

0.1 Stochastic Differential Calculus

(Lesson 8 of
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We want now to generalize the ordinary calculus rules to the stochastic case.

We obtained a stochastic differential equation from the Master Equation. Then, to understand the underlying physics, we introduced the *Langenvin equation* (in the Overdamped limit):

$$dx(t) = f(x(t), t) dt + \sqrt{2D(x(t), t)} dB(t)$$

with $f = F_{\text{ext}}/\gamma$.

The meaning of $dB(t)$ is clear only passing to *finite differences*, where $dB \rightarrow \Delta B$ is distributed according to:

$$\Delta B \frac{1}{\sqrt{2\pi\Delta t}} \sim \exp\left(-\frac{(\Delta B)^2}{2\Delta t}\right)$$

If we consider a *stochastic integral*:

$$\int g dB$$

we need to discretize it and take the continuum limit (in the mean square convergence).

In the more general case, we want to integrate wrt $(dB)^n$. We will now show that:

$$(dB)^n = \begin{cases} dB & n = 1 \\ dt & n = 2 \\ 0 & n > 2 \end{cases}$$

Example 1 (Integral in dB^2):

Consider a *non-anticipating* function $G(\tau)$, and the following integral:

$$I = \int G(\tau)(dB(\tau))^2$$

With *non-anticipating* we mean that $G(\tau)$ does not depend on $B(s) - B(\tau) \forall s > \tau$, i.e. it does not depend on the *future*. Discretizing:

$$I = \text{ms} - \lim \sum_{i=1}^n G(t_{i-1}) \Delta B_i^2$$

For simplicity, we will denote $G_{i-1} \equiv G(t_{i-1})$, $\Delta B_i \equiv B_i - B_{i-1} \equiv B(t_i) - B(t_{i-1})$ and $t_i - t_{i-1} = \Delta t_i$.

We want to prove that:

$$\int_0^t G(\tau)(dB(\tau))^2 \stackrel{?}{=} \int_0^t G(\tau) d\tau = \lim_{n \rightarrow \infty} \sum_i G(t_{i-1}) \Delta t_i$$

So we take their average difference squared and compute the limit:

$$\left\langle \left(\sum_{i=1}^n G_{i-1} \Delta B_i^2 - \sum_{i=1}^n G_{i-1} \Delta t_i \right)^2 \right\rangle \xrightarrow[n \rightarrow \infty]{?} 0$$

Then, proceeding similarly to last lecture:

$$\begin{aligned} &= \left\langle \left[\sum_{i=1}^n G_{i-1} ((\Delta B_i)^2 - \Delta t_i) \right]^2 \right\rangle = \left\langle \sum_{i=1}^n G_{i-1} ((\Delta B_i)^2 - \Delta t_i) \sum_{j=1}^n G_{j-1} ((\Delta B_j)^2 - \Delta t_j) \right\rangle = \\ &= \sum_{i=1}^n \langle G_{i-1}^2 ((\Delta B_i)^2 - \Delta t_i)^2 \rangle + 2 \sum_{i < j} \langle G_{i-1} ((\Delta B_i)^2 - \Delta t_i) G_{j-1} ((\Delta B_j)^2 - \Delta t_j) \rangle \end{aligned}$$

Note that the yellow term *does not depend* on $\Delta B_j = B_j - B_{j-1} = B(t_j) - B(t_{j-1})$ (recall that G is *non-anticipating*). Thus, the yellow and blue terms are *independent* of each other, and so we can factorize the average:

$$= \sum_{i=1}^n \langle G_{i-1}^2 ((\Delta B_i)^2 - \Delta t_i)^2 \rangle + 2 \sum_{i < j} \langle G_{i-1} ((\Delta B_i)^2 - \Delta t_i) G_{j-1} \rangle \langle (\Delta B_j)^2 - \Delta t_j \rangle$$

However, recall that:

$$\langle (\Delta B_j)^2 - \Delta t_j \rangle = \langle (\Delta B_j)^2 \rangle - \Delta t_j = 0$$

and so only the first term remains. Again, we can separate two independent terms:

$$= \left\langle \sum_{i=1}^n G_{i-1}^2 ((\Delta B_i)^2 - \Delta t_i)^2 \right\rangle = \sum_{i=1}^n \langle G_{i-1}^2 \rangle \langle ((\Delta B_i)^2 - \Delta t_i)^2 \rangle$$

Then, focus on a single term:

$$\begin{aligned} \langle ((\Delta B_i)^2 - \Delta t_i)^2 \rangle &= \langle \Delta B_i^4 - 2\Delta t_i (\Delta B_i)^2 \rangle + \Delta t_i^2 = \\ &= \underbrace{\langle (\Delta B_i)^4 \rangle}_{3(\Delta t_i)^2} - 2\Delta t_i \underbrace{\langle (\Delta B_i)^2 \rangle}_{\Delta t_i} + \Delta t_i^2 = 2\Delta t_i^2 \end{aligned}$$

Substituting back into the sum:

$$\begin{aligned} &= \left\langle \left[\sum_{i=1}^n G_{i-1} ((\Delta B_i)^2 - \Delta t_i) \right]^2 \right\rangle = 2 \sum_{i=1}^n G_{i-1}^2 \Delta t_i^2 \\ &\leq 2 \left(\max_{i \leq j \leq n} \Delta t_j \right) \sum_{i=1}^n G_{i-1}^2 \Delta t_i \xrightarrow[n \rightarrow \infty]{} 2 \cdot 0 \cdot \int_0^t G^2 d\tau = 0 \end{aligned}$$

This proves that $(dB)^2 = dt$.

Example 2 (The case with $n > 2$):

We want now to show that:

$$\int_0^t G(\tau) (dB(\tau))^n = \text{ms} - \lim \sum_{i=1}^n G(t_{i-1}) \Delta B_i^n = 0$$

Again, we discretize and consider the average distance squared:

$$\left\langle \left(\sum_{i=1}^n G_{i-1} (\Delta B_i)^n \right)^2 \right\rangle = \sum_{i=1}^n G_{i-1}^2 (\Delta B_i)^{2n} + 2 \sum_{i < j} \langle G_{i-1} G_{j-1} (\Delta B_i)^n (\Delta B_j)^n \rangle$$

Suppose, for simplicity, that G is bounded, i.e. $|G| < K$. Then, as G is non-anticipating, we can factorize the averages (as we did before). If n is **odd**, the second term vanishes:

$$= \sum_i \langle G_{i-1}^2 \rangle (\Delta t_i)^n \frac{(2n)!}{2^n n!} \leq \frac{K^2 (2n)!}{2^n n!} \sum_{i=1}^n (\Delta t_i)^n \leq \frac{K^2 (2n)!}{2^n n!} \max_{i \leq j \leq n} (\Delta t)^{n-1} \underbrace{\sum_{i=1}^n \Delta t_i}_t \xrightarrow{n \rightarrow \infty} 0$$

Otherwise, if n is even, the following holds instead:

$$\leq 2K^2 \left(\frac{n!}{2^{n/2} (n/2)!} \right)^2 \sum_{i < j} \Delta t_i^{n/2} \Delta t_j^{n/2} \leq 2K^2 \left(\frac{n!}{2^{n/2} (n/2)!} \right)^2 \left(\max_{i \leq l \leq n} \Delta t_l \right)^{2(n/2-1)} \sum_{i < j} \underbrace{\Delta t_i \Delta t_j}_{\leq t^2}$$

and by taking the limit $n \rightarrow \infty$ it goes to 0, proving the thesis.

Example 3 (Other cases):

Consider now:

$$\int G(\tau) dB(\tau) d\tau = 0$$

In fact, as $(dB)^2 = d\tau$, $dB du = 0$ because $(dB)^n = 0 \forall n > 2$.

Consider now:

$$\begin{aligned} dB(t)^n &= (B(t) + dB(t))^n - (B(t))^n = \\ &= \sum_{k=1}^n \binom{n}{k} (dB(t))^k B(t)^{n-k} \end{aligned}$$

We consider the non-trivial case, i.e when $n \geq 2$:

$$= n dB(t) B^{n-1}(t) + \frac{n(n-1)}{2} \underbrace{(dB(t))^2}_{dt} B^{n-2}(t) + 0$$

If we set $n-1 = m$ we arrive at:

$$(m+1)B^m(t) dB(t) = dB(t)^{m+1} - \frac{m(m+1)}{2} B^{m-1}(t) dt$$

Then:

$$\int_0^\tau B^m(t) dB(t) = \frac{1}{m+1} \int_0^\tau d(B(t))^{m+1} - \frac{m}{2} \int_0^\tau B^{m-1}(t) dt$$

Note that, if $m = 1$:

$$\int_0^\tau B(\tau) dB(t) = \frac{1}{2} \int_0^\tau d(B(t))^2 - \frac{1}{2} \int_0^\tau dt$$

and so we arrive again at a formula we already seen:

$$\int_0^\tau B(t) dt = \frac{1}{2}(B^2(\tau) - B^2(0)) - \frac{\tau}{2}$$

In the general case, for $m > 0$:

$$\begin{aligned} \int_0^\tau B^m(t) dB(t) &= \frac{1}{m+1} \int_0^\tau d(B(t))^{m+1} - \frac{m}{2} \int_0^\tau B^{m-1}(t) dt = \\ &= \frac{B^{m+1}(\tau) - B^{m+1}(0)}{m+1} - \frac{m}{2} \int_0^\tau B^{m-1}(t) dt \end{aligned}$$

Note that we can generalize this to a generic function f of $B(t)$:

$$\begin{aligned} df(B(t)) &= f(B(t) + dB(t)) - f(B(t)) = \\ &= f'(B(t)) dB(t) + \frac{1}{2} f''(B(t)) (dB(t))^2 + 0 \end{aligned}$$

and so:

$$df = f' dB + \frac{1}{2} f'' dt + O(dt^{3/2})$$

0.2 Derivation of the Fokker-Planck equation

Starting from the Master Equation and taking the continuum limit we arrived at the Fokker-Planck equation:

$$\dot{\mathbb{P}}(x, t) = -\frac{\partial}{\partial x} \left[f(x, t) \mathbb{P}(x, t) - \frac{\partial}{\partial x} \mathbb{P}(x, t) D(x, t) \right]$$

($\mathbb{P}(x, t) \equiv W(x, t)$).

We want to derive from that the Langenvin equation:

$$dx(t) = f(x(t), t) + \sqrt{2D(x(t), t)} dB(t)$$

Note that in the Langenvin eq. $x(t)$ appears because we are talking about a *single trajectory*, while in $\mathbb{P}(x, t)$ x and t are independent variables, and $dx \mathbb{P}(x, t)$ is the *density* of trajectories passing in $[x, x + dx]$ at time t . So, to compute $\mathbb{P}(x, t)$, we need to *generate many trajectories* with the Langenvin equation, and count the

ones satisfying the appropriate conditions (crossing $[x, x + dx]$ at time t). We can do this by writing:

$$\mathbb{P}(x, t) = \langle \delta(x(t) - x) \rangle$$

In fact, if we integrate in $[x, x + \Delta x]$:

$$\int_x^{x+\Delta x} \mathbb{P}(x', t) dx' = \langle \int_x^{x+\Delta x} \delta(x(t) - x') dx' \rangle$$

Now, all trajectories that pass through $[x, x + \Delta x]$ at time t (i.e. such that $x(t) \in (x, x + \Delta x)$) *contribute* to that integral, and the others do not. So, if we take the average over the set of *all trajectories*, we find the *density* of trajectories passing through that *gate*, which is exactly $\mathbb{P}(x, t)$.

Consider now a generic function h . Its average over the trajectory is defined as:

$$\langle h(x(t)) \rangle = \int dx \mathbb{P}(x, t) h(x)$$

Differentiating:

$$d\langle h(x(t)) \rangle = \langle h(x(t + dt)) - h(x(t)) \rangle = dt \int dx \dot{\mathbb{P}}(x, t) h(x)$$

We have now all the results we need to derive the Fokker-Planck equation. Start with:

$$\begin{aligned} dh &= h(x(t) + dx(t)) - h(x(t)) = \\ &= h'(x(t)) dx(t) + \frac{1}{2} h''(x(t)) (dx(t))^2 + O((dx)^3) \end{aligned}$$

Using Langenvin:

$$\begin{aligned} (dx(t))^2 &= dt^2 f^2 + 2D \overbrace{(dB)^2}^{dt} + \cancel{f\sqrt{2D} dB dt} \\ (dx(t))^3 &= O((dt)^2) \end{aligned}$$

Substituting back:

$$\begin{aligned} dh &= h'(f dt + \sqrt{2D} dB) + \frac{1}{2} h'' 2D dt + O(dt^2) = \\ &= dt [h' f + h'' D] + h' \sqrt{2D} dB \end{aligned}$$

$$\langle dt (h' f + h'' D) \rangle + \langle h' \sqrt{2D} dB \rangle$$

Focus on the last term. Factorizing the average (ad D is non-anticipating...)

$$2\langle h'(x(t)) D(x(t), t) dB(t) \rangle = 2\langle h' D \rangle \underbrace{\langle dB \rangle}_{=0}$$

Thus:

$$\begin{aligned}\frac{d}{dt}\langle h(x(t)) \rangle &= \langle h'(x(t))f(x(t), t) + h''(x(t))D(x(t), t) \rangle = \\ &= \int dx \mathbb{P}(x, t) [h'(x)f(x, t) + h''(x)D(x, t)]\end{aligned}$$

Now:

$$\frac{d}{dt}\langle h(x(t)) \rangle = \int dx \dot{\mathbb{P}}(x, t)h(x)$$

We want to compare the two different expressions to get an expression for $\dot{\mathbb{P}}(x, t)$.
Integrating by parts:

$$\int dx \dot{\mathbb{P}}(x, t)h(x, t) = \left[\mathbb{P}h \right]_{-\infty}^{+\infty} - \int dx h(x) \frac{\partial}{\partial x} (\mathbb{P}(x, t)f(x, t)) + \left[D\mathbb{P}h' \right]_{-\infty}^{+\infty} - \int dx h' \frac{\partial}{\partial x} (\mathbb{P}D) dx$$

Integrating by parts again:

$$- \int dx h' \frac{\partial}{\partial x} (\mathbb{P}D) dx = \left[-h \frac{\partial}{\partial x} (\mathbb{P}D) \right]_{-\infty}^{+\infty} + \int dx h \frac{\partial^2}{\partial x^2} (\mathbb{P}D) dx$$

$h(x, t)$ can be chosen arbitrarily, as a *test function*. We can then choose h to have a narrow peak centered on a certain x , so that all highlighted terms vanish. Then:

$$\begin{aligned}\frac{d}{dt}\langle h(x(t)) \rangle &= \int dx \dot{\mathbb{P}}(x, t)h(x) = \\ &= \int dx h(x) \left[-\frac{\partial}{\partial x} (\mathbb{P}f) + \frac{\partial^2}{\partial x^2} (\mathbb{P}D) \right]\end{aligned}$$

This proves that the Langenvin equation is *equivalent* to the Fokker-Planck equation:

$$\begin{array}{ll}\text{F-P} & \dot{\mathbb{P}}(x, t) = -\frac{\partial}{\partial x} \left[f(x, t)\mathbb{P}(x, t) - \frac{\partial}{\partial x} (\mathbb{P}(x, t)D(x, t)) \right] \\ \text{L} & dx(t) = dt f(x(t), t) + \sqrt{2D(x(t), t)} dB(t)\end{array}$$

0.3 The role of temperature

Recall the definition of $f(x, t)$:

$$f(x, t) = \frac{F_{\text{ext}}}{\gamma}; \quad \gamma = 6\pi a\eta; \quad F_{\text{ext}}(x) = -\frac{\partial V}{\partial x}(x)$$

Assuming D independent of x :

$$\dot{\mathbb{P}}(x, t) = \frac{\partial}{\partial x} \left[\frac{1}{\gamma} \frac{\partial V}{\partial x} \cdot \mathbb{P} + D \frac{\partial \mathbb{P}}{\partial x} \right]$$

We expect that a particle will reach, after some time, an equilibrium described by the Maxwell-Boltzmann distribution:

$$\mathbb{P}(x, t) \xrightarrow{t \rightarrow \infty} P_{\text{eq}}(x) = \frac{e^{-\beta V(x)}}{z} \quad z = \int dx e^{-\beta V(x)}; \quad \beta = \frac{1}{k_B T}$$

At equilibrium, $\dot{\mathbb{P}} = 0$ (stationary solution):

$$\frac{\partial}{\partial x} \left[\frac{1}{\gamma} \mathbb{P}^* + D \frac{\partial}{\partial x} \mathbb{P}^* \right] = 0$$

We would like that $\mathbb{P}^* = P_{\text{eq}}$. We then ask: how many stationary solution are there? If there is only one, does it coincide always with P_{eq} ?

Note that the terms in the parenthesis must be constant:

$$\Rightarrow [\dots] = \text{constant}$$

We assume that \mathbb{P}^* vanishes at infinity (as it is normalized), and also its first derivative:

$$\mathbb{P}^*, \frac{\partial}{\partial x} \mathbb{P}^* \xrightarrow{x \rightarrow \infty} 0$$

Then $[\dots]$ must be 0:

$$\frac{\partial}{\partial x} \mathbb{P}^* = -\frac{1}{\gamma D} \frac{\partial V}{\partial x} \mathbb{P}^*$$

Dividing both sides by \mathbb{P}^* and integrating:

$$\frac{\partial}{\partial x} \ln \mathbb{P}^*(x) = -\frac{1}{\gamma D} \frac{\partial V}{\partial x} \Rightarrow \ln \mathbb{P}^*(x) = -\frac{1}{\gamma D} V(x) + \text{const}$$

Exponentiating:

$$\mathbb{P}^*(x) = \exp\left(-\frac{1}{\gamma D} V(x)\right) \cdot K$$

Comparing to the Maxwell-Boltzmann distribution:

$$\frac{1}{\gamma D} = \beta \Rightarrow \gamma D = k_B T \Rightarrow D = \frac{k_B T}{6\pi\eta a}$$

This is the Einstein relation (*fluctuation-dissipation relationship*), found in 1905. As $D(x, t) \propto T$, the amplitude of stochastic oscillations in the Langevin eq. is proportional to \sqrt{T} .

However, are we sure that this is true *for every initial condition*?