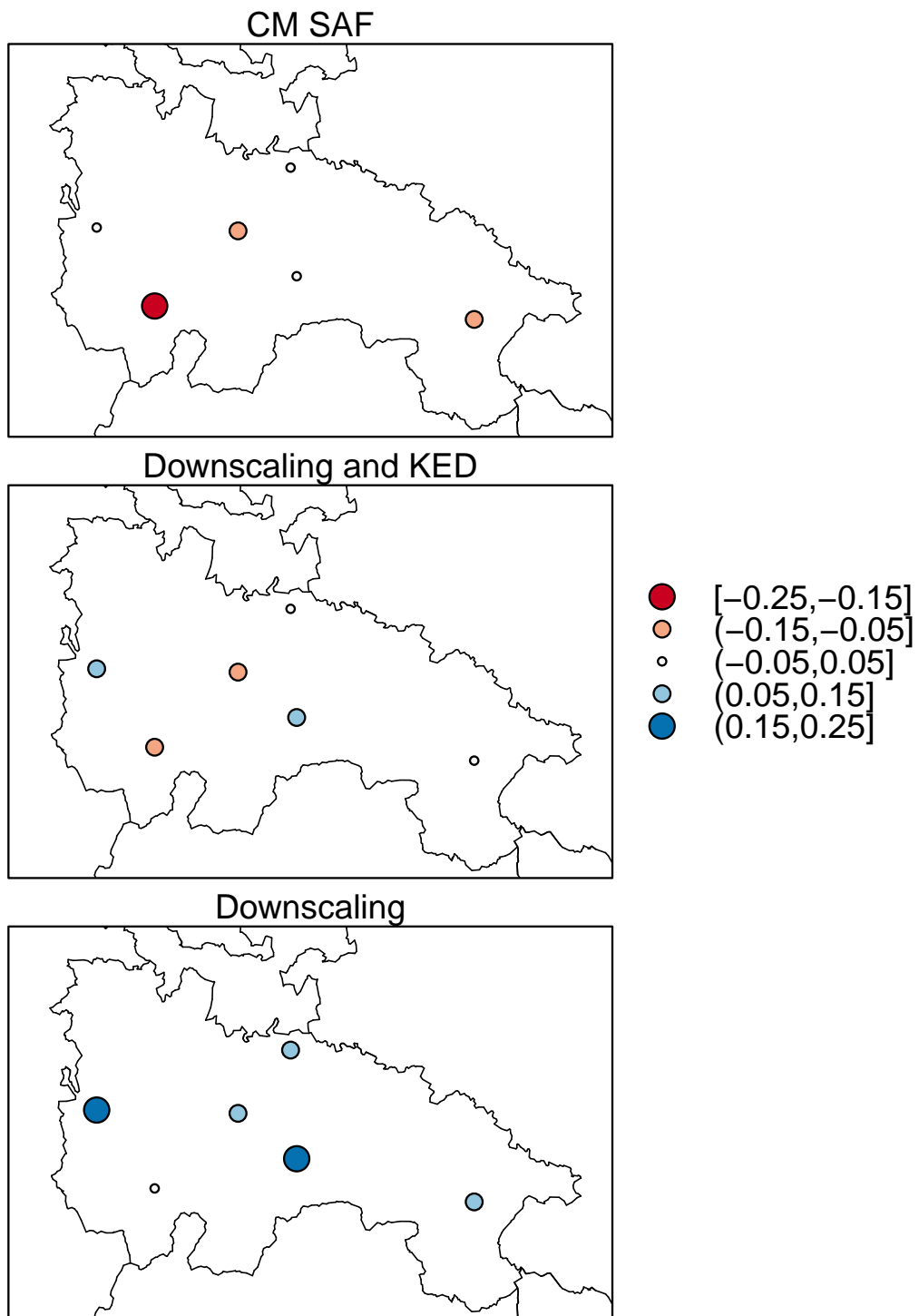


Figure 6: Annual relative differences evaluated with station measurements.

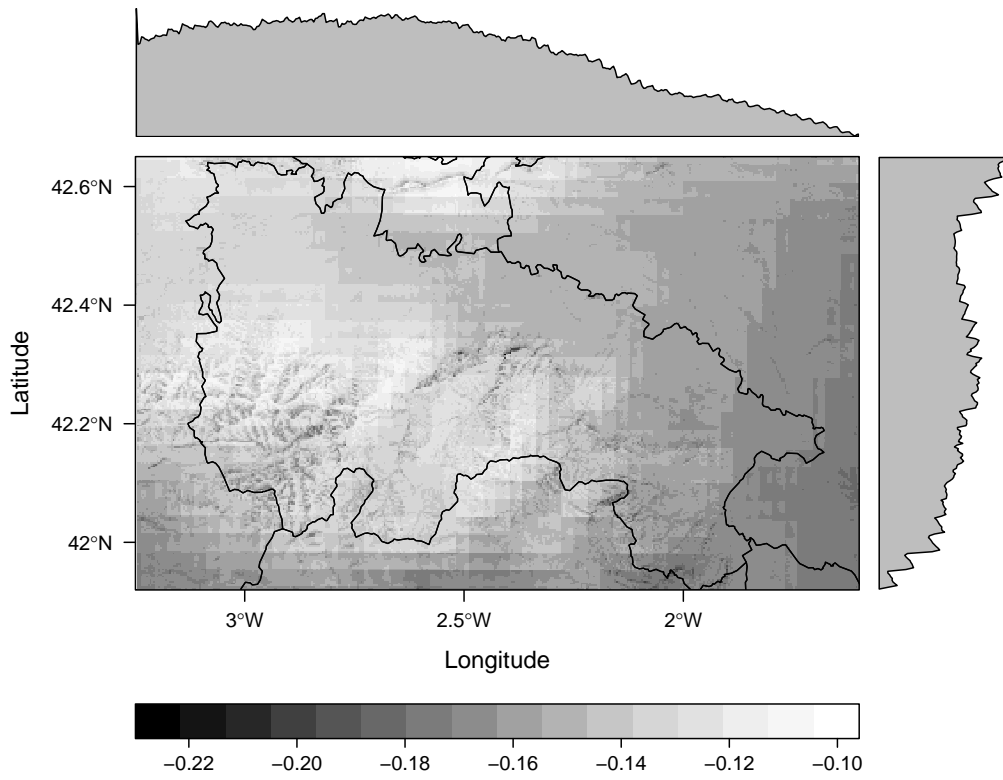


Figure 7: Relative difference of  $GHI_{KED}$  and  $GHI_{CMSAF,down}$  related to  $GHI_{CMSAF,down}$

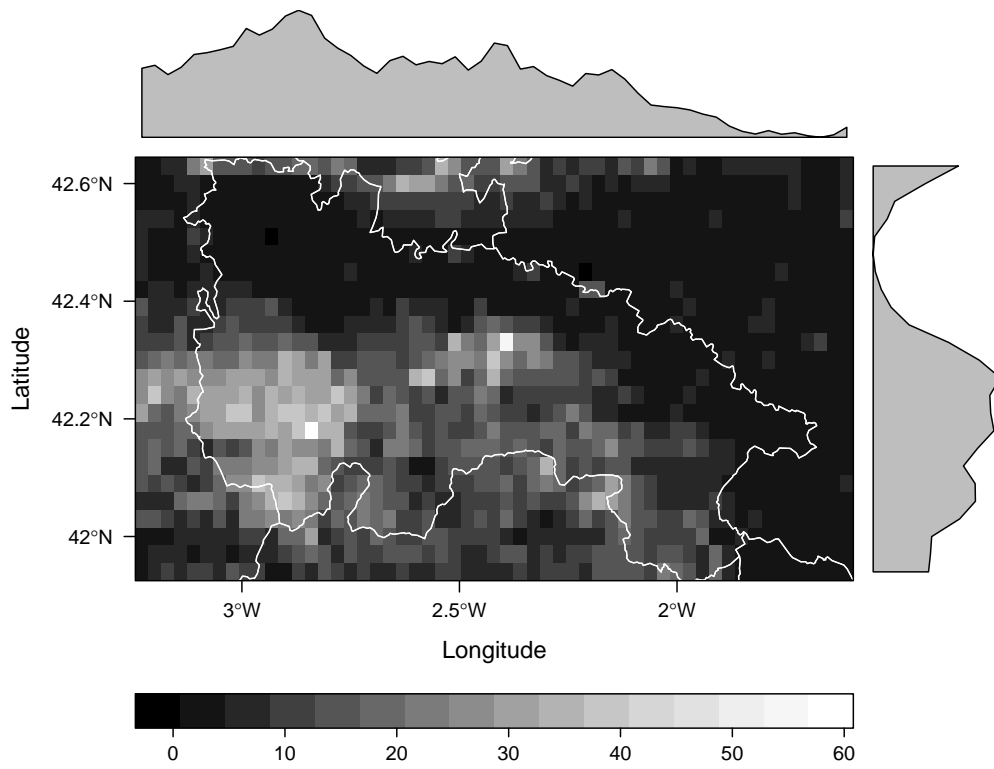


Figure 8: Difference of zonal standard deviations ( $kWh/m^2$ ) between downscaling without KED and with KED.

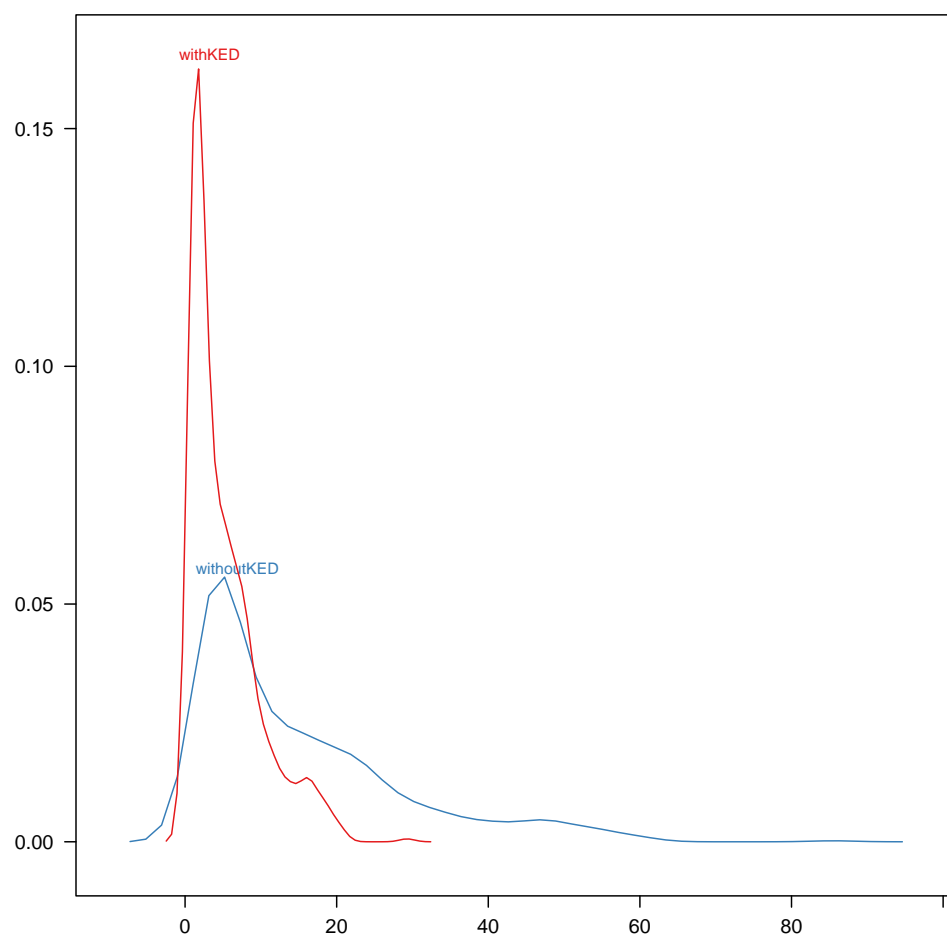


Figure 9: Density plot of zonal standard deviations between CM SAF and downscaling.