

Binary neutron star mergers of quark matter based equations of state

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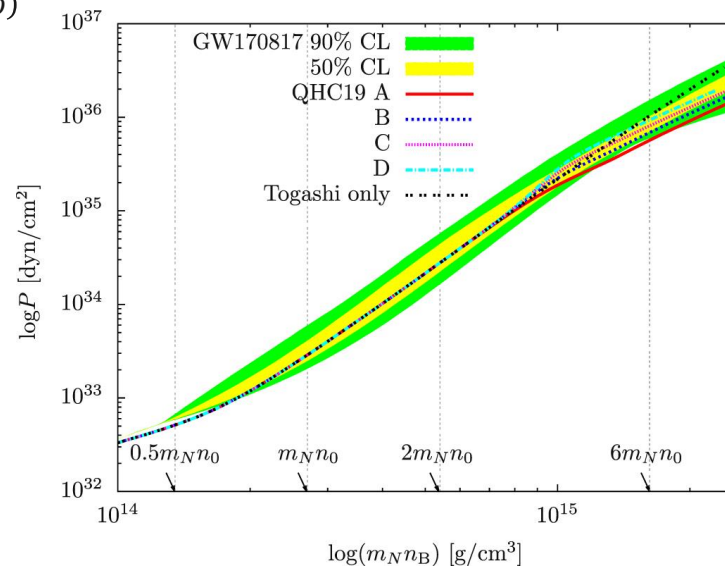
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Dense matter Equation of State

Equation of state (EoS) in the form of Pressure v. density $P(\rho)$



Baym, Furusawa, Hatsuda et al. *ApJ* (2019) 885:42

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Quark-Hadron Crossover (QHC19A-D) EoS

Baym et al. 2019, 2018, Other's by Kurkela et al. *ApJ* (2014) 789:127

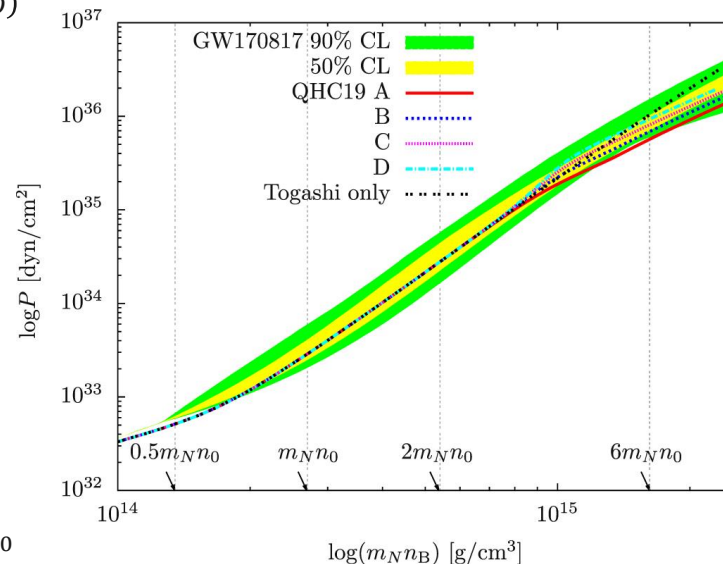
Cold EoS ($T = 0$)

Hadronic EoS: Togashi EoS (extension of APR) for densities
 $< 2n_0 \approx 5.4 \times 10^{14} \text{ gm/cm}^3$

Quark EoS: Nambu-Jona-Lasinio model

$> 5n_0 \approx 1.3 \times 10^{15} \text{ gm/cm}^3$

Smooth fitting in the intermediate Crossover region: $2n_0 - 5n_0$



Baym, Furusawa, Hatsuda et al. *ApJ* (2019) 885:42

Piecewise polytropic fitting scheme

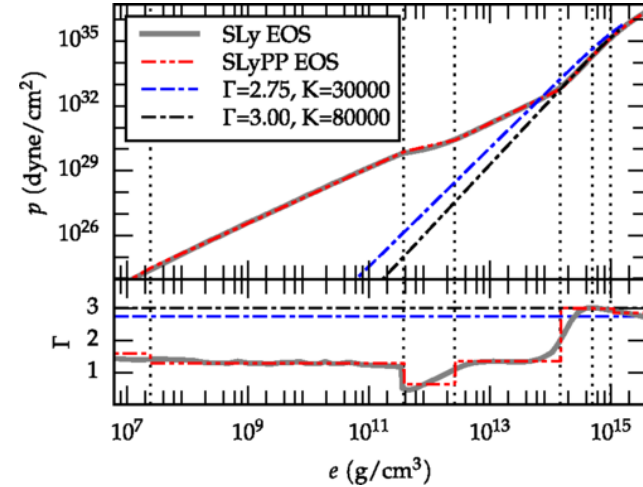
Fitting conditions:

- Pressure-density relation: $P(\rho) = K_i \rho^{\Gamma_i}$
- Internal energy density $\epsilon = a_i \rho + \frac{K_i}{\Gamma_i - 1} \rho^{\Gamma_i}$
- Continuity of pressure, internal energy is imposed.

Read, Lackey, Owen, Friedman *PRD* (2009) 79, 124032.

3 piece piecewise scheme (high density) capture all essential structure of the EoS, + SLy EoS at low densities.

Non-zero temperature corrections by $\Gamma_{thermal}$ contribution.



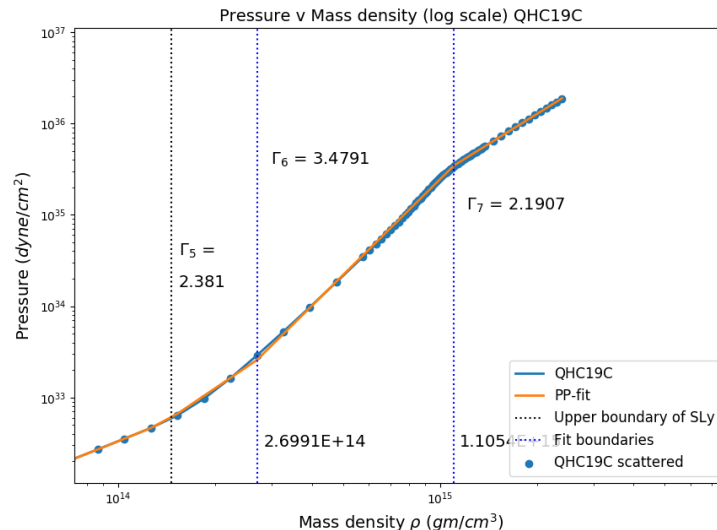
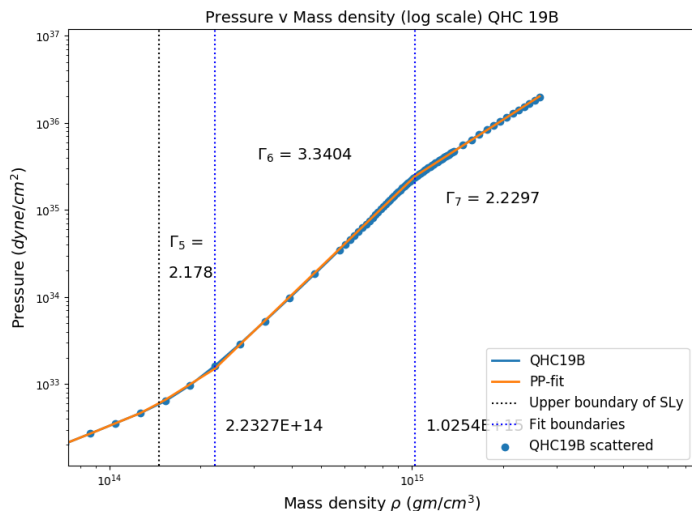
Pietri, Feo, Maione, Löffler *PRD* (2016) 93, 064047

Parameterized QHC EoS piecewise

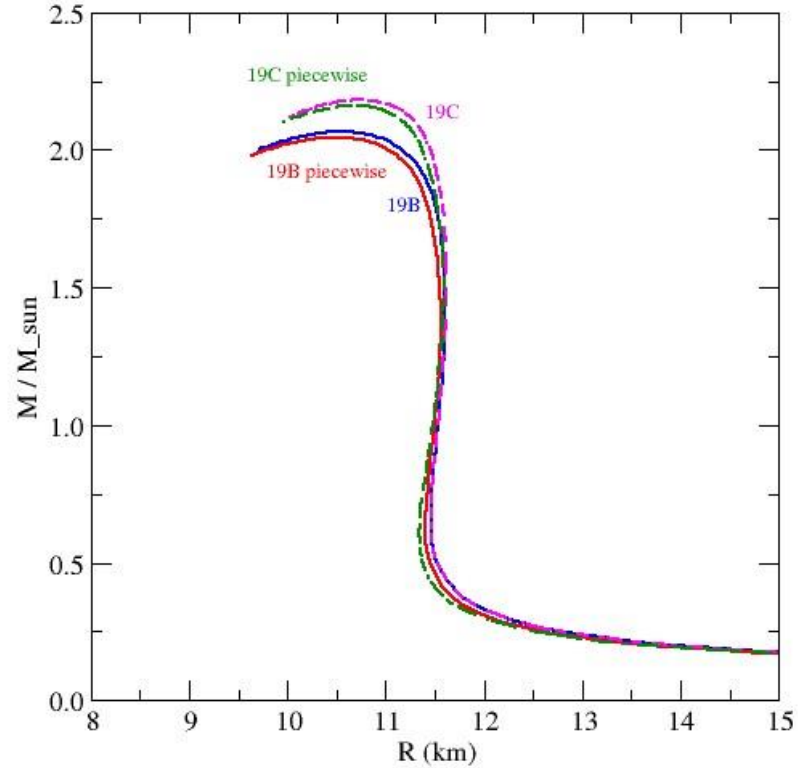
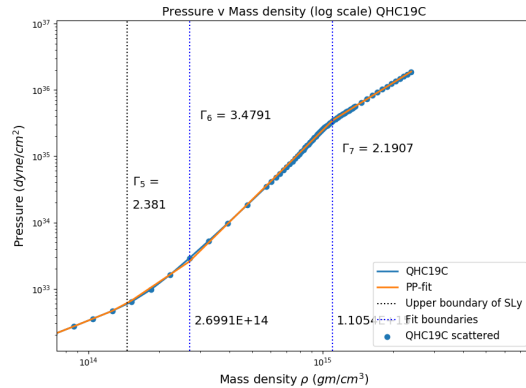
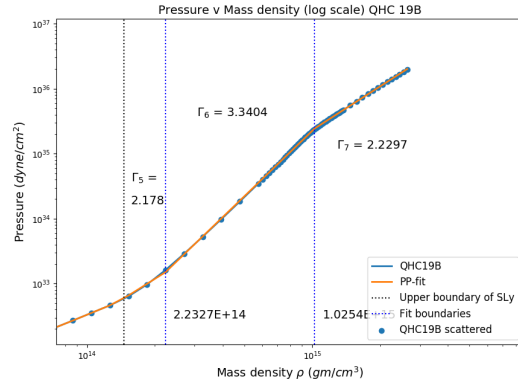
We fit the High density QHC region using this scheme and obtain the EoS Pieces:

Degrees of freedom in fitting ($\Gamma_5, \rho_5, \Gamma_6, \rho_6, \Gamma_7$)

$$\text{RMS Residual} = \sqrt{\frac{1}{m} (\sum_i (\text{Pieces}) \sum_j (\text{Densities}) (\log P_j - \log K_i - \Gamma_i \log \rho_j)^2)} = 0.014 \text{ i.e. } \log P = 35 \pm 0.014.$$

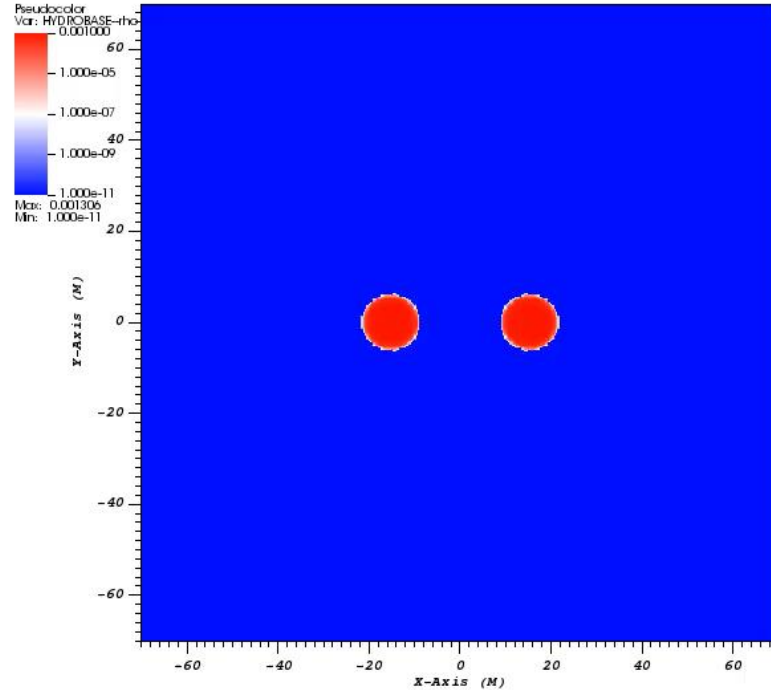


Mass-Radius relations of QHC19



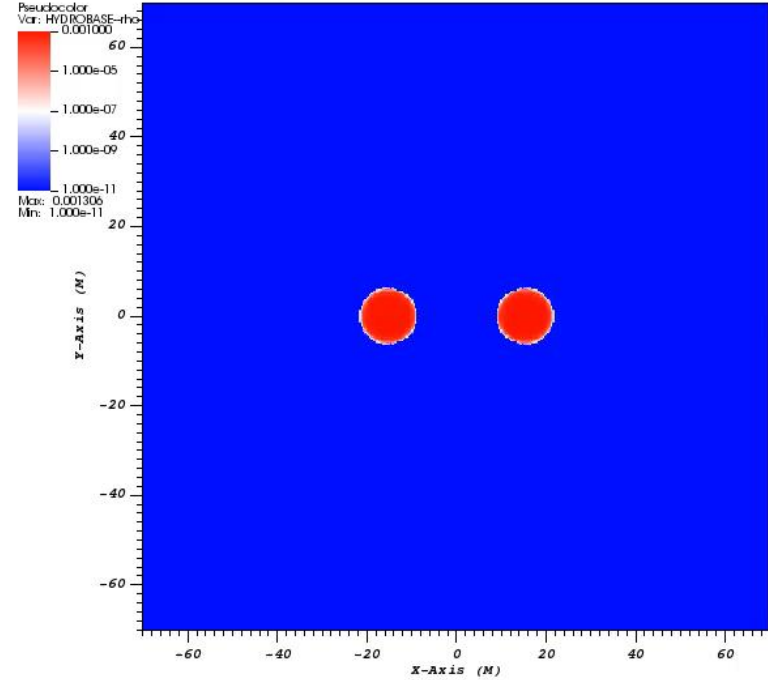
Merger for QHC19C

DB: rho.xy.h5 Rest Mass Density [$1/\text{Msun}^2$]
Cycle: 0 Time: 0



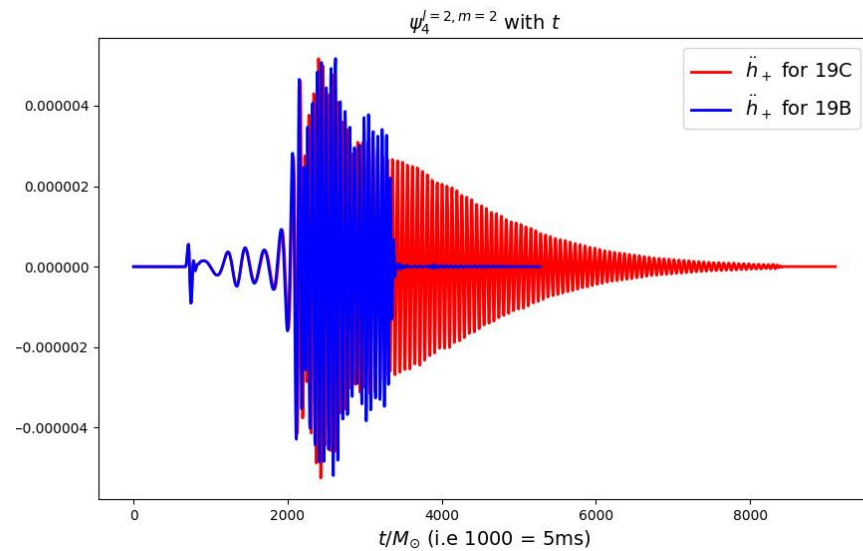
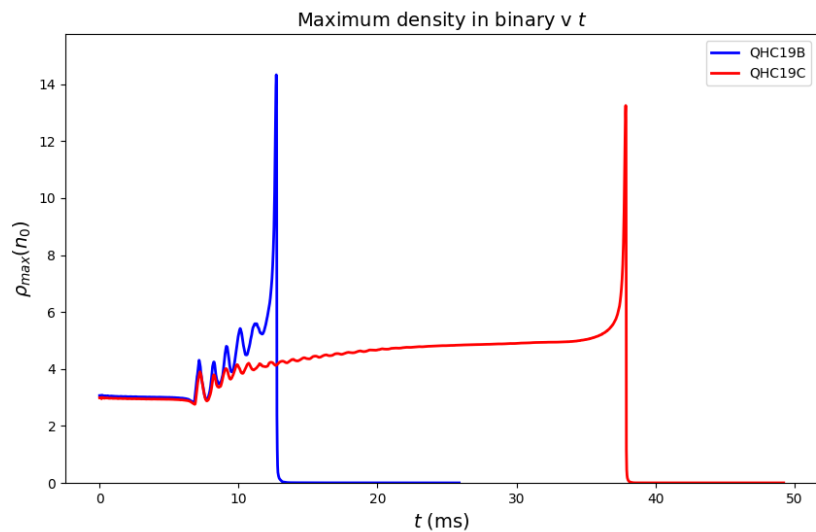
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DB: rho.xz.h5 Rest Mass Density [$1/\text{Msun}^2$]
Cycle: 0 Time: 0

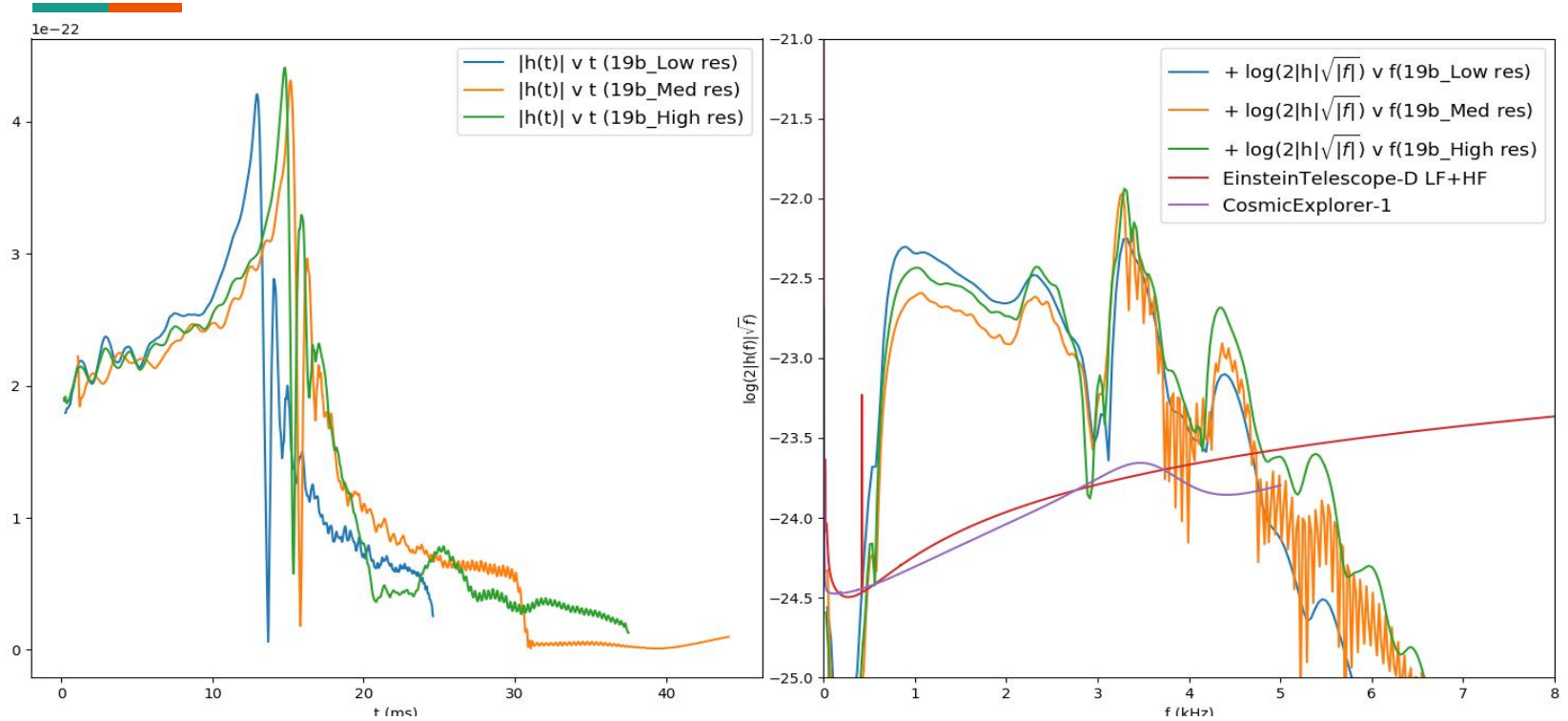


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Mergers of 19B v 19C



Convergence and Spectrum of 19B



Summary and Conclusion



Summary :

- The high density EoS is uncertain, and can be studied via BNSMs.
- QHC19 EoS endeavors to explain the behaviour at the phase transition and quark matter state.
- We parameterized the QHC19B-C EoS and performed mergers to see stark differences in the post merger evolution.

Future work :

- Convergence tests for the evolution for massive NS simulations.
- Detailed Gravitational waves analysis.
- Evolutions of all four QHC19A-D

Thank you! Questions?

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