CISM 2.1.1 Documentation

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Part I User Documentation

Chapter 1

Introduction and Overview

1.1 Introduction

CISM, the Community Ice Sheet Model, originates from the Glimmer and Glimmer–CISM projects (Rutt et al., 2009)¹. The current name reflects the project's evolution from a standalone ice sheet model to a fully supported, coupled component of the Community Earth System Model, or CESM. CISM is a numerical model—a collection of software libraries, utilities and drivers— used to simulate ice sheet evolution. CISM is modular in design and coded almost entirely in standards-compliant Fortran 95. It currently supports two different dynamical cores ("dycores"), which solve the equations for conservation of mass, energy, and momentum. As with previous versions of Glimmer and Glimmer-CISM, the current version of CISM supports a serial, shallow-ice representation of ice dynamics. New with CISM2 are support for "higher-order" ice dynamics, scalable parallelism, and software links for coupling to modern, robust, C++ based, third-party solver libraries.

1.2 Overview

CISM consists of several components:

- cism_driver: the high-level driver (i.e., the executable) that is used to run the ice sheet model. Unlike the drivers in previous versions of Glimmer and Glimmer-CISM, cism_driver is used to run the code in all model configurations (e.g., for idealized test cases with simplified climate forcing and for model runs based on realistic geometries and climate forcing data).
- Glide: the dynamical core based on shallow-ice dynamics. This component is responsible for solving the governing conservation equations and determining ice velocities, internal ice temperature, and ice geometry evolution (see Chapter 4). Apart from minor changes, this is the same shallow-ice dynamical core used by Glimmer and Glimmer-CISM.
- Glissade: the dynamical core based on a first-order-accurate approximation of the Stokes equations for ice flow. This dycore, like Glide, solves the governing conservation equations. Unlike Glide, Glissade is fully parallel in order to take advantage of modern, multiprocessor, high-performance architectures (see Chapter 6).
- Glint: the climate model interface. Glint allows the core ice sheet model to be coupled to a variety of global climate models, or indeed any source of time-varying climate data on a lat-lon grid.

 $^{^1}Glimmer$ was originally an acronym for GENIE Land Ice Model with Multiply Enabled Regions, reflecting the project's origin within the GENIE (Grid ENabled Integrated Earth-system) model.

- **Test cases:** idealized test cases for the Glide and Glissade dynamical cores and for the Glint climate interface. These are used to (1) confirm that the model is working as expected and (2) provide a range of simple model configurations from which new users can learn about model options and create their own configurations (see Chapter 8).
- Shared code: a number of modules shared by different parts of the code. Examples include modules for defining derived types, physical constants, and model parameters, and modules that handle parsing of the configuration file and data input/output (I/O), as discussed below.

Each component is configured using a configuration file (*.config) similar to Windows .ini files. At run-time, the model configuration is written to a log file.

1D, 2D, and 3D data are written and read to and from netCDF files using the CF (Climate-Forecast) metadata convention². NetCDF is a scientific data format for storing multidimensional data in a platform- and language-independent binary format. The CF conventions specify the metadata used to describe the file contents. Many programs (e.g., Python, Matlab, OpenDX, Ferret, and IDL) can process and visualize netCDF data.

 $^{^2}$ http://cfconventions.org/

Chapter 2

Installing CISM

2.1 Getting and Installing CISM

CISM¹ is a relatively complex system of libraries and programs which build on other libraries. This section documents how to download and install CISM and its prerequisites. Many common problems and questions can be addressed using the user discussion boards². A CISM users mailing list is also available and can be subscribed to by sending an email to cism-users@googlegroups.com³. Please report unresolved problems using the bug reporting facility at the CISM Github website⁴.

CISM is distributed as source code, and a reasonably complete build environment is therefore required to compile the model. For UNIX and LINUX based systems (including Mac), the CMake build system is used to build the model. Sample build scripts for a number of standard architectures are included, as are working build scripts for a number of large-scale, high-performance-computing architectures (e.g., Cheyenne (CISL), Titan (OLCF), and Edison (NERSC)).

There are two ways to get the source code:

- 1. Download⁵ a released version of the code as an archive (.zip or .tar.gz file).
- 2. Clone the code from the CISM Github repository⁶ using the following command: git clone https://github.com/CISM/cism.git (It is also possible to clone the repository using the SSH protocol if you have an SSH keypair generated on your computer and attached to your GitHub account. See the CISM Github repository webpage and Github's help pages for more information.)

For beginners, downloading a zip archive of the latest release tag is recommended. More experienced users may want to download directly from the repository, as it will make updating the code easier in the future.

In either case, a Fortran90 compiler is required. Other software dependencies include the netCDF library (used for data I/O) and a Python distribution (used to analyze dependencies and to automatically generate parts of the code) with a number of specific Python modules. Users who want to run the code in parallel will need to install MPI, and users who want access to the *Trilinos* solver library will need to download and build *Trilinos*, and link to it when

¹https://cism.github.io/

²http://forum.cgd.ucar.edu/forums/ice-sheet-modeling-cism

³In order to subscribe from a non-Google email address, you should first make sure to completely log out from any Google sites (e.g., Gmail) before sending your request. If you do not, it will automatically try to associate your request with your Gmail account instead.

⁴https://github.com/CISM/cism/issues

⁵https://github.com/CISM/cism/releases

⁶https://github.com/CISM/cism

building CISM. Finally, you will need CMake and Gnu Make to compile the code and link to the various third-party libraries.

If you have not done so already, clone a tagged version of CISM or download an archive of the code, as noted above. Store the code or the unzipped/untarred archive in a location of your choice. More detailed build instructions, including instructions for the installation of supporting software, are given below.

2.2 Installing Supporting Software for Basic (Serial) CISM

Because the build process can be fairly complicated, we describe it in detail below, relying on the use of a package manager to handle many of the standard software dependencies. For each step we give specific instructions for both Mac OS X using MacPorts (in red boxes) and Linux (in blue boxes). For the former the instructions work for Mac OS High Sierra. For the latter, we have verified that the instructions work for Ubuntu 12.10 and 14.04 LTS. For different but related systems, hopefully these instructions can be used as a guide.

CISM can be installed in either a serial or parallel configuration. The parallel mode allows the model to be run on multiple processors which can greatly speed up execution. This is a common configuration to use on supercomputing clusters, but can also be convenient on modern desktops and laptops which often have four or more cores available. However, the parallel build requires additional supporting software, so we first detail how to build serial CISM. For new users, it is recommended to first build and successfully run serial CISM before moving on to the parallel build.

Note: Glide, the shallow-ice dycore, can only run on a single processor, even when the code is built with full parallel support. This is also true of the SLAP solver routines.

The instructions below assume the user has administrative privileges for installing new software (note the extensive use of sudo). If you are working on a shared machine without administrative privileges, you might proceed by assuming all needed packages are present and continue to the CISM installation section. If you encounter problems, you can refer back to this section to determine which packages might be missing or problematic before contacting your system administrator.

Mac OS X

As mentioned above, we will take advantage of MacPorts, a software package manager for Macs. This will allow us to install most of the base level software libraries needed by CISM with few complications.

Go to http://www.macports.org/install.php, where you will find a range of ".pkg" installs available, including those for Mountain Lion, Lion, and Mavericks versions of Mac OS X.

Installing MacPorts requires installing the Xcode developer toolset provided by Apple. Details of how to obtain Xcode vary by version of OS X. See MacPorts installation instructions and this link for details. Once Xcode is installed, you may need to additionally download the "command line tool" from the Preferences / Downloads menu of Xcode.

Depending on computer security settings at your institution (firewalls, etc.), you may need to add proxy information so that Macports can communicate and download software from the outside world. All Macports software will be installed under /opt/local/by default. To add proxy information, after installing Macports, edit the configuration file at /opt/local/etc/macports/macports.conf. (Note this is probably a read-only file that requires superuser permission to edit, so you will need to edit the file with something like: sudo vim /opt/local/etc/macports/macports.conf). By searching for the text string "proxy", you will find the lines like proxy_http hostname:12345

near the bottom of the file. Enter your proxy information here as appropriate (e.g., hostname:your_host_info_here).

If you have previously installed Macports but not updated it recently, it's generally a good idea to do so. Ideally, this should be done with admin or root privileges (you will be prompted to enter your password) using:

sudo port selfupdate

You will then be prompted to update any installed ports that are outdated, which you can do using:

sudo port upgrade outdated

To search for available software in Macports, type:

port search software-name

Software is installed through Macports using the command:

sudo port install software-name

Additional Macports tips will follow inline below. Extensive documentation for Macports can be found at the Macports website.

Ubuntu 12.10

This Ubuntu instructions describe setting up supporting software and CISM in a Linux environment. These instructions were written using a fresh installation of Ubuntu 12.10 but steps should be very similar in other versions of Ubuntu or other distributions of Linux. Instructions make use of the command line tool for installing packages that comes with Ubuntu, apt-get. Other package management tools (e.g., Software Center) could also be used.

It's generally a good idea to synchronize your local package index files before installing new software using <code>apt-get</code>:

sudo apt-get update

To search for available packages, type:

apt-cache search software-name

And to see detailed information about a package, type:

apt-cache show software-name

Packages are installed through apt-get using the command:

sudo apt-get install software-name

Some users have reported that BLAS and LAPACK libraries need to installed explicitly, for example when starting from a "clean" machine. To do this, use the following two commands:

sudo apt-get install libblas-dev

sudo apt-get install liblapack-dev

Additional apt-get tips will follow inline below. Extensive documentation for apt-get can be found at the <u>Ubuntu</u> website and through man pages (man apt-get).

2.2.1 Install git version control software

If you intend to download the CISM code as a git repository, you will need the git package installed. If you prefer to download a zipped archive of the code, this step can be skipped.

Mac OS X

Install git with:

sudo port install git

Ubuntu 12.10

```
Install git with:
    sudo apt-get install git
```

2.2.2 Install the GCC compiler suite

The GCC compiler suite contains compilers for C, C++, and, optionally, Fortran. Fortran and C compilers are required for serial CISM, and a C++ compiler is also needed for parallel CISM. CISM is known to work with GNU gfortran compilers, Intel ifort, and PGI. In these instructions we will use GNU compilers because they have been extensively tested with CISM and are freely available. Advanced users are welcome to use other compilers of their choosing.

CISM has been tested extensively with gfortran version 6. Newer (or older) versions may also work.

Mac OS X

Searching for gcc with port search gcc will return:

```
apple-gcc40 @5495 (lang)
  Apple's version of gcc 4.0
```

in addition to a lot of other information on available Macports installs related to the GCC (Gnu) compiler suite.

Where possible, we want to make sure that all other software we build and install with Macports uses the version of GCC we choose to install. To date, we've had success with GCC 6.4.0 (others may work as well but have not been tested). To install GCC 6.4.0 type:

```
sudo port install gcc6
```

You will see some verbose output telling you what is happening (downloading packages, expanding them, building, installing, checking, etc.). When the install is complete, you can type:

```
port installed
```

to see what packages you currently have installed. You should see something like gcc6 6.4.0_0 (active). (The minor version numbers after the "4.6" may differ as MacPorts makes updates to the port.) You will likely also see other packages that have been installed (software dependencies for GCC that were automatically installed by MacPorts and/or other ports you have manually installed).

The "(active)" description identifies which version of a particular package Macports currently thinks you want to use (e.g., you could also have another older GCC suite installed). To make sure the newly installed version is active, you would type:

```
port select gcc
which will return something like:
Available versions for gcc:
   mp-gcc6 (active)
   none
```

This confirms that GCC 6 is active (the mp indicates a Macports version). It is possible that gcc6 will be listed as active when you type port installed, but that mp-gcc6 will not be listed as active when you type port select gcc. If mp-gcc6 is not active as shown above, then you will need to select it using:

```
sudo port select --set gcc mp-gcc6
```

This will ensure that any generic call to gcc, gfortran, g++, will point to the libraries just installed.

Ubuntu 12.10

GNU compilers may have come with your Linux distribution. If not, they need to installed. Ubuntu 12.10 comes with gcc installed but not gfortran.

Install gfortran with:

sudo apt-get install gfortran

2.2.3 Install build tools

Additional tools are needed for managing the build process. make (specifically, GNU's gmake) usually comes with Mac and Linux distributions, but if not it should be installed. Additionally, CISM uses the CMake build utility (a cross-platform, open-source build system).

Mac OS X

While you probably already have a version of make on your system, it may be out of date or conflict with other Macports installed software. The required versions for CISM can be installed through Macports with this command:

```
sudo port install gmake cmake
```

In addition to the software installed above, you should now see something like the following when you type port installed:

```
gmake @4.2.1_0 (active) cmake @3.10.2_0 (active)
```

Ubuntu 12.10

On Ubuntu (and other Debian systems) there is usually a package called build-essential that includes a large collection of tools and libraries that are typically necessary for compiling code. Install these tools and CMake with:

sudo apt-get install build-essential cmake cmake-curses-gui

2.2.4 Install netCDF

NetCDF stands for "network Common Data Form" libraries, which are a machine-independent format for representing scientific data. This is required by CISM for performing input/output. The netCDF package you install must include Fortran libraries for CISM to compile (in some package managers, the Fortran libraries are in a separate package). There are substantial

differences between versions 3.x and 4.x of netCDF, but both version series should work with CISM. It is also possible to download and compile netCDF libraries manually, which may be preferred by advanced users wanting to use a specific version.

It is also recommended that you install optional tools for working with netCDF datafiles. ncview is a convenient tools for viewing netCDF files. (Some alternatives are to write Python or Matlab scripts or to use another tool like Paraview or Ferret.) NCO ("netCDF Operators") is a toolkit of command line tools for manipulating and analyzing data stored in netCDF-accessible formats.

Mac OS X

To install NetCDF, use sudo port install netcdf-fortran +gcc6. Note that there are other versions of NetCDF available to install. It is important to choose the one with the "Fortran" extension. The "gcc6" syntax specifies a port "variant". This tell Macports that, if there is a version of the selected software to install that is consistent with the GCC 6 compiler suite, then it should choose that one. Typing port installed should now include:

```
netcdf @4.4.1.1_1+dap+gcc6+netcdf4 (active)
netcdf-fortran @4.4.4_1+gcc6 (active)
```

The "dap+gcc6+netcdf4" comes along automatically.

Optional but recommended: Tools for working with netCDF data files. sudo port install neview neo

If you encounter an 'unable to open display' error when running ncview, you may need to install a newer version of the X Window System than the one provided by Apple. We have had success using the latest version of XQuartz: http://xquartz.macosforge.org

Ubuntu 12.10

Install netCDF libraries with:

sudo apt-get install libnetcdf-dev

Optional but recommended: Tools for working with netCDF data files. sudo apt-get install netcdf-bin ncview nco

2.2.5 Install Python and related modules

Python is used by CISM to autogenerate I/O code during compilation, and is also used by most test case scripts to set up initial conditions and analyze and plot results. Only Python 2.7 has been tested. Python 3 may work for some uses but is likely to generate errors due to extensive changes between versions 2 and 3. Also, CISM uses a number of python modules:

- numpy required for generating many test case initial conditions
- matplotlib used by some plotting scripts. Not strictly necessary but required for those scripts to work properly.
- a python netCDF I/O module. Options are netCDF4, Scientific.IO.NetCDF, or PyCDF. netCDF4 is the ideal choice, but it is often not available through Linux package managers and must be installed through a python package manager like pip, or manually. PyCDF is the least recommended option here because it is not entirely compatible with the others. Scientific.IO.NetCDF is usually available through Linux package managers.

Mac OS X

While Mac OS X already comes with a working Python distribution, we will need additional modules that can sometimes be tricky to get working together correctly. We have successfully used both the Enthought Python distribution (which is free for people associated with a university) and a version installed using Macports. To obtain and install Enthought, click on the link above and follow their directions. To install version 2.7 using Macports, along with the necessary additional modules, do the following:

sudo port install python 27 py27-numpy py27-matplotlib py27-scientific py27-netcdf4

The existence of two versions of python on your system can lead to confusion. It is important that you leave the version of python that came supplied by Apple so that your system has access to it. However, you will want to be sure that CISM has access to the new, more modern version of python you have installed. In our experience, this can be one of the most problematic parts of the installation process. You can use port select:

sudo port select python python27

You can check that Macports python is used by default by typing:

which python

and you should see: <code>/opt/local/bin/python</code>. If you instead see <code>/usr/bin/python</code> then the default Apple python is still the version that is being used on the command line. If this happens, or if you encounter errors with this setup, an alternative approach is to modify the PATH variable in your <code>.bashrc</code> or similar environment settings script to make sure that <code>/opt/local/bin</code> is before <code>/usr/bin</code> in your path.

Ubuntu 12.10

Python generally comes with most Linux distributions. If it is not present, it must be installed. Often, there is an additional python development package that is necessary when working with compiled code (tpyically called python-dev on Ubuntu).

Install python modules with:

 $\verb|sudo| apt-get| in \verb|stall| python-dev| python-numpy| python-matplotlib| \\ python-scientific|$

Optional: Installing netCDF4 python module.

Install pip (a tool for installing and managing Python packages):

sudo apt-get install pipa

Next, install ${\tt HDF5}$ using pip:

sudo apt-get install libhdf5-dev hdf5-tools hdf5-helpers flex

Finally, install netCDF4 using pip:

sudo -E pip install netcdf4

^aFor Ubuntu 14.04 there is a known issue with pip and handling freetypes. This can be fixed using sudo apt-get install libfreetype6-dev libxft-dev.

2.3 Building Serial CISM

At this point we are ready to build a *serial* version of CISM and its linked libraries. While we ultimately want to build a version of the code that also runs in parallel, it is often useful to stop at this step to make sure everything is working. Then, if problems occur during the parallel build process (as they sometimes do), we know those problems have occurred only during the last step of the process.

If you have not already done so, obtain the source code following the instructions above

in section 2.1. Below, all paths starting with ./ indicate the root level of the source code directory. E.g., if you expanded your tar.gz archive into a main source code directory with the path /usr/JohnDoe/CISM, the ./ refers to the path /usr/JohnDoe/CISM/.

Unlike previous versions of the code, the build system is now entirely based on CMake (Autotools is no longer used).

Build scripts are provided that should work for most standard Mac and Linux setups, as well as some supercomputing platforms on which CISM is commonly run. All build scripts are located in the ./builds directory from the root level of the code. In general, change to the subdirectory that most closely matches your system and intended build.

If you encounter an error when using the included scripts, you may need to modify some details, such as the location of your NetCDF libraries or your compiler names. Other errors you might encounter may indicate that some of the supporting software (above) is missing.

Mac OS X

On a Mac, you should be able to build the code by doing the following:

- 1. Change to the ./builds/mac-gnu directory from the root level of the code.
- 2. Configure the build using source mac-gnu-serial-cmake
- 3. Build the code using make -j X, where the "X" refers to the number of processors available for use in the build (or just make if you have only one processor).

Ubuntu 12.10

On a Linux platform, you should be able to build the code by doing the following:

- 1. Change to the ./builds/linux-gnu directory from the root level of the code.
- 2. Configure the build using source linux-gnu-serial-cmake
- 3. Build the code using make -j X, where the "X" refers to the number of processors available for use in the build (or just make if you have only one processor).

When the build completes, you can check for the executable driver by typing ls cism_driver from within the current build directory (here, from within the ./builds/mac-gnu or ./builds/linux-gnu/directory). The file cism_driver is the executable you will link to when running the model, which is generally done using a symbolic link. For example, from the ./tests/shelf/directory, one would link to this executable using,

ln -s ../../builds/mac-gnu/cism_driver/cism_driver ./

Chapter 8 discusses running the executable for standard test cases.

Advanced users may want more control over the build scripts. There are a number of build options used by CMake to customize the build. You can manually modify the build scripts included with the code, or use the tool ccmake to interactively adjust build options (type ccmake ../../ from any build directory after having run the configure script once). The available options are listed in Table 2.1. Many of these options pertain to the parallel build which is discussed in more detail below.

Also, there are standard CMake options that can be set (e.g., CMAKE_C_COMPILER, CMAKE_Fortran_COMPILER, etc.). Many of these are explained in the CMake documentation.

CISM_BUILD_CISM_DRIVER	Toggle to build cism_driver, on by default
CISM_BUILD_EXTRA_EXECUTABLES	Toggle to build other executables, off by default
CISM_COUPLED	Toggle to build CISM for use with CESM, off by
	default
CISM_ENABLE_BISICLES	Toggle to build a BISICLES-capable cism_driver, off
	by default
CISM_FORCE_FORTRAN_LINKER	Toggle to force using a Fortran linker for building
	executables, off by default
CISM_GNU	Toggle to set compilation flags needed for the gnu
	compiler, off by default
CISM_INCLUDE_IMPLICIT_LINK_LIB	Toggle to explicitly include the
	CMAKE_Fortran_IMPLICIT_LINK_LIBRARIES on
	the link line, on by default
CISM_MPI_MODE	Toggle to configure with MPI, on by default
CISM_NETCDF_LIBS	netCDF library name(s)
CISM_NO_EXECUTABLE	Set to ON to just build libraries, off by default
CISM_SERIAL_MODE	Toggle to configure in serial mode: off by default
CISM_SOURCEMOD_DIR	Path to SourceMod directory of F90 files to replace
	CISM files
CISM_STATIC_LINKING	Toggle to set static linking for executables, off by
	default
CISM_USE_DEFAULT_IO	Toggle to use default i/o files rather than running
	python script, off by default
CISM_USE_GPTL_INSTRUMENTATION	Toggle to use GPTL instrumentation, on by default
CISM_USE_MPI_WITH_SLAP	Toggle to use mpi when using SLAP solver, only rel-
	evant if CISM_SERIAL_MODE=ON, off by default
CISM_USE_TRILINOS	Toggle to use Trilinos external solver libraries, on by
	default
CMAKE_VERBOSE_CONFIGURE	Verbose CMake configuration, on by default

Table 2.1: Available ${\tt CMake}$ settings for configuring the CISM build process

2.4 Installing Supporting Software for Parallel CISM

To build parallel CISM, MPI compilers and libraries are required. Only the higher-order dycore (Glissade) can run in parallel. (There is also a higher-order dycore called Glam that can be run in parallel, but it is used for development and testing and is not supported for scientific applications.)

In addition, you may choose to include the Trilinos package of external solver libraries. These are not required, but for some problems Trilinos may provide better performance and stability than the native solvers. Trilinos can also technically be used with a serial build, but this configuration is not supported or recommended.

2.4.1 Install MPI

MPI (Message Passing Interface) libraries and compilers are necessary for compiling parallel CISM. These libraries are used for handling parallel communications when running the code on multiple processors. A more complete description of parallel model configurations is given in Chapter 7. (For example, some test cases and configurations when running the shallowice dycore are not fully supported in parallel). OpenMPI and MPICH are two common MPI implementations.

Mac OS X

It is likely that you already have versions of MPI installed on your system, but they may be out of date or not compatible with the other libraries we have and will be installing. Using Macports, the MPICH version of MPI is known to work when building CISM. (OpenMPI may also work, but we've seen more consistent success on Macs with MPICH.)

First, check Macports for available versions of MPICH using port search mpich*. We want the version that is compatible with our GCC compiler suite, so we type:

```
sudo port install mpich-devel-gcc6 +fortran
```

To make sure this is active, type

port installed mpi*

which should return

mpich-devel-gcc6 @3.3a2_3+fortran (active)

As when installing the GCC compilers, we want to make sure any generic call to MPI points to MPICH. This can be done with the following command:

sudo port select --set mpi mpich-devel-gcc6-fortran

Ubuntu 12.10

Either OpenMPI or MPICH are likely to work with CISM on Linux machines. On Linux machines, we have tested OpenMPI more thoroughly. Install OpenMPI with: sudo apt-get install openmpi-bin

2.4.2 Install Trilinos solver libraries

Trilinos is a modern, open source, C++ based library of parallel nonlinear and linear solvers, preconditioning and mesh-partitioning tools, and much more. It can be downloaded here⁷. (The software is free, but you are required to enter your email address to download it.) The documentation below assumes that you are working with version 11.10.* and was specifically tested using version 11.10.2.

Building Trilinos requires CMake version 2.8 or later, which ideally you have already installed as discussed above. Trilinos is not needed to run the default parallel, higher-order dycore (Glissade), but it may be useful for more difficult problems or for debugging in cases where the native Fortran solvers fail to converge.

The build instructions for Trilinos on Mac and Linux are very similar, so users of both systems can follow the primary instructions below, except where noted.

Trilinos requires both (1) an "out-of-source build" and (2) an "out-of-build installation". This means that you cannot build the code in the same directory where the source code lives, and you cannot install the libraries in the same directory where you build the code. (Older versions of Trilinos required an out-of-source build but not an out-of-build installation.) The easiest way to satisfy this requirement is to have separate "source", "build" and "install" directories in the location where you want to install the code. For example, in /usr/local/, you could set up the following three directories:

```
trilinos-11.10.2-Source/
trilinos-11.10.2-Build/
trilinos-11.10.2-Install/
```

The "source" directory will be created on its own when you uncompress the tar.gz archive that you download. You do not have to keep the source code where you build and install the Trilinos libraries, but you will need to remember the path to where that source code lives on your computer.

To configure the Trilinos build, you will need to execute a CMake configure script. Sample configure scripts for a number of standard platforms are included in the "sampleScripts" directory under the root level of the Trilinos source code. Also, the CISM code includes examples of Trilinos configure scripts ("do-configure") for use with CISM for both Linux and Mac platforms in the ./utils/trilinos_config_scripts_examples directory. We recommend starting with one of those scripts and modifying it as necessary to work on your system⁸.

The paths to both the "source" and "install" directories are specified within the "doconfigure" scripts. In these instructions, those directories are both assumed to live within /usr/local/, but other locations are fine to use too (e.g., in your home/User directory).

Mac OS X

Also note the explicit path in the MPI lines, e.g.,

```
-D MPI_EXEC="/opt/local/mpiexec" \
```

Since some Macs may come with their own pre-installed OpenMPI libraries, it is important here to specify the path to the version we previously installed using Macports.

Find the example script most appropriate for your system, copy it to the trilinos-11.10.2-Build directory, and modify it if necessary (e.g., adjust paths, compiler locations, etc.). Execute it

⁷http://trilinos.org/download/

⁸If you are following the above installation instructions for Mac exactly, then the configure script ./utils/trilinos_config_scripts_examples/do-configure-Trilinos-11.10.2-for-Mac-10.9.4 should work with few modifications.

with:

```
source ./do-cmake9
```

from within your trilinos-11.10.2-Build directory. Depending on where you are building and installing the code, you may need to have administrative privileges (in which case you would type sudo source ./do-cmake). If the configure step was successful, you should see the following displayed on your screen:

```
Processing enabled package: [PACKAGE NAME]
...

Exporting library dependencies ...

Finished configuring Trilinos!

-- Configuring done
-- Generating done
```

-- Build files have been written to: /usr/local/trilinos-11.10.2-Build

It is a good idea to scan the output while the "do-cmake" script is executing, for example to ensure the configure process is picking up the compilers you specified (e.g., it is using the Macports versions as opposed to some Mac default versions that might also be on your system). Once the code is configured successfully, build the libraries from within the trilinos-11.10.2-Build directory by typing:

```
make (or sudo make if necessary)
```

For multiprocessor machines, the build process can be sped up significantly using the "-j" command as described above for building serial CISM:

```
make -j X
```

where "X" is the number of cores available on your machine (e.g., make -j 4 for a 2-processor, dual-core machine).

Building Trilinos can take a long time (e.g., an hour or more), depending on your machine, the number of processors used for the build, and the number and type of libraries you are installing. Once you have built the code, we highly recommend testing it using:

```
make test
```

(The Trilinos_ENABLE_TESTS:BOOL variable in the do-cmake script can be set to "OFF" to disable building of the tests.) Screen output will tell you if and how many tests failed. We have seen a few tests fail while still having a perfectly good and working Trilinos library. In general, if the number of tests passed is above 90%, the library will likely work fine with CISM. Query the CISM users or developers lists 10 if you have questions about specific Trilinos tests failing.

Mac OS X

On a Mac, MPI tests have been known to trigger a dialog box from the firewall. With more than 300 tests, these messages popping up continually can make it impossible to use your computer until the tests complete. To keep them from appearing, you can temporarily turn off your firewall under "System Preferences" (Security > Firewall > Stop). Be sure to turn the firewall back on when the tests are complete!

After running the tests, you will need to install Trilinos using:

⁹Here we have assumed that the name of the configure script is "do-cmake". The script name may differ depending on what you have called it or if you copied and modified one of the scripts from ./utils/trilinos_config_scripts_examples.

 $^{^{10} {\}rm cism-users+subscribe@googlegroups.com} \ - {\rm or-} \ {\rm cism-devel+subscribe@googlegroups.com}$

make install

This will build the actual Trilinos libraries in the path specified in the

-D CMAKE_INSTALL_PREFIX:PATH=/path

line of your "do-cmake" script (above). For this example, those libraries will be installed in: /usr/local/trilinos-11.10.2-Install

After successfully building Trilinos, create an environment variable called CISM_TRILINOS_DIR so the CISM build process can find the Trilinos installation. For example, if you are using the bash shell and your current directory is the Trilinos install directory, you can do:

```
export CISM_TRILINOS_DIR=$PWD
```

You may prefer to modify your .bashrc or .bash_profile (or similar) to set this environment variable on every login.

Alternatively, you can modify the CISM parallel build script (below) so that the line:

-D CISM_TRILINOS_DIR=\$CISM_TRILINOS_DIR \

is set to the Trilinos installation directory.

2.5 Building Parallel CISM

The procedure for building parallel CISM is nearly identical to the serial build (above). The build script for parallel CISM for a Mac is located at builds/mac-gnu/mac-gnu-cmake, while the build script for Linux is located at builds/linux-gnu/linux-gnu-cmake. From the appropriate directory, run:

source mac-gnu-cmake or source linux-gnu-cmake

Once the configuration step completes successfully, you can compile the code as before with:

make

or

make -j 4

if you have 4 processors available (or as many processors as you would like to use). See Section 2.3 for details about customizing the build process.

Building a parallel version of CISM that includes Trilinos requires setting the

-D CISM_USE_TRILINOS

flag to ON in the builds/mac-gnu/mac-gnu-cmake script.

2.6 Next Steps

If you make any changes to the source code, you only need to re-run make from your build directory to generate an updated executable. One exception is that if you edit the lists of input/output netCDF variables in *_vars.def files (see Appendices A and B), you need to first re-source the configuration script (e.g., source mac-gnu-cmake) before re-running make.

Now that you have successfully built the code, you can proceed to Chapters 3 and 5 to learn more detailed information about ice sheet modeling, to Chapters 4 and 6 to learn more about the various model approximations available through CISM, or you can proceed to Chapter 8 to learn how to run and examine some standard model test cases.

Chapter 3

Introduction to Ice Sheet Modeling: Derivation of Field Equations

In this chapter we give an introduction to ice dynamics and the other conservation equations that must be accounted for when simulating glacier and ice sheet evolution.

Ice sheets are key components of Earth's climate system. They contain nearly all of the planet's fresh water; changes in their volume have an immediate effect on sea level; changes in their area and surface characteristics affect global albedo; and they play a role in the circulation of both the atmosphere and the ocean. Through the latter half of the Quaternary Period, ice sheets have modulated the planetary response to orbitally-driven insolation cycles. Looking forward, the Greenland and Antarctic ice sheets have the potential to play important roles in climate change.

A numerical model is a discrete approximation of a continuous process. The approximation is discrete, due to both the finite nature of a computer's precision and the finite problem size that can be tackled using a computer. The underlying process is continuous because it is commonly formulated in terms of ordinary or partial differential equations (ODEs and PDEs, respectively). Numerical models cannot be "solved" until the boundary and initial conditions are specified. Such models may be very simple, such as a harmonic oscillator, or very complex, as in a complete Earth System Model (ESM). In general, ESMs are composed of a number of component models, of which land ice (encompassing glaciers, ice caps, and large ice sheets) is but one component (with atmosphere, ocean, sea ice, and land surface models being the other primary components).

3.1 Conservation Equations

For the majority of the physical systems encountered in Earth science problems, the first step in modeling is a mathematical description of the conservation of energy, momentum, and mass. Only after that description is laid out do we turn to the question of approximating those equations in a form that can be solved on a computer.

3.1.1 Integral form

The mathematical formulation of conservation can be arrived at by considering the change in a quantity ϕ that is known within a **control** volume V. The control volume is enclosed by a surface S, with the outward positive unit vector \hat{n} , normal to S.

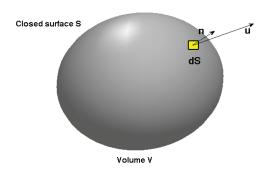


Figure 3.1: Diagram of control volume and associated quantities.

The value of ϕ within V may change over time if:

- 1. There is a flux of ϕ through S. The flux is partitioned into two parts, one due to diffusion and another due to advection.
- 2. ϕ is created or destroyed within V.

Formally, the time rate of change in ϕ within V is written:

$$\frac{\partial}{\partial t} \int_{V} \phi dV = -\int_{S} \mathbf{F} \cdot \hat{n} \ dS - \int_{S} \phi \mathbf{u} \cdot \hat{n} \ dS + \int_{V} R dV \tag{3.1}$$

where **F** represents the flux due to diffusion ($\mathbf{F} \propto \nabla \phi$), $\phi \mathbf{u}$ represents a velocity field *advecting* ϕ , and R represents a source (or sink) of ϕ . Vector quantities are represented in **boldface**. The negative signs in front of the first two terms on the right-hand side indicate that an outward flux results in a decrease in ϕ within the volume enclosed by S.

This statement of conservation of ϕ in the unit volume V is always true, independent of the size of V and even if the fields enclosed by S are not continuous (this is the case because we integrate over V). It is important to note that that any information on spatial scales smaller than V is lost in the process of integration.

3.1.2 Derivative form

Numerical models are often easier to formulate from the derivative form of the conservation equation, and this is more often the form in which conservation equations are written. This requires the derivatives of ϕ to exist within V, which in turn allows the integral form of the conservation equation to be written as partial differential equations, which are upheld within the control volume.

Begin with the terms describing diffusive and advective fluxes into or out of the control volume. The divergence theorem states that

$$\int_{S} \mathbf{F} \cdot \hat{n} \ dS = \int_{V} \nabla \cdot \mathbf{F} \ dV. \tag{3.2}$$

Here and below, we will alternate between standard vector notation (as above) and index notation, where a single subscript indicates a component of a vector, two different subscripts

indicate a tensor quantity, and two repeated subscripts indicate summation. Thus, an identical way of writing equation (3.2) is

$$\int_{S} \mathbf{F}_{j} \ n_{j} \ dS = \int_{V} \frac{\partial \mathbf{F}_{j}}{\partial x_{j}} \ dV. \tag{3.3}$$

Using the divergence theorem, the surface integrals over fluxes in (3.1) may be replaced by

$$-\int_{S} \mathbf{F} \cdot \hat{n} dS - \int_{S} \phi \mathbf{u} \cdot \hat{n} dS = -\int_{V} \nabla \cdot (\mathbf{F} + \phi \mathbf{u}) dV.$$
 (3.4)

Assuming that the coordinate system is stationary with respect to the velocity field \mathbf{u} (i.e, assuming an Eularian reference frame), it is possible to write

$$\frac{\partial}{\partial t} \int_{V} \phi \ dV = \int_{V} \frac{\partial \phi}{\partial t} \ dV \tag{3.5}$$

The integral form of the conservation equation can now be written as

$$\int_{V} \left\{ \frac{\partial \phi}{\partial t} + \nabla \cdot (\mathbf{F} + \phi \mathbf{u}) - R \right\} dV = 0$$
 (3.6)

Because V is an arbitrary volume, this equation can be true only if the term in brackets is zero for the volume. Hence, for any volume having continuously differentiable ϕ ,

$$\frac{\partial \phi}{\partial t} + \nabla \cdot (\mathbf{F} + \phi \mathbf{u}) - R = 0. \tag{3.7}$$

This is the general form for all conservation laws in continuum mechanics. Below, we apply this equation to the three quantities conserved in an ice sheet (and ideally, in an ice sheet model): mass, energy, and momentum.

3.2 Applications of the General Conservation Equation

3.2.1 Conservation of momentum

Starting from Newton's second law of motion, conservation of momentum is expressed as

$$\frac{d}{dt} \int_{V} \rho u_i \ dV = \int_{V} \frac{\partial \sigma_{ij}}{\partial x_j} \ dV + \int_{V} \rho g_i \ dV \tag{3.8}$$

where t is time, ρ is density, u is velocity, σ_{ij} is the stress tensor, g is the acceleration due to gravity, V is the volume of an arbitrary fluid element, and $(i,j) = \{x,y,z\}$ in a Cartesian coordinate system. Equation (3.8) tells us that a fluid element of arbitrary size experiences a "body force" $\rho g_i \delta V$ due to gravity, which is balanced by stress divergence $\frac{\partial \sigma_{ij}}{\partial x_j} \delta V$ and acceleration of the fluid in the volume δV .

Making the assumptions that we have continuous fields and that ice is incompressible (i.e., its density ρ does not change under conditions of interest), we can write

$$\rho \frac{Du_i}{Dt} = \frac{\partial \sigma_{ij}}{\partial x_i} + \rho g_i \tag{3.9}$$

in which D is a material derivative. Because the Froude number for ice flow is extremely small, the acceleration term (the first term on the left-hand side) can be neglected, leaving the steady-state form,

$$\frac{\partial \sigma_{ij}}{\partial x_j} + \rho g_i = 0. ag{3.10}$$

Equation (3.10) states that the body force (the gravitational driving force) is balanced by forces resulting from gradients in the stress tensor σ_{ij} . All models of ice-flow dynamics are based on solving this equation in some form. Chapters 4 and 6 provide additional details on the approximations to this equation that are solved by CISM.

The stress tensor σ_{ij} has nine components in a three-dimensional, Cartesian coordinate system,

$$\sigma = \begin{vmatrix} \sigma_{xx} & \sigma_{xy} & \sigma_{xz} \\ \sigma_{yx} & \sigma_{yy} & \sigma_{yz} \\ \sigma_{zx} & \sigma_{zy} & \sigma_{zz}. \end{vmatrix}$$
(3.11)

Since σ_{ij} is symmetric, only six of these components are independent. The components along the diagonal are called normal stresses, and the off-diagonal components are called shear stresses. Deformation results not from the full stress but from the deviatoric stress,

$$\tau_{ij} = \sigma_{ij} - \frac{1}{3}\sigma_{kk}\delta_{ij}, \tag{3.12}$$

in which δ_{ij} is the Kroneker delta (or the identity tensor). For shear stresses, (3.12) indicates that the full and deviatoric stresses are identical.

Constitutive relationship

To relate the stress tensor to fluid motion, we introduce the strain rate tensor,

$$\dot{\epsilon}_{ij} = \frac{1}{2} \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right), \quad i, j = x, y, z, \tag{3.13}$$

where u_i are the velocity vector components. The strain rate tensor $\dot{\epsilon}_{ij}$, and hence gradients in the velocity field, are related to the stress tensor τ_{ij} by a constitutive relation. For a Newtonian fluid, this can be expressed as

$$\tau_{ij} = \eta \dot{\epsilon}_{ij}, \tag{3.14}$$

which states that the strain rate is proportional to the stress, with the ice viscosity η serving as the constant of proportionality. Ice does not behave as a Newtonian fluid, and instead exhibits a power-law rheology, so that it becomes more fluid (less viscous) the faster it deforms. This relationship can be expressed through Nye's generalization of Glen's flow law,

$$\tau_{ij} = A(T^*)^{\frac{-1}{n}} \dot{\epsilon}_e^{\frac{1-n}{n}} \dot{\epsilon}_{ij} \tag{3.15}$$

in which T^* is the absolute temperature corrected for the pressure dependence of the melt temperature, $\dot{\epsilon}_e$ is the second invariant (a norm) of the stress tensor, and the power-law exponent n is commonly taken as 3. A comparison of (3.14) and (3.15) indicates that one can define an "effective" ice viscosity for (3.14) as

$$\eta_e = A(T^*)^{\frac{-1}{n}} \dot{\epsilon}_e^{\frac{1-n}{n}}.$$
(3.16)

The temperature-dependent rate factor A follows the Arrhenius relationship

$$A(T^*) = EA_o e^{-Q/RT^*}, (3.17)$$

in which A_o is a constant, Q is the activation energy for crystal creep, R is the gas constant, and E is a tuning parameter, which can be used to account for the effects of impurities and anisotropic ice fabrics. The homologous temperature is

$$T^* = T + \rho g H \Phi, \tag{3.18}$$

in which Φ is 9.8×10^{-8} K Pa⁻¹, or about 8.7×10^{-4} K m⁻¹ in ice. The pressure-dependent melt temperature is simply the triple point temperature minus the product $\rho g H \Phi$.

3.2.2 Conservation of energy

The first law of thermodynamics is used to make a basic statement of conservation of energy in a volume of ice V enclosed within a surface S:

$$\frac{d}{dt} \int_{V} E \ dV = -\int_{S} \mathbf{F} \cdot \hat{n} \ dS - \int_{S} E \mathbf{u} \cdot \hat{n} \ dS + \int_{V} W dV$$
 (3.19)

in which E is the total energy within the volume, F_i is the energy flux due to diffusion, and W represents any sources or sinks of energy within the volume. The term Eu_i is an energy flux through S due to advection. Following the steps laid out earlier, we use the divergence theorem and the assumptions of continuous fields and incompressibility to obtain

$$\frac{dE}{dt} + \nabla \cdot (F_i + Eu_i) - W = 0 \tag{3.20}$$

Our goal is to use the first law of thermodynamics to compute the temperature of the ice and any changes in that temperature over time.

The energy E is the product of density and the specific internal energy of the ice e, which is itself the product of the specific heat capacity c_p and temperature T (because there is no transfer between internal energy and pressure for an incompressible fluid). Thus,

$$\frac{dE}{dt} = \frac{d(\rho e)}{dt}
= \rho \frac{de}{dt} + e \frac{d\rho}{dt}
= \rho c_p \frac{dT}{dt}$$
(3.21)

The heat flux due to diffusion follows Fourier's "law" for heat conduction,

$$\nabla \cdot F_i = \nabla \cdot (-k \nabla T)
= -k \nabla^2 T,$$
(3.22)

in which k is the thermal conductivity of ice and we assume gradients in its magnitude to be negligible.

Using progress made above, we can write the advection term

$$\nabla \cdot (Eu_i) = \rho c_p \ u_i \cdot \nabla T. \tag{3.23}$$

In the expansion of the terms on the left-hand side of (3.23) (using the product rule), we have implicitly ignored the term involving $\nabla \cdot u_i$ because it is small with respect to the other terms retained on the right-hand side.

Two energy sources must be considered: the work done on the system by internal deformation and the latent heat associated with phase changes. The former is the product of the strain rate and deviatoric stress, $\dot{\epsilon}_{ij}\tau_{ij}$. The latter is the product of the latent heat of fusion and the amount of material (ice) subject to melting (or freezing) per unit volume, per unit time, $L_f M_f$.

At last, we can write equation (3.20) in terms of temperature:

$$\frac{\partial T}{\partial t} = \frac{k}{\rho c_p} \nabla^2 T - u_i \cdot \nabla T + \frac{1}{\rho c_p} \dot{\epsilon}_{ij} \tau_{ij} + \frac{1}{\rho c_p} L_f M_f. \tag{3.24}$$

It is often the case that horizontal terms $\frac{\partial^2 T}{\partial x^2}$ and $\frac{\partial^2 T}{\partial y^2}$ are small enough to be ignored.

3.2.3 Conservation of mass

In this case, ϕ from our general conservation equation represents the mass M, or more conveniently $M = \int_V \rho dV$, the integral of the density over the volume. Assuming that there are no sources or sinks of mass in the volume (R=0), the mass conservation equation is written

$$\int_{V} \frac{\partial \rho}{\partial t} dV + \int_{V} \nabla \cdot \rho \mathbf{u} dV = 0$$
(3.25)

Ice is incompressible (the density does not change in time), which provides the equation for local mass continuity,

$$\nabla \cdot \mathbf{u} = 0. \tag{3.26}$$

Equation (3.26) says that the velocity field is divergence-free. Applying the ∇ operator in Cartesian coordinates gives

$$\frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} + \frac{\partial u_z}{\partial z} = 0. {(3.27)}$$

To make use of this statement, we need to integrate from the base b to the upper surface s of the ice mass,

$$\int_{b}^{s} \left(\frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} + \frac{\partial u_z}{\partial z} \right) dz = 0.$$
 (3.28)

The integral of $\frac{\partial u_z}{\partial z}$ is simply the difference between the vertical component of the velocity at the upper and lower surfaces, so

$$u_{z}(s) - u_{z}(b) = -\int_{b}^{s} \frac{\partial u_{x}}{\partial x} dz - \int_{b}^{s} \frac{\partial u_{y}}{\partial y} dz.$$
 (3.29)

Changing the order of integration and differentiation using Leibniz's rule, we obtain

$$u_{z}(s) - u_{z}(b) = -\frac{\partial}{\partial x} \int_{b}^{s} u_{x} dz + u_{x}(s) \frac{\partial s}{\partial x} - u_{x}(b) \frac{\partial b}{\partial x} - \frac{\partial}{\partial y} \int_{b}^{s} u_{y} dz + u_{y}(s) \frac{\partial s}{\partial y} - u_{y}(b) \frac{\partial b}{\partial x}.$$
(3.30)

The vertical velocity at the upper surface $u_z(s)$ is the result of motion parallel to the surface slope, the rate of new ice accumulation B_s , and any time change in the surface height,

$$u_z(s) = \frac{\partial s}{\partial t} + u_x(s)\frac{\partial s}{\partial x} + u_y(s)\frac{\partial s}{\partial y} - B_s,$$
 (3.31)

recognizing that a negative accumulation rate indicates ablation. Similarly, the vertical velocity at the lower surface is

$$u_z(b) = \frac{\partial b}{\partial t} + u_x(b)\frac{\partial b}{\partial x} + u_y(b)\frac{\partial b}{\partial y} - M_b$$
 (3.32)

in which M_b is the basal melt rate ($M_b < 0$ for freeze-on). Substituting (3.31) and (3.32) into (3.30), we find that many terms cancel:

$$\frac{\partial s}{\partial t} - B_s - \frac{\partial b}{\partial t} + M_b = -\frac{\partial}{\partial x} \int_{s}^{s} u_x dz - \frac{\partial}{\partial y} \int_{s}^{s} u_y dz.$$
 (3.33)

Finally, making the substitution that the ice thickness H = s - b, we obtain

$$\frac{\partial H}{\partial t} = -\frac{\partial}{\partial x} \int_{b}^{s} u_{x} dz - \frac{\partial}{\partial y} \int_{b}^{s} u_{y} dz + B_{s} - M_{b}. \tag{3.34}$$

Integrating in the vertical gives

$$\frac{\partial H}{\partial t} = -\nabla \cdot (U_i H) + B_s - M_b, \qquad (3.35)$$

in which U_i is the vertically averaged velocity, i.e., $U_i = \frac{1}{H} \int_b^s u_i dz$. Equation (3.35) is prognostic; we can use the current velocity and geometry of the ice to compute the rate of change in the geometry.

 $26 CHAPTER\ 3.\ INTRODUCTION\ TO\ ICE\ SHEET\ MODELING:\ DERIVATION\ OF\ FIELD\ EQUATIONS$

Chapter 4

law:

Shallow Ice Dynamics: Glide Dynamical Core

This chapter describes the numerical implementation of mass, momentum, and energy conservation in the Glide dynamical core, which uses a Shallow Ice Approximation for ice flow. For a model governed by shallow-ice dynamics, the solutions for the conservation of mass and momentum are intimately linked.

4.1 Ice Thickness Evolution

The evolution of the ice thickness, H, stems from the continuity equation and can be expressed as

$$\frac{\partial H}{\partial t} = -\nabla \cdot (\overline{\boldsymbol{u}}H) + B,\tag{4.1}$$

where $\overline{\boldsymbol{u}}$ is the vertically averaged ice velocity, B is the surface mass balance and ∇ is the horizontal gradient operator (Payne and Dongelmans, 1997).

For some regions of large-scale ice sheets, such as the slow-moving interior, or for simulations run at coarse spatial resolution, a model governed by the *shallow ice approximation* (SIA) may be appropriate. Further, for very long time integrations, such as those required in paleoclimate studies, a model governed by the SIA may be the only computationally practical approach.

Based largely on the assumption that bedrock and ice surface slopes are sufficiently small (Hutter, 1983), the SIA neglects all stress components other than those associated with vertical shearing in the horizontal directions. These stresses, τ_{xz} and τ_{yz} , are approximated by

$$\tau_{xz}(z) = -\rho g(s-z) \frac{\partial s}{\partial x},$$

$$\tau_{yz}(z) = -\rho g(s-z) \frac{\partial s}{\partial y},$$
(4.2)

where ρ is the density of ice, g the acceleration due to gravity and s = H + b the ice surface. Strain rates $\dot{\epsilon}_{ij}$ of polycrystalline ice are related to the stress tensor by the nonlinear flow

$$\dot{\epsilon}_{iz} = \frac{1}{2} \left(\frac{\partial u_i}{\partial z} + \frac{\partial u_z}{\partial i} \right) = A(T^*) \tau_*^{(n-1)} \tau_{iz} \qquad i = x, y, \tag{4.3}$$

where τ_* is the effective shear stress defined by the second invariant of the stress tensor, n the flow law exponent and A the temperature–dependent flow law coefficient. T^* is the absolute temperature corrected for the dependence of the melting point on pressure $(T^* = T + 8.7 \cdot 10^{-4} (s - z), T$ in Kelvin, Huybrechts, 1986). The parameters n and A are determined experimentally; n

is usually taken to be 3 and A depends primarily on temperature and secondarily on factors such as crystal size and orientation and ice impurities. Experiments suggest that A follows the Arrhenius relationship:

$$A(T^*) = fae^{-Q/RT^*}, (4.4)$$

where a is a temperature–independent material constant, Q is the activation energy for creep and R is the universal gas constant (Paterson, 1994). The tuning parameter f may be used to speed up the ice flow, accounting for the effects of ice impurities and the development of anisotropic ice fabrics (Payne, 1999; Tarasov and Peltier, 1999, 2000; Peltier et al., 2000).

Integrating (4.4) with respect to z gives the vertical profile of the horizontal velocity in each column:

$$\boldsymbol{u}(z) - \boldsymbol{u}(b) = -2(\rho g)^n |\nabla s|^{n-1} \nabla s \int_b^z A(s-z)^n dz, \tag{4.5}$$

where u(b) is the basal sliding velocity. Integrating (4.5) again with respect to z gives an expression for the vertically averaged ice velocity:

$$\overline{\boldsymbol{u}}H = -2(\rho g)^n |\nabla s|^{n-1} \nabla s \int_b^s \int_b^z A(s-z)^n dz dz'. \tag{4.6}$$

The vertical ice velocity can be derived from the conservation of mass for an incompressible material:

$$\frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} + \frac{\partial u_z}{\partial z} = 0. \tag{4.7}$$

Integrating (4.7) with respect to z gives the vertical profile of the vertical velocity in each column:

$$w(z) = -\int_{b}^{z} \nabla \cdot \boldsymbol{u}(z)dz + w(b), \tag{4.8}$$

with lower kinematic boundary condition

$$w(b) = \frac{\partial b}{\partial t} + \boldsymbol{u}(b) \cdot \boldsymbol{\nabla} b - M_b, \tag{4.9}$$

where M_b is the basal melt rate given by (4.55). The upper kinematic boundary is given by the surface mass balance and must satisfy

$$w(s) = \frac{\partial s}{\partial t} + \boldsymbol{u}(s) \cdot \boldsymbol{\nabla} s - B_s. \tag{4.10}$$

4.1.1 Numerical grid

The continuous equations describing ice physics have to be discretized in order to be solved by a computer (which is inherently finite). This section describes the finite—difference grids used by the model.

Horizontal grid

The modelled region $(x \in [0, L_x], y \in [0, L_y])$ is discretized using a regular grid so that $x_i = (i-1)\Delta x$ for $i \in [1, N]$ (and similarly for y_j). The model uses two staggered horizontal grids in order to improve stability. Both grids use the same grid spacing, Δx and Δy , but are offset by half a grid cell (see Fig. 4.1). Quantities calculated on the staggered (r, s)-grid are denoted with

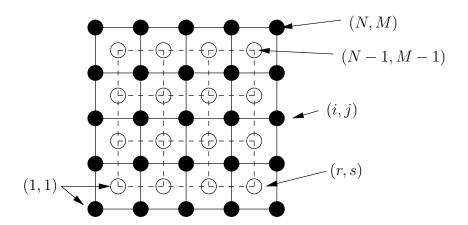


Figure 4.1: Horizontal Grid.

a tilde, i.e., \tilde{F} . Quantities are transformed between grids by averaging over the surrounding nodes; i.e., a quantity in the (i,j)-grid becomes in the (r,s)-grid:

$$\tilde{F}_{r,s} = \tilde{F}_{i+\frac{1}{2},j+\frac{1}{2}} = \frac{1}{4} (F_{i,j} + F_{i+1,j} + F_{i+1,j+1} + F_{i,j+1}), \tag{4.11a}$$

and similarly for the reverse transformation:

$$F_{i,j} = F_{r-\frac{1}{2},s-\frac{1}{2}} = \frac{1}{4} (\tilde{F}_{r-1,s-1} + \tilde{F}_{r,s-1} + \tilde{F}_{r,s} + \tilde{F}_{r-1,s}). \tag{4.11b}$$

In general, horizontal velocities and associated quantities like the diffusivity are calculated on the (r,s)-grid. Ice thickness, temperatures and vertical velocities are calculated on the (i,j)-grid.

Horizontal gradients are calculated on the (r, s)-grid; i.e., surface gradients are

$$\left(\frac{\partial s}{\partial x}\right)_{r,s} = \tilde{s}_{r,s}^x = \frac{s_{i+1,j} - s_{i,j} + s_{i+1,j+1} - s_{i,j+1}}{2\Delta x},\tag{4.12a}$$

$$\left(\frac{\partial s}{\partial y}\right)_{r,s} = \tilde{s}_{r,s}^{y} = \frac{s_{i,j+1} - s_{i,j} + s_{i+1,j+1} - s_{i+1,j}}{2\Delta y}.$$
(4.12b)

Ice thickness gradients, $\tilde{H}_{r,s}^x$ and $\tilde{H}_{r,s}^y$, are formed analogously. Gradients in the (r,s)-grid are formed in a similar way:

$$\left(\frac{\partial u}{\partial x}\right)_{i,j} = u_{i,j}^x = \frac{\tilde{u}_{r,s-1} - \tilde{u}_{r-1,s-1} + \tilde{u}_{r,s} - \tilde{u}_{r-1,s}}{2\Delta x}.$$
(4.13)

Periodic boundary conditions

The model can be run with horizontal periodic boundary conditions, i.e. with the western edge of the modelled region joined to the eastern edge. Figure 4.2 illustrates the numeric grid when the model is run in torus mode.

These boundary conditions are enforced by exchanging points for the temperature and vertical velocity calculations. The ice thicknesses are calculated explicitly at the ghostpoints.

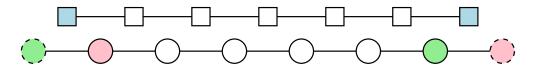


Figure 4.2: A row of the numeric grid when the model is used in torus mode. Circles indicate points in (i, j)-grid and squares indicate points in the (r, s)-grid. Points with the same color are logically the same.

σ -coordinate system

The vertical coordinate, z, is scaled by the ice thickness analogous to the s-coordinate in numerical weather simulations (e.g., Holton, 1992). A new vertical coordinate, σ , is introduced so that the ice surface is at $\sigma = 0$ and the ice base at $\sigma = 1$ (see Fig. 4.3), i.e.

$$\sigma = \frac{s - z}{H}.\tag{4.14}$$

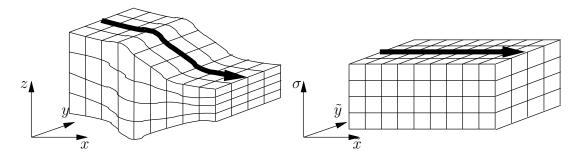


Figure 4.3: Vertical scaling of the ice sheet model. The vertical axis is scaled to unity. The horizontal coordinates are not changed.

The derivatives of a function f in (x, y, z, t) become in the new $(\tilde{x}, \tilde{y}, \sigma, \tilde{t})$ system:

$$\frac{\partial f}{\partial x} = \frac{\partial f}{\partial \tilde{x}} + \frac{1}{H} \Delta_{\tilde{x}} \frac{\partial f}{\partial \sigma}, \tag{4.15a}$$

$$\frac{\partial f}{\partial y} = \frac{\partial f}{\partial \tilde{y}} + \frac{1}{H} \Delta_{\tilde{y}} \frac{\partial f}{\partial \sigma}, \qquad (4.15b)$$

$$\frac{\partial f}{\partial t} = \frac{\partial f}{\partial \tilde{t}} + \frac{1}{H} \Delta_{\tilde{t}} \frac{\partial f}{\partial \sigma}, \qquad (4.15c)$$

$$\frac{\partial f}{\partial z} = -\frac{1}{H} \frac{\partial f}{\partial \sigma}, \qquad (4.15d)$$

$$\frac{\partial f}{\partial t} = \frac{\partial f}{\partial \tilde{t}} + \frac{1}{H} \Delta_{\tilde{t}} \frac{\partial f}{\partial \sigma}, \tag{4.15c}$$

$$\frac{\partial f}{\partial z} = -\frac{1}{H} \frac{\partial f}{\partial \sigma},\tag{4.15d}$$

where the geometric factors, $\Delta_{\tilde{x}}$, $\Delta_{\tilde{y}}$ and $\Delta_{\tilde{t}}$, are defined by

$$\Delta_{\tilde{x}} = \left(\frac{\partial s}{\partial \tilde{x}} - \sigma \frac{\partial H}{\partial \tilde{x}}\right),\tag{4.16a}$$

$$\Delta_{\tilde{y}} = \left(\frac{\partial s}{\partial \tilde{y}} - \sigma \frac{\partial H}{\partial \tilde{y}}\right),\tag{4.16b}$$

$$\Delta_{\tilde{t}} = \left(\frac{\partial s}{\partial \tilde{t}} - \sigma \frac{\partial H}{\partial \tilde{t}}\right). \tag{4.16c}$$

The integral of z becomes in the σ -coordinate system:

$$\int_{b}^{z} f dz = -H \int_{1}^{\sigma} f d\sigma. \tag{4.17}$$

The vertical coordinate can be discretized using an irregular grid spacing to reflect the fact that ice flow is more variable at the bottom of the ice column. In the vertical the index k is used.

4.1.2 Ice sheet equations in σ -coordinates

The horizontal velocity, (4.5), becomes in the σ -coordinate system

$$\boldsymbol{u}(\sigma) = -2(\rho g)^n H^{n+1} |\nabla s|^{n-1} \nabla s \int_{1}^{\sigma} A\sigma^n d\sigma + \boldsymbol{u}(1). \tag{4.18}$$

The vertically averaged velocity is

$$\overline{\boldsymbol{u}}H = H \int_{0}^{1} \boldsymbol{u}d\sigma + \boldsymbol{u}(1)H. \tag{4.19}$$

The vertical velocity (4.8) becomes

$$w(\sigma) = -\int_{1}^{\sigma} \left(\frac{\partial \boldsymbol{u}}{\partial \sigma} \cdot (\boldsymbol{\nabla} s - \sigma \boldsymbol{\nabla} H) + H \boldsymbol{\nabla} \cdot \boldsymbol{u} \right) d\sigma + w(1), \tag{4.20}$$

with lower boundary condition

$$w(1) = \frac{\partial b}{\partial t} + \boldsymbol{u}(1) \cdot \boldsymbol{\nabla} b - M_b. \tag{4.21}$$

4.1.3 Calculating the horizontal velocity and diffusivity

Horizontal velocity and diffusivity calculations are split up into two parts:

$$\boldsymbol{u}(\sigma) = c\boldsymbol{\nabla}s + \boldsymbol{u}(1),\tag{4.22a}$$

$$D = H \int_{0}^{1} cd\sigma, \tag{4.22b}$$

$$q = D\nabla s + Hu(1), \tag{4.22c}$$

with

$$c(\sigma) = -2(\rho g)^n H^{n+1} |\nabla s|^{n-1} \int_{1}^{\sigma} A\sigma^n d\sigma.$$
 (4.22d)

Quantities \boldsymbol{u} and D are found on the velocity grid. Integrating from the ice base (k = N - 1), the discretized quantities become

$$\tilde{c}_{r,s,N} = 0, (4.23a)$$

$$\tilde{c}_{r,s,k} = -2(\rho g)^n H_{r,s}^{n+1} \left((\tilde{s}_{r,s}^x)^2 + (\tilde{s}_{r,s}^y)^2 \right)^{\frac{n-1}{2}}$$

$$\sum_{\kappa=N-1}^k \frac{A_{r,s,\kappa} + A_{r,s,\kappa+1}}{2} \left(\frac{\sigma_{\kappa+1} + \sigma_{\kappa}}{2} \right)^n (\sigma_{\kappa+1} - \sigma_{\kappa}), \quad (4.23b)$$

$$\tilde{D}_{r,s} = H_{r,s} \sum_{k=0}^{N-1} \frac{\tilde{c}_{r,s,k} + \tilde{c}_{r,s,k+1}}{2} (\sigma_{k+1} - \sigma_k). \tag{4.23c}$$

Expressions for $\boldsymbol{u}_{i,j,k}$ and $\boldsymbol{q}_{i,j}$ are straightforward.

4.1.4 Solving the ice thickness evolution equation

Equation (4.1) can be rewritten as a diffusion equation, with nonlinear diffusion coefficient D:

$$\frac{\partial H}{\partial t} = -\nabla \cdot D\nabla s + B = -\nabla \cdot \boldsymbol{q} + B, \tag{4.24}$$

where $B = B_s - M_b$ is the total mass balance. This nonlinear partial differential equation can be linearized by using the diffusion coefficient from the previous time step. The diffusion coefficient is calculated on the (r, s)-grid, i.e. staggered in both x and y directions. Figure 4.4 illustrates the staggered grid. Using finite differences, the fluxes in the x direction, q^x , become

$$q_{i+\frac{1}{2},j}^x = -\frac{1}{2}(\tilde{D}_{r,s} + \tilde{D}_{r,s-1})\frac{s_{i+1,j} - s_{i,j}}{\Delta x},$$
(4.25a)

$$q_{i-\frac{1}{2},j}^{x} = -\frac{1}{2}(\tilde{D}_{r-1,s} + \tilde{D}_{r-1,s-1})\frac{s_{i,j} - s_{i-1,j}}{\Delta x},$$
(4.25b)

and the fluxes in the y direction are

$$q_{i,j+\frac{1}{2}}^{y} = -\frac{1}{2}(\tilde{D}_{r,s} + \tilde{D}_{r-1,s})\frac{s_{i,j+1} - s_{i,j}}{\Delta y}, \tag{4.25c}$$

$$q_{i,j-\frac{1}{2}}^{y} = -\frac{1}{2}(\tilde{D}_{r,s-1} + \tilde{D}_{r-1,s-1})\frac{s_{i,j} - s_{i,j-1}}{\Delta y}.$$
(4.25d)

ADI Scheme

The alternating–direction implicit (ADI) method uses the concept of operator splitting where (4.24) is solved first in the x-direction and then in the y-direction (Press et al., 1992). The time step Δt is divided into two partial steps $\Delta t/2$. The discretized version of Equation (4.24) becomes (Huybrechts, 1986):

$$2\frac{H_{i,j}^{t+\frac{1}{2}} - H_{i,j}^{t}}{\Delta t} = -\frac{q_{i+\frac{1}{2},j}^{x,t+\frac{1}{2}} - q_{i-\frac{1}{2},j}^{x,t+\frac{1}{2}}}{\Delta x} - \frac{q_{i,j+\frac{1}{2}}^{y,t} - q_{i,j-\frac{1}{2}}^{y,t}}{\Delta y} + B_{i,j},$$
(4.26a)

$$2\frac{H_{i,j}^{t+1} - H_{i,j}^{t+\frac{1}{2}}}{\Delta t} = -\frac{q_{i+\frac{1}{2},j}^{x,t+\frac{1}{2}} - q_{i-\frac{1}{2},j}^{x,t+\frac{1}{2}}}{\Delta x} - \frac{q_{i,j+\frac{1}{2}}^{y,t+1} - q_{i,j-\frac{1}{2}}^{y,t+1}}{\Delta y} + B_{i,j}.$$
 (4.26b)

Gathering all $t+\frac{1}{2}$ terms on the left side, (4.26a) can be expressed as a tridiagonal set of equations for each row j:

$$-\alpha_{i,j}H_{i-1,j}^{t+\frac{1}{2}} + (1-\beta_{i,j})H_{i,j}^{t+\frac{1}{2}} - \gamma_{i,j}H_{i+1,j}^{t+\frac{1}{2}} = \delta_{i,j}, \tag{4.27}$$

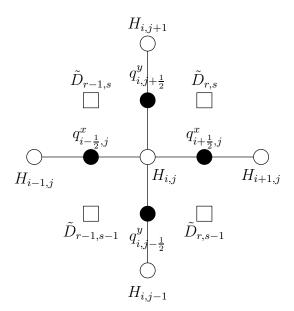


Figure 4.4: Illustration of the staggered grid used to calculate ice thicknesses, diffusivities and mass fluxes.

with

$$\alpha_{i,j} = \frac{\tilde{D}_{r-1,s} + \tilde{D}_{r-1,s-1}}{4\Delta x^2} \Delta t,$$

$$\beta_{i,j} = -\frac{\tilde{D}_{r,s} + 2\tilde{D}_{r-1,s} + \tilde{D}_{r-1,s-1}}{4\Delta x^2} \Delta t = -(\alpha_{i,j} + \gamma_{i,j}),$$

$$\gamma_{i,j} = \frac{\tilde{D}_{r,s} + \tilde{D}_{r,s-1}}{4\Delta x^2} \Delta t.$$
(4.28a)
$$(4.28a)$$

$$\beta_{i,j} = -\frac{\tilde{D}_{r,s} + 2\tilde{D}_{r-1,s} + \tilde{D}_{r-1,s-1}}{4\Delta x^2} \Delta t = -(\alpha_{i,j} + \gamma_{i,j}), \tag{4.28b}$$

$$\gamma_{i,j} = \frac{\tilde{D}_{r,s} + \tilde{D}_{r,s-1}}{4\Delta x^2} \Delta t. \tag{4.28c}$$

The RHS is given by

$$\delta_{i,j} = H_{i,j}^t - \frac{\Delta t}{2\Delta y} \left(q_{i,j+\frac{1}{2}}^{y,t} - q_{i,j-\frac{1}{2}}^{y,t} \right) + \frac{\Delta t}{2} B_{i,j} + \alpha_{i,j} h_{i-1,j} - \beta_{i,j} h_{i,j} + \gamma_{i,j} h_{i+1,j}.$$
 (4.28d)

A similar tridiagonal system is found for each column i of (4.26b).

Linearised semi-implicit scheme

Using the Crank-Nicolson scheme, the semi-implicit temporal discretisation of (4.24) is

$$\frac{H_{i,j}^{t+1} - H_{i,j}^{t}}{\Delta t} = \frac{q_{i+\frac{1}{2},j}^{x,t+1} - q_{i-\frac{1}{2},j}^{x,t+1}}{2\Delta x} + \frac{q_{i,j+\frac{1}{2}}^{y,t+1} - q_{i,j-\frac{1}{2}}^{y,t+1}}{2\Delta y} + \frac{q_{i,j+\frac{1}{2}}^{x,t} - q_{i-\frac{1}{2},j}^{x,t}}{2\Delta x} + \frac{q_{i,j+\frac{1}{2}}^{y,t} - q_{i,j+\frac{1}{2}}^{y,t}}{2\Delta x} + B_{i,j}. \quad (4.29)$$

The superscripts t and $^{t+1}$ indicate at what time the ice thickness H is evaluated. Collecting all H^{t+1} terms of (4.29) on the LHS and moving all other terms to the RHS, we can rewrite (4.29) as

$$-\alpha_{i,j}H_{i-1,j}^{t+1} - \beta_{i,j}H_{i+1,j}^{t+1} - \gamma_{i,j}H_{i,j-1}^{t+1} - \delta_{i,j}H_{i,j+1}^{t+1} + (1 - \epsilon_{i,j})H_{i,j}^{t+1} = \zeta_{i,j}, \tag{4.30}$$

with the RHS

$$\zeta_{i,j} = \alpha_{i,j} H_{i-1,j}^t + \beta_{i,j} H_{i+1,j}^t + \gamma_{i,j} H_{i,j-1}^t + \delta_{i,j} H_{i,j+1}^t + (1 + \epsilon_{i,j}) H_{i,j}^t + 2(\alpha_{i,j} h_{i-1,j} + \beta_{i,j} h_{i+1,j} + \gamma_{i,j} h_{i,j-1} + \delta_{i,j} h_{i,j+1} + \epsilon_{i,j} h_{i,j}) + B_{i,j} \Delta t.$$
(4.31)

Here the elements of the sparse matrix are

$$\alpha_{i,j} = \frac{\tilde{D}_{r-1,s} + \tilde{D}_{r-1,s-1}}{4\Lambda r^2} \Delta t, \tag{4.32a}$$

$$\beta_{i,j} = \frac{\tilde{D}_{r,s} + \tilde{D}_{r,s-1}}{4\Delta x^2} \Delta t, \tag{4.32b}$$

$$\gamma_{i,j} = \frac{\tilde{D}_{r,s-1} + \tilde{D}_{r-1,s-1}}{4\Delta y^2} \Delta t, \tag{4.32c}$$

$$\delta_{i,j} = \frac{\tilde{D}_{r,s} + \tilde{D}_{r-1,s}}{4\Delta u^2} \Delta t, \tag{4.32d}$$

$$\epsilon_{i,j} = -(\alpha_{i,j} + \beta_{i,j} + \gamma_{i,j} + \delta_{i,j}). \tag{4.32e}$$

This matrix equation is solved using an iterative solver for non-symmetric sparse matrices. The solver used here is the bi–conjugate gradient method with incomplete LU decomposition preconditioning provided by the SLAP package.

Nonlinear scheme

The nonlinearity of (4.24) arises from the dependance of D on s. A nonlinear scheme for (4.24) can be formulated using Picard iteration, which consists of two iterations: an outer, nonlinear and an inner, linear equation. The scheme starts with the diffusivity from the previous time step:

$$D^{(0),t+1} = D^t. (4.33a)$$

Equation (4.30) becomes

$$-\alpha_{i,j}^{(\xi),t+1}H_{i-1,j}^{t+1} - \beta_{i,j}^{(\xi),t+1}H_{i+1,j}^{(\xi+1),t+1} - \gamma_{i,j}^{(\xi),t+1}H_{i,j-1}^{(\xi+1),t+1} - \delta_{i,j}^{(\xi),t+1}H_{i,j+1}^{(\xi+1),t+1} + (1 - \epsilon_{i,j}^{(\xi),t+1})H_{i,j}^{(\xi+1),t+1} = \zeta_{i,j}^{(0),t}. \quad (4.33b)$$

Equation (4.33b) is iterated over ξ until the maximum ice thickness residual is smaller than some threshold:

$$\max\left(\left|H^{(\xi+1),t+1} - H^{(\xi),t+1}\right|\right) < H_{\text{res}}.$$
 (4.34)

4.1.5 Calculating vertical velocities

Grid velocity

The vertical grid moves as a result of using a σ -coordinate system. The grid velocity is

$$w^{\text{grid}}(\sigma) = \frac{\partial s}{\partial t} + \boldsymbol{u} \cdot \boldsymbol{\nabla} s - \sigma \left(\frac{\partial H}{\partial t} + \boldsymbol{u} \cdot \boldsymbol{\nabla} H \right). \tag{4.35}$$

The numerical implementation of (4.35) is straightforward.

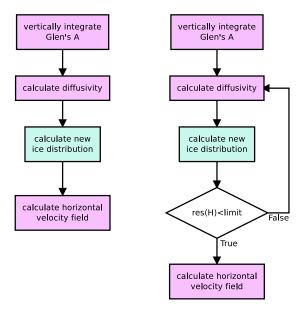


Figure 4.5: Flow diagram showing how the linearised solver (on the left) and the non–linear solver work. The inner, linear iteration is contained within the box labeled "calculate new ice distribution".

Vertical velocity

The discretized version of the vertical velocity equation (4.20) is slightly more compilicated because the horizontal velocities are calculated on the (r,s) grid. The vertical velocity at the ice base is $w_{i,j,N} = w_{i,j,N}^{\text{grid}} - M_b i, j$, where $M_b i, j$ is the basal melt rate. Integrating from the bottom, the vertical velocity is then

$$w_{i,j,k} = -\sum_{\tilde{k}=N-1}^{1} \left\{ \mathcal{H}_{i,j} \left(\frac{u_{i,j,k}^{x} + u_{i,j,k+1}^{x}}{2} + \frac{v_{i,j,k}^{y} + v_{i,j,k+1}^{y}}{2} \right) (\sigma_{k+1} - \sigma_{k}) + (\tilde{u}_{i,j,k+1} - \tilde{u}_{i,j,k}) \left(\tilde{s}_{i,j}^{x} - \frac{1}{2} (\sigma_{k+1} + \sigma_{k}) \tilde{H}_{i,j}^{x} \right) + (\tilde{v}_{i,j,k+1} - \tilde{v}_{i,j,k}) \left(\tilde{s}_{i,j}^{y} - \frac{1}{2} (\sigma_{k+1} + \sigma_{k}) \tilde{H}_{i,j}^{y} \right) \right\} + w_{i,j,N},$$

$$(4.36)$$

with the weighted ice thickness

$$\begin{split} \mathcal{H}_{i,j} &= \frac{4H_{i,j} + 2(H_{i-1,j} + H_{i+1,j} + H_{i,j-1} + H_{i,j+1})}{16} \\ &\quad + \frac{H_{i-1,j-1} + H_{i+1,j-1} + H_{i+1,j+1} + H_{i-1,j+1}}{16}. \end{split}$$

This scheme produces vertical velocities at the ice divide which are too small. The vertical velocities on the ice surface are given by the upper kinematic boundary condition (4.10). Equation (4.36) can be corrected with

$$w_{i,j,k}^* = w_{i,j,k} - (1 - \sigma_k)(w_{i,j,k} - w_{s_{i,j}}), \tag{4.37}$$

where $w_{si,j}$ is the vertical velocity at the ice surface given by (4.10). Figure 4.6 shows the different vertical velocities at the ice surface. The difference between the vertical velocities calculated by the model and the vertical velocities given by (4.10) at the ice margin are due

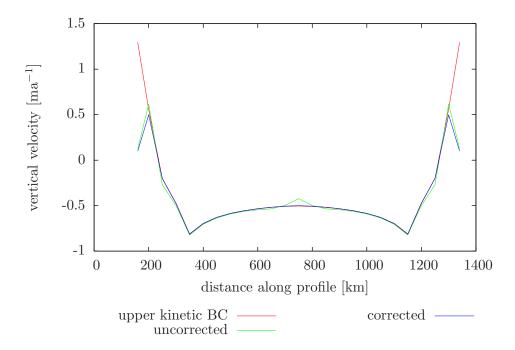


Figure 4.6: Vertical ice surface velocities of the EISMINT-1 moving margin experiment.

to the fact that temperatures and velocities are only calculated when the ice is thicker than a certain threshold value which is not met at the ice margin.

Figure 4.7 shows vertical profiles of the vertical velocity at the ice divide and a point halfway between the divide and the domain margin. A corresponding temperature profile is also shown since the vertical velocity determines the vertical temperature advection (see Section 4.2.4).

4.2 Temperature Solver

The flow law coefficient A in (4.3) depends on the ice temperature. Thus it is necessary to determine how the distribution of ice temperatures changes with a changing ice sheet configuration. The thermal evolution of the ice sheet is described by

$$\frac{\partial T}{\partial t} = \frac{k}{\rho c} \nabla^2 T - \boldsymbol{u} \cdot \boldsymbol{\nabla} T + \frac{\Phi}{\rho c} - w \frac{\partial T}{\partial z}, \tag{4.38}$$

where T is the absolute temperature, k is the thermal conductivity of ice, c is the specific heat capacity, and Φ is the heat generated by internal friction. In the σ -coordinate system, (4.38) becomes

$$\frac{\partial T}{\partial t} = \frac{k}{\rho c H^2} \frac{\partial^2 T}{\partial \sigma^2} - \boldsymbol{u} \cdot \boldsymbol{\nabla} T + \frac{\sigma g}{c} \frac{\partial \boldsymbol{u}}{\partial \sigma} \cdot \boldsymbol{\nabla} s + \frac{1}{H} \frac{\partial T}{\partial \sigma} \left(w - w_{\text{grid}} \right). \tag{4.39}$$

The terms on the RHS represent (1) vertical diffusion, (2) horizontal advection, (3) internal heat generation due to friction, and (4) vertical advection and a correction due to the σ -coordinate system. We rewrite (4.39) as

$$\frac{\partial T}{\partial t} = a \frac{\partial^2 T}{\partial \sigma^2} + b(\sigma) + \Phi(\sigma) + c(\sigma) \frac{\partial T}{\partial \sigma}, \tag{4.40}$$

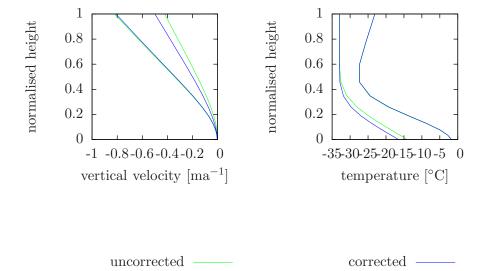


Figure 4.7: Vertical velocity and temperature distribution for columns at the ice divide and a point halfway between the divide and the domain margin.

where

$$a = \frac{k}{\rho c H^2}$$

$$b(\sigma) = -\mathbf{u} \cdot \nabla T$$
(4.41a)
$$(4.41b)$$

$$b(\sigma) = -\boldsymbol{u} \cdot \boldsymbol{\nabla} T \tag{4.41b}$$

$$\Phi(\sigma) = \frac{\sigma g}{c} \frac{\partial \mathbf{u}}{\partial \sigma} \cdot \nabla s \tag{4.41c}$$

$$\Phi(\sigma) = \frac{\sigma g}{c} \frac{\partial \mathbf{u}}{\partial \sigma} \cdot \nabla s \tag{4.41c}$$

$$c(\sigma) = \frac{1}{H} (w - w_{\text{grid}}) \tag{4.41d}$$

(Note the different meanings for b and c here relative to usage elsewhere in this document.)

Vertical diffusion 4.2.1

Discretization of $\partial^2 T/\partial \sigma^2$ is slightly complicated because the vertical grid is irregular. Using Taylor series the central difference formulas are

$$\left. \frac{\partial T}{\partial \sigma} \right|_{\sigma_{k-1/2}} = \frac{T_k - T_{k-1}}{\sigma_k - \sigma_{k-1}} \tag{4.42a}$$

and

$$\left. \frac{\partial T}{\partial \sigma} \right|_{\sigma_{k+1/2}} = \frac{T_{k+1} - T_k}{\sigma_{k+1} - \sigma_k}.$$
(4.42b)

The second partial derivative is then, also using central differences:

$$\frac{\partial^2 T}{\partial \sigma^2}\Big|_{\sigma_k} = \frac{\partial T/\partial \sigma|_{\sigma_{k+1/2}} - \partial T/\partial \sigma|_{\sigma_{k-1/2}}}{1/2 \left(\sigma_{k+1} - \sigma_{k-1}\right)}.$$
(4.42c)

Inserting (4.42a) and (4.42b) into (4.42c), we obtain

$$\frac{\partial^2 T}{\partial \sigma^2}\Big|_{\sigma_k} = \frac{2(T_{k+1} - T_k)}{(\sigma_{k+1} - \sigma_k)(\sigma_{k+1} - \sigma_{k-1})} - \frac{2(T_k - T_{k-1})}{(\sigma_k - \sigma_{k-1})(\sigma_{k+1} - \sigma_{k-1})}.$$
(4.42d)

Finally, the terms of equation (4.42d) are rearranged:

$$\frac{\partial^2 T}{\partial \sigma^2} \Big|_{\sigma_k} = \frac{2T_{k-1}}{(\sigma_k - \sigma_{k-1})(\sigma_{k+1} - \sigma_{k-1})} - \frac{2T_k}{(\sigma_{k+1} - \sigma_k)(\sigma_k - \sigma_{k-1})} + \frac{2T_{k+1}}{(\sigma_{k+1} - \sigma_k)(\sigma_{k+1} - \sigma_{k-1})}.$$
(4.43)

4.2.2 Horizontal advection

The horizontal advection term, $-\mathbf{u} \cdot \nabla T$, is found using an upwinding scheme. We start with the 1D case. (The method can readily be extended to 2D.) As always, the temperature function is expressed as a Taylor series:

$$T(x + \Delta x) = T(x) + \Delta x T'(x) + \frac{\Delta x^2}{2} T''(x) + \dots$$
 (4.44a)

If we replace Δx with $2\Delta x$, (4.44a) becomes

$$T(x + 2\Delta x) = T(x) + 2\Delta x T'(x) + 2\Delta x^2 T''(x) + \dots$$
 (4.44b)

From (4.44a) and (4.44b) we can construct a difference formula where the $\mathcal{O}(\Delta x^2)$ error is cancelled, by multiplying (4.44a) with 4 and substracting the result from (4.44b):

$$T'_{+}(x) = \frac{4T(x + \Delta x) - T(x + 2\Delta x) - 3T(x)}{2\Delta x},$$
(4.45a)

and similarly for the backward difference:

$$T'_{-}(x) = -\frac{4T(x - \Delta x) - T(x - 2\Delta x) - 3T(x)}{2\Delta x}.$$
 (4.45b)

So the horizontal advection term in 1D becomes

$$b_x = -u_x \frac{\partial T}{\partial x} = \frac{-u_x}{2\Delta x} \begin{cases} -(4T_{i-1} - T_{i-2} - 3T_i) & \text{when } u_x > 0\\ 4T_{i+1} - T_{i+2} - 3T_i & \text{when } u_x < 0 \end{cases}$$
(4.46)

A similar expression is found for b_y by substituting y for x. The combined horizontal advection term is

$$b = -\mathbf{u} \cdot \nabla T = -\left(u_x \frac{\partial T}{\partial x} + u_y \frac{\partial T}{\partial y}\right) = b_x + b_y = b_1 + b_2 T_i. \tag{4.47}$$

4.2.3 Heat generation

Taking the derivative of (4.18) with respect to σ , we get

$$\frac{\partial u_x}{\partial \sigma} = -2(\rho g)^n H^{n+1} |\nabla s|^{n-1} \frac{\partial s}{\partial x} A(T^*) \sigma^n. \tag{4.48}$$

Thus,

$$\Phi(\sigma) = \frac{\sigma g}{c} \frac{\partial \mathbf{u}}{\partial \sigma} \cdot \nabla s = \frac{\sigma g}{c} \left(\frac{\partial u_x}{\partial \sigma} \frac{\partial s}{\partial x} + \frac{\partial u_y}{\partial \sigma} \frac{\partial s}{\partial y} \right)
= -2(\rho g)^n H^{n+1} |\nabla s|^{n-1} \frac{\sigma g}{c} A(T^*) \sigma^n \left(\left(\frac{\partial s}{\partial x} \right)^2 + \left(\frac{\partial s}{\partial y} \right)^2 \right)
= -\frac{2}{c\rho} (g\sigma\rho)^{n+1} (H|\nabla s|)^{n+1} A(T^*).$$
(4.49)

The constant factor $\frac{2}{c\rho}(g\sigma\rho)^{n+1}$ is calculated during initialization in the subroutine glide_init_temp. This factor is assigned to array c1(1:upn). c1 also includes various scaling factors and the factor 1/16 to normalise \mathcal{A} .

The factor $(H|\nabla s|)^{n+1}$ is calculated in subroutine glide_finddisp:

$$c_{2i,j} = \left(\tilde{H}_{i,j} \sqrt{\tilde{S}_{xi,j}^2 + \tilde{S}_{yi,j}^2}\right)^{n+1}.$$
 (4.50)

The final factor is found by averaging over the neighboring nodes:

$$\mathcal{A}_{i,j} = 4A_{i,j} + 2(A_{i-1,j} + A_{i+1,j} + A_{i,j-1} + A_{i,j+1}) + (A_{i-1,j-1} + A_{i+1,j-1} + A_{i+1,j+1} + A_{i-1,j+1}). \tag{4.51}$$

4.2.4 Vertical advection

The vertical advection term $\partial T/\partial \sigma$ is computed from the central difference formula for unevenly spaced nodes:

$$\frac{\partial T}{\partial \sigma} = \frac{T_{k+1} - T_{k-1}}{\sigma_{k+1} - \sigma_{k-1}}. (4.52)$$

4.2.5 Boundary conditions

At the upper boundary, ice temperatures are set to the surface temperature, T_{surf} . The ice at the base is heated by the geothermal heat flux and sliding friction:

$$\left. \frac{\partial T}{\partial \sigma} \right|_{\sigma=1} = -\frac{GH}{k} - \frac{H\boldsymbol{\tau}_b \cdot \boldsymbol{u}(1)}{k},\tag{4.53}$$

where $\tau_b = -\rho g H \nabla s$ is the basal shear stress and u(1) is the basal ice velocity. Ice temperatures are held constant if they reach the pressure melting point of ice, i.e.

$$T^* = T_{\text{pmp}} \quad \text{if } T \ge T_{\text{pmp}}. \tag{4.54}$$

Excess heat is then used to formulate a melt rate, M_b :

$$M_b = \frac{k}{\rho L} \left(\frac{\partial T^*}{\partial z} - \frac{\partial T}{\partial z} \right), \tag{4.55}$$

where L is the latent heat of fusion. Finally, basal temperatures are held constant if the ice is floating:

$$\frac{\partial T(1)}{\partial t} = 0. (4.56)$$

4.2.6 Putting it all together

Equation (4.39) is solved for each ice column. The horizontal dependency of the horizontal advection term, (4.41b), is resolved by iterating the vertical solution. Combining the individual terms using a fully explicit finite-difference scheme, (4.40) becomes

$$\frac{T_{k,t+1} - T_{k,t}}{\Delta t} = \left(\frac{2aT_{k-1,t}}{(\sigma_k - \sigma_{k-1})(\sigma_{k+1} - \sigma_{k-1})} - \frac{2aT_{k,t}}{(\sigma_{k+1} - \sigma_k)(\sigma_k - \sigma_{k-1,t})} + \frac{2aT_{k+1,t}}{(\sigma_{k+1} - \sigma_k)(\sigma_{k+1} - \sigma_{k-1})}\right) + b_{1k,t} + b_{2k}T_{k,t} + \Phi_k + c_k \frac{T_{k+1,t} - T_{k-1,t}}{\sigma_{k+1} - \sigma_{k-1}}.$$
(4.57a)

Similarly, the fully implicit scheme is

$$\frac{T_{k,t+1} - T_{k,t}}{\Delta t} = \left(\frac{2aT_{k-1,t+1}}{(\sigma_k - \sigma_{k-1})(\sigma_{k+1} - \sigma_{k-1})} - \frac{2aT_{k,t+1}}{(\sigma_{k+1} - \sigma_k)(\sigma_k - \sigma_{k-1,t+1})} + \frac{2aT_{k+1,t+1}}{(\sigma_{k+1} - \sigma_k)(\sigma_{k+1} - \sigma_{k-1})}\right) + b_{1k,t+1} + b_{2k}T_{k,t+1} + \Phi_k + c_k \frac{T_{k+1,t+1} - T_{k-1,t+1}}{\sigma_{k+1} - \sigma_{k-1}}.$$
(4.57b)

Taking the average of (4.57a) and (4.57b) gives the Crank-Nicolson scheme. The resulting equation is then rearranged, with terms of $T_{k-1,t+1}$, $T_{k,t+1}$ and $T_{k+1,t+1}$ combined to give the tridiagonal system

$$\alpha_k T_{k-1,t+1} + \beta_k T_{k,t+1} + \gamma_k T_{k+1,t+1} = \delta_k, \tag{4.58}$$

where for k=2 to N-1 we have

$$\alpha_k = -\frac{1}{2} \frac{2a\Delta t}{(\sigma_k - \sigma_{k-1})(\sigma_{k+1} - \sigma_{k-1})} + \frac{1}{2} \frac{c_k \Delta t}{\sigma_{k+1} - \sigma_{k-1}},$$
(4.59a)

$$\beta_{k} = 1 + \frac{1}{2} \frac{2a\Delta t}{(\sigma_{k+1} - \sigma_{k})(\sigma_{k} - \sigma_{k-1})} - \frac{1}{2} b_{2k} \Delta t = 1 - \alpha_{k} - \gamma_{k} - \frac{1}{2} b_{2k} \Delta t,$$

$$\gamma_{k} = -\frac{1}{2} \frac{2a\Delta t}{(\sigma_{k+1} - \sigma_{k})(\sigma_{k+1} - \sigma_{k-1})} - \frac{1}{2} \frac{c_{k} \Delta t}{\sigma_{k+1} - \sigma_{k-1}},$$

$$(4.59b)$$

$$\gamma_k = -\frac{1}{2} \frac{2a\Delta t}{(\sigma_{k+1} - \sigma_k)(\sigma_{k+1} - \sigma_{k-1})} - \frac{1}{2} \frac{c_k \Delta t}{\sigma_{k+1} - \sigma_{k-1}},\tag{4.59c}$$

$$\delta_k = -\alpha_k T_{k-1,t} + (2 - \beta_k) T_{k,t} - \gamma_k T_{k+1,t} + \frac{1}{2} (b_{1k,t} + b_{1k,t+1}) \Delta t + \Phi_k \Delta t. \tag{4.59d}$$

Boundary conditions

At the upper boundary we have

$$\alpha_1 = 0, \quad \beta_1 = 1, \quad \gamma_1 = 0, \quad \delta_1 = T_{\text{surf}}.$$
 (4.59e)

The lower boundary condition is more complicated. Here we consider only the case when the temperature is below the pressure melting point of ice. The BCs for floating ice and temperatures at the pressure melting point of ice are trivial. The geothermal heat flux is applied at the lower boundary; i.e., (4.42b) becomes

$$\left. \frac{\partial T}{\partial \sigma} \right|_{\sigma_{k+1/2}} = -\frac{GH}{k}.\tag{4.60}$$

Assuming that $\sigma_k - \sigma_{k-1} = \sigma_{k+1} - \sigma_k = \Delta \sigma$ and inserting (4.42a) and (4.60) into (4.42c), the second partial derivative becomes

$$\left. \frac{\partial^2 T}{\partial \sigma^2} \right|_{\sigma_N} = \left(-\frac{GH}{k} - \frac{T_N - T_{N-1}}{\Delta \sigma} \right) / \Delta \sigma = -\frac{GH}{k\Delta \sigma} - \frac{T_N - T_{N-1}}{\Delta \sigma^2}. \tag{4.61}$$

Inserting the new conduction term and replacing the derivative of the vertical advection term with the Neumann boundary condition, (4.57a) becomes

$$\frac{T_{N,t+1} - T_{N,t}}{\Delta t} = -a \left(\frac{GH}{k\Delta \sigma} + \frac{T_{N,t} - T_{N-1,t}}{\Delta \sigma^2} \right) + b_{1N,t} + b_{2N} T_{N,t} + \Phi_N - c_N \frac{GH}{k}, \quad (4.62a)$$

and similarly for (4.57b):

$$\frac{T_{N,t+1} - T_{N,t}}{\Delta t} = -a \left(\frac{GH}{k\Delta\sigma} + \frac{T_{N,t+1} - T_{N-1,t+1}}{\Delta\sigma^2} \right) + b_{1N,t+1} + b_{2N}T_{N,t+1} + \Phi_N - c_N \frac{GH}{k}.$$
(4.62b)

The elements of the tridiagonal system at the lower boundary are then

$$\alpha_N = -\frac{a\Delta t}{2(\sigma_N - \sigma_{N-1})^2},\tag{4.63a}$$

$$\beta_N = 1 - \alpha_N + \frac{1}{2}b_{2N}\Delta t, \tag{4.63b}$$

$$\gamma_N = 0, \tag{4.63c}$$

$$\delta_{N} = -\alpha_{N} T_{N-1,t} + (2 - \beta_{N}) T_{N,t} - a \frac{GH\Delta t}{k(\sigma_{N} - \sigma_{N-1})} + \frac{1}{2} (b_{1N,t} + b_{1N,t+1}) \Delta t + \Phi_{N} \Delta t - c_{N} \frac{GH\Delta t}{k}.$$
(4.63d)

4.3 Basal Boundary Condition

This section describes the formulation of the basal boundary condition. An interface for the upper boundary condition (atmospheric BC) is easily defined by the surface temperature and mass balance. Similarly, the basal boundary consists of mechanical and thermal boundary conditions. The complications arise because the thermal and mechanical boundary conditions depend on each other. The interface of the basal boundary can be described with the following fields (see also Fig. 4.8):

- 1. basal traction: This field specifies a parameter which is used to allow basal sliding.
- 2. **basal heat flux:** This is the heat flux entering the ice sheet from below.
- 3. basal water depth: The presence of basal melt water affects the basal ice temperature.

Also, the ice sheet model calculates a melt/freeze rate based on the temperature gradient and basal water depth. This is handled by Glide.

4.3.1 Mechanical boundary conditions

If the ice is not frozen to the bed, basal décollement may occur. This can be parameterized by a traction factor, t_b . Within the ice sheet model, t_b can be used to calculate basal sliding velocities u_b in the case of zero-order physics. That is,

$$\mathbf{u}_b = t_b \boldsymbol{\tau}_b, \tag{4.64}$$

where τ_b is the basal shear stress. Alternatively, t_b can be used as part of the stress-balance calculations when the model is used with higher order physics. In simple models t_b may be uniform or prescribed as a spatial variable. More complex models may wish to make t_b dependent on other variables, such as basal melt rate. Typically t_b will depend on the presence of basal water

The second mechanical boundary condition, basal melting/freeze–on M_b , is handled within the ice sheet model. The details are described in Section 4.3.2.

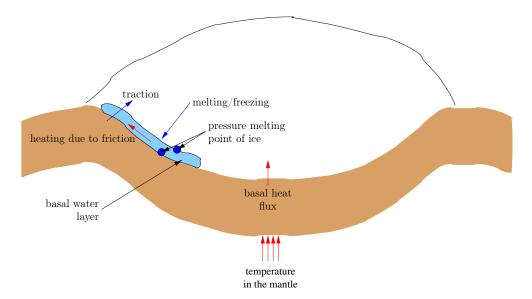


Figure 4.8: Basal boundary condition.

4.3.2 Thermal boundary conditions

The thermal boundary condition at the ice base is more complicated than the mechanical BC. The ice is heated from below by the geothermal heat flux. Heat is generated by friction with the bed. Furthermore, the ice temperature is constrained to be less than or equal to the pressure melting point of ice. The thermal boundary condition is a flux condition if there is no water present. If there is water, the basal temperature is set to the pressure melting temperature. (If it were lower, there would be no water, and if it were higher, there would be no ice.)

Basal melting and freezing

At the ice base, z = b, we can define outgoing and incoming heat fluxes, H_o and H_i :

$$H_o = -k_{\rm ice} \left. \frac{\partial T}{\partial z} \right|_{z=b^+} \tag{4.65a}$$

and

$$H_{i} = -k_{\text{rock}} \left. \frac{\partial T}{\partial z} \right|_{z=b^{-}} + \boldsymbol{u}_{b} \cdot \boldsymbol{\tau}_{b} + \begin{cases} \rho_{\text{ice}} M_{b} / L & \text{when } M_{b} < 0\\ 0 & \text{otherwise} \end{cases}$$
(4.65b)

where $k_{\rm ice}$ and $k_{\rm rock}$ are the thermal conductivities of ice and rock, $u_b \cdot \tau_b$ is the heat generated by friction with the bed, and L is the latent heat of fusion. The basal melt/freeze—on rate M_b can then be calculated from the difference between the incoming and outgoing heat fluxes:

$$M_b = \frac{H_i - H_o}{\rho_{ice}L} \tag{4.66}$$

There is freeze–on if $M_b < 0$, and basal melting if $B_b > 0$.

Geothermal heat flux

The heat flux across the basal boundary depends on past temperature variations since temperature perturbations penetrate the bedrock if the ice is frozen to the ground (Ritz, 1987). The

heat equation for the bedrock layer is given by the diffusion equation

$$\frac{\partial T}{\partial t} = \frac{k_{\text{rock}}}{\rho_{\text{rock}} c_{\text{rock}}} \nabla^2 T = \frac{k_{\text{rock}}}{\rho_{\text{rock}} c_{\text{rock}}} \left(\frac{\partial^2 T}{\partial x^2} + \frac{\partial^2 T}{\partial y^2} + \frac{\partial^2 T}{\partial z^2} \right), \tag{4.67}$$

where $k_{\rm rock}$ is the thermal conductivity, $\rho_{\rm rock}$ the density, and $c_{\rm rock}$ the specific heat capacity of the bedrock layer.

Initial conditions for the temperature field T are found by applying the geothermal heat flux G to an arbitrary surface temperature T_0 :

$$T(x, y, z) = T_0 + \frac{G}{k_{\text{rock}}} z.$$
 (4.68)

This ensures that initially the geothermal heat flux experienced by the ice sheet is equal to the regional heat flux. The basal boundary condition of the bedrock layer is kept constant, i.e.

$$T(x, y, H_{\text{rock}}) = T_0 + \frac{G}{k_{\text{rock}}} H_{\text{rock}}.$$
(4.69)

Lateral boundary conditions are given by

$$\left. \frac{\partial T}{\partial x} \right|_{x=0} = \left. \frac{\partial T}{\partial x} \right|_{x=L_x} = \left. \frac{\partial T}{\partial y} \right|_{y=0} = \left. \frac{\partial T}{\partial y} \right|_{y=L_y} = 0. \tag{4.70}$$

At the upper boundary, the heat flux of the rock layer has to be matched with the heat flux in the basal ice layer when the ice is frozen to the bed, i.e.

$$k_{\text{rock}} \left. \frac{\partial T}{\partial z} \right|_{z=-0} = k_{\text{ice}} \left. \frac{\partial T}{\partial z} \right|_{z=+0}.$$
 (4.71)

Otherwise the temperature of the top bedrock layer is set to the surface temperature (if the cell has been occupied by ice, but there is no ice present) or the basal ice temperature (if there is ice). Equation (4.71) is automatically fulfilled if we set the top bedrock temperature to the basal ice temperature everywhere and then calculate the geothermal heat flux to be used as boundary condition for (4.38).

4.3.3 Numerical solution

The horizontal grid of the bedrock layer is described in Section 4.1.1. The vertical bedrock grid is irregular like the vertical grid of the ice sheet. However, it is not scaled. Also for now, we ignore topography or isostatic adjustment, i.e. the bedrock layer is assumed to be flat and constant.

The horizontal second derivative in (4.67) becomes using finite-differences

$$\left. \frac{\partial^2 T}{\partial x^2} \right|_{x_i, y_i, z_i} = T_{xx, i, j, k} = \frac{T_{i+1, j, k} - 2T_{i, j, k} + T_{i-1, j, k}}{\Delta x}, \tag{4.72}$$

and similarly for $\partial^2 T/\partial y^2$. The vertical second derivative $\partial^2 T/\partial z^2$ is similar to (4.43):

$$\frac{\partial^2 T}{\partial z^2}\Big|_{x_i, y_i, z_i} = T_{zz, i, j, k} = \frac{2T_{i, j, k-1}}{(z_k - z_{k-1})(z_{k+1} - z_{k-1})} - \frac{2T_{i, j, k}}{(z_{k+1} - z_k)(z_k - z_{k-1})} + \frac{2T_{i, j, k+1}}{(z_{k+1} - z_k)(z_{k+1} - z_{k-1})}.$$
(4.73)

Using the Crank-Nicholson scheme, (4.67) becomes

$$\frac{T_{i,j,k}^{t+1} - T_{i,j,k}^t}{\Delta t} = D \left\{ \frac{T_{xx,i,j,k}^{t+1} + T_{xx,i,j,k}^t}{2} + \frac{T_{yy,i,j,k}^{t+1} + T_{yy,i,j,k}^t}{2} + \frac{T_{zz,i,j,k}^{t+1} + T_{zz,i,j,k}^t}{2} \right\}, \quad (4.74)$$

with $D = k_{\rm rock}/(\rho_{\rm rock}c_{\rm rock})$. Equation (4.74) is solved by gathering all T^{t+1} terms on the LHS and all other terms on the RHS. The index (i,j,k) is linearized using $\iota = i + (j-1)N + (k-1)NM$. The resulting matrix system is solved using the same bi–conjugate gradient solver as for the ice thickness evolution.

4.3.4 Basal hydrology

It is clear from the discussion above that the presence of basal water plays a crucial role in specifying both the mechanical and thermal boundary conditions. However, the treatment of basal water can vary greatly. Basal water is, therefore, left as an unspecified interface. Glide does provide a simple local water balance model which can be run in the absence of more complex models.

4.3.5 Putting it all together

The basal boundary consists of the individual components described in the previous sections. All components are tightly linked with each other. Figure 4.9 illustrates how the modules are linked and in what order they are resolved. The order of executions is then:

- 1. Find the basal heat flux by either solving the equation describing the thermal evolution of the lithosphere, (4.67), or by using the geothermal heat flux directly. The upper boundary condition of (4.67) is the same as the lower boundary condition of the thermal evolution of the ice sheet.
- 2. Either (1) the lower boundary condition for the thermal evolution of the ice sheet is given by the basal heat flux from *Step 1*; or (2) if melt water is present, the basal temperature is set to the pressure melting point of ice.
- 3. Calculate the temperature distribution within the ice sheet given the boundary condition found during Step 2 and the atmospheric BC.
- 4. Calculate a melt/freeze-on rate using Equation (4.66) given the outgoing heat flux calculated during Step 3, friction with the bed (calculated during the previous Step 7) and the incoming heat flux from Step 1. Freezing occurs only when there is basal water.
- 5. Track basal water. This is a user supplied module which can take any complexity. Inputs will typically be the melt/freeze—on rate determined during *Step 4*.
- 6. Calculate the basal traction parameter. Again, this is a user supplied module which typically will involve the presence of basal water (calculated during *Step 5*).
- 7. Solve the mechanical ice equations given basal traction parameter from Step 6.

Clearly, this scheme has the problem that heat is lost if the basal heat flux is such that more water could be frozen than is available. This might be avoided by iterating the process. On the other hand, the heat loss may be negligible if time steps are fairly small.

4.4 Isostatic Adjustment

The ice sheet model includes simple approximations for calculating isostatic adjustment. These approximations depend on how the lithosphere and the mantle are treated. For each subsystem there are two models. The lithosphere can be described as a

local lithosphere: the flexural rigidity of the lithosphere is ignored, i.e. this is equivalent to ice floating directly on the asthenosphere; or

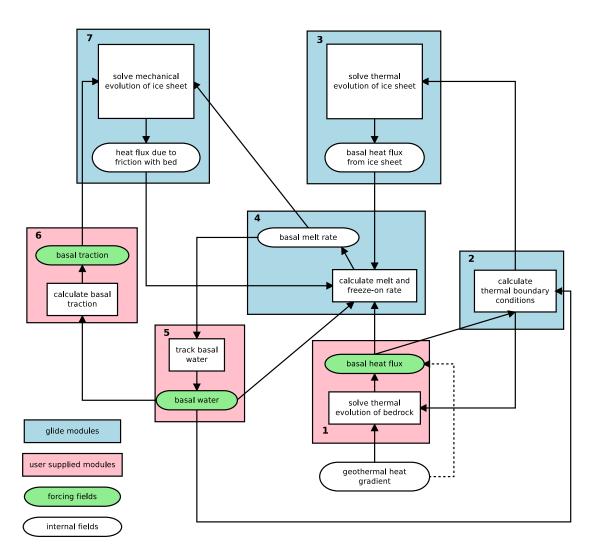


Figure 4.9: Flow diagram illustrating how the various modules communicate with each other by exchanging data fields.

elastic lithosphere: the flexural rigidity is taken into account;

while the mantle is treated as a

fluid mantle: the mantle behaves like a non-viscous fluid, with isostatic equilibrium reached instantaneously; or

relaxing mantle: the flow within the mantle is approximated by an exponentially decaying hydrostatic response function, i.e. the mantle is treated as a viscous half space.

4.4.1 Calculation of ice-water load

At each isostasy time step, the load of ice and water is calculated as an equivalent mantle-depth (L). If the basal elevation is above sea level, the load is simply due to the ice:

$$L = \frac{\rho_i}{\rho_m} H,\tag{4.75}$$

where H is the ice thickness, with ρ_i and ρ_m being the densities of the ice and mantle respectively. In the case where the bedrock is below sea level, the load is calculated is that due to a change in sea level and/or the presence of non-floating ice. When the ice is floating $(\rho_i H < \rho_o(z_0 - h))$, the load is due only to sea-level changes:

$$L = \frac{\rho_o}{\rho_m} z_0,\tag{4.76}$$

whereas when the ice is grounded, it displaces the water and adds an additional load:

$$L = \frac{\rho_i H + \rho_o b_r}{\rho_m}. (4.77)$$

Here, ρ_o is the density of sea water, z_0 is the change in sea-level relative to a reference level, and b_r is the bedrock elevation relative to the same reference level. The value of b_r will be negative for submerged bedrock, hence the plus sign in (4.77).

4.4.2 Elastic lithosphere model

The elastic model is selected by setting lithosphere = 1 in the configuration file. By simulating the deformation of the lithosphere, the deformation seen by the asthenosphere beneath is calculated. In the absence of this model, the deformation is that due to Archimedes' Principle, as though the load were floating on the asthenosphere.

The elastic lithosphere model is based on work by Lambeck and Nakiboglu (1980), and its implementation is fully described in Hagdorn (2003). The lithosphere model affects only the geometry of the deformation; the timescale for isostatic adjustment is controlled by the asthenosphere model.

The load due to a single (rectangular) grid point is approximated as being applied to a disc of the same area. The deformation due to a disc of ice of radius A and thickness H is given by the following expressions. For r < A:

$$w(r) = \frac{\rho_i H}{\rho_m} \left[1 + C_1 \operatorname{Ber} \left(\frac{r}{L_r} \right) + C_2 \operatorname{Bei} \left(\frac{r}{L_r} \right) \right], \tag{4.78}$$

and for $r \geq A$:

$$w(r) = \frac{\rho_i H}{\rho_m} \left[D_1 \operatorname{Ber} \left(\frac{r}{L_r} \right) + D_2 \operatorname{Bei} \left(\frac{r}{L_r} \right) + D_3 \operatorname{Ker} \left(\frac{r}{L_r} \right) + D_4 \operatorname{Kei} \left(\frac{r}{L_r} \right) \right], \tag{4.79}$$

where Ber(x), Bei(x), Ker(x) and Kei(x) are Kelvin functions of zero order, $L_r = (D/\rho_m g))^{1/4}$ is the radius of relative stiffness, and D is the flexural rigidity. The constants C_i and D_i are given by

$$C_1 = a \operatorname{Ker}'(a),$$

 $C_2 = -a \operatorname{Ker}'(a),$
 $D_1 = 0,$
 $D_2 = 0,$
 $D_3 = a \operatorname{Ber}'(a),$
 $D_4 = -a \operatorname{Ber}'(a).$

$$(4.80)$$

Here, the prime indicates the first spatial derivative of the Kelvin functions.

4.4.3 Relaxing asthenosphere model

If a fluid mantle is selected, it adjusts instantly to changes in lithospheric loading. However, a relaxing mantle is also available.

4.5 Time Step Ordering

Relative to Glimmer-CISM1.x, the order of operations on each time step has been somewhat reorganized in CISM2.x. This was done for consistency with the ordering of operations used in Glissade and to eliminate discrepancies between time levels applied in the model and the time levels written in output files. These changes have only a minor impact on model results, but do result in output that will not be an exact match between the two model versions for the same configuration and initial conditions.

While unlikely to be of much interest to the average user, particularly if you are not migrating from Glimmer-CISM1.x, the details of these changes follow. Specifically, in Glimmer-CISM1.x, for a given time step, H is advanced in time relative to T and v, and, because a complete solve occurs even at the initial time, the output for the initial time is different than the input for the prognostic variables H and T, even though no time step has occurred at that point.

In Glimmer-CISM1.x, the time-stepping loop is organized as follows:

```
glide_initialise(): init T, H, v
do while (t < tend)
   glide_tstep_p1(): solve T
   glide_tstep_p2(): solve v, H
   glide_tstep_p3(): calculate some diagnostic variables, write output for time t
   advance time
end do</pre>
```

This results in the following relation between the state of model variables and the time level at which they are output:

Output time:	0	1	2
T	1	2	3
H	1	2	3
v	0	1	2

In CISM2.x, the time-stepping loop is now organized as follows:

```
initialize modules
initial diagnostic solve of v (and any other diagnostic variables like upper surface)
write output for time 0
do while (t <= tend)
   advance time</pre>
```

```
solve column physics: T = f(H,T) solve prognostic variables: advection of H=f(H,v), T=f(H,v) solve diagnostic variables: v=f(H,T) write output for time t end do
```

This results in the following relation between the state of model variables and the time level at which they are output:

Output time:	0	1	2
\overline{T}	0	1	2
H	0	1	2
v	0	1	2

Chapter 5

Introduction to Higher-Order Ice Dynamics

5.1 Higher-Order Dynamics in CISM

5.1.1 Basics

The main distinction between so-called higher-order models and 0-order (or "shallow ice") models is that higher-order models attempt a closer approximation to solving the nonlinear Stokes equations. This usually means incorporating some approximation of horizontal-stress gradients (along-flow stretching or compression and across-flow shearing) in addition to the vertical stress gradients that are accounted for in shallow ice models (Figure 5.1). This is important for several reasons:

- For the ice-sheet regions of greatest interest e.g. ice streams, ice shelves, and other regions of fast flow horizontal-stress gradients are at least as important as vertical stress gradients. To model the flow accurately in these regions, higher-order models are required.
- The shallow-ice approximation, applied to situations in which there is basal sliding, gives rise to a singularity in the vertical velocity. Models compute the vertical velocity by integrating the incompressibility equation:

$$\frac{\partial w}{\partial z} = -\frac{\partial u}{\partial x} - \frac{\partial v}{\partial y}. (5.1)$$

If there is a jump from no sliding in one grid cell to sliding in an adjacent cell, the horizontal velocity gradients at the bed will depend on the grid spacing; the horizontal gradients (and through incompressibility, the vertical velocity gradient and thus the vertical velocity) will become increasingly large as the grid spacing decreases. Obviously, this should be avoided.

- Near the grounding line (the boundary between grounded and floating ice), horizontal rather than vertical stresses often control the flow. Incomplete knowledge of the stresses makes it unlikely that shallow ice models will ever be able to accurately simulate grounding-line advance and retreat.
- In some regions of very slow flow, like ice divides, horizontal-stress gradients are important or dominant. Ice cores are often recovered at ice divides, and flow modeling is important for interpreting ice core records and using information (such as layer thickness) to infer the past flow history in the region. In order to model that flow correctly, one must include horizontal stresses. (At an ice divide the surface slope is ~0, in which case vertical stress

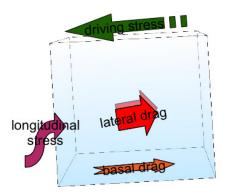


Figure 5.1: The gravitational stress available to move the ice is the driving stress, indicated in green. Because the ice is assumed to be in equilibrium, the sum of the other stresses is equal to (i.e., must balance) the gravitational driving stress. In the 0-order model (shallow-ice approximation) the driving stress is assumed to be balanced by basal drag alone. In higher-order models, this restriction is relaxed and the balance of stresses now includes lateral and/or longitudinal stresses. Because these stresses must be computed based on conditions outside the local ice column, the model becomes significantly more complex.

gradients that drive deformation in 0-order models are also \sim 0. In reality, deformation is not 0 at ice divides, but is controlled by horizontal stretching rather than vertical shearing).

The term "higher-order" comes from scaling analyses of the Stokes equations for which a scaling parameter $\lambda = H/L$ (the ratio of the thickness to the horizontal length scale of interest) is used to assign importance to the various terms. Shallow ice models retain only terms of order 0 while higher-order models also retain terms of order 1 (and possibly more) (Figure 5.2).

5.1.2 Available schemes

- The most fundamental higher-order scheme is the full set of nonlinear Stokes equations. Because of the computational burden, many large-scale 3D models solve lower-order approximations of the Stokes equations (Figure 5.3). However, several groups have made significant advances in Stokes modeling (e.g. the ELMER-Ice effort; Gagliardini et al. (2013)). Current DOE-funded efforts include implementation of a nonlinear Stokes model on unstructured grids (Leng et al., 2012).
- Probably the most long-lived higher-order approximation in glaciology is the shallow-shelf approximation (or SSA) describing flow within an ice shelf. While not truly "higher-order" when only applied to ice shelf flow, it can provide nearly equivalent results to "true" higher-order models in cases where basal drag is small but non-zero (e.g., for modeling the flow of ice streams). The SSA was developed and made popular by Douglas MacAyeal (e.g., MacAyeal (1989)). Its main disadvantage is that it is not fully 3D, as it assumes uniform velocity throughout the ice thickness driven only by horizontal stress gradients. It is adequate for describing fast flow in many parts of ice sheets, such as ice shelves and some ice streams. In these regions, not resolving vertical gradients is a computational advantage.
- The SSA equations are actually a depth-averaged form of a more general higher-order model, commonly referred to as the "Blatter-Pattyn model" (Blatter (1995); Pattyn

$$\boldsymbol{\tau} - P\boldsymbol{I} = 2\eta \begin{pmatrix} \frac{\partial u}{\partial x} & \frac{1}{2} \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) & \frac{1}{2} \left(\frac{\partial u}{\partial z} + \frac{\partial w}{\partial x} \right) \\ \frac{1}{2} \left(\frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} \right) & \frac{\partial v}{\partial y} & \frac{1}{2} \left(\frac{\partial v}{\partial z} + \frac{\partial w}{\partial y} \right) \\ \frac{1}{2} \left(\frac{\partial w}{\partial x} + \frac{\partial u}{\partial z} \right) & \frac{1}{2} \left(\frac{\partial w}{\partial y} + \frac{\partial v}{\partial z} \right) & \frac{\partial w}{\partial z} \end{pmatrix}$$

$$\boldsymbol{\tau} - P\boldsymbol{I} = 2\eta \begin{pmatrix} \frac{\partial u}{\partial x} & \frac{1}{2} \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) & \frac{1}{2} \left(\frac{\partial u}{\partial z} \right) \\ \frac{1}{2} \left(\frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} \right) & \frac{\partial v}{\partial y} & \frac{1}{2} \left(\frac{\partial v}{\partial z} \right) \\ \frac{1}{2} \left(\frac{\partial u}{\partial z} \right) & \frac{1}{2} \left(\frac{\partial v}{\partial z} \right) & -\frac{\partial u}{\partial x} - \frac{\partial v}{\partial y} \end{pmatrix}$$

Figure 5.2: Stokes-flow (top) and first-order (bottom) constitutive equations, which relate stresses to strain rates (and in turn to velocity gradients). τ is the deviatoric stress tensor, I is the identity tensor, P is the pressure, and η is the effective ice viscosity. The strain rate tensor is given by the term in parentheses on the right-hand side of the equation. For the first-order approximation (bottom), the bold terms from the Stokes equations are assumed negligible.

(2003)). Blatter-Pattyn dynamics, synonymous with and more formally described as the "first-order accurate Stokes approximation", are implemented in CISM's default higher-order dynamical core, Glissade.

- Another formally first-order accurate Stokes approximation, which might be considered "quasi" three-dimensional, is the so-called "L1L2" approximation (Schoof and Hindmarsh, 2010). This model combines a two-dimensional SSA solve and a one-dimensional (column) SIA solve in a mathematically rigorous way. The model solution, from a combination of SSA (for approximating membrane stresses) and SIA (for vertical shear stresses) solutions, mimics a fully 3D, higher-order model solution but at a fraction of the computational cost.
- Several "hybrid" schemes exist that are also computationally cheaper than the Blatter-Pattyn model. These are similar to the L1L2 model in spirit, but combine solutions to the SSA and SIA models in a heuristic fashion. Bueler and Brown (2009) and Pollard and Deconto (2009) describe large-scale models using distinct hybrid approaches.

Review papers by Schoof and Hewitt (2013) and Kirchner et al. (2011) go into a fair amount of detail about the various ice flow modeling approximations, their derivations, and their applicability.

5.1.3 Shallow-ice vs. higher-order models: practical differences

There are large differences between shallow-ice and higher-order models in (1) how the momentum equations (the dynamics, or stress balance) are solved and (2) how the information derived from that solution (the kinematics, or velocity fields) is used to evolve the ice sheet geometry in time. There are two main reasons for these differences:

¹See Schoof and Hindmarsh (2010) and Dukowicz et al. (2010) for a more complete scaling analysis and derivation of the first-order Stokes approximation.

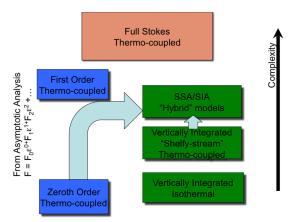


Figure 5.3: The relationship between several common varieties of ice sheet modesl. Complexity increases along the vertical axis.

- The numerical solution of the dynamical equations is fundamentally different. For the shallow-ice case, we need only local information (surface elevation and thickness) to solve for the velocity as a function of depth in a single column of ice. We do this pointwise for each location on the model domain (in map view), which is a relatively easy numerical problem. Each column of velocities leads to a tridiagonal matrix that is easy to invert (i.e., to solve for the velocities). This problem is embarrassingly parallel, since each column of unknown velocities results in a tridiagonal matrix that can be solved on a single processor. For higher-order models, however, the solution at any point also depends on the solution at neighboring points (in map view), leading to an elliptic system of equations in which every velocity must be solved for simultaneously with every other velocity. The result is a much larger system of equations, which are more difficult to solve on one processor and far more difficult to solve on multiple processors. Because large-scale higher-order model applications (e.g., whole-ice-sheet models) require efficient, robust solution and parallelization techniques, this is an active area of current research.
- The equations governing dynamics AND evolution in a shallow-ice model can be recast as a nonlinear diffusion equation for ice thickness. A single system of equations is solved to calculate the velocity field and evolve the ice sheet geometry. For higher-order models, we must first solve the momentum balance equations to obtain the velocity field, and then use some other scheme to evolve the ice thickness.

These differences imply that a higher-order model must be built in a fundamentally different way than a shallow-ice model. Most recent development work on CISM has focused on higher-order dynamics.

Using the constitutive law from the lower panel of Figure 5.2, the equations describing 3D higher-order flow² are:

$$x: \frac{\partial}{\partial x} \left(2\eta \left(2\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) \right) + \frac{\partial}{\partial y} \left(\eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right) + \frac{\partial}{\partial z} \left(\eta \frac{\partial u}{\partial z} \right) = \rho_i g \frac{\partial s}{\partial x},$$

$$y: \frac{\partial}{\partial y} \left(2\eta \left(2\frac{\partial v}{\partial y} + \frac{\partial u}{\partial x} \right) \right) + \frac{\partial}{\partial x} \left(\eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right) + \frac{\partial}{\partial z} \left(\eta \frac{\partial v}{\partial z} \right) = \rho_i g \frac{\partial s}{\partial y}.$$

$$(5.2)$$

²Specifically, the first-order accurate Stokes approximation discussed above. These will be derived in more detail in Section 5.2.

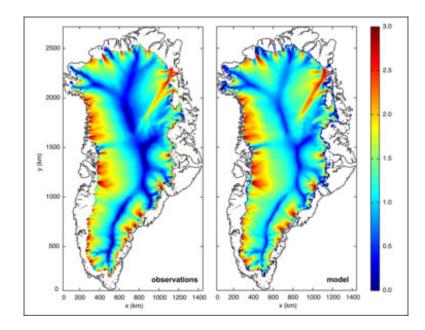


Figure 5.4: Surface ice speed (m/yr, log scale) for the Greenland ice sheet. (a) Observed speed (Joughin et al., 2010). (b) Simulated speed at the end of a 50 kyr spin-up without floating ice shelves using DIVA and a pseudo-plastic sliding law described in section 6.5. The model was inverted to match observed ice thickness.

In these equations we can point out several important differences between shallow-ice and higher-order models:

- 1. The third term on the left-hand side represents the vertical diffusion of horizontal velocities. In a shallow-ice model, this term alone accounts for all ice dynamics and balances the entire body force in the x direction (the term on the right-hand side). Since the velocity gradients are only in the vertical direction, the solution is local (in map view).
- 2. The first-order equations must be solved for each of a set of horizontal layers (the number of which is defined by the resolution in the vertical, here dz). These layers communicate with each other through the vertical diffusion term.
- 3. The first and second sets of terms on the left-hand side represent the resistance to the body-force term from "membrane stresses" (i.e., stresses arising from horizontal velocity gradients). These horizontal gradients make the problem elliptic and non-local. The absence of membrane stresses in the shallow-ice momentum balance is the primary reason for their failure to realistically simulate the flow in outlet glaciers, ice streams, and ice shelves.

Other complications arise when we account for boundary conditions at the lateral margins of the domain and at the upper and lower ice surfaces. We describe both the governing equations and the boundary conditions in more detail below.

5.2 Higher-Order Momentum Balance

CISM'S higher-order dynamics scheme (some output from which is shown in Figure 5.4) is discussed in more detail in the following sections. First we describe the derivation of the equations themselves, followed by a discussion and derivation of the boundary conditions. We

then describe generic solution methods for the nonlinear system of equations, followed by a brief discussion of the solution of the thickness evolution equation.

The higher-order scheme is based on the 3D first-order accurate Stokes approximation (often referred to as the "Blatter-Pattyn" model). The starting point is the nonlinear Stokes equations:

$$x: \frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} - \frac{\partial P}{\partial x} = 0,$$

$$y: \frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yy}}{\partial y} + \frac{\partial \tau_{yz}}{\partial z} - \frac{\partial P}{\partial y} = 0,$$

$$z: \frac{\partial \tau_{xz}}{\partial x} + \frac{\partial \tau_{zy}}{\partial y} + \frac{\partial \tau_{zz}}{\partial z} - \frac{\partial P}{\partial z} = \rho_i g,$$

$$(5.3)$$

where P is the pressure and τ is the deviatoric stress tensor. The latter is given by $\tau_{ij} = \sigma_{ij} + P\delta_{ij}$, where σ is the full stress tensor.

There are a number of ways to argue that, due to the shallowness of ice sheets (i.e., the small value of H/L, where H is the thickness and L is a relevant horizontal length scale), the Stokes equations can be reduced to the following first-order approximation:

$$x: \quad \frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} - \frac{\partial P}{\partial x} = 0,$$

$$y: \quad \frac{\partial \tau_{yy}}{\partial y} + \frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yz}}{\partial z} - \frac{\partial P}{\partial y} = 0,$$

$$z: \quad \frac{\partial \tau_{zz}}{\partial z} - \frac{\partial P}{\partial z} = \rho g.$$
(5.4)

The arguments supporting this reduction are fairly complex and are based on either a variational analysis or an asymptotic analysis (see Schoof and Hindmarsh (2010) and Dukowicz et al. (2010) for details).

The third (vertical) balance equation in (5.4) can be integrated vertically to give an expression for the pressure:

$$P = \rho g \left(s - z \right) + \tau_{zz}(z). \tag{5.5}$$

This is simply a statement that the full vertical normal stress is balanced by the hydrostatic pressure (the so-called hydrostatic assumption). This expression can be substituted into the horizontal pressure gradient terms above to remove pressure from the equations. For example, for the x component of velocity we have

$$\frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} = \frac{\partial}{\partial x} \left[\rho g \left(s - z \right) + \tau_{zz}(z) \right]
\frac{\partial \tau_{xx}}{\partial x} - \frac{\partial \tau_{zz}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} = \rho g \frac{\partial s}{\partial x}$$
(5.6)

Using the incompressibility constraint (5.1) and the assumption that stress and strain are aligned we can write

$$\tau_{zz} = -\tau_{xx} - \tau_{yy}. (5.7)$$

Taking the gradient of (5.7) with respect to x, we can rewrite the vertical normal deviatoric stress in terms of horizontal normal deviatoric stresses:

$$-\frac{\partial \tau_{zz}}{\partial x} = -\frac{\partial}{\partial x} \left(-\tau_{xx} - \tau_{yy} \right) = \frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{yy}}{\partial x}.$$
 (5.8)

Substituting (5.8) into (5.6), we obtain

$$\frac{\partial}{\partial x} \left(2\tau_{xx} + \tau_{yy} \right) + \frac{\partial \tau_{xy}}{\partial y} + \frac{\partial \tau_{xz}}{\partial z} = \rho g \frac{\partial s}{\partial x}. \tag{5.9}$$

Similarly, the y horizontal balance equation is

$$\frac{\partial}{\partial y} \left(2\tau_{yy} + \tau_{xx} \right) + \frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yz}}{\partial z} = \rho g \frac{\partial s}{\partial y}. \tag{5.10}$$

At this point we have removed the vertical balance equation entirely; it has been incorporated into the horizontal balance equations through incompressibility.

Next we write these equations in terms of the velocities for which we are ultimately solving. Stresses and velocities are linked through the constitutive law for ice, which relates strain rates to stresses (here we assume Nye's generalization of Glen's law), and through the definition of the strain rate tensor, which relates strain rates to velocity gradients:

1.
$$\tau_{ij} = B\dot{\varepsilon}_e^{\frac{1-n}{n}}\dot{\varepsilon}_{ij}, \quad B = B(T)$$
2.
$$\dot{\varepsilon}_{ij} = \frac{1}{2}\left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i}\right)$$
3.
$$2\dot{\varepsilon}_e = \dot{\varepsilon}_{ij}\dot{\varepsilon}_{ij}$$
4.
$$\eta \equiv \frac{1}{2}B\dot{\varepsilon}_e^{\frac{1-n}{n}}$$
5.
$$\tau_{ij} = 2\eta\dot{\varepsilon}_{ij}$$
(5.11)

In order, the five expressions above give:

- 1. Glen's flow law³, where $B = A^{\frac{-1}{n}}$ is the temperature dependent rate factor
- 2. The definition of the strain-rate tensor in terms of velocity gradients
- 3. The definition of the effective strain rate, $\dot{\varepsilon}_e$, a norm of the strain-rate tensor
- 4. A definition of the "effective viscosity" (after rearranging some terms in (1))
- 5. Items (1)-(4) allow us to write the relationship between stress and strain in a standard "Newtonian" way, but with a non-Newtonian (nonlinear) viscosity.

Taking these definitions into the stress balance equations (5.9) and (5.10) and expanding in terms of strain rates and effective viscosity, we have (for the x direction only):

$$x: \quad 2\frac{\partial}{\partial x}\left(2\eta\dot{\varepsilon}_{xx}\right) + \frac{\partial}{\partial x}\left(2\eta\dot{\varepsilon}_{yy}\right) + \frac{\partial}{\partial y}\left(2\eta\dot{\varepsilon}_{xy}\right) + \frac{\partial}{\partial z}\left(2\eta\dot{\varepsilon}_{xz}\right) = \rho g\frac{\partial s}{\partial x},\tag{5.12}$$

Replacing strain-rate components with velocity gradients, we obtain

$$x: \quad \frac{\partial}{\partial x} \left(4\eta \frac{\partial u}{\partial x} + 2\eta \frac{\partial v}{\partial y} \right) + \frac{\partial}{\partial y} \left[\eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right] + \frac{\partial}{\partial z} \left(\eta \frac{\partial u}{\partial z} \right) = \rho g \frac{\partial s}{\partial x}. \tag{5.13}$$

An analogous expression gives the y-direction momentum balance:

$$y: \quad \frac{\partial}{\partial y} \left(4\eta \frac{\partial v}{\partial y} + 2\eta \frac{\partial u}{\partial x} \right) + \frac{\partial}{\partial x} \left[\eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right] + \frac{\partial}{\partial z} \left(\eta \frac{\partial v}{\partial z} \right) = \rho g \frac{\partial s}{\partial y}. \tag{5.14}$$

These are the basic equations to be discretized and solved.

 $^{^3}$ Technically, we are using the *inverse* form of the flow-law here.

5.3 Higher-Order Model Boundary Conditions

We now carry out an approximate derivation of the boundary conditions in a higher-order scheme. By "approximate" we mean that some of the derivation is guided by physical intuition and reasonable arguments, rather than rigorous mathematics. In the end, we derive the same set of equations as when following the more rigorous approach (e.g., Dukowicz et al. (2010)). We will consider the boundary conditions in three parts: (1) the free surface BC, (2) the basal traction BC, and (3) lateral BCs.

5.3.1 Stress-free surface

At the upper ice surface, a stress-free boundary condition is applied. The traction vector T_i must be continuous across the ice sheet surface and, assuming that atmospheric pressure and surface tension are small, we have

$$T_i = -T_{i(\text{boundary})} \approx 0,$$

$$T_i = \sigma_{ii} n_i = \sigma_{i1} n_1 + \sigma_{i2} n_2 + \sigma_{i3} n_3 = 0,$$
(5.15)

where the n_i are the components of the outward-facing, unit normal vector in Cartesian coordinates.

For a function F(x,y,z) = z - f(x,y) = 0, where z = f(x,y) defines the surface, the gradient of F(x,y,z) gives the components of the surface normal vector. For the case of the ice sheet surface, s = f(x,y) and the surface normal is given by

$$n_i = \left(-\frac{\partial s}{\partial x}, -\frac{\partial s}{\partial y}, 1\right) \frac{1}{a},\tag{5.16}$$

where

$$a = \sqrt{\left(\frac{\partial s}{\partial x}\right)^2 + \left(\frac{\partial s}{\partial y}\right)^2 + 1}. (5.17)$$

Eq. (5.15) gives three equations that must be satisfied at the free surface:

$$\begin{split} i &= x: \quad T_x = \sigma_{xx} n_x + \sigma_{xy} n_y + \sigma_{xz} n_z = 0, \\ i &= y: \quad T_y = \sigma_{yx} n_x + \sigma_{yy} n_y + \sigma_{yz} n_z = 0, \\ i &= z: \quad T_z = \sigma_{zx} n_x + \sigma_{zy} n_y + \sigma_{zz} n_z = 0. \end{split} \tag{5.18}$$

Expanding the z equation and expressing stresses in terms of strain rates and pressures, where η is the effective viscosity defined above, gives

$$(2\eta \dot{\varepsilon}_{zx}) n_x + (2\eta \dot{\varepsilon}_{zy}) n_y + (2\eta \dot{\varepsilon}_{zz} - P) n_z = 0.$$
 (5.19)

Solving for the pressure gives

$$Pn_z = (2\eta \dot{\varepsilon}_{zz}) \, n_z + (2\eta \dot{\varepsilon}_{zx}) \, n_x + (2\eta \dot{\varepsilon}_{zy}) \, n_y. \tag{5.20}$$

Expanding in terms of velocity gradients and normal vector components, we find

$$P = 2\eta \frac{\partial w}{\partial z} - \left(\eta \frac{\partial u}{\partial z}\right) \frac{\partial s}{\partial x} - \left(\eta \frac{\partial v}{\partial z}\right) \frac{\partial s}{\partial y}, \tag{5.21}$$

where we have made the usual first-order approximation that

$$\frac{\partial w}{\partial x} = \frac{\partial w}{\partial y} \approx 0. \tag{5.22}$$

and also assumed $a \approx 1$.

We use this expression for the pressure and expand the two horizontal boundary condition expressions

$$i = x: \quad T_x = \sigma_{xx} n_x + \sigma_{xy} n_y + \sigma_{xz} n_z = 0,$$

$$i = y: \quad T_y = \sigma_{yx} n_x + \sigma_{yy} n_y + \sigma_{yz} n_z = 0,$$
(5.23)

in terms of velocity gradients and the effective viscosity to obtain

$$\hat{x}: -2\eta \frac{\partial u}{\partial x} \frac{\partial s}{\partial x} - \eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \frac{\partial s}{\partial y} + \eta \frac{\partial u}{\partial z} = -2\eta \frac{\partial w}{\partial z} \frac{\partial s}{\partial x} + \eta \frac{\partial u}{\partial z} \left[\frac{\partial s}{\partial x} \frac{\partial s}{\partial x} \right] + \eta \frac{\partial v}{\partial z} \left[\frac{\partial s}{\partial y} \frac{\partial s}{\partial x} \right],$$

$$\hat{y}: -2\eta \frac{\partial v}{\partial y} \frac{\partial s}{\partial y} - \eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \frac{\partial s}{\partial x} + \eta \frac{\partial v}{\partial z} = -2\eta \frac{\partial w}{\partial z} \frac{\partial s}{\partial y} + \eta \frac{\partial u}{\partial z} \left[\frac{\partial s}{\partial x} \frac{\partial s}{\partial y} \right] + \eta \frac{\partial v}{\partial z} \left[\frac{\partial s}{\partial y} \frac{\partial s}{\partial y} \right].$$

$$(5.24)$$

In both these expressions, the terms in square brackets are ~ 0 (because slopes on ice sheets are small and the slope squared is exceedingly small). From continuity, we also have

$$\frac{\partial w}{\partial z} = -\frac{\partial u}{\partial x} - \frac{\partial v}{\partial y}. ag{5.25}$$

Using this expression for the normal vertical velocity gradient and removing the terms in square brackets, our two horizontal BC expressions become

$$\hat{x}: \quad -2\eta \frac{\partial u}{\partial x} \frac{\partial s}{\partial x} - \eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \frac{\partial s}{\partial y} + \eta \frac{\partial u}{\partial z} = 2\eta \left(\frac{\partial u}{\partial x} \frac{\partial s}{\partial x} + \frac{\partial v}{\partial y} \frac{\partial s}{\partial x} \right),$$

$$\hat{y}: \quad -2\eta \frac{\partial v}{\partial y} \frac{\partial s}{\partial y} - \eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \frac{\partial s}{\partial x} + \eta \frac{\partial v}{\partial z} = 2\eta \left(\frac{\partial u}{\partial x} \frac{\partial s}{\partial y} + \frac{\partial v}{\partial y} \frac{\partial s}{\partial y} \right).$$

$$(5.26)$$

Moving terms to one side and dividing through by the effective viscosity η , we arrive at the final form of the first-order, free surface boundary conditions

$$\hat{x}: \quad 4\frac{\partial u}{\partial x}\frac{\partial s}{\partial x} + \frac{\partial u}{\partial y}\frac{\partial s}{\partial y} + 2\frac{\partial v}{\partial y}\frac{\partial s}{\partial x} + \frac{\partial v}{\partial x}\frac{\partial s}{\partial y} - \frac{\partial u}{\partial z} = 0,$$

$$\hat{y}: \quad 4\frac{\partial v}{\partial y}\frac{\partial s}{\partial y} + \frac{\partial v}{\partial x}\frac{\partial s}{\partial x} + 2\frac{\partial u}{\partial x}\frac{\partial s}{\partial y} + \frac{\partial u}{\partial y}\frac{\partial s}{\partial x} - \frac{\partial v}{\partial z} = 0.$$
(5.27)

5.3.2 Specified basal traction

Derivation of the basal boundary condition follows that above for the free surface except that

1. The right-hand side of the equation is not zero, but rather consists of an assumed basal traction vector with components

$$\tau_{bi} = (\tau_{bx}, \tau_{by}). \tag{5.28}$$

- 2. The outward-facing normal vector components now consist of horizontal gradients in the basal surface (with a resulting sign switch on the x, y, z components relative to the upper surface).
- 3. the effective viscosity does not disappear from the equations when we divide through.

Making these substitutions, we obtain the first-order basal boundary conditions for a specified basal traction:

$$\hat{x}: \quad 4\frac{\partial u}{\partial x}\frac{\partial b}{\partial x} + \frac{\partial u}{\partial y}\frac{\partial b}{\partial y} + 2\frac{\partial v}{\partial y}\frac{\partial b}{\partial x} + \frac{\partial v}{\partial x}\frac{\partial b}{\partial y} - \frac{\partial u}{\partial z} = \frac{\tau_{bx}}{\eta},$$

$$\hat{y}: \quad 4\frac{\partial v}{\partial y}\frac{\partial b}{\partial y} + \frac{\partial v}{\partial x}\frac{\partial b}{\partial x} + 2\frac{\partial u}{\partial x}\frac{\partial b}{\partial y} + \frac{\partial u}{\partial y}\frac{\partial b}{\partial x} - \frac{\partial v}{\partial z} = \frac{\tau_{by}}{\eta}.$$

$$(5.29)$$

We assume that basal traction is linearly related to basal sliding velocity through a positive traction parameter, β , such that

$$\tau_{bx} = -\beta u|_{z=b}, \quad \tau_{by} = -\beta v|_{z=b}.$$
(5.30)

The minus sign in front of the β indicates that the traction vector is oriented parallel to and in the opposite direction of the sliding velocity vector.

Substituting this expression into (5.29) gives the final horizontal boundary conditions at the ice sheet base:

$$\hat{x}: \quad 4\frac{\partial u}{\partial x}\frac{\partial b}{\partial x} + \frac{\partial u}{\partial y}\frac{\partial b}{\partial y} + 2\frac{\partial v}{\partial y}\frac{\partial b}{\partial x} + \frac{\partial v}{\partial x}\frac{\partial b}{\partial y} - \frac{\partial u}{\partial z} + \left(\frac{\beta}{\eta}\right)u = 0,$$

$$\hat{y}: \quad 4\frac{\partial v}{\partial y}\frac{\partial b}{\partial y} + \frac{\partial v}{\partial x}\frac{\partial b}{\partial x} + 2\frac{\partial u}{\partial x}\frac{\partial b}{\partial y} + \frac{\partial u}{\partial y}\frac{\partial b}{\partial x} - \frac{\partial v}{\partial z} + \left(\frac{\beta}{\eta}\right)v = 0.$$

$$(5.31)$$

The use of a linear traction parameter as in (5.30) can be modified to implement basal friction laws that are a function of effective pressure (ice overburden pressure minus water pressure at the bed) and generally are nonlinear in the sliding velocity. Two such basal friction laws are implemented in CISM. For friction laws that are nonlinear in sliding velocity, β in (5.30) becomes a function of sliding velocity. This is implemented by lagging the value of sliding velocity used to calculate an "effective" β in the fixed-point iterations used to solve the momentum balance.

The first is a Weertman-style friction law (Weertman, 1957, 1964) modified to include an effective pressure dependence (Bindschadler, 1983; Budd et al., 1979). Notation here follows the summary by Cuffey and Paterson (2010, p.240, eq. 7.17). The form of the friction law is:

$$\boldsymbol{u_b} = k\boldsymbol{\tau_b}^p N^{-q},\tag{5.32}$$

which can be rearranged for τ_b as:

$$\tau_b = k^{\frac{-1}{p}} u_b^{\frac{1}{p}} N^{\frac{q}{p}}, \tag{5.33}$$

where k is a friction coefficient based on thermal and mechanical properties of ice and is inversely related to bed roughness. Cuffey and Paterson (2010) suggest values of p=3 and q=1, making the basal friction nonlinear in velocity.

One problem with (5.32) is that it allows for unbounded basal traction. CISM also includes a physically-based basal friction law for sliding over hard beds that allows for cavitation and bounded basal drag (Schoof, 2005). In this law, the friction is related to the velocity and effective pressure by

$$\tau_{\mathbf{b}} = C \left(\frac{\mathbf{u_b}}{\mathbf{u_b} + N^n \Lambda} \right)^{1/n} N, \quad \Lambda = \frac{\lambda_{max} A}{m_{max}}.$$
(5.34)

In Eq. 5.34, C is a Coulomb friction coefficient, and λ_{max} and m_{max} are the wavelength (m) and maximum slope, respectively, of the dominant bedrock bumps. Near cavitation (i.e., when effective pressure approaches 0 at high water pressure), the friction law becomes a Coulomb friction law of the form $\tau_b = CN$. Alternatively, at large effective pressures (i.e., low water pressure), the friction law takes a power law form, $\tau_b \propto u_b^{1/n}$. The representation of bump geometry in the friction law is at a sub-grid scale and is not explicitly resolved by the model.

The release version of CISM does not currently include a hydrology model that computes effective pressure, so at present the effective pressure fields must be input as data to use these two basal friction laws.

5.3.3 Specified basal yield stress

The above implementation of a specified basal traction can be altered to simulate sliding over a sediment-covered bed for which the sediment has a plastic or "Coulomb friction" rheology. That is, sliding does not occur below the specified yield stress of the underlying material (e.g., a water-saturated subglacial till). When the yield stress, τ_0 is reached, sliding occurs at a rate that maintains but does not exceed that yield stress. Consider the x-direction boundary condition in (5.29), written out in terms of basal stress components:

$$\hat{x}: 4\eta \frac{\partial u}{\partial x} \frac{\partial b}{\partial x} + \eta \frac{\partial u}{\partial y} \frac{\partial b}{\partial y} + 2\eta \frac{\partial v}{\partial y} \frac{\partial b}{\partial x} + \eta \frac{\partial v}{\partial x} \frac{\partial b}{\partial y} - \eta \frac{\partial u}{\partial z} = \tau_{bx},$$

$$\tau_{bx} \approx \tau_0 = \tau_0 \left(\frac{u}{|\mathbf{u}|}\right) = \tau_0 \left(\frac{u}{\sqrt{u_0^2 + v_0^2 + \gamma}}\right).$$
(5.35)

In this expression, u is the x component of the basal sliding velocity from the current iteration, u_0 and v_0 are components from the previous iteration, and γ is a regularization constant (a small number to avoid division by zero during early iterations). As the solution converges, velocities at the current and previous iteration are nearly equal, and the expression in parentheses approaches 1, in which case the basal stress and yield stress are approximately equal.

We can accommodate this plastic behavior in our earlier framework by treating β as nonlinear and dependent on the velocity:

$$\tau_{bx} = -\beta u,
\beta \equiv \frac{\tau_0}{\sqrt{u_0^2 + v_0^2 + \gamma}},
\tau_{bx} = -\left(\frac{\tau_0}{\sqrt{u_0^2 + v_0^2 + \gamma}}\right) u.$$
(5.36)

This expression for nonlinear β can be substituted in the expression (5.31) for the basal boundary conditions. Figure 5.5 shows an example of this type of boundary condition for the case of an idealized ice stream simulated by CISM.

5.3.4 Lateral boundary conditions

Currently, the only significant lateral boundary condition implemented in CISM is that for floating ice; the depth-averaged stress resulting from an adjacent column of ocean water is applied at the location of an ice shelf (or ice tongue) front. The derivation largely follows those for the free surface and specified basal traction boundary conditions, except that the surface normal vectors, n_x and n_y , are taken as lying entirely in the x, y plane (i.e., they are perpendicular to a vertical shelf front). Thus we have

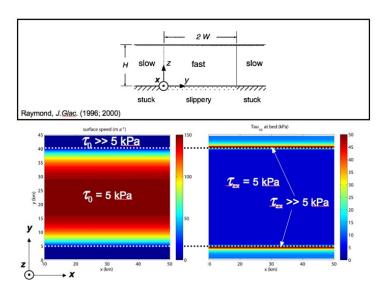


Figure 5.5: Idealized ice stream with plastic-bed sliding. Top panel shows a schematic of an idealized ice stream, frozen at the margins and thawed within the ice stream (flow is out of the page). Bottom panel (color) shows a modeled, idealized ice stream (flow is from left to right) with a yield stress of 5 kPa within the ice stream and much larger than 5 kPa outside of it. Bottom left shows the ice stream speed (m/yr) and the bottom right shows the basal drag (kPa). Within the ice stream, basal drag is equal to the yield stress. Outside the ice stream, stress transfer to the lateral margins results in basal drag that is much larger than the yield stress (and also much larger than the driving stress).

$$\hat{x}: \quad 4\eta \frac{\partial u}{\partial x} n_x + \eta \frac{\partial u}{\partial y} n_y + \eta \frac{\partial u}{\partial z} n_z = -2\eta \frac{\partial v}{\partial y} n_x - \eta \frac{\partial v}{\partial x} n_y + \rho g (s - z) n_x - S_x,$$

$$\hat{y}: \quad 4\eta \frac{\partial v}{\partial y} n_y + \eta \frac{\partial v}{\partial x} n_x + \eta \frac{\partial v}{\partial z} n_z = -2\eta \frac{\partial u}{\partial x} n_y - \eta \frac{\partial u}{\partial y} n_x + \rho g (s - z) n_y - S_y.$$

$$(5.37)$$

Here, S_x and S_y are source terms from the pressure due to ocean water (to be defined below), and $\rho g(s-z)$ comes from including the first-order vertical stress balance:

$$\frac{\partial \sigma_{zz}}{\partial z} = \frac{\partial \tau_{zz}}{\partial z} - \frac{\partial P}{\partial z} = \rho g \quad \text{(integrate w.r.t. } z),$$

$$P = \rho g (s - z) + \tau_{zz},$$

$$P = \rho g (s - z) + 2\eta \dot{\varepsilon}_{zz} = \rho g (s - z) + 2\eta \left(-\frac{\partial u}{\partial x} - \frac{\partial v}{\partial y} \right),$$
(5.38)

with ρ being the ice density. In the last step above we have used the constitutive relation and incompressibility to expand the vertical, normal-deviatoric stress in terms of the effective viscosity and horizontal-normal strain rates. We calculate the source terms S_x and S_y as the depth-averaged stress at the ice shelf front due to the pressure of ocean water there. This value is given by

$$S_{i} = \left[\frac{1}{H} \frac{1}{2} \rho_{w} g \left(H - h_{f}\right)^{2}\right] n_{i} = \left[\frac{1}{2} \rho_{w} g H \left(\frac{\rho}{\rho_{w}}\right)^{2}\right] n_{i}, \qquad h_{f} = H \left(1 - \frac{\rho}{\rho_{w}}\right), \qquad (5.39)$$

where $i=x,y,\ n_i$ is the shelf-front normal vector, H is the ice thickness, h_f is the "freeboard", or ice thickness above flotation, and ρ_w is the density of ocean water. Because we have taken a depth average for this source term, we take a depth average of the term $\rho g (s-z)$ above, which is $\frac{1}{2}\rho gH$.

Combining these two terms and inserting them in the horizontal boundary condition expressions above gives

$$\hat{x}: 4\eta \frac{\partial u}{\partial x} n_x + \eta \frac{\partial u}{\partial y} n_y + \eta \frac{\partial u}{\partial z} n_z =$$

$$-2\eta \frac{\partial v}{\partial y} n_x - \eta \frac{\partial v}{\partial x} n_y + \left[-\frac{1}{2} H \left(\frac{\rho}{\rho_w} \right)^2 \rho_w g + \frac{1}{2} \rho g H \right] n_x,$$

$$\hat{y}: 4\eta \frac{\partial v}{\partial y} n_y + \eta \frac{\partial v}{\partial x} n_x + \eta \frac{\partial v}{\partial z} n_z =$$

$$-2\eta \frac{\partial u}{\partial x} n_y - \eta \frac{\partial u}{\partial y} n_x + \left[-\frac{1}{2} H \left(\frac{\rho}{\rho_w} \right)^2 \rho_w g + \frac{1}{2} \rho g H \right] n_y,$$

$$(5.40)$$

which can be rearranged to

$$\hat{x}: \quad 4\frac{\partial u}{\partial x}n_x + \frac{\partial u}{\partial y}n_y + \frac{\partial u}{\partial z}n_z + 2\frac{\partial v}{\partial y}n_x + \frac{\partial v}{\partial x}n_y = \frac{\rho gH}{2\eta}\left(1 - \frac{\rho}{\rho_w}\right)n_x,$$

$$\hat{y}: \quad 4\frac{\partial v}{\partial y}n_y + \frac{\partial v}{\partial x}n_x + \frac{\partial v}{\partial z}n_z + 2\frac{\partial u}{\partial x}n_y + \frac{\partial u}{\partial y}n_x = \frac{\rho gH}{2\eta}\left(1 - \frac{\rho}{\rho_w}\right)n_y.$$
(5.41)

For an ice shelf, the surface and basal velocities are equal, in which case the vertical velocity gradient terms are ~ 0 , giving the final form of the lateral boundary conditions:

$$\hat{x}: \quad 4\frac{\partial u}{\partial x}n_x + \frac{\partial u}{\partial y}n_y + 2\frac{\partial v}{\partial y}n_x + \frac{\partial v}{\partial x}n_y = \frac{\rho gH}{2\eta} \left(1 - \frac{\rho}{\rho_w}\right)n_x,$$

$$\hat{y}: \quad 4\frac{\partial v}{\partial y}n_y + \frac{\partial v}{\partial x}n_x + 2\frac{\partial u}{\partial x}n_y + \frac{\partial u}{\partial y}n_x = \frac{\rho gH}{2\eta} \left(1 - \frac{\rho}{\rho_w}\right)n_y.$$
(5.42)

5.3.5 Summary

All the boundary conditions above have a similar form. In fact, all that differs among the equations for the free surface, the specified basal traction, and the lateral shelf boundary condition is (1) the definition of the normal vectors and (2) the existence and definition of a source term. Here are the x-direction equations again for the three cases:

$$\hat{x}: \quad 4\frac{\partial u}{\partial x}\frac{\partial s}{\partial x} + \frac{\partial u}{\partial y}\frac{\partial s}{\partial y} + 2\frac{\partial v}{\partial y}\frac{\partial s}{\partial x} + \frac{\partial v}{\partial x}\frac{\partial s}{\partial y} - \frac{\partial u}{\partial z} = 0,
\hat{x}: \quad 4\frac{\partial u}{\partial x}\frac{\partial b}{\partial x} + \frac{\partial u}{\partial y}\frac{\partial b}{\partial y} + 2\frac{\partial v}{\partial y}\frac{\partial b}{\partial x} + \frac{\partial v}{\partial x}\frac{\partial b}{\partial y} - \frac{\partial u}{\partial z} = \frac{\tau_{bx}}{\eta},
\hat{x}: \quad 4\frac{\partial u}{\partial x}n_x + \frac{\partial u}{\partial y}n_y + 2\frac{\partial v}{\partial y}n_x + \frac{\partial v}{\partial x}n_y = \frac{\rho gH}{2\eta}\left(1 - \frac{\rho}{\rho_w}\right)n_x.$$
(5.43)

In the first equation (free surface), the normals are related to the ice sheet surface slope and the source term is zero (which allows us to divide through by the effective viscosity and remove it from the equations). In the second equation (specified basal traction), the normals are related to the bedrock slopes and the source term is related to the assumed relationship between the basal shear stress and the basal sliding rate. In the last equation, the normals are defined by the shape of the ice-shelf front in map view, the vertical velocity gradient terms are absent, and the source term is related to the pressure from the adjacent column of ocean water.

In CISM's earlier developmental dycore, Glam, these boundary conditions were implemented using finite differences, leading to significant complexity. The current dycore, Glissade, solves for ice velocity using a finite-element approach, in which boundary conditions are handled more simply. For example, the surface boundary condition is a "natural" BC that is satisfied automatically. See, e.g., Dukowicz et al. (2010) for details.

5.4 Numerical Solution of Higher-Order Equations

This section describes numerical methods for solving the higher-order equations discussed above. More specific details are given in Chapter 6.

5.4.1 Governing matrix equations

We would like to solve the stress-balance equations (5.13) and (5.14) for the horizontal velocity components u and v. Any number of standard methods (e.g., finite differences, finite volumes, or finite elements) can be used to discretize these governing equations: i.e., to re-formulate them from their continuous, PDE form to a discontinuous, piecewise, algebraic approximation based on a specific computational mesh. In the limit, as the mesh spacing goes to zero, the difference between the solution to the actual PDE and its algebraic approximation also goes to zero. The Glissade dynamical core uses the Finite Element Method (FEM) to discretize the governing equations, as discussed in Chapter 6. The result of the discretization is a set of linear algebraic equations that can be written in matrix form. An example is given below, where the coefficients of u and v are contained in block matrices \mathbf{A} , given by

$$\begin{bmatrix} \mathbf{A_{uu}} & \mathbf{A_{uv}} \\ \mathbf{A_{vu}} & \mathbf{A_{vv}} \end{bmatrix} \begin{bmatrix} \mathbf{u} \\ \mathbf{v} \end{bmatrix} = \begin{bmatrix} \mathbf{b_{u}} \\ \mathbf{b_{v}} \end{bmatrix},$$

$$\mathbf{A_{uu}u} + \mathbf{A_{uv}v} = \mathbf{b_{u}}, \qquad \mathbf{A_{vu}u} + \mathbf{A_{vv}v} = \mathbf{b_{v}}.$$

$$(5.44)$$

Here the **uu** subscript denotes block matrices containing coefficients for gradients of u in the equation for the x component of velocity (i.e., u). The subscript **uv** denotes block matrices containing coefficients for gradients of v in the equation for the x component of velocity (and similarly for the **vu** and **vv** subscripts). On the right-hand side, the subscripts **u** and **v** are attached to the geometric source terms for the x and y components of velocity, respectively.

5.4.2 Treating nonlinearity through a fixed-point iteration

The nonlinearity of the equations—the fact that the matrix coefficients (in particular, the effective viscosity) depend on the velocity (or more specifically, the velocity gradients)—is handled through a fixed-point iteration. A general fixed-point iteration for a vector of unknowns u can be written as

$$u^k = \mathbf{B}\left(u^{k-1}\right),\tag{5.45}$$

where k is the current iteration index and \mathbf{B} is a matrix operation performed on the components of u obtained at the previous iteration, k-1. The fixed point occurs when the values of u at k and k-1 are equal to within some given tolerance, at which point the iteration process is halted. For a Picard iteration, which is used in Glissade, the matrix coefficients with a velocity dependence are simply based on the velocities from the previous iteration. In other words, velocities obtained during iteration k-1 are used to calculate the strain rates appearing in the expression for effective viscosity at iteration k.

Accounting for the Picard iteration on the effective viscosity, the final form of the matrix equations solved by Glissade becomes

$$\begin{bmatrix} \mathbf{A}_{\mathbf{u}\mathbf{u}}^{k-1} & \mathbf{A}_{\mathbf{u}\mathbf{v}}^{k-1} \\ \mathbf{A}_{\mathbf{v}\mathbf{u}}^{k-1} & \mathbf{A}_{\mathbf{v}\mathbf{v}}^{k-1} \end{bmatrix} \begin{bmatrix} \mathbf{u}^k \\ \mathbf{v}^k \end{bmatrix} = \begin{bmatrix} \mathbf{b}_{\mathbf{u}}^k \\ \mathbf{b}_{\mathbf{v}}^k \end{bmatrix},$$

$$\mathbf{A}_{\mathbf{u}\mathbf{u}}^{k-1}\mathbf{u}^k + \mathbf{A}_{\mathbf{u}\mathbf{v}}^{k-1}\mathbf{v}^k = \mathbf{b}_{\mathbf{u}}^k, \qquad \mathbf{A}_{\mathbf{v}\mathbf{u}}^{k-1}\mathbf{u}^k + \mathbf{A}_{\mathbf{v}\mathbf{v}}^{k-1}\mathbf{v}^k = \mathbf{b}_{\mathbf{v}}^k,$$

These equations form a linear system; for the solution at any particular iteration k, the matrix coefficients that depend on the velocity components u and v are held frozen during the solution. This linear system can be solved using a variety of methods. For large, sparse systems, the preconditioned conjugate gradient (PCG) method or some related method (e.g., preconditioned BiCG or GMRES) is generally the most efficient. (Glissade forms a symmetric positive-definite matrix, as required for the PCG method. $\mathbf{A}_{\mathbf{u}\mathbf{v}}$ and $\mathbf{A}_{\mathbf{v}\mathbf{v}}$ are symmetric, and $\mathbf{A}_{\mathbf{u}\mathbf{v}} = \mathbf{A}_{\mathbf{v}\mathbf{u}}^{\mathbf{T}}$.) In this case the linear system is solved not exactly, but to within some small tolerance of the true solution.

5.5 Thickness Evolution in Higher-Order Models

5.5.1 Conservation of mass

As mentioned previously, conservation of mass for an ice sheet can be expressed by

$$\frac{\partial H}{\partial t} = -\nabla \cdot (UH) + B,\tag{5.46}$$

where U = (U, V) is the 2D, depth-integrated velocity vector, H is the ice thickness, UH is the 2D ice flux vector, and B is the sum of surface and basal mass balance terms. The change in thickness per unit time is given by the negative of the flux divergence plus a source term. The minus sign in front of the divergence (terms in parentheses on the right-hand side) ensures sensible behavior: consider a section of the ice sheet where the thickness is nearly uniform and there is no accumulation or ablation. If the velocity gradient along the flow is negative (the ice is slowing down), then we expect local thickness (left-hand side of the equation is > 0), and if the ice is speeding up along flow, we expect local thinning. This is the equation that needs to be solved to calculate the evolution of the ice sheet geometry.

5.5.2 A diffusive approach

For the case of a shallow-ice model, U and V are expressed in terms of ice thickness and elevation gradients, in which case the problem can be recast as a diffusion equation in ice thickness. In 1D, (5.46) becomes

$$\frac{\partial H}{\partial t} = \frac{\partial}{\partial x} \left(D \frac{\partial s}{\partial x} \right) + B, \quad D = \frac{2A}{n+2} \left(\rho g \right)^n H^{n+2} \left| \frac{\partial s}{\partial x} \right|^{n-1}, \tag{5.47}$$

where the diffusivity D is nonlinear (because it depends on the solution H), A is the rate factor in Glen's flow law, n is the power-law exponent, s is the ice surface elevation, and ρ and g are the ice density and the acceleration due to gravity. Importantly, we need only local information in order to solve the above equation. If the velocity cannot be solved locally, as in the case of higher-order models, we cannot easily use the above formulation to solve ice sheet evolution. In an attempt to use this form and retain a diffusion-solution approach to the problem (diffusion problems generally have favorable numerical properties), we could try the following approach (again, in 1D):

$$\frac{\partial H}{\partial t} = \frac{\partial}{\partial x} \left(D \frac{\partial s}{\partial x} \right) + B, \quad D = UH \left(\frac{\partial s}{\partial x} \right)^{-1}, \tag{5.48}$$

where the U in the expression for D is the depth-integrated velocity field from the higher-order model. This is the approach initially taken by Pattyn (2003) in one of the first higher-order modeling studies to deal with this particular problem. Notice that ice sheet surface slope is in the denominator of the diffusivity. As slopes become smaller (as they tend to do in fast-flowing ice streams and ice shelves), the diffusivity grows larger, approaching infinity as the slope goes to zero. The faster velocities in these regions appear in the numerator of the diffusivity, also making it larger. This is a severe restriction because, for explicit schemes, the diffusive CFL condition requires that

$$\Delta t < \frac{\left(\Delta x\right)^2}{2D},\tag{5.49}$$

where Δt is the time step required for numerical stability and Δx is the grid spacing. As the diffusivity goes to infinity (i.e., for faster flows and shallower slopes), the stable time step goes to zero. Thus in practice, this approach has proven very difficult to use for calculating ice sheet evolution in the dynamically interesting areas of ice sheets. An alternate approach is needed.

5.5.3 Advection schemes

An alternate approach is to solve the evolution equation using an advection scheme. Numerically, advection schemes can be more problematic than diffusion schemes, but in cases like this one, they are difficult to avoid. One of the simplest advection schemes is a first-order, upwind-advection scheme (as above, "first-order" here refers to first-order accurate), and we will outline the implementation of such a scheme below.

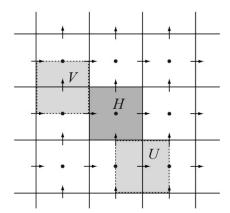


Figure 5.6: Staggered grid in two dimensions, showing scalars (like thickness, H) at grid cell centers and velocity components, U and V, at grid cell edges. This is a "C" grid. Another staggered-grid possibility is a "B" grid, in which both velocity components live at the grid cell corners. While CISM uses a B grid, averaging of corner velocities to edges allows one to express them on a C grid, if necessary.

Most ice sheet models (and fluid dynamics models in general) perform calculations on a "staggered" grid of the type shown below in Figure 5.6, where scalar components (e.g., temperature, pressure, and thickness) live on one grid, and velocity components live on a grid that is staggered by 1/2 grid cell in the horizontal direction. (This is essential for numerical stability for reasons not discussed here).

A control volume (or finite volume) approach to solving the equation

$$\frac{\partial H}{\partial t} = -\left(\frac{\partial (UH)}{\partial x} + \frac{\partial (VH)}{\partial y}\right) + B \tag{5.50}$$

would be to integrate the equation over a control volume centered on the scalar values (Figures 5.6 and 5.7). Ignoring source terms for now (assume B=0), and assuming flow in the x direction only (assume V=0) we have

$$\frac{\partial H}{\partial t} = -\frac{1}{\Delta y \Delta x} \int_{s}^{n} \int_{w}^{e} \frac{\partial (UH)}{\partial x} dx dy = -\frac{1}{\Delta y \Delta x} (HU_{e} - HU_{w}) \Delta y.$$
 (5.51)

The "east" and "west" (subscripts e and w) faces of the control volume are shown in Figure 5.7.

In the above equation, we have deliberately left it vague as to which value of H is being advected across the east and west interfaces, into or out of the volume. This is where the term "upwinding" comes in – we choose the scalar value to advect across the interface based on an upwinding criterion. If, for example, the flow at interface e is from left to right (U>0), we would advect the value of H at P out of the volume centered at P and into the volume centered at E. If, on the other hand, the flow at e was from right to left (U<0), we would advect the value of E and E into the volume at E and out of the volume at E.

The product of velocity, thickness, and the grid spacing at each interface gives a volume flux in units of cubic meters per year. The sum of the volume fluxes over the total number of faces being considered (two in this case, but four for the 2-dimensional case) gives the total volume flux in (total>0) or out (total<0) of the volume. When this number is divided by the area of the volume, the result is the mean thickness change in that volume, per unit time, that is required to maintain conservation of mass.

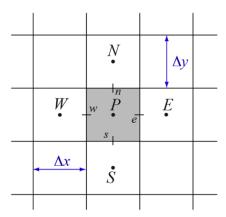


Figure 5.7: Staggered grid in two dimensions, showing locations of interfaces and control volume dimensions. Interfaces e, w, n and s connect the volume centered at P with those volumes to the east, west, north, and south (E, W, N, A).

5.5.4 Explicit vs. implicit evolution schemes

Is this an explicit or implicit scheme? If we discretize the right-hand side of the primary equation in time we get

$$\frac{\partial H}{\partial t} = -\nabla \cdot (UH) + B,
\frac{H_{t=1} - H_{t=0}}{\Delta t} \approx -\nabla \cdot (UH) + B.$$
(5.52)

This can be rearranged to

$$H_{t_{t=1}} \approx H_{t=0} + [-\nabla \cdot (UH) + B]_{t=0} \Delta t.$$
 (5.53)

The thickness at the new time step t = 1 is a function only of variables evaluated at the previous time step t = 0. Thus, the scheme is *explicit* and, as such, is subject to the advective CFL condition for stability:

$$\Delta t < \frac{\Delta x}{U}.\tag{5.54}$$

Essentially, this means that we cannot take such a large time step that we advect mass through more than one grid cell per time step.

There is not much more to a first-order advection scheme other than extending it to two dimensions for non-zero V. The finite volume formulation guarantees that it will conserve mass (i.e., the mass moving out of one grid cell across a given cell face is equal to the mass moving into an adjacent grid cell across that face). There are many "higher-order" advection schemes (e.g., the incremental remapping scheme used in CISM and described in Chapter 6), but they are mainly based on the principles outlined here and rely on corrections to the simple upwind assumption in order to gain more accuracy.

Chapter 6

Higher-Order Ice Dynamics: Glissade Dynamical Core

6.1 Introduction

CISM includes a parallel, higher-order dynamical core called Glissade¹, the successor to the Glide shallow-ice dycore. Like Glide, Glissade solves equations for conservation of momentum (i.e., an appropriate approximation of Stokes flow), mass, and internal energy. Glissade numerics, however, differ substantially from Glide numerics:

- Velocity: Glide is limited to the shallow-ice approximation, whereas Glissade can solve several Stokes approximations, as described in Section 6.2. These include a 3D first-order Blatter-Pattyn solver (the most accurate and complex scheme) as well as simpler solvers for the shallow-ice approximation (SIA), shallow-shelf approximation (SSA), and a vertically integrated higher-order approximation. Unlike Glide, which computes the velocity using finite-difference techniques, Glissade uses a more robust and flexible finite-element method.
- Temperature: To evolve the ice temperature, Glide solves a prognostic equation that incorporates horizontal advection as well as vertical heat diffusion and internal dissipation. In Glissade, temperature advection is handled by the transport scheme, and a separate module solves for vertical diffusion and internal dissipation. The vertical solver, described in Section 6.3, is local; each ice column calculation is independent of other column calculations.
- Mass and tracer transport: Glide solves a diffusion equation for transport of mass (i.e., ice thickness); this equation incorporates the local shallow-ice velocities. Since Glissade solves for higher-order flow that may have a large advective (as oppposed to diffusive) component, a different approach is needed. Glissade has two mass- and energy-conserving transport options, described in Section 6.4: incremental remapping (the more complex and accurate scheme) and first-order upwind. These schemes transport mass and internal energy in the horizontal direction, followed by a vertical remapping to sigma coordinates.

Glissade numerics are described in detail below.

¹The name "Glissade" was originally an acronym, but as with "Glimmer" and "Glide" the acronym is rarely used. It can be pronounced either American–style (gliss–AID) or French–style (gliss–AHD).

6.2 Velocity Solver

Glissade computes the ice velocity by solving an appropriate approximation of the Stokes equations, given the 2D surface elevation and thickness fields, the 3D temperature field, and relevant boundary conditions. The directory libglissade contains the primary higher-order velocity module, glissade_velo_higher.F90, along with supporting modules for the various solver options. Section 6.2.1 describes assembly and solution of the matrix problem for the 3D first-order Blatter-Pattyn approximation. The following sections discuss simpler approximations, including the SIA, SSA, and a depth-integrated-viscosity approximation (DIVA, based on Goldberg (2011)).

6.2.1 Blatter-Pattyn approximation

The basic equations of the Blatter-Pattyn approximation in Cartesian coordinates, repeated from Section 5.2, are

$$x: \frac{\partial}{\partial x} \left(2\eta \left(2\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) \right) + \frac{\partial}{\partial y} \left(\eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right) + \frac{\partial}{\partial z} \left(\eta \frac{\partial u}{\partial z} \right) = \rho g \frac{\partial s}{\partial x},$$

$$y: \frac{\partial}{\partial y} \left(2\eta \left(2\frac{\partial v}{\partial y} + \frac{\partial u}{\partial x} \right) \right) + \frac{\partial}{\partial x} \left(\eta \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right) + \frac{\partial}{\partial z} \left(\eta \frac{\partial v}{\partial z} \right) = \rho g \frac{\partial s}{\partial y},$$

$$(6.1)$$

where u and v are the components of horizontal velocity, η is the effective viscosity, s is the ice surface elevation, ρ is the density of ice (assumed constant), and g is gravitational acceleration.

As in Glide, the equations are discretized on a structured 3D mesh. In the map plane the mesh consists of rectangular cells. These cells form an unstaggered 2D grid of dimension (nx, ny) (thus the number of cells is the product (nx)(ny)), together with a staggered grid of dimension (nx-1,ny-1). The corners of each cell (where four rectangles meet) are called vertices. The vertical levels of the mesh are based on a terrain-following sigma coordinate system. We define $\sigma = (s-z)/H$, where H is the ice thickness, with $\sigma = 0$ at the top surface and $\sigma = 1$ at the bed. There are nz levels in the vertical direction, with nz-1 layers between these levels. An element is the region associated with a particular cell and layer; there are (nx)(ny)(nz-1) elements on the mesh. A node is a point where eight elements intersect (or where four elements intersect at the upper or lower surface). There are (nx-1)(ny-1)(nz) nodes on the mesh.

Scalar 2D fields such as H and s are defined for each cell. Scalar 3D fields such as ice temperature T lie at the center of each element (i.e., at the midpoint of each layer associated with each cell). Gradients of 2D scalar fields (e.g., the surface slope ∇s) are defined at vertices. The velocity components u and v live at nodes.

For problems solved in parallel, the domain is partitioned among multiple processors. Let $nx_{\rm local}$ and $ny_{\rm local}$ be the number of locally owned cells on each processor. Each processor holds data for two rows of halo cells (i.e., cells belonging to other processors) surrounding the block of locally owned cells. Thus the 2D grid on a given processor has dimensions $(nx_{\rm local}+2, ny_{\rm local}+2)$. Each locally owned cell is associated with a locally owned vertex lying at the northeast (upper right) corner of the cell.

An active cell is a cell that borders a locally owned vertex, whose ice thickness exceeds a minimum threshold. Each active cell is associated with a column of nz-1 active elements. (All the data in a given column resides on a single processor.) An active vertex is any vertex of an active cell. Each active vertex is associated with nz active nodes, including nodes at the surface and bed.

The effective viscosity is defined in each active element by

$$\eta \equiv \frac{1}{2} A^{\frac{-1}{n}} \dot{\varepsilon}_e^{\frac{1-n}{n}},\tag{6.2}$$

where A is the temperature-dependent rate factor in Glen's flow law, and $\dot{\varepsilon}_e$ is the effective strain rate, given in the Blatter-Pattyn approximation by

$$\dot{\varepsilon}_e^2 = \dot{\varepsilon}_{xx}^2 + \dot{\varepsilon}_{yy}^2 + \dot{\varepsilon}_{xx}\dot{\varepsilon}_{yy} + \dot{\varepsilon}_{xy}^2 + \dot{\varepsilon}_{xz}^2 + \dot{\varepsilon}_{yz}^2, \tag{6.3}$$

where

$$\dot{\varepsilon}_{ij} = \frac{1}{2} \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right). \tag{6.4}$$

for the horizontal terms $(\dot{\varepsilon}_{xx}, \dot{\varepsilon}_{xy} \text{ and } \dot{\varepsilon}_{yy})$, but the vertical terms are simplified as $\dot{\varepsilon}_{xz} = \frac{1}{2} \frac{\partial x}{\partial z}$ and $\dot{\varepsilon}_{yz} = \frac{1}{2} \frac{\partial y}{\partial z}$. To compute the rate factor A, we assume an Arrhenius relationship:

$$A(T^*) = ae^{-Q/RT^*}, (6.5)$$

where T^* is the absolute temperature corrected for the dependence of the melting point on pressure ($T^* = T + 8.7 \cdot 10^{-4} (s - z)$, T in Kelvin), a is a temperature–independent material constant (given by Paterson and Budd (1982)), Q is the activation energy for creep, and R is the universal gas constant.

Given T, s, H, and an initial guess for u and v, the problem is to solve Eq. (6.1) for u and v at each active node. (At inactive nodes we set u = v = 0.) This problem can be written in the form

$$\mathbf{A}\mathbf{x} = \mathbf{b},\tag{6.6}$$

or more fully,

$$\begin{bmatrix} \mathbf{A}_{\mathbf{u}\mathbf{u}} & \mathbf{A}_{\mathbf{u}\mathbf{v}} \\ \mathbf{A}_{\mathbf{v}\mathbf{u}} & \mathbf{A}_{\mathbf{v}\mathbf{v}} \end{bmatrix} \begin{bmatrix} \mathbf{u} \\ \mathbf{v} \end{bmatrix} = \begin{bmatrix} \mathbf{b}_{\mathbf{u}} \\ \mathbf{b}_{\mathbf{v}} \end{bmatrix}. \tag{6.7}$$

Eq. (6.7) shows the four parts of the global matrix \mathbf{A} , with the solution and right-hand-side vectors separated into u and v components. In Glissade, \mathbf{A} is always symmetric and positive-definite.

Since **A** depends (through η) on u and v, the problem is nonlinear and must be solved iteratively (see Section 5.4.2). For each nonlinear iteration, Glissade computes the 3D η field based on the current guess for the velocity field and solves a linear problem of the form (6.7). Then η is updated and the process is repeated until the solution converges to within a given tolerance. This procedure is known as Picard iteration.

The following sections describe how the matrix equations are assembled and solved.

Assembly

The coupled PDEs (6.1) are discretized using the finite-element method. Here we give a detailed but non-rigorous description of the method as applied to the Blatter-Pattyn approximation on the CISM mesh. We refer the reader to standard texts (e.g., Hughes, 2000; Huebner et al., 2001) for a full discussion of finite elements.

The PDEs, with appropriate boundary conditions, are converted to a system of algebraic equations by dividing the full domain into subdomains (i.e., elements), representing the velocity solution on each element, and integrating over elements. On the CISM mesh, the elements are hexahedra (rectangles in map view), each of which has eight nodes shared with its neighbors. The solution is approximated as a sum over basis functions φ . Each active node is associated with a basis function whose value is $\varphi=1$ at that node, with $\varphi=0$ at all other nodes. The solution at a point within an element can be expanded in terms of basis functions and nodal values:

$$u(x,y,z) = \sum_{n} \varphi_n(x,y,z)u_n, \quad v(x,y,z) = \sum_{n} \varphi_n(x,y,z)v_n, \tag{6.8}$$

where the sum is over the nodes of the element; u_n and v_n are nodal values of the solution; and φ_n varies smoothly between 0 and 1 within the element. The fact that $\varphi = 0$ except in a small region ensures that the sum includes only as many terms as there are nodes per element.

Glissade uses standard finite-element techniques to represent the PDE on each element and assemble the element equations into a global set of algebraic equations of the form (6.7). The scheme is formally equivalent to that described by Perego et al. (2012) (henceforth PGB). Eq. (6.1) can be written as

$$-\nabla \cdot (2\eta \dot{\varepsilon}_1) = -\rho g \frac{\partial s}{\partial x},$$

$$-\nabla \cdot (2\eta \dot{\varepsilon}_2) = -\rho g \frac{\partial s}{\partial y},$$
(6.9)

where

$$\dot{\varepsilon}_{1} = \begin{bmatrix} 2\dot{\varepsilon}_{xx} + \dot{\varepsilon}_{yy} \\ \dot{\varepsilon}_{xy} \\ \dot{\varepsilon}_{xz} \end{bmatrix}, \quad \dot{\varepsilon}_{2} = \begin{bmatrix} \dot{\varepsilon}_{xy} \\ \dot{\varepsilon}_{xx} + 2\dot{\varepsilon}_{yy} \\ \dot{\varepsilon}_{yz} \end{bmatrix}. \tag{6.10}$$

(These are Eqs. 12 and 13 in PGB.) We rewrite the equations in *weak form* (see PGB Eq. 15), which is obtained by multiplying (6.9) by the basis functions and integrating over the domain, using integration by parts to eliminate the second derivative:

$$x: \int_{\Omega} 2\eta \dot{\varepsilon}_{1}(u,v) \cdot \nabla \varphi_{1} \, d\Omega + \int_{\Gamma_{B}} \beta u \varphi_{1} \, d\Gamma + \int_{\Gamma_{L}} p n_{1} \varphi_{1} \, d\Gamma + \int_{\Omega} \rho g \frac{\partial s}{\partial x} \varphi_{1} \, d\Omega = 0,$$

$$y: \int_{\Omega} 2\eta \dot{\varepsilon}_{2}(u,v) \cdot \nabla \varphi_{2} \, d\Omega + \int_{\Gamma_{B}} \beta u \varphi_{2} \, d\Gamma + \int_{\Gamma_{L}} p n_{2} \varphi_{2} \, d\Gamma + \int_{\Omega} \rho g \frac{\partial s}{\partial y} \varphi_{2} \, d\Omega = 0,$$

$$(6.11)$$

where Ω represents the domain volume, Γ_B denotes the lower boundary, Γ_L denotes the lateral boundary (e.g., the calving front of an ice shelf), β is a basal traction parameter, p is the pressure at the lateral boundary, and n_1 and n_2 are components of the normal to Γ_L . These equations can also be obtained from a variational principle as described by Dukowicz et al. (2010).

The four terms on the LHS of (6.11) describe internal ice stresses, basal friction, lateral pressure, and the gravitational driving force, respectively. Next we describe how these terms are summed over elements and assembled into the global matrix $\bf A$ and right-hand side vector $\bf b$.

Internal ice stresses. We start with the internal stress term, which is the most complex. We rewrite the first term on the LHS of (6.11) in terms of velocity components:

$$x: \int_{\Omega} 2\eta \left[2\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} - \frac{1}{2} \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) - \frac{1}{2} \frac{\partial u}{\partial z} \right] \begin{Bmatrix} \frac{\partial \varphi}{\partial x} \\ \frac{\partial \varphi}{\partial y} \\ \frac{\partial \varphi}{\partial z} \end{Bmatrix},$$

$$y: \int_{\Omega} 2\eta \left[\frac{1}{2} \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) - 2\frac{\partial v}{\partial y} + \frac{\partial u}{\partial x} - \frac{1}{2} \frac{\partial v}{\partial z} \right] \begin{Bmatrix} \frac{\partial \varphi}{\partial x} \\ \frac{\partial \varphi}{\partial y} \\ \frac{\partial \varphi}{\partial y} \end{Bmatrix},$$

$$(6.12)$$

where brackets are used for row vectors and braces for column vectors. Glissade evaluates (6.12) for each active element. Recall that hexahedral elements have eight nodes, with u and v to be determined at each active node. Inserting the velocity expressions (6.8) in (6.12), we obtain

$$x: \int_{\Omega} 2\eta \left(\frac{\partial \varphi_{i}}{\partial x} \left(2 \left[\frac{\partial \varphi_{j}}{\partial x} \right] \left\{ u_{j} \right\} + \left[\frac{\partial \varphi_{j}}{\partial y} \right] \left\{ v_{j} \right\} \right) + \frac{\partial \varphi_{i}}{\partial y} \left(\frac{1}{2} \left[\frac{\partial \varphi_{j}}{\partial y} \right] \left\{ u_{j} \right\} + \frac{1}{2} \left[\frac{\partial \varphi_{j}}{\partial x} \right] \left\{ v_{j} \right\} \right) + \frac{\partial \varphi_{i}}{\partial z} \left(\frac{1}{2} \left[\frac{\partial \varphi_{j}}{\partial z} \right] \left\{ u_{j} \right\} \right)$$

$$y: \int_{\Omega} 2\eta \left(\frac{\partial \varphi_{i}}{\partial y} \left(2 \left[\frac{\partial \varphi_{j}}{\partial y} \right] \left\{ v_{j} \right\} + \left[\frac{\partial \varphi_{j}}{\partial x} \right] \left\{ u_{j} \right\} \right) + \frac{\partial \varphi_{i}}{\partial x} \left(\frac{1}{2} \left[\frac{\partial \varphi_{j}}{\partial x} \right] \left\{ v_{j} \right\} + \frac{1}{2} \left[\frac{\partial \varphi_{j}}{\partial y} \right] \left\{ u_{j} \right\} \right) + \frac{\partial \varphi_{i}}{\partial z} \left(\frac{1}{2} \left[\frac{\partial \varphi_{j}}{\partial z} \right] \left\{ v_{j} \right\} \right)$$

$$(6.13)$$

Each row or column vector has eight terms, one for each node of the element. These terms can be evaluated to form a set of four 8x8 element matrices. Each row of an element matrix is associated with u or v at a given node. The columns in this row contain terms linking the node to u or v at the other nodes of the same element (with the diagonal term linking the node to itself).

In the x component of (6.13), the terms that multiply u_i are given by

$$\int_{\Omega} \eta \left(4 \frac{\partial \varphi_i}{\partial x} \frac{\partial \varphi_j}{\partial x} + \frac{\partial \varphi_i}{\partial y} \frac{\partial \varphi_j}{\partial y} + \frac{\partial \varphi_i}{\partial z} \frac{\partial \varphi_j}{\partial z} \right) d\Omega.$$
 (6.14)

Letting i and j range from 1 to 8, (6.14) gives the 64 terms of the 8x8 element matrix $\mathbf{K_{uu}}$, which links the u value at each node to the u values at all eight nodes. Similarly, the 64 terms of element matrix $\mathbf{K_{uv}}$, which links u at each node to v at each of the eight nodes, are given by

$$\int_{\Omega} \eta \left(2 \frac{\partial \varphi_i}{\partial x} \frac{\partial \varphi_j}{\partial y} + \frac{\partial \varphi_i}{\partial y} \frac{\partial \varphi_j}{\partial x} \right) d\Omega. \tag{6.15}$$

Likewise, two 8x8 matrices are associated with the y component of (6.13). The terms of $\mathbf{K_{vu}}$, which connects v at each node to u at each of eight nodes, are given by

$$\int_{\Omega} \eta \left(2 \frac{\partial \varphi_i}{\partial y} \frac{\partial \varphi_j}{\partial x} + \frac{\partial \varphi_i}{\partial x} \frac{\partial \varphi_j}{\partial y} \right) d\Omega. \tag{6.16}$$

Finally, the terms of $\mathbf{K}_{\mathbf{v}\mathbf{v}}$, which links v at each node to v at each of eight nodes, are given by

$$\int_{\Omega} \eta \left(4 \frac{\partial \varphi_i}{\partial y} \frac{\partial \varphi_j}{\partial y} + \frac{\partial \varphi_i}{\partial x} \frac{\partial \varphi_j}{\partial x} + \frac{\partial \varphi_i}{\partial z} \frac{\partial \varphi_j}{\partial z} \right) d\Omega. \tag{6.17}$$

Because of the symmetry of the underlying PDEs, $\mathbf{K_{uu}}$ and $\mathbf{K_{vv}}$ are symmetric, and $\mathbf{K_{uv}} = \mathbf{K_{vu}}^T$. Note that $\mathbf{K_{vv}}$ can be obtained from $\mathbf{K_{uu}}$, and $\mathbf{K_{vu}}$ from $\mathbf{K_{uv}}$, by exchanging x and y. The terms containing z (i.e., the vertical shear stresses associated with the shallow-ice approximation) appear only in $\mathbf{K_{uu}}$ and $\mathbf{K_{vv}}$. The terms containing x and y (i.e., the membrane stresses) appear in all four element matrices.

Eqs. (6.14)–(6.17) lie at the heart of the code. Together with the expressions for the effective viscosity η (discussed below), these equations contain the physical content of the Blatter-Pattyn approximation.

In the weak form of the equations, each of the 64 coefficients in each element matrix must be integrated over the element. (Since φ varies over the element, the integrands in (6.14)–(6.17) have a different value at each point.) This is done for a given element by evaluating the integrand at each of eight *quadrature points* and summing over quadrature points. We first specify the form of the basis functions, then transform the basis functions to the geometry of

the element (which is irregular in the vertical because of the sigma coordinate) and evaluate the basis function derivatives at the quadrature points.

Glissade uses trilinear basis functions defined on a reference cube. This cube is centered at the origin (0,0,0) in local reference coordinates $(\hat{x},\hat{y},\hat{z})$. The eight nodes of the reference cube are located at $(\hat{x},\hat{y},\hat{z})=(\pm 1,\pm 1,\pm 1)$. By convention, nodes 1–4 are the nodes of the lower face, proceeding counterclockwise from the southwest corner $(\hat{x},\hat{y})=(-1,-1)$. Nodes 5–8 are the nodes of the upper face, also moving counterclockwise from the southwest corner. Thus we have

$$\varphi_{1} = (1 - \hat{x})(1 - \hat{y})(1 - \hat{z})/8,
\varphi_{2} = (1 + \hat{x})(1 - \hat{y})(1 - \hat{z})/8,
\varphi_{3} = (1 + \hat{x})(1 + \hat{y})(1 - \hat{z})/8,
\varphi_{4} = (1 - \hat{x})(1 + \hat{y})(1 - \hat{z})/8,
\varphi_{5} = (1 - \hat{x})(1 - \hat{y})(1 + \hat{z})/8,
\varphi_{6} = (1 + \hat{x})(1 - \hat{y})(1 + \hat{z})/8,
\varphi_{7} = (1 + \hat{x})(1 + \hat{y})(1 + \hat{z})/8,
\varphi_{8} = (1 - \hat{x})(1 + \hat{y})(1 + \hat{z})/8.$$
(6.18)

For each n we have $\varphi_n = 1$ at a single node, with $\varphi_n = 0$ at the other nodes.

The integrands in (6.14)–(6.17) are written in terms of real Cartesian coordinates (x, y, z) rather than reference coordinates $(\hat{x}, \hat{y}, \hat{z})$. Spatial derivatives in real coordinates are related to derivatives in reference coordinates by

where [J] is the Jacobian of the transformation between coordinate systems. Given the finite-element expansion

$$x = \sum_{n} \varphi_n x_n, \tag{6.20}$$

along with the spatial derivatives of φ at $(\hat{x}, \hat{y}, \hat{z})$ (which are easily derived from (6.18)), we can compute $[J(\hat{x}, \hat{y}, \hat{z})]$ as

$$[J] = \begin{bmatrix} \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{x}} x_n & \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{x}} y_n & \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{x}} z_n \\ \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{y}} x_n & \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{y}} y_n & \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{y}} z_n \\ \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{z}} x_n & \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{z}} y_n & \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial \hat{z}} z_n \end{bmatrix}.$$
(6.21)

We then invert (6.19) to obtain the basis function derivatives in terms of (x, y, z):

The left-hand side of (6.22) contains the spatial derivatives appearing in (6.14)–(6.17).

Eqs. (6.14)–(6.17) also contain the viscosity η , which is computed at each quadrature point. In the Blatter-Pattyn approximation, η is given by (6.2); it is a function of the flow factor A and the effective strain rate defined by (6.3). We approximate A by its value at the element center. The (squared) effective strain rate, $\dot{\varepsilon}_e^2$, is evaluated at each quadrature point by summing over strain-rate components. The x components are given by

$$\frac{\partial u}{\partial x} = \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial x} u_n, \quad \frac{\partial v}{\partial x} = \sum_{n=1}^{8} \frac{\partial \varphi_n}{\partial x} v_n, \tag{6.23}$$

and similarly for the y and z components. The nodal velocities in (6.23) are values from the previous iteration; otherwise the resulting system of equations would be nonlinear.

We now have the information needed to compute the integrands (6.14)–(6.17) at quadrature points. To integrate over a hexahedron, we take a weighted sum of the values at each of eight quadrature points. These points are located at reference coordinates $(\hat{x}, \hat{y}, \hat{z}) = (\pm 1/\sqrt{3}, \pm 1/\sqrt{3}, \pm 1/\sqrt{3})$. To evaluate an integral of the form

$$\int_{\Omega} \eta \left(\frac{\partial \varphi_i}{\partial z} \frac{\partial \varphi_j}{\partial z} \right) d\Omega \tag{6.24}$$

over element volume Ω , we compute the sum over quadrature points

$$\sum_{p=1}^{8} w_p \eta_p \left(\frac{\partial \varphi_i}{\partial z} \frac{\partial \varphi_j}{\partial z} \right)_p |J_p|, \tag{6.25}$$

where |J| is the determinant of the Jacobian (6.21). For this choice of quadrature points, each point has $w_p = 1$.

The terms of the element matrices $\mathbf{K_{uu}}$, $\mathbf{K_{vu}}$, $\mathbf{K_{vu}}$ and $\mathbf{K_{vv}}$ are then inserted into the corresponding global matrices $\mathbf{A_{uu}}$, $\mathbf{A_{uv}}$, $\mathbf{A_{vu}}$ and $\mathbf{A_{vv}}$. This is mostly a matter of bookkeeping. For example, the first row of $\mathbf{K_{uu}}$ corresponds to a particular node of element (k,i,j) (specifically, the node with indices (k-1,i-1,j-1), given our convention for numbering nodes within elements). This row corresponds to a row of the global matrix $\mathbf{A_{uu}}$, and each of the eight terms in the row is associated with a column of $\mathbf{A_{uu}}$. Glissade determines the correct column and adds the $\mathbf{K_{uu}}$ term to the corresponding term in $\mathbf{A_{uu}}$. This process proceeds until the code has looped over all the active elements and filled the global matrices.

If written in full, each global matrix would have as many rows and columns as there are active nodes. These matrices, however, are sparse, with a maximum of 27 nonzero terms per row (corresponding to a node and its 26 nearest neighbors in a 3D hexahedral lattice). Glissade therefore assembles and stores arrays of dimension (27, nz, nx - 1, ny - 1). The 27 terms of the first array dimension are indexed such that each index has a geometric meaning. Suppose we are filling columns for the matrix row corresponding to node (k, i, j). Then, by convention, index 1 refers to the node with coordinates (k-1, i-1, j-1), index 14 refers to the node itself (i.e., the diagonal term of the row), and index 27 refers to the node at (k+1, i+1, j+1) (and similarly for the other indices). After assembly, these matrices can be converted as needed to the form required by a particular solver.

The remaining assembly consists of evaluating the other terms in (6.11) (i.e., the basal and lateral boundary conditions and the gravitational forcing) and implementing Dirichlet boundary conditions, if applicable. We consider these in turn.

Basal boundary conditions. At the basal boundary we assume a friction law of the form

$$\tau_{\mathbf{b}} = \beta \mathbf{u}_{\mathbf{b}}.\tag{6.26}$$

The coefficient β is defined at each vertex and can vary spatially. If β depends on the velocity, as in some friction laws, then it is calculated using the velocity from the previous iteration. See Sections 5.3.2 and 5.3.3 for a detailed discussion of basal traction in higher-order models.

The basal boundary terms to be evaluated in (6.11) are

$$x: \int_{\Gamma_B} \beta u \varphi_1 d\Gamma,$$

$$y: \int_{\Gamma_B} \beta v \varphi_2 d\Gamma.$$
(6.27)

The basal face of each cell is a rectangle. To integrate over a rectangle, we sum over terms at four quadrature points. These points lie at $(\hat{x}, \hat{y}) = (\pm 1/\sqrt{3}, \pm 1/\sqrt{3})$ in a reference square with center (0,0) and vertices $(\pm 1, \pm 1)$. (This reference square is the 2D analog of the reference cube discussed above.) We define four bilinear basis functions on the square (cf. (6.18)):

$$\varphi_{1} = (1 - \hat{x})(1 - \hat{y})/4,
\varphi_{2} = (1 + \hat{x})(1 - \hat{y})/4,
\varphi_{3} = (1 + \hat{x})(1 + \hat{y})/4,
\varphi_{4} = (1 - \hat{x})(1 + \hat{y})/4.$$
(6.28)

Given these basis functions and their spatial derivatives, we can compute the Jacobian for the transformation between the reference square and the rectangular cell face, using the 2D versions of (6.21) and (6.22):

$$[J] = \begin{bmatrix} \sum_{n=1}^{4} \frac{\partial \varphi_n}{\partial \hat{x}} x_n & \sum_{n=1}^{4} \frac{\partial \varphi_n}{\partial \hat{x}} y_n \\ \sum_{n=1}^{4} \frac{\partial \varphi_n}{\partial \hat{y}} x_n & \sum_{n=1}^{4} \frac{\partial \varphi_n}{\partial \hat{y}} y_n \end{bmatrix}, \tag{6.29}$$

The integrand at a quadrature point has the form

$$\beta \varphi_i \varphi_j,$$
 (6.31)

where the second φ term arises from the finite-element expansion of u at a quadrature point:

$$u = \sum_{n=1}^{4} u_n \varphi_n. \tag{6.32}$$

We determine β at quadrature points from the values at cell vertices:

$$\beta = \sum_{n=1}^{4} \beta_n \varphi_n. \tag{6.33}$$

The integral of (6.31) over a cell is then computed as a sum over quadrature points:

$$\sum_{p=1}^{4} w_p \beta_p(\varphi_i \varphi_j)_p |J_p|, \tag{6.34}$$

where $w_p = 1$ for each point. This procedure generates a 4x4 matrix describing the connections between each node and its neighbors in the cell. Since the x term in (6.27) contains u but not v, and the y term contains v but not u, we form 2D matrices $\mathbf{K}_{\mathbf{u}\mathbf{u}}$ and $\mathbf{K}_{\mathbf{v}\mathbf{v}}$, but not $\mathbf{K}_{\mathbf{u}\mathbf{v}}$ and $\mathbf{K}_{\mathbf{v}\mathbf{u}}$. Each term of $\mathbf{K}_{\mathbf{u}\mathbf{u}}$ is then inserted into the global matrix $\mathbf{A}_{\mathbf{u}\mathbf{u}}$, and similarly for $\mathbf{K}_{\mathbf{v}\mathbf{v}}$ and $\mathbf{A}_{\mathbf{v}\mathbf{v}}$.

This assembly procedure tends to smooth the β field. If it is desirable to resolve sharp discontinuities in β , as in the stream test problem of Section 8.2.3, Glissade supports an alternate assembly procedure in which the terms in a given row of the matrix depend only on β at the vertex associated with that particular row of the matrix. See option which_ho_assemble_beta in Section 7.3.

Lateral boundary conditions. The lateral boundary terms in (6.11) are

$$x: \int_{\Gamma_L} pn_1 \varphi_1 \, d\Gamma,$$

$$y: \int_{\Gamma_L} pn_2 \varphi_2 \, d\Gamma.$$
(6.35)

Since these terms are independent of u and v, they contribute to the load vectors $\mathbf{b_u}$ and $\mathbf{b_v}$ on the right-hand side of (6.7). They are integrated over the lateral faces of floating cells that border the ocean. (Grounded cells are assumed to have no lateral forcing.)

The lateral faces bordering the ocean are quadrilaterals that can be mapped to a reference square. The integral over each face is found by summing over four quadrature points. Basis functions are given by (6.28), and the Jacobian of the reference square is computed using (6.29). We evaluate the ice thickness H at each quadrature point using

$$H = \sum_{n=1}^{4} H_n \varphi_n, \tag{6.36}$$

where the H_n are nodal values. The integrands have the form $p_n\varphi$, where p_n is the vertically averaged net pressure normal to the ice edge. (We use the vertically averaged pressure to avoid dealing with vertical shear at the ice edge.) The net pressure is equal to the pressure directed outward from the ice toward the ocean by the weight of the ice, minus the (smaller) pressure directed inward from the ocean to the ice by the hydrostatic water pressure. The outward pressure is obtained by integrating $\rho_i g(s-z)dz$ from s-H to s and then dividing by H; it is given by

$$p_{\text{out}} = \frac{\rho_i g H}{2}.\tag{6.37}$$

The inward pressure is found by integrating $(-\rho_w gzdz)$ from s-H to 0 and then dividing by H-s; it is given by

$$p_{\rm in} = \frac{\rho_w g(s-H)^2}{2H} \tag{6.38}$$

Assuming hydrostatic balance, we have $s - H = (\rho_i/\rho_w)H$. Thus (6.38) becomes

$$p_{\rm in} = \frac{\rho_i g H}{2} \frac{\rho_i}{\rho_w} \tag{6.39}$$

Combining (6.37) and (6.39) gives

$$p_{\text{net}} = \frac{\rho_i g H}{2} \left(1 - \frac{\rho_i}{\rho_w} \right), \tag{6.40}$$

directed from the ice to the ocean. The integral of the pressure terms over a lateral face is computed as a sum over quadrature points:

$$\sum_{p=1}^{4} \pm w_p(p_{\text{net}})_p(\varphi_i)_p |J_p|, \tag{6.41}$$

where the sign depends on the orientation of the face. The resulting pressure terms are inserted into the load vector (either $\mathbf{b_u}$ or $\mathbf{b_v}$, depending on the orientation) in the rows associated with each of the four nodes of the face.

Gravitational driving stress. The gravitational forcing terms in (6.11) are

$$x: \int_{\Omega} \rho g \frac{\partial s}{\partial x} \varphi_1 \, d\Omega,$$

$$y: \int_{\Omega} \rho g \frac{\partial s}{\partial y} \varphi_2 \, d\Omega.$$
(6.42)

To compute these terms we evaluate $\frac{\partial s}{\partial x}$ and $\frac{\partial s}{\partial y}$ at each active vertex. A standard centered approximation at vertex (i,j) is

$$\frac{\partial s}{\partial x}(i,j) = \frac{s(i+1,j+1) + s(i+1,j) - s(i,j+1) - s(i,j)}{2\Delta x},$$
(6.43)

and similarly for $\frac{\partial s}{\partial y}$. This is equivalent to computing a gradient in the x direction on the edges above and below the vertex, and then averaging the two gradients to the vertex. A centered approximation works well when the ice geometry is fixed but can cause problems when the geometry is evolving. These problems arise because checkerboard noise in s (which is common on structured meshes with the velocity at cell vertices) is invisible to the momentum balance; it is canceled out by the centered averaging in (6.43). Checkerboard noise can therefore persist and grow. To damp this noise, Glissade can use an upstream average:

$$\frac{\partial s}{\partial x}(i,j) = \frac{1.5(s(i+1,j+1) - s(i,j+1)) - 0.5(s(i+1,j+2) - s(i,j+2))}{\Delta x}.$$
 (6.44)

Here, "upstream" means in the direction of increasing surface elevation. Both (6.43) and (6.44) are second-order accurate. The default is (6.43), but (6.44) can be chosen by setting which_ho_gradient = 1 in the config file (see Section 7.3).

Eqs. (6.43) and (6.44) are ambiguous at the ice margin, where at least one of the four cells neighboring a vertex is ice-free. (Cells with very thin ice, H < thklim, are considered ice-free by the velocity solver. CISM's default value is thklim = 100 m, which is appropriate for Glide, but Glissade is typically run with thklim = 1 m.) One option is to include all cells, including ice-free cells, in the gradient. This generally works well for land-based ice but gives large gradients and excessive ice speeds at floating shelf margins. A second option is to include only ice-covered cells in the gradient. For example, suppose cells (i,j) and (i,j+1) have ice, but cells (i+1,j) and (i+1,j+1) are ice-free. Then, lacking the required information to compute x gradients at the adjacent edges, we would set $\frac{\partial s}{\partial x} = 0$. The y gradient would be one-sided: $\frac{\partial s}{\partial y} = (s(i,j+1)-s(i,j))/\Delta y$. This option works well at shelf margins but tends to underestimate gradients at land margins. A third, hybrid option is to compute a gradient at edges where either (1) both adjacent cells are ice-covered, or (2) one cell is ice-covered and grounded, and lies higher in elevation than the neighboring ice-free cell. Thus, gradients are masked out at edges where a floating cell is adjacent to an ocean cell, and where a grounded ice-covered cell lies below

ice-free land (i.e., a nunatak). Since this option (which_ho_gradient_margin = 1) works well for both land and shelf margins, it is the default.

The integrals in (6.42) are over 3D elements. Hence each hexahedral element is mapped to a reference cube as described above. Given $\frac{\partial s}{\partial x}$ at the nodes of a cell, the surface slope terms at quadrature points are

$$\frac{\partial s}{\partial x} = \sum_{n=1}^{8} \varphi_n \left(\frac{\partial s}{\partial x}\right)_n, \quad \frac{\partial s}{\partial y} = \sum_{n=1}^{8} \varphi_n \left(\frac{\partial s}{\partial y}\right)_n, \tag{6.45}$$

where the basis functions φ are given by (6.18) and the spatial derivatives are derived from (6.21) and (6.22). The integral of $\rho g \frac{\partial s}{\partial x} \varphi$ over an element is evaluated as a sum over quadrature points:

$$\sum_{p=1}^{8} w_p \rho g \left(\frac{\partial s}{\partial x} \right)_p (\varphi_i)_p |J_p|, \tag{6.46}$$

and similarly for the $\frac{\partial s}{\partial y}$ term. Glissade inserts these terms into the load vectors $\mathbf{b_u}$ and $\mathbf{b_v}$.

Dirichlet boundary conditions. Once the matrix has been assembled, it may need to be adjusted for Dirichlet boundary conditions (i.e., prescribed values of the velocity at certain nodes). The most common Dirichlet condition is to set u = v = 0 at the bed to enforce a no-slip boundary condition. (A no-slip condition can also be enforced by setting the basal traction coefficient β to a very large value, but formally this is not a Dirichlet condition.) Also, it may be desirable to set u and v to observed values at certain locations, as in the Ross Ice Shelf test case (Section 8.2.6).

Suppose that at node (k, i, j) we have $u = u_c$ and $v = v_c$, where u_c and v_c are prescribed values. Let nr be the row of $\mathbf{A_{uu}}$ associated with this node, and let nc range over the columns with nonzero entries in this row. To enforce the Dirichlet condition, we set $\mathbf{A_{uu}}(nr, nc) = \mathbf{A_{vv}}(nr, nc) = 0$ for all values of nc except nc = nr (the diagonal term); we set $\mathbf{A_{uu}}(nr, nr) = \mathbf{A_{vv}}(nr, nr) = 1$. In addition, we set $\mathbf{A_{uv}}(nr, nc) = \mathbf{A_{vu}}(nr, nc) = 0$ for all nc, since these two matrices do not contain any terms on the diagonal of the full global matrix (i.e., \mathbf{A} in (6.6)). On the right-hand side, we set $\mathbf{b_u}(nr) = u_c$ and $\mathbf{b_v}(nr) = v_c$. These operations convert the matrix rows associated with node (k, i, j) to the equations $1 \cdot u = u_c, 1 \cdot v = v_c$, which clearly have the desired solutions u_c and v_c .

A further step is needed to maintain matrix symmetry, as required when using a preconditioned conjugate gradient solver. Consider the term $\mathbf{A}_{\mathbf{u}\mathbf{u}}(nr,nc)$, where nc is a specific column associated with a neighboring node (say, node (k+1,i+1,j+1)). We have already set $\mathbf{A}_{\mathbf{u}\mathbf{u}}(nr,nc)=0$, so we need to set $\mathbf{A}_{\mathbf{u}\mathbf{u}}(nc,nr)=0$ to maintain symmetry. The Dirichlet condition is $u(nr)=u_c$. In the matrix-vector product, the product $\mathbf{A}_{\mathbf{u}\mathbf{u}}(nc,nr)u(nr)=\mathbf{A}_{\mathbf{u}\mathbf{u}}(nc,nr)u_c$ contributes to $\mathbf{b}_{\mathbf{u}}(nc)$. Thus we can replace $\mathbf{b}_{\mathbf{u}}(nc)$ with $\mathbf{b}_{\mathbf{u}}(nc)-\mathbf{A}_{\mathbf{u}\mathbf{u}}(nc,nr)u_c$ and set $\mathbf{A}_{\mathbf{u}\mathbf{u}}(nc,nr)=0$ without altering the problem. We do this for all the terms in the columns associated with node (k,i,j) (i.e., all the terms multiplied by u_c or v_c in the matrix-vector product). Thus, both the rows and the columns associated with node (k,i,j) are filled with zeros, except for the diagonal term, and the full global matrix remains symmetric.

Solving the linear system

Once the matrices and right-hand side vectors have been assembled, we solve the linear problem (6.7). Glissade supports three kinds of solvers:

- A native Fortran 90 preconditioned conjugate gradient (PCG) solver
- Links to the Sparse Linear Algebra Package (SLAP), with options for the generalized minimum residual (GMRES) and biconjugate gradient methods

• Links to Trilinos, with options for PCG and GMRES, preconditioned by incomplete lowerupper (ILU) factorization or a multigrid method

We describe each solver in turn.

Preconditioned conjugate gradient solver. The native PCG solver works directly with the assembled matrices A_{uu} , A_{uv} , A_{vu} , and A_{vv} , along with the right-hand side vectors b_u and $\mathbf{b_v}$. These matrices are passed to one of two PCG solvers: a "standard" solver or a Chronopolous-Gear solver. The standard method for solving $\mathbf{A}\mathbf{x} = \mathbf{b}$ can be summarized as follows:

• Let x_0 denote the initial guess for x ($x_0 = 0$ if no guess is available); r = b - Ax is the residual vector; d is a conjugate direction vector; M is a preconditioning matrix: α, β, η_0 , η_1 and η_2 are scalars; q and z are work vectors; and (r,z) is the dot product of two vectors.

```
• r_0 = b - Ax_0, d_0 = 0, \eta_0 = 1
```

• Do until converged:

```
* Solve Mz = r for z (the preconditioning step)
\star \eta_1 = (r, z)
\star \beta = \eta_1/\eta_0
\star d = z + \beta d
\star \eta_0 = \eta_1
\star q = Ad
\star \eta_2 = (d,q)
\star \alpha = \eta_1/\eta_2
\star x = x + \alpha d
\star r = r - \alpha q (or periodically, set r = b - Ax and check for convergence)
```

There are two dot products, one matrix-vector product, and one preconditioning step per iteration. The convergence condition is $\sqrt{(r,r)}/\sqrt{(b,b)} < \epsilon$, where ϵ is a small tolerance (10⁻⁸ by default). The convergence check is done every five iterations, as a compromise between the expense of an extra matrix-vector multiply and that of unnecessary iterations.

This solver works either in serial or in parallel. If run in parallel, the dot products require global sums, and a halo update for d is needed once per iteration. Halo updates in CISM operate on 2D arrays with horizontal indices i and j (or 3D/4D arrays with additional indices). This is the main reason for leaving the global matrices and vectors in standard Fortran arrays, instead of converting to a sparse matrix storage format such as compressed row storage. In array form, it is easy to do halo updates of vectors u, v, r_u, r_v, d_u, d_v , etc.

The PCG solver has two preconditioning options:

- Diagonal (also known as Jacoby) preconditioning, in which the preconditioning matrix M consists of the diagonal terms of A, so that M is trivial to invert. Convergence with diagonal preconditioning can be slow.
- Shallow-ice-based preconditioning (for the 3D Blatter-Pattyn approximation only). In this case M includes only the terms in A that link a given node to itself and its immediate neighbors above and below. The matrix M is then tridiagonal, as in the shallow-ice approximation, and can be inverted efficiently. This preconditioner works well when the physics is dominated by vertical shear and the horizontal terms are small, but not as well when membrane stresses are important. Since the preconditioner is local (it consists of independent column solves), it scales well with an increasing number of processors.

Other preconditioning options could be added in the future. If the existing options are inefficient for a given problem, it may be better to use Trilinos (see below).

For small problems on a modest number of processors, the dominant cost is the matrix-vector multiply required during each iteration. For large problems on many processors, however, the global sums become increasingly expensive. To reduce the cost of global sums, the standard PCG algorithm can be replaced by the Chronopoulos-Gear algorithm, which rearranges operations such that both global sums are done with a single MPI call. We omit the details here, but refer the reader to the code comments in module glissade_velo_higher_pcg.F90.

SLAP. The Sparse Linear Algebra Package (SLAP) is a set of Fortran routines for solving sparse systems of linear equations. SLAP was part of the original Glimmer release and is still used to solve the thickness evolution and temperature advection problems in Glide. It can also be used to solve higher-order systems in Glissade, using either GMRES or a biconjugate gradient solver. SLAP, however, is limited to a single processor and thus is unsuited for large problems.

In order for Glissade to use CISM's SLAP routines, it must convert the global matrices and the right-hand-side and solution vectors to a SLAP-friendly format. First (before any matrix assembly), Glissade assigns a unique integer ID to each active node. Following assembly, the code loops through the matrices $\mathbf{A_{uv}}$, $\mathbf{A_{vv}}$, $\mathbf{A_{vu}}$ and $\mathbf{A_{vv}}$, putting each nonzero entry into a Fortran derived type containing the row, column and value of such entry. Similarly, the current solution and the right-hand side are placed in SLAP-friendly vectors. This information is passed to the SLAP solver and ultimately to the various SLAP subroutines. SLAP returns the velocity solution and residual vectors, which are copied back into Glissade data structures.

Trilinos. Trilinos is a set of solver packages and related software developed at Sandia National Laboratories. Under the DOE ISICLES project, the Glam higher-order dycore was parallelized and linked to Trilinos. The Trilinos links developed for Glam were later adapted for Glissade. Trilinos has been tested extensively on parallel architectures and has been installed on DOE high-performance computers where CISM is often run. See Section 2.4.2 for instructions on building CISM with Trilinos.

A C++ file, libglimmer-trilinos/trilinosGlissadeSolver.cpp, contains procedures that link Glissade to Trilinos:

- initializetgs sends Trilinos a global ID for each active node.
- insertrowtgs sends Trilinos a row of the matrix (specifically, the global row index, the number of potentially nonzero column entries, the index and value of each such entry, and the associated right-hand side term).
- solvevelocitytgs returns the velocity solution.

Various Glissade subroutines gather the information needed by Trilinos. Before assembly, Glissade computes global IDs for each node (k,i,j) on each processor, along with unique indices for each unknown (u and v) on each active node. Glissade also computes a logical array of dimension (27, nz, nx-1, ny-1)—the same dimension as the global matrices—whose value is .true. for the potential nonzero entries. After assembly of $\mathbf{A_{uu}}$, $\mathbf{b_{u}}$, etc., the nonzero matrix entries are passed to Trilinos, one row at a time. After solution, the Trilinos velocity result is copied to Glissade data structures.

Trilinos solver options are set in a file called trilinosOptions.xml in the directory where the code is executed. Among the key settings are the solver type and preconditioner type. The default solver is Block GMRES, and the default preconditioner is Ifpack, which applies ILU preconditioning. ILU generally works well for problems where shallow-ice physics is dominant, but can struggle for shelf-type problems where membrane stresses are dominant. For these problems the ML preconditioner type, which uses a multigrid method for preconditioning, is

likely to work better. Optimal Trilinos settings for large, complex problems are an area of active research.

Choosing among the various solver options is something of an art. The default is native PCG using the Chronopoulos-Gear algorithm. The PCG solver runs efficiently on most platforms and scales well. On the other hand, its preconditioning options are limited, so convergence may be slow for problems with dominant membrane stresses. SLAP solvers are generally robust and efficient, but are limited to one processor. Trilinos is typically slower than the other solvers, in part because of the extra cost of setting up Trilinos data structures, as well as the slower performance of C++ code relative to Fortran on many platforms. On the other hand, Trilinos contains a vast range of solver options (only a few of which have so far been tested with CISM), some of which may be more flexible and robust than those implemented in SLAP and the native PCG solver.

Solving the nonlinear system

Each call to a linear solver (native PCG, SLAP, or Trilinos) returns a solution vector, along with the number of iterations and the error, defined as $\sqrt{(r,r)}/\sqrt{(b,b)}$. If the linear solution fails to converge after the maximum allowed number of iterations (typically 200), the solver exits with the last computed solution (which may be adequate despite being unconverged).

Then a new global matrix is assembled, using the latest velocity solution to compute the effective viscosity. The right-hand is adjusted as needed to incorporate Dirichlet boundary conditions. Glissade then computes the new residual b - Ax. If the L_2 norm of the residual (defined as $\sqrt{(r,r)}$) is smaller than a desired threshold (10^{-4} by default), the nonlinear system of equations is considered solved. (An absolute threshold of 10^{-4} may be too stringent for large problems. Alternatively, Glissade can use a relative threshold, based on the ratio of the L2 norm of the residual to the L2 norm of b; see Section 7.3.) Otherwise the linear solver is called again, until either the solution converges or the maximum number of nonlinear iterations (typically 100) is reached.

As mentioned above, the procedure of updating the matrix with the latest guess for the solution is known as Picard iteration. The older Glam dycore includes an option to solve the nonlinear system using a Jacobian-Free Newton-Krylov (JFNK) method. JFNK requires an extra residual evaluation per nonlinear iteration, but generally converges in fewer iterations than does the Picard method. Glissade does not yet have a JFNK option.

In addition to the 3D Blatter-Pattyn approximation, Glissade can solve the simpler shallow-ice and shallow-shelf approximations, as well as a depth-integrated higher-order approximation (DIVA; Goldberg (2011)). These are described next.

6.2.2 Shallow-ice approximation

SIA: matrix form

By setting which ho_approx = 0 in the config file, Glissade's finite-element solver can be used as an SIA solver. The shallow-ice equations follow from the Blatter-Pattyn equations if the horizontal-plane stress terms are neglected. The SIA analogs of (6.1) are

$$x: \frac{\partial}{\partial z} \left(\eta \frac{\partial u}{\partial z} \right) = \rho g \frac{\partial s}{\partial x},$$

$$y: \frac{\partial}{\partial z} \left(\eta \frac{\partial v}{\partial z} \right) = \rho g \frac{\partial s}{\partial y},$$
(6.47)

leading to the following internal stress terms in weak form:

$$x: \int_{\Omega} 2\eta \frac{\partial \varphi_i}{\partial z} \left(\frac{1}{2} \left[\frac{\partial \varphi_j}{\partial z} \right] \{ u_j \} \right),$$

$$y: \int_{\Omega} 2\eta \frac{\partial \varphi_i}{\partial z} \left(\frac{1}{2} \left[\frac{\partial \varphi_j}{\partial z} \right] \{ v_j \} \right).$$
(6.48)

These terms are integrated over elements using (6.18) and (6.21) for the 3D basis functions and Jacobians. The resulting 8x8 element matrices $\mathbf{K_{uu}}$ and $\mathbf{K_{vv}}$ link each node to the eight nodes of the element, including itself. Since there are no horizontal stress terms, $\mathbf{K_{uv}} = \mathbf{K_{vu}} = 0$. In the expression for effective viscosity, the effective strain rate is given by

$$\dot{\varepsilon}_e^2 = \dot{\varepsilon}_{xz}^2 + \dot{\varepsilon}_{yz}^2. \tag{6.49}$$

The basal and lateral boundary conditions and gravitational loading are handled just as for the Blatter-Pattyn case.

The resulting SIA global matrices contain about half as many nonzero terms as the BP matrices, since $A_{uv} = A_{vu} = 0$. Each ice column, however, cannot be solved independently of the others; rather, each node is linked to its horizontal neighbors by terms that arise during element assembly. As a result, the matrix-based SIA solver is not dramatically faster than the BP solver. This solver can be useful for code testing but is not practical for productions runs, since a practical SIA solver should be many times faster than a BP solver. Glissade has a much faster alternative SIA solver, described next.

SIA: Local form

Glissade's local shallow-ice solver (in glissade_velo_sia.F90), is distinct from the finite-element solvers described in this section. It is local in the sense that u and v in each ice column are found independently of u and v in all other columns. It resembles the Glide shallow-ice solver described in Section 4.1. Glide, however, incorporates the velocity solution in a diffusion equation for ice thickness. Glissade's local SIA solver computes u and v only, with thickness evolving separately as described in Section 6.4.

The local SIA solver first computes the basal velocity given the basal traction coefficient t_b , which is defined as in Glide:

$$\mathbf{u_b} = t_b \tau_{\mathbf{b}}.\tag{6.50}$$

Note that $t_b = 1/\beta$, where the higher-order traction coefficient β is defined as in (6.26). There are several options for setting t_b : no sliding ($t_b = 0$); uniform traction ($t_b = \text{constant}$); uniform traction where basal water is present (and no sliding elsewhere); and uniform traction where the bed is at the pressure melting point, $T_b = T_{pmp}$ (and no sliding elsewhere). As in Glide, the basal velocity is proportional to t_b and the gravitational driving stress:

$$\mathbf{u_b} = -t_b \rho_i g \bar{H} \nabla s, \tag{6.51}$$

where \bar{H} is the ice thickness interpolated to cell vertices and ∇s is the surface slope at vertices. To find the interior velocities, we first compute a vertically integrated factor $c(\sigma)$ for each level at each cell vertex:

$$c(\sigma) = -2(\rho g)^n \bar{H}^{n+1} |\nabla s|^{n-1} \int_{1}^{\sigma} \bar{A}\tilde{\sigma}^n d\tilde{\sigma}, \qquad (6.52)$$

where \bar{A} is the flow factor interpolated to vertices, and a tilde distinguishes σ at a particular level from the integration variable $\tilde{\sigma}$. (The same term appears in the Glide SIA calculation; see (4.22).) Discretized in the vertical, this becomes

$$c(k) = -2(\rho g)^n \bar{H}^{n+1} |\nabla s|^{n-1} \sum_{l=n,r-1}^k \bar{A}_l \left(\frac{\sigma_l + \sigma_{l+1}}{2}\right)^n (\sigma_{l+1} - \sigma_l).$$
 (6.53)

The factors c(k) are interpolated from cell vertices to edges. The u and v components of velocity are computed at cell edges as a function of the surface slope:

$$\Delta u_E(k,i,j) = \sum_{l=nz-1}^k \left(\frac{c(k,i,j) + c(k,i,j-1)}{2} \right) (s(i+1,j) - s(i,j)),$$

$$\Delta v_N(k,i,j) = \sum_{l=nz-1}^k \left(\frac{c(k,i,j) + c(k,i-1,j)}{2} \right) (s(i,j+1) - s(i,j)).$$
(6.54)

Here, Δu_E is the *u* velocity component on the east edge of a cell, relative to the basal velocity, and similarly for Δv_N on the north edge of the cell.

Finally, we average Δu_E and Δv_N to vertices and add the bed velocities (6.51) to obtain the velocity in the ice column at each vertex:

$$u(k, i, j) = u_b(i, j) + \left(\frac{\Delta u_E(k, i, j) + \Delta u_E(k, i, j+1)}{2}\right),$$

$$v(k, i, j) = v_b(i, j) + \left(\frac{\Delta v_N(k, i, j) + \Delta v_N(k, i+1, j)}{2}\right).$$
(6.55)

For serial problems there are no special advantages to using Glissade's local SIA solver in place of Glide. Glissade's local SIA solver, however, can run on multiple processors, whereas Glide cannot. For large SIA problems, parallel Glissade can hold more data in memory and may be faster.

6.2.3 Shallow-shelf approximation

Glissade's finite-element solver can also be used as a shallow-shelf solver. The SSA equations can be derived by vertically integrating the 3D Blatter-Pattyn equations. They are valid when the basal shear stress is small or zero and the velocity is (to a good approximation) vertically uniform. The shallow-shelf analog of (6.1) is

$$x: \frac{\partial}{\partial x} \left(2\bar{\eta}H \left(2\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) \right) + \frac{\partial}{\partial y} \left(\bar{\eta}H \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right) = \rho g H \frac{\partial s}{\partial x},$$

$$y: \frac{\partial}{\partial y} \left(2\bar{\eta}H \left(2\frac{\partial v}{\partial y} + \frac{\partial u}{\partial x} \right) \right) + \frac{\partial}{\partial x} \left(\bar{\eta}H \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right) = \rho g H \frac{\partial s}{\partial y},$$

$$(6.56)$$

where $\bar{\eta}$ is the vertically averaged effective viscosity. The SSA equations in weak form resemble (6.11), but with internal stress terms

$$x: \int_{\Omega} 2\bar{\eta} H\left(\frac{\partial \varphi_{i}}{\partial x} \left(2\left[\frac{\partial \varphi_{j}}{\partial x}\right] \left\{u_{j}\right\} + \left[\frac{\partial \varphi_{j}}{\partial y}\right] \left\{v_{j}\right\}\right) + \frac{\partial \varphi_{i}}{\partial y} \left(\frac{1}{2}\left[\frac{\partial \varphi_{j}}{\partial y}\right] \left\{u_{j}\right\} + \frac{1}{2}\left[\frac{\partial \varphi_{j}}{\partial x}\right] \left\{v_{j}\right\}\right)\right),$$

$$y: \int_{\Omega} 2\bar{\eta} H\left(\frac{\partial \varphi_{i}}{\partial y} \left(2\left[\frac{\partial \varphi_{j}}{\partial y}\right] \left\{v_{j}\right\} + \left[\frac{\partial \varphi_{j}}{\partial x}\right] \left\{u_{j}\right\}\right) + \frac{\partial \varphi_{i}}{\partial x} \left(\frac{1}{2}\left[\frac{\partial \varphi_{j}}{\partial x}\right] \left\{v_{j}\right\} + \frac{1}{2}\left[\frac{\partial \varphi_{j}}{\partial y}\right] \left\{u_{j}\right\}\right)\right).$$

$$(6.57)$$

The resulting element matrices are similar to (6.14)–(6.17), except that the terms containing $\frac{\partial \varphi_i}{\partial z}$ and $\frac{\partial \varphi_j}{\partial z}$ are missing, and the viscosity term η is replaced by $\bar{\eta}H$. The effective viscosity is defined as in (6.2), but with a vertically averaged flow factor and with (6.3) replaced by

$$\dot{\varepsilon}_e^2 = \dot{\varepsilon}_{xx}^2 + \dot{\varepsilon}_{yy}^2 + \dot{\varepsilon}_{xx}\dot{\varepsilon}_{yy} + \dot{\varepsilon}_{xy}^2. \tag{6.58}$$

The integrals in (6.57) are taken over 2D cells rather than 3D elements, with basis functions and Jacobians given by (6.28) and (6.29).

The basal boundary terms are handled as in the 3D Blatter-Pattyn approximation. The lateral boundary and gravitational forcing terms are computed initially in 3D (as for the Blatter-Pattyn case), but then are summed in the vertical before being inserted into the 2D right-hand vectors $\mathbf{b_u}$ and $\mathbf{b_v}$. The solution procedure is the same as for the 3D case, except that the problem has no vertical dimension. When using the native PCG solver, the shallow-ice-based preconditioner is inappropriate and the diagonal preconditioner should be used instead.

In short, the SSA problem is similar to the Blatter-Pattyn problem except for the missing vertical shear terms and the reduction to 2D. The solution is much faster, because the SSA matrices have roughly 3*nz times fewer nonzero entries than the Blatter-Pattyn matrices. The factor of 3 arises from the fact that the BP equations at each level include connections to the levels above and below, whereas the SSA equations are solved at a single level.

6.2.4 Depth-integrated-viscosity approximation

Goldberg (2011) derived a higher-order stress approximation that is similar in accuracy to Blatter-Pattyn but much less expensive to solve, because (as with the SSA) the matrix system is assembled and solved in two dimensions. Since the stress balance equations use a depth-integrated effective viscosity in place of a vertically varying viscosity, we refer to this scheme as the depth-integrated-viscosity approximation, or DIVA. Here we summarize the DIVA stress balance and describe Glissade's solution method.

Dukowicz et al. (2010) showed that the Blatter-Pattyn equations (6.1) and associated boundary conditions can be derived by taking the first variation of a functional and setting it to zero. The functional includes terms that depend on the effective strain rate (6.3), which can be written as

$$\dot{\varepsilon}_{BP}^2 = u_x^2 + v_y^2 + u_x v_y + \frac{1}{4} (u_y + v_x)^2 + \frac{1}{4} u_z^2 + \frac{1}{4} v_z^2.$$
 (6.59)

Goldberg (2011) derived an alternate set of equations using a similar functional, but with the horizontal velocity gradients in the effective strain rate replaced by their vertical averages:

$$\dot{\varepsilon}_{\text{DIVA}}^2 = \bar{u}_x^2 + \bar{v}_y^2 + \bar{u}_x \bar{v}_y + \frac{1}{4} (\bar{u}_y + \bar{v}_x)^2 + \frac{1}{4} u_z^2 + \frac{1}{4} v_z^2, \tag{6.60}$$

where

$$\bar{u} = \frac{1}{H} \int_{b}^{s} u(z)dz, \quad \bar{v} = \frac{1}{H} \int_{b}^{s} v(z)dz.$$
 (6.61)

The resulting equations of motion, analogous to (6.1), are

$$x: \frac{1}{H} \frac{\partial}{\partial x} \left(2H\bar{\eta} \left(2\frac{\partial \bar{u}}{\partial x} + \frac{\partial \bar{v}}{\partial y} \right) \right) + \frac{1}{H} \frac{\partial}{\partial y} \left(H\bar{\eta} \left(\frac{\partial \bar{u}}{\partial y} + \frac{\partial \bar{v}}{\partial x} \right) \right) + \frac{\partial}{\partial z} \left(\eta \frac{\partial u}{\partial z} \right) = \rho g \frac{\partial s}{\partial x},$$

$$y: \frac{1}{H} \frac{\partial}{\partial y} \left(2H\bar{\eta} \left(2\frac{\partial \bar{v}}{\partial y} + \frac{\partial \bar{u}}{\partial x} \right) \right) + \frac{1}{H} \frac{\partial}{\partial x} \left(H\bar{\eta} \left(\frac{\partial \bar{u}}{\partial y} + \frac{\partial \bar{v}}{\partial x} \right) \right) + \frac{\partial}{\partial z} \left(\eta \frac{\partial v}{\partial z} \right) = \rho g \frac{\partial s}{\partial y},$$

$$(6.62)$$

where the depth-integrated effective viscosity is given by

$$\bar{\eta} = \frac{1}{H} \int_{b}^{s} \eta(z) dz. \tag{6.63}$$

Since the horizontal stress terms are depth-independent, (6.62) can be integrated in the vertical to give

$$x: \frac{\partial}{\partial x} \left(2H\bar{\eta} \left(2\frac{\partial \bar{u}}{\partial x} + \frac{\partial \bar{v}}{\partial y} \right) \right) + \frac{\partial}{\partial y} \left(H\bar{\eta} \left(\frac{\partial \bar{u}}{\partial y} + \frac{\partial \bar{v}}{\partial x} \right) \right) - \beta u_b = \rho g H \frac{\partial s}{\partial x},$$

$$y: \frac{\partial}{\partial y} \left(2H\bar{\eta} \left(2\frac{\partial \bar{v}}{\partial y} + \frac{\partial \bar{u}}{\partial x} \right) \right) + \frac{\partial}{\partial x} \left(H\bar{\eta} \left(\frac{\partial \bar{u}}{\partial y} + \frac{\partial \bar{v}}{\partial x} \right) \right) - \beta v_b = \rho g H \frac{\partial s}{\partial y},$$

$$(6.64)$$

where the boundary conditions at b and s have been used to evaluate the vertical stress terms, with a sliding law of the form (6.26).

Note that (6.64) has the same form as the SSA stress balance (6.56), but with the addition of basal stress terms. Thus the methods used to assemble and solve the SSA equations can also be applied to the DIVA equations, to obtain the mean velocity components \bar{u} and \bar{v} . The 2D DIVA system can be solved much more cheaply than the 3D BP system.

In order to solve (6.64), however, the basal stress terms must be rewritten in terms of \bar{u} and \bar{v} . This is done as follows (considering only the x component of velocity; the y component is analogous). The x component of (6.64) can be rearranged and divided by H:

$$\frac{\beta u_b}{H} = \frac{1}{H} \frac{\partial}{\partial x} \left(2H\bar{\eta} \left(2\frac{\partial \bar{u}}{\partial x} + \frac{\partial \bar{v}}{\partial y} \right) \right) + \frac{1}{H} \frac{\partial}{\partial y} \left(H\bar{\eta} \left(\frac{\partial \bar{u}}{\partial y} + \frac{\partial \bar{v}}{\partial x} \right) \right) - \rho g \frac{\partial s}{\partial x}. \tag{6.65}$$

From (6.62), the RHS is just $-(\eta u_z)_z$, giving

$$\frac{\beta u_b}{H} = -\frac{\partial}{\partial z} \left(\eta \frac{\partial u}{\partial z} \right). \tag{6.66}$$

Integrating (6.66) from z to the surface s gives

$$\eta \frac{\partial u}{\partial z} = \frac{\beta u_b \left(s - z\right)}{H}.\tag{6.67}$$

Dividing by η and integrating from b to z gives u(z) in terms of u_b :

$$u(z) = u_b + \frac{\beta u_b}{H} \int_{b}^{z} \frac{(s - z')}{\eta(z')} dz'.$$
 (6.68)

Finally, integrating u(z) from the bed to the surface gives the mean velocity \bar{u} in terms of the bed velocity:

$$\bar{u} = u_b \left(1 + \frac{\beta \omega}{H} \right), \tag{6.69}$$

where

$$\omega \equiv \frac{1}{H} \int_{b}^{s} \int_{b}^{z} \frac{(s-z')}{\eta(z')} dz' dz. \tag{6.70}$$

We can think of ω/H as a depth-integrated inverse viscosity. The less viscous the ice, the greater the value of ω/H , and the greater the difference between u(z) and u_b .

Given (6.69), we can replace βu_b and βv_b in (6.64) with $\beta_{\text{eff}}\bar{u}$ and $\beta_{\text{eff}}\bar{v}$, respectively, where

$$\beta_{\text{eff}} = \frac{\beta}{1 + \frac{\beta \omega}{H}}.\tag{6.71}$$

For a frozen bed $(u_b = v_b = 0$, with nonzero basal stress τ_{bx}), the βu_b term on the RHS of (6.68) is replaced by τ_{bx} , leading to

$$\bar{u} = \frac{\tau_{bx}\omega}{H}.\tag{6.72}$$

Then the basal stress terms in (6.64) can be replaced by $\beta_{\text{eff}}^{\text{frozen}}\bar{u}$ and by $\beta_{\text{eff}}^{\text{frozen}}\bar{v}$, where

$$\beta_{\text{eff}}^{\text{frozen}} = \frac{H}{\omega} \tag{6.73}$$

With these replacements, (6.64) is fully analogous to the SSA stress balance (6.56).

In order to compute $\bar{\eta}$, the effective viscosity $\eta(z)$ must be evaluated at each level. We use (6.2) with the effective strain rate (6.60). The horizontal terms in $\dot{\varepsilon}_{\text{DIVA}}^2$ are found using the mean velocities from the previous iteration. We compute the vertical shear terms using (6.67):

$$u_z(z) = \frac{\tau_{bx}(s-z)}{\eta(z)H}, \quad v_z(z) = \frac{\tau_{by}(s-z)}{\eta(z)H},$$
 (6.74)

where both $\eta(z)$ and τ_b are from the previous iteration.

The DIVA solution procedure can be summarized as follows:

- 1. Begin with the current guess for \bar{u} and \bar{v} .
- 2. Assemble the DIVA solution matrix in analogy to the SSA solution matrix, with the following DIVA-specific computations:
 - For each quadrature point, compute $\eta(z)$ using (6.2), (6.60) and (6.74).
 - Given $\eta(z)$, compute the depth-averaged $\bar{\eta}$.
 - Compute the vertical integrals (including ω) in (6.68) and (6.70).
 - Interpolate ω to vertices and compute β_{eff} .
- 3. Solve the matrix system for \bar{u} and \bar{v} .
- 4. Given (\bar{u}, \bar{v}) at each vertex, use (6.69) to find (u_b, v_b) , and use (6.68) to find u(z) at each level
- 5. Iterate to convergence.

DIVA is about as accurate as the BP equations in regions of fast sliding. Where there is little or no sliding, DIVA remains accurate provided that the horizontal gradients of the membrane stresses are uniform in the vertical. It is less accurate in regions of slow sliding and rough bed topography, where the horizontal gradients of the membrane stresses vary strongly with height. As shown by Goldberg (2011), DIVA is less accurate than BP for ISMIP-HOM Test A, especially for short wavelengths of bed topography.

6.2.5 L1L2 approximation

Glissade also includes the L1L2 approximation, which can be derived by vertically integrating the Blatter-Pattyn equations with the assumption that the horizontal gradients of the membrane stresses are uniform with depth. (The term "L1L2" is based on the classification scheme of Hindmarsh (2004).) This assumption leads to a 2D matrix system (like the SSA and DIVA systems), which can be solved much more cheaply than the 3D BP system. In regions of little or no sliding, L1L2 (unlike DIVA) reduces to the shallow-ice approximation. Thus it is accurate

where the SIA is accurate, but not in regions of slow sliding and rough bed topography. L1L2 performs poorly for ISMIP-HOM Test A, which has no sliding.

The L1L2 scheme in Glissade is based on Perego et al. (2012). Although formally of similar accuracy to DIVA, it has proven less accurate and robust in practice. For this reason it is not scientifically supported, and may be removed from future versions of CISM. It is described here for completeness.

Consider the Blatter-Pattyn equations (6.1), written in terms of deviatoric stresses instead of strain rates:

$$x: \frac{\partial}{\partial x} (2\tau_{xx} + \tau_{yy}) + \frac{\partial}{\partial y} (\tau_{xy}) + \frac{\partial}{\partial z} (\tau_{xz}) = \rho g \frac{\partial s}{\partial x},$$

$$y: \frac{\partial}{\partial y} (2\tau_{yy} + \tau_{xx}) + \frac{\partial}{\partial x} (\tau_{xy}) + \frac{\partial}{\partial z} (\tau_{yz}) = \rho g \frac{\partial s}{\partial y}.$$

$$(6.75)$$

Assuming that τ_{xx} , τ_{yy} and τ_{xy} are depth-independent, (6.75) can be integrated from b to s to obtain

$$x: \frac{\partial}{\partial x} \left(\int_{b}^{s} (2\tau_{xx} + \tau_{yy}) dz \right) + \frac{\partial}{\partial y} \left(\int_{b}^{s} \tau_{xy} dz \right) - \beta u_{b} = \rho g H \frac{\partial s}{\partial x},$$

$$y: \frac{\partial}{\partial y} \left(\int_{b}^{s} (2\tau_{yy} + \tau_{xx}) dz \right) + \frac{\partial}{\partial x} \left(\int_{b}^{s} \tau_{xy} dz \right) - \beta v_{b} = \rho g H \frac{\partial s}{\partial y},$$

$$(6.76)$$

where u_b and v_b are the components of basal velocity, and we have used Leibniz's rule to move the partial derivatives outside the integrals and eliminate several boundary terms. Since the deviatoric stresses are depth-independent, they can be written in terms of the basal strain rates. Taking these strain rates outside the integral, we obtain

$$x: \quad 2\frac{\partial}{\partial x}\left(\bar{\eta}H\left(2\frac{\partial u_b}{\partial x} + \frac{\partial v_b}{\partial y}\right)\right) + \frac{\partial}{\partial y}\left(\bar{\eta}H\left(\frac{\partial u_b}{\partial y} + \frac{\partial v_b}{\partial x}\right)\right) - \beta u_b = \rho gH\frac{\partial s}{\partial x},$$

$$y: \quad 2\frac{\partial}{\partial y}\left(\bar{\eta}H\left(2\frac{\partial v_b}{\partial y} + \frac{\partial u_b}{\partial x}\right)\right) + \frac{\partial}{\partial x}\left(\bar{\eta}H\left(\frac{\partial u_b}{\partial y} + \frac{\partial v_b}{\partial x}\right)\right) - \beta v_b = \rho gH\frac{\partial s}{\partial y},$$

$$(6.77)$$

where the vertically averaged effective viscosity is

$$\bar{\eta} = \frac{1}{H} \int_{1}^{s} \eta dz. \tag{6.78}$$

This stress balance closely resembles the SSA stress balance (6.56) and the DIVA stress balance (6.62).

The effective viscosity, however, is treated differently for L1L2. (Here we follow Perego et al. (2012).) The parallel norm $|\cdot|_{||}$ is defined as

$$\left|\dot{\varepsilon}\right|_{\parallel}^{2} = \dot{\varepsilon}_{xx}^{2} + \dot{\varepsilon}_{yy}^{2} + \dot{\varepsilon}_{xx}\dot{\varepsilon}_{yy} + \dot{\varepsilon}_{xy}^{2} \tag{6.79}$$

and the perpendicular norm $|\cdot|_{\perp}$ as

$$|\dot{\varepsilon}|_{\perp}^2 = \dot{\varepsilon}_{xz}^2 + \dot{\varepsilon}_{yz}^2.$$
 (6.80)

The constitutive law can be written as

$$\dot{\varepsilon}_{ij} = A \tau_e^2 \tau_{ij},\tag{6.81}$$

where we assume n=3. For the L1L2 approximation the effective stress is given by

$$\tau_e^2 = |\tau|_{\parallel}^2 + \tilde{\tau}_{\perp}^2, \tag{6.82}$$

with $\tilde{\tau}_{\perp}$ estimated based on the SIA:

$$\tilde{\tau}_{\perp}^2 = \left[\rho g(s-z)\nabla s\right]^2. \tag{6.83}$$

Inserting (6.82) in (6.81) and taking the parallel norm gives

$$|\dot{\varepsilon}|_{\parallel} = A(|\tau|_{\parallel}^2 + |\tau|_{\perp}^2)|\tau|_{\parallel}.$$
 (6.84)

Given $|\dot{\varepsilon}|_{\parallel}$ from (6.79), equation (6.84) can be written in the form $x^3 + ax + b = 0$, a cubic equation that can be solved for $|\tau|_{\parallel}$ using standard techniques. (Of the three roots, the one of interest is real and positive; the other two are complex conjugates. Glissade computes $|\tau|_{\parallel}$ only for n=3.) The effective viscosity at depth z is then given by

$$\eta = \frac{1}{2A\tau_e^2},\tag{6.85}$$

where τ_e is found from (6.82), (6.83) and (6.84). Equation (6.85) follows from (6.81) and the relation

$$\tau_{ij} = 2\eta \dot{\varepsilon}_{ij},\tag{6.86}$$

which defines the effective viscosity.

After solving the 2D L1L2 system for the basal velocity, the velocity profile at each vertex can be found by vertical integration. First we estimate τ_{xz} and τ_{yz} in each layer. Since the depth-independent membrane stresses are determined by the basal velocity, τ_{xz} and τ_{yz} are now the only unknowns in (6.75). These equations can be integrated from the top surface (assumed to be stress-free) to depth z to give

$$x: \quad \tau_{xz}(z) = -\rho g(s-z) \frac{\partial s}{\partial x} + 2 \frac{\partial}{\partial x} \left[\left(2 \frac{\partial u_b}{\partial x} + \frac{\partial v_b}{\partial y} \right) \left(\int_z^h \eta dz' \right) \right] + 2 \frac{\partial}{\partial y} \left[\left(\frac{\partial u_b}{\partial y} + \frac{\partial v_b}{\partial x} \right) \left(\int_z^h \eta dz' \right) \right],$$

$$y: \quad \tau_{yz}(z) = -\rho g(s-z) \frac{\partial s}{\partial y} + 2 \frac{\partial}{\partial y} \left[\left(2 \frac{\partial v_b}{\partial y} + \frac{\partial u_b}{\partial x} \right) \left(\int_z^h \eta dz' \right) \right] + 2 \frac{\partial}{\partial x} \left[\left(\frac{\partial u_b}{\partial y} + \frac{\partial v_b}{\partial x} \right) \left(\int_z^h \eta dz' \right) \right],$$

$$(6.87)$$

where again we use Leibniz's rule to move the partial derivatives outside the integrals and remove boundary terms. The bracketed terms are evaluated at cell centers, and then the gradient terms are computed at nodes using finite-difference formulas (e.g., (6.43)). Given τ_{xz} and τ_{yz} at each level, along with $|\dot{\varepsilon}|_{\parallel}^2$ from (6.79), the effective stress is given by

$$\tau_e^2 = 2\eta \, |\dot{\varepsilon}|_{\parallel}^2 + \tau_{xz}^2 + \tau_{yz}^2. \tag{6.88}$$

The velocity components can then be integrated upward from the bed using

$$x: \frac{1}{2} \frac{\partial u}{\partial z} = \dot{\varepsilon}_{xz} = A \tau_e^{n-1} \tau_{xz},$$

$$y: \frac{1}{2} \frac{\partial v}{\partial z} = \dot{\varepsilon}_{yz} = A \tau_e^{n-1} \tau_{yz},$$
(6.89)

which imply

$$x: \quad u(z) = u_b + 2 \int_b^z A \tau_e^{n-1} \tau_{xz} dz,$$

$$y: \quad v(z) = v_b + 2 \int_b^z A \tau_e^{n-1} \tau_{yz} dz.$$
(6.90)

6.3 Temperature Solver

As discussed in Section 4.2, the thermal evolution of the ice sheet is given by

$$\frac{\partial T}{\partial t} = \frac{k}{\rho c} \nabla^2 T - \mathbf{u} \cdot \nabla T - w \frac{\partial T}{\partial z} + \frac{\Phi}{\rho c}, \tag{6.91}$$

where T is the temperature in ${}^{\circ}$ C, k is the thermal conductivity of ice, c is the specific heat of ice, ρ is the density, and Φ is the rate of heating due to internal deformation and dissipation. This equation describes the conservation of internal energy under horizontal and vertical diffusion (the first term on the RHS), horizontal and vertical advection (the second and third terms, respectively), and internal heat dissipation (the last term). Glide solves this equation in module ${\tt glide_temp.F90}$. Glissade takes a different approach, dividing the temperature evolution into separate advection and diffusion/dissipation components. Module ${\tt glissade_temp.F90}$ solves for diffusion and internal dissipation:

$$\frac{\partial T}{\partial t} = \frac{k}{\rho c} \nabla^2 T + \frac{\Phi}{\rho c},\tag{6.92}$$

as described in this section. The advective part of (6.91) is described in Section 6.4.

Glissade's vertical discretization of temperature is also different from that of Glide. In Glide, T is located at each of nz vertical levels. In Glissade, internal temperatures are located at the midpoints of the nz-1 layers. Temperature is also defined at the upper and lower ice surface, giving a total of nz+1 temperature points in each column. The upper surface temperature is denoted by T_0 and the lower surface temperature by T_{nz} .

This convention makes it fairly straightforward to advect temperature conservatively. The total internal energy in a column is the sum over layers of $\rho cT\Delta z$, where Δz is the layer thickness. This internal energy is conserved under transport (see Section 6.4). T_0 and T_{nz} , which are determined by the boundary conditions, are associated with infinitesimally thin layers that do not contain any internal energy.

The following sections describe how the terms in (6.92) are computed, how the boundary conditions are specified, and how the equation is solved.

6.3.1 Vertical diffusion

Computing the vertical diffusion term requires a discretization for $\nabla^2 T$. As in Glide (Section 4.2.1), we assume that horizontal diffusion is negligible compared to vertical diffusion:

$$\nabla^2 T \simeq \frac{\partial^2 T}{\partial z^2} = \frac{1}{H^2} \frac{\partial^2 T}{\partial \sigma^2}, \tag{6.93}$$

where the last equality follows from $\sigma = (s - z)/H$.

In σ -coordinates, the central difference formulas for first derivatives at the upper and lower interfaces of layer k are

$$\frac{\partial T}{\partial \sigma} \Big|_{\sigma_k} = \frac{T_k - T_{k-1}}{\tilde{\sigma}_k - \tilde{\sigma}_{k-1}},
\frac{\partial T}{\partial \sigma} \Big|_{\sigma_{k+1}} = \frac{T_{k+1} - T_k}{\tilde{\sigma}_{k+1} - \tilde{\sigma}_k},$$
(6.94)

where $\tilde{\sigma}_k$ is the value of σ at the midpoint of layer k, halfway between σ_k and σ_{k+1} :

$$\tilde{\sigma}_k = \frac{\sigma_{k+1} - \sigma_k}{2}.\tag{6.95}$$

The second partial derivative, defined at the midpoint of layer k, is

$$\frac{\partial^2 T}{\partial \sigma^2} \Big|_{\tilde{\sigma}_k} = \frac{\frac{\partial T}{\partial \sigma} \Big|_{\sigma_{k+1}} - \frac{\partial T}{\partial \sigma} \Big|_{\sigma_k}}{\sigma_{k+1} - \sigma_k}$$
(6.96)

Inserting (6.94) in (6.96), we obtain the required vertical diffusion term:

$$\frac{\partial^{2} T}{\partial \sigma^{2}}\Big|_{\tilde{\sigma}_{k}} = \frac{T_{k-1}}{\left(\tilde{\sigma}_{k} - \tilde{\sigma}_{k-1}\right)\left(\sigma_{k+1} - \sigma_{k}\right)} - T_{k} \left(\frac{1}{\left(\tilde{\sigma}_{k} - \tilde{\sigma}_{k-1}\right)\left(\sigma_{k+1} - \sigma_{k}\right)} + \frac{1}{\left(\tilde{\sigma}_{k+1} - \tilde{\sigma}_{k}\right)\left(\sigma_{k+1} - \sigma_{k}\right)}\right) + \frac{T_{k+1}}{\left(\tilde{\sigma}_{k+1} - \tilde{\sigma}_{k}\right)\left(\sigma_{k+1} - \sigma_{k}\right)}. \quad (6.97)$$

6.3.2 Heat dissipation

In higher-order models the internal heating rate Φ in (6.92) is given by the tensor product of strain rate and stress:

$$\Phi = \dot{\varepsilon}_{ij} \tau_{ij}. \tag{6.98}$$

The effective strain rate and effective stress (cf. (6.3)) are defined by

$$\dot{\varepsilon}_e^2 = \frac{1}{2}\dot{\varepsilon}_{ij}\dot{\varepsilon}_{ij}, \quad \tau_e^2 = \frac{1}{2}\tau_{ij}\tau_{ij}.$$
 (6.99)

It follows from (6.98) and (6.99) that

$$\Phi = 2\dot{\varepsilon}_e \tau_e. \tag{6.100}$$

Eq. (6.86), which defines the effective viscosity, implies

$$\dot{\varepsilon}_e = \frac{\tau_e}{2n},\tag{6.101}$$

which can be substituted in (6.100) to give

$$\Phi = \frac{\tau_e^2}{\eta}.\tag{6.102}$$

Both terms on the RHS of (6.102) are available to the temperature solver, since the higher-order velocity solver computes η during matrix assembly and diagnoses τ_e from η and $\dot{\varepsilon}_{ij}$ at the end of the calculation.

6.3.3 Boundary conditions

The temperature T_0 at the upper boundary is set to $\min(T_{\text{air}}, 0)$, where the surface air temperature T_{air} is usually specified from data or passed from a climate model. The diffusive heat flux at the upper boundary (defined as positive up) is

$$F_d^{\text{top}} = \frac{k}{H} \frac{T_1 - T_0}{\tilde{\sigma}_1}.\tag{6.103}$$

(The denominator contains just one term because $\sigma_0 = 0$.) Optionally, this flux can be returned to the climate model as a lower boundary condition in the land-surface model.

The lower ice boundary is more complex. For grounded ice there are three heat sources and sinks. First, the diffusive flux from the bottom surface to the ice interior (positive up) is

$$F_d^{\text{bot}} = \frac{k}{H} \frac{T_{nz} - T_{nz-1}}{1 - \tilde{\sigma}_{nz-1}}.$$
 (6.104)

Second, there is a geothermal heat flux F_g to the lower boundary. This is typically prescribed as a constant ($\sim 0.05~{\rm W~m^{-2}}$) or read from an input file. Finally, there is a frictional heat flux associated with basal sliding, given by (Cuffey and Paterson, 2010, p. 418)

$$F_f = \tau_{\mathbf{b}} \cdot \mathbf{u_b},\tag{6.105}$$

where $\tau_{\mathbf{b}}$ and $\mathbf{u}_{\mathbf{b}}$ are 2D vectors of basal shear stress and basal velocity, respectively. With a friction law of the form (6.26), this becomes

$$F_f = \beta(u_b^2 + v_b^2). \tag{6.106}$$

If the basal temperature $T_{nz} < T_{\text{pmp}}$ (where $T_{\text{pmp}} = T + 8.7 \cdot 10^{-4} (s - z)$ is the pressure melting point temperature), then the fluxes at the lower boundary must balance:

$$F_g + F_f = F_d^{\text{bot}}. ag{6.107}$$

In other words, the energy supplied by geothermal heating and sliding friction is equal to the energy removed by vertical diffusion. If, on the other hand, $T_{nz} = T_{pmp}$, then the lower surface temperature is fixed and the net flux is used to melt or freeze ice at the boundary:

$$M_b = \frac{F_g + F_f - F_d^{\text{bot}}}{\rho L},\tag{6.108}$$

where M_b is the melt rate and L is the latent heat of melting. Melting generates basal water, which may either stay in place, flow downstream (possibly replaced by water from upstream), or simply disappear from the system, depending on the parameterization chosen in the config file (basal_water; see 7.3). We hold $T_{nz} = T_{\rm pmp}$ as long as basal water is present.

For floating ice the basal boundary condition is simpler; T_{nz} is simply set to the freezing temperature T_f of seawater. Optionally, a melt rate can be prescribed at the lower surface, as in the MISOMIP experiments (Asay-Davis et al., 2016).

6.3.4 Vertical temperature solution

Eq. (6.92) can be discretized for an ice layer k as

$$\frac{T_k^{n+1} - T_k^n}{\Delta t} = \frac{k}{\rho c H^2} \left[a_k T_{k-1}^{n+1} - (a_k + b_k) T_k^{n+1} + b_k T_{k+1}^{n+1} \right] + \frac{\Phi_k}{\rho c},\tag{6.109}$$

where the coefficients a_k and b_k are given by (6.97), n is the current time level, and n+1 is the new time level. The vertical diffusion terms are evaluated at the new time level, making the discretization backward Euler (i.e., fully implicit) in time. A Crank-Nicolson formulation, in

which the temperature terms are evaluated at time n+1/2, is also available. Although Crank–Nicolson is second-order-accurate in time (compared to first-order for backward Euler), it can lead to temperature oscillations in thin ice. For this reason, backward Euler is the default.

Eq. (6.109) can be rewritten as

$$-\alpha_k T_{k-1}^{n+1} + (1 + \alpha_k + \beta_k) T_k^{n+1} - \beta_k T_{k+1}^{n+1} = T_k^n + \frac{\Phi_k \Delta t}{\rho c}, \tag{6.110}$$

where

$$\alpha_k = \frac{ka_k \Delta t}{\rho c H^2}, \quad \beta_k = \frac{kb_k \Delta t}{\rho c H^2}.$$
 (6.111)

At the upper surface, $T_0 = T_{\text{air}}$. At the lower surface we have either a temperature boundary condition ($T_{nz} = T_{\text{pmp}}$ for grounded ice, or $T_{nz} = T_f$ for floating ice) or a flux boundary condition:

$$F_f + F_g - \frac{k}{H} \frac{T_{nz}^{n+1} - T_{nz-1}^{n+1}}{1 - \tilde{\sigma}_{nz-1}} = 0, \tag{6.112}$$

which can be rearranged to give

$$-T_{nz-1}^{n+1} + T_{nz}^{n+1} = \frac{(F_f + F_g)H(1 - \tilde{\sigma}_{nz-1})}{k}.$$
 (6.113)

In each ice column the above equations form a tridiagonal system that is easily solved for T_k .

Occasionally, the solution T_k in one or more layers will exceed $T_{\rm pmp}$ for the layer. If so, we set $T_k = T_{\rm pmp}$ and use the extra energy to melt ice internally. This melt is assumed (not very realistically) to drain immediately to the bed.

If (6.108) applies, we compute M_b and adjust the basal water depth. When the basal water goes to zero, T_{nz} is set to the temperature of the lowest layer (below T_{pmp}), so that the flux boundary condition will apply during the next time step.

6.3.5 Enthalpy model

CISM includes an alternative vertical temperature solver based on enthalpy. In this scheme, water that melts internally can be retained in the ice instead of draining instantly to the bed. The retained water reduces the ice viscosity compared to that of pure ice. Since this scheme has not yet been scientifically validated, we defer discussion for a future version of the documentation.

6.4 Mass and Tracer Transport

Ice sheet models must solve a transport equation for ice thickness H:

$$\frac{\partial H}{\partial t} + \nabla \cdot (H\mathbf{U}) = B,\tag{6.114}$$

where U is the vertically averaged 2D velocity and B is the total surface mass balance. Eq. (6.114) describes the conservation of ice volume under horizontal transport. With the assumption of uniform density, volume conservation is equivalent to mass conservation. There is a similar conservation equation for the internal energy in each ice layer:

$$\frac{\partial(hT)}{\partial t} + \nabla \cdot (hT\mathbf{u}) = 0, \tag{6.115}$$

where h is the layer thickness, T is the temperature (defined at the layer midpoint), and \mathbf{u} is the horizontal velocity. If other tracers (e.g., internal water fraction or ice age) are present,

their transport is described by equations of the same form as (6.115). In the discussion below, everything said about temperature and internal energy applies also to other tracers and the associated conserved quantities.

Unlike Glide, which solves for mass transport and temperature advection independently, Glissade solves (6.114) and (6.115) in a coordinated way, one layer at a time. (Advection of tracers other than temperature is carried out in conjunction with temperature advection.) This coordination makes it possible to avoid numerical oscillations with undesirable effects, such as raising T above the melting point. The non-advective terms of the energy equation (6.91) (i.e., vertical diffusion and internal dissipation) are handled in a separate vertical column solve, as described in Section 6.3.

Glissade has two horizontal transport schemes: a first-order upwind scheme and a more accurate incremental remapping (IR) scheme (Dukowicz and Baumgardner, 2000; Lipscomb and Hunke, 2004). The transport driver is in glissade_transport.F90, and the IR scheme is in glissade_remap.F90. The IR scheme was originally implemented in the Los Alamos sea ice model, CICE, and has been adapted for CISM. It is fairly complex and is described in detail in Section 6.4.1.

After horizontal transport, the mass balance is applied at the top and bottom ice surfaces. Any energy that is available for melt after an ice column has entirely melted is discarded, but in coupled applications could be returned to the climate model if needed to conserve energy.

After the transport scheme and mass-balance update have been applied, the new layer thicknesses generally do not have the desired spacing in σ coordinates. A vertical remapping scheme is used to transfer ice volume, internal energy, and other conserved quantities between layers to restore σ layers in a way that conserves mass and energy. This scheme computes overlaps between the new layers and the target σ layers. If, for example, part of new layer k overlaps with target layer k+1, the mass and internal energy are computed for the overlap region, with the assumption that temperature is uniform within each layer. (A more accurate reconstruction of temperature may be added in the future.) Mass and internal energy are then transferred between layers, and the new layer temperatures are derived from the ratio of internal energy to mass.

6.4.1 Incremental remapping

Next we describe the incremental remapping scheme, which has several desirable features:

- It conserves the quantities being transported (mass and energy).
- It is non-oscillatory; that is, it does not create spurious ripples in the transported fields.
- It preserves tracer monotonicity. That is, it does not create new extrema in temperature; the values at time m + 1 are bounded by the values at time m.
- It is second-order accurate in space and therefore is less diffusive than first-order schemes. The accuracy may be reduced locally to first order to preserve monotonicity.
- It is efficient for large numbers of tracers. Much of the work is geometrical and is performed only once per cell edge instead of being repeated for each quantity being transported. (This is more of an issue for CICE than for CISM, but for some applications it might be desirable to carry additional tracers in CISM.)

The upwind scheme, like IR, is conservative, non-oscillatory, and monotonicity-preserving, but because it is first-order it is highly diffusive.

The IR time step is limited by the requirement that trajectories projected backward from grid cell corners are confined to the four surrounding cells; this is what is meant by incremental remapping as opposed to general remapping. This requirement leads to a Courant–Friedrichs–Lewy (CFL) condition,

$$\frac{\max|\mathbf{u}|\Delta t}{\Delta x} \le 1\tag{6.116}$$

For highly divergent velocity fields the maximum time step must be reduced by a factor of two to ensure that trajectories do not cross. However, ice-sheet velocity fields usually have small divergence per time step relative to the grid size.

The remapping algorithm can be summarized as follows:

- 1. Given mean values of the ice thickness and tracer fields in each grid cell, construct linear approximations of these fields. Limit the field gradients to preserve monotonicity.
- 2. Given ice velocities at grid cell vertices, identify departure regions for the fluxes across each cell edge. Divide these departure regions into triangles and compute the coordinates of the triangle vertices.
- 3. Integrate the thickness and tracer fields over the departure triangles to obtain the mass and energy transported across each cell edge.
- 4. Given these transports, update the state variables.

These steps are carried out for each of nz-1 ice layers, as described below.

Reconstructing area and tracer fields

First, using the known values of the state variables, the ice thickness and tracer fields are reconstructed in each grid cell as linear functions of x and y. For each field we compute the value at the cell center, along with gradients in the x and y directions. The gradients are limited to preserve monotonicity. When integrated over a grid cell, the reconstructed fields must have mean values equal to the known state variables, denoted by \bar{h} for ice thickness and \tilde{T} for temperature.

Consider first the ice thickness in a given layer. We contruct a field $h(\mathbf{r})$ whose mean is \bar{h} , where $\mathbf{r} = (x, y)$ is the position vector relative to the cell center. That is, we require

$$\int_{A} h \, dh = \bar{h} \, A,\tag{6.117}$$

where $A = \int_A dA$ is the grid cell area. Eq. (6.117) is satisfied if $h(\mathbf{r})$ has the form

$$h(\mathbf{r}) = \bar{h} + \alpha_h \langle \nabla h \rangle \cdot (\mathbf{r} - \bar{\mathbf{r}}), \tag{6.118}$$

where $\langle \nabla h \rangle$ is a centered estimate of the thickness gradient within the cell, α_h is a limiting coefficient that enforces monotonicity, and $\bar{\mathbf{r}}$ is the cell centroid in a local coordinate system whose origin lies at the cell center:

$$\bar{\mathbf{r}} = \frac{1}{A} \int_{A} \mathbf{r} \, dA.$$

On CISM's structured rectangular grid the cell center and centroid coincide, so that $\bar{\mathbf{r}} = 0$ and $h(\bar{\mathbf{r}}) = \bar{h}$.

Next consider the ice temperature field in a layer of thickness h. The reconstructed temperature must satisfy

$$\int_{A} h T dA = \bar{h} \tilde{T} A, \tag{6.119}$$

where $\tilde{T} = T(\tilde{\mathbf{r}})$ is the temperature at the center of ice mass. Eq. (6.119) is satisfied when $T(\mathbf{r})$ is given by

$$T(\mathbf{r}) = \tilde{T} + \alpha_T \langle \nabla T \rangle \cdot (\mathbf{r} - \tilde{\mathbf{r}}), \tag{6.120}$$

where $\langle \nabla T \rangle$ is a centered estimate of the thickness gradient, α_T is a limiting coefficient, and the center of ice mass $\tilde{\mathbf{r}}$ is given by

$$\tilde{\mathbf{r}} = \frac{1}{\bar{h} A} \int_{A} h \, \mathbf{r} \, dA. \tag{6.121}$$

Evaluating the integrals, we find that the components of $\tilde{\mathbf{r}}$ are

$$\tilde{x} = \frac{h_c \overline{x} + h_x \overline{x^2} + h_y \overline{xy}}{\overline{h}},$$

$$\tilde{y} = \frac{h_c \overline{y} + h_x \overline{xy} + h_y \overline{y^2}}{\overline{h}}.$$
(6.122)

where $h_c = \overline{h}$ is the thickness at the cell center, (h_x, h_y) is the limited thickness gradient, and the terms with overbars are geometric means. On a rectangular mesh, only the terms proportional to $\overline{x^2}$ and $\overline{y^2}$ are nonzero. From (6.120), the temperature at the cell center is given by

$$T_c = \tilde{T} - T_x \tilde{x} - T_u \tilde{y},$$

where (T_x, T_y) is the limited gradient of temperature.

We preserve monotonicity by limiting the gradients. If $\bar{\phi}(i,j)$ denotes the mean value of some field in grid cell (i,j), we first compute centered gradients of $\bar{\phi}$ in the x and y directions, then check whether these gradients give values of ϕ within cell (i,j) that lie outside the range of $\bar{\phi}$ in the cell and its eight neighbors. Let $\bar{\phi}_{\max}$ and $\bar{\phi}_{\min}$ be the maximum and minimum values of $\bar{\phi}$ over the cell and its neighbors, and let ϕ_{\max} and ϕ_{\min} be the maximum and minimum values of the reconstructed ϕ within the cell. Since the reconstruction is linear, ϕ_{\max} and ϕ_{\min} are located at cell corners. If $\phi_{\max} > \bar{\phi}_{\max}$ or $\phi_{\min} < \bar{\phi}_{\min}$, we multiply the unlimited gradient by $\alpha = \min(\alpha_{\max}, \alpha_{\min})$, where

$$\alpha_{\text{max}} = (\bar{\phi}_{\text{max}} - \bar{\phi})/(\phi_{\text{max}} - \bar{\phi}),$$

$$\alpha_{\text{min}} = (\bar{\phi}_{\text{min}} - \bar{\phi})/(\phi_{\text{min}} - \bar{\phi}).$$
(6.123)

Otherwise the gradient need not be limited.

Locating departure triangles

The method for locating departure triangles is discussed in detail by Dukowicz and Baumgardner (2000). The basic idea is illustrated in Figure 6.1, which shows a shaded quadrilateral departure region whose contents are transported to the target or home grid cell, labeled H. The neighboring grid cells are labeled by compass directions: NW, N, NE, W, and E. The four vectors point along the velocity field at the cell corners, and the departure region is formed by joining the starting points of these vectors. Instead of integrating over the entire departure region, it is convenient to compute fluxes across cell edges. We identify departure regions for the north and east edges of each cell, which are also the south and west edges of neighboring cells. Consider the north edge of the home cell, across which there are fluxes from the neighboring NW and N cells. The contributing region from the NW cell is a triangle with vertices abc, and that from the N cell is a quadrilateral that can be divided into two triangles with vertices acd and ade. Focusing on triangle abc, we first determine the coordinates of vertices b and c relative to the cell corner (vertex a), using Euclidean geometry to find vertex c. Then we translate the three vertices to a coordinate system centered in the NW cell. This translation is needed in order to integrate fields in the coordinate system where they have been reconstructed.

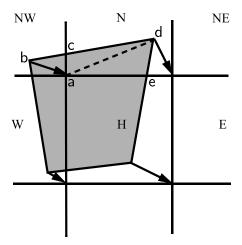


Figure 6.1: In incremental remapping, conserved quantities are remapped from the shaded departure region, a quadrilateral formed by connecting the backward trajectories from the four cell corners, to the grid cell labeled H. The region fluxed across the north edge of cell H consists of a triangle (abc) in the NW cell and a quadrilateral (two triangles, acd and ade) in the N cell.

Repeating this process for the north and east edges of each grid cell, we compute the vertices of all the departure triangles associated with each cell edge.

Figure 6.2, reproduced from Dukowicz and Baumgardner (2000), shows all possible triangles that can contribute fluxes across the north edge of a grid cell. There are 20 triangles, which can be organized into five groups of four mutually exclusive triangles as shown in Table 6.1. In this table, (x_1, y_1) and (x_2, y_2) are the Cartesian coordinates of the departure points relative to the northwest and northeast cell corners, respectively. The departure points are joined by a straight line that intersects the west edge at $(0, y_a)$ relative to the northwest corner and intersects the east edge at $(0, y_b)$ relative to the northeast corner. The east cell triangles and selecting conditions are identical except for a rotation through 90 degrees.

Departure triangles across a given cell edge are computed in a local coordinate system whose origin lies at the midpoint of the edge and whose vertices are at (-0.5, 0) and (0.5, 0). Intersection points are computed assuming Cartesian geometry with cell edges meeting at right angles. Let CL and CR denote the left and right vertices, which are joined by line CLR. Similarly, let DL and DR denote the departure points, which are joined by line DLR. Also, let IL and IR denote the intersection points $(0, y_a)$ and $(0.y_b)$ respectively, and let IC = $(x_c, 0)$ denote the intersection of CLR and DLR. It can be shown that y_a, y_b , and x_c are given by

$$y_{a} = \frac{x_{CL}(y_{DM} - y_{DL}) + x_{DM}y_{DL} - x_{DL}y_{DM}}{x_{DM} - x_{DL}},$$

$$y_{b} = \frac{x_{CR}(y_{DR} - y_{DM}) - x_{DM}y_{DR} + x_{DR}y_{DM}}{x_{DR} - x_{DM}},$$

$$x_{c} = (x_{DL} - y_{DL}) \frac{(x_{DR} - x_{DL})}{(y_{DR} - y_{DL})}.$$
(6.124)

Each departure triangle is defined by three of the seven points (CL, CR, DL, DR, IL, IR, IC). In Dukowicz and Baumgardner (2000), departure points are defined by projecting cell corner velocities directly backward. That is,

$$\mathbf{x_D} = -\mathbf{u}\,\Delta t,\tag{6.125}$$

where \mathbf{x}_D is the location of the departure point relative to the cell corner and \mathbf{u} is the velocity

Triangle	Triangle	Selecting logical
group	label	condition
1	NW	$y_a > 0$ and $y_1 \ge 0$ and $x_1 < 0$
	NW1	$y_a < 0$ and $y_1 \ge 0$ and $x_1 < 0$
	W	$y_a < 0 \text{ and } y_1 < 0 \text{ and } x_1 < 0$
	W2	$y_a > 0$ and $y_1 < 0$ and $x_1 < 0$
2	NE	$y_b > 0$ and $y_2 \ge 0$ and $x_2 > 0$
	NE1	$y_b < 0$ and $y_2 \ge 0$ and $x_2 > 0$
	\mathbf{E}	$y_b < 0 \text{ and } y_2 < 0 \text{ and } x_2 > 0$
	E2	$y_b > 0$ and $y_2 < 0$ and $x_2 > 0$
3	W1	$y_a < 0$ and $y_1 \ge 0$ and $x_1 < 0$
	NW2	$y_a > 0$ and $y_1 < 0$ and $x_1 < 0$
	E1	$y_b < 0 \text{ and } y_2 \ge 0 \text{ and } x_2 > 0$
	NE2	$y_b > 0$ and $y_2 < 0$ and $x_2 > 0$
4	H1a	$y_a y_b \ge 0$ and $y_a + y_b < 0$
	N1a	$y_a y_b \ge 0$ and $y_a + y_b > 0$
	H1b	$y_a y_b < 0$ and $\tilde{y}_1 < 0$
	N1b	$y_a y_b < 0$ and $\tilde{y}_1 > 0$
5	H2a	$y_a y_b \ge 0$ and $y_a + y_b < 0$
Ŭ.	N2a	$y_a y_b \ge 0$ and $y_a + y_b < 0$ $y_a y_b \ge 0$ and $y_a + y_b > 0$
	H2b	$y_a y_b \le 0$ and $\tilde{y}_a + \tilde{y}_b > 0$
	N2b	$y_a y_b < 0$ and $\tilde{y}_2 > 0$
	1.20	gago (cana g ₂) (

Table 6.1: Evaluation of contributions from the 20 triangles across the north cell edge. The coordinates x_1, x_2, y_1, y_2, y_a , and y_b are defined in the text. We define $\tilde{y}_1 = y_1$ if $x_1 > 0$, else $\tilde{y}_1 = y_a$. Similarly, $\tilde{y}_2 = y_2$ if $x_2 < 0$, else $\tilde{y}_2 = y_b$.

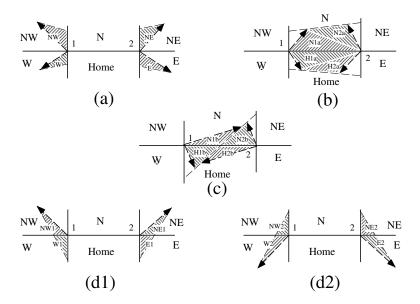


Figure 6.2: The 20 possible triangles that can contribute fluxes across the north edge of a grid cell.

at the corner. This approximation is only first-order accurate. In CISM, accuracy is improved by estimating the velocity at the midpoint of the trajectory.

Integrating fields

Next, we integrate the reconstructed fields over the departure triangles to find the total volume and internal energy transported across each cell edge. Volume transports are easy to compute since the ice thickness is linear in x and y. Given a triangle with vertices $\mathbf{x_i} = (x_i, y_i), i \in \{1, 2, 3\}$, the triangle area is

$$A_T = \frac{1}{2} |(x_2 - x_1)(y_3 - y_1) - (y_2 - y_1)(x_3 - x_1)|.$$
 (6.126)

The integral F_a of any linear function $f(\mathbf{r})$ over a triangle is given by

$$F_a = A_T f(\mathbf{x_0}),\tag{6.127}$$

where $\mathbf{x}_0 = (x_0, y_0)$ is the triangle midpoint,

$$\mathbf{x}_0 = \frac{1}{3} \sum_{i=1}^{3} \mathbf{x}_i. \tag{6.128}$$

Thus, to compute the volume transport, we evaluate the thickness at the midpoint,

$$h(\mathbf{x}_0) = h_c + h_x x_0 + h_y y_0, \tag{6.129}$$

and multiply by A_T . By convention, northward and eastward transport is positive, while southward and westward transport is negative.

Eq. (6.127) cannot be used for energy transport, because the reconstructed internal energy is a quadratic function of position. (It is the product of two linear functions, for thickness

and temperature.) The integral of a quadratic polynomial over a triangle requires function evaluations at three points,

$$F_h = \frac{A_T}{3} \sum_{i=1}^{3} f(\mathbf{x}_i'), \qquad (6.130)$$

where $\mathbf{x}'_i = (\mathbf{x}_0 + \mathbf{x}_i)/2$ are points lying halfway between the midpoint and the three vertices.

Updating state variables

Finally, we compute new values of the state variables in each ice layer of each grid cell. The new ice thickness h'(i, j) is given by

$$h'(i,j) = h(i,j) + \frac{F_E(i-1,j) - F_E(i,j) + F_N(i,j-1) - F_N(i,j)}{A(i,j)}$$
(6.131)

where $F_E(i,j)$ and $F_N(i,j)$ are the volume transports across the east and north edges, respectively, of cell (i,j), and A(i,j) is the grid cell area. All transports added to one cell are subtracted from a neighboring cell; thus (6.131) conserves total ice volume.

The new internal energy in each layer is computed analogously. New temperatures are then given by the ratio of energy to volume. (Other tracers, if present, are updated in the same way.) Tracer monotonicity is ensured because

$$T' = \frac{\int_A h \, T \, dA}{\int_A h \, dA},$$

where T' is the new-time temperature, given by integrating the old-time thickness and temperature over a Lagrangian departure region with area A. That is, the new-time temperature is a weighted averages over old-time values, with non-negative weights h. Thus the new-time values must lie between the maximum and minimum of the old-time values.

6.4.2 CFL checks

As mentioned above, the time step used for explicit advection is limited by the advective CFL condition (6.116). Furthermore, for ice flux parallel to ∇s , ice thickness evolution is diffusive, giving rise to an additional diffusive CFL condition (Bueler and Brown, 2009):

$$(\max D)\Delta t < 0.5\Delta x^2 \tag{6.132}$$

Flow governed by the shallow-ice approximation is always subject to this diffusive CFL, but some component of higher-order flow may not be because the ice velocity may not be parallel to ∇s (Bueler and Brown, 2009).

For these reasons, Glissade's transport scheme checks both the advective and diffusive CFL conditions and writes a warning to the log file if either is violated. These warnings indicate the i, j indices (on the global domain) of the point where the worst violation occurred, as well as the maximum allowable time step given the CFL condition(s).

The diffusivity is approximated using the component of velocity in the downslope direction:

$$D = \frac{(\boldsymbol{u} \cdot \nabla s)H}{|\nabla s|} \tag{6.133}$$

Future work may identify a less restrictive diffusivity, so the diffusive CFL warnings currently written may be overly restrictive. In other words, it is possible for a stable simulation to generate a diffusive CFL warning.

Advective CFL violations generally lead to stability problems with thickness evolution. However, since CFL violations may occur in portions of the domain that are not of interest to the

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user, these violations are not set to be a fatal error. Note, however, that the IR scheme may generate a "departure points out of bounds" error when advective CFL violations occur. This is a failure in the method that does result in a fatal error.

Eventually, CISM may include an adaptive time-stepping scheme that adjusts the time step based on the advective and diffusive CFL conditions. At present, however, the time step must be set manually and remain constant over a simulation. To aid in determining an appropriate time step, the adv_cfl_dt and diff_cfl_dt variables can be added to an output file. These variables are the maximum allowable time steps at each time in the model based on the advective and diffusive CFL conditions, respectively. In some cases, stability might require taking half of the reported values.

6.5 Basal physics

Glissade assumes a basal friction law of the form

$$\tau_{\mathbf{b}} = \beta \mathbf{u}_{\mathbf{b}},\tag{6.134}$$

where $\mathbf{u_b}$ is the basal velocity, τ_b is the basal shear stress, and $\beta > 0$. If β were independent of velocity, then (6.26) would be a simple linear sliding law. Allowing β to depend on velocity, however, allows more complex and physically realistic sliding laws.

The sliding law used by Glissade is determined by the value of which_ho_babc. The following options are supported:

- Spatially uniform β .
- Very slow sliding, enforced by setting β to a large value of 10^{10} Pa yr/m. This is a special case of spatially uniform β and is the default.
- \bullet No sliding, enforced by a Dirichlet boundary condition $\mathbf{u_b} = 0$ during finite-element assembly.
- A 2D β field specified from an external file. Typically, β is chosen at each vertex to optimize the fit to a surface velocity data set.
- β is set to a large value where the bed is frozen $(T_b < T_{pmp})$ and a lower value where the bed is thawed $(T_b = T_{pmp})$.
- Power-law sliding, based on the classic sliding relation of Weertman (1957). Basal velocity is given by

$$\mathbf{u_b} = k\tau_b^{\mathbf{p}} N^{-q},\tag{6.135}$$

where N is the effective pressure (the difference between the ice overburden pressure and the basal water pressure) and k is an empirical constant. This can be rearranged to give

$$\tau_{\mathbf{b}} = k^{-1/p} N^{q/p} \mathbf{u_b}^{1/p}, \tag{6.136}$$

which determines the form of β . Following Cuffey and Paterson (2010), the default exponents are p=3 and q=1.

• Coulomb friction, as described by Schoof (2005), with notation and default values from Pimentel et al. (2010). The form of the sliding law is

$$\tau_{\mathbf{b}} = C|\mathbf{u}|^{\frac{1}{n}-1}\mathbf{u}\left(\frac{N^n}{\kappa |\mathbf{u}| + N^n}\right)^{\frac{1}{n}}$$
(6.137)

where C is a dimensionless constant, $\kappa = m_{\text{max}}/(\lambda_{\text{max}}A_b)$, m_{max} is the maximum bed obstacle slope, λ_{max} is the wavelength of bedrock bumps, and A_b is a basal flow-law parameter. The default values are C = 0.42, $m_{\text{max}} = 0.5$, $\lambda_{\text{max}} = 2.0$ m, and $A_b =$ $10^{-16} \text{ Pa}^{-n} \text{ m}^{-1}$.

Eq. (6.137) has two asymptotic behaviors. In the interior, where the ice is thick and slow-moving, $\kappa |\mathbf{u}| \ll N^n$ and the basal friction is independent of N:

$$\tau_b \approx C|\mathbf{u}|^{\frac{1}{n}-1}\mathbf{u},\tag{6.138}$$

In the Coulomb-friction limit, where the ice is thin and fast-moving, $\kappa |\mathbf{u}| \gg N^n$, giving

$$\tau_b \approx \frac{C}{\kappa^{\frac{1}{n}}} N \frac{\mathbf{u}}{|\mathbf{u}|}.$$
 (6.139)

• Pseudo-plastic sliding, as described by Aschwanden et al. (2013) and implemented in the Parallel Ice Sheet Model (PISM). The basal friction law is

$$\tau_{\mathbf{b}} = -\tau_c \frac{\mathbf{u_b}}{|\mathbf{u_b}|^{1-q} u_0^q},\tag{6.140}$$

where τ_c is the yield stress, q is a pseudo-plastic exponent (default = 0.5), and u_0 is a threshold speed (default = 100 m/yr). This law incorporates linear (q = 1), plastic (q = 0) and power-law (0 < q < 1) behavior in a single expression. The yield stress is proportional to effective pressure:

$$\tau_c = \tan(\phi)N,\tag{6.141}$$

where ϕ is a friction angle that can vary with bed elevation. Following Aschwanden et al. (2016), the default is to set $\phi = 5^{\circ}$ at elevations more than 700 m below sea level and $\phi = 40^{\circ}$ at elevations more than 700 m above sea level, with a linear ramp in between.

The power-law, Coulomb-friction and pseudo-plastic sliding laws give lower friction and faster sliding as the effective pressure N decreases from the overburden pressure to lower values. CISM offers several options for computing N at the bed:

- $N = \rho_i g H$, the full overburden pressure. That is, the water pressure $p_w = 0$ at the bed.
- N is reduced as the basal temperature approaches the pressure melting point, reaching a small fraction of overburden pressure (typically $\delta = 0.02$) when $T = T_{pmp}$.
- N is reduced as the basal water depth increases, reaching a small fraction of overburden pressure when the water depth exceeds a prescribed threshold. (This option is of limited use, however, since CISM does not yet have a realistic basal hydrology model.)
- Following Leguy et al. (2014), N is reduced where the bed is below sea level, to account for partial or full connectivity of the basal water system to the ocean. The effective pressure is given by

$$N = \rho_i g H \left(1 - \frac{H_f}{H} \right)^p, \tag{6.142}$$

where $H_f = \max\left(0, \frac{\rho_w}{\rho_i}b\right)$ is the flotation thickness, b is the depth of the bed, and $0 \le$ $p \le 1$. For p = 0 there is no connectivity and N is the full overburden pressure. For p = 1there is full connectivity, and the basal water pressure is equal to the ocean pressure at the same depth.

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6.6 Calving

Glide includes several simple schemes for iceberg calving. Three of these schemes are supported in Glissade:

- Calve all floating ice.
- Calve ice where the current bedrock lies below a certain elevation.
- Calve ice where the relaxed bedrock (i.e., the bedrock toward which a viscous asthenosphere is relaxing) lies below a certain elevation.

In addition, there is a more complex scheme to calve floating ice thinner than a user-defined threshold H_C . As discussed by Albrecht et al. (2011), this is not as simple as it sounds. Suppose $H_C = 100$ m, and an ice shelf with ice thicker than 100 m is advancing. During the next time step, ice will be transported to grid cells just downstream of the initial calving front. Provided the time step is not too long, these cells will have $H < H_C$. If this thin ice immediately calves, then the calving front cannot advance.

Glissade avoids this problem by defining a protected ring of thin ice at the margin of floating ice. If ice in a grid cell has $H < H_C$ but shares an edge with a cell where $H \ge H_C$, then it is protected from calving. It can then thicken due to transport from neighboring cells, while remaining dynamically inactive. Once the ice in this cell thickens to $H \ge H_C$, it becomes dynamically active, and its thin-ice neighbors become part of the protected ring.

Suppose, on the other hand, that ice is thinning at the calving front. A cell might initially have $H \geq H_C$, with one or more protected thin-ice neighbors. If the ice in this cell thins to $H < H_C$, it becomes a dynamically inactive cell in the protected ring. Its thin-ice neighbors (assuming they have no other neighbors with $H \geq H_C$) are no longer protected and are allowed to calve. In this way the calving front can retreat.

The thickness-based calving scheme is more realistic than the simple schemes listed above, but still is likely to give large errors in calving-front location. A more accurate damage-based scheme is under development and will be included in future versions of CISM. In this scheme, a scalar damage parameter evolves as the ice undergoes strain, and ice calves when the damage in a grid cell exceeds a prescribed threshold.

By default, any floating ice that meets the calving criterion (e.g., it is thin and lies outside the protected ring, or the bedrock topography lies below the threshold elevation) will calve on a given time step. A user might prefer, however, to limit the extent of calving. By changing the value of calving_domain, the user can specify that ice calves (1) only when it lies at the marine margin (i.e., it borders an ice-free ocean cell), or (2) only when it is connected to the ocean through other cells that meet the calving criterion. In this way, cells in the interior of an ice shelf can be prevented from calving.

The user can also prescribe a calving timescale. If this timescale is longer than the model time step, then only a fraction of the ice that meets the calving criterion will calve in a single time step. For example, if the model time step is 1 yr and the calving timescale is 2 yr, then half the ice thickness in each calving-eligible cell will calve.

A related option is to remove so-called floating ice islands. A floating island is a connected region where every cell is floating and thus has no basal traction. These islands can be present in initial data sets or can arise occasionally during simulations, depending on the calving scheme. Finding the velocity for such a region is an ill-posed problem, so it is best to remove floating islands as soon as they occur. This is done by first marking all the grounded ice cells, and then using a boundary-fill algorithm to mark every cell that is connected to a grounded cell. (That is, one could travel from a grounded cell to the cell in question along a path at least one cell wide, without leaving the ice pack.) For parallel simulations, the algorithm is repeated enough times to connect cells on all processors. Any floating cells remaining are ice islands that immediately calve.

When CISM is coupled to a climate model, calving ice is typically routed to the ocean as a solid ice flux. In the absence of an explicit iceberg model, the ocean model supplies the heat needed to melt the ice. The ocean cools and freshens accordingly.

6.7 Other model physics

The Glissade modules for velocity, temperature, transport, and calving replace the corresponding Glide modules when dycore = 2 is set in the config file. Other model physics, however, remains similar to Glide. In particular:

- Isostasy is treated in the same way as in Glide (see Section 4.4). The lithosphere can be treated as local or elastic, and the asthenosphere is either fluid or relaxing. The elastic lithosphere calculation is non-local and has not been parallelized. Thus it is not possible to simulate an elastic lithosphere when running on more than one processor.
- Typically, the geothermal heat flux is set to a constant or prescribed from an input file. As in Glide, however, it is also possible to compute the geothermal heat flux from a model of heat flow in bedrock (see Section 4.3.2). This model includes non-local horizontal diffusion which, like the elastic lithosphere, has not been parallelized and therefore will not work on multiple processors.
- The basal water options in Glissade are also the same as those in Glide. These are all local except for basal_water = 2, which computes a flux of basal water based on a steady-state routing calculation. The routing calculation is not parallel and has not been tested with Glissade.

A sophisticated basal hydrology model is under development and will likely be included in future versions of CISM. This model will output the effective pressure, which is a required input for the Coulomb friction law (which_ho_babc = 10) recently introduced in CISM. For now the effective pressure is either read from an input file or parameterized simply (using basal_water = 4). The Coulomb friction law is not yet scientifically supported and should be used at your own risk.

In summary, although CISM has a sophisticated higher-order dycore, the physics parameterizations remain fairly simple, especially for runs with multiple processors. This is an area of active development.

Chapter 7

Running CISM

7.1 Overview of Running CISM

Assuming you successfully completed the Installation instructions in Chapter 2, the executable for running the model, cism_driver, can be found in your build directory in a subdirectory called cism_driver (e.g., ./builds/mac-gnu/cism_driver/cism_driver).

The build system creates the executable at this path but does not automatically make it available to other locations on your system. How you choose to do so depends on your situation. See the introduction to Chapter 8 for an overview of how to make the executable available to other locations on your system (e.g., symlinking, copying, or modifying your PATH environment variable).

CISM2.1 runs with a single executable, cism_driver, for all configurations. cism_driver can be invoked with a single argument specifying a CISM .config file to run CISM as a standalone ice sheet model without Glint climate forcing, or with two arguments (a CISM config file and a Glint config file) to run CISM with Glint climate forcing:

```
Call cism_driver with either 1 or 2 arguments. Examples: cism_driver ice_sheet.config climate.config climate.config
```

The available options for the CISM configuration file and for the Glint climate interface configuration file are described in detail below.

To perform a parallel run with the parallel build of CISM, you must use the MPI run command, which is typically mpirun or mpiexec but may vary among MPI versions and installations. A standard syntax that is likely to work on most installations is

```
mpirun -np N cism_driver ice_sheet.config climate.config where N is the number of processors you want to use, ice_sheet.config is the name of the CISM configuration file, and the optional argument climate.config is the name of the climate configuration file. For example:
```

```
mpirun -np 4 cism_driver dome.config
would run the dome test case on four processors.
```

When CISM runs, some basic information about its operation will be output to the screen (standard out). More verbose information about the run will be written to a log file which is named <code>ice_sheet.config.log</code>, where <code>ice_sheet.config</code> is the name of the .config file used to perform the run. (For example, if running the model with <code>./cism_driver dome.config</code>, the log file will be called <code>dome.config.log</code>.) The log file is an important reference, especially for debugging runs that do not behave as expected. For example, this file includes a list of configuration options and parameter values, which can be useful in diagnosing problems like typos in your .config file. The log file also indicates what files were used for input and output and

at which times I/O occurred. The log file may contain warnings about potentially problematic configuration combinations or model behavior, such as the use of configurations settings that are not scientifically validated, or a CFL violation during advection. In contrast, fatal errors will kill the model and the error message will be written to both the screen and the log file.

Optionally, the log file also contains diagnostic information about the global state of the ice sheet (e.g., the total ice area and volume, the maximum surface and basal speeds, and the max and min temperatures), along with vertical profiles of speed and temperature at a user-specified grid point. This information is written at intervals specified by the config file variable dt_diag, for the diagnostic point (idiag, jdiag).

In addition to the log file, the model will create any netCDF output files requested in the config file (see Section 7.4 below for details). If the model dies for some reason midway through a simulation, the netCDF files will still include output for the part of the simulation that was completed.

CISM2 (like its predecessor, Glimmer-CISM) has been coupled to the Community Earth System Model¹ and will be included in future CESM releases. Lipscomb et al. (2013) described the initial implementation of CISM in CESM, with one-way forcing of the Greenland ice sheet (using the shallow-ice approximation) by the surface mass balance computed in CESM's land model. Interactive coupling between ice sheets and other climate components has recently been implemented and is now being tested. This document, however, does not provide instructions for running CISM within CESM or other climate models. Please see here² for guidance on running CISM in CESM.

7.2 Overview of Configuration Files

Running CISM is managed through configuration files (*.config) that enable desired model features and control input of initial conditions and forcing and output of model results. This chapter summarizes the configuration options available for running CISM and is divided into sections on general Model Configuration, Input/Ouput Configuration, and optional Climate Forcing Configuration.

The format of CISM configuration files is taken from that used by the ConfigParser module in Python 2.x, which is similar to Windows .ini files and contains sections. Each section contains key/value pairs.

Comments: Empty lines, or lines starting with a #, ; or ! are ignored. Comments can also be added on the same line as a key/value pair using these delimiters.

Sections: A new section starts with the section name enclosed by square brackets, [] and can be up to 50 characters long, e.g., [grid].

Key/Value Pairs: Keys are separated from their associated values by = or :. The names can be up to 50 characters long. Values can be up to 400 characters long.

Sections and keys are case-sensitive and may contain white space. However, the configuration parser is very simple and thus the number of spaces within a key or section name also matters. Sensible defaults are used when a specific key is not found; defaults are shown in bold in the tables below.

Here is an example configuration file:

;a comment
[a section]
an_int : 1

 $^{^{1}}$ www2.cesm.ucar.edu

²http://www.cesm.ucar.edu/models/cesm2/land-ice/

```
a_float = 2.0
a_char = hey, this is rather cool
an_array = 10. 20. -10. 40. 100.

[another section]
! more comments
foo : bar
```

7.3 Model Configuration

General model configuration options specify the grid and time-stepping used by the model, the dynamical core (dycore) used, and control various optional physics packages and parameter values. The [grid] and [time] configuration sections are always required. Also, while not required, in almost all situations [options] and (if using a higher-order dycore) [ho_options] will be included in .config files. The [parameters] and [sigma] sections are also commonly used. The [GTHF], [isostasy], and [projection] sections are needed only if the associated features are desired. Details of each of these sections, what they control, and the available options for each section are listed in the tables below.

[grid]		
Define model grid.		
ewn	(integer) number of nodes in x-direction	
nsn	(integer) number of nodes in y-direction	
upn	(integer) number of nodes in z -direction	
dew	(real) node spacing in x -direction (m)	
dns	(real) node spacing in y -direction (m)	
nx_block	0 (int) user specific block size to handle active-blocks option	
ny_block	0 (int) user specific block size to handle active-blocks option	
global_bc	boundary conditions for the edges of the global domain	
	0 periodic	
	1 outflow	
	2 no penetration	
	3 no ice	
sigma	method for specifying sigma coordinates:	
	0 Use Glimmer's default spacing	
	$\sigma_i = \frac{1 - (x_i + 1)^{-n}}{1 - 2^{-n}}$ with $x_i = \frac{\sigma_i - 1}{\sigma_n - 1}, n = 2.$	
	1 use sigma coordinates defined in external file (named sigma_file)	
	2 use sigma coordinates given in configuration file	
	3 use evenly spaced sigma levels (required by the Glam dycore)	
	4 use Pattyn sigma levels	
[grid ocn]		

grid_ocn	
Define ocean model grid for ocean data input.	
nbasin	(int) Number of basins
nzocn	(int) Number of ocean levels
dzocn (real) Thickness of ocean levels	

sigma

Define the sigma levels used in the vertical discretization (sigma=2 above). This is an alternative to using a separate file (specified in section [grid] above). If neither is used, the levels are calculated as described above.

continued on next page

continued from previous	page
sigma_levels	(real) list of sigma levels, in ascending order, separated by
_	spaces. These run between 0.0 and 1.0.
[time]	
Configure time steps and diagnostic specifications	

[time]		
Configure time steps and diagnostic specifications		
tstart	(real) start time of the model in years	
tend	(real) end time of the model in years	
$\mathtt{dt_input_option}$	0 Input dt in year	
	1 Input dt in steps per year	
dt	(real) size of time step in years	
nsteps_per_year	(int) Number of time steps per year (for dt_input_option=1)	
subcyc	(int) number of time steps to subcycle evolution within dt	
	using a steady velocity field	
adaptive_cfl_threshold	(real) adaptively subcycle the advection when advective	
	CFL exceeds this value	
ntem	(real) time step multiplier setting the ice temperature up-	
	date interval	
$\mathtt{dt}_{\mathtt{-}}\mathtt{diag}$	(real) writing diagnostic variables to log file every dt_diag	
	yrs (for dt_input_option=0)	
ndiag	(int) writing diagnostic variables to log file every ndiag time	
	steps (for $dt_{input_option} = 1$)	
idiag	(int) x direction index for diagnostic grid point in log file	
jdiag	(int) y direction index for diagnostic grid point in log file	

[options]

Parameters set in this section determine how various components of the ice sheet model are treated. Configuration number options with a † are specific to the higher-order dycores (e.g., Glissade). Options with a ? are working, but are currently not scientifically supported, and are therefore for use at your own risk. Options with a ! are in development and will be supported in future code releases.

dycore	0 Glide (1-processor, 3d, shallow-ice-approximation
	dycore)
	1†? Glam (parallel, 3d, FDM, 1st-order-accurate dy-
	core)
	2† Glissade (parallel, 3d, FEM, 1st-order-accurate
	dycore)
	3†! FELIX (parallel, 3d, FEM, 1st-order-accurate dycore)
	4†! BISÍCLES (parallel, quasi-3d, FVM, L1L2 dycore)
evolution (ice thickness)	0 pseudo-diffusion (Glide only)
	1 ADI scheme (Glide only)
	2 diffusion (Glide only)
	3† incremental remapping
	4† first-order upwind
	5† evolve without changing ice thickness (Useful for
	running with a fixed geometry, e.g. for a temper-
	ature spinup. On each time step, geometry and
	tracers are evolved using incremental remapping,
	after which geometry is reset to its initial value.
	This evolution scheme is still subject to the advec-
	tive CFL condition.)
	continued on next page

continued from previous p	0 Set each ice column to local surface air tempera-
cemperature	ture
	1 prognostic temperature calculation
	2 hold temperature steady at initial value
	3! prognostic temperature calculation using
	enthalpy-based formulation
$ exttt{temp_init}$	0 initial temperatures isothermal at 0°C
	1 initial column temperatures set to atmos. temperature
	2 initial column temperatures linearly interpolated
	between atmos. temperature and pressure melting
	point
	3 initialize temperature with an advective-diffusive
	balance in each column
	4 initialize temperature from external file
flow_law	o set equal to $1 \times 10^{-16} \text{Pa}^{-n} \text{yr}^{-1}$
T T O W _ T C W	1 temperature-dependent, Paterson and Budd
	(1982) $(T = -5^{\circ}\text{C})$
	2 temperature-dependent, Paterson and Budd
	(1982) (function of variable T)
	3 Read flow_law from file
slip_coeff	slip coefficient (Glissade local SIA and Glide only)
siip_coeii	0 zero (no sliding)
	1 set to a non-zero constant everywhere
	· /
	4 set proportional to basal melt rate
	5 tanh function of basal water (bwat)
basal_water	0 none
	1 local water balance
	2? compute the steady-state, routing-based, basal
	water flux and water layer thickness (NOTE: not
	supported for > 1 processor)
	3 use a constant basal water layer thickness every-
	where, to enforce $T=T_{pmp}$ everywhere
bmlt_float	0 none
	depth-dependent basal melt rate for floating ice as
	specified for MISMIP+ (Asay-Davis et al., 2016)
	2 constant value for floating ice (with option to se-
	lectively mask out melting)
	3 depth-dependent basal melt rate for floating ice
	4 external basal melt rate for floating ice field (from
	input file or coupler)
	5? basal melt rate for floating ice from MISOMIP
	(Asay-Davis et al., 2016) ocean forcing with plume
	model
	6 basal melt rate for floating ice derived from ocean
	thermal forcing
	continued on next pa

continued from previous page	
bmlt_float_thermal_forcing_param	used with config option bmlt_float = 6 only
	0 Quadratic parameterization to compute basal
	melting from thermal forcing
	1 ISMIP6 local quadratic parameterization to com-
	pute basal melting from thermal forcing
	2 ISMIP6 nonlocal quadratic parameterization to
	compute basal melting from thermal forcing
	3 ISMIP6 nonlocal quadratic parameterization with
	slope dependence
bmlt_float_ismip6_magnitude	used with config option bmlt_float = 6
	0 Lowest level of forcing (e.g., pct5)
	1 Median level of forcing
	2 High level of forcing (e.g., pct95)
ocean_data_domain	o ocean data on ocean domain only; extrapolate
	data to shelf cavities)
	1 ocean data available everywhere; already extrapolated to shelf cavities
	2 ocean data applied where CISM has ice-free ocean;
	extrapolated to shelf cavities
enable_bmlt_anomaly	false no anomaly applied to bmlt_float
Chapte_bmrt_anomary	true apply a prescribed anomaly to bmlt_float
basal_mass_balance	0 ignore basal melt rate in mass balance calculation
	1 include basal melt rate in mass balance calculation
smb_input	0 SMB input in units of m/yr ice (same as acab)
•	1 SMB input in units of mm/yr water equivalent
smb_input_function	0 SMB(x,y); input as a function of horizontal loca-
-	tion only
	1 SMB(x,y) + $dSMB/dz(x,y) * dz$; input SMB and
	its vertical gradient
	2 SMB(x,y,z); input SMB at multiple elevations
$artm_input_function$	0 artm(x,y); input as a function of horizontal loca-
	tion only
	1 $\operatorname{artm}(x,y) + \operatorname{dartm}/\operatorname{dz}(x,y) * \operatorname{dz}$; input air tem-
	perature and its vertical gradient
	2 artm(x,y,z); input air temperature at multiple el-
	evations
nlev_smb	(int) 1 (number of vertical levels SMB is provided
enable_acab_anomaly	false no anomaly applied to SMB true apply a prescribed anomaly to SMB
enable_artm_anomaly	true apply a prescribed anomaly to SMB false no anomaly applied to air temperature
enable at om-anomaly	true apply a prescribed anomaly to air temperature
overwrite_acab	0 no overwrite of SMB anywhere
0 1 01 WI I 00 - a0 a0	1 overwrite SMB where input SMB = 0
	2 overwrite SMB where input sMB = 0
	threshold value
	continued on next page
	puge

continued from previous page	
gthf	o prescribed, uniform geothermal heat flux
	1 read 2d geothermal heat flux field from input file
	2 calculate geothermal heat flux using 3d diffusion
	model
isostasy	0 no isostatic adjustment
	1 compute isostatic adjustment using lithosphere /
	asthenosphere model (see below for available op-
	tions)
marine_margin	0 ignore marine margin
	1 set thickness to zero if floating
	2 lose fraction of ice from edge cells
	3 set thickness to zero if relaxed bedrock is below a
	given depth (variable "mlimit" in glide_types)
	4 set thickness to zero if present-day bedrock
	is below a given depth (variable "mlimit" in
	glide_types) 5 set thickness to zero based on grid location (field
	()
	"calving_mask")
	6 set thickness to zero if ice at marine margin
	is thinner than a certain value (variable "calv-
	ing_minthck" in glide_types)
	7 set thickness to zero based on stress (eigencalving)
	criterion
	8 calve ice that is sufficiently damaged
	9? Huybrechts calving scheme
calving_init	0 do not calve at initialization
	1 calve at initialization
calving_domain	0 calve only at ocean edge
	1 calve wherever the calving criterion is met
apply_calving_mask	false no calving mask applied
	true apply a calving mask to prevent calving-front ad-
	vance
remove_icebergs	false no icebergs removal
	true identify and remove icebergs after calving
	Note: these are connected regions with zero basal traction
	and no connection to grounded ice. It is safer to make it
	true, but not necessary for all applications.
remove_isthmuses	false no isthmuses removal
Temove_T2 ciniluses	
	true identify and remove ice isthmuses after calving
	Note: these are narrow bridges connecting two regions of
	floating ice. False by default, but may need to be true for
	the FORCE_RETREAT_FLOATING_ICE option.
expand_calving_mask	false do not expand the calving mask
	true expand the calving mask to include all ice that is
	floating at initialization
	continued on next page

continued from previous page	
	Note: by default, this is done for a hardwired set of ISMIP6
	basins, excluding large shelves
limit_marine_cliffs	false marine cliffs remain unchanged
	true thin marine-based cliffs based on a thickness
	threshold
cull_calving_front	false no calving front culling
_	true cull calving_front cells at initialization
	Note: this can make the run more stable by removing long,
	thin peninsulas
adjust_input_thickness	false no ice thickness adjustment
	true adjust ice thickness to maintain upper surface, in-
	stead of deriving upper surface from bed topogra-
	phy and ice thickness
smooth_input_topography	false no bed topography smoothing
	true apply Laplacian smoothing to the topography at
	initialization
adjust_input_topography	false no bed topography adjustment
	true adjust the input topography in a selected region
	at initialization
read_lat_lon	false no read of lat-lon fields
	true read lat and lon fields from the input file and write
	to restarts
dm_dt_diag	0 write dmass/dt diagnostic in units of kg/s
	1 write dmass/dt diagnostic in units of Gt/yr
diag_minthck	0 include cells with H; 0 in global diagnostics
	1 include cells with H ; thklim in global diagnostics
vertical_integration	(Glide only)
	o standard integration (to obtain vertical velocity
	profile) 1 constrained to obey kinematic velocity at upper
	surface boundary
restart	Note: alternate keyword hotstart is retained for back-
1 CS dai t	wards compatibility.
	0 normal start (initial values taken from input file
	or, if absent, using default options)
	1 restart model using input from previous run; spe-
	cific fields required for restart are dependent on
	chosen options (add "restart" to the variable list
	in the [CF output] section of the .config file to
	automatically save the appropriate restart fields.)
restart_extend_velo	0 write velocities to restart file on standard stag-
	gered mesh
	1 write uvel_extend and vvel_extend to restart file
	on extended staggered mesh (required if restart
	velocities are nonzero on global boundaries)
	continued on next page

continued from previous page	(string) name of filet-i-itCDE I/O
ioparams	(string) name of file containing netCDF I/O configuration.
	The main configuration file is searched for I/O related sections if no file name is given (default). In other words,
	you can remove sections CF input, CF output, and CF
	forcing from the primary configuration file and place them
	in a separate file, the path to which is specified here.
[he ontional	The state of the s
[ho_options] Ontions set in this section date	rmine how various components of the higher-order extensions
	Glissade) are treated. Defaults are indicated in bold. These
	shallow-ice (Glide) dycore. In this section, options with a?
1 -	not scientifically supported (and are therefore for use at your
	th a * apply only to a serial build (or a parallel build if run
on 1 processor). Options mark	xed with a! are under development and will be supported in
future versions of the code (he	ence, these are also for use at your own risk).
which_ho_efvs	0 use a constant value for the effective viscosity (i.e.,
	linear viscosity). The default value is 2336041 Pa
	yr (as used by ISMIP-HOM Test F).
	1 set the effective viscosity to a value based on the
	flow rate factor: efvs = $0.5 * A^{-1/n}$ 2 use the effective strain rate to compute the effec-
	tive viscosity (i.e., full nonlinear treatment)
which_ho_disp	-1 no dissipation term included in temperature equa-
wiitoninolarsp	tion
	0 calculate dissipation in temperature equation as-
	suming SIA ice dynamics
	1 calculate dissipation in temperature equation as-
	suming first-order ice dynamics
which_ho_thermal_timestep	0 vertical thermal solve before transport solve
	1 vertical thermal solve after transport solve
	2 vertical thermal solve split; both before and after
which_ho_babc	transport solve Implementation of basal boundary condition in higher-
wiiten_no_babe	order dycore
	0 spatially uniform value of "beta" (low value of 10
	Pa/yr by default)
	1 large value for frozen bed, lower value for bed at
	pressure melting point (hardcoded, mainly useful
	for debugging)
	2 read map of yield stress (in Pa) from input field
	"mintauf" to simulate sliding over a plastic sub-
	glacial till (Picard-based solution)
	3 pseudo-plastic basal sliding law; can model linear,
	power-law or plastic behavior (virtually) no slip everywhere in domain ("beta"
	set to very large value)
	5 read map of "beta" from .nc input file using stan-
	dard I/O
	6 no slip everywhere in domain (using Dirichlet basal
	BC)
	7! read map of yield stress (in Pa) from input field
	"mintauf" to simulate sliding over a plastic sub-
	glacial till (Newton-based solution)
	8* Spatial field of "beta" required for ISMIP-HOM
	Test C (avoids interpolation error associated with option 5; works for a single processor only)
	9 power law
	10 Coulomb friction law using effective pressure, with
	flwa from lowest ice layer
	11 Coulomb friction law (Fg. 5.34)

11 Coulomb friction law (Eq. 5.34)
12 basal stress is the minimum of Coulomb and

continued from previous page	
$compute_blocks$	0 compute on all blocks in the global domain; one
	task per block
	1 compute on active blocks only; one task per active
	block, no task for some or all inactive blocks
	2 inquire how many active blocks there are, in prep
	for resubmitting with option 1
use_c_space_factor	Multiplying "beta" by a 2D scalar field (used for MISMIP3d experiments (Pattyn et al., 2013))
	false no multiplication by a scalar field
	true multiplication by a scalar field
which_ho_beta_limit	Setting minimum "beta" beneath grounded ice
which_no_beta_limit	
	0 absolute limit given by beta_grounded_min
	1 limit using beta_grounded_min, then multiplied by grounded ice fraction (f_ground)
ahiah ha an inananian	
which_ho_cp_inversion	Basal inversion options: invert for Cp = powerlaw_c
	Note: Cp inversion is currently supported for
	which_ho_babc = 9 and 11 only
	0 no inversion
	1 invert for basal friction parameter Cp
	2 apply Cp from a previous inversion
which_ho_bmlt_inversion	Basal inversion options: invert for bmlt_float
	0 no inversion
	1 invert for basal melt rate (bmlt_float)
	2 apply basal melt rate from previous inversion
which_ho_bmlt_basin_inversion	Inversion of basin-based basal melting parameters
	0 no inversion
	1 invert for basin-based melting parameters
	2 apply basin-based melting parameters from previ-
	ous inversion
which_ho_bwat	Basal water depth
	O Set to zero everywhere.
	1 Set to constant everywhere, to force $T = Tpmp$
	2 Local basal till model with constant drainage
which_ho_effecpress	Effective pressure calculation for HO dyn core
	0 N = overburden pressure, rhoi*g*H
	1 N is reduced where the bed is at or near the pres-
	sure melting point
	2 N is reduced where there is melting at the bed
	3 N is reduced due to connection of subglacial water
	to the ocean
	4 N is reduced where basal water is present
	continued on next page

continued from previous page which_ho_resid	
wnicn_no_resid	Residual calculation method for higher-order velocity
	solvers (e.g., Glissade). Nonlinear iterations are halted once
	the residual falls below a specified value.
	0? use the maximum value of the normalized velocity
	vector update, defined by $r = \frac{ vel_{k-1} - vel_k }{vel_k}$
	1? as in option 0 but omitting the basal velocities
	from the comparison (useful in cases where an ap-
	prox. no slip basal BC is enforced)
	2? as in option 0 but using the mean rather than the
	max
	3 use the L2 norm of the system residual, defined by $r = Ax - b$
	4 use L2 norm of residual relative to rhs, $ Ax-b / b $
which_ho_nonlinear	Method for solving the nonlinear iteration when solving the
	first-order momentum balance
	0 treat nonlinearity in momentum balance using Pi-
	card iteration
	1? treat nonlinearity in momentum balance us-
	ing Jacobian-Free Newton-Krylov iteration (Glam
	only)
which_ho_sparse	Method for solving the sparse linear system that arises from
_	the higher-order solver
	-1* solve sparse linear system using SLAP with incom-
	plete Cholesky preconditioned conjugate gradient
	method
	0 solve sparse linear system using SLAP with in-
	complete LU-preconditioned biconjugate gradient
	method
	1* solve sparse linear system using SLAP with incom-
	plete LU-preconditioned GMRES method
	2 solve sparse linear system using preconditioned
	conjugate gradient method: standard algorithm
	(Glissade only)
	3 solve sparse linear system using preconditioned
	conjugate gradient method: Chronopoulos-Gear
	algorithm (Glissade only)
	4 solve sparse linear system using <i>Trilinos</i> , incom-
	plete LU-preconditioned GMRES method (Trili-
	nos-compatible build only)
	continued on next pag

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which_ho_approx	Stokes-flow approximation to use with Glissade dycore	
	Note 1: this option is not valid for other dycores	
	Note 2:There are two SIA options:	
	· Option -1 uses module glissade_velo_sia to compute local	
	SIA velocities, similar to Glide	
	· Option 0 uses module glissade_velo_higher to compute SIA	
	velocities via an iterative solve	
	-1 local shallow-ice approximation, Glide-type calcu-	
	lation (uses glissade_velo_sia)	
	0 3d matrix shallow-ice approximation, vertical-	
	shear stresses only (uses glissade_velo_higher)	
	1 shallow-shelf approximation (SSA) with	
	horizontal-plane stresses only (uses glis-	
	sade_velo_higher; requires which_ho_precond	
	<=1)	
	2 Blatter-Pattyn with both vertical-shear	
	and horizontal-plane stresses (uses glis-	
	sade_velo_higher)	
	3 depth-integrated (L1L2) approximation, with	
	both vertical shear and horizontal-plane	
	stresses (uses glissade_velo_higher; requires	
	which_ho_precond <=1)	
	4 depth-integrated viscosity approximation based on	
	Goldberg (2011) (uses glissade_velo_higher)	
which_ho_precond	Preconditioner to use in the linear PCG solve of the Glis-	
_	sade dycore	
	0 no preconditioner	
	1 diagonal preconditioner	
	2 physics-based (shallow-ice) preconditioner (not	
	valid for SSA and L1L2)	
	2 physics-based shallow-ice preconditioner	
	3 local tridiagonal preconditioner	
	4 global tridiagonal preconditioner	
which_ho_gradient	Which spatial gradient operator to use in the Glissade dy-	
	core	
	0 centered gradient	
	1 first order upstream gradient (damps checkerboard	
	noise in prognostic simulations)	
	2 second order upstream gradient	
which_ho_gradient_margin	Spatial gradient operator to use in the Glissade dycore at	
	ice sheet margins.	
	0 use information from all neighboring cells, ice-	
	covered or ice-free	
	1 use information from ice-covered and/or land cells,	
	but not ice-free ocean cells	
	2 use information from ice-covered cells only	
	continued on next page	

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which_ho_vertical_remap	Order of accuracy for vertical remapping
	0 first-order accurate in the vertical direction
	1 second-order accurate in the vertical direction
which_ho_assemble_taud	Finite-element assembly method for driving stress term
	0 Standard finite-element calculation, which effec-
	tively applies a smoothing to driving stress
	1 Apply the local driving stress value at each vertex
	(no smoothing)
which_ho_assemble_beta	Finite-element assembly method for basal boundary condi-
	tions that use "beta" field
	0 Standard finite-element calculation, which ef-
	fectively applies a smoothing to "beta" (and
	"mintauf")
	1 Apply the local "beta" (or "mintauf") value at
	each vertex (no smoothing)
which_ho_assemble_bfric	Finite-element assembly method for basal friction heat flux
	conditions
	0 Standard finite-element calculation, summing over
	each quadrature points
	1 Apply the local "beta"* $(u^2 + v^2)$ (or
	"mintauf" * $(u^2 + v^2)$) value at each vertex
	(no smoothing)
which_ho_assemble_lateral	Finite-element assembly method for lateral stress terms O Standard finite-element calculation of H and usrf.
	O Standard finite-element calculation of H and usrf, at quadrature points
	1 Apply the local cell-center value of H and usrf on
	each face (no smoothing)
which_ho_calving_front	Use of subgrid calving front parameterization
willow no conving from	0 no subgrid calving front parameterization
	1 subgrid calving front parameterization
which_ho_ground	Computation of the grounded fractions of each cell in the
	glissade dycore only
	0 a cell is either floating or grounded (based on flota-
	tion condition)
	1 grounding line parameterization: a cell is partially
	floating/grounded (0 \leq f_ground \leq 1)
	2 similar to option 1, but f_ground is summed over
	quadrants rather than staggered cells
which_ho_fground_no_glp	Indicate how to identify grounded and floating vertices
	when running without a GLP
	0 a cell is grounded if any neighbor cell is grounded
	or land, otherwise it is floating
	1 the cell is grounded depending on the staggered
	flotation function
	continued on next page

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which_ho_ground_bmlt	Computation of the melt (bmlt_float) in partly grounded cells (Leguy et al., 2020)
	o apply melt in floating cells only (based on floating_mask)
	1 weigh the melt by the floating fraction of the cell if $(0 < f_{ground} < 1)$
	2 set melt (bmlt_float) to zero in partly grounded cells (f_ground > 0)
which_ho_flotation_function	Computation of flotation function at and near vertices (glis-
	sade dycore only) 0 f_flotation = (-rhow*b/rhoi*H) = f_pattyn (<=1
	for grounded ice, > 1 for floating ice) (Pattyn, 2006)
	$1 f_{\text{-flotation}} = (\text{rhoi*H})/(\text{-rhow*b}) = 1/f_{\text{-pattyn}}$
	(>=1 for grounded ice, < 1 for floating ice)
	2 f_flotation = -rhow*b - rhoi*H = ocean cavity
	thickness (≤ 0 for grounded ice, > 0 for floating ice)
	3 Modified version of option 2, without extrapolation to ice-free ocean
	4 As option 3 but with standard deviation correction
	for bed topography
block_inception	Blocking ice inception away from the main ice sheet
•	Note: if set to true, allows for code speed up: this option
	partitions ocean cells in a domain over which computation
	does not take place. Expert use only!
	false allowing ice inception away from ice sheet
	true not allowing ice inception away from ice sheet
remove_ice_caps	Removing ice caps and adding them to the calving flux false no ice caps removal
	true ice caps removal
force retreat	Forcing ice retreat using ice_fraction_retreat_mask
10100_1001000	Note: this is one of the retreat option used in ISMIP6 (Now-
	icki et al., 2016)
	0 no forced retreat
	1 forced retreat
which_ho_ice_age	Computation of 3d ice age tracer
	0 No ice age computation
m]:	1 ice age computation
glissade_maxiter	100 maximum number of nonlinear (Picard) iterations in the Glissade dycore
linear_solve_ncheck	5 check the linear solver for convergence every linear_solve_ncheck iterations
linear_maxiter	200 maximum number of linear iterations before quit-
	ting
	continued on next page

<pre>continued from previou linear_tolerance</pre>	1e-8 error tolerance for linear solver
[parameters]	
	s parameters. Parameters with a † are specific to the higher-order
dycores (e.g., Glissad	- · · · · · · · · · · · · · · · · · · ·
rhoi	(real) ice density (default = 917.0 kg m^{-3})
rhoo	(real) ocean density (default = 577.6 kg m^{-3})
grav	(real) gravitational acceleration (default = 9.81 m s^{-2})
shci	(real) Specific heat capacity of ice (default = 2009.0
SHCI	J kg $^{-1}$ K $^{-1}$)
lhci	(real) Latent heat of melting of ice (default = 335.0 J kg^{-1})
trpt	(real) Triple point of water (default = 273.16 K)
log_level	(int) set to a value between 0, no messages, and 6, all
J	messages are displayed to stdout. By default, messages
	are only logged to a file. The format for this filename is
	"configuration-file-name.config.log"
thk_gradient_ramp	(real) thickness scale over which gradients increase from
G	zero to full value (default $= 0.0$); higher-order only
ice_limit	(real) below this limit ice is only accumulated/ablated; ice
	dynamics are switched on once the ice thickness is above
	this value. (default = 100.0 m)
ice_limit_temp †	(real) minimum thickness for computing vertical tempera-
icc_iimic_ccmp	ture (default = 1.0 m)
pmp_offset	(real) offset of initial bed temperature from pressure melt-
pmp_orrsec	ing point temperature (default = 5.0 deg C)
nmn +hroahold	(real) bed is assumed thawed where Tbed := pmptemp -
pmp_threshold	pmp_threshold (default = 1e-3 deg C)
maa+hammal	,
geothermal	(real) constant geothermal heat flux, positive down by con-
£3 £+	vention (hence < 0). (default = -0.05 W m^{-2})
flow_factor	(real) the flow law rate factor is multiplied by this factor
	(default = 1.0 ; in previous versions of Glimmer-CISM the
63 6 4 63	default value was 3.0)
flow_factor_float	(real) flow enhancement parameter for floating ice (default
	= 1.0; for marine simulations a smaller value may be needed
	to match observed shelf speeds)
$default_flwa$	flow law parameter A to use in isothermal experiments
	(flow law set to 0) (default value is $10^{-16} \text{ Pa}^{-n} \text{ yr}^{-1}$). This
	overrides any temperature dependence.
$efvs_constant \dagger$	Constant value of effective viscosity when using
	which_ho_efvs=0 (default value is 2336041 Pa yr, as
	in ISMIP-HOM Test F).
effstrain_min	(real) min value of effective strain for computation of effec-
	tive viscosity (default value is 1e-8 yr ⁻ 1; generally works
	well for SSA and DIVA, but a larger value (1e-6) may be
	needed for BP)
max_slope	(real) maximum surface slope allowed in Glissade dycore
1	(default = 1.0 (unitless)); Note: It may be necessary to re-
	duce max_slope to 0.1 to prevent huge velocities in regions
	of rough coastal topography

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Parameters to adjust exter	rnal forcing
acab_factor	(real) adjustment factor for external acab field (default = 1.0 (unitless))
bmlt_float_factor	(real) adjustment factor for external bmlt_float field (default = 1.0)
Calving parameters	,
marine_limit	(real) all ice is assumed lost (calved) once water depths reach this value (for marine_margin=3 or 4 in [options] above). Note, water depth is negative. (default = -200.0 m)
${\tt calving_fraction}$	(real) fraction of ice lost due to calving (for marine_margin=2). (default = 0.8)
calving_minthck	(real) minimum thickness of floating ice at marine edge before it calves (default $= 0.0$ m). If used, must be set to a nonzero value. Used with [options] marine_margin $= 6.7.8$ only.
lateral_rate_max	(real) maximum lateral calving rate for damaged ice (default = 3000 m yr^{-1}); used with [options] marine_margin = 8 only .
eigencalving_constant	(real) eigencalving constant per unit stress (default = 1e-2 m yr ⁻¹ Pa ⁻¹); used with [options] marine_margin = 7 only.
eigen2_weight	(real) weight given to the second eigenvalue relative to the first eigenvalue of 2D horizontal stress tensor in the effective calving stress (default = 1.0 unitless)
${\tt damage_constant}$	(real) rate of change of damage (default = $1e-7 \text{ yr}^{-1} \text{ Pa}^{-1}$); used with [options] marine_margin = 8 only .
${\tt damage_threshold}$	(real) threshold at which ice column is deemed sufficiently damaged to calve, assuming that $0 = \text{no damage}$, $1 = \text{total damage (default} = 0.75)$; used with [options] marine_margin = 8 only.
taumax_cliff	(real) yield stress for marine-based ice cliffs. (default = 1e6 Pa)
cliff_timescale	(real) time scale for limiting marine cliffs (default = 10.0 yr)
ncull_calving_front	(int) number of times to cull calving front cells at initialization (default = 0). If [options] cull_calving_front is True T, set to a large value to remove peninsulas.
calving_timescale	(real) time scale for calving (default = 0.0 yr). If calving_timescale = 0, then the full column calves at once. GLISSADE-only parameter
calving_front_x	(real) calve ice wherever $abs(x) > calving_front_x$ (default = 0.0 m). Used with [options] marine_margin=5 only. This option is applied only if calving_front_x > 0.
${\tt calving_front_y}$	(real) calve ice wherever abs(y) > calving_front_y (default = 0.0 m). Used with [options] marine_margin=5 only. This option is applied only if calving_front_y > 0.
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Basal traction parameters	
basal_tract_const	constant basal traction parameter. You can load a .nc file with a variable called soft if you want a spatially varying bed softness parameter (Glissade local SIA and Glide only)
basal_tract_max	max value for basal traction when using slip_coeff=4.
basal_tract_slope	slope value for basal traction relation when using slip_coeff=4. (Relation also uses basal_tract_const.)
basal_tract_tanh	$(real(5))$ basal traction factors. Basal traction is set to $B = tanh(W)$ with the parameters (1) width of the tanh curve (2) W at midpoint of tanh curve [m] (3) B minimum $[ma^{-1}Pa^{-1}]$ (4) B maximum $[ma^{-1}Pa^{-1}]$ (5) multiplier for marine sediments
Basal friction/sliding parameter	ers
beta_grounded_min	(real) minimum value of beta for grounded ice (default = 1.0 Pa yr m^{-1}). Scaled during initialization. GLISSADE-only parameter.
ho_beta_const †	(real) spatially uniform beta used when which_ho_babc = 0. (default = $1000.0 \text{ Pa yr m}^{-1}$)
ho_beta_small †	(real) small beta for sliding over a thawed bed when $\mbox{which_ho_babc} = 1$. (default = 1000.0 Pa yr m ⁻¹)
ho_beta_large †	(real) large beta to enforce (virtually) no slip when $\mbox{which_ho_babc} = 1$. (default = 1e10 Pa yr m ⁻¹)
friction_powerlaw_k †	(real) friction coefficient k used for which_ho_babc = 9 (Eq. 5.33) (default = 8.4e-9 m y ⁻¹ Pa ⁻² , from Bindschadler (1983) converted to CISM units)
coulomb_c †	(real) Coulomb friction coefficient (unitless), C , used for which_ho_babc = 10 (Eq. 5.34) (default = 0.42, from Pimentel et al. (2010))
coulomb_bump_max_slope †	(real) maximum slope (unitless) of the dominant bedrock bumps, m_{max} , used for which ho_babc = 10 (Eq. 5.34) (default = 0.5 m, from Pimentel et al. (2010))
coulomb_bump_wavelength †	(real) wavelength (m) of the dominant bedrock bumps, λ_{max} , used for which_ho_babc = 10 (Eq. 5.34) (default = 2.0 m, from Pimentel et al. (2010))
flwa_basal	(real) Glen's A at the bed for Schoof (2005) Coulomb friction law (default = 1e-16 Pa $^{-n}$ yr $^{-1}$)
powerlaw_c	(real) friction coefficient in power law (default = 1e4 Pa $\rm m^{-1/3}~\rm yr^{1/3})$
powerlaw_m	(real) exponent in power law (default $= 3$)
beta_powerlaw_umax	(real) upper limit of ice speed when evaluating powerlaw beta (default = 0 m yr^{-1}). Where u > umax, let u = umax when evaluating beta(u)
pseudo_plastic_q	(real) exponent for pseudo-plastic law (default = 0.5 unitless).
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continued from previous page	Note: $0 \le q \le 1$; $q = 1 \Longrightarrow$ linear sliding law; $q = 0$	
	Note. 0 <= q <= 1, q = 1 => linear shding law, q = 0	
pseudo_plastic_u0	(real) threshold velocity for pseudo-plastic law (default =	
	100.0 m yr ⁻¹)	
pseudo_plastic_phimin	(real) min(phi) in pseudo-plastic law, for topg <=	
	pseudo_plastic_bedmin (default = 5.0 degrees, $0 < phi <$	
	90)	
pseudo_plastic_phimax	(real) max(phi) in pseudo-plastic law, for topg <=	
	pseudo_plastic_bedmin (default = 40.0 degrees, $0 < phi <$	
	90)	
pseudo_plastic_bedmin	(real) bed elevation below which phi =	
	pseudo_plastic_phimin (default = -700.0 m)	
pseudo_plastic_bedmax	(real) bed elevation above which phi =	
	pseudo_plastic_phimax (default = 700.0 m)	
Effective pressure parameters	, , , , , , , , , , , , , , , , , , , ,	
p_ocean_penetration	(real) p-exponent in ocean penetration parameterization for	
F	$(\text{basal_water} = 4 \text{ (default} = 0.0)$	
effecpress_delta	(real) multiplier for effective pressure N where the bed is	
01100p1088_u010u	saturated and/or thawed (default = 0.02)	
effecpress_bpmp_threshold	(real) basal temperature range over which N ramps from a	
ellechless_bbmb_thleshold	small value to full overburden (default = 0.1 deg C)	
offormage hml+ +hroshold	(real) basal melting range over which N ramps from a small	
effecpress_bmlt_threshold		
D 1 /	value to full overburden (default = 1e-3 m yr^{-1})	
Basal water parameters	(1)	
const_bwat	(real) constant basal water depth (default = 10.0 m)	
bwat_till_max	(real) maximum water depth in till (default = 2.0 m)	
c_drainage	(real) uniform drainage rate (default = 1e-3 m yr ⁻¹)	
Ocean data parameters		
gamma0	(real) (default 0.0)	
thermal_forcing_anomaly	(real) thermal forcing anomaly, applied everywhere (default	
	0.0 deg C)	
thermal_forcing_anomaly_tstart	(real) starting time for applying or phasing in the anomaly	
	(default 0.0 a)	
thermal_forcing_anomaly_timescale	(real) number of years over which the anomaly is phased	
	in linearly. If set to zero, then the full anomaly is applied	
	immediately (default 0.0)	
thermal_forcing_anomaly_basin	(int) basin where anomaly is applied (default 0 (apply to	
	all basins))	
Input topography adjustment parameters		
adjust_topg_xmin	(real) minimum x value of the region with adjusted topog-	
	raphy (default 0.0 m)	
adjust_topg_xmax	(real) maximum x value of the region with adjusted topog-	
	raphy (default 0.0 m)	
adjust_topg_ymin	(real) minimum y value of the region with adjusted topog-	
	raphy (default 0.0 m)	
adjust_topg_ymax	(real) maximum y value of the region with adjusted topog-	
J	raphy (default 0.0 m)	
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adjust_topg_no_adjust	(real) elevation beyond which there is no adjustment of topg
	(default 0.0 m)
adjust_topg_max_adjust	(real) elevation beyond which there is full adjustment of
	topg (default 0.0 m)
adjust_topg_delta	(real) maximum elevation adjustment (default 0.0 m). Ap-
	plied where topg $>$ adjust_topg_max_adjust
Anomaly parameters	
overwrite_acab_value	(real) acab value to apply in grid cells where over-
	$write_acab_mask = 1 (default 0.0 m yr^{-1} ice)$
overwrite_acab_minthck	(real) overwrite acab where thek <= over-
	write_acab_minthck (default 0.0 m)
acab_anomaly_timescale	(real) number of years over which the acab/smb anomaly
	is phased in linearly. If set to zero, then the anomaly is
	applied immediately (default 0.0 yr)
artm_anomaly_timescale	(real) number of years over which the artm anomaly is
	phased in linearly. If set to zero, then the anomaly is ap-
	plied immediately (default 0.0 yr)
bmlt_anomaly_timescale	(real) number of years over which the bmlt_float anomaly
	is phased in linearly. If set to zero, then the anomaly is
	applied immediately (default 0.0 yr)
Basal melting parameters	
bmlt_cavity_h0	(real) scale for reducing melting in sub-shelf cavities (de-
	fault = 0.0 m
	Note: this parameter similar to bmlt_float_h0, which is used
	specifically for MISMIP+
bmlt_float_omega	(real) adjustment factor for external bmlt_float field (de-
	fault = 1.0)
bmlt_float_h0	(real) scale for sub-shelf cavity thickness (default $= 75 \text{ m}$)
bmlt_float_z0	(real) scale for ice draft, relative to sea level (default $=$ -100
	m)
bmlt_float_const	(real) constant melt rate (default = 0 m a^{-1})
bmlt_float_xlim	(real) scale for ice draft, relative to sea level (default $= 0$
	m)
bmlt_float_depth_meltmax	(real) maximum melt rate at depth (default = 10 m a^{-1})
bmlt_float_depth_frzmax	(real) maximum freezing rate near surface (default $= 0$ m
-	a^{-1}
bmlt_float_depth_zmeltmax	(real) depth below which bmlt_float =
-	bmlt_float_depth_meltmax (default = -500 m)
bmlt_float_depth_zmelt0	(real) depth where $bmlt_float = 0$ (default = -200 m)
bmlt_float_depth_zfrzmax	(real) depth above which bmlt_float = -
-	bmlt_float_depth_frzmax (default = -100 m)
bmlt_float_depth_meltmin	(real) minimum melt rate in warm ocean (default = 0 m
•	a ⁻¹) (only applies where the field warm_ocean_mask=1)
bmlt_float_depth_zmeltmin	(real) depth above which bmlt_float =
1	bmlt_float_depth_meltmin (default = 0 m) (only ap-
	plies where the field warm_ocean_mask=1)
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Basal inversion parameters	
inversion_thck_flotation_buffer	(real) if usrf_obs implies H near the flotation thickness, set to thck_flotation +/- inversion_thck_flotation_buffer (default = 1.0 m)
inversion_thck_threshold	(real) removes ice thinner than this threshold (default = 0.0 m)
powerlaw_c_max	(real) maximum value of powerlaw_c (default = 10^5 Pa (m a^{-1}) $^{(-1/3)}$
powerlaw_c_min	(real) minimum value of powerlaw_c (default = 10^2 Pa (m a^{-1}) $^{(-1/3)}$
inversion_babc_timescale	(real) inversion timescale (default = 500.0 a); must be > 0
inversion_babc_thck_scale	(real) thickness inversion scale (default = 100.0 m); must be > 0
inversion_bmlt_timescale	(real) time scale for relaxing toward observed thickness (default = 0.0 yr)
inversion_bmlt_max_melt	(real) maximum melting rate allowed from inversion; ignored when set to 0 (default = 0.0 m a^{-1})
inversion_bmlt_max_freeze	(real) maximum freezing rate allowed from inversion; ignored when set to 0 (default = 0.0 m a^{-1})
inversion_nudging_factor_min	(real) minimum value of nudging factor between wean_tstart and wean_tend (default $= 0.0$)
inversion_wean_bmlt_float_tstart	(real) starting time for weighted nudging of bmlt_float (default = 0.0 a)
inversion_wean_bmlt_float_tend	(real) end time for weighted nudging of bmlt_float (default = 0.0 a)
inversion_wean_bmlt_float_timescale	(real) time scale for weaning of bmlt_float (default = 0.0)
inversion_dbmlt_dtemp_scale	(real) scale for rate of change of bmlt w.r.t temperature (default = 10.0 m/yr/deg)
inversion_bmlt_basin_timescale	(real) timescale for adjusting deltaT_basin (default = 10.0 a)
<pre>inversion_bmlt_basin_flotation_threshold</pre>	(real) threshold for counting ice as lightly floating/grounded (default = 500.0 m)
inversion_bmlt_basin_mass_correction	(real) optional mass correction for a selected basin (default = 0 Gt)
$inversion_bmlt_basin_number_mass_correction$	(integer) ID for the basin receiving the correction (default $= 0$)
ISMIP-HOM parameters	
periodic_offset_ew†	(real) vertical offset between east and west edges of the global domain. (default = 0.0 m) (Primarily used for ISMIP-HOM and Stream test cases.)
periodic_offset_ns†	(real) vertical offset between north and south edges of the global domain. (default = 0.0 m) (Primarily used for ISMIP-HOM and Stream test cases.)
[GTHF]	
Options related to lithospheric temperature and unless $gthf = 1$.	geothermal heat calculation. Ignored
num_dim	can be either 1 for 1D calculations or 3 for 3D calculations.
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nlayer	number of vertical layers (default: 20).
surft	initial surface temperature (default 2°C).
rock_base	depth below sea-level at which geothermal heat gradient is
	applied (default: -5000m).
numt	number time steps for spinning up GTHF calculations (de-
	fault: 0).
rho	The density of lithosphere (default: 3300kg m^{-3}).
shc	specific heat capcity of lithosphere (default: $1000 \mathrm{J} \ \mathrm{kg}^{-1}$
	K^{-1}).
con	thermal conductivity of lithosphere (3.3 W m^{-1} K ⁻¹).
[isostasy]	

[isostasy]		
Options related to isostasy model. Ignored unless isostasy = 1. Options marked with a		
* work only with a serial build (or a parallel build if run on 1 processor).		
lithosphere	0 local lithosphere, equilibrium bedrock depression is	
	found using Archimedes' principle	
	1* elastic lithosphere, flexural rigidity is taken into ac-	
	count	
asthenosphere	0 fluid mantle, isostatic adjustment happens instantaneously	
	1 relaxing mantle, mantle is approximated by a half-space	
whichrelaxed	0 relaxed topography is read from a separate input variable,	
	$\operatorname{rel}\mathbf{x}$	
	1 first time slice of input topography is assumed to be relaxed	
	2 first time slice of input topography is assumed to be in	
	isostatic equilibrium with ice thickness	
relaxed_tau	characteristic time constant of relaxing mantle (default:	
	4000.a)	
lithosphere_period	lithosphere update period (default: 100.a)	
flexural_rigidity	flexural rigidity of the lithosphere (default: $0.24e25 \text{ Pa m}^3$)	

[projection]		
Specify map projection for reference. The reader is referred to Snyder J.P. (1987) Map		
Projections - a working	Projections - a working manual. USGS Professional Paper 1395.	
type	string that specifies the projection type (LAEA, AEA, LCC or	
	STERE).	
centre_longitude	(real) central longitude $(default = 0.0 degrees East)$	
centre_latitude	(real) central latitude $(default = 0.0 degrees North)$	
false_easting	(real) false easting $(default = 0.0 m)$	
false_northing	(real) false northing (default $= 0 \text{ m}$)	
standard_parallel	location of standard parallel(s) in degrees north. Up to	
	two standard parallels may be specified (depending on the	
	projection).	
scale_factor	non-dimensional; relevant only for the stereographic pro-	
	jection	
compute_area_factor	false does not compute area distortion factors	
	True compute area distortion factors	

7.4 Input/Output Configuration

NetCDF I/O can be configured in the main configuration file or in a separate file (see ioparams in the [options] section description above). Any number of input, forcing, and output files can be specified.

Input files are processed in the same order they appear in the configuration file, thus potentially overwriting previously loaded fields. The configuration section for an input file specifies which time slice from the input file should be used as the initial condition. (This is an integer specifying the time level, not the actual time.)

Input files can contain any of a large number of model fields that are specified as being "loadable". These fields are marked with asterisks in the netCDF model variable table in Appendix A. The option for a given variable to be "loadable" or not can be changed within the *_vars.def files (most commonly, glide_vars.def), as described in Appendix B.1. The standard test cases discussed in Chapter 8 give examples of variables that might be specified for standard model setups.

CISM2.1 can run with prescribed forcing files. These are input files that are read on every time step to allow time-dependent forcing to be applied during a simulation. Any input fields specified in *_vars.def (again, most commonly, glide_vars.def) can be included in forcing files. Forcing files should have a "time" field which is used to assign values to each field in the file during the simulation. Forcing is applied in a piecewise constant fashion; the most recent time slice in the forcing file prior to the current model time is used on each time step. (Linear interpolation of forcing may be available in the future but is not yet implemented.) If a field is present in both an input file and a forcing file at the start time, the value in the forcing file will overwrite the value from the input file because forcing files are read after input files. Forcing files are processed in the same order they appear in the configuration file on each time step, thus potentially overwriting previously loaded fields from other forcing files. The "dome" test case (tests/dome) includes an optional setup of how to implement time-dependent forcing.

One special input field is the kinbcmask field, which is used to specify locations (i.e., indices in the horizontal grid plane) on the staggered (velocity) grid where Dirichlet boundary conditions are to be applied. This input field is used in conjunction with the uvel and vvel fields. At any location where kinbcmask = 1, the input field values of uvel and vvel for that same column will be applied as Dirichlet boundary conditions on the velocity solution. If uvel and vvel are not included in the input file, zero velocity will be specified throughout the column for any location where kinbcmask is set to 1. An example of this application can be seen in the confined shelf test case described in Chapter 8.2. The construction of the relevant input kinbcmask field is done by the python script that constructs the other input fields for this test case (the shelf-confined.py script). Also, the dome test case described in Chapter 8.2 includes a script and a config file demonstrating how kinbcmask, uvel, and vvel can be used in a forcing file to provide time-varying Dirichlet boundary conditions.

The following tables describe netCDF sections and parameters in the .config file.

[CF default]	
This section contains metadata describing the experiment. Any of these parameters can	
be optionally included in the [CF output] section for a specific output file, overriding the	
default values	specified here.
title	Title of the experiment
institution	Institution at which the experiment was run
references	References that might be useful
comment	Comments further describing the experiment
[CF input]	
Any number of input files can be specified in separate [CF input] sections. They are pro-	
cessed in the order they appear in the configuration file, potentially overwriting previously	
loaded variables with the same names from previous input files.	
name	Name of the netCDF file to be read. Typically, netCDF files end with
	.nc.
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time	Time slice (not actual time) to be read from the netCDF file. The first	
	time slice is read by default.	
[CF forcin	g]	
	r of forcing files can be specified in separate [CF forcing] sections. They	
_	d in the order they appear in the configuration file, potentially overwriting	
	paded variables. Each forcing file needs a "time" dimension and variable that	
-	e model time associated with each time slice in the file.	
name	Name of the netCDF file to be read. Typically, netCDF files end with	
	.nc.	
[CF output		
This section	controls how often selected variables are written to files.	
name	Name of the output netCDF file. Typically, netCDF files end with .nc.	
start	(real) Start writing to file when this time (years) is reached (default:	
	first time slice).	
stop	(real) Stop writing to file when this time (years) is reached (default: last	
fmaguanar	time slice). (real) The time interval in years, determining how often selected vari-	
frequency	ables are written to file.	
xtype	Set the floating point representation used in netCDF file. xtype can be	
JF-	one of real, double (default: real). (If the name restart is included	
	in the list of variables, xtype is automatically set to double to ensure	
	bit-reproducible restarts, overriding the value set here.)	
variables	List of variables to be written to file. See Appendix A for a list of	
	possible variables. Names should be separated by at least one space. The	
	variable names are case-sensitive. The name restart selects all variables	
	necessary for a restart based on the specified model configuration. (The	
	name hot is also retained for this purpose for backwards compatability.)	
	If the name restart is included, the xtype option is automatically set	
	to double to ensure bit-reproducible restarts.	

7.5 Climate Forcing Configuration

The core ice sheet model is connected to the climate via the surface mass balance (acab field) and air temperature (artm field) and (optionally) a scalar value for eustatic sea level. This climate forcing can come from:

- climate schemes included in the CISM code:
 - EISMINT 1 and 2
 - annual and daily PDD schemes
- data input into the model:
 - directly as user-supplied acab and artm fields
 - from a global or regional climate model (e.g., CESM, GENIE, either from data or coupled) throught the Glint interface

7.5.1 EISMINT climate forcing

Like previous versions of Glimmer, CISM includes a set of idealized climate forcings used in the European Ice Sheet Modeling INiTiative (EISMINT) Phase 1 and 2 series of experiments. These

forcings consist of surface mass balance and air temperature fields for predefined experiments. See Chapter 8 for details of how to run the individual tests, and see the EISMINT publications for a more detailed description of the tests and the forcings associated with each. Huybrechts et al. (1996) describe EISMINT Phase 1, and Payne et al. (2000) describe EISMINT Phase 2.

Configuration

The various EISMINT climate forcings are enabled by adding one of the following sections to the configuration file used for running CISM. See the files associated with the EISMINT test cases (Chapter 8) for examples of their use for each of the specific experiment setups described in Huybrechts et al. (1996) and Payne et al. (2000).

[EISMINT-1 fi	xed margin]	
EISMINT 1 fixe	EISMINT 1 fixed margin scenario. Some of the EISMINT-1 fixed margin tests use periodic,	
time-varying for	time-varying forcing.	
temperature	(real(2 values)) Temperature forcing	
massbalance period	$T_{\rm surface}=t_1+t_2d$ where $d=\max\{ x-x_{\rm summit} , y-y_{\rm summit} \}$ (real) Mass balance forcing (real) Period of time–dependent forcing (switched off when set to 0) $\Delta T=10\sin\frac{2\pi t}{T}$	
	and $\Delta M = 0.2 sin \frac{2\pi t}{T}$	
mb_amplitude	(real) Amplitude of the surface mass balance when $\mathtt{period} > 0$	

[EISMINT-1 moving margin]

EISMINT 1 moving margin scenario. Some of the EISMINT-1 moving margin tests use periodic, time-varying forcing.

temperature	(real(2 values)) Temperature forcing
	$T_{ m surface} = t_1 - t_2 H$
	where H is the ice thickness
massbalance	(real(3 values)) Mass balance forcing
	$M = \min\{m_1, m_2(m_3 - d)\}$
	where $d = \sqrt{(x - x_{\text{summit}})^2 + (y - y_{\text{summit}})^2}$
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period	(real) Period of time–dependent forcing (switched off when set to 0)
	$\Delta T = 10 \sin \frac{2\pi t}{T}$
	and $M = \min \left\{ m_1, m_2 \left(m_3 + 100 sin \frac{2\pi t}{T} - d \right) \right\}$
mb_amplitude	(real) Amplitude of the surface mass balance when $\mathtt{period} > 0$
[EISMINT-2]	

[EISMINT-2]		
EISMINT 2 climate forcing. Both surface mass balance and air temperature depend solely		
on position in t	on position in the map plane and not on ice-surface elevation.	
temperature	(real(2 values)) Temperature forcing	
	$T_{\rm surface}=t_1-t_2d$ where d is the distance from the summit, $d=\sqrt{(x-x_{\rm summit})^2+(y-y_{\rm summit})^2}$	
massbalance	(real(3 values)) Mass balance forcing	
	$M = \min\{m_1, m_2(m_3 - d)\}$	
	where d is the distance from the summit,	
	$d = \sqrt{(x - x_{\text{summit}})^2 + (y - y_{\text{summit}})^2}$	

7.5.2 Glint driver

Glint (originally an acronym for "Glimmer interface") allows CISM to be run with forcing from an external climate model or global data sets. Glint was originally developed as an interface between Glide and the GENIE Earth-system model, but is designed to be flexible enough to be used with a wide range of global climate models. In older versions of Glint, it was assumed that forcing from the climate model would include the temperature and precipitation fields required to force a positive-degree-day (PDD) scheme for the surface mass balance (SMB). In CISM2, PDD forcing is still supported, but Glint can also receive the SMB (typically computed in multiple elevation classes on the relatively coarse grid of a climate model) and downscale it directly to the ice sheet model grid.

A distinctive feature of Glint is the way it uses the object-oriented Glide architecture to enable multiple ice models to be coupled to the same climate model. This means that regional ice models can potentially run at high resolution over several parts of the globe (e.g., Greenland and Antarctica), without the expense of running a global ice sheet model.

Glint automates the processes required in coupling regional models to a global model, particularly the downscaling and upscaling of the fields that form the interface between the two models. The user may specify map projection parameters for each of the ice sheet models (known as *instances*). The different time steps of the global model, mass-balance scheme, and ice sheet model are handled automatically by temporal averaging or accumulation of quantities as appropriate.

Prerequisites

Glint users should bear the following in mind:

- Global input fields must be supplied on a latitude-longitude grid. The grid does not have to be uniform in latitude, meaning that Gaussian grids may be used. Irregular grids (e.g., icosahedral grids) are not currently supported. The boundaries of the grid boxes may be specified; if not, they are assumed to lie halfway between the grid points in lat-lon space.
- In the global field arrays, latitude must be indexed from north to south. That is, the first row of the array is the northernmost one. (Some flexibility might be introduced here in the future.)
- The global grid must not have grid points at either of the poles. This restriction is not expected to be permanent, and in the meantime can probably be overcome by moving the location of the polar points to be fractionally short of the pole (e.g. at 89.9° and -89.9°).

Initializing and calling

The easiest way to learn how Glint is used is by way of an example. Glint should be built automatically as part of CISM, and we assume here that this has been done successfully.

Typically, Glint will be called from the main program body of a climate model. (In CESM, Glint subroutines are called from a wrapper layer called **glc**.) To make this possible, the compiler needs to be told to use the Glint code, with "use" statements like the following:

```
use glint_main
```

The next task is to declare a variable of type glint_params, which holds everything relating to the model, including any number of ice-sheet instances:

```
type(glint_params) :: ice_sheet
```

Before the ice-sheet model may be called from the climate model, it must be initialized. This can be done with the following subroutine call³:

```
call initialise_glint(ice_sheet,lats,lons,time_step,paramfile)
```

These are the required arguments; many optional arguments are also possible. The required arguments are defined as follows:

- ice_sheet is the variable of type glint_params defined above;
- lats and lons are one-dimensional arrays giving the locations of the global grid-points in latitude and longitude, respectively;
- time_step is the intended interval between calls to Glint, in hours. This is known as the forcing timestep.
- paramfile is the name of the Glint configuration file. The contents of this file are discussed below.

If Glint is to be forced with an externally computed surface mass balance (rather than the fields that would drive a PDD scheme), the climate model should instead call the subroutine initialise_glint_gcm, which has the same required arguments but a different set of optional arguments.

After Glint is initialised, it may be called as part of the main climate model time-step loop:

```
call glint(ice_sheet,time,temp,precip,orog)
```

 $^{^3}$ The spelling of some subroutine names reflects the British origins of the code.

where

- ice_sheet is the variable of type glint_params defined above;
- time is the current model time, in hours;
- temp is the daily mean 2 m global air temperature field, in °C;
- precip is the global daily accumulated precipitation field, in mm (water equivalent, making no distinction between rain, snow, etc.);
- orog is the global orography field, in m.

Many optional arguments may also be specified.

The latter three compulsory fields are needed to drive a PDD scheme. Glint includes two such schemes. One of these calculates the mass-balance for the whole year (the annual PDD scheme), while the other calculates on a daily basis (the daily PDD scheme). The annual scheme incorporates a stochastic temperature variation to account for diurnal and other variations, which means that if this scheme is used, Glint should be called such that short-term variations have been removed. In practice, this means calling Glint on a monthly basis, with monthly mean temperatures. For the daily scheme, no such restriction exists, and the scheme should be called at least every 6 hours.

If the SMB is computed externally by a climate model, the call to Glint would resemble this one:

```
call glint_gcm(ice_sheet,time,qsmb,tsrf,topo)
```

where the last three arguments (all compulsory) are defined as follows:

- qsmb is the surface mass balance in kg m^{-2} s⁻¹;
- tsrf is the surface ground temperature in ${}^{\circ}C$, used as an upper thermal boundary condition for the ice sheet; and
- topo is the surface elevation in m.

These fields contain two horizontal dimensions, along with a third dimension for elevation class. CESM typically computes these quantities in 10 elevation classes per glaciated grid cell in the climate model.

Among the optional arguments are the following output fields (returned by Glint to the climate model):

- gfrac is the fractional grid cell area covered by ice;
- gtopo is the mean surface elevation in m;
- ghflx is the upwelling heat flux at the ice sheet surface, in W m²;
- grofi is the solid runoff flux (i.e., the calving flux) in kg m^{-2} s⁻¹;
- grof1 is the liquid runoff flux in kg m⁻² s⁻¹. This includes basal and possibly internal melting, but not surface melting (which has already been computed by the climate model).

These fields are required by CESM to update its land-surface types and topography and to conserve heat and water when CISM is coupled interactively to the climate model. The first three fields are computed for each elevation class in each climate model grid cell, and the two runoff fluxes are averaged over each grid cell. The initial "g" denotes that these fields have been upscaled to the global climate grid.

Finishing off

After the desired number of time steps have been run, Glint may have some tidying up to do. To do this, the subroutine end_glint must be called:

call end_glint(ice_sheet)

Configuration

Glint uses the same configuration file format as the rest of CISM. If there is only one ice sheet instance, all the configuration data for Glint and Glide (or Glissade) can reside in the same file. If two or more instances are used, a top-level file specifies the number of model instances and the name of a configuration file for each one. Configuration sections specific to Glint are as follows:

[Glint]	
Section specifying number of instances.	
n_instance	(integer) Number of instances (default=1)
[Glint instance]	
Specifies the name	of an instance-specific configuration file. Unnecessary if we only have one
instance whose cor	afiguration data is in the main config file.
name	Name of instance-specific config file (required).
[Glint climate]	
Glint climate confi	guration.
evolve_ice	specify whether or not the ice sheet evolves in time: 0 Do not evolve ice sheet (hold fixed in time). This setting is appropriate if we want to analyze the SMB downscaled to the observed ice sheet geometry, without the complexity of evolution. 1 Allow the ice sheet to evolve
precip_mode	Method of precipitation downscaling: 1 Use large-scale precipitation rate 2 Use parameterization of Roe and Lindzen
acab_mode	Mass-balance model to use: 0 The surface mass balance is computed externally (e.g., by a climate model) 1 Annual PDD mass-balance model (Section 7.7.1) 2 Annual accumulation only 3 Hourly energy-balance model (NOTE: not supported) 4 Daily PDD mass-balance model (Section 7.7.2)
ice_albedo lapse_rate	Albedo of ice—may be used for coupling to climate model (default=0.4) Atmospheric temperature lapse-rate, used to correct the atmospheric temperature onto the ice model orography. This should be <i>positive</i> for temperature falling with height (K km ⁻¹) (default=8.0)
data_lapse_rate	Atmospheric temperature vertical lapse rate, to be used in the calculation of temperature at sea level. The variable lapse_rate is then used to adjust the temperature to the surface of the local ice sheet topography. If data_lapse_rate is not set, it is set to the value of lapse_rate by default. **continued on next page**
	communication next page

continued from previous page		
ice_tstep_multiply	Ice time-step multiplier: allows asynchronous climate-ice coupling. See	
	below for full explanation of Glint time-stepping. $(default = 1)$	
mbal_accum_time	Mass-balance accumulation time (in years, default is equal to mass-	
	balance or ice time step, whichever is larger). See below for full ex-	
	planation of Glint time-stepping.	

CISM can run with climate forcings defined on a global array without being coupled to a climate model. With this configuration a second configuration file relative to the global forcing needs to be provided. CISM would then be called using 2 configuration files, one specific to the ice sheet model configuration properties and one relative to the climate forcing properties. The configuration options specific to global forcings are as follow:

[GLEX climate]	
Specifies the global climate forcing properties.	
days_in_year	(integer) Number of days in a year available in the forcing file (default:
	365)
total_years	(integer) Length of run in years (default: 10)
climate_tstep	(integer) Climate time step in hours (default: 6)
gcm_smb	(logical) If .true., pass SMB instead of PDD (default: .false.)
[GLEX precip]	
Specifies the enti	ry of global precipitation forcing.
filename	File name containing precipitation data (required).
variable	(unit: mm/s) Name of variable to be read from the filename containing
	precipitation data (required).
[GLEX temps]	
Specifies the enti	ry of global temperature forcing.
filename	File name containing precipitation data (required).
variable	Name of variable to be read from the filename containing temperature
	data (required).
kelvin	(logical) Temperature assumed to be in kelvin if .true. or in degrees C
	if .false. (default: true).
[GLEX orog]	
Specifies the entry of global orography.	
filename	File name containing orography (required).
variable	(unit: m) Name of variable to be read from the filename containing
	orography data (required).
	continued on next page

When running with glint, the simulation time span is managed by the configuration entry "total_years" in [GLEX climate] defined in the global forcing configuration file. Consequently, the time specific entries "tstart" and "tend" in the "[time]" entry of the model configuration file do not have any impact on the simulation. A specific example on how to run Glint with global forcings is the subject of section 7.6.

Glint timestepping — an explanation

By default, the model accepts input on each forcing timestep (as specified in the call to initialise_glint). Input fields are accumulated over the course of a mass-balance timestep, whereupon the mass-balance model is called. The output from the mass-balance model is accumulated over the course of an ice sheet model time-step, and finally the ice sheet model (Glide or Glissade) is called.

This default behaviour can be altered, in two ways:

- 1. The number of ice sheet time steps executed for each accumulated mass-balance field may be increased—thus accelerating the ice sheet relative to the forcing. To do this, set ice_tstep_multiply in the [Glint climate] config section; this must be an integer. This acceleration is possible only if the mass balance is accumulated over an integer number of years.
- 2. The mass-balance accumulation period can be altered by setting mbal_accum_time in the [Glint climate] config section; this is a floating-point value in years.

The interaction of these two parameters is fairly complex, and permits sophisticated control of how the ice sheet model is forced. Various checks are made at run-time to make sure that sensible values are selected. Most importantly, all relevant time-steps must divide into one another appropriately. The model will (or should) stop if an un-sensible combination of values is detected.

When using global climate forcing the first time stamp in a forcing file should correspond to the climate time step value entered in [GLEX climate] section of the global configuration file. The remaining entries in the forcing file should have entries defined at a time interval equal to that of the climate time step as well.

Glint timestepping — further examples

To aid understanding of the time-stepping controls, here are some examples. First, suppose we have these time-step values:

forcing time-step: 6 hours mass-balance time-step: 1 day ice time-step: 0.5 year

By default, the model will accumulate 6 months of mass-balance calculations, and force the ice sheet model based on the 6-month average. This might not be desirable, so you could set:

```
mbal_accum_time = 1.0
```

This would make Glint accumulate 1 year of mass-balance output before forcing the ice sheet (at which point it would execute *two* ice sheet time-steps of 0.5 years each).

Having done that, you could accelerate the ice model by a factor of ten, by setting

```
ice_tstep_multiply = 10
```

In this scenario, 20 ice sheet time-steps of 0.5 years each would be done after each 12-month accumulation of mass-balance data.

For the second example, we consider the contrasting situation where we do not want to calculate a mass balance on all the available data (perhaps to save time). Consider these timestep values:

forcing time-step: 6 hours mass-balance time-step: 1 day ice time-step: 10 years

(Clearly this is a numerically stable and/or low-resolution ice sheet). To avoid running the daily PDD scheme c.3600 times (depending on the value of days_in_year), we can choose to use only the first two years of data:

```
mbal_accum_time = 2.0
```

Glint accumulates the mass balance for 2 years, then waits for 8 years (ignoring incoming data during this time) before calling the ice sheet. Ice sheet acceleration may be enabled with ice_tstep_multiply as before.

7.6 Glint: Using glint-example

To run CISM using the Glint climate driver, go to directory tests/glint_example and see the README.md file. This file gives directions for downloading a tar file, glint-example.1.0.0.tar.gz, that contains three data files used to force Glint:

- ncep-doe_6h_climate.64x32.nc consists of 6-hourly precipitation and surface temperature data on a coarse global grid;
- orog.igcmgrid.64x32.nc consists of surface elevation data on a global grid (these data are used to downscale temperature as a function of elevation); and
- gland20.input.nc includes thickness and topography data for the Greenland ice sheet on a 20-km mesh.

These files should be placed in directory tests/glint_example. This directory already includes the following config files:

- greenland_20km.config.pdd and greenland_20km.config.smb are ice sheet configuration files for PDD and SMB forcing, respectively;
- glint_example.config.pdd and glint_example.config.smb are the corresponding climate configuration files.

To apply the PDD scheme, type

./cism_driver greenland_20km.config.pdd glint-example.config.pdd and to apply the SMB scheme, type

./cism_driver greenland_20km.config.smb glint-example.config.smb

The 6-hourly temperature and precipitation data are tailored for a PDD scheme. The same data, however, can also supply a crude surface mass balance and surface temperature in multiple elevation classes, imitating the forcing from a climate model such as CESM. (Accumulation is set equal to precipitation, and ablation is assumed to be linearly related to the downscaled temperature. This scheme is convenient for testing but obviously is not appropriate for science.) By default, two files are output at a specified frequency: a history file containing variables of interest (e.g., thickness, velocity, and SMB) is output every 10 years, and a "hot" file containing the variables required for exact restart is output every 1000 years.

7.7 Supplied mass-balance schemes

Users are free to supply their own mass-balance model for use with Glide. However, Glint includes two positive-degree-day models for mass balance, one annual and one daily. This section describes how to configure and call these models.

NOTE: CISM2 development has focused on CESM-style coupling where the SMB is computed externally. The two PDD schemes (and the following description) are identical to those in the original Glimmer code, but the schemes have not been extensively tested. Users should proceed with caution.

7.7.1 Annual PDD scheme

The annual PDD scheme is contained in the f90 module glimmer_pdd, and the model parameters are contained in the derived type glimmer_pdd_params. Configuration data is contained in a standard CISM config file, which needs to be read before initializing the mass-balance model. The model is initialized by calling the subroutine glimmer_pdd_init, and the mass-balance may be calculated annually by calling glimmer_pdd_mbal.

Example of use:

```
use glimmer_pdd
use glimmer_config
....

type(glimmer_pdd_params) :: pdd_scheme
type(ConfigSection),pointer :: config
....

call glimmer_pdd_init(pdd_scheme,config)
....

call glimmer_pdd_mbal(pdd_scheme,artm,arng,prcp,ablt,acab)
```

In the subroutine call to glimmer_pdd_mbal, apart from the parameter variable pdd_scheme, there are three input fields (artm, arng and prcp), which are, respectively, the annual mean air temperature, annual temperature half-range, and annual accumulated precipitation fields. The final two arguments are output fields — annual ablation (ablt) and annual mass-balance (acab). Temperatures are in degrees Celsius, and precipitation, ablation and mass-balance are measured in m of water equivalent.

Degree-day calculation

The greater part of the information in the glimmer_pdd_params derived type comprises a look-up table (the *PDD table*). The model is implemented this way for computational efficiency.

The table has two dimensions: mean annual air temperature (T_a) (as the second index) and annual air temperature half range (i.e., from July's mean to the annual mean ΔT_a) (as the first index). Following *Huybrechts and others* [1991], daily air temperatures (T'_a) are assumed to follow a sinusoidal cycle

$$T_a' = T_a + \Delta T_a \cos\left(\frac{2\pi t}{A}\right) + \mathbf{R}(0, \sigma) \tag{7.1}$$

where A is the period of a year and R is a random fluctuation drawn from a normal distribution with mean 0 °C and standard deviation σ °C. Huybrechts and others [1991] indicate that the number of positive degree days (D, °C days) for this temperature series can be evaluated as

$$D = \frac{1}{\sigma\sqrt{2\pi}} \int_{0}^{A} \int_{0}^{T_a' + 2.5\sigma} T_a \times \exp\left(\frac{-(T_a - T_a')^2}{2\sigma^2}\right) dT dt$$
 (7.2)

where t is time. The table is completed by evaluating this integral using a public-domain algorithm (Romberg integration) by Bauer [1961]. The inner and outer integrals are coded as

two subroutines (inner_integral and pdd_integrand), which call the Romburg integration recursively.

The main parameter needed is the assumed standard deviation of daily air temperatures, which can be set in the configuration file (the default is 5 °C).

The positive-degree days are then looked up in the table (as a function of T_a and ΔT_a). We take care to check that this look up is in done within the bounds of the table. The final value of P is determined using bilinear interpolation given the four nearest entries in the table to the actual values of T_a and ΔT_a .

The remainder of the loop completes the calculation of the ablation and accumulation given this value for P.

Mass balance calculation

We use the following symbols: a is total annual ablation; a_s is potential snow ablation; b_0 is the capacity of the snowpack to hold meltwater by refreezing; the total number of positive degree days (D); degree-day factors for snow and ice $(f_s \text{ and } f_i)$; and the fraction of snowfall that can be held in the snowpack as refrozen meltwater (W_{max}) . The degree-day factors have been converted from ice to water equivalents using the ratio of densities.

First, determine the depth of superimposed ice (b_0) that would have to be formed before runoff (mass loss) occurs as a constant fraction (W_{max}) of precipitation (P):

$$b_0 = W_{\text{max}} P. \tag{7.3}$$

Now determine the amount of snow melt by applying a constant degree-day factor for snow to the number of positive degree-days:

$$a_s = f_s D. (7.4)$$

We now compare the potential amount of snow ablation with the ability of the snow layer to absorb the melt. Three cases are possible. First, all snow melt is held within the snowpack and no runoff occurs (a=0). Second, the ability of the snowpack to hold meltwater is exceeded but the potential snow ablation is still less than the total amount of precipitation so that $a=a_s-b_0$. Finally, the potential snow melt is greater than the precipitation (amount of snow available), so that ice melt (a_i) has to be considered as well. The total ablation is therefore the sum of snow melt (total precipitation minus meltwater held in refreezing) and ice melt (from the total number of degree-days, deduct the number of degree-days needed to melt all snowfall and convert to ice melt):

$$a = a_s + a_i = P - b_0 + f_i \left(D - \frac{P}{f_s} \right).$$
 (7.5)

We now have a total annual ablation, and can find total net mass balance as the difference between the annual precipitation and the annual ablation.

This methodology is fairly standard and stems from a series of Greenland papers by Huybrechts, Letreguilly and Reeh in the early 1990s.

Configuration

The annual PDD scheme is configured using a single section in the configuration file:

[GL	[GLIMMER annual pdd]		
Spe	cifies parameters for the PDD table and mass-balance calculation		
dx	Table spacing in the x-direction ($^{\circ}$ C) (default=1.0)		
dy	Table spacing in the y-direction ($^{\circ}$ C) (default=1.0)		
ix	Lower bound of x-axis ($^{\circ}$ C) (default=0.0)		
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iy	Lower bound of y-axis ($^{\circ}$ C) (default=-50.0)	
nx	Number of values in x-direction (default=31)	
ny	Number of values in x-direction (default=71)	
wmax	Fraction of melted snow that refreezes (default=0.6)	
$pddfac_ice$	PDD factor for ice (m day $^{-1}$ °C $^{-1}$) (default=0.008)	
pddfac_snow	PDD factor for snow (m day ⁻¹ $^{\circ}$ C ⁻¹) (default=0.003)	

References

Bauer (1961) Comm. ACM 4, 255.

Huybrechts, Letreguilly and Reeh (1991) Palaeogeography, Palaeoclimatology, Palaeoecology (Global and Planetary Change) 89, 399-412.

Letreguilly, Reeh and Huybrechts (1991) Palaeogeography, Palaeoclimatology, Palaeoecology (Global and Planetary Change) 90, 385-394.

Letreguilly, Huybrechts and Reeh (1991) Journal of Glaciology 37, 149-157.

7.7.2 Daily PDD scheme

The other PDD scheme is a daily scheme. This is simpler than the annual scheme in that it does not incorporate any stochastic variations. The mass balance is calculated on a daily basis, given the daily mean temperature and half-range, and assuming a sinusoidal diurnal cycle. Consequently, the firn model is more sophisticated than with the annual scheme, and includes a snow-densification parameterization.

Configuration

The daily PDD scheme is configured using a single section in the configuration file:

[GLIMMER daily	[GLIMMER daily pdd]			
Specifies parameter	ers for the PDD table and mass-balance calculation			
wmax	Fraction of melted snow that refreezes (default=0.6)			
pddfac_ice	PDD factor for ice (m day $^{-1}$ °C $^{-1}$) (default=0.008)			
pddfac_snow	PDD factor for snow (m day $^{-1}$ °C $^{-1}$) (default=0.003)			
$rain_threshold$	Temperature above which precipitation is held to be rain (°C) (de-			
	fault=1.0)			
whichrain	Which method to use to partition precipitation into rain and snow:			
	1 Use sinusoidal diurnal temperature variation			
	2 Use mean temperature only			
tau0	Snow densification timescale (s) (default=10 years)			
constC	Snow density profile factor C (m ⁻¹) (default=0.0165)			
firnbound	Ice-firm boundary as fraction of density of ice (default=0.872)			
snowdensity	Density of fresh snow $(kg m^{-3})$ (default=300.0)			

Chapter 8

Test Cases

Test cases for CISM include experiments with analytic solutions, standardized experiments without analytic solutions but for which community benchmarks are available, and some experiments specific to CISM which have been well characterized by CISM developers.

Here we organize test cases based on the velocity solver that is most appropriate for each test. Any velocity solver can be used with any test if the .config file settings are adjusted manually. In some cases, however, the results may be difficult to interpret.

Each test directory includes a README.md file with some technical details about how to run the test. Many tests have python scripts that are used to set up the initial condition and, in some cases, execute the model. Some tests have an additional python script for analyzing the CISM output.

The user must manually provide each test with access to the CISM executable. There are several ways to do this:

• Softlink the executable into the directory, e.g.:

```
ln -s ../../builds/mac-gnu/cism_driver/cism_driver ./
```

This is the recommended procedure during development so that the test will always be using the most up-to-date version of the executable.

• Use the -e command line option to point to include an explicit path for the executable (for test case run scripts that support this option), e.g.:

```
./runDome.py -e ../../builds/mac-gnu/cism_driver/cism_driver
```

This is useful for quickly trying a different version of CISM (e.g., comparing serial and parallel executables).

- Add the directory containing the CISM executable to your environment PATH.
- Copy the executable into the directory. This is typically not the most efficient approach, but may make sense in some situations.

The python scripts generally have useful command line options that control their execution. Typically, you can see details by using the --help (or -h) command line option, e.g.:

```
./runDome.py --help
```

The various tests are described below.

8.1 Shallow-Ice Test Cases

These tests are primarily useful for testing the shallow-ice approximation (SIA) dynamical core, Glide (see Chapter 4). The Glissade dycore also has a shallow-ice option (see Section 6.2.2).

8.1.1 Halfar dome

The Halfar test case describes the time evolution of a parabolic dome of ice, as described by Halfar (1983). For a flat-bedded SIA problem, this case has an analytic solution for the time varying ice thickness. We start with the general SIA ice evolution equation,

$$\frac{\partial H}{\partial t} = \nabla \cdot (\Gamma H^{n+2} |\nabla H|^{n-1} \nabla H), \tag{8.1}$$

where n is the exponent in the Glen flow law, commonly taken as 3, and Γ is a positive constant:

$$\Gamma = \frac{2}{n+2} A(\rho g)^n. \tag{8.2}$$

For n = 3, the time-dependent solution is

$$H(t,r) = H_0 \left(\frac{t_0}{t}\right)^{\frac{1}{9}} \left[1 - \left(\left(\frac{t_0}{t}\right)^{\frac{1}{18}} \frac{r}{R_0}\right)^{\frac{4}{3}} \right]^{\frac{3}{7}}, \tag{8.3}$$

where

$$t_0 = \frac{1}{18\Gamma} \left(\frac{7}{4}\right)^3 \frac{R_0^4}{H_0^7},\tag{8.4}$$

and H_0 , R_0 are the central height of the dome and its radius at time $t = t_0$. For more details, see Halfar (1983), Bueler et al. (2005), and this link¹.

Provided files

Our implementation of the Halfar dome test has an initial radius of $R_0 = 21.2$ km and an initial thickness of H = 707.1 m. These values can be changed by editing the halfarDome function in runHalfar.py.

- README.md
 Information about the test case, including technical details about running it.
- halfar.config
 This is the config file defining CISM options. It is set up to run Glide.
- halfar-HO.config

 This alternative config file is set up to run Glissade using the Blatter-Pattyn approximation
 of Stokes flow. By manually setting which_ho_approx under [ho_options], users can
 choose other approximations (see Section 7.3).
- runHalfar.py
 This python script generates the dome initial condition and runs CISM.
- halfar_results.py
 This script compares model results to the analytic solution.

 $^{^{1}} http://www.giseis.alaska.edu/input/carl/teaching/extra/bueler_notes.pdf$

Running the test

One script sets up the initial condition and runs the model:

./runHalfar.py

Note that to run the test with the halfar-HO.config settings, you can use the -c commandline option for specifying a configuration file:

./runHalfar.py -c halfar-HO.config

Another script analyzes and plots the results:

./halfar_results.py

Results

With the default .config settings, this simulation should only take a few seconds and is a good first test for a working Glide dycore. With Glissade, the Blatter-Pattyn option takes a few minutes, but the SIA and L1L2 settings are much faster. As the dome of ice evolves, its margin advances and its thickness decreases (there is no surface mass balance to add new mass). The script halfar_results.py will plot the modeled and analytic thickness at a specified time (Figure 8.1), and also report error statistics. Invoke halfar_results.py --help for details on its use.

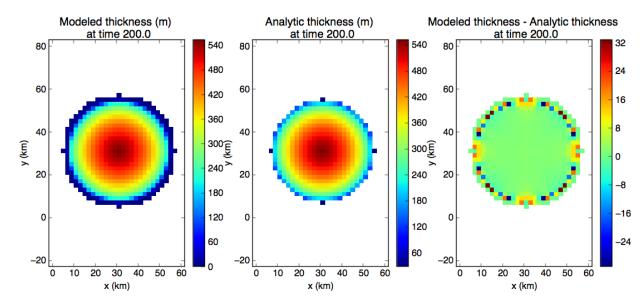


Figure 8.1: Halfar test case results (using Glide) after 200 years of dome evolution. This figure is generated by halfar_results.py.

8.1.2 EISMINT-1

This test case is from phase 1 of the European Ice Sheet Modelling INiTiative intercomparison experiments. These experiments are described in more detail here² and in Huybrechts et al. (1996).

Provided files

• README.md Information about the test case, including technical details about running it.

 $^{^2} http://homepages.vub.ac.be/~phuybrec/eismint.html\\$

• *.config

There are .config files for each of the six experiments in EISMINT 1: three fixed margin experiments (fm) and three moving margin experiments (mm).

Running the test

There is no script for running these experiments. They are simply run manually, e.g. using: ./cism_driver e1-fm.1.config

Results

These experiments are meant to be run to steady-state, and the supplied .config files are set up to run for long enough to do this. These simulations take more than a few minutes to complete. As the ice sheet evolves, its shape eventually equilibrates with the imposed surface mass balance. Currently there is no script for analyzing the model results. However, users can visually compare their results to those in Huybrechts et al. (1996).

8.1.3 EISMINT-2

This test case is from phase 2 of the European Ice Sheet Modelling INiTiative intercomparison experiments. These experiments are described in more detail here³ and in Payne et al. (2000).

Provided files

- README.md
 Information about the test case, including technical details about running it.
- *.config
 There are 11 .config files: one for each of the experiments described in Payne et al. (2000).
- mound.nc, trough.nc
 These are input netCDF files used by the EISMINT-2 experiments.

Running the test

There is no script for running these experiments. They are run manually, e.g. using: ./cism_driver e2.a.config

Results

These experiments are meant to be run to steady-state, and the supplied .config files are set up to run for long enough to do this. Some experiments use the final state of a previous experiment as the initial condition (e.g., most experiments following experiment A use the final state from A as an initial condition). See the experimental descriptions in Payne et al. (2000) for details. These simulations take more than a few minutes to complete. As the ice sheet evolves, its shape equilibrates with the imposed surface mass balance. There is no script for analyzing model results, but users can visually compare their results to those in Payne et al. (2000), and also compare model diagnostics (e.g., ice area, volume, and maximum thickness) to Table 4 of that paper.

8.1.4 Glint example

Section 7.6 explains how to set up and run the Glint example test case. We summarize that information here.

³http://homepages.vub.ac.be/~phuybrec/eismint.html

Provided files

- README.md
 - Information about the test case, including technical details about downloading the required data files and running the case.
- greenland_20km.config.pdd and glint_example.config.pdd

 These config files are used to run Glide with the surface mass balance computed by Glint's positive-degree-day scheme.
- greenland_20km.config.smb and glint_example.config.smb

 These config files are used to run Glide with the surface mass balance passed directly to
 Glint. (In real applications the SMB would be provided by a climate model, but here it
 is crudely approximated from temperature and precip data.)
- ncep-doe_6h_climate.64x32.nc, orog.igcmgrid.64x32.nc, gland20.input.nc
 These are data files that must be downloaded and placed in the glint_example test
 directory by the user. Users may also supply their own input data.

Running the test

There is no script for running these experiments. They are run manually, e.g. using: ./cism_driver greenland_20km.config.pdd glint_example.config.pdd

The two config files specify the ice sheet and climate configurations, respectively.

Results

The config files are set up to run Glide on a coarse Greenland grid for 10 model years, so that the default experiments complete quickly. The results are written to netCDF files. This test illustrates the application of CISM to a whole ice sheet but is not expected to be scientifically accurate.

8.2 Higher-Order Test Cases

The higher-order test cases are designed to test various aspects of higher-order dycores. By default they use the Glissade dycore with the Blatter-Pattyn approximation or DIVA (see Sections 6.2.1 and 6.2.4). As other higher-order dycores become available, they can be applied to these same tests. Additional test cases will be added as needed.

8.2.1 Dome

The dome test case is based on a parabolic dome of ice, similar to the Halfar test case. By default the dome has the same radius and center thickness as the Halfar case. However, it uses a simple square root function to define thickness as a function of distance from the dome center, resulting in a somewhat steeper profile. The dome test has been widely used for day-to-day testing because it is simple and relatively fast to run. It is a good test to confirm that basic higher-order model physics is working correctly, but does not strenuously test the physics and boundary conditions, or analytically verify the model.

Provided files

• README.md
Information about the test case, including technical details about running it.

- runDome.py
 The script to set up and run the test.
- dome.config

 The default configuration settings for running CISM with the test case.
- dome-forcing.config
 An example configuration script that can be used to generate a forcing file and run CISM with it.

Running the test

One script sets up the initial condition and runs the model:

./runDome.py
There is no script for analyzing the results.

The dome test case can also be used to set up an example of CISM's time-dependent forcing capability. (See Section 7.4 for details.) By passing the runDome.py script a config file which has [CF Forcing] section, as found in dome-forcing.config, it will generate a dome.forcing.nc file that contains time-varying fields and then run CISM using this generated netCDF file. Run:

./runDome.py -c dome-forcing.config

Results

There is not an analytic solution for this test, nor is there a script to analyze the results. You can manually inspect the results using a tool such as ncview.⁴ Sample output is shown in Figure 8.2.

 $^{^4 \}mathrm{See}$ section 2.2.4

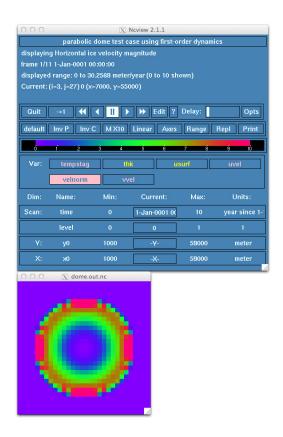


Figure 8.2: Dome velnorm (i.e., the ice speed in m/yr) field at time 0 using default dome.config settings. This figure is a screenshot of neview.

8.2.2 **ISMIP-HOM**

The Ice Sheet Model Intercomparison Project for Higher-Order Models (ISMIP-HOM) prescribes a set of experiments for testing the implementation of higher-order physics. For more information, see here⁵ and the ISMIP-HOM description paper by Pattyn et al. (2008).

The python scripts provided (runISMIP_HOM.py and plotISMIP_HOM.py, referred to here as the ISMIP-HOM scripts) were created to run experiments A through F using CISM and compare the results with results from other models.

Note: The README.md file gives many additional details about running and analyzing the test case that are not described here.

Provided files

- README.md Information about the test case, including technical details about running it.
- ismip-hom.config
 A default configuration file used as a template for generating the .config file for each test.
 If you wish to run the tests with different solver settings, for example, you should edit this file.
- runISMIP_HOM.py

⁵http://homepages.ulb.ac.be/~fpattyn/ismip/

The script used for running any/all of the ISMIP-HOM experiments. Invoke with '--help' to see the many command line options for controlling execution.

• plotISMIP_HOM.py

The script used for analyzing/plotting any/all of the ISMIP-HOM experiments. Invoke with '--help' to see the many command line options for controlling execution.

Running the test

One script sets up the initial condition and runs the model:

./runISMIP_HOM.py

and another is used to analyze the results:

./plotISMIP_HOM.py

Results

The plotISMIP_HOM.py script will plot results relative to other models. None of the ISMIP-HOM tests have a useful analytic solution, so these tests are used as community benchmarks rather than actual model verification tests. The Pattyn et al. (2008) paper is useful for interpreting model results. An example output plot is shown in Figure 8.3.

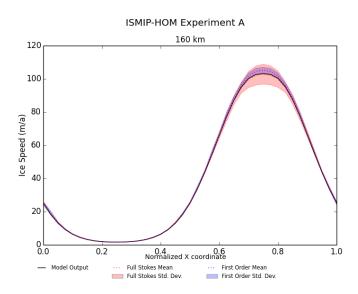


Figure 8.3: An example of the ISMIP-HOM test case output for test A at size L=160 km. CISM output is shown with a black line, and the range of output from other models is shown by colored bars. This figure was generated with ./plotISMIP_HOM.py -e a -s 160 after running the test with ./runISMIP_HOM.py -e a -s 160. Additional options (e.g., running and plotting for multiple tests simultaneously) are described in the README.md and by invoking the '--help' option at the command line.

8.2.3 Stream

The stream test case simulates flow over an idealized ice stream underlain by a subglacial till with a known and specified yield stress distribution (see discussion in Section 5.3). For the two yield stress distributions specified in this test case, analytical solutions are available from Raymond (2000) and Schoof (2006).

For the Raymond test case, the yield stress within the ice stream is given a uniform value below the driving stress, and outside the ice stream it is given a uniform value much higher than the driving stress (i.e., the yield stress distribution is approximated by a step function). For the School test case, the till yield stress across the ice stream is given by a continuously varying function.

In both cases, the basal properties vary in the across-flow direction only and are symmetric about the ice stream centerline. As a result, the velocity solutions are also uniform along flow and symmetric about the centerline.

Provided files

- README.md
 - Information about the test case, including technical details about running it.
- runStream.py

The script to set up and run the test.

• stream.config

The default configuration settings for running CISM with the test case. Note that this file is parsed by the runStream.py script. Most of the relevant options that might be changed for this test case (e.g., grid spacing) can be done so from the command line, without having to edit the .config files (use ./runStream.py --help for a description of available options). The choice of yield stress distribution (Raymond or Schoof) can be toggled by editing line 122 of runStream.py.

Running the test

One script sets up the initial condition and runs the model:

./runStream.py

and another is used to analyze the results:

./plotStream.py

Results

The plotStream.py script will plot model output relative to the analytical solutions in Raymond (2000) and Schoof (2006). The choice of analytical solution for comparison is automatic, based on the yield stress chosen in the runStream.py script. Figure 8.4 shows example output for both test cases.

Note: The excellent agreement between the CISM results and the Raymond analytical solution requires setting which_ho_assemble_beta = 1 in the config file (see Section 7.3). With this setting, the step change in yield stress is well resolved. Otherwise there is some smoothing of the traction parameter β over neighboring nodes during finite-element assembly (Section 6.2.1), and the results are less accurate.

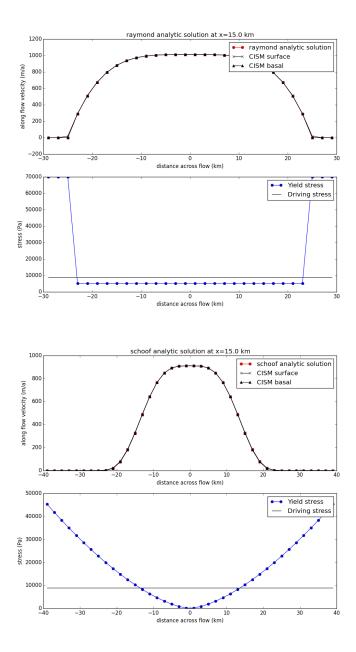


Figure 8.4: Comparison between CISM model output (black) and analytic solution (red) for the Raymond (top) and Schoof (bottom) stream test cases. This figure was generated with ./plotStream.py after running the test with ./runStream.py. Additional runtime options are described in the README.md and by invoking the '--help' option at the command line.

8.2.4 Confined shelf

The confined shelf test is based on tests 3 and 4 of the idealized (i.e., not Ross) EISMINT shelf test cases. It simulates the flow within an idealized, 500 m thick ice shelf in a confined, rectangular embayment. Grounded ice is not explicitly modeled but included in the model setup as Dirichlet boundary conditions for velocity along the ice shelf edges. More detailed

information on this test case can be found here ⁶ in the "shelf-descr.pdf" document.

Note that the confined shelf and circular shelf experiments are both in the tests/shelf directory and share some files.

Provided files

- README.md
 Information about the test case, including technical details about running it.
- runShelfConfined.py
 The script to set up and run the test.
- shelf-confined.config

 The default configuration settings for running CISM with the test case.

Running the test

One script sets up the initial condition and runs the model:

./runShelfConfined.py

Results

There is no script for analyzing the results. See the URL above for information about assessing the model output. You can manually inspect the results using a tool such as ncview. Figure 8.5 shows an example.

 $^{^6 \}rm http://homepages.vub.ac.be/~phuybrec/eismint/iceshelf.html$

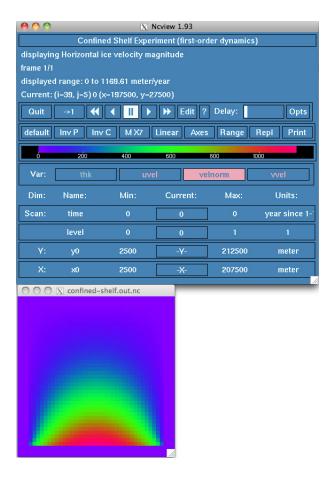


Figure 8.5: Confined shelf velnorm field using default shelf-confined.config settings. This figure is a screenshot of noview.

8.2.5 Circular shelf

The circular shelf test case is a variant on the confined shelf discussed above. It simulates the flow within a circular ice shelf with a uniform thickness of 1000 m, which is grounded at a single grid point at its center. This test case confirms a working "floating ice" boundary condition in 2D (i.e., in map plane view) and also confirms radial symmetry.

Note that the confined shelf and circular shelf experiments are both in the tests/shelf directory and share some files.

Provided files

- README.md
 Information about the test case, including technical details about running it.
- runShelfCircular.py
 The script to set up and run the test.
- shelf-circular.config

 The default configuration settings for running CISM with the test case.

Running the test

One script sets up the initial condition and runs the model:

./runShelfCircular.py

Results

There is no script for analyzing the results. You can manually inspect the results using a tool such as ncview. Figure 8.6 shows an example.

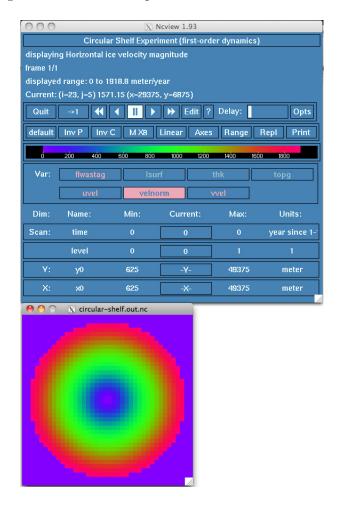


Figure 8.6: Circular shelf velnorm field using default shelf-circular.config settings. This figure is a screenshot of noview.

8.2.6 Ross Ice Shelf

The Ross experiment is designed to simulate the flow of the Ross Ice Shelf of Antarctica under idealized conditions (e.g., constant and uniform flow-law rate factor). For more information about the experiment and its results, see here ⁷. Also, see MacAyeal et al. (1996) for a discussion of the official model intercomparison results.

Using the Glissade solver with the Blatter-Pattyn approximation, this experiment typically takes about 10 minutes to run on a single processor. Using the shallow-shelf approximation instead, the results are very similar and are found more quickly (in less than a minute). For this reason, the SSA (which_ho_approx = 1) is the default.

 $^{^7 \}rm http://homepages.vub.ac.be/^{\sim}phuybrec/eismint/iceshelf.html$

Provided files

- README.md
 Information about the test case, including technical details about running it.
- runRoss.py
 The script to set up and run the test.
- ross.config
 The default configuration settings for running CISM with the test case.
- plotRoss.py
 The script to plot the test results.

Running the test

One script sets up the initial condition and runs the model:

./runRoss.py -r

(Without the "-r" flag, the script will set up the initial condition but not run CISM.) Another script can be used to visualize the results:

./plotRoss.py

Results

The plotRoss.py script will generate a figure of the velocity field calculated for the Ross Ice Shelf. The results should look very similar to Figures 8.7 and 8.8. You can compare these with similar figures in the paper by MacAyeal et al. (1996).

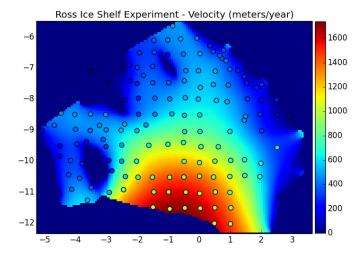


Figure 8.7: Ross Ice Shelf velocity field calculated by CISM. This figure is generated by plotRoss.py.

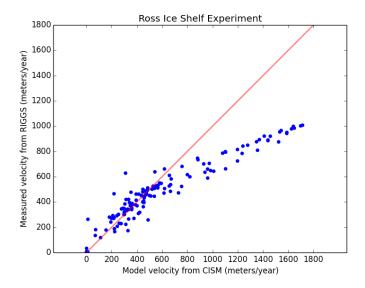


Figure 8.8: CISM-modeled Ross Ice Shelf speeds vs. those from observations. This figure is generated by plotRoss.py.

8.2.7 MISMIP

The MISMIP (Marine Ice Sheet Model Intercomparison Project) experiments (Pattyn et al., 2012) are designed to study the behavior of marine ice-sheet models under idealized conditions while modifying the flow-rate factor. The goal is to study the advance and retreat of the ice sheet and ideally, for a set flow-rate factor, the steady-state grounding line position should be the same in the advance state as in the retreat state. Pattyn et al. (2012) and Leguy et al. (2014) showed the conditions under which the hysteresis is minimal while using different resolution, Stokes approximations and basal water pressure at the grounding line. The experiments are proposed to be run over two different bed topographies: linear and polynomial. Using the Glissade solver with the SSA approximation, this experiment typically takes about a couple of hours to run on 36 processors and at a resolution of 8 km. While running with a linear bed, the initial time step should be chosen as 1 y for the first 3 flow rate values and can then be relaxed to 2 and 4 yr later on. Of course this will vary based on the chosen resolution.

Provided files

- README.mismip
 Information about the test case, including technical details on how to run it.
- mismipSetup.py
 The script to set up the test.
- mismipRun.py
 The script to run the test.
- mismipWriteGL.py

 The script to write the file containing the grounding line location.
- mismipPlotGL.py
 The script to plot the test results.

- mismip.config.template

 The default configuration template settings for running CISM with the test case. Inputs
 will vary based on the options used with the setup file.
- runCISM.cheyenne.template
 The template script to submit a job on Cheyenne HPC system. The wall time, number of nodes and processors will need to be adjusted depending on the need of the simulation.

Running the test

For detailed instruction, follow the instruction from the README.mismip file. In short the order of operations are:

Set up the experiment (add options if need be):

> python mismipSetup.py

Run the experiment (add options if need be):

> python mismipRun.py

Write the grounding line position to a file (add options if need be):

> python mismipWriteGL.py

Plot the results (add options if need be):

> python mismipPlotGL.py

Results

The mismipPlotGL.py script will generate a figure of the grounding line position as a function of the flow-rate factor inverse at the end of each experiment and the semi-analytic solution from Schoof (2007). The results should look very similar to Figures 8.9 (for the linear bed) and 8.10 (for the polynomial bed) when using the "powerlaw" (Schoof, 2007) basal friction law. You can compare these with similar figures in the papers by Pattyn et al. (2012) and Leguy et al. (2014).

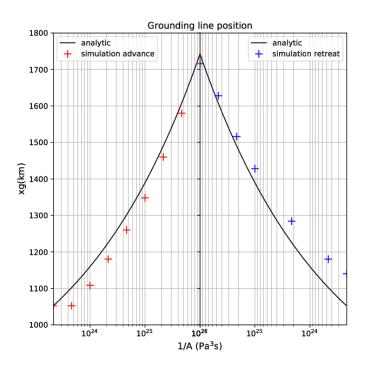


Figure 8.9: Steady state grounding line positions for the MISMIP experiments and a linear bed. This figure is generated using mismipPlotGL.py.

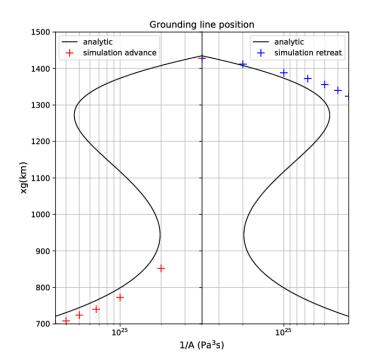


Figure 8.10: Steady state grounding line positions for the MISMIP experiments and a polynomial bed. This figure is generated using mismipPlotGL.py with the option --bed poly.

8.2.8 MISMIP3d

The MISMIP3d experiments (Pattyn et al., 2013) study the effect of a spatially variable perturbation in the basal shear stress factor of marine ice-sheet models and consists in 3 steps. First, during the spinup, the model is ran to steady state. The basal friction law is then modified by a Gaussian perturbation that introduces slipperiness at the grounding line at the center of the domain. This perturbation is maintained over 100 y. The third step consists in resetting the basal friction to its original value and ran again to steady state. At the end of the simulation the model should reverse back to the same marine ice sheet profile than the one at the end of the spinup. Using the Glissade solver with the SSA approximation, this experiment typically takes less than 10 minutes to run on 36 processors and at a resolution of 8 km. The initial time step can be chosen as 4 initially but the perturbation experiment requires a smaller time step initially of 1 y in order to run without CFL violation. Of course time steps will vary based on the chosen resolution.

Provided files

- README.mismip3d Information about the test case, including technical details on how to run it.
- mismip3dSetup.py
 The script to set up the test.
- mismip3dRun.py
 The script to run the test.

- mismip3dWriteGL.py
 The script to write the file containing the grounding line location.
- mismip3dPlotGL.py
 The script to plot the test results.
- mismip3d.config.template

 The default configuration template settings for running CISM with the test case. Inputs
 will vary based on the options used with the setup file.
- runCISM.cheyenne.template

 The template script to submit a job on Cheyenne HPC system. The wall time, number of nodes and processors will need to be adjusted depending on the need of the simulation.

Running the test

For detailed instruction, follow the instruction from the README.mismip3d file. In short the order of operations are:

Set up the experiment (add options if need be):

> python mismip3dSetup.py

Run the experiment (add options if need be):

> python mismip3dRun.py

Write the grounding line position to a file (add options if need be):

> python mismip3dWriteGL.py

Plot the results:

> python mismip3dPlotGL.py

Results

The mismip3dPlotGL.py script will generate a figure of the grounding line positions for each step of the experiments as a function of spatial location. Note that the grounding line position for the non-perturbed experiment can be compared to the semi-analytic solution from Schoof (2007). The results should look very similar to Figure 8.11 when using the "powerlaw" (Schoof, 2007) basal friction law. You can compare these with similar figures in Pattyn et al. (2013) (in the paper the results are plotted over half of the y-axis due to symmetry of the perturbation).

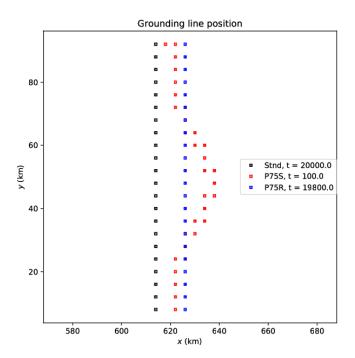


Figure 8.11: Grounding line positions calculated by CISM for the MISMIP3d experiments with a resolution of 4 km. This figure is generated using mismip3dPlotGL.py.

8.2.9 MISMIP+

The MISMIP+ (Marine Ice Sheet and Model Intercomparison Project third phase) experiments (Asay-Davis et al., 2016) are designed to study the transient behavior of marine ice sheet model while undergoing a perturbation in basal melting under ice shelves and ice calving. The bed topography is represented by a trough which bottom is represented by a polynomial function. This way the flowing ice is subject to buttressing. This experimental setup and perturbations are designed to more closely mimics Antarctic marine ice stream behavior (compared to MISMIP and MISMIP3d). Using the Glissade solver with the SSA approximation, this experiment typically takes about 25 minutes to run on 36 processors and at a resolution of 4 km. A time step of 2 y is adequate for the spinup but it then needs to be reduced to 1 y for some perturbation experiments. Of course this will vary based on the chosen resolution.

Provided files

- README.mismip+ Information about the test case, including technical details on how to run it.
- mismip+Setup.py
 The script to set up the test.
- mismip+Run.py
 The script to run the test.
- mismip+WriteGL.py
 The script to write the file containing the grounding line location.

- mismip+PlotGL.py
 The script to plot the test results.
- mismip+.config.template

 The default configuration template settings for running CISM with the test case. Inputs
 will vary based on the options used with the setup file.
- runCISM.cheyenne.template

 The template script to submit a job on Cheyenne HPC system. The wall time, number of nodes and processors will need to be adjusted depending on the need of the simulation.

Running the test

For detailed instruction, follow the instruction from the README.mismip+ file. In short the order of operations are:

Set up the experiment (add options if need be):

> python mismip+Setup.py

Run the experiment (add options if need be):

> python mismip+Run.py

Write the grounding line position to a file (add options if need be):

> python mismip+WriteGL.py

Plot the results:

> python mismip+PlotGL.py

Results

The mismip+PlotGL.py script will generate a figure of the grounding line positions and the grounded area for each step of the control and perturbed experiments as a function of spatial location and time respectively. The results should look very similar to Figure 8.12 when using the "Schoof2005" (Schoof, 2005; Leguy et al., 2014) basal friction law. You can compare it with similar figures in Asay-Davis et al. (2016).

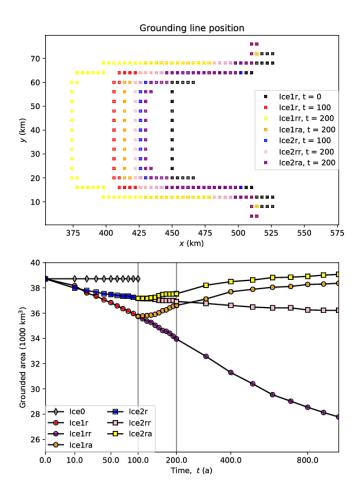


Figure 8.12: Grounding line positions and grounded area calculated by CISM for the MISMIP+ experiments with a resolution of 4 km. This figure is generated using mismip+PlotGL.py.

8.2.10 Other tests

Other higher-order tests that are still in development (e.g., "slab") are also included in the ./tests/ directory. Instructions for running these tests are included in each directory. Since these tests have not yet been validated, they are for use at your own risk.

8.3 The build and test structure (BATS)

A build and test structure (BATS) has been included in the ./tests/regression/ directory. BATS is capable of automatically building CISM and then running a set of regression tests on a number of platforms. BATS is primarily intended to allow users and developers of CISM to quickly generate a set of regression tests for use with the Land Ice Verification and validation toolkit(LIVVkit)⁸. This allows users to quickly and easily build confidence in their installation

 $^{^{8}}$ https://github.com/LIVVkit/LIVVkit

of CISM and, by extension, their scientific results.

Provided files

- README.md
 Information about the build and test structure, including technical details about using it.
- build_and_test.py
 The main script to build CISM, and run a set of regression tests.
- setup_hopper.bash and setup_titan.bash Scripts to load the needed modules on the higher-performance-computing architectures Hopper (NERSC) and Titan (OLCF).
- util/
 A directory containing a set of helper python modules and files for build_and_test.py.

Using BATS

BATS works very similar to how you would build CISM and run one or more CISM tests. BATS will build a version of CISM, and then either run a set of regression tests if you are using a personal computer (PC), or setup a series of regression tests and generate a job submission script if you are using a high-performance-computing architecture (HPC).

For example, on the HPC Titan, if you wanted to build CISM using the gnu compiler, and run a set of regression tests, you would run these commands:

```
cd tests/regression/
source setup_titan.bash
./build_and_test.py -p titan -c gnu -b ./build
```

which will result in BATS generating all the CMake build files into a new directory called build and building CISM into the directory build/cism_driver. BATS will then setup a set of CISM's higher-order tests:

- Dome (at a variety of resolutions and processor counts)
- Circular and confined shelf
- ISMIP-HOM a and c (at 20 and 80 km resolutions)
- ISMIP-HOM f
- Stream

All the files associated with each test will be output to a new directory called reg_test/titan-gnu, which has a directory structure that mirrors CISM's test directory structure:

```
reg_test/
|-- PLATFORM-COMPILER/
|-- CMakeCache.txt
|-- higher-order/
| |--dome/
| |-- dome.RESO.pPRC.*
| |-- timing/dome-t?.RESO.pPRC.*
| |-- ismip-hom
| |-- ismip-hom-a.RESO.pPRC.*
| |-- ismip-hom-c.RESO.pPRC.*
```

where * and ? are POSIX regular expressions metacharacters, RESO is a number indicating the resolution the test was run at (the meaning of the number is test dependent), and pPRC is a number, prefixed by a p, indicating the number of processors used when running the model. This reg_test directory will be formatted to used with LIVVkit directly. Note: everything below the dashed line will only appear on HPC systems (tests are immediately run on PCs). submit_all_jobs.bash will submit all the jobs in the jobs/ directory and clean_timing.bash will clean out any */timing/ directory such that only the timing files remain (to be used after all jobs finish).

Advanced usage

BATS is designed to be flexible and work with any LIVVkit usage scenario. In order to do that, BATS provides a number of options to configure which system you are using, when/where/how CISM is built, the destination of the output directory, and which tests are run. For more information, see the LIVVkit wiki ⁹.

 $^{^9 \}rm https://github.com/LIVVkit/LIVVkit/wiki$

$\begin{array}{c} {\rm Part~II} \\ {\bf Appendix} \end{array}$

Appendix A

NetCDF Variables

The following list shows all the netCDF variable names used by CISM. Only variables marked with * are loaded (if present) by the input routines.

A.1 Glide/Glissade Variables

Name	Description	Units
level	sigma layers	1
	CF name: land_ice_sigma_coordinate	
lithoz	vertical coordinate of lithosphere layer	meter
staglevel	stag sigma layers	1
	CF name: land_ice_stag_sigma_coordinate	
stagwbndlevel	stag sigma layers with boundaries	1
	${ m CF\ name:\ land_ice_stag_sigma_coordinate_with_branch }$	nd
х0	Cartesian x-coordinate, velocity grid	meter
x1*	Cartesian x-coordinate	meter
у0	Cartesian y-coordinate, velocity grid	meter
y1*	Cartesian y-coordinate	meter
C_space_factor*	spatial factor for basal shear stress	1
D_{plume}^*	plume thickness	meter
_	CF name: plume_thickness	
S_ambient	ambient ocean salinity	psu
	CF name: ambient ocean salinity	
S_basal	basal ice salinity	psu
	CF name: basal_ice_salinity	
S_{plume}^*	plume salinity	psu
	CF name: plume_salinity	
T_{-} ambient	ambient ocean temperature	degree_Celsius
	CF name: ambient_ocean_temperature	
T_basal	basal ice temperature	degree_Celsius
	CF name: basal_ice_temperature	
T_plume*	plume temperature	degree_Celsius
	CF name: plume_temperature	
acab*	accumulation, ablation rate	meter/year ice
	${ m CF\ name:\ land_ice_surface_specific_mass_balanc}$	е
$acab_anomaly^*$	accumulation, ablation rate anomaly	meter/year
	contin	ued on next page

continued from previous page Name	Description	Units
Name		
	CF name: land_ice_surface_specific_mass_balanc	
acab_applied	applied accumulation, ablation rate	meter/year
	CF name: land_ice_surface_specific_mass_balanc	
$acab_applied_tavg$	applied accumulation, ablation rate (time average)	meter/year
	CF name: land_ice_surface_specific_mass_balanc	
$\mathtt{acab}_{ extsf{-}}corrected$	corrected accumulation, ablation rate	meter/year
	CF name: land_ice_surface_specific_mass_balanc	$e_corrected$
adv_cfl_dt	advective CFL maximum time step	years
area_factor	area scale factor for stereographic map projection	unitless
artm*	annual mean air temperature	degree_Celsius
	CF name: surface_air_temperature	
basal_mbal_flux	basal mass balance flux	m kg/m2/s
	CF name: land_ice_basal_specific_mass_balance_:	flux
basal_mbal_flux_tavg	basal mass balance flux (time average)	kg/m2/s
	CF name: land_ice_basal_specific_mass_balance_:	flux
beta*	higher-order bed stress coefficient	Pa yr/m
beta_internal	weighted higher-order bed stress coefficient	Pa yr/m
bfricflx*	basal friction heat flux	watt/meter2
bheatflx*	downward basal heat flux	watt/meter2
bmlt	basal melt rate	meter/year
	CF name: land_ice_basal_melt_rate	
bmlt_applied	applied basal melt rate	meter/year
bmr o_apprioa	CF name: land_ice_basal_melt_rate_applied	lifeter/year
bmlt_applied_tavg	applied basal melt rate (time average)	meter/year
bmrt_appried_tavg	CF name: land_ice_basal_melt_rate_applied	meter/year
bmlt_float	basal melt rate for floating ice	meter/year
biii10_110at	CF name: floating_ice_basal_melt_rate	meter/year
bmlt_float_anomaly*	basal melt rate anomaly for floating ice	meter/year
bmit_iioat_anomaiy	CF name: floating_ice_basal_melt_rate_anomaly	meter/year
bmlt_float_external*		
bmIt_IIoat_external	external basal melt rate for floating ice	meter/year
11	CF name: floating_ice_basal_melt_rate_external	
bmlt_float_inversion*	basal melt rate for floating ice from inversion	meter/year
bmlt_float_prescribed*	prescribed basal melt rate for floating ice	meter/year
bmlt_ground	basal melt rate for grounded ice	meter/year
	CF name: grounded_ice_basal_melt_rate	1 01:
bpmp	basal pressure melting point temperature	degree_Celsius
	CF name: pressure_melting_point_temperature	
btemp	basal ice temperature	degree_Celsius
	CF name: land_ice_temperature	
btract*	basal traction (magnitude)	Pa
btractx*	basal traction (x-direction comp)	Pa
btractx_extend*	basal traction (x-direction comp)	Pa
btracty*	basal traction (y-direction comp)	Pa
$\mathtt{btracty_extend}^*$	basal traction (y-direction comp)	Pa
btrc	basal slip coefficient	meter/pascal/yea
bwat*	basal water depth	meter
bwatflx	basal water flux	meter3/year
calving_flux	calving flux	kg/m2/s
5	CF name: land_ice_specific_mass_flux_due_to_cal	
		ontinued on next pag

continued from previous Name	Description	Units
calving_flux_tavg	calving flux (time average)	kg/m2/s
carving_rrux_cavg	CF name: land_ice_specific_mass_flux_due_to_cal	
calving_lateral	lateral calving rate	meter/year
calving_mask*	calving mask	1
calving_rate	rate of mass loss by calving	meter/year
calving_rate_tavg	rate of mass loss by calving (time average)	meter/year
calving_thck	thickness of calving ice	meter
cell_area*	cell area of cism grid	meter2
damage*	ice damage	unitless [0,1]
detrainment	detrainment rate	meter/year
decraimment	CF name: detrainment_rate	meter/year
diff_cfl_dt	diffusive CFL maximum time step	vonre
diffu	apparent diffusivity	years meter2/year
dissip*	dissipation rate (W m-3) divided by rhoi Ci	deg C/yr
dissipstag*	dissipation rate (W m-3) divided by rhoi Ci	deg C/yr deg C/yr
divDu_plume	divergence of D_plume*u_plume	meter/year
aivbu_prume	CF name: plume divergence rate	meter/year
divu*	horizontal divergence	1 /****
		1/year
drho_plume	plume density diffrence with ambient ocean	kg/m3
J+1-1- J+	CF name: plume_density_difference	mant on /rra an
dthck_dt	tendency of ice thickness (NOTE: Glissade only) dH/dt used for inversion	meter/year
dthck_dt_inversion*	tendency of ice thickness (NOTE: Glissade only)	meter/year
dthck_dt_tavg	(time average)	meter/year
dthckdtm	tendency of ice thickness (NOTE: Glide only)	meter/year
dusrfdtm	rate of upper ice surface elevation change (NOTE:	meter/year
dustidum	Glide only)	meter/year
effecpress*	effective pressure	Pa
efvs*	effective pressure	Pascal * years
enthalpy	specific enthalpy	J/m3
entrainment	entrainment rate	meter/year
encrainmenc	CF name: entrainment_rate	meter/year
eps_eff	effective strain rate	1/year
	xx component of strain rate tensor	1/year
eps_xx	xy component of strain rate tensor	1/year
eps_xy eps_xz	xz component of strain rate tensor	1/year
	yy component of strain rate tensor	1/year
eps_yy	yz component of strain rate tensor yz component of strain rate tensor	1/year
eps_yz	global average sea level	meter
eus	CF name: global_average_sea_level_change	meter
f_flotation	flotation function	unitless
f_ground	grounded ice fraction	unitless [0,1]
1_ground	CF name: grounded_fraction	unitiess [0,1]
f_ground_cell	grounded ice fraction in cells	unitless [0,1]
T-RIOMMG-cell	CF name: grounded_fraction_cell	umness [0,1]
floating magic	mask for floating ice	1
floating_mask	Pre-exponential flow law parameter	pascal**(-n) year**(-1
fl:73*	TITE-EXPONERMAN HOW DAY ATTRIBUTED	pascai (-11) veai (-1
flwa*		
flwa* flwastag* gl_flux	Pre-exponential flow law parameter grounding line flux	pascal**(-n) year**(-1 kg/m/s

continued from previous page Name	Description	Units
Name		Units
	CF name: land_ice_mass_flux_at_grounding_line	
gl_flux_east	grounding line flux eastward	kg/m/s
	CF name: land_ice_mass_flux_at_grounding_line_e	
gl_flux_north	grounding line flux northward	kg/m/s
	CF name: land_ice_mass_flux_at_grounding_line_r	
gl_flux_tavg	grounding line flux (time average)	kg/m/s
	CF name: land_ice_mass_flux_at_grounding_line	
gravity	gravitational acceleration	meter/s/s
	CF name: gravity	
grounded_mask	mask for grounded ice	1
iarea	area covered by ice	m2
iareaf	area covered by floating ice	m2
iareag	area covered by grounded ice	m2
ice_age*	ice age	year
3	CF name: land_ice_age	
ice_cap_mask	mask for ice caps	1
ice_mask	mask for ice (1) or no ice (0)	1
ice_mask_stag	mask on staggered grid for ice (1) or no ice (0)	1
ice_sheet_mask	mask for ice sheet	1
ice_specific_heat	ice specific heat	J/kg/K
ice_specific_neat	CF name: ice_specific_heat	3/ Kg/ K
ice_thermal_conductivity	ice thermal conductivity	J/(K kg)
ice_thermai_conductivity	CF name: ice_thermal_conductivity	3/(IX Kg)
imass	ice mass	lea
	ice mass above flotation	kg
imass_above_flotation		kg
ivol	ice volume	m3
kinbcmask*	Mask of locations where uvel, vvel value should be	1
	held constant	1 011
litho_temp*	lithosphere temperature	degree_Celsius
load*	bedrock deflection from applied load	meter
lsurf	ice lower surface elevation	meter
$no_advance_mask^*$	cells where advance is not allowed	1
	CF name: land_ice_no_advance_mask	
overwrite_acab_mask*	cells where acab is overwritten	1
	CF name: land_ice_overwrite_acab_mask	
powerlaw_c_inversion*	spatially varying C for powerlaw sliding	Pa $(m/yr)**(-1/3)$
powerlaw_c_prescribed*	prescribed spatially varying C for powerlaw sliding	Pa $(m/yr)**(-1/3)$
relx*	relaxed bedrock topography	meter
resid_u	u component of residual Ax - b (NOTE: Glissade only)	Pa/m
resid_v	v component of residual Ax - b (NOTE: Glissade	Pa/m
, .	only)	1 /
rho_ice	ice density	kg/meter3
	CF name: rho_ice	1 /
rho_seawater	seawater density	kg/meter3
	CF name: rho_seawater	
rhs_u	u component of b in $Ax = b$	Pa/m
rhs_v	v component of b in $Ax = b$	Pa/m
seconds_per_year	seconds per year	s/yr

continued from previous page		
Name	Description	Units
	CF name: seconds_per_year	
sfc_mbal_flux	surface mass balance flux	kg/m2/s
	CF name: land_ice_surface_specific_mass_balanc	
sfc_mbal_flux_tavg	surface mass balance flux (time average)	kg/m2/s
	CF name: land_ice_surface_specific_mass_balanc	
smb*	surface mass balance	mm/year water equivaler
	CF name: land_ice_surface_specific_mass_balanc	
soft*	bed softness parameter	meter/pascal/year
stagthk	staggered ice thickness	meter
20080	CF name: stag_land_ice_thickness	
tau_c	yield stress for plastic sliding	kilopascal
tau_eff	effective stress	Pa
tau_eff_calving	effective stress for calving	Pa
		Pa
tau_eigen1*	first eigenvalue of horizontal stress tensor	Pa Pa
tau_eigen2*	second eigenvalue of horizontal stress tensor	
tau_xx	xx component of deviatoric stress tensor	Pa
tau_xy	xy component of deviatoric stress tensor	Pa
tau_xz	xz component of deviatoric stress tensor	Pa
tau_yy	yy component of deviatoric stress tensor	Pa
tau_yz	yz component of deviatoric stress tensor	Pa
taudx*	driving stress (x-direction comp)	Pa
taudy*	driving stress (y-direction comp)	Pa
tauf*	higher-order basal yield stress	Pa
taux	basal shear stress in x direction (NOTE: Glide only)	kilopascal
tauy	basal shear stress in y direction	kilopascal
temp*	ice temperature	degree_Celsius
	CF name: land_ice_temperature	
tempstag*	ice temperature on staggered vertical levels with	degree_Celsius
	boundaries	
	CF name: land_ice_temperature_stag	
tempunstag*	ice temperature on unstaggered vertical levels	degree_Celsius
	CF name: land_ice_temperature_unstag	
thck_obs*	observed ice thickness	meter
thk*	ice thickness	meter
	CF name: land_ice_thickness	
thkmask*	mask	1
topg*	bedrock topography	meter
10	CF name: bedrock_altitude	
topg_obs*	observed bed topography	meter
total_bmb_flux	total basal mass balance flux	kg/s
total_bmb_flux_tavg	total basal mass balance flux (time average)	kg/s
total_calving_flux	total calving mass balance flux	kg/s
total_calving_flux_tavg	total calving mass balance flux (time average)	kg/s
total_gl_flux	total grounding line flux	kg/s
	total grounding line flux (time average)	
total_gl_flux_tavg	total grounding line nux (time average) total surface mass balance flux	kg/s
total_smb_flux		kg/s
total_smb_flux_tavg	total surface mass balance flux (time average)	kg/s
u_plume	plume velocity in x direction on B grid	meter/second
	CF name: plume_x_velocity	
		continued on next pag

$continued\ from\ previous$		
Name	Description	Units
u_plume_Cgrid	plume velocity in x direction on C grid	meter/second
1 0	CF name: plume_x_velocity_Cgrid	,
ubas	basal slip velocity in x direction	meter/year
	CF name: land_ice_basal_x_velocity	7.0
uflx	flux in x direction (NOTE: Glide only)	meter2/year
unstagbeta*	higher-order bed stress coefficient on the unstaggered	Pa yr/m
Ü	grid (NOTE: this will overwrite beta if both are in-	0 /
	put)	
usfc	surface velocity in x direction	meter/year
	CF name: land_ice_surface_x_velocity	, ,
usrf_inversion*	surface elevation used for inversion	meter
usrf_obs*	observed surface elevation	meter
ustar_plume	plume friction velocity	meter/second
1	CF name: plume_friction_velocity	, , , , , , , , ,
usurf*	ice upper surface elevation	meter
	CF name: surface_altitude	
uvel*	ice velocity in x direction	meter/year
	CF name: land_ice_x_velocity	
uvel_2d*	vertically averaged ice velocity in x direction	meter/year
	CF name: land_ice_x_velocity_2d	7,5
uvel_2d_extend*	vertically averaged ice velocity in x direction (ex-	meter/year
	tended grid)	7,5
	CF name: land_ice_x_velocity_2d	
uvel_extend*	ice velocity in x direction (extended grid)	meter/year
	CF name: land_ice_x_velocity	7.0
uvel_mean	vertical mean ice velocity in x direction	meter/year
	CF name: land_ice_vertical_mean_x_velocity	, ,
v_plume	plume velocity in y direction on B grid	meter/second
•	CF name: plume_y_velocity	,
v_plume_Cgrid	plume velocity in y direction on C grid	meter/second
1 0	CF name: plume_y_velocity_Cgrid	/
vbas	basal slip velocity in y direction	meter/year
	CF name: land_ice_basal_y_velocity	7.0
velnorm	Horizontal ice velocity magnitude	meter/year
vflx	flux in x direction (NOTE: Glide only)	meter2/year
vsfc	surface velocity in y direction	meter/year
	CF name: land_ice_surface_y_velocity	7.5
vvel*	ice velocity in y direction	meter/year
	CF name: land_ice_y_velocity	7.5
vvel_2d*	vertically averaged ice velocity in y direction	meter/year
	CF name: land_ice_y_velocity_2d	7,5
vvel_2d_extend*	vertically averaged ice velocity in y direction (ex-	meter/year
	tended grid)	7,5
	CF name: land_ice_y_velocity_2d	
vvel_extend*	ice velocity in y direction (extended grid)	meter/year
	CF name: land_ice_y_velocity	
vvel_mean	vertical mean ice velocity in y direction	meter/year
51_110411	CF name: land_ice_vertical_mean_y_velocity	1110001/3001
warm_ocean_mask*	warm ocean mask	1
CIM_COCUII_MUDIX		$\frac{1}{tinued \ on \ next \ pag}$

continued from previous page			
Name	Description	Units	
waterfrac*	internal water fraction	unitless [0,1]	
wbas	basal velocity in z direction	meter/year	
	CF name: land_ice_basal_z_velocity		
wgrd*	Vertical grid velocity	meter/year	
wsfc	surface velocity in z direction	meter/year	
	CF name: land_ice_surface_z_velocity		
wvel*	vertical ice velocity	meter/year	

A.2 Glint Variables

Name	Description	Units
ablt	ablation	meter (water)/year
arng	air temperature half-range	degreeC
global_orog	orographic elevation provided by global model	meter
	CF name: surface_altitude	
hflx_tavg*	heat flux to ice surface	W m-2
inmask	downscaling mask	1
local_orog	orographic elevation provided by local model	meter
	CF name: surface_altitude	
outmask*	upscaling mask	1
prcp	precipitation	meter (water)/year
	CF name: lwe_precipitation_rate	
rofi_tavg*	solid calving flux	kg m-2 s-1
rofl_tavg*	liquid runoff flux	kg m-2 s-1
siced*	superimposed ice depth	meter
snowd*	snow depth	meter
	CF name: surface_snow_thickness	

Appendix B

Input and Output (I/O)

B.1 NetCDF I/O

The netCDF¹ library is used for platform-independent, binary file I/O. CISM uses the f90 netCDF interface. Several source files are automatically generated from template files and a variable definition file using a python script. The netCDF files adhere to the CF convention² for naming climate variables. The netCDF files also store parameters used to define the geographic projection.

B.1.1 Data structures

Information associated with each dataset is stored in the glimmer_nc_stat type. Variable and dimension IDs are retrived from the data set by using the relevant netCDF library calls. Metadata (such as title, institution and comments) are stored in the derived type glimmer_nc_meta.

Input and output files are managed by two separate linked lists. Elements of the input file list contain the number of available time slices and information describing which time slice(s) should be read. Output file elements describe how often data should be written and the current time.

B.1.2 Code generator

Much of the code needed to do netCDF I/O is repetitive and can therefore be automatically generated. The code generator, <code>generate_ncvars.py</code>, is written in python and produces source files from a template <code>ncdf_template.in</code> and the variable definition file (Section B.1.3). The templates are valid source files; the generator simply replaces special comments with the code generated from the variable file.

B.1.3 Variable definition file

All netCDF variables are defined in control files, MOD_vars.def, where MOD is the name of the model subsystem. Variables can be modified/added by editing these files to provide I/O functionality for additional variables beyond those listed in Appendix A. A common example would be to modify the file glide_vars.def in ./libglide so that additional standard variables can be treated as "loadable" on input, or so that other standard variables internal to the code can be written to NetCDF output. Note that if MOD_vars.def is modified, the CMake configure script must be re-sourced before running "make" in order for those changes to be applied.

¹ http://www.unidata.ucar.edu/packages/netcdf/

²http://cfconventions.org/

The file is read using the python <code>ConfigParser</code> module. The format of the file is similar to Windows .ini files. Lines beginning with <code>#</code> or ; or empty lines are ignored. These files must have a definition section <code>[VARSET]</code> (see Table <code>B.1</code>). A new variable definition block starts with the variable name in square brackets <code>[]</code>. Variables are further specified by parameter name/value pairs which are separated by : or <code>=</code>. Parameter names and their meanings are summarized in Table <code>B.2</code>. All parameter names not recognised by the code generator (i.e., not in Table <code>B.2</code>) are added as variable attributes.

name	description	
name	Name of the model subsystem, e.g. glide. The f90 file is renamed	
	based on this name. The f90 module and module procedures are	
	prefixed with this name.	
datatype	The name of the f90 type on which the netCDF variables depend.	
datamod	The name of the f90 module in which the f90 type, datatype,	
	defined.	

Table B.1: Each variable definition file must have a section, called [VARSET], containing the parameters described above.

name	description		
dimensions	List of comma-separated dimension names of the variable. C no-		
	tation is used here, i.e. the slowest-varying dimension is listed		
	first.		
data	The variable to be stored/loaded. The f90 variable is assumed		
	to be one dimension smaller than the netCDF variable, i.e. f90		
	variables are always snapshots of the present state of the model.		
	Variables which do not depend on time are not handled automat-		
	ically. Typically, these variables are filled when the netCDF file		
	is created.		
factor	Variables are multiplied with this factor on output and divided by		
	this factor on input. Default: 1.		
load	Set to 1 if the variable can be loaded from file. Default: 0.		
average	Set to 1 if the variable should also be available as a mean over the		
	write-out interval. Averages are calculated only if required. To		
	store the average in a netCDF output file, append _tavg to the		
	variable name.		
units	UDUNITS compatible unit string describing the variable units.		
long_name	A more descriptive name of the variable.		
standard_name The corresponding standard name defined by the CF st			

Table B.2: List of accepted variable definition parameters.

Appendix C

Commonly Used Notation

The following table is a partial list of the notation for commonly used variables in this document.

Symbol	Variable
H	ice thickness
s	ice surface elevation
b	ice bottom elevation
b_r	bedrock elevation
B_s	surface mass balance (positive for accumulation)
M_b	basal melt rate (positive for melt)
B	combined surface and basal mass balance
u	x-component of ice velocity
v	y-component of ice velocity
w	z-component of ice velocity
η	effective viscosity

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