

# Hunting dark matter signals with deep learning at the LHC

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@ Feno-Exp

# Plan

## ◆ Models and sample generation

- Simplified models
- Kinematic features
- Benchmark models

## ◆ Neural Network algorithms

- Event-by-event data
- Data as 2D histograms → DNNs and CNNs

## ◆ Flexibility of the method

- Performance invariance with the number of background events
- Testing under incorrect training hypothesis:
  - coupling values
  - benchmark model

## ◆ Multimodel Classifiers

- Binary classifier
- Multiclass classifier

## ◆ Conclusions

# What we want

- Study several simplified **dark matter** models and their signatures at the **LHC** using **Neural Networks**
  - monojet plus missing transverse energy channel.
- Determine the **viability** of deep learning methods and analyze **flexible** they are
  - different neural networks and different data representations

# Models and sample generation

- Simplified models
- Kinematic features
- Benchmark models

# Simplified models

- **DM with a spin-0 mediator**
- **DM with a spin-1 mediator**
- **DM with a spin-2 mediator**
  
- **Axion-Like Particle (ALP) as DM**

$m_{\text{DM}}$ : DM mass

$m_\gamma$ : mediator mass

DM-mediator couplings

SM-mediator coupling

$m_a$ : axion mass (DM)

$f_a$ : axion scale

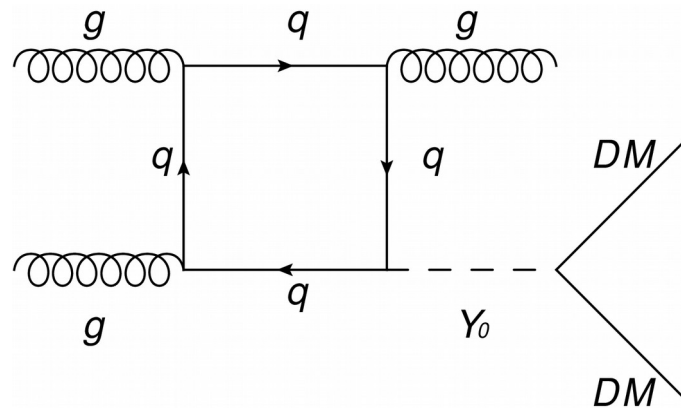
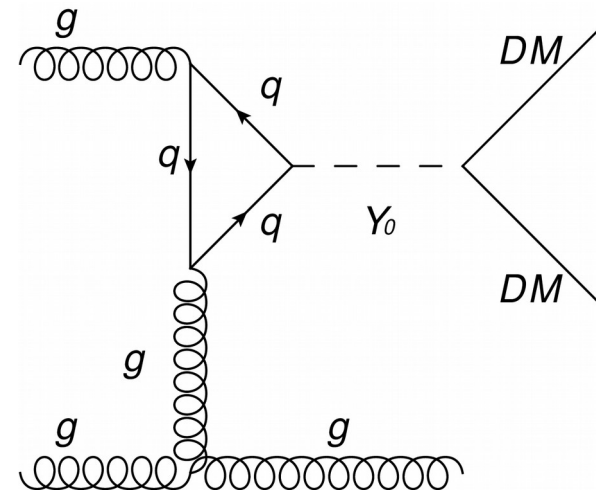
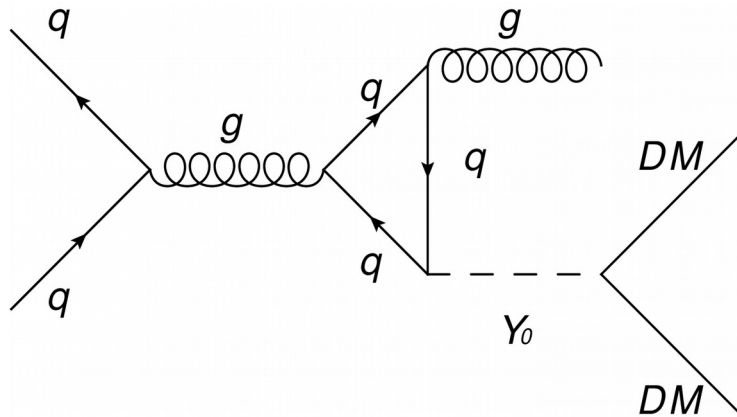
SM-axion couplings

# Simplified models

monojet plus  
missing transverse  
energy channel

- DM with a spin-0 mediator**

$$pp \rightarrow DM DM j$$

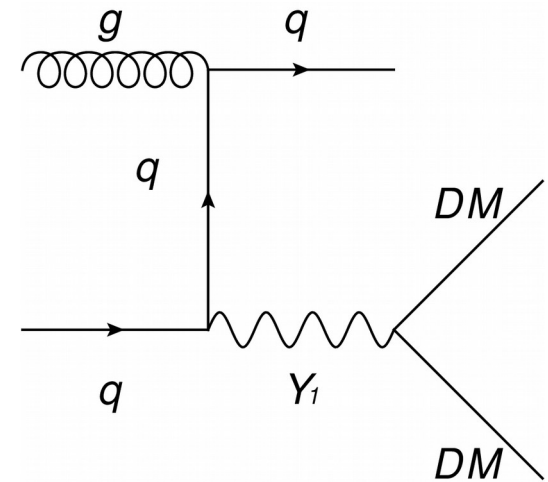
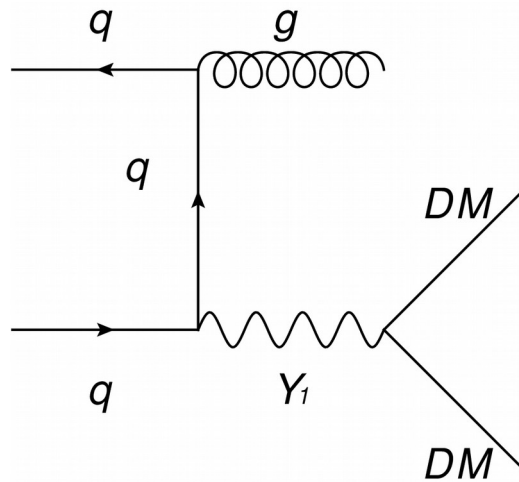
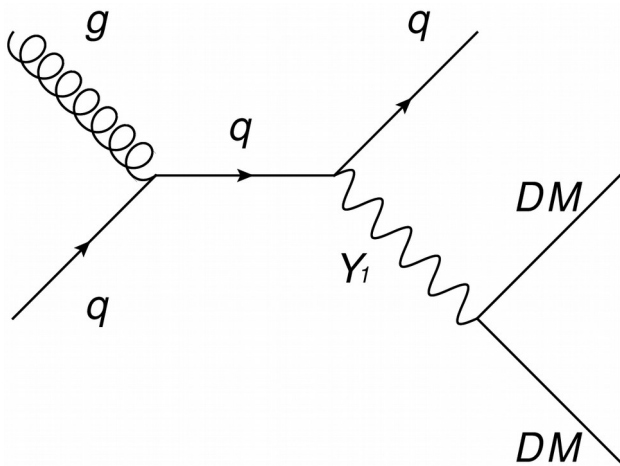


# Simplified models

monojet plus  
missing transverse  
energy channel

- DM with a spin-1 mediator**

$$pp \rightarrow DM DM j$$

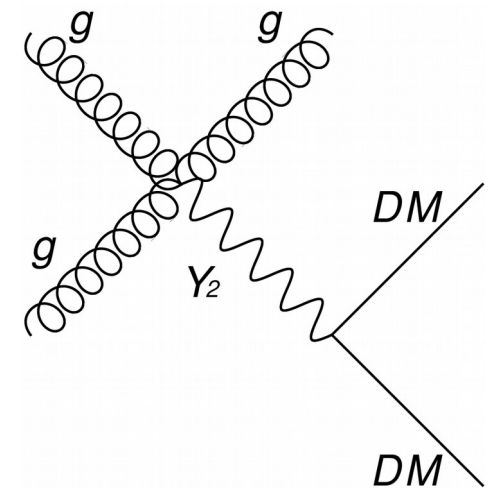
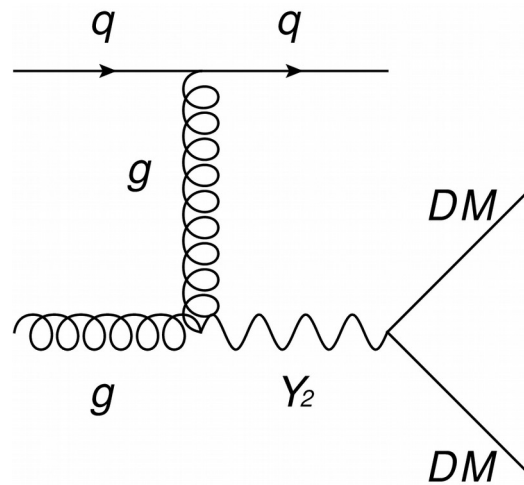
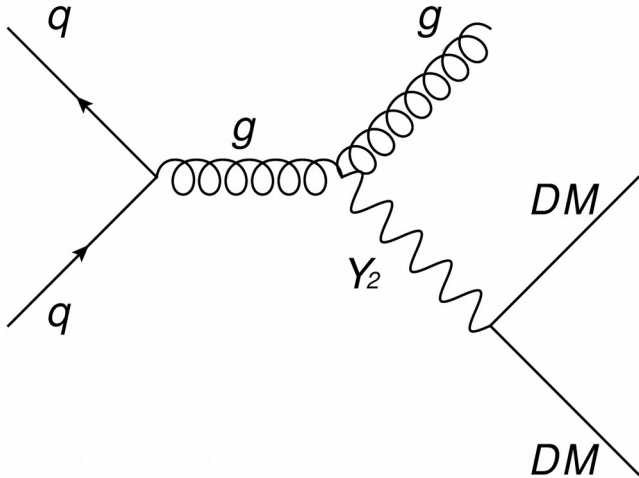


# Simplified models

monojet plus  
missing transverse  
energy channel

- **DM with a spin-2 mediator**

$$pp \rightarrow DM DM j$$



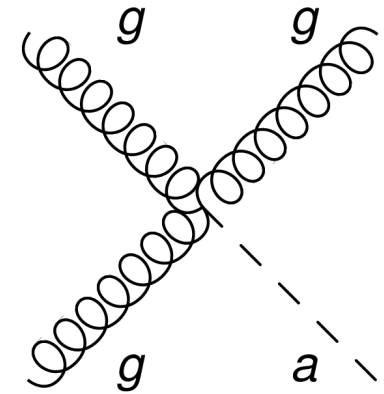
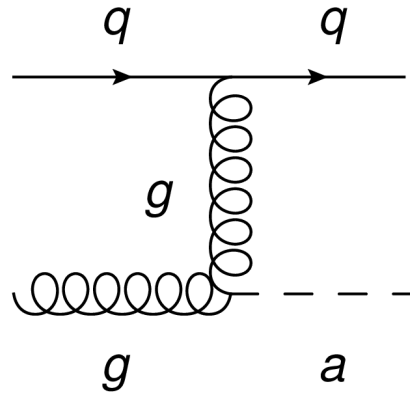
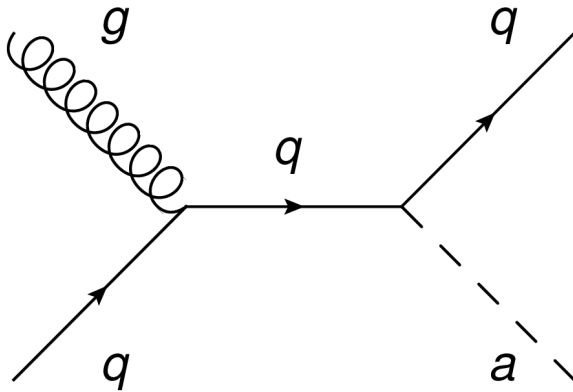


# Simplified models

monojet plus  
missing transverse  
energy channel

- **Axion-Like Particle (ALP) as DM**

$$pp \rightarrow a j$$

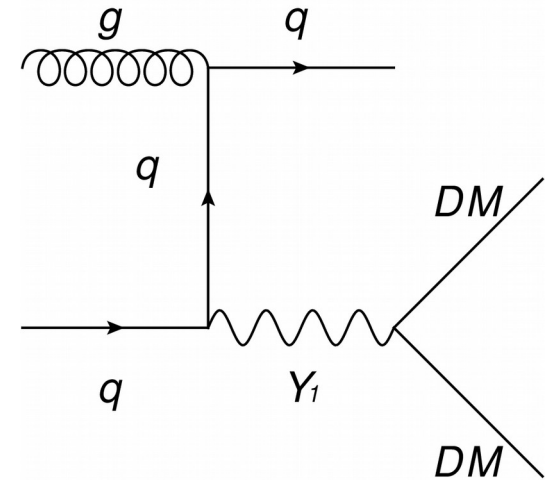
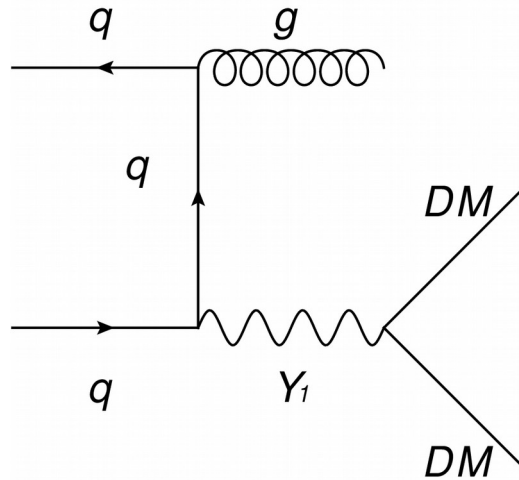
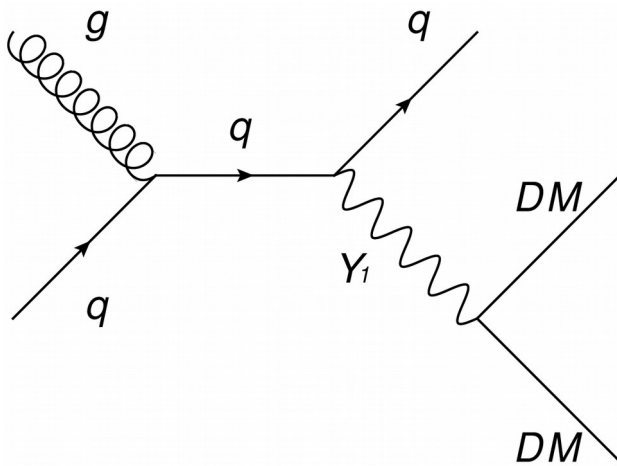


# Simplified models

monojet plus  
missing transverse  
energy channel

- **SM background**

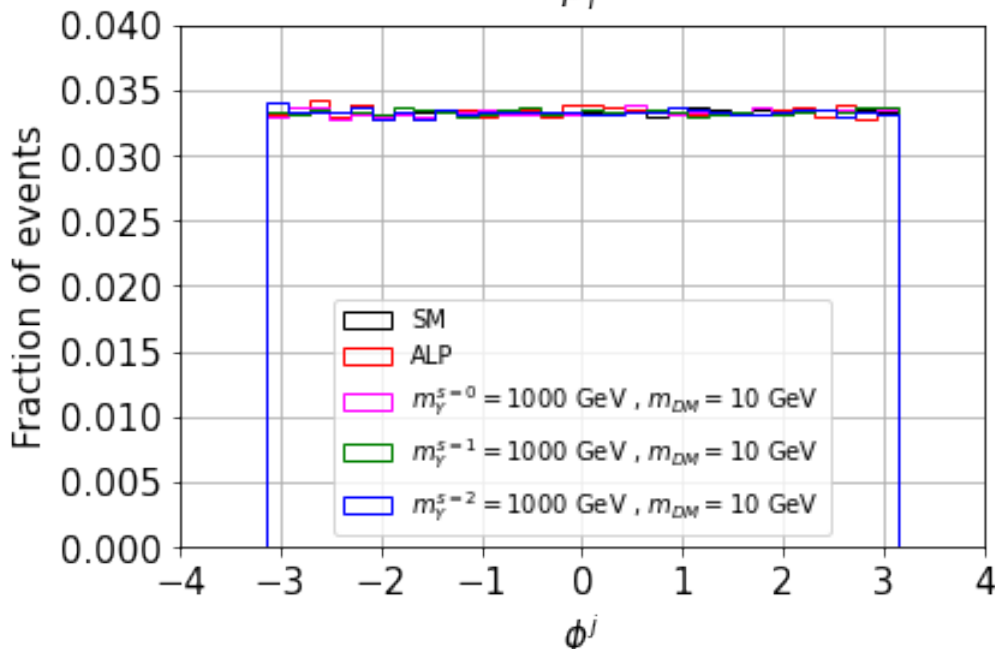
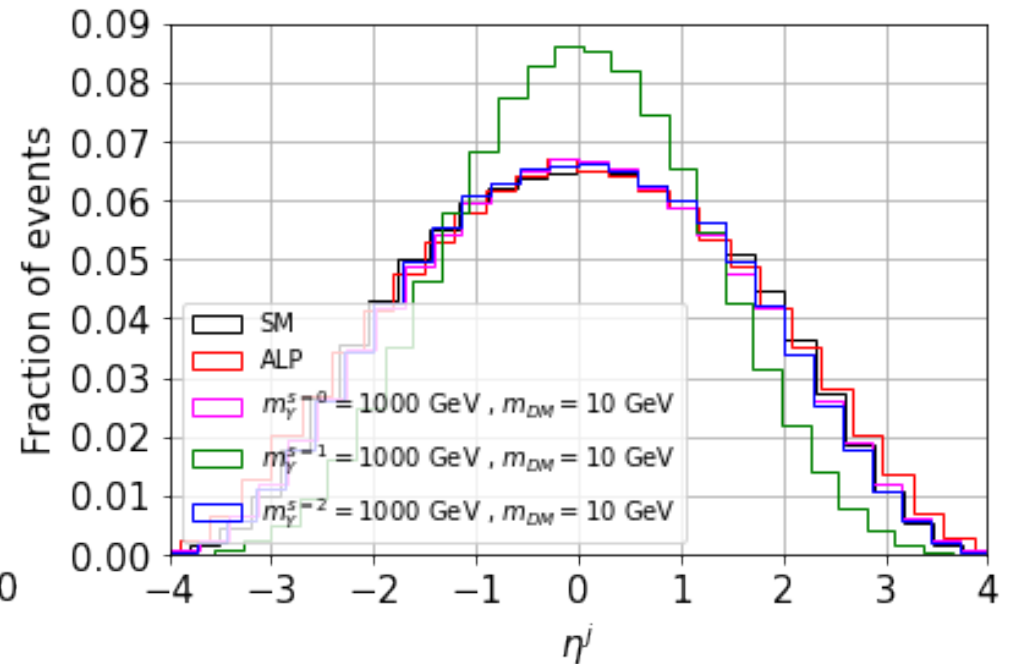
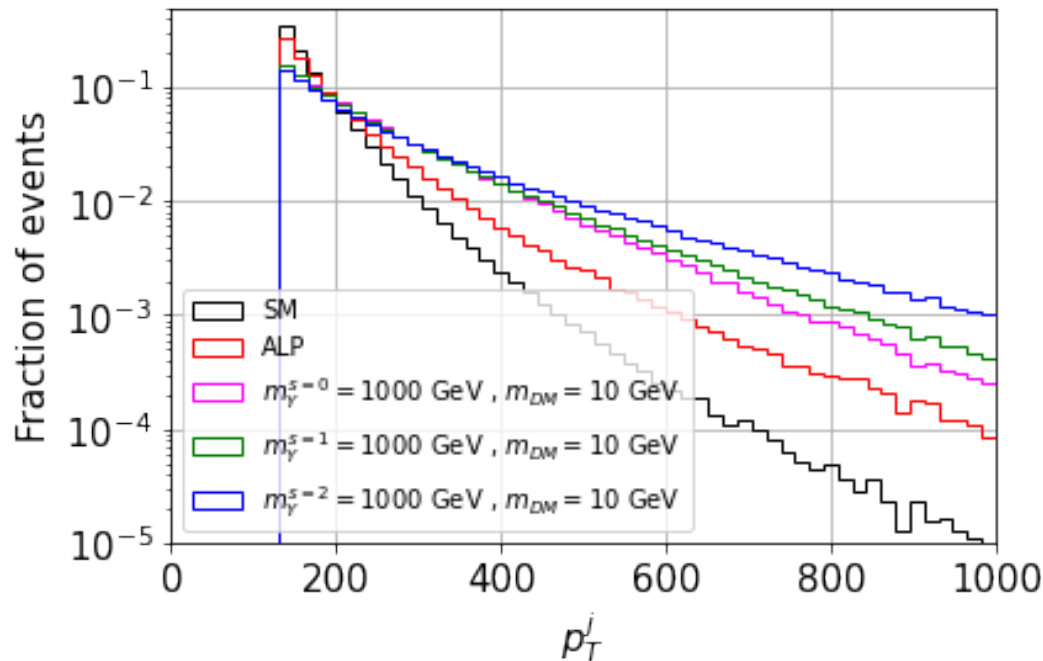
$$pp \rightarrow Z j \ (Z \rightarrow \nu \nu)$$



$$\begin{array}{l} Y \rightarrow Z \\ DM \rightarrow \nu \end{array}$$

# Kinematic distributions

Input data for our deep learning algorithms



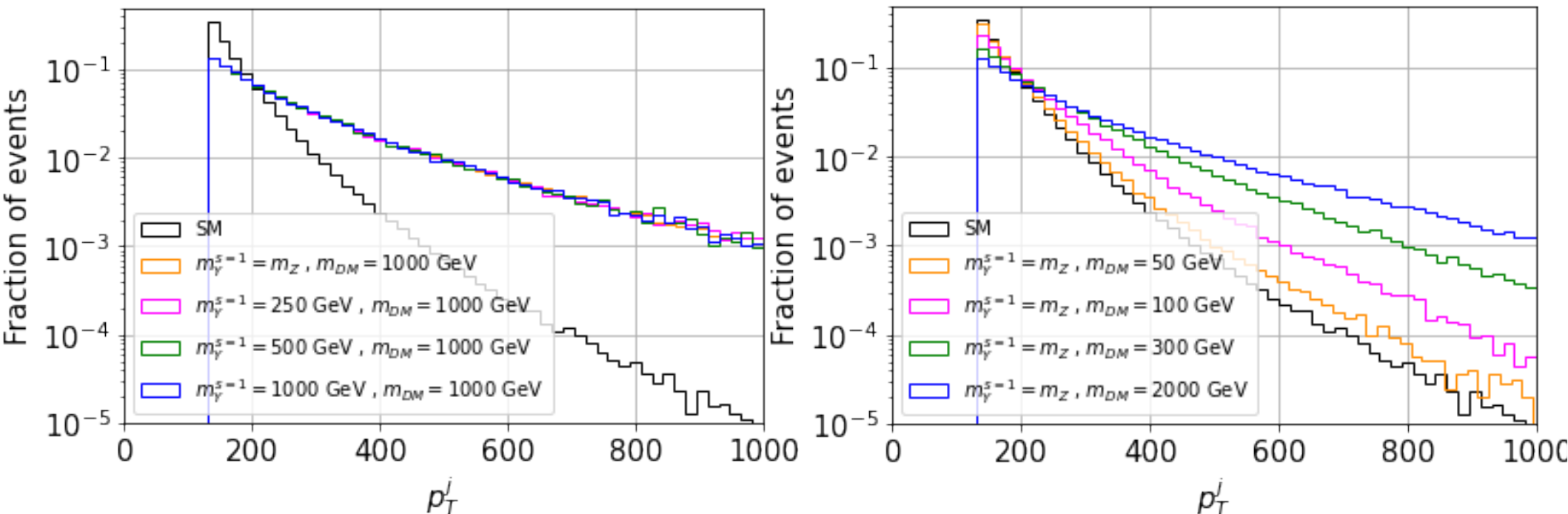
- 1.** The coupling values do not modify the kinematic distributions  
 → *models with a mediator defined by  $(m_{DM}, m_Y)$*
- 2.** ALPs distributions are independent of  $m_a$  and  $f_a$ .
- 3.** The azimuthal angle distribution does not show any useful structure.

# Kinematic distributions

## Models with a mediator

1. The coupling values do not modify the kinematic distributions  
 → *models with a mediator defined by  $(m_{DM}, m_Y)$*

But if the parameter space is divided according to the mediator status:



**Mediator status**

Off-shell

On-shell

Off-shell PS (by phase space)

**Relevant parameter**

$m_{DM}$

$m_Y$

$m_{DM}$

# Benchmark models

Benchmark Model		Label		
		Individual	Multimodel	
		Binary	Binary	Multiclass
<div>1.5M events</div> <div>0.5M events</div>	SM only	0	0	0
	ALPs	1	1	1
	Spin-0 mediator	$m_Y^{s=0} = m_Z, m_{DM} = 300 \text{ GeV}$	1	2
		$m_Y^{s=0} = 1000 \text{ GeV}, m_{DM} = 10 \text{ GeV}$	1	3
		$m_Y^{s=0} = 10 \text{ TeV}, m_{DM} = 10 \text{ GeV}$	1	-
	Spin-1 mediator	$m_Y^{s=1} = m_Z, m_{DM} = 300 \text{ GeV}$	1	4
		$m_Y^{s=1} = 1000 \text{ GeV}, m_{DM} = 10 \text{ GeV}$	1	-
		$m_Y^{s=1} = 10 \text{ TeV}, m_{DM} = 10 \text{ GeV}$	1	5
	Spin-2 mediator	$m_Y^{s=2} = m_Z, m_{DM} = 300 \text{ GeV}$	1	-
		$m_Y^{s=2} = 1000 \text{ GeV}, m_{DM} = 10 \text{ GeV}$	1	6
		$m_Y^{s=2} = 10 \text{ TeV}, m_{DM} = 10 \text{ GeV}$	1	7

**MadGraph5\_aMC@NLO** to generate events with monojets plus missing energy at parton level.  
 Parton shower and hadronization are performed with **Pythia**.  
 Detector-level data is simulated using **Delphes** with the default ATLAS card.

$\sqrt{s} = 14\text{TeV}$

generation level cuts:  $p_T \geq 130\text{GeV}$  and  $|\eta_j| \leq 5$  for the leading jet.

# Neural Networks algorithms

Event-by-event data

Data as 2D histograms → DNNs and CNNs

# DNN: Deep Neural Networks

All the algorithms are constructed in **Python** using the library **Keras** along with **TensorFlow** for backend implementation

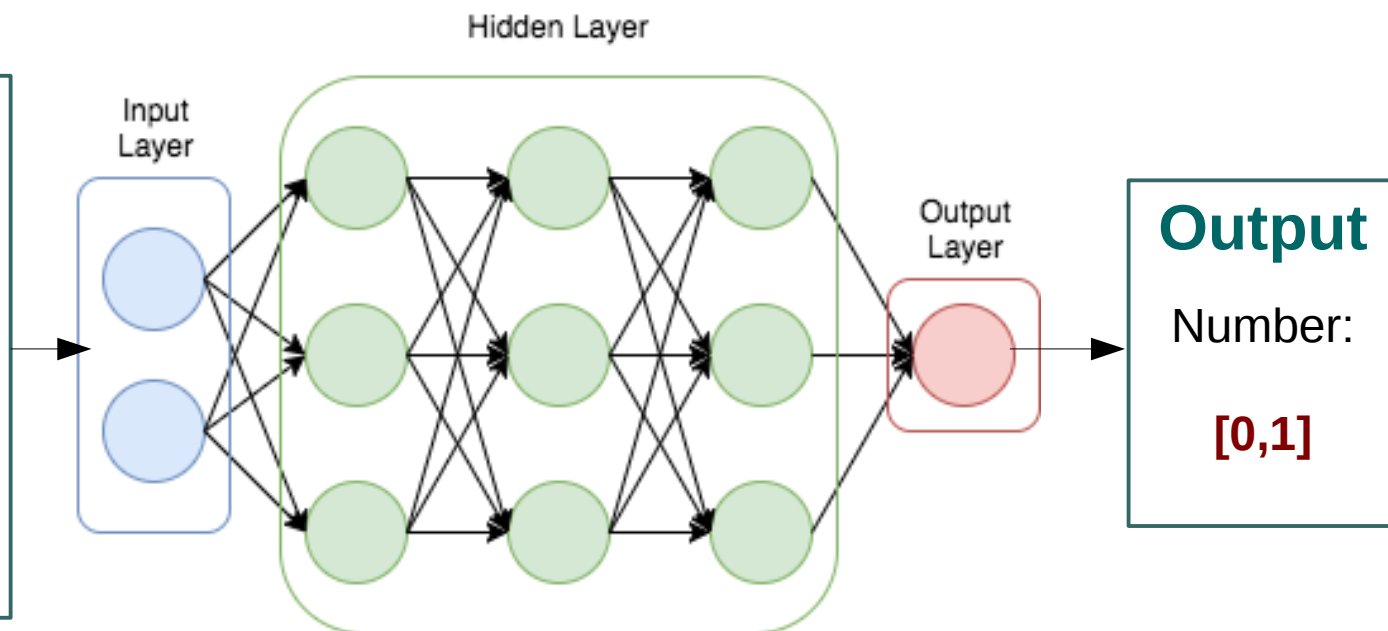
## Supervised Learning

### Training

2 Data sets:

No Signal → SM background  
→ **labeled '0'**

Signal → New Physics  
→ **labeled '1'**



Trained to discriminate data samples with Signal and No Signal

Trained each benchmark model vs SM **individually**.

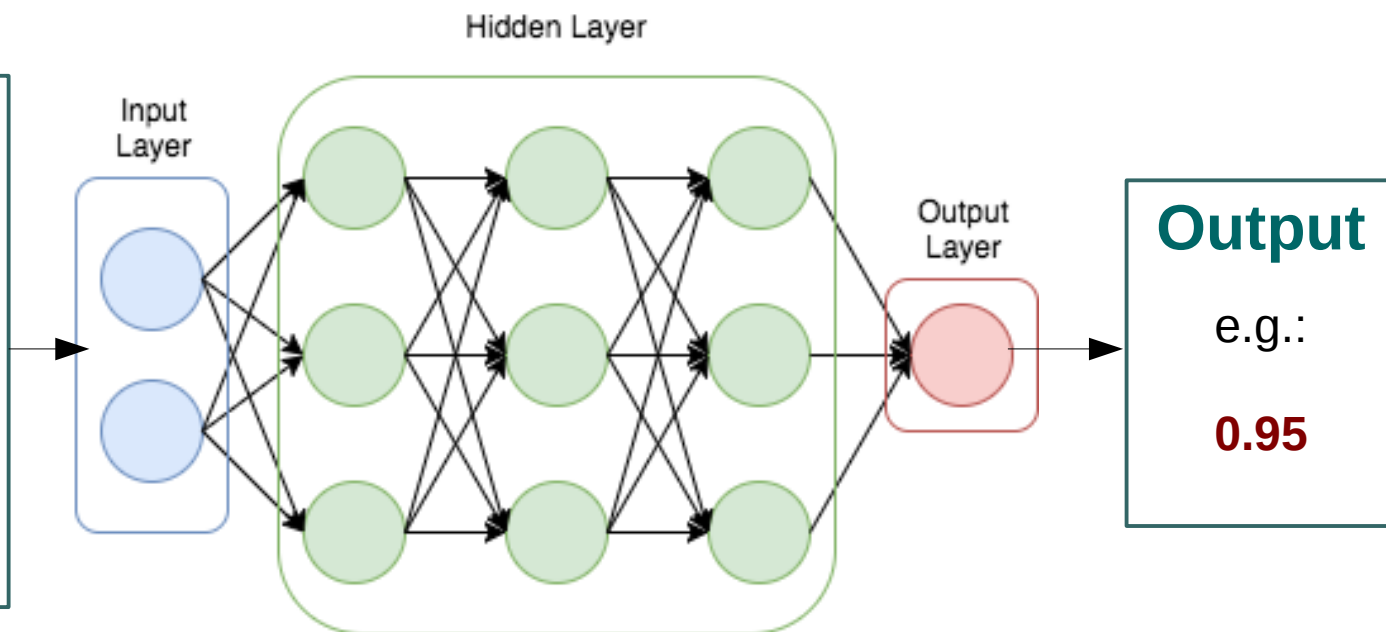
# DNN: Deep Neural Networks

All the algorithms are constructed in **Python** using the library **Keras** along with **TensorFlow** for backend implementation

## Supervised Learning

### Testing

1 data sample  
that I do not know if it is  
Signal or No Signal:



Threshold: 0.5 → Output > threshold

→ The DNN predicted **'1'**  
Signal → New Physics



# DNN with Event-by-event data

We simulated **1.5M SM events** and **0.5M New Physics events**

Each event has the monojet kinematic information  $(p_T^j, \eta^j, \Phi^j)$

## Input

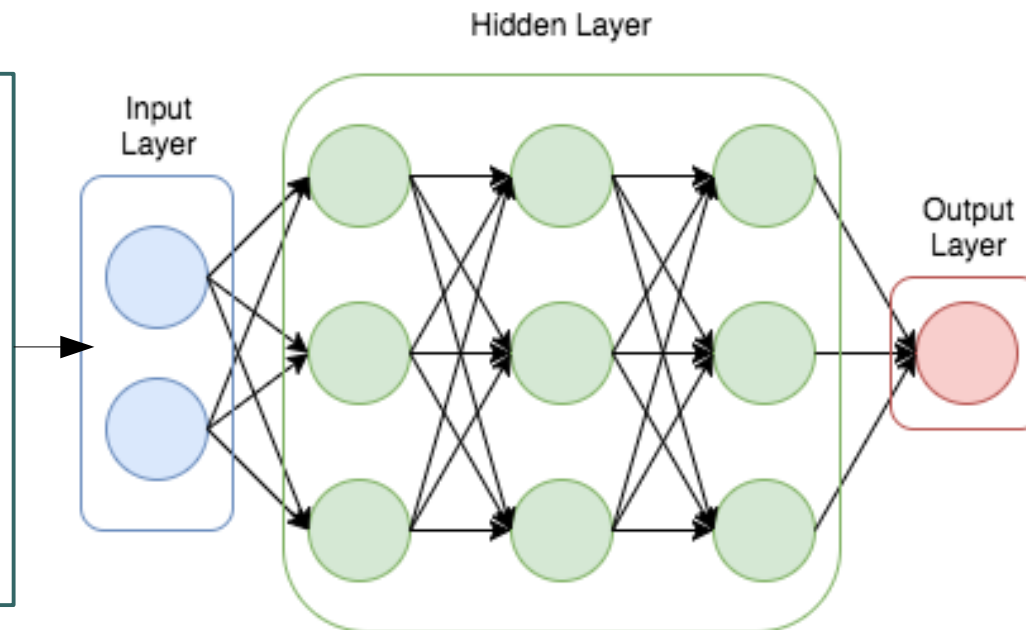
Each data sample is a single event:

Event 1  $\rightarrow (p_T^j, \eta^j, \Phi^j)$

Event 2  $\rightarrow (p_T^j, \eta^j, \Phi^j)$

...

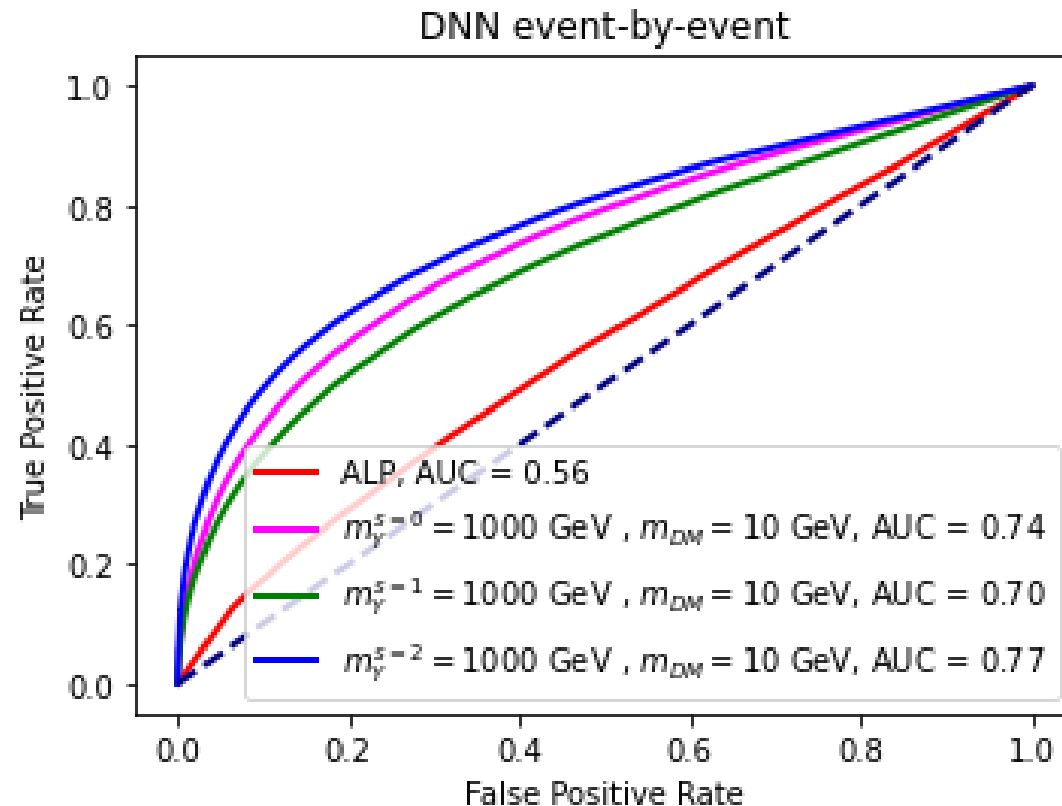
Event N  $\rightarrow (p_T^j, \eta^j, \Phi^j)$



Data samples are divided with a 0.64:0.20:0.16 train-test-validation ratio

# DNN with Event-by-event data

Receiver Operating Characteristic (ROC) curves:

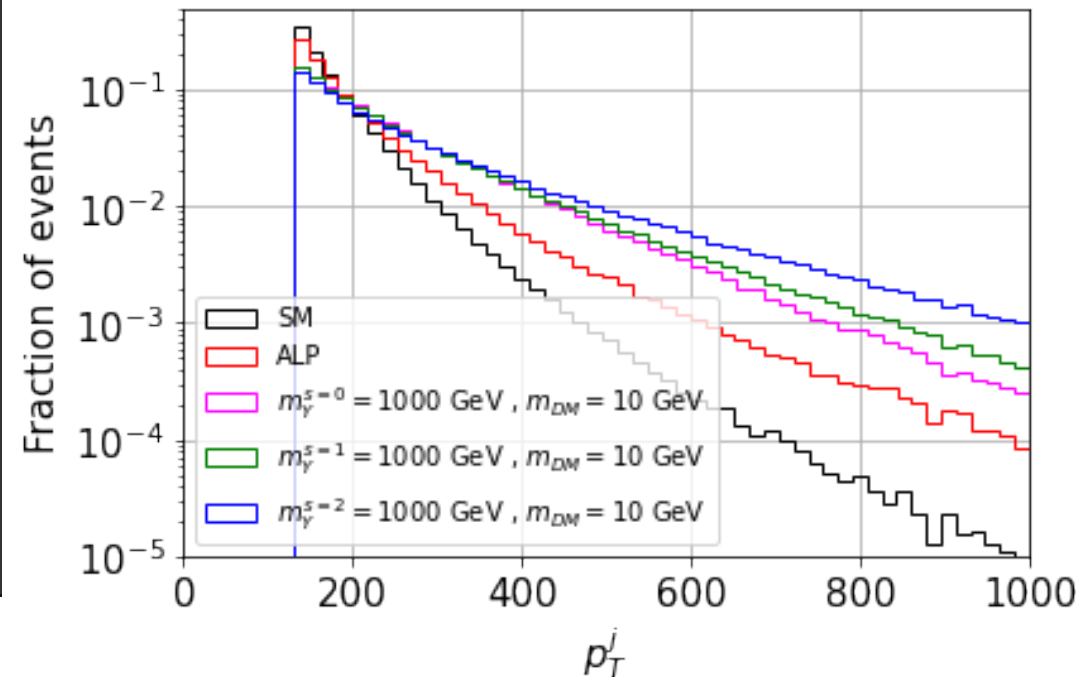
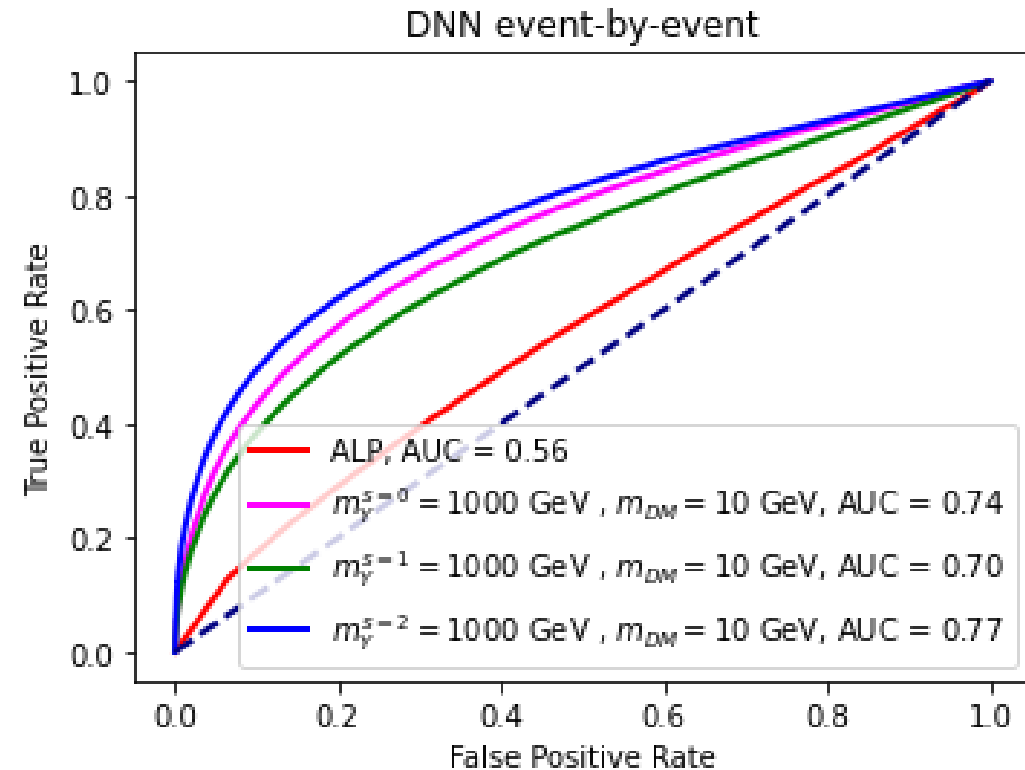


The area under the ROC curve (AUC), a conventional metric to test the performance of binary classifiers

**AUC=1** is a perfect classifier, and **AUC=0.5** represents a random classifier

# DNN with Event-by-event data

Receiver Operating Characteristic (ROC) curves:



The area under the ROC curve (AUC), a conventional metric to test the performance of binary classifiers

**AUC=1** is a perfect classifier, and **AUC=0.5** represents a random classifier

# DNN with data as 2D histograms

S: # NP events  
B: # SM events

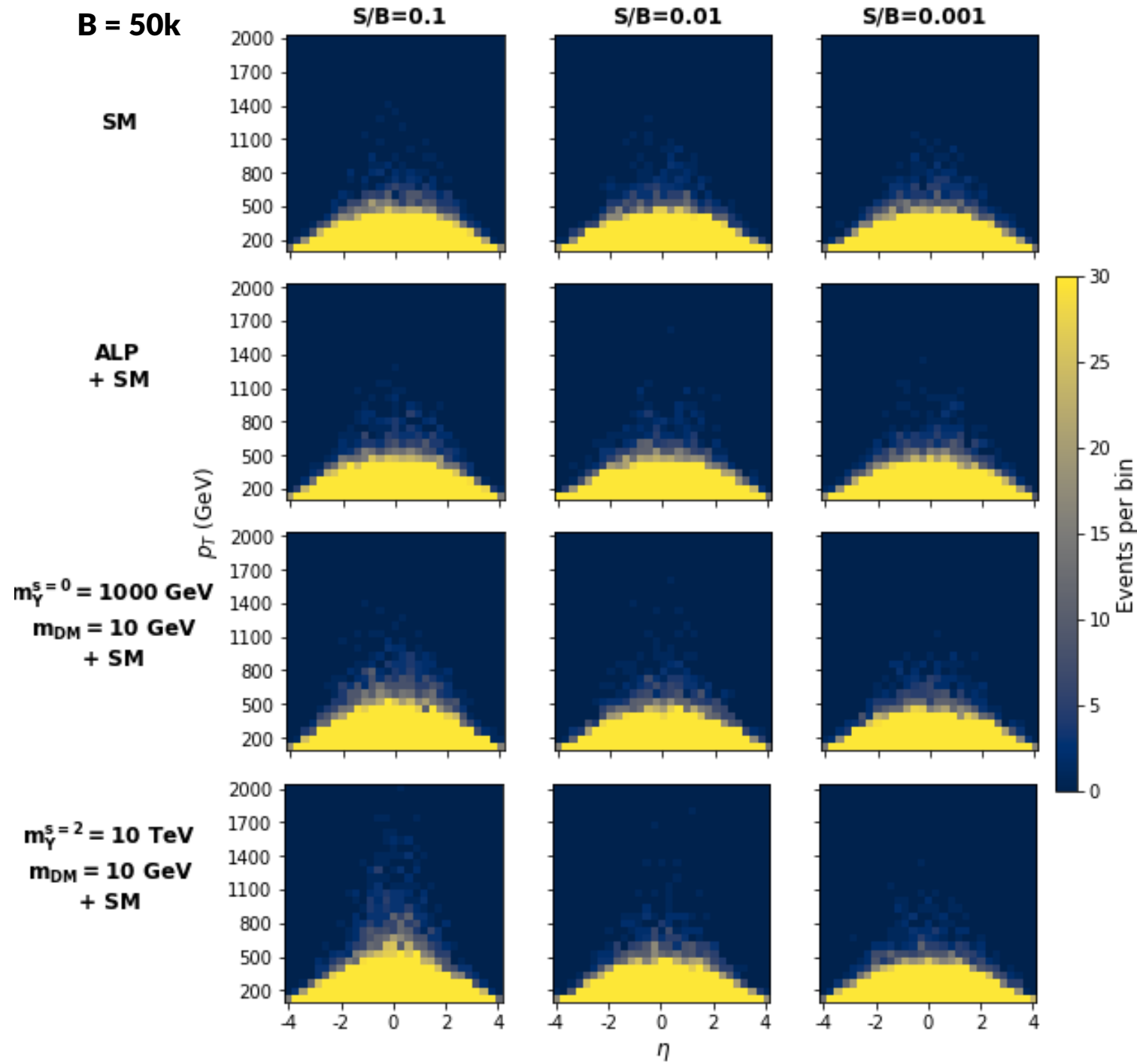
The jet azimuthal angle  $\Phi^j$  does not provide any useful information.

We can construct 2D histograms made from the pair  $(p_T^j, \eta^j)$

**. 20k histograms**  
with only SM events

**. 20k histograms**  
with NP + SM events

↓  
per benchmark model  
and per S/B ratio



# DNN with data as 2D histograms

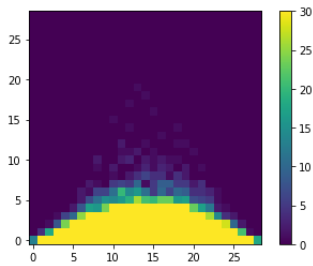
We simulated

**20k SM only histograms** and

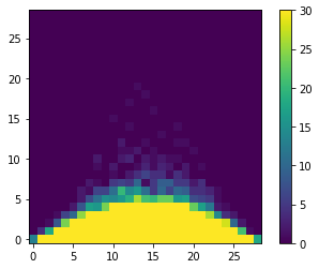
**20k New Physics + SM histograms** (per benchmark model and per S/B ratio)

## Input

Each data sample is a single histogram:



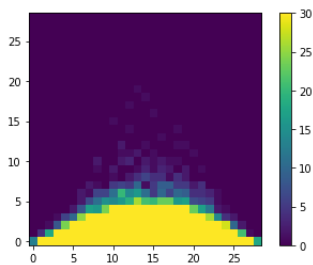
Data sample 1  
**SM only**  
**labeled '0'**



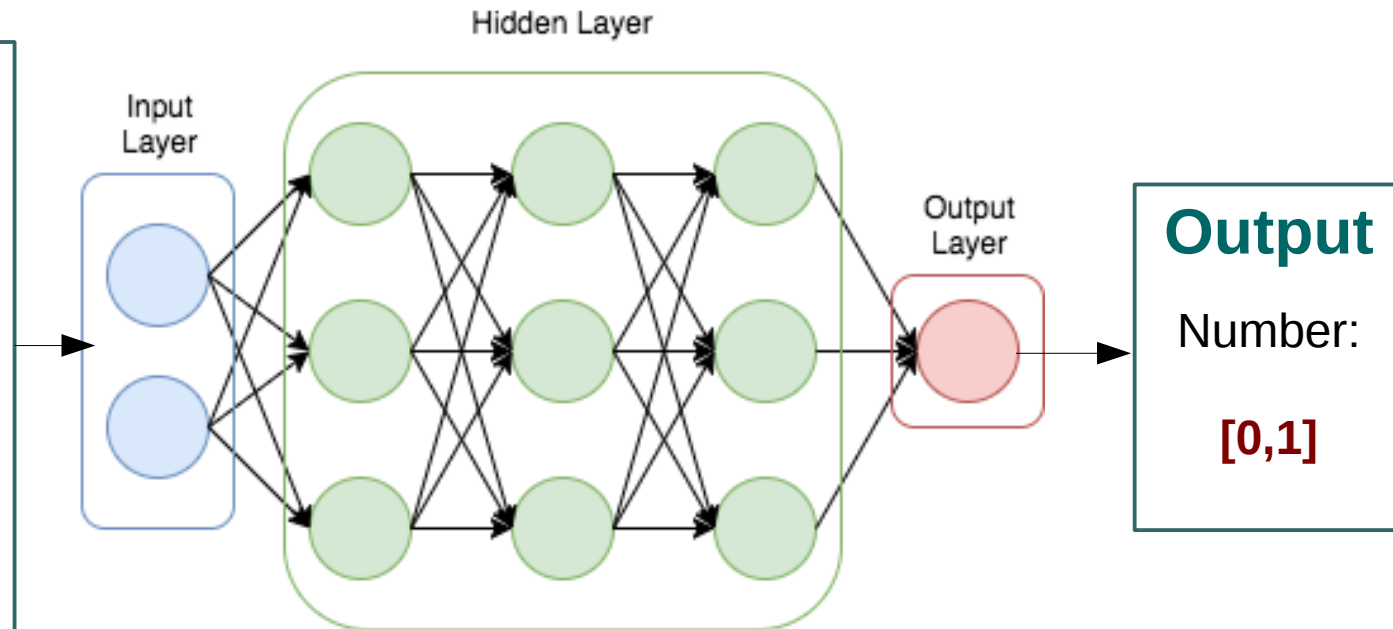
Data sample 2  
**NP + SM**  
**labeled '1'**

...

...



Data sample N  
**SM only**  
**labeled '0'**



DNN trained to discriminate:

**histograms** with SM only events vs  
**histograms** with NP+SM events

Trained each benchmark model vs SM  
**individually.**

# DNN with data as 2D histograms

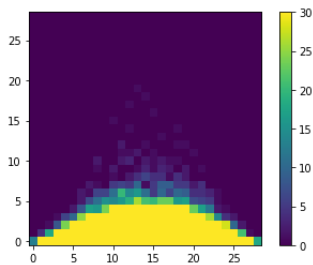
We simulated

**20k SM only histograms** and

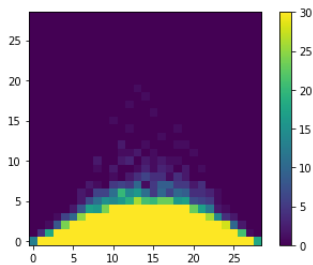
**20k New Physics + SM histograms** (per benchmark model and per S/B ratio)

## Input

Each data sample is a single histogram:

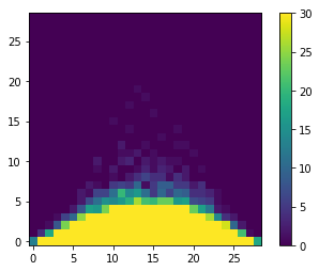


Data sample 1  
**SM**  
labeled '0'

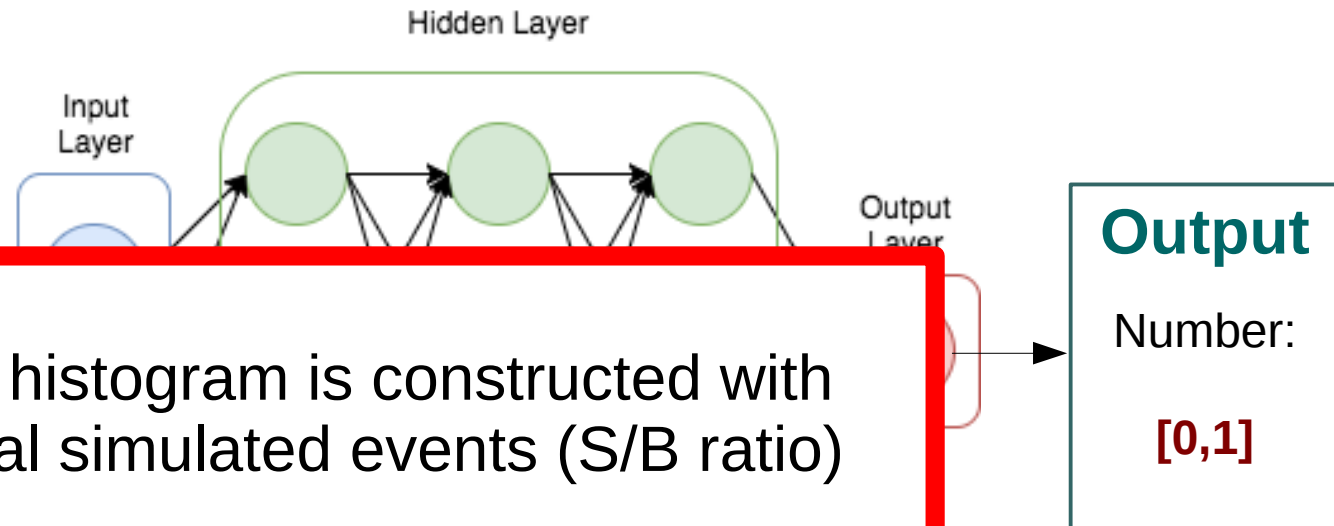


Data sample 2  
**NP**  
labeled '1'

...



Data sample N  
**SM only**  
labeled '0'



## Output

Number:

**[0,1]**

Each histogram is constructed with several simulated events (S/B ratio)

BUT now the DNN considers each histogram as a single entity

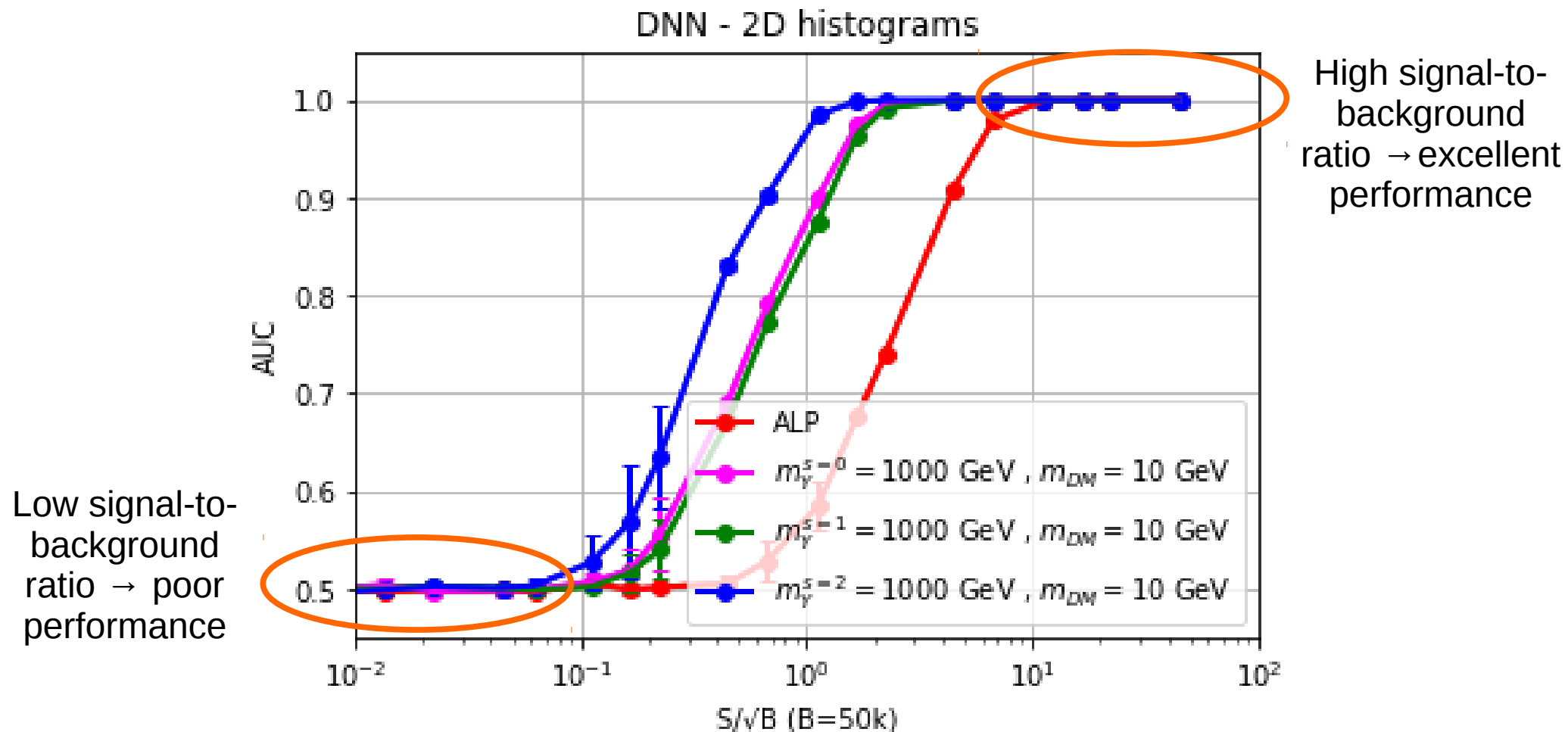
**1 data sample = 1 histogram**

rate:  
events vs  
events

Trained each benchmark model vs SM  
**individually.**

# DNN with data as 2D histograms

Each point represents a DNN trained with a data set with a specific benchmark, S and B

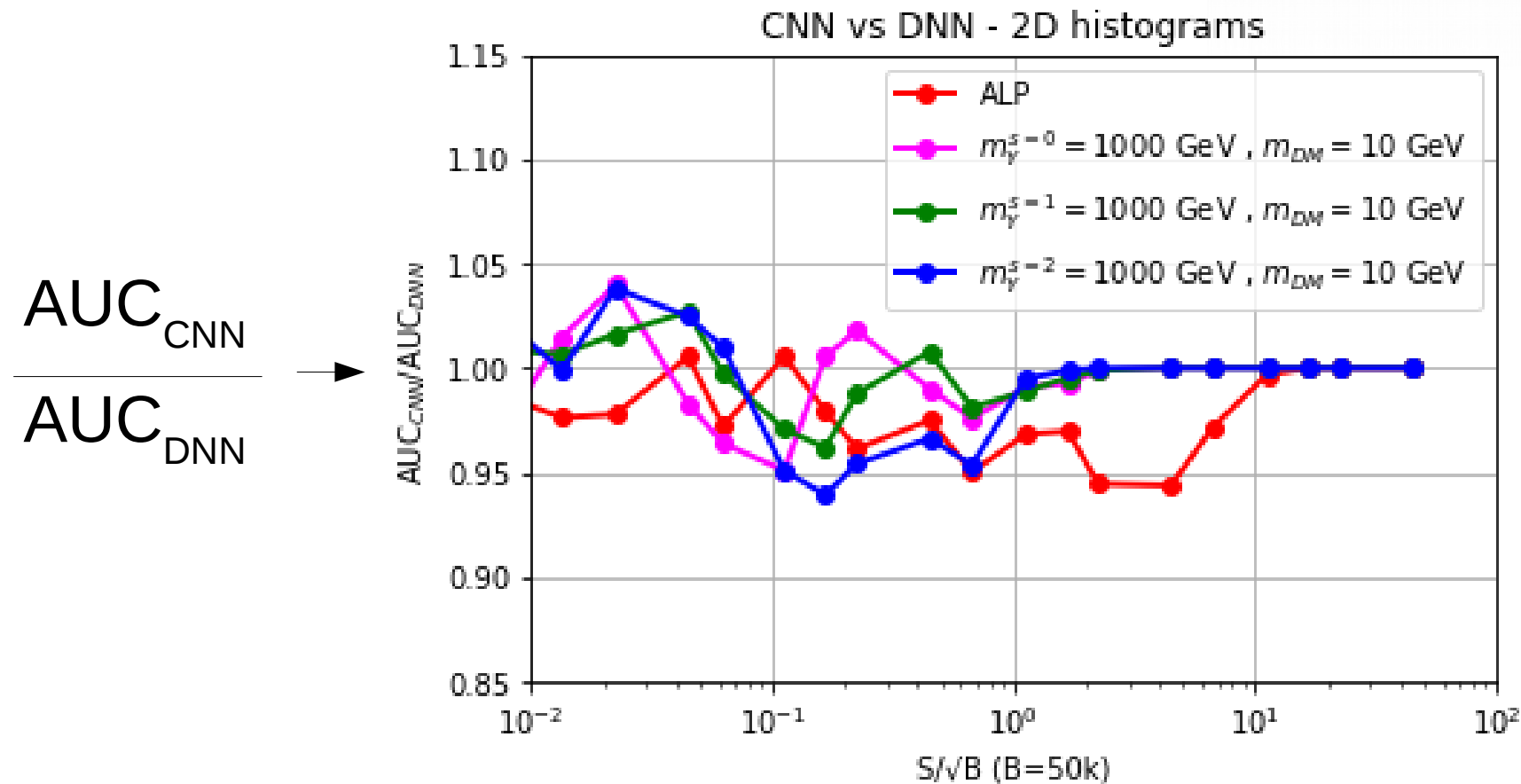
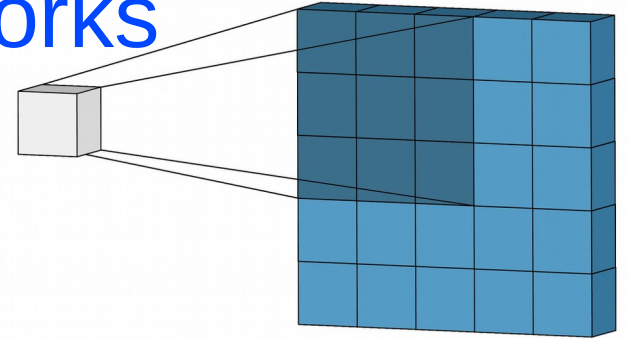


**Huge performance boost** w.r.t. event-by-event representation!

**AUC=1** is a perfect classifier, and **AUC=0.5** represents a random classifier

# CNN: Convolutional Neural Networks

- Image recognition.
- Have other structure (kernel layers, etc).
- Scan the image by parts looking for patterns.
- Demand significantly more computing time than DNN



Differences are within  $\sim 5\%$  for all the models considered

→ no improvement with respect to the DNN is found

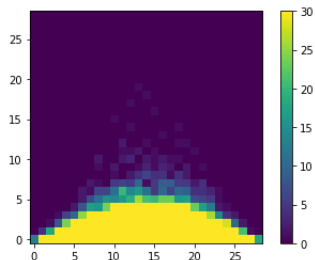


# Flexibility of the method

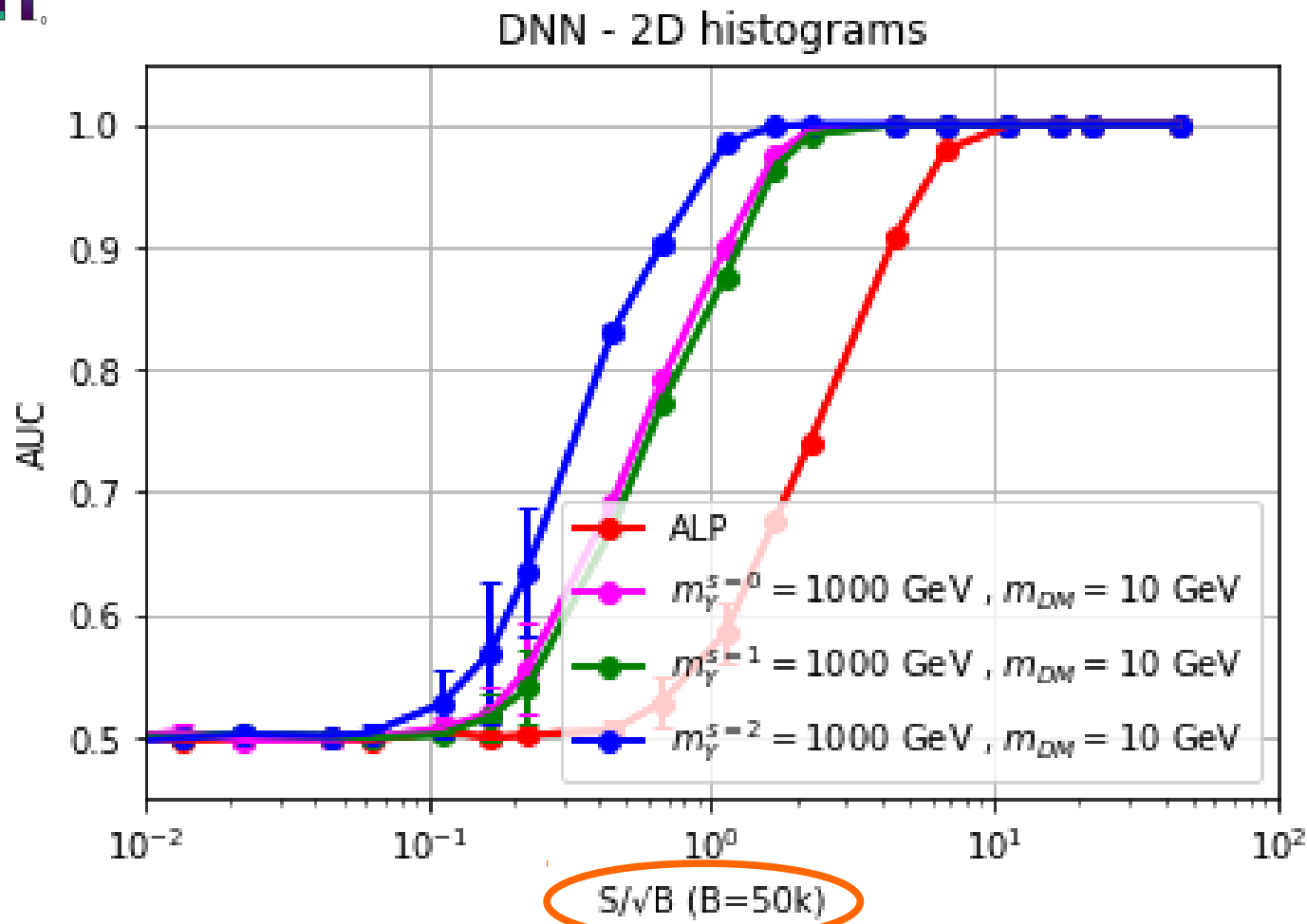
Performance invariance with the number of background events  
Testing under incorrect training hypothesis: - coupling values  
- benchmark model

# Performance invariance with B

S: # NP events  
B: # SM events

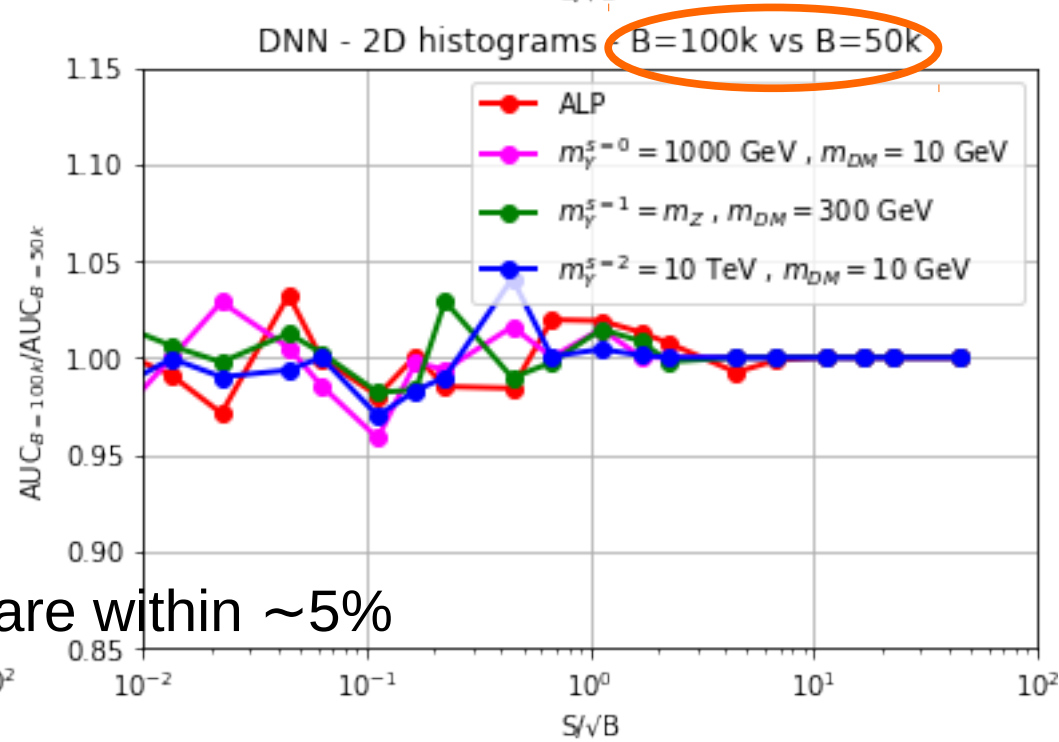
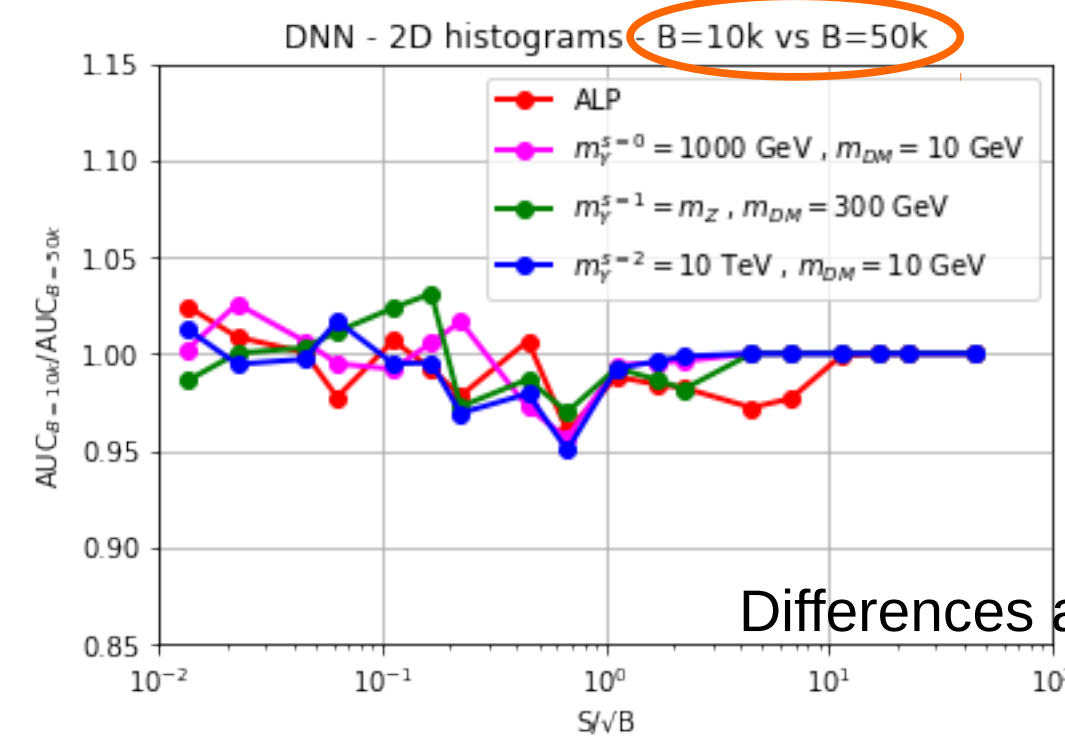
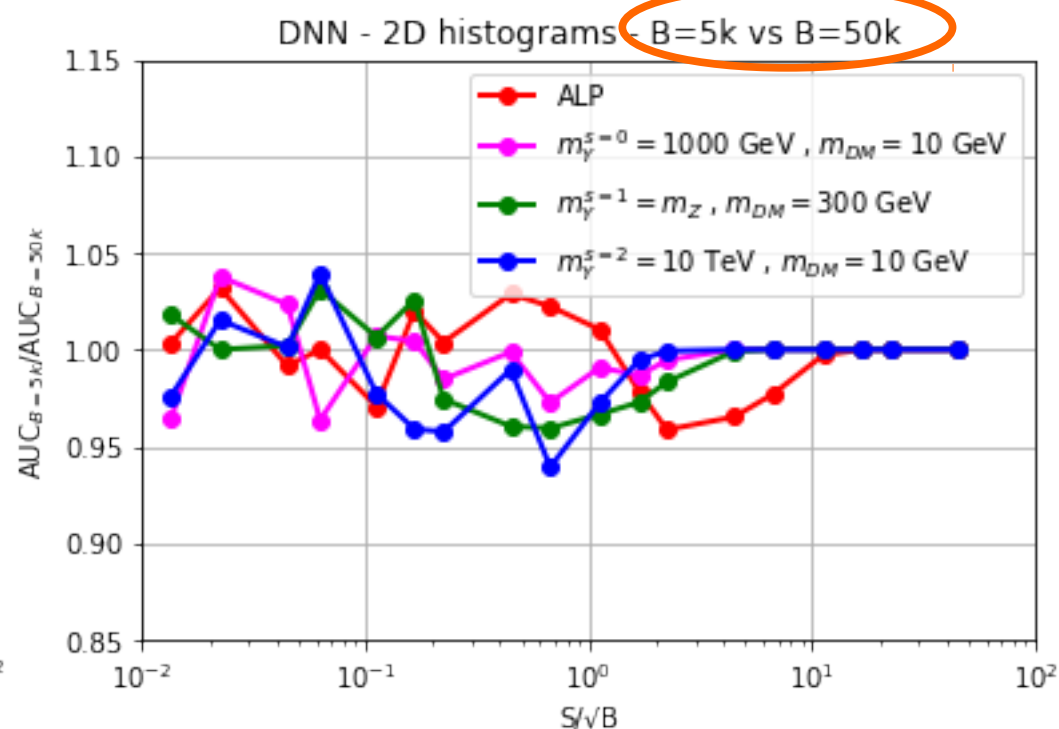
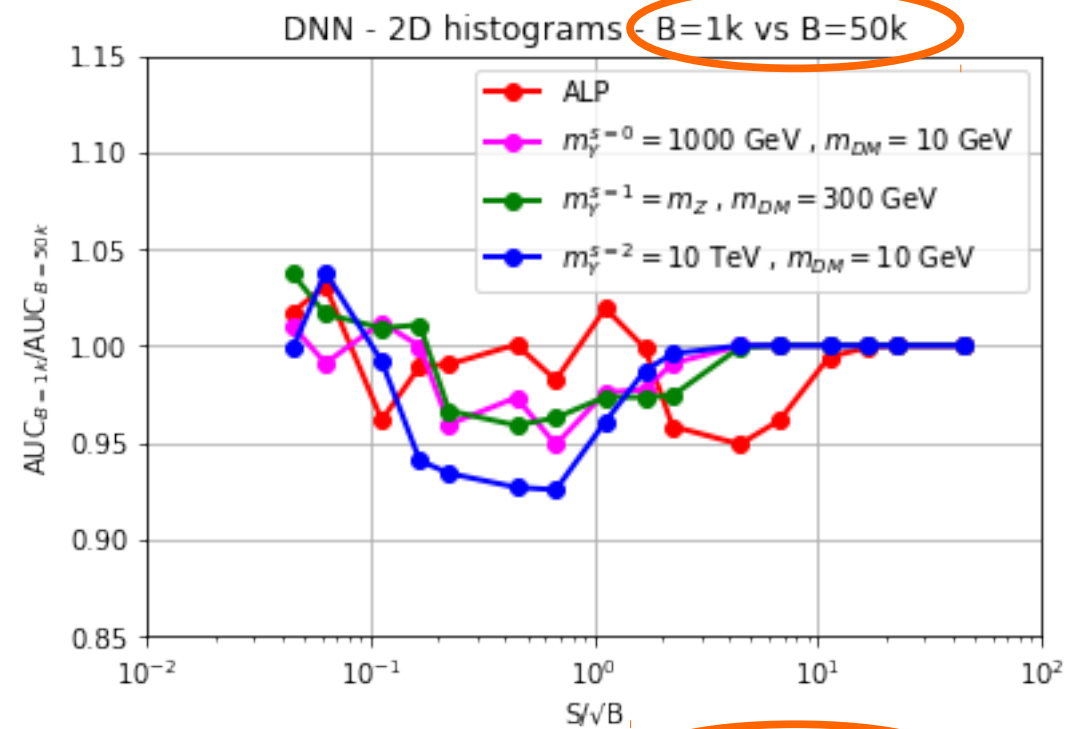


Before, each 2D histogram constructed with  
**B = 50k** SM events + **S** New Physics events



Performance is not modified significantly for different values of B, if the results are presented as a function of  **$S/\sqrt{B}$** .

# Performance invariance with B



Differences are within ~5%

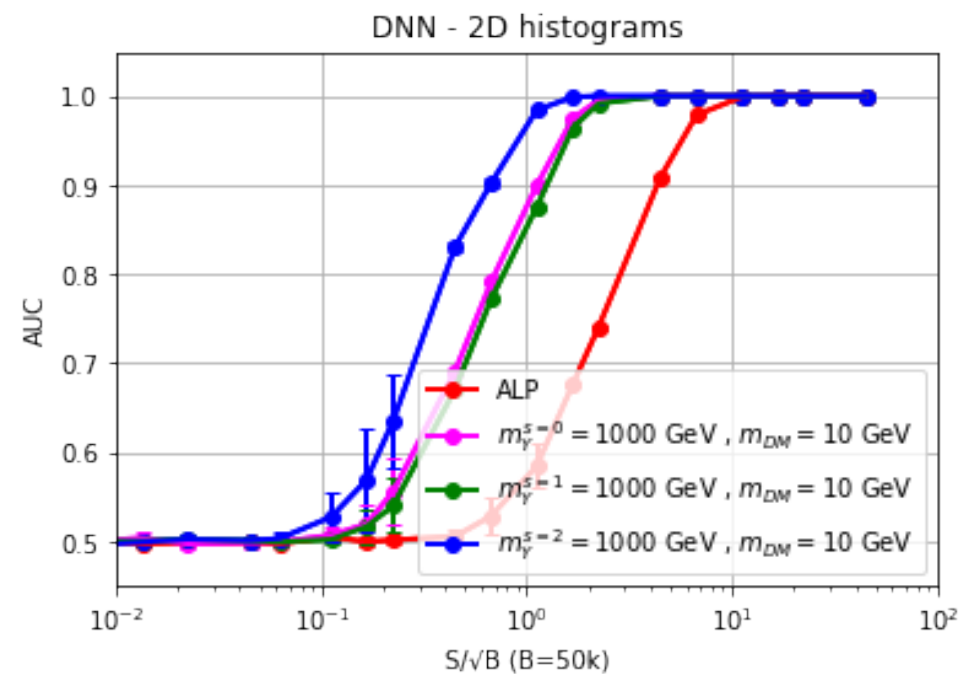
# Performance invariance with B

To know if a DNN or CNN with 2D histograms could distinguish a particular new physics model from the SM background, we only need to:

- **Identify the curve of the corresponding benchmark model**
- **Calculate the model cross section for the chosen couplings**
- **Calculate the SM background cross section**
- **Calculate  $S/\sqrt{B}$  for any luminosity, and check the corresponding AUC**

Also, we can have an idea of the luminosity needed to obtain a given efficiency. Change the last step for:

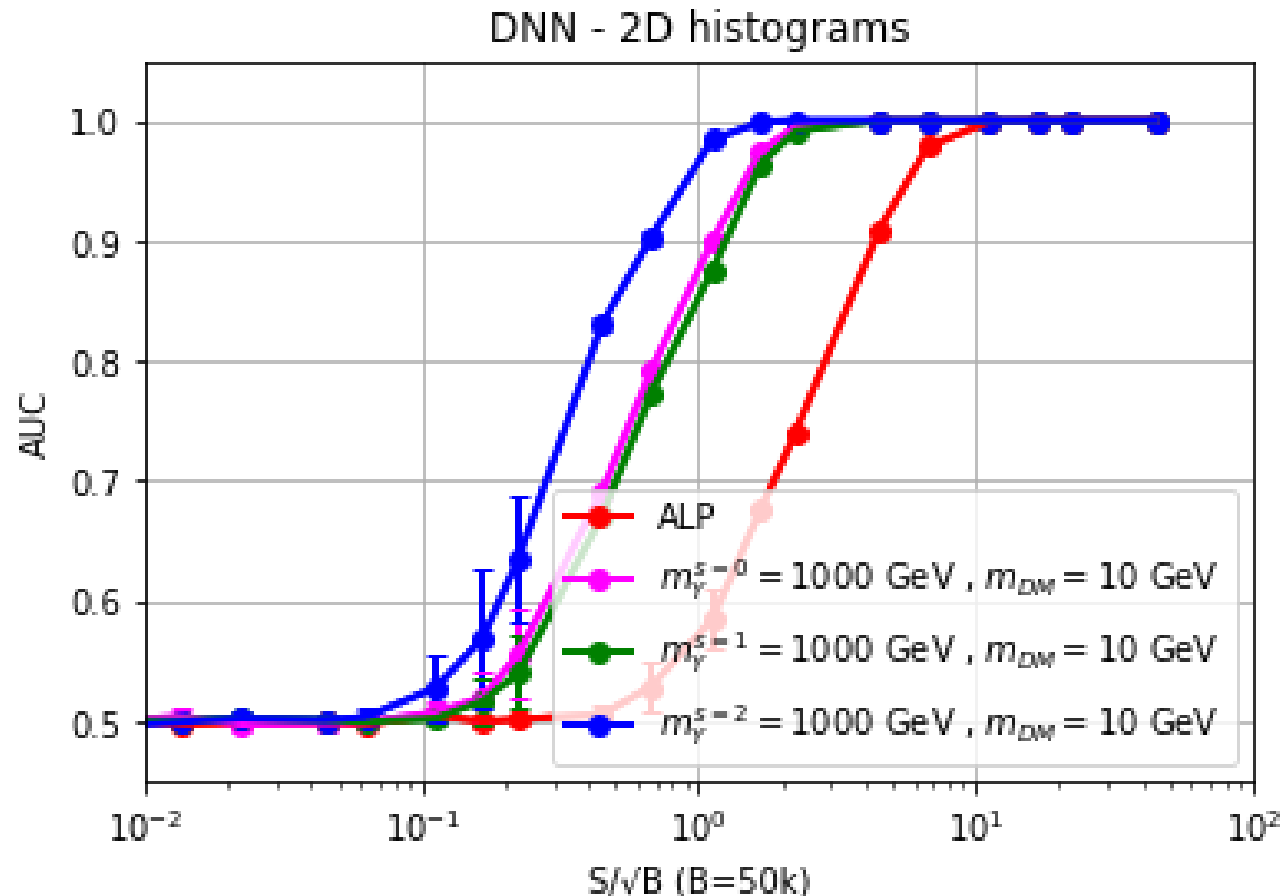
- **Identify the  $S/\sqrt{B}$  value for the corresponding AUC you would like to get and calculate the luminosity needed**



$$\# \text{ events} = \text{cross section} * \text{luminosity} * \text{detector efficiency}$$

# Performance invariance with B

S: # NP events  
B: # SM events



$S/\sqrt{B}$  is a histogram parameter



each 2D histogram constructed with  
**B** SM events + **S** New Physics events

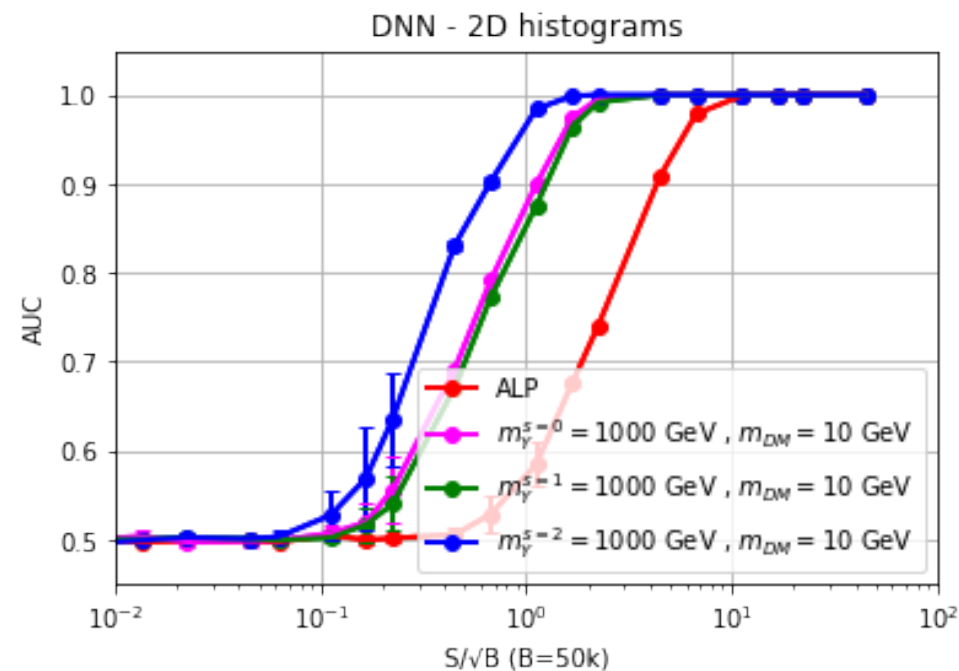
It is not the significance in an  
usual counting experiment

( 1 data sample = 1 histogram )

# Preguntas abiertas

¿Cuál es la relación entre AUC y la significancia de excluir la hipótesis nula?

¿Cómo estimar la significancia del método?



¿Cómo es el proceso de aplicación del método?

¿Tiene sentido construir más de 1 histograma con los datos experimentales?

Disminuye  $S/\sqrt{B} \rightarrow$  i.e. disminuye el AUC de la red, pero ¿tengo más estadística?

# Testing under incorrect training hypothesis

A DNN is prepared to handle the same kind of data that it has been trained with, i.e. matching train and test data samples (same underlying model).

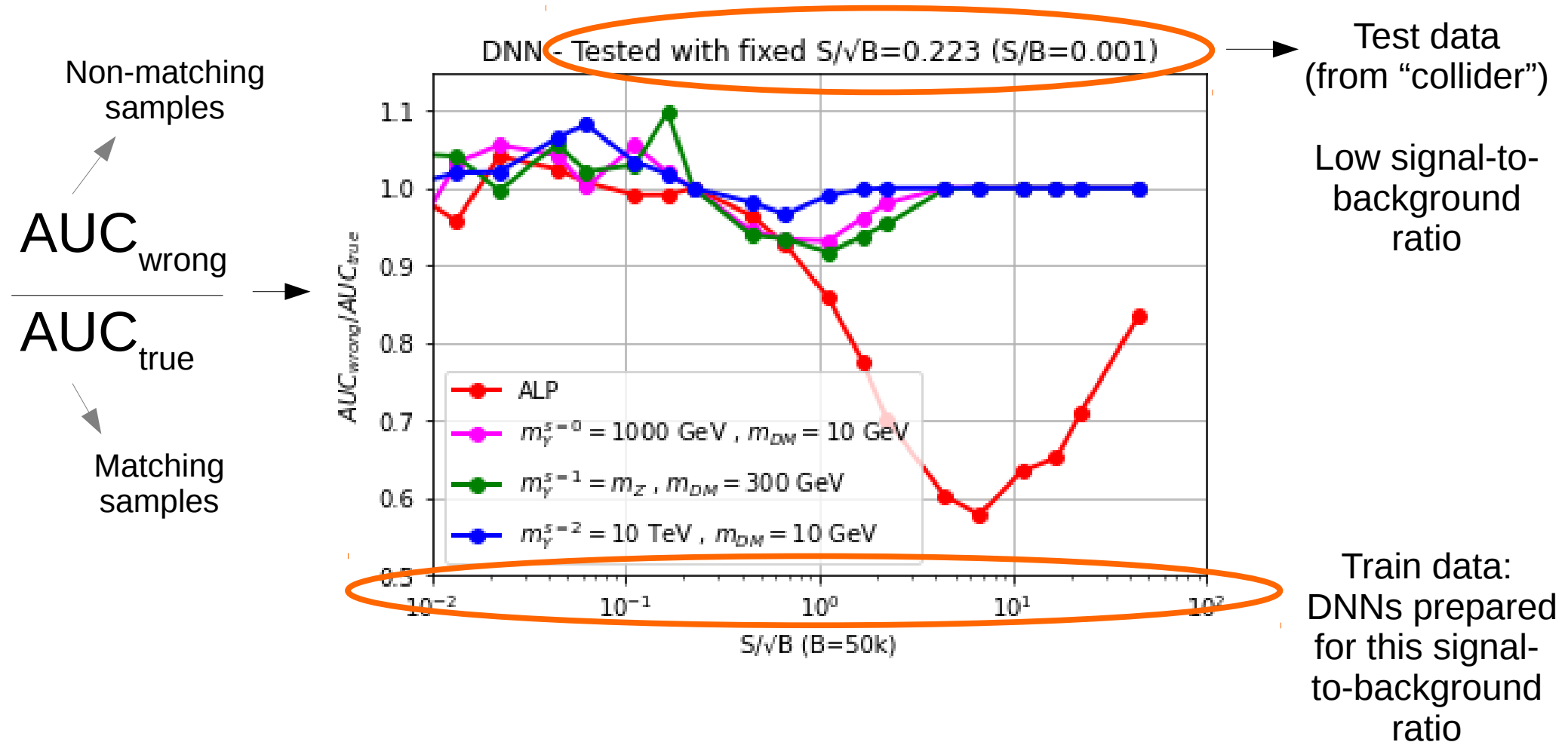
**What is the performance when we apply a test data set to a DNN trained with a different underlying model?**

DNN results with matching data samples are called  $AUC_{\text{true}}$

DNN results with non-matching data samples are called  $AUC_{\text{wrong}}$

# Testing under incorrect training hypothesis: *coupling values*

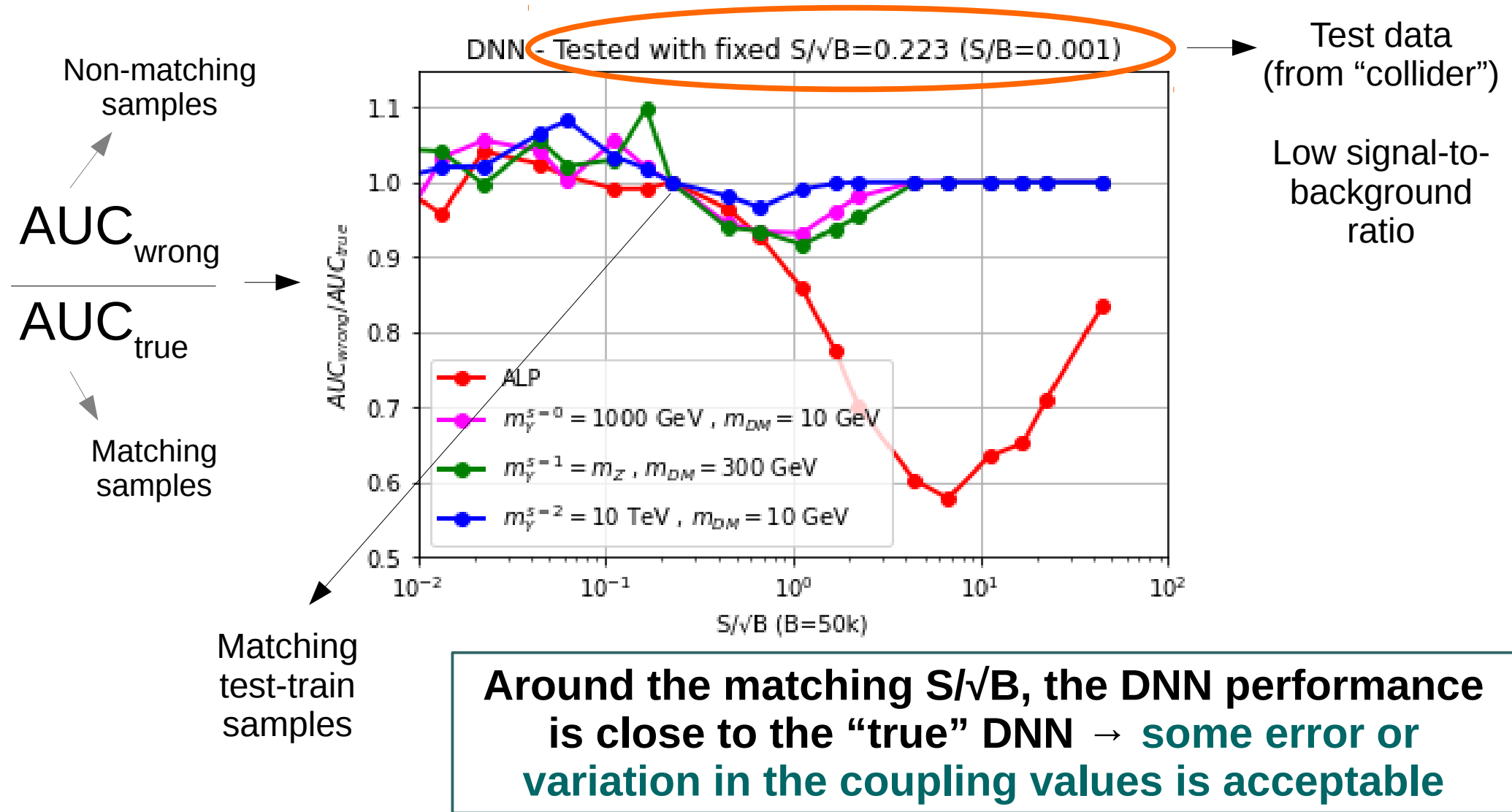
Train and test data samples are generated with the **same benchmark models**, but with **different S/B ratios**, i.e. a same model with **different coupling values**.





# Testing under incorrect training hypothesis: *coupling values*

Train and test data samples are generated with the **same benchmark models**, but with **different S/B ratios**, i.e. a same model with **different coupling values**.



# Testing under incorrect training hypothesis: *benchmark model*

Train and test data samples are generated with **different benchmark models**, but with the **same S/B ratios**.

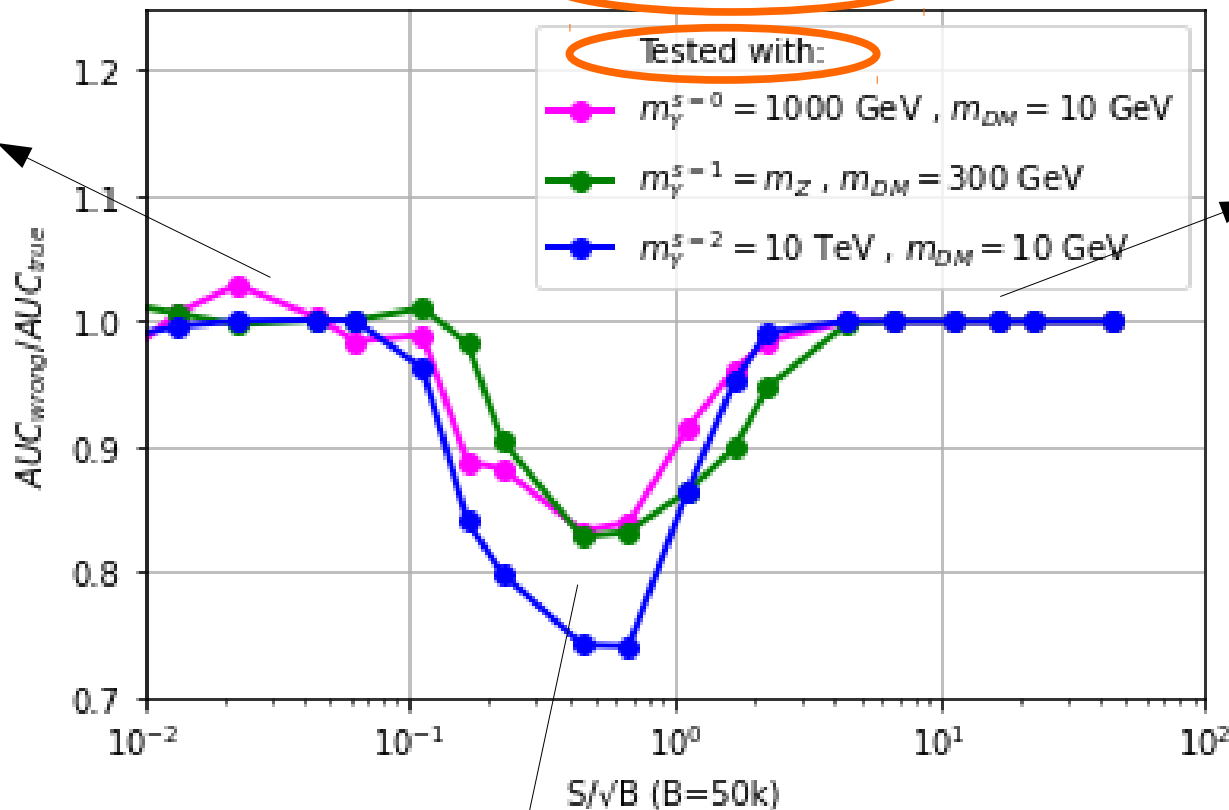
DNN - Trained with ALP

Tested with:

$m_Y^{S=0} = 1000 \text{ GeV}, m_{DM} = 10 \text{ GeV}$

$m_Y^{S=1} = m_Z, m_{DM} = 300 \text{ GeV}$

$m_Y^{S=2} = 10 \text{ TeV}, m_{DM} = 10 \text{ GeV}$



Low signal-to-background ratio

Can NOT discriminate NP vs SM regardless the underlying model.  
“Few signal”

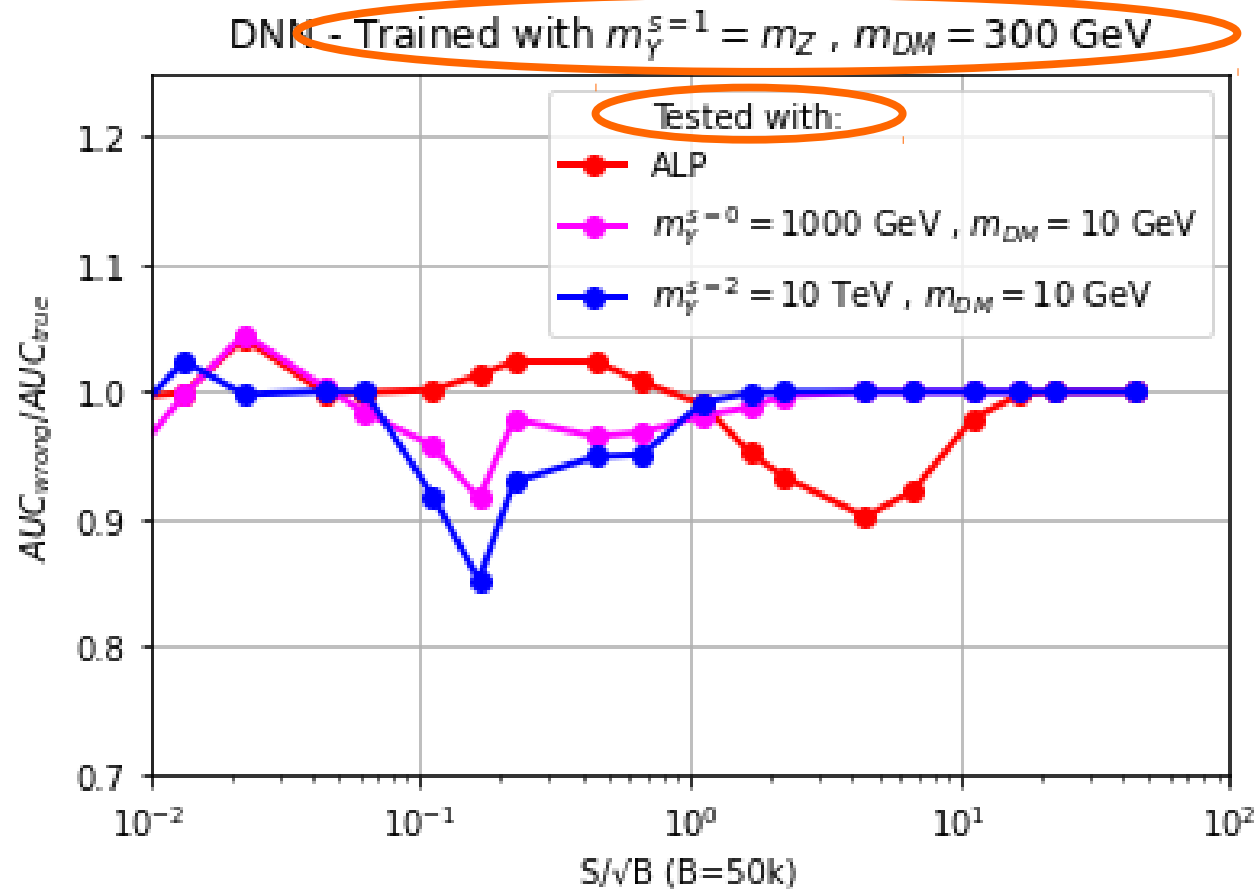
High signal-to-background ratio

Can discriminate NP vs SM regardless the underlying model.  
“A lot of signal”

Intermediate S/√B values  
discrepancies can be huge, up to 25%

# Testing under incorrect training hypothesis: *benchmark model*

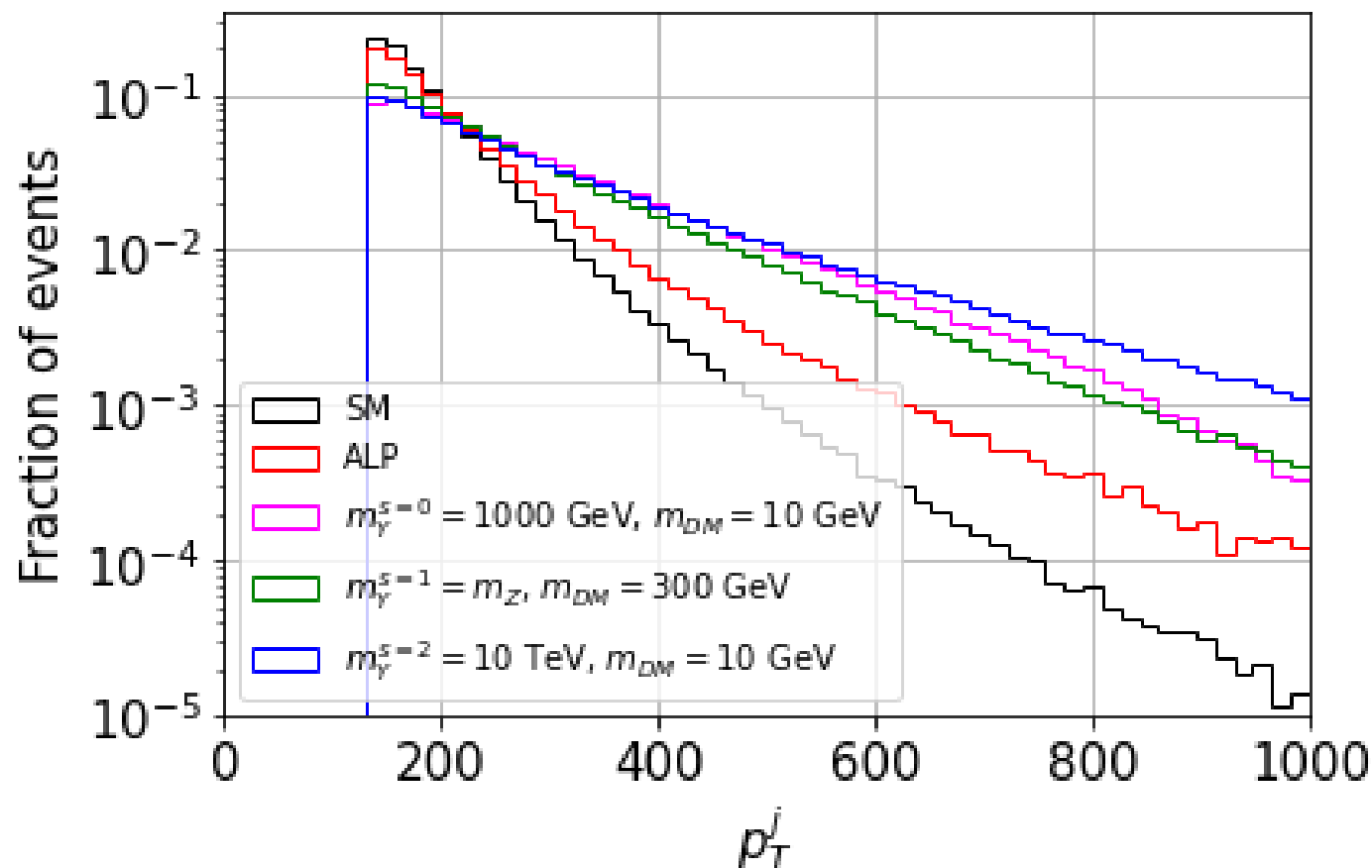
Another example



Performance variations within  $\sim 10\%$

**The DNN classifies “kinematic distributions” not “models”**

The DNN classifies “kinematic distributions” not “models”



**Benchmark models with “similar” kinematic distributions result in “similar” DNN performances**

# Multimodel classifiers

- Binary classifier
- Multiclass classifier

# Binary classifier

A single DNN per S/B value to discriminate, regardless the underlying model, between

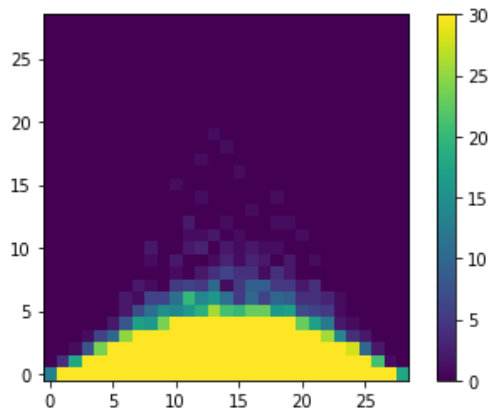
**SM background**  
(labeled '0')

vs

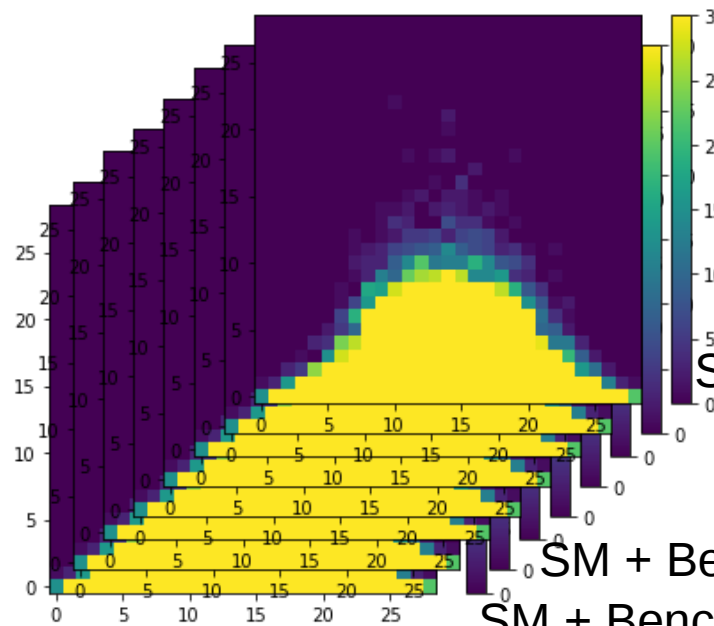
**non-SM processes**  
(labeled '1')

One DNN trained with several benchmark models

## Training the DNN



SM only → "0"



SM + Benchmark 2 → "1"  
SM + Benchmark 1 → "1"

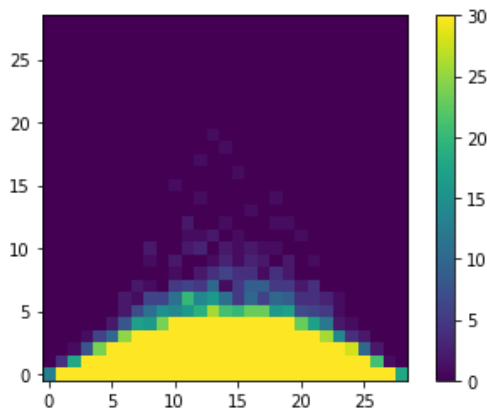
10k histograms per  
benchmark and per  
S/B ratio

SM + Benchmark 7 → "1"

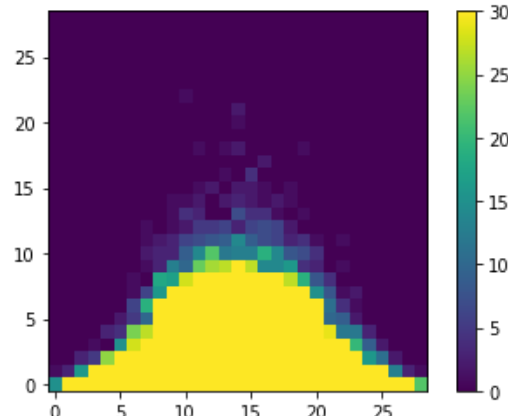
DNN

# Binary classifier

## Test the DNN



SM only → "0"



SM + Benchmark i → "1"

DNN

Value between [0,1]

NP + SM  
or  
SM only

ROC curve → AUC as performance

Test the network  
with only **one  
benchmark model  
at a time** to  
evaluate its  
performance,  
denoted  
**AUC<sub>7-models</sub>**

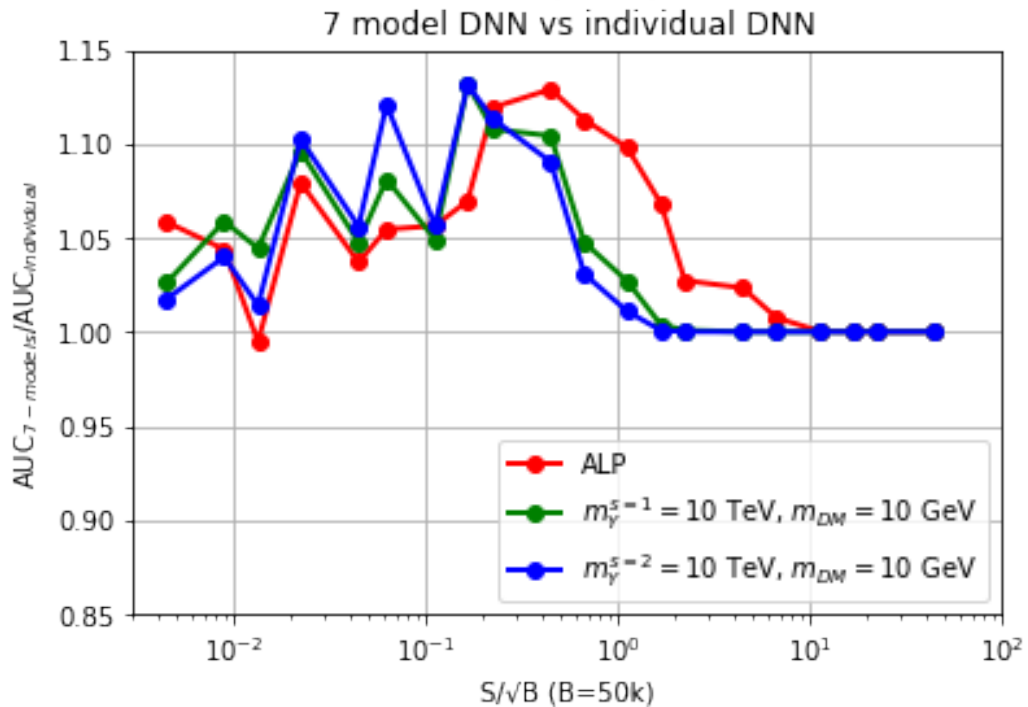
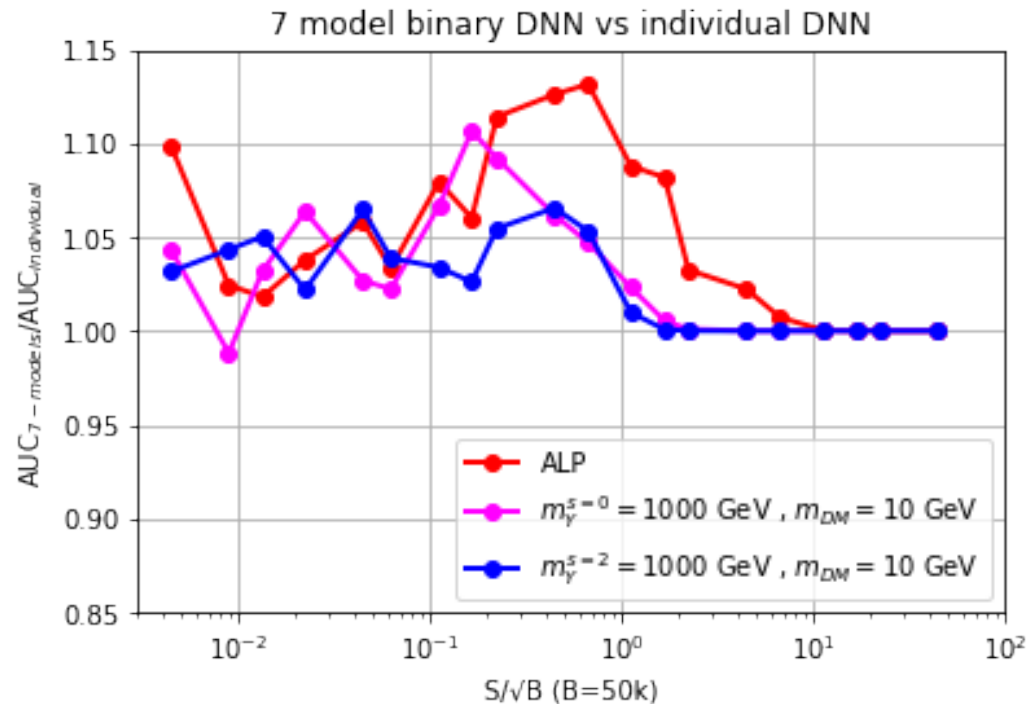
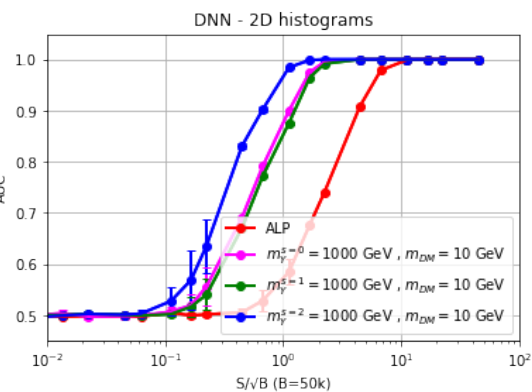
# Binary classifier

Trained with 7  
benchmark  
models, tested  
with only one

$AUC_{7\text{-models}}$

$AUC_{\text{individual}}$

Trained and  
tested with only  
1 benchmark  
model



Testing with  
training models

Some  
**improvement**,  
up to  $\sim 15\%$ , w.r.t.  
individual DNNs,  
but in regions  
with **low AUC**



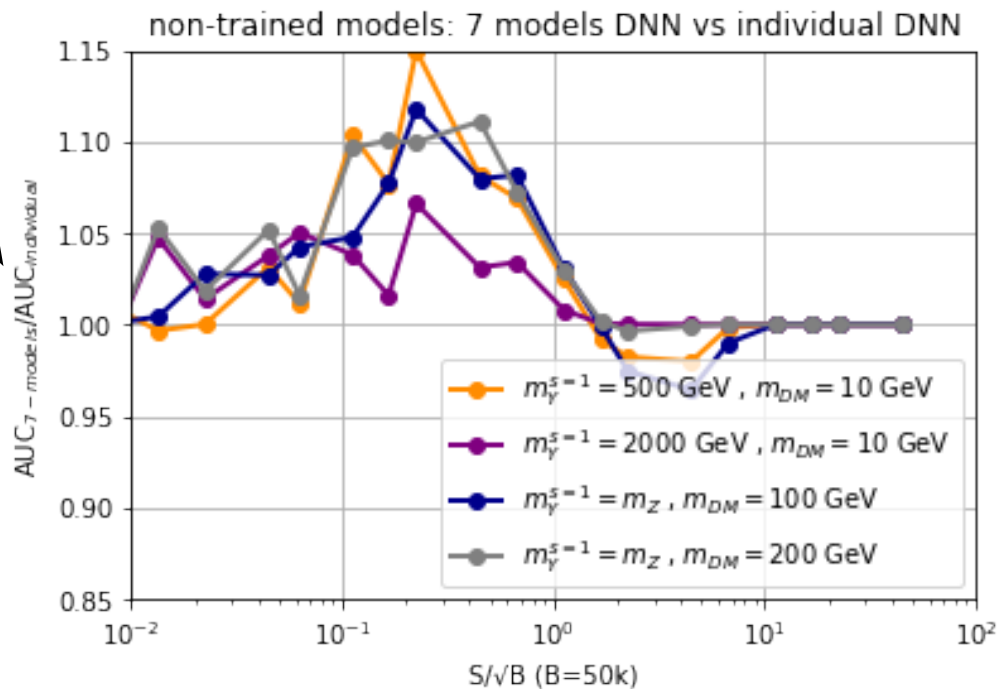
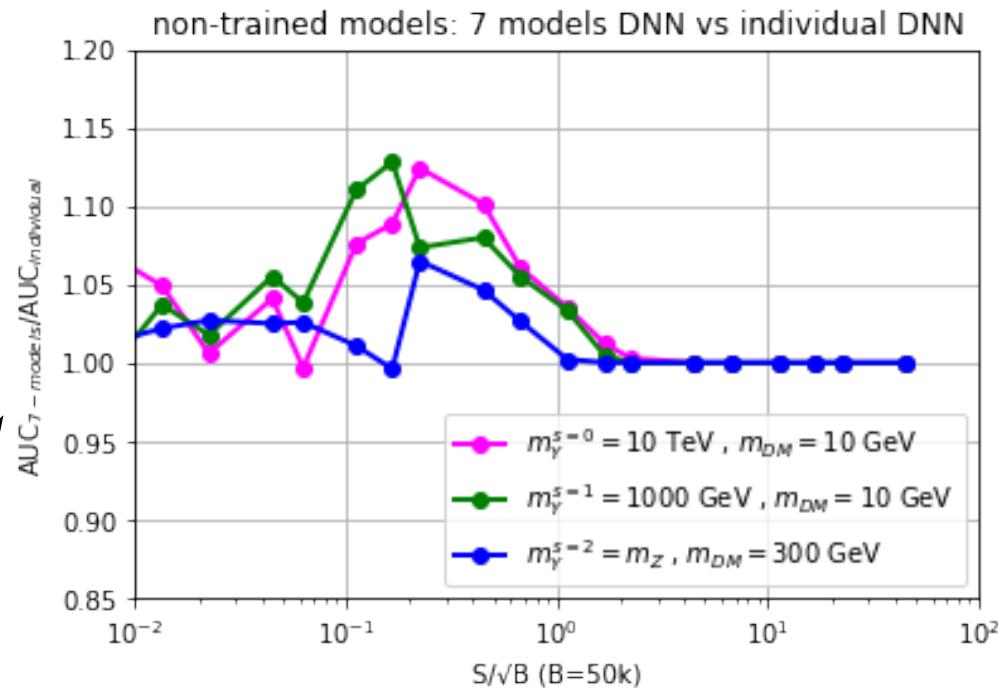
# Binary classifier

Trained with 7  
benchmark  
models, tested  
with only one

$AUC_{7\text{-models}}$

$AUC_{\text{individual}}$

Trained and  
tested with only  
1 benchmark  
model



Models completely  
new to the DNN

Testing with  
non-training  
models

Some  
**improvement**,  
up to  $\sim 15\%$ , w.r.t.  
individual DNNs,  
but in regions  
with **low AUC**

Info about many  
models  
→ helps the  
discrimination of  
NP + SM vs SM only

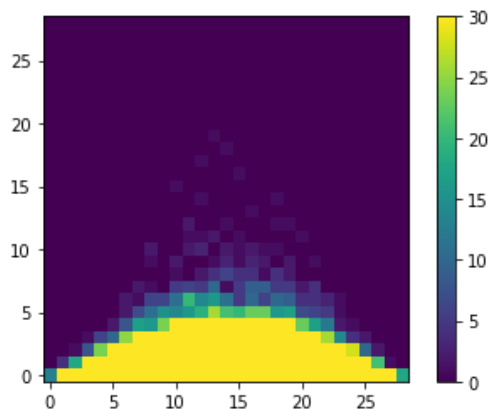
# Multiclass classifier

A single DNN per S/B value to discriminate between:

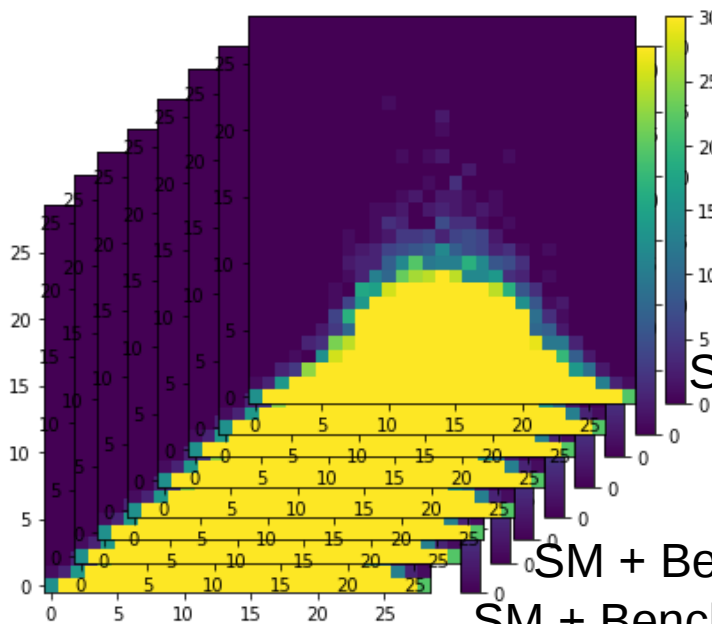
**SM only** (labeled '0') vs **Benchmark 1 + SM** (labeled '1') vs ... vs **Benchmark N + SM** (labeled 'N')

a single DNN trained with **several** new physics models

## Training the DNN



SM only → "0"



SM + Benchmark 1 → "1"

SM + Benchmark 2 → "2"

SM + Benchmark 7 → "7"

10k histograms per benchmark and per S/B ratio

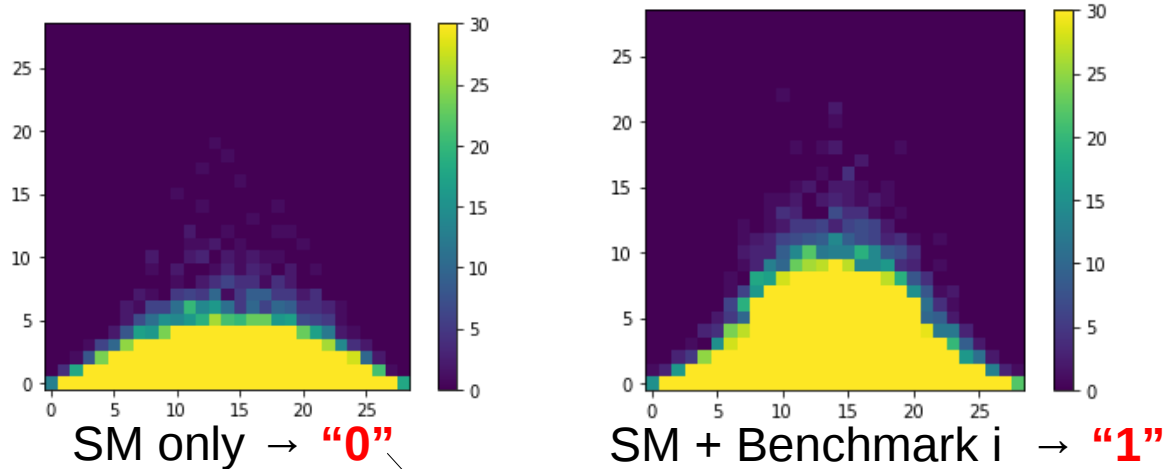
**DNN**

(In this work, 7 NP+SM models and SM only)

# Multiclass classifier

## Test the DNN

Test the network  
with only **one  
benchmark model  
at a time** to  
evaluate its  
performance,  
denoted  
**AUC<sub>7-models</sub>**



**DNN**

Array of 8 elements between **[0,1]**

Each element is the probability that the  
histogram belongs to a training model

(trained with **7 NP+SM** models  
and **SM only**)

For example

Label	<b>0</b>	<b>1</b>	<b>2</b>	<b>3</b>	<b>4</b>	<b>5</b>	<b>6</b>	<b>7</b>
Array	[ 0.09	0.03	0.01	<b>0.84</b>	0.005	0.005	0.01	0.01 ]
	↓			↓				
	Prob(SM)			Prob(benchmark model 3 + SM)				

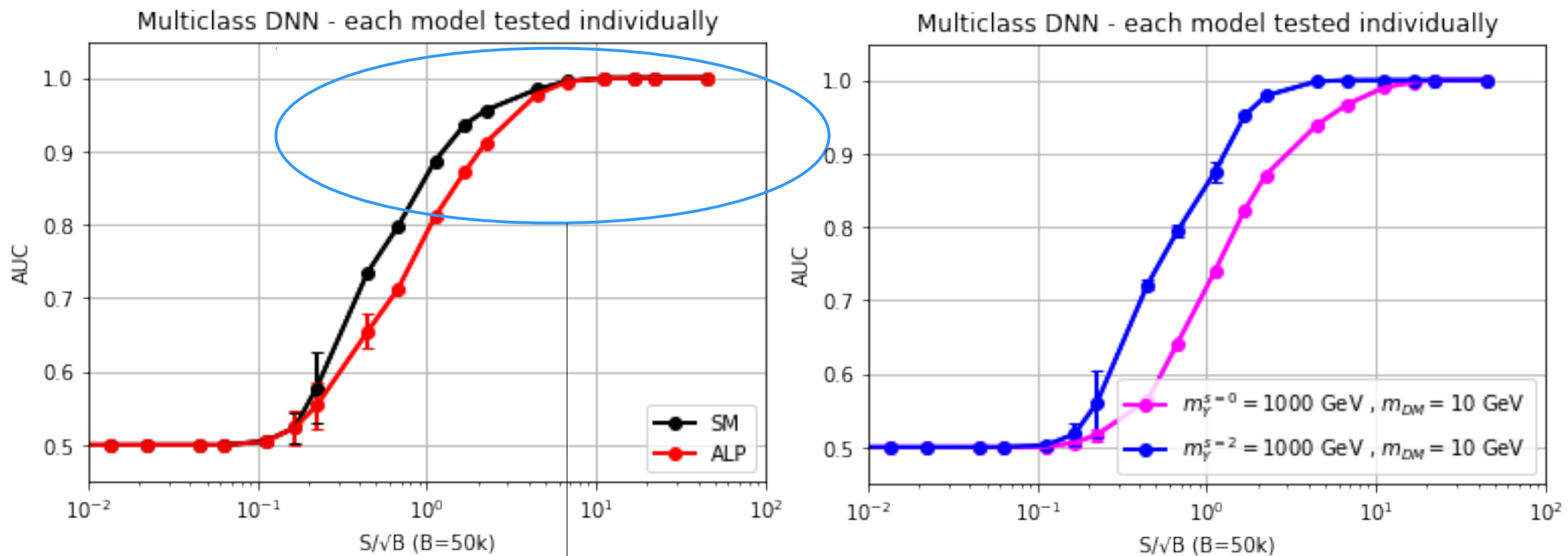
# Multiclass classifier

Convert the output array in a **binary result**:

**positive class** → the element corresponding to the model we are testing,

**negative classes** → the other elements

The discrimination power is between a particular model and the rest of the 7 selected models



Can discriminate **SM only** histograms  
(or **ALP + SM** histograms) from the rest

Which model is predicted by the DNN?

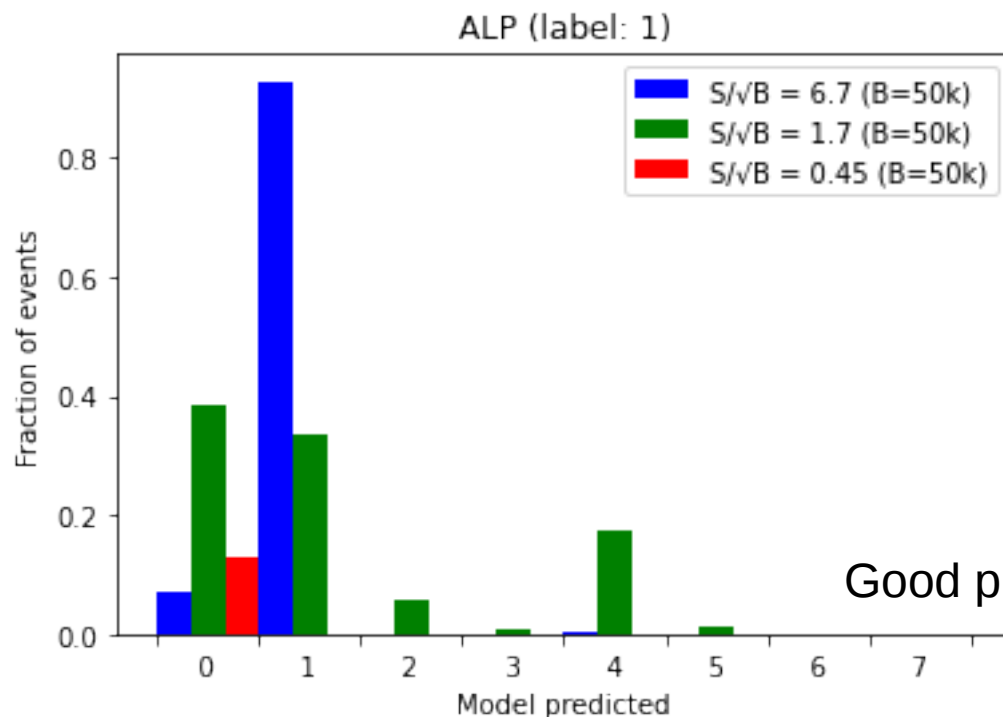
# Multiclass classifier

One count is assigned to the output array element if its probability fulfills two conditions:

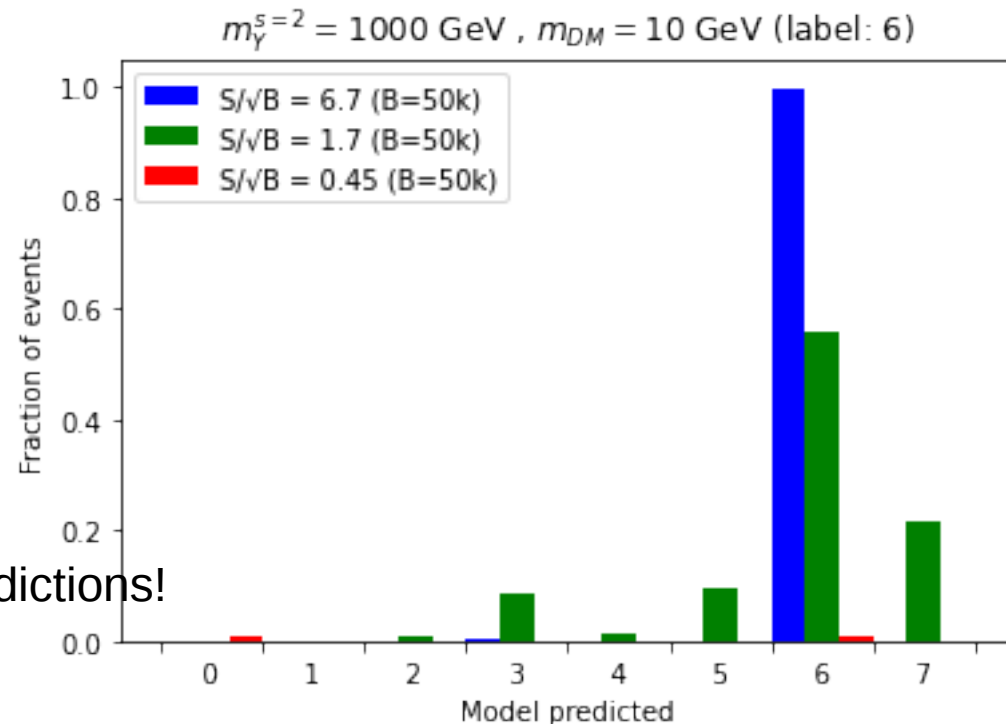
1. it is the **element with the highest value**,
2. its **value is above** a threshold equal to **0.25**, defined as two times the probability that would be obtained for a completely random classifier.  
(there are 8 training models  $\rightarrow 2 * 1/8$ )

Then, a histogram of the frequency of occurrence can be constructed

## Testing with training models



Good predictions!

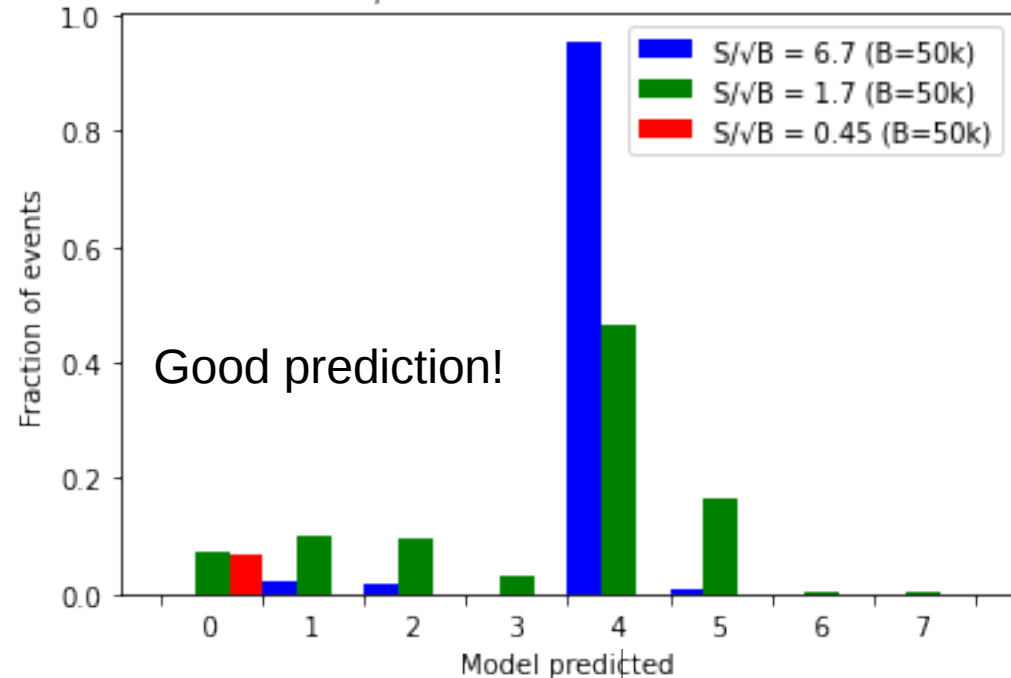


# Multiclass classifier

## Testing with non-training models

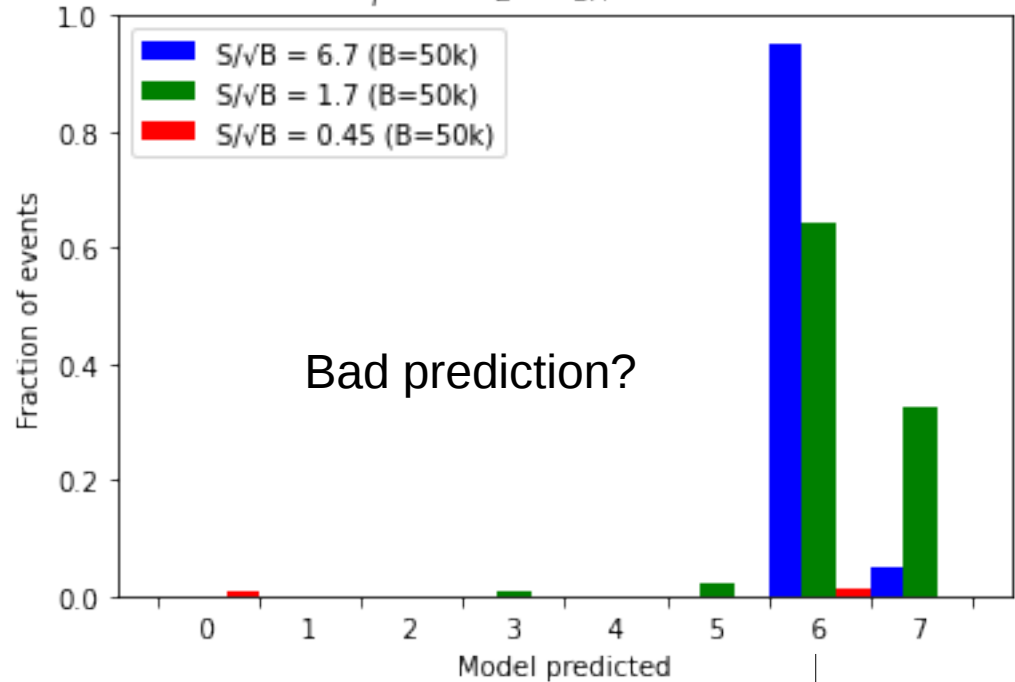
Models completely new to the DNN

$m_Y^{s=1} = m_Z, m_{DM} = 200 \text{ GeV}$



$m_Y^{s=1} = m_Z, m_{DM} = 300 \text{ GeV}$

$m_Y^{s=2} = m_Z, m_{DM} = 300 \text{ GeV}$



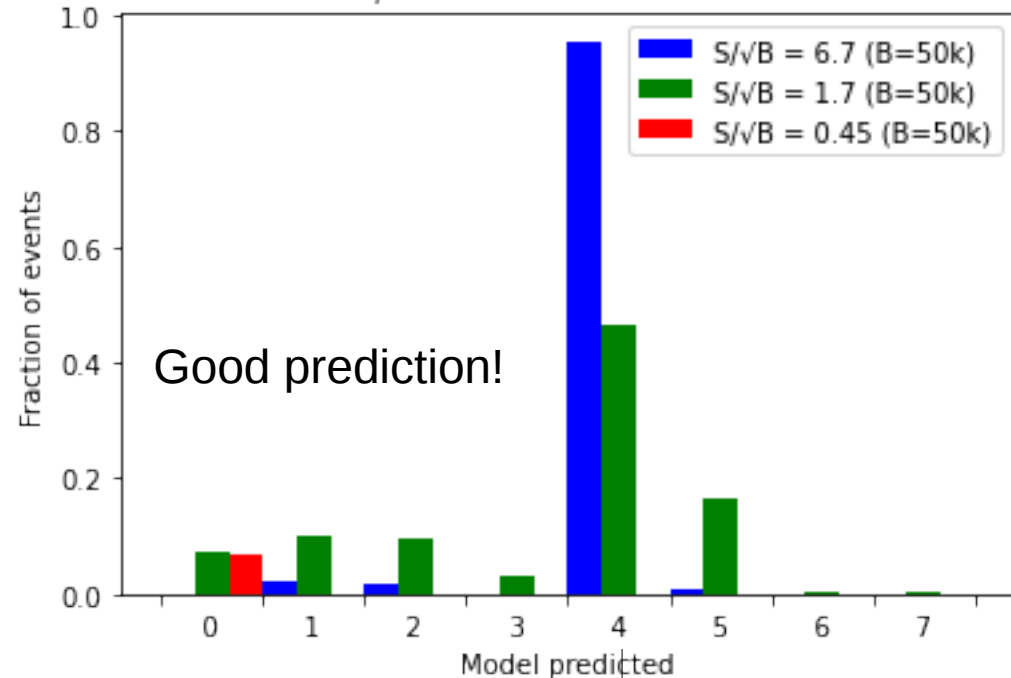
$m_Y^{s=2} = 1000 \text{ GeV}, m_{DM} = 10 \text{ GeV}$

# Multiclass classifier

## Testing with non-training models

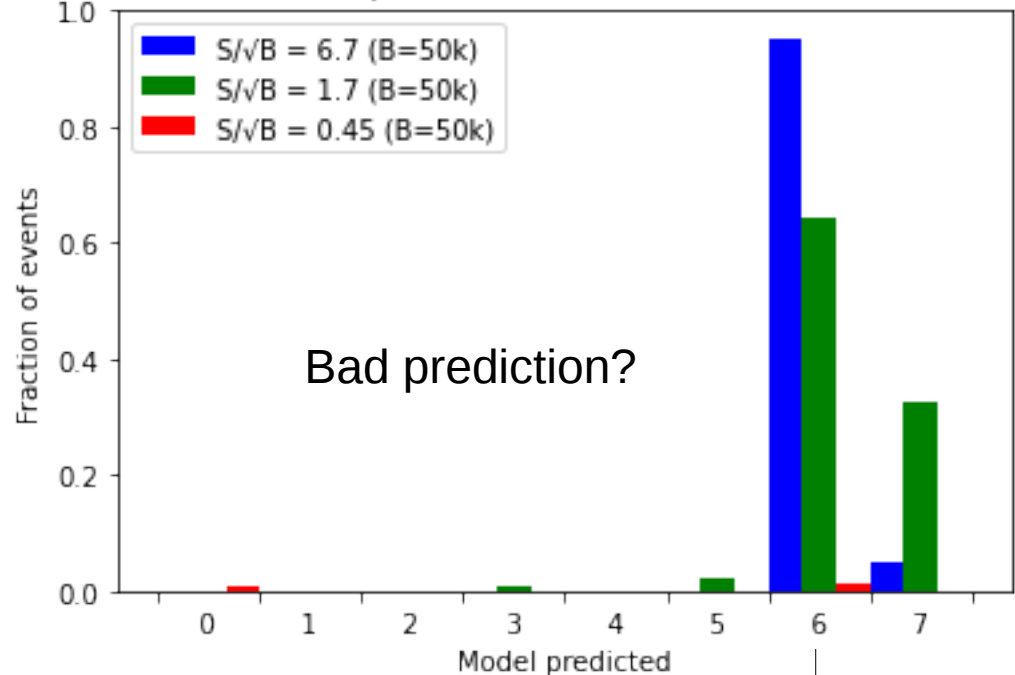
Models completely new to the DNN

$m_Y^{s=1} = m_Z, m_{DM} = 200 \text{ GeV}$



$m_Y^{s=1} = m_Z, m_{DM} = 300 \text{ GeV}$

$m_Y^{s=2} = m_Z, m_{DM} = 300 \text{ GeV}$



$m_Y^{s=2} = 1000 \text{ GeV}, m_{DM} = 10 \text{ GeV}$

The DNN classifies “kinematic distributions” not “models”

Predicts compatible kinetic distribution of the underlying model.

# Conclusions



# Conclusions

Search for dark matter signatures  
at the LHC using deep learning

## Models:

- ♦ Monojet plus missing transverse energy channel of four simplified dark matter frameworks:  
**ALP and spin-0, spin-1, and spin-2 mediator models**
- ♦ One usual drawback of supervised techniques: the need of a specific data set per model  
We take **kinematic features as input data**
  - *Independent of the coupling values*
  - In models with a mediator, *considering its status is key to reduce the free parameters*  
→ **we describe a family of models with a single data set**

## Neural Networks (individual classifiers):

- ♦ Discerning new physics signatures from SM background, two data representations:
  - event-by-event data
  - **2D histograms**
  - poor performance
  - **huge performance boost**
- ♦ CNN no significant improvement w.r.t. DNN results
- ♦ The **method is quite robust**.
  - The DNN can handle small errors in the coupling values,
  - and test an incorrect model as long as their kinematic distributions are similar,

# Conclusions

Search for dark matter signatures  
at the LHC using deep learning

- ◆ Histogram approach drawback: construct histograms for every S/B ratio
  - DNN performance independent of the number of background event if we choose  $S/\sqrt{B}$  as variable

**Easy to check** if a DNN could discriminate a particular model from the SM, for any luminosity.  
Or to **estimate the luminosity needed** to achieve a certain performance level.

## Multimodel classifiers:

- ◆ Supervised algorithms but instead of training one benchmark model per DNN, several ones are used.
  - a **multimodel binary classifier**, prepared to discriminate between:  
SM only histograms vs SM background plus any kind of new physics events
  - a **multiclass classifier** prepared to identify between:  
SM only and several benchmark models, pointing out the most likely underlying model
- ◆ A more challenging task, but a good performance is achieved.
  - misleading model properties can be predicted (e.g. incorrect dark sector masses or spin).
  - But result points towards a compatible kinetic distribution, a key tool to guide further analysis

Thank you!



Back-ups

# NNs

parameters

	DNN		CNN
Input data			
Format	event-by-event	2D histogram	
Features	$(p_T^j, \eta^j, \phi^j)$	$p_T^j$ vs $\eta^j$	
Sample size	400k per model	20k histograms per model	
Train:Test:Validation	0.64:0.20:0.16		
Neural Network			
Layers	3 dense layers Hidden layers nodes = 20 Dropout in every hidden layer: 0.2	Convolutional 2D layer (3×3, 32) Max pooling 2D layer (2×2) Convolutional 2D layer (3×3, 64) Max pooling 2D layer (2×2) Flatten layer Dense layer	
Activation function	Hidden layers: <i>relu</i> Output layer: <i>sigmoid</i>		
Compilation			
Loss function	<i>binary_crossentropy</i>		
Optimizer	<i>adam</i> (initial learning rate = 0.0001)		
Metric	<i>accuracy</i>		
Batch size	128		
Max epochs	1500-2500		
Patience	100-300 epochs		

# SM background

Process	$\sigma$ (pb)	D	PRC
$Zj (Z \rightarrow \nu\bar{\nu})$	53.6	0.78	1
$Wj (W \rightarrow \tau\nu)$	24.8	0.20	0.12
$Wj (W \rightarrow l\nu), l = e, \mu$	49.7	0.049	0.058
$Zj (Z \rightarrow ll), l = e, \mu, \tau$	19.1	0.013	0.006
$t\bar{t}$	217	0.021	0.11
		0.014 ( $b_{tag} \leq 1$ )	0.073
		0.004 ( $b_{tag} = 0$ )	0.021
diboson (WW, WZ, ZZ)	5.18	0.12	0.014

Same  
diagrams  
replacing  
 $Z \rightarrow W$   
 $\nu \rightarrow l$

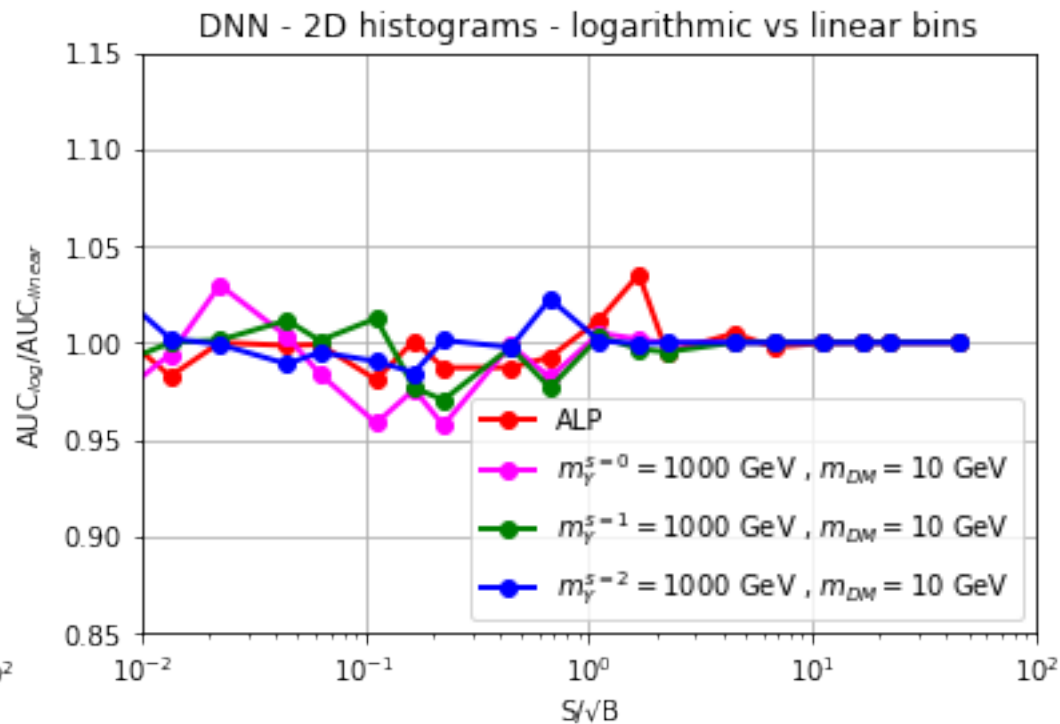
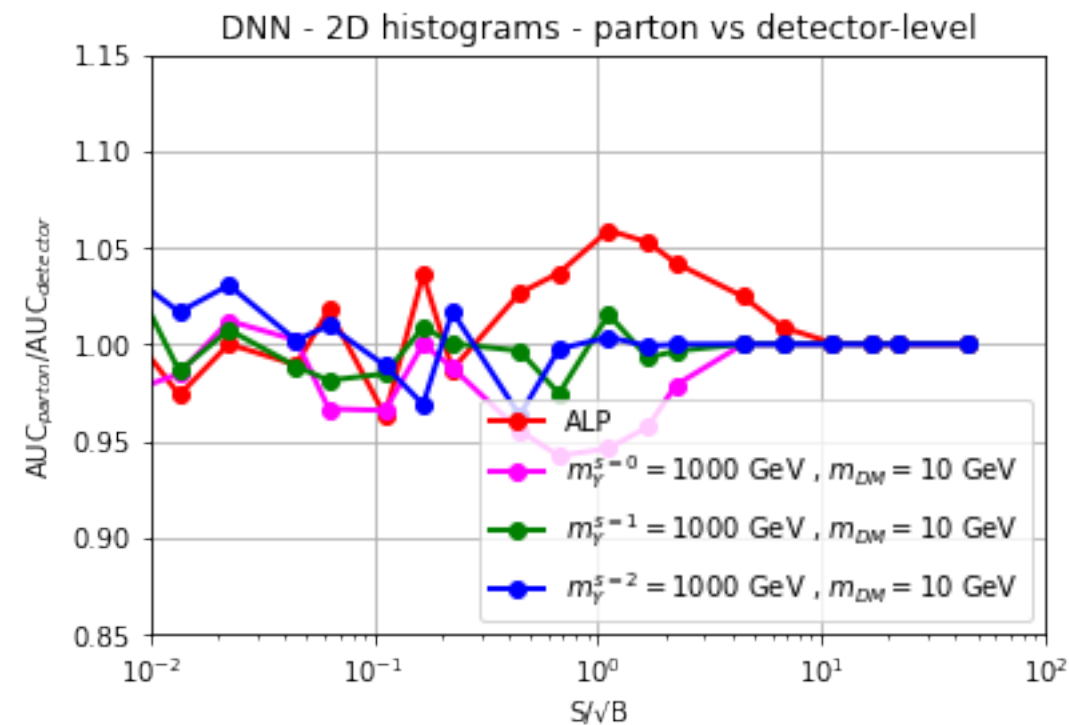
$$PRC = \frac{N_{\text{process}}}{N_{pp \rightarrow Zj(Z \rightarrow \nu\bar{\nu})}},$$

$$N_{\text{process}} = L \sigma_{\text{process}} D_{\text{process}},$$

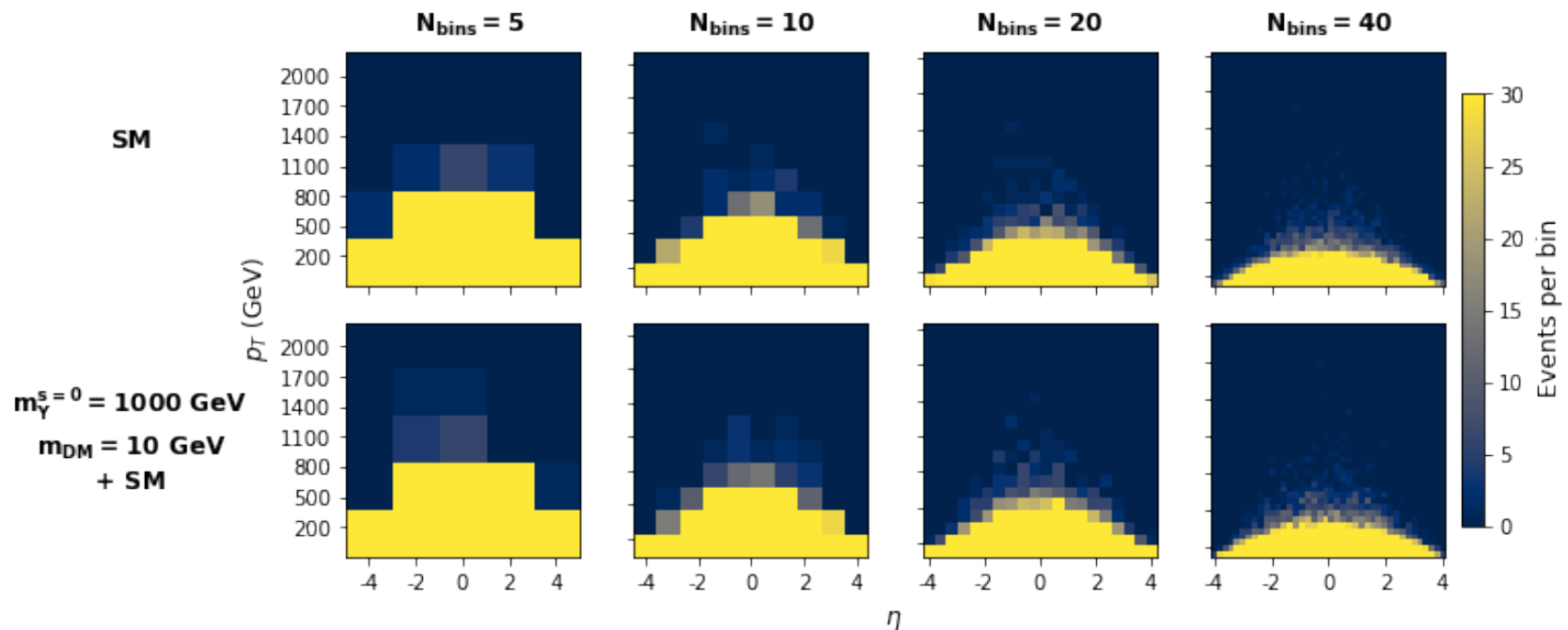
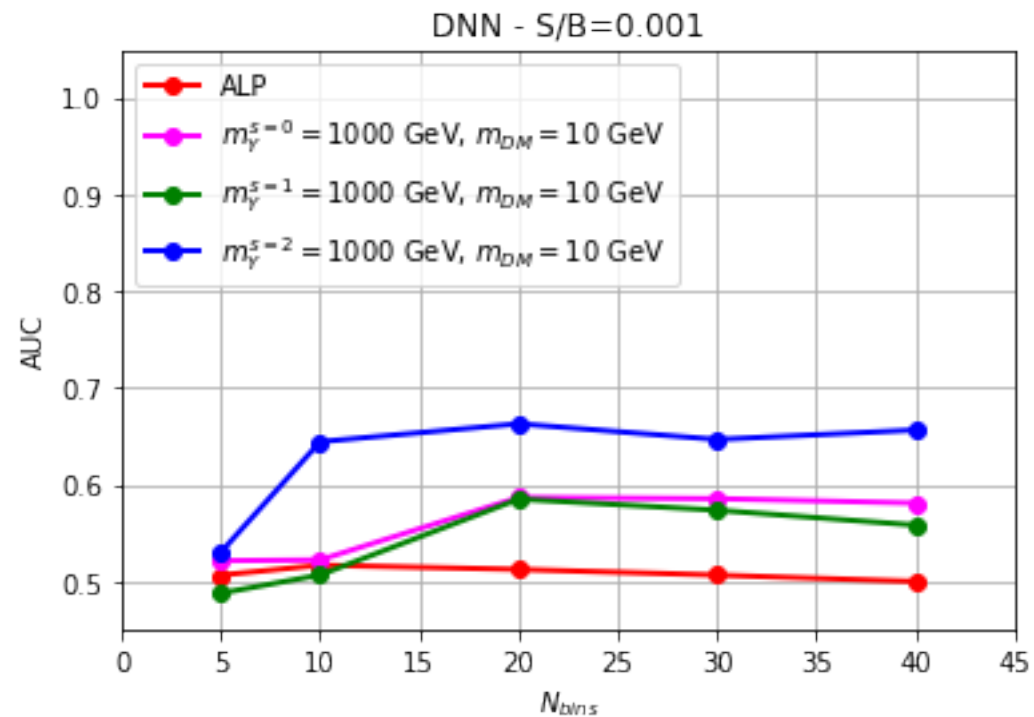
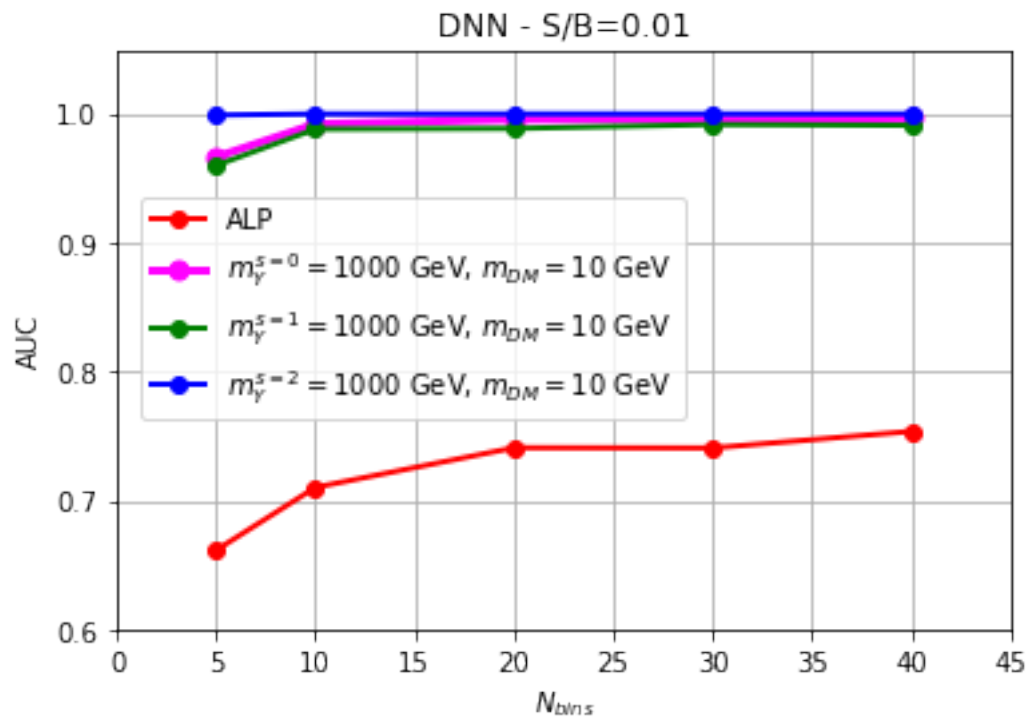
number of events      luminosity      cross section      fraction of events identified as jets + missing  $E_T$  at detector level

# Parton vs detector-level data

# Logarithmic vs linear bins



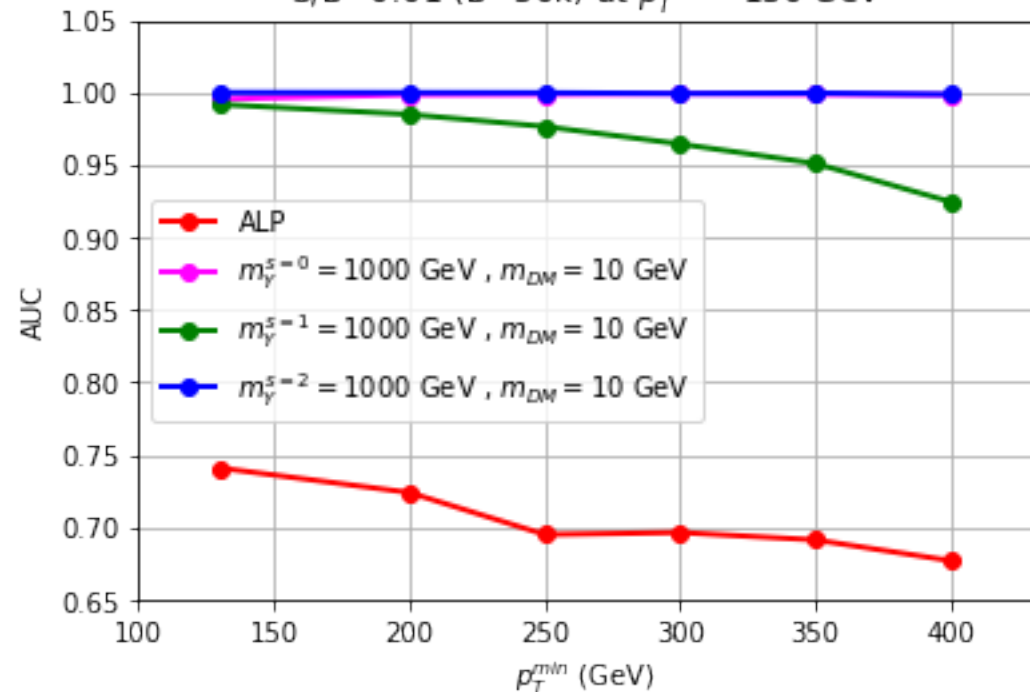
# Bin number per histogram



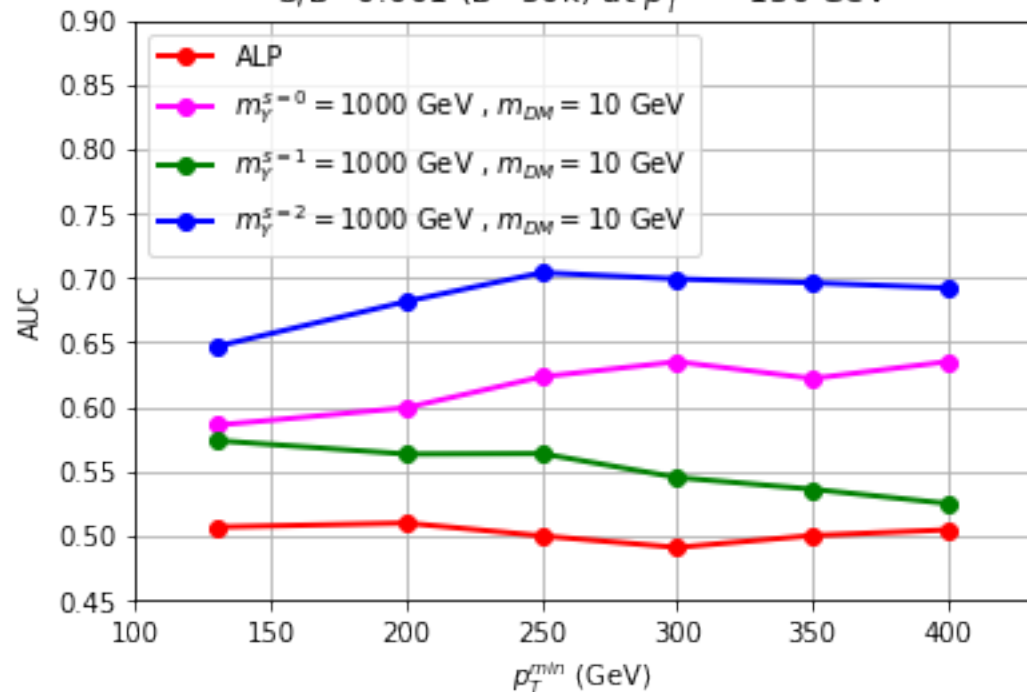


# Minimum jet transverse momentum ( $p_T^j$ cut)

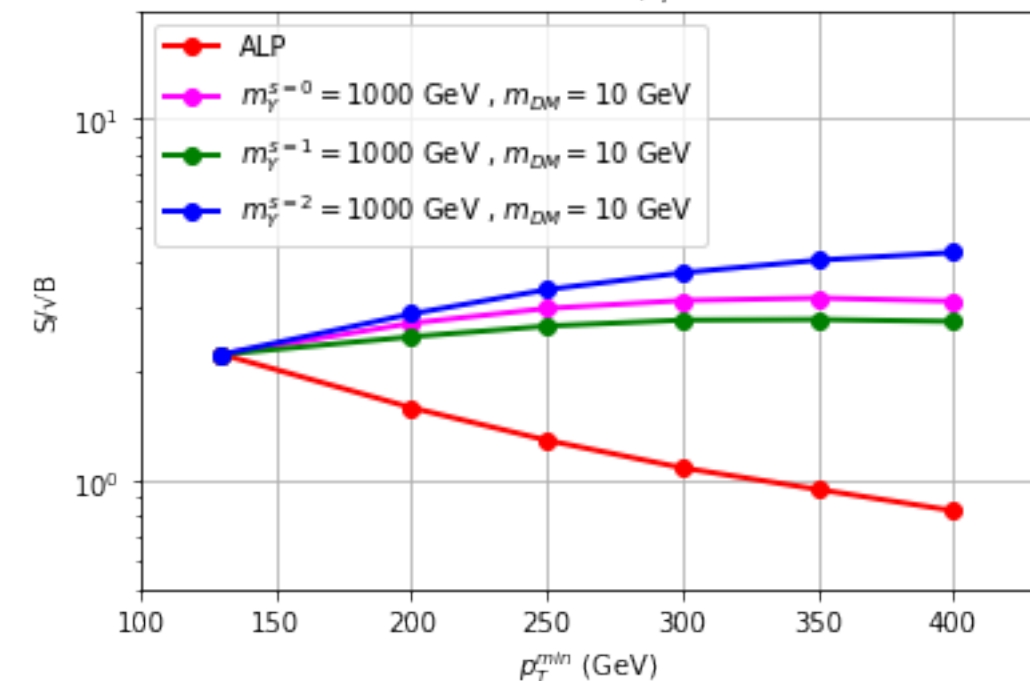
$S/B=0.01$  ( $B=50k$ ) at  $p_T^{min} = 130$  GeV



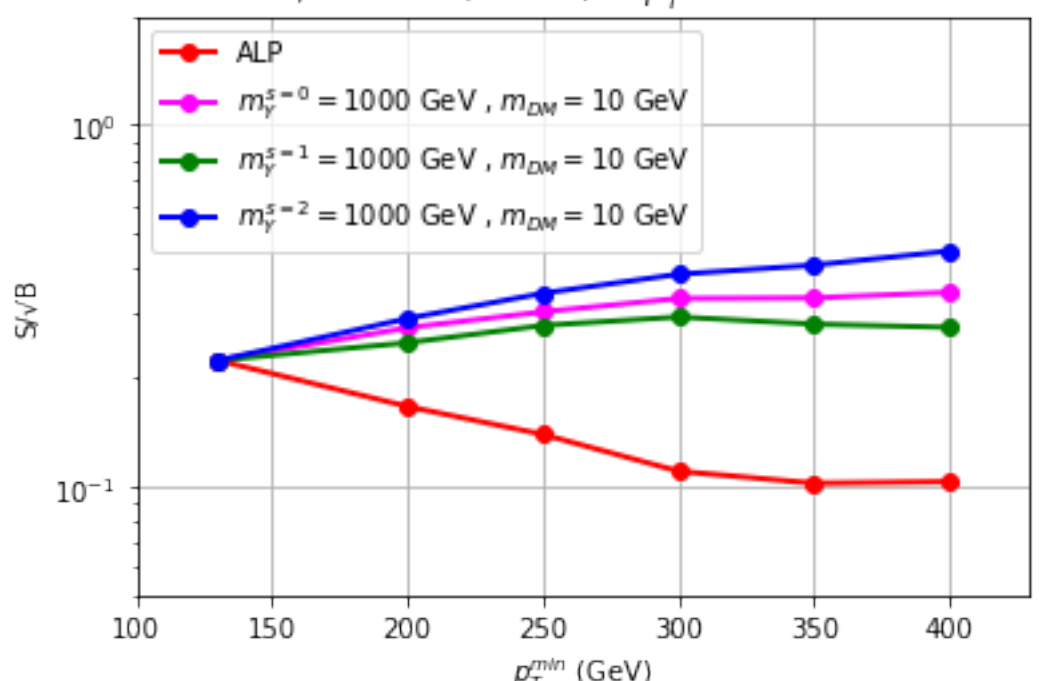
$S/B=0.001$  ( $B=50k$ ) at  $p_T^{min} = 130$  GeV



$S/B=0.01$  ( $B=50k$ ) at  $p_T^{min} = 130$  GeV



$S/B=0.001$  ( $B=50k$ ) at  $p_T^{min} = 130$  GeV



# Discerning between DM models

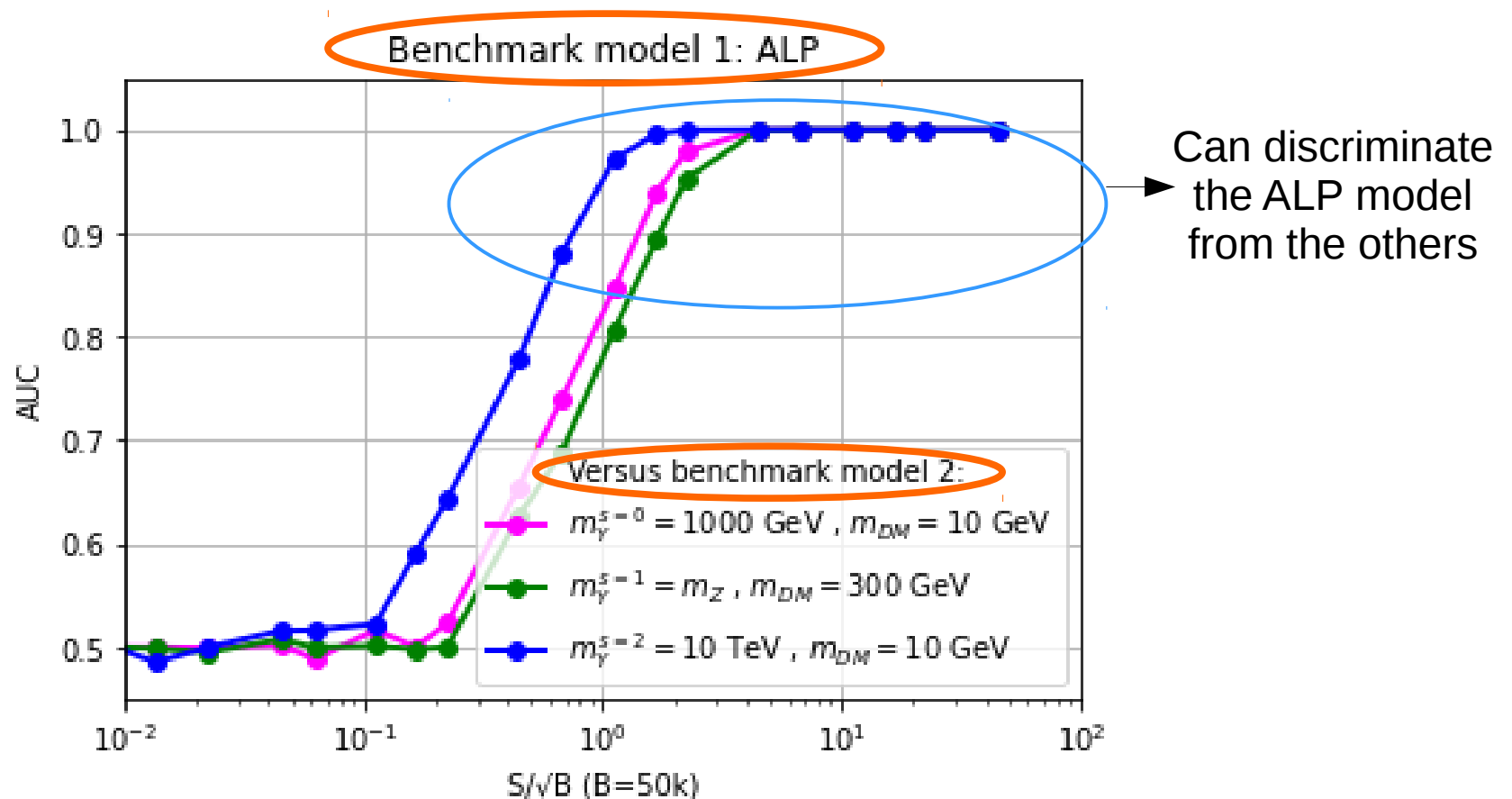
| With individual DNNs

# Discerning between DM models

If signal indicates the presence of non-SM processes, we would like to identify the underlying new physics model.

DNN performance to distinguish between histograms with:

**Benchmark model 1 plus SM** vs **Benchmark model 2 plus SM**

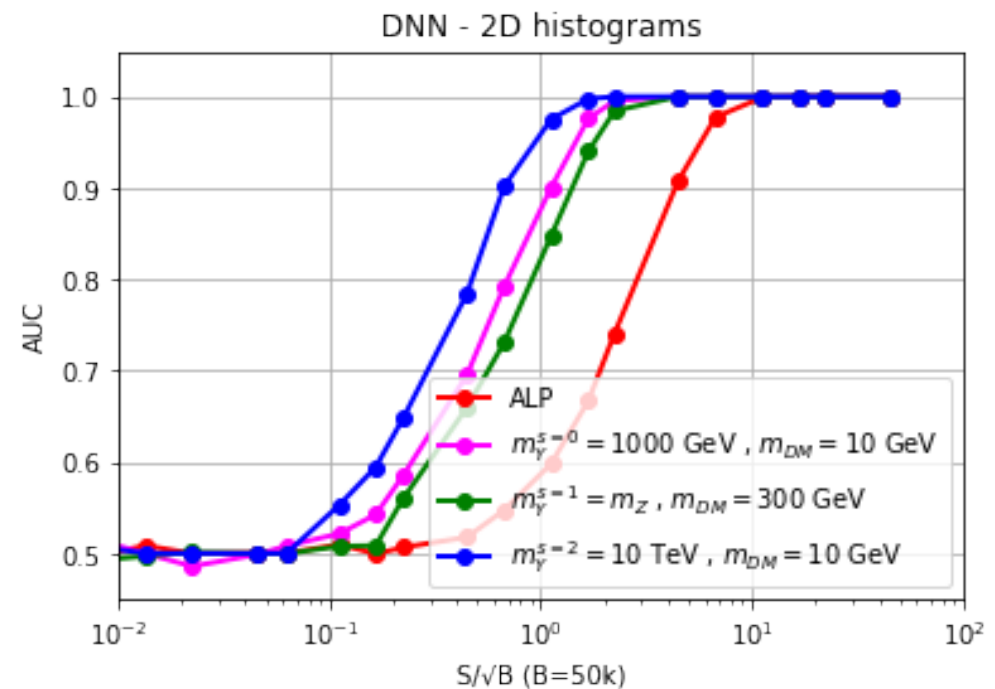
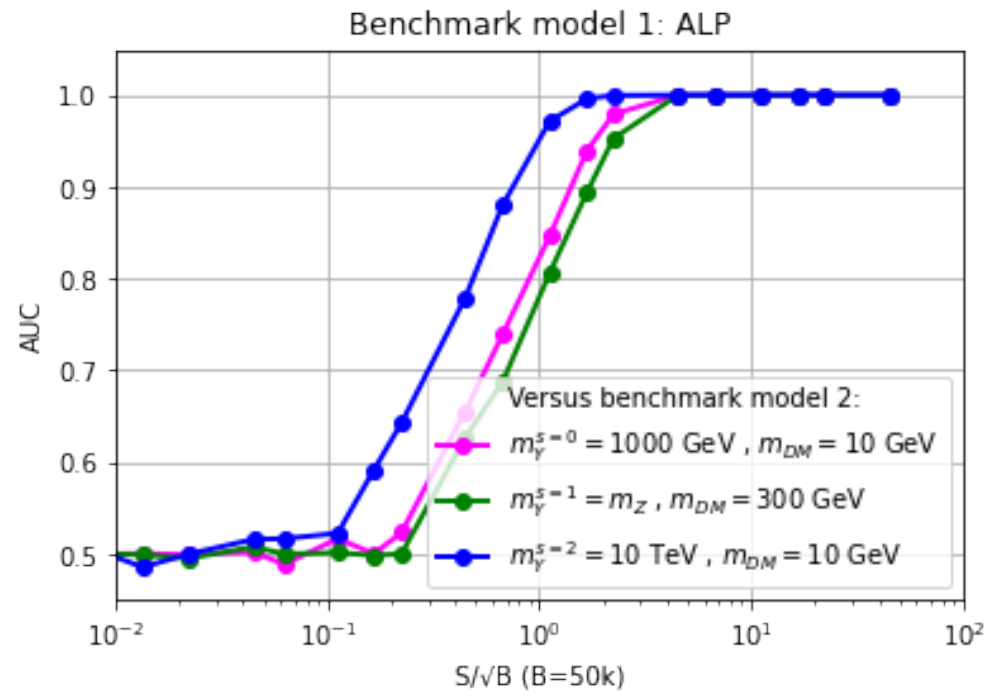


# Discerning between DM models

**ALP plus SM**  
VS  
**the other benchmark models plus SM**

The ALP model takes the role of the SM, because both kinematic distributions are similar.

**SM only**  
VS  
**Benchmark model plus SM**

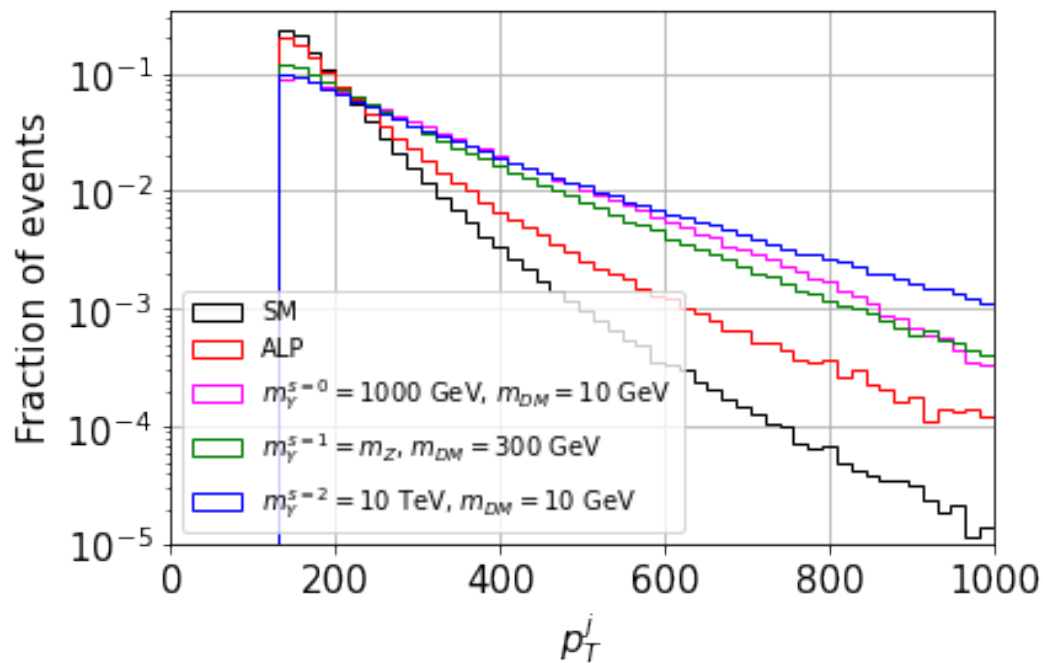
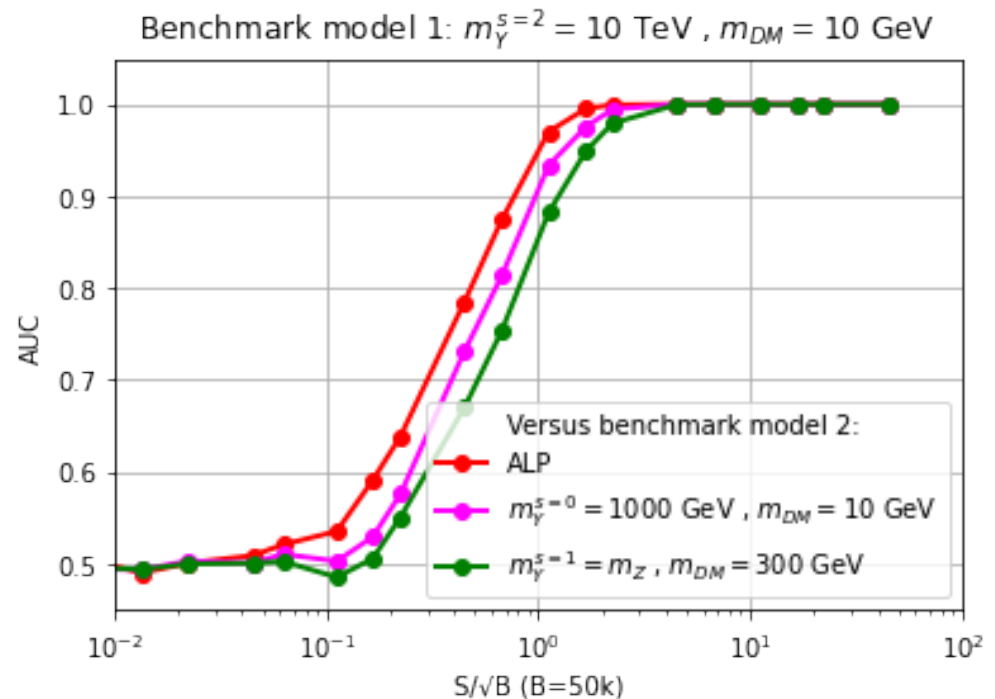


# Discerning between DM models

**Spin-2 mediator plus SM**  
VS  
**the other benchmark models plus SM**

The **spin-2 model** has the hardest spectrum

→ highest efficiency achieve when we try to disentangle between the **spin-2 model** and the model with the softer spectrum, **ALP**.



# Monojet and dijet + MET processes

Monojet  $\rightarrow$  3 kinematic variables:  $p_T^j$ ,  $\eta^j$ ,  $\Phi^j$

Dijet  $\rightarrow$  8 kinematic variables:

$p_T^{j1}$ ,  $\eta^{j1}$ , (j1: leading jet)

$p_T^{j2}$ ,  $\eta^{j2}$ , (j2: sub-leading jet)

$p_T^{\text{MET}}$  missing transverse momentum

$\Delta\Phi^{j1 j2}$ ,  $\Delta\Phi_{\text{MET}}^{j1}$ ,  $\Delta\Phi_{\text{MET}}^{j2}$

