



# Ground Tests Commanding Manual

KU Leuven PLATO Team

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# Changelog

**12/02/2024 — 0.11**

- Added a warning that loading/submitting a Setup doesn't propagate this new Setup to all components in the system, see [Chapter 6](#)

**29/06/2023 — 0.10**

- Added table with N-FEE modes and their register values for PFM, see [Table 3](#)

**28/06/2023 — 0.9**

- Added an appendix to explain the difference between GitHub issues, NCRs and PVS, see [Appendix 17.E](#)
- Added a section on actions on the Setup when a new camera arrives, see [Chapter 9](#). (This is taken from a technical note by Pierre PLATO-KUL-PL-TN-0022)
- updated CCD layout images in a TAB for EM and FM, see [Chapter 12](#).
- Added section on SFT analysis to section on Camera Switch ON, see [Chapter 10](#).
- minor editorial changes like fixing links, typos, minor rewording, etc.

**20/06/2023 — v0.8**

- Added explanation about Camera Switch ON / OFF, see [Chapter 10](#)

**19/06/2023 — v0.7**

- bringing the commanding manual up-to-date with the current implementation of the Common-EGSE, the test scripts and the situation in the test houses.
  - Section on the plato-test-prep repository has been removed (was Section 3.3)
  - Data Acquisition: updates in all sections
  - Configuration and Setups: updates in all sections
  - Common-EGSE startup, shutdown and sleep: updates in all sections
  - Utility functions: updates in all sections
  - Operating the N-FEE: updates in all sections
  - Operating the TCS EGSE: only editorial updates
  - Operating the OGSE: updates in all sections
  - Operating the tests, system states: only editorial changes
  - Appendix A: removed unimplemented ALT mode for `ccd_side` and updated other commands with correct parameters
  - Appendix B: updated commands for visited positions
- added a backlink to the CGSE Documentation web site for your convenience. It's at the top of the HTML page.

▼ *Click to see older changelog entries*

**12/06/2023 — v0.6**

- move the commanding manual to the CGSE documentation page and converted into



asciidoc.

- Update section on data structure

#### **14/06/2021 — v0.5**

- Update section 11.3 (rem. col\_end from BB signatures)

#### **03/05/2021 — v0.4**

- Update sections 3.2, 3.3, 7.3, 7.4
- Sections 11.1 & 11.3 Change ccd\_side & EF convention
- New sections 5.3, 6.3, 11.4 and 14
- Update sections 3.2, 3.3, 7.3, 7.4
- Sections 11.1 & 11.3 Change ccd\_side & EF convention
- New sections 5.3, 6.3, 11.4 and 14



# Colophon

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1<sup>st</sup> Edition — February 2023

This manual is written in PyCharm using the AsciiDoc plugin. The PDF Book version is processed with asciidoctor-pdf.

The manual is available as HTML from [ivs-kuleuven/github.io](https://ivs-kuleuven.github.io/plato-cgse-doc/). The HTML pages are generated with Hugo which is an OSS static web-pages generator. From this site, you can also download the PDF books.

The source code is available in a GitHub repository at [ivs-kuleuven/plato-cgse-doc](https://github.com/ivs-kuleuven/plato-cgse-doc).

When you find an error or inconsistency or you have some improvements to the text, feel free to raise an issue or create a pull request. Any contribution is greatly appreciated and will be mentioned in the acknowledgement section.



# Conventions used in this Book

We try to be consistent with the following typographical conventions:

## Italic

Indicates a new term or ...

## Constant width

Used for code listings, as well as within paragraphs to refer to program elements like variable and function names, data type, environment variables (**ALL\_CAPS**), statements and keywords.

## Constant width between angle brackets <text>

Indicates `text` that should be replaced with user-supplied values or by values determined by context. The brackets should thereby be omitted.

When you see a `$ ...` in code listings, this is a command you need to execute in a terminal (omitting the dollar sign itself). When you see `>>> ...` in code listings, that is a Python expression that you need to execute in a Python REPL (here omitting the three brackets).

## Setup versus setup

I make a distinction between Setup (with a capital C) and setup (with a small c). The Setup is used when I talk about the object as defined in a Python environment, i.e. the entity itself that contains all the definitions, configuration and calibration parameters of the equipment that make up the complete test setup (notice the small letter 'c' here).

(sometimes you may find setup in the document which really should be 'set up')

## Using TABs

Some of the manuals use TABs in their HTML version. Below, you can find an example of tabbed information. You can select between FM and EM info and you should see the text change with the TAB.

### NOTE

This feature is only available in the HTML version of the documents. If you are looking at the PDF version of the document, the TABs are shown in a frame where all TABs are presented successively.

### FM

In this TAB we present **FM specific information**.

### EM

In this TAB we present **EM specific information**.

## Using Collapse

Sometimes, information we need to display is too long and will make the document hard to read. This happens mostly with listings or terminal output and we will make that information collapsible. By default, the info will be collapsed, press the small triangle before the title (or the



title itself) to expand it.

**NOTE** In the PDF document, all collapsible sections will be expanded.

▼ *A collapsible listing*

```
plato-data@strawberry:/data/CSL1/obs/01151 CSL1_chimay$ ls -l
total 815628
-rw-r--r-- 1 plato-data plato-data      7961 Jun 20 10:38 01151_CSL1_chimay_AEU-AWG1_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data      9306 Jun 20 10:38 01151_CSL1_chimay_AEU-AWG2_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     309375 Jun 20 10:38 01151_CSL1_chimay_AEU-CRIO_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     42950 Jun 20 10:38 01151_CSL1_chimay_AEU-PSU1_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     43239 Jun 20 10:38 01151_CSL1_chimay_AEU-PSU2_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     42175 Jun 20 10:38 01151_CSL1_chimay_AEU-PSU3_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     42327 Jun 20 10:38 01151_CSL1_chimay_AEU-PSU4_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     42242 Jun 20 10:38 01151_CSL1_chimay_AEU-PSU5_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     42269 Jun 20 10:38 01151_CSL1_chimay_AEU-PSU6_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     67149 Jun 20 10:38 01151_CSL1_chimay_CM_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data    20051 Jun 20 10:38 01151_CSL1_chimay_DAQ6510_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data      105 Jun 20 10:38 01151_CSL1_chimay_DAS-DAQ6510_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data    19721 Jun 20 10:38 01151_CSL1_chimay_DPU_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data    22833 Jun 20 10:38 01151_CSL1_chimay_FOV_20230620_095819.csv
-rw-rw-r-- 1 plato-data plato-data 833754240 Jun 20 10:34 01151_CSL1_chimay_N-FEE_CCD_00001_20230620_cube.fits
-rw-r--r-- 1 plato-data plato-data    292859 Jun 20 10:38 01151_CSL1_chimay_N-FEE-HK_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     8877 Jun 20 10:38 01151_CSL1_chimay_OGSE_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data    19841 Jun 20 10:38 01151_CSL1_chimay_PM_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data   188419 Jun 20 10:38 01151_CSL1_chimay_PUNA_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data     7662 Jun 20 10:38 01151_CSL1_chimay_SMC9300_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data    19781 Jun 20 10:38 01151_CSL1_chimay_SYN_20230620_095819.csv
-rw-r--r-- 1 plato-data plato-data   147569 Jun 20 10:38 01151_CSL1_chimay_SYN-HK_20230620_095819.csv
plato-data@strawberry:/data/CSL1/obs/01151 CSL1_chimay$
```



# Purpose

This document describes the concepts and some practicalities relative to the commanding of the PLATO Cameras (AEU, FEE) and various GSEs to be controlled during the ground-based testing of the PLATO Cameras.

The document shall be used to train members of the PLATO camera test team in operating the Common-EGSE system and write commanding scripts for camera-level tests.

The present version still contains a few ‘TBW’, some of which will be filled thanks to inputs from the Test Houses (TH), describing the location-specific GSE controls.



# Documents and Acronyms

## Applicable documents

ID	Document Title	Document Number	Issue

## Reference Documents

ID	Document Title	Document Number	Issue
RD-01	CAM Test EGSE User Manual	PT-EVO-SYS-MA-0261-1	1
RD-02	Common-EGSE User Manual	PLATO-KUL-PL-MAN-0001	0.1
RD-03	Test Specification	PLATO-KUL-PL-TS-0001	2.7
RD-04	Common-EGSE ICD	PLATO-KUL-PL-ICD-0002	1.0
RD-05	PLATO-AEU CAM TEST EGSE TMTC ICD	PTO-ECO-SYS-ICD-0188	2B
RD-06	PLATO N-FEE ICD	PLATO-MSSL-PL-ICD-0002	9.0
RD-07	Reverse clocking for N-CAMs	PLATO-MSSL-PL-TN-0015	1.0
RD-08	N-FEE Readout Operations	PLATO-MSSL-PL-TN-0012	2.1
RD-09	PLATO PL TCGSE UNIT Subsystems and Equipment ICD and IDS	PTO-AST-PL-TCGSE-ICD-0012	10.0
RD-10	PLATO Instrument Coordinate Systems	PLATO-OHB-PL-LI-0009	05
RD-11	PLATO Room temperature collimator user manual	PLATO-UOL-PL-RP-0004	1
RD-12	R1_CalibrationReport	PLATO-PL-CSL-RP-0031	3.0
RD-13	R2_CalibrationReport	PLATO-PL-CSL-RP-0032	2.1
RD-14	N-FEE Conceptualised Operation Specification	PLATO-MSSL-PL-SP-0015	2.0
RD-15	PLATO N-FEE PFM Register Map	PLATO-MSSL-PL-FI-0004	1.0

## Acronyms

AEU	Ancillary Electronics Unit
API	Application Programming Interface
CAM	Camera
CCD	Charged-Coupled Device
CGSE	Common-EGSE
CSL	Centre Spatial de Liège
CSV	Comma-Separated Values



COT	Commercial off-the-shelf
CTI	Charge Transfer Inefficiency
CTS	Consent to Ship
DPU	Data Processing Unit
DSI	Diagnostic SpaceWire Interface
EGSE	Electrical Ground Support Equipment
EOL	End Of Life
FAQ	Frequently Asked Questions
FEE	Front End Electronics
FITS	Flexible Image Transport System
FPA	Focal Plane Assembly/Array
GSE	Ground Support Equipment
GUI	Graphical User Interface
HDF5	Hierarchical Data Format version 5 (File format)
HK	Housekeeping
IAS	Institut d'Astrophysique Spatiale
ICD	Interface Control Document
LDO	Leonardo space, Italy
MGSE	Mechanical Ground Support Equipment
MMI	Man-Machine Interface
NCR	Non-Conformance Report
NRB	Non-Conformance Review Board
OBSID	Observation Identifier
OGSE	Optical Ground Support Equipment
OS	Operating System
PDF	Portable Document Format
PID	Process Identifier
PLATO	PLAnetary Transits and Oscillations of stars
PPID	Parent Process Identifier
PLM	Payload Module
PVS	Procedure Variation Sheet
REPL	Read-Evaluate-Print Loop, e.g. the Python interpreter prompt
RMAP	Remote Memory Access Protocol
SFT	Short Functional Test



SpW	SpaceWire
SQL	Structured Query Language
SRON	Stichting Ruimte-Onderzoek Nederland
SUT	System Under Test
SVM	Service Module
TBC	To Be Confirmed
TBD	To Be Decided or To Be Defined
TBW	To Be Written
TC	Telecommand
TCS	Thermal Control System
TH	Test House
TM	Telemetry
TOU	Telescope Optical Unit
TS	Test Scripts
TUI	Text-based User Interface
TV	Thermal Vacuum
UM	User Manual
USB	Universal Serial Bus
YAML	YAML Ain't Markup Language



# 1. Introduction

## 1.1. Contents

This document describes the usage of the PLATO Common-EGSE to command the PLATO camera (SUT) and the ground support equipment (GSE) during the PLATO camera ground testing.

Throughout this document, we use the following notation:

- a command to be executed in a terminal on either the egse-server or on the client machine.

```
$ command
```

- a command to be executed in a Python session. This Python session will always be running on the client machine.

```
>>> command
```

## 1.2. EGSE commanding software environment

The user interface to the PLATO camera EGSE is a software application developed at KU Leuven in collaboration with the test houses for site specific components. It provides graphical user interfaces to control the software processes in the system, monitor telemetry parameters, provide the operator with quick-look analysis images of the camera detectors, etc. It also features a commanding interface in Python. The commanding logic is entirely defined in Python. The hardware interfaces have been implemented in Python (possibly accessing a library written in the C programming language), and the user is accessing these by executing higher-level Python functions in a Python interpreter.

## 1.3. How different user profiles use the software

- **test-operator**: running a test, launching scripts and inspecting results of quick look analysis script at the test house. The test operator is knowledgeable of the Common-EGSE, Test Scripts, all GSE and basic operation of the PLATO Camera (SUT).
- **test-developer**: translating test specification into commanding scripts, writing quick-look analysis scripts.
- **offline-analyst**: loads the test data from the archive and process this data for validation and feedback.
- **site-operator**: manages the test-infrastructure, i.e. the test-environment in the TH, administration of egse-server and client machines.
- **site-developer**: implements test house specific software for test equipment used at test houses. This includes but is not limited to device drivers, GUI applications, local setups.

Typical flow of events and responsibilities:



- Long before the tests are executed, the site-developer implements the interfaces to the test-equipment in the Common-EGSE.
- Months before the test, the test-developer turns the test-specification into a commanding script.
- Minutes before the test, the site-operator starts up the system, activates all connections, launches GUIs, loads the proper system configuration into the system, and gives a go-ahead to the test-operator.
- At t0, the test-operator launches the execution of the commanding script and follows it in real-time.
- At the end of the execution, the analyst gets hold of the data and starts the analysis.

This commanding manual is mainly targeted at the test developers, the test operators and the analysts. The site-operators and site-developers will find most of the information they need in the [Common-EGSE installation manual](#) and [user manual](#).



## 2. Software Overview

The PLATO commanding during ground-based testing resides upon three GitHub repositories:

- [plato-common-egse](#) ([link to documentation](#))
- [plato-test-scripts](#) ([link to documentation](#))
- [plato-cgse-conf](#)

Amongst others, the documentation will point you to the installation guides, list the low-level commands that are available in the Common-EGSE, list the low-level building blocks implemented in the test scripts etc.

### 2.1. The plato-common-egse GitHub repository

This repository contains the background infrastructure to interface with the actual hardware in the test houses. The complete documentation can be found in the link above, and in RD-02, see [Reference Documents](#).

In a nutshell, it provides access to all the necessary functions to

- Operate the camera under test
- Operate all devices in the test-environment (in or out of the TVAC chamber)
- Record all telemetry
  - From the camera (image data & HK)
  - From the test-environment (HK)
- Maintain the setup and calibration file information under configuration control (see ‘setup’ below).
- Command the camera and the test-environment devices from Python scripts (see plato-test-scripts)

The THs are expected to contribute to this repository, to implement the interfaces to their own environment and devices, so they should fork this repository in order to be able to create pull requests.

The test-operators are not expected to contribute to this repository, so they should clone the repository.

### 2.2. The plato-test-scripts GitHub repository

This repository contains the Python scripts

- for commanding the tests as defined in the test specification,
- for analysing the data

The commanders and the THs are expected to contribute to this repository, so everyone should fork it.



## 2.3. The plato-cgse-conf GitHub repo

This repository holds the setups saved during the tests under configuration control. The concept of setup is explained in [Chapter 6](#).

No one is expected to contribute directly to this repository as it is maintained by the Common-EGSE. The repository can be cloned for inspection.

## 2.4. Software installation

Operators use the software installation at the test house on the operational machine. The installation is read only and under configuration control. Only official releases shall be installed on the operational machines. The installation is maintained by the site-operator. Please refer to the [on-line documentation of the Common-EGSE](#) for full installation details.

```
$ git clone https://github.com/IvS-KULeuven/plato-common-egse.git
$ git checkout tags/<release-tag> -b <release-tag>-branch
$ python3.8 setup.py install --home=/cgse
```

The test-developer uses an installation that is more suited for development of scripts and can be changed. The test-developer has forked the repo on the GitHub website to her personal GitHub account and clones the repo from that account. She works in a virtual environment. Please refer to the [Common-EGSE on-line docs](#) for full details.

```
$ git clone https://github.com/<github-username>/plato-common-egse.git
$ source venv/bin/activate
$ python3.8 setup.py develop
```

Please note that the installation examples above are simplified and serve as a reminder. The full installation process is detailed in the on-line Common-EGSE documentation.

## 2.5. Naming Convention

The table below summarises the coding style that we have adopted for this project. More detailed information can be found in the on-line documentation at [github.io: Development notions: Style Guide](#)

*Table 1. Summary of the adopted code style.*

Type	Style	Examples
Variables, building_block parameters, Entries in setup files	lowercase with underscores	key, last_value, model, index, user_info
Methods, functions, building_blocks	lowercase with underscores	get_value, set_mask, create_image
Classes	CapWords (no underscores)	ProcessManager, ImageViewer, CommandList, Observation, MetaData



Type	Style	Examples
Constants	UPPERCASE with underscores	MAX_LINES, BLACK, COMMANDING_PORT
Modules & Packages	lowercase without underscores	dataset, commanding, multiprocessing

### Top-level scripts

Top-level scripts start with the test identifier from the test specification, followed by a descriptive name in lowercase, words separated by underscores, e.g., `cam_tvpt_010_best_focus_determination.py`.

### Building blocks

Building blocks are normal function definitions that are decorated with `@building block` to identify them as separated by one underscore, i.e., `snake_case`.

### User utility functions

User functions are normal Python functions and follow normal Python function naming conventions, i.e., all lower case and words separated by underscores, i.e., `snake_case`.



# 3. Test script architecture

## 3.1. Overview

At the user level, the PLATO commanding resides upon a few key concepts

- **Building-block:** a building block is a Python function implementing the commands corresponding to a logical entity within a test (a Lego©-bloc) or an entire test (the Lego©-house, called “test-script”).
- **Test-script:** a test-script is a building block encapsulating all the commands necessary for a given test. The name of a test-script should identify it in the test-specification document (RD-03). The test-script should be unique, i.e. the test script for any given test should be useable in all test-houses.
- **Execution:** building blocks, and only building blocks, can be executed, i.e. run on the operational machine inside the test-environment.
- **Observation:** executing a building block triggers the creation of an “observation”. An observation is defined by a unique “observation identifier” (obsid) and lasts as long as the execution of the corresponding test (building block).
- **Setup:** a Setup encapsulates the complete configuration of the test-environment (test-equipment) and of the camera (test-item) as well as all calibration files associated either to the hardware, or necessary for a given test.

More information on some of these elements are given in the next sections. The figure below describes the generic software architecture in place for the PLATO commanding:

- At the top level
  - a test script implements an entire test.
- At an intermediate level
  - The test script can call Python functions (e.g. “calculate the next FOV position to visit”) or lower level building blocks (e.g. “go to the next FOV position” or “acquire images over a 5x5 dither pattern”)
  - Building blocks can call other building blocks, or regular Python functions (recursion is forbidden within building blocks)
  - Functions can call other functions, but should not call building blocks
- At the lowest level, the Common-EGSE software provides all the “atomic commands” necessary to interface with the hardware. This layer allows to provide the users with user-friendly commands (e.g. human-readable parameter names, and no bit-field or hexadecimal numbers to provide).
- The active test-setup is available at all levels to provide all necessary information with respect to hardware or calibration.

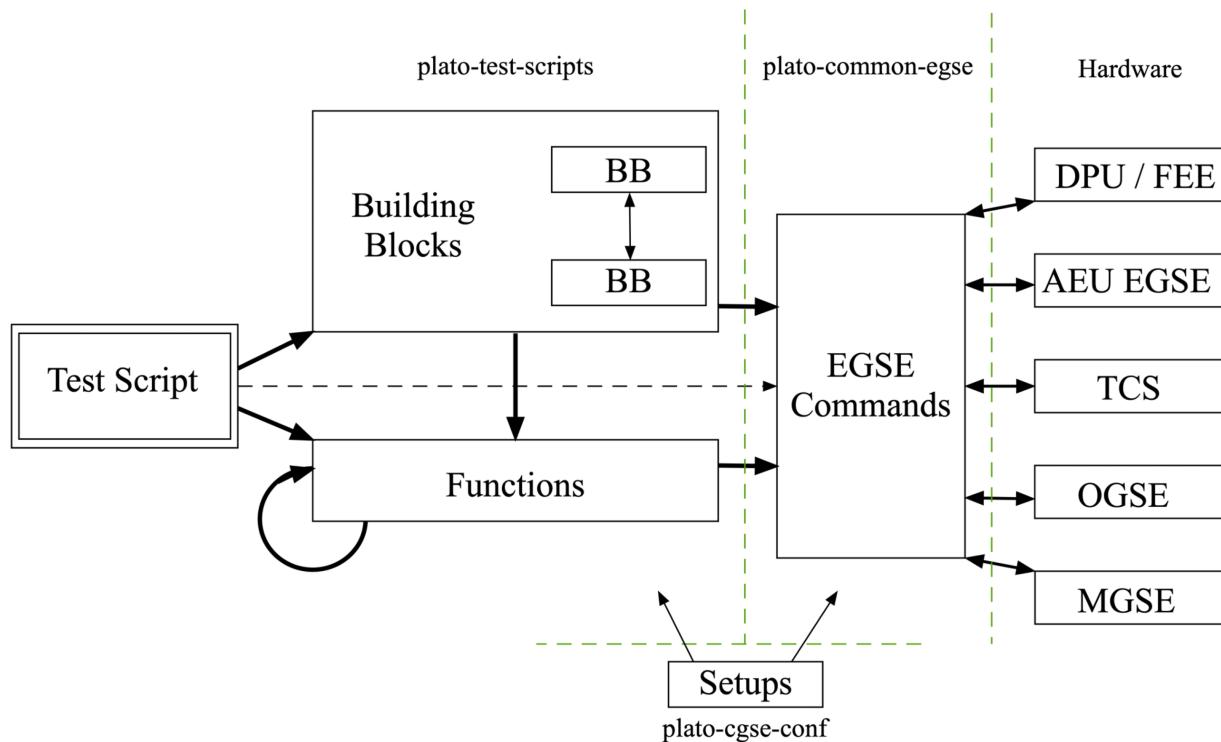


Figure 1. Test Scripts architecture and its relation to the Common-EGSE and the GSE and SUT.

## 3.2. Building block

### 3.2.1. Definition of a building block

A building block is a Python function marked with a specific decorator in the code. That simply means that the line directly above the definition of the building block in the code should be `@building_block`:

Example of a building block definition:

```
@building_block
def move_filter_wheel(position=None):
    # your code comes here
    return True
```

All parameters of a building block are keyword parameters. That means the parameter name and its value must be specified when calling the building block.

```
...
move_filter_wheel(position="A")
...
```

### 3.2.2. Building block justification

**Execution:** Building blocks are the only entity in the code that will be accepted for execution on the operational machines in the test-environments (THs).

**Structure:** a unique identifier (BBID) is automatically attributed to every building block entering the



plato-test-scripts repository. At every moment during a test execution, the current BBID is recorded in the data (TBC).

Organising a test script into a logical structure and implementing the underlying building blocks accordingly will help structure the data, which in turn will be precious to ease the analysis.

For instance, if a test-script (BB1) is calling building blocks BB2 and BB3 sequentially, the telemetry will contain a sequence of timestamps where the individual building-blocks start, or end.

Note that, as a further convenience, a building block counter (BBCOUNT) will also be recorded, which is a natural number incremented by one every time the active BBID is changing, hence running from 0 to 4 in the example above (TBC).

### 3.2.3. The properties of a building block

Building blocks implement some safeguards, imposing a number of limitations on the code

- Building blocks cannot be called recursively. Beware of building blocks calling other building blocks. Avoid too many layers. Avoid functions calling building blocks
- At run time, the names and the values of every arguments of a building blocks must be explicitly given. Building blocks with many parameters are hence strongly discouraged
- `def my_block(param1, param2=3)`: is forbidden for 2 reasons
  - Positional arguments are forbidden (`param1`)
  - All arguments must be given the default value `None` (`param2`). The reason for this is that we want, when the building block is called, all arguments to be passed explicitly, not implicitly.

### 3.2.4. Using default argument values in a building block

#### **WARNING**

we strongly discourage using this way of '*work around default parameters*'. It is currently still implemented in the core system, but this behaviour might/will be removed in the future.

In the code definition, it is not possible (prev. subsection).

It is nevertheless possible via an input file. The input file must

- be in YAML format
- bear the same name as the building block it corresponds to
- be put in directory camtest/commanding/input

Example: building block `camtest/commanding/hk_only.py`

```
@building_block
def hk_only(wait_time=None):
    time.sleep(wait_time)
```

Corresponding input file `camtest/commanding/input/hk_only.yaml`



```
# 'hk_only' - Building Block
Args:
    wait_time:      10          # System idle time [seconds]
```

### 3.2.5. Utility building blocks

A collection of low-level, general purpose building blocks is already provided in the commanding section of the plato-test-scripts ([camtest/commanding](#)), to help the test-developer, for instance to manipulate some hardware device, tune the OGSE light intensity, set the FEEs in different operating modes, acquire a number of full-frame images etc.



# 4. Test execution

## 4.1. Test execution: execute

The execution of a test is triggered by the Python function `execute()`. Like `@building_block`, it is a core functionality implemented by plato-test-scripts. It can be used with the following syntax:

```
>>> execute(building_block_name, param1=value1, param2=value2)
```

All parameters must be specified by their name. No positional argument is allowed; hence the order of the parameters is not important. Note that the building block name does not have the parenthesis '`()`', only the name of the building block is given.

The `execute` command will start an observation and end the observation when the building block returns. Starting and ending an observation is an expensive operation in the sense that it notifies underlying mechanisms like the configuration and storage manager of the observation so they can take action. Therefore, although any building block can be executed using the `execute(..)` command, this should really only be used for higher level building blocks and from the Python prompt. Never use the `execute(..)` function inside a building block.

## 4.2. Preview the command sequence

**WARNING**

This functionality will be removed in the future as it has only limited applicability and doesn't give a full consistent view of the command sequence.

It is possible to perform a dry-run of a building-block by running `generate_command_sequence()` instead of `execute()`, with the same syntax. Be aware that the duration of the dry run may be as long as the execution itself!

```
>>> generate_command_sequence(building_block_name, param1=value1, param2=value2)
```

This feature will execute all building blocks and functions without actually sending command strings to the test equipment. The current implementation does not take return values from device queries into account which makes it less suited for test scripts that need this feedback for conditional processing, e.g., waiting until a temperature is reached.

Note that `execute` and `generate_command_sequence` will only work under a set of restrictive conditions

- Disposing of software simulators for every piece of equipment addressed by the test. Simulators exist in the EGSE for the N-FEEs and most of the equipment to be used in CSL.
- Emulating an operational environment on your machine. We refer you to the EGSE documentation for the details of what this entails



## 4.3. Executing short building blocks individually

The execute command will trigger the creation of an obsid and an associated data stream. In the case were the operator (e.g. while setting up and testing the TH environment) want to execute short commands or building blocks outside of the scope of a test-script, this is an overkill and will make the analysis of the resulting data very cumbersome, because the data will be distributed over many very very short obsids.

To work around this, it is possible to manually start and stop an observation

```
>>> start_observation("Running a few examples")
```

Will start an observation, attribute it an obsid and the associated data stream, just like what happens at the start of an **execute**.

After that any command passed (individually or within a function/building block) and any data generated will be recorded as part of the running obsid:

```
>>> command(args)
>>> func(args)
>>> building_block(args)
```

Finally

```
>>> end_observation()
```

Will close the observation (or do nothing if none is running).



# 5. Data acquisition and storage

All device housekeeping data and all camera telemetry data (image data and FEE HK) are automatically stored by the Storage Manager. You don't need to collect and save data from within your scripts. Data is stored in a location on the egse-server and accessible from the desktop client machines running the tests and analysis scripts. This location is defined in the environment variable **PLATO\_DATA\_STORAGE\_LOCATION**, which points to a `/data` directory.

**NOTE**

Since release 2023.6.0+CGSE, the environment variable **PLATO\_DATA\_STORAGE\_LOCATION** will not include the SITE ID anymore.

Several types of data are generated during a test campaign

- Housekeeping telemetry (saved in CSV files)
  - From the test-item (e.g. TOU TRP-1 temperature)
  - From the test-environment (e.g. OGSE filter wheel position)
- SpaceWire packets and register maps received from the FEE (saved in HDF5 files)
- Image data from the camera (saved in FITS files)
- Command history logging the test executions

## 5.1. Housekeeping telemetry

The acquisition of housekeeping telemetry is automatically started by the egse-server. In practice that means that the telemetry from all devices connected to the active egse-server is automatically, and constantly recorded (see the Common-EGSE documentation for further details).

**IMPORTANT**

Please note that **all timestamps** saved in housekeeping files are UTC.

The housekeeping telemetry is structured this way

- One file per day saved in a sub-folder of 'daily'. The sub-folder is the **YYYYMMDD** of that day.
- One CSV file per source of telemetry, i.e. for example one from the TCS, one for the hexapod (controller), one for the FEE etc.
- Each CSV file contains a timestamp in the first columns, and the various entries in the subsequent columns.
- All TM are recorded at 1Hz by default. This is configurable on a per-device-controller basis (= per CSV file)

The csv filenames have the following structure:

`YYYYMMDD_SITE_DEVICE.csv`

Where:



- **YYYYMMDD** is the day the telemetry was taken
- **SITE** is the identifier for the test facility: CSL, IAS, INTA, SRON, KUL, ESA, ...
- **DEVICE** is a mnemonic referring to the device controller issuing these data

For example, the daily file at CSL for the Hexapod PUNA housekeeping on June 8th, 2023 is 20230608\_CSL\_PUNA.csv. The header of the CSV files contains explicit column-names. The complete information on the content and format of the telemetry files is contained in the ICD (RD-04, see [Reference Documents](#)).

▼ *Excerpt of the PUNA hexapod telemetry file 20230608\_CSL\_PUNA.csv:*

```

timestamp,GCSL1_HEX_USER_T_X,GCSL1_HEX_USER_T_Y,GCSL1_HEX_USER_T_Z,GCSL1_HEX_USER_R_X,GCSL1_HEX_USER_R_Y,GCSL1_HE
X_USER_R_Z,GCSL1_HEX_MACH_T_X,GCSL1_HEX_MACH_T_Y,GCSL1_HEX_MACH_T_Z,GCSL1_HEX_MACH_R_X,GCSL1_HEX_MACH_R_Y,GCSL1_H
EX_MACH_R_Z,GCSL1_HEX_ALEN_1,GCSL1_HEX_ALEN_2,GCSL1_HEX_ALEN_3,GCSL1_HEX_ALEN_4,GCSL1_HEX_ALEN_5,GCSL1_HEX_ALEN_6
,GCSL1_HEX_HOMED,GCSL1_HEX_IN_POS
...
2023-06-08T10:00:01.560+0000,0.014144539424,-0.003925761937,-3.489246984,-0.013989085157,0.0010419456108,-
0.00799891817,0.390713812,0.1455886605,17.7970682,0.03226852454,0.06991046997,0.2646041152,205.93219583,206.16533
51,205.52915657,205.97228441,205.88274269,206.14508725,True,True
2023-06-08T10:00:02.560+0000,0.014144539424,-0.003925761937,-3.489246984,-0.013989085157,0.0010419456108,-
0.00799891817,0.390713812,0.1455886605,17.7970682,0.03226852454,0.06991046997,0.2646041152,205.93219583,206.16533
51,205.52915657,205.97228441,205.88274269,206.14508725,True,True
2023-06-08T10:00:03.563+0000,0.014144539424,-0.003925761937,-3.489246984,-0.013989085157,0.0010419456108,-
0.00799891817,0.390713812,0.1455886605,17.7970682,0.03226852454,0.06991046997,0.2646041152,205.93219583,206.16533
51,205.52915657,205.97228441,205.88274269,206.14508725,True,True
2023-06-08T10:00:04.562+0000,0.014144539424,-0.003925761937,-3.489246984,-0.013989085157,0.0010419456108,-
0.00799891817,0.390713812,0.1455886605,17.7970682,0.03226852454,0.06991046997,0.2646041152,205.93219583,206.16533
51,205.52915657,205.97227986,205.88274269,206.14508725,True,True
2023-06-08T10:00:05.562+0000,0.014144539424,-0.003925761937,-3.489246984,-0.013989085157,0.0010419456108,-
0.00799891817,0.390713812,0.1455886605,17.7970682,0.03226852454,0.06991046997,0.2646041152,205.93219583,206.16533
51,205.52915657,205.97227986,205.88274269,206.14508725,True,True
2023-06-08T10:00:06.581+0000,0.014144539424,-0.003925761937,-3.489246984,-0.013989085157,0.0010419456108,-
0.00799891817,0.390713812,0.1455886605,17.7970682,0.03226852454,0.06991046997,0.2646041152,205.93219583,206.16534
276,205.52915967,205.97228296,205.88274579,206.14508725,True,True

```

## 5.2. Observation

Image data are usually only generated during a test execution. In order to facilitate the selection of data pertaining to a particular test execution, we are using the concept of ‘observation’ interchangeably with the name “test”.

An observation corresponds to a single execution of a given test, i.e. a single use of the execute() command, or all commands executed between a **start\_observation** and an **end\_observation**.

A unique obsid is automatically attributed to every observation. The filenames of housekeeping files are constructed from the unique obsid and a limited set of useful metadata about the test execution, with the following structure:

```
TEST_SITE_SETUP_DEVICE_YYYYMMDD_HHMMSS.csv
```

where:

- **TEST** is a unique test identifier, this is the first part of a full OBSID,



- **SITE** is the identifier for the test facility: CSL, IAS, INTA, SRON, KUL, ESA, ...
- **SETUP** is the unique setup\_id (section below)
- **DEVICE** is a mnemonic referring to the source of the data, e.g. the device controller issuing these data, the TCS EGSE, the N-FEE, etc.
- **YYYYMMDD\_HHMMSS** points to the start of the test execution (UTC)

The first three items in the above list form the full OBSID. The OBSID can be abbreviated to the TEST-ID and the SITE-ID thereby omitting the SETUP-ID which can be automatically recovered from the meta data.

For instance, CSV files for observation 00757 performed at CSL could be:

- **00757\_CSL\_00009\_PUNA\_20201013\_162037.csv**
- **00757\_CSL\_00009\_CM\_20201013\_162037.csv**

(PUNA points to a hexapod controller and CM refers to the Configuration Manager of the Common-EGSE).

For FITS files with the camera image data, the naming convention is similar (though we leave out the timestamp), with a sequencing number when the data volume justifies splitting of the data over multiple files.

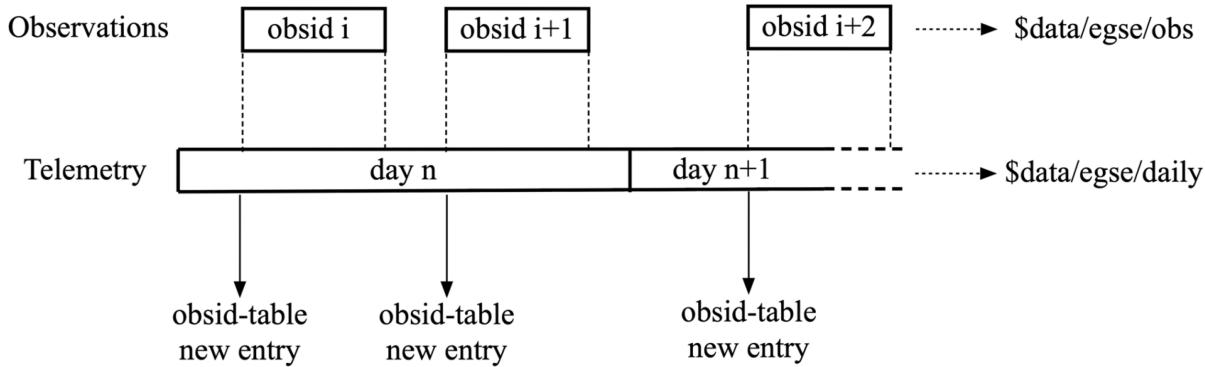
- **00757\_CSL\_00009\_N-FEE\_CCD\_00001\_20201013\_cube.fits**
- **00757\_CSL\_00009\_N-FEE\_CCD\_00002\_20201013\_cube.fits**

Note that we used to have two types of FITS files: FITS files in which the data is organised in a flat structure, and FITS files in which the data is organised in cubes (with filenames as described above). The former will have "image" in their names (rather than "cube") but will be removed once the corresponding cube-structure FITS file has been created.

All data corresponding to the obsid are gathered in a dedicated directory named ‘obs’. That means that the fraction of the regular housekeeping data acquired during the execution of the observation is duplicated in that directory, keeping the same format as in the “daily” directory.

The data is hence stored in two different directories:

- A “daily” directory, accumulating housekeeping telemetry data all day long (e.g. for trend analysis). Each day has its own sub-folder.
- An “obs” directory, recording only observation-specific data and housekeeping in sub-folders containing the camera name also, e.g. **01061\_CSL1\_duvel** is observation 1061 taken in clean room 1 at CSL for the FM#3, i.e. Duvel.



*Figure 2. Timeline illustrating how data is generated and stored. Telemetry for all devices is continuously generated and stored in the daily files. When an observation is started, telemetry is also stored in the 'obs' area under the sub-folder for the current observation. These telemetry are a mirror of the daily telemetry for that observation. Time goes from left to right.*

## 5.3. Image data

Image data are saved for each test execution as FITS files, with the corresponding housekeeping in csv files (see above).

Images are saved in FITS format, as UNSIGNED integers. You will have to cast them to float explicitly in python to avoid negative numbers to be wrongly interpreted.

**Splitting:** The DPU monitors a set of "crucial parameters" (see further) and each time at least one of these changes, a new FITS file will be created, provided the FEE is in full-image mode (or full-image pattern mode). Additionally, a maximum number of extensions is specified in the configuration (TBD). When that number is achieved, the image data are split, i.e. a new FITS file is opened to store the next block of image data.

**Slicing:** In a future version of the software, it will be possible to insert a given command in the test scripts in order to force the creation of a new FITS file. This will allow for a more flexible slicing of the data, i.e. enforce a clearer structure of the data, matching the commanding logic, and will facilitate the data analysis and interpretation.

**Crucial parameters:** As mentioned before, the DPU monitors (for changes in) crucial parameters. These are:

- the first and last row that will be transmitted (`v_start` and `v_end`),
- the last column that will be transmitted (`h_end`),
- the number of rows that will be dumped after the requested number of rows have been transmitted (`rows_final_dump`),
- the readout order of the CCDs (`ccd_readout_order`),
- and the FEE mode (`ccd_mode_config`), which should become/stay full-image mode or full-image pattern mode.



### 5.3.1. Data products

During the camera tests, the DPU will be configured such that a specific part of the E- and/or F-side of the selected CCDs will be transmitted (in the form of SpW packets) for a specified number of cycles (as explained in [Section 12.4.2](#)). The following information will be reconstructed from these SpW packets and stored in FITS files:

- the transmitted image data from the selected side(s) of the selected CCDs, for all cycles,
- the transmitted serial pre-scan data of the selected side(s) of the selected CCDs, for all cycles,
- the transmitted serial over-scan data of the selected side(s) of the selected CCDs, for all cycles,
- the transmitted parallel over-scan data of the selected side(s) of the selected CCDs, for all cycles.

While the SpW packets come in, the individual exposures are stored in individual extensions in FITS files that carry "images" in their name. This type of data arrangement is called a "flat structure". When there's a change in crucial parameters, a new FITS file will be constructed (with "cube" in its name), based on the flat-structure FITS file, in which the exposures are aggregated into cubes. The original, flat-structure file will be removed from the system.

For analysis, only the FITS files with the cubes will be available, and therefore only the structure of these will be discussed in the section below.

In case both sides of a CCD are selected, the image data of both sides will be stitched together before storing it in the FITS file. This will also be done for the parallel over-scan data (if present). For the serial pre-scan and the serial over-scan, the information is stored per CCD side.

### 5.3.2. Internal structure

Each of the extensions (apart from the PRIMARY extension) will occur only once and comprise a 3D data array and a header with the metadata that is specific to that extension. The name of an extension will reflect what type of data product it comprises (e.g. image data of the F-side of CCD2, serial pre-scan data of the E-side of CCD3, etc.).

Although we have provided a set of convenience functions (see below) to extract information from a FITS file without being exposed to the internal details, we want to explain the internal structure in more detail.

The following extensions can be included in the FITS files:

Extension name	Content
PRIMARY	Contains only header information, with metadata that pertains to the whole data product (site name, etc.). This extension always be present.
SPRE_<1/2/3/4>_<E/F>	Serial pre-scan data for the E-/F-side of CCD1/2/3/4. This extension will occur once (with all exposures included), if that side of that CCD was transmitted.
SOVER_<1/2/3/4>_<E/F>	Serial over-scan data for the E-/F-side of CCD1/2/3/4. If this information is transmitted, this extension will occur once.



Extension name	Content
POVER_<1/2/3/4>_<E/F>	Parallel over-scan data for the E-/F-side of CCD1/2/3/4. If this information is transmitted, this extension will occur once.
IMAGE_<1/2/3/4>_<E/F>	Image data for the E-/F-side of CCD1/2/3/4. This extension will occur once (with all exposures included), if that-side of that CCD was transmitted.
WCS-TAB_<1/2/3/4>_<E/F>	Table with a single column (with the name TIME) in which the relative time (in seconds) w.r.t. the first exposure (of any CCD (side)) in the file is listed.

**Example 1:** 4CCDs, full frame, 10 cycles

Relevant FEE parameters (see [Section 12.4.2](#)):

- `ccd_readout = [1, 2, 3, 4];`
- `ccd_side = BOTH;`
- `num_cycles = 10;`
- `row_start = 0;`
- `row_end = 4510 + 30 - 1` (i.e. 4510 rows from the image + 30 rows from the parallel over-scan);

Structure/extensions of the FITS file:

- **PRIMARY:** The primary extension (PrimaryHDU object). This does not contain any transmitted data.
- For CCD1:
  - 1 extension **SPRE\_1\_E**: The serial pre-scan for the E-side of CCD1 (10 frames)  
→ dimensions: 25 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SPRE\_1\_F**: The serial pre-scan for the F-side of CCD1 (10 frames)  
→ dimensions: 25 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SOVER\_1\_E**: The serial over-scan for the E-side of CCD1 (10 frames)  
→ dimensions: 15 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SOVER\_1\_F**: The serial over-scan for the F-side of CCD1 (10 frames)  
→ dimensions: 15 columns x (4510 + 30) rows x 10 layers
  - 1 extension **POVER\_1\_E**: The parallel over-scan for the E-side of CCD1 (10 frames)  
→ dimensions: 2255 columns x 30 rows x 10 layers
  - 1 extension **POVER\_1\_F**: The parallel over-scan for the F-side of CCD1 (10 frames)  
→ dimensions: 2255 columns x 30 rows x 10 layers
  - 1 extension **IMAGE\_1\_E**: The image data for the E-side of CCD1 (10 frames)  
→ dimensions: 2255 columns x 4510 rows x 10 layers
  - 1 extension **IMAGE\_1\_F**: The image data of CCD1 (10 frames)  
→ dimensions: 2255 columns x 4510 rows x 10 layers
- For CCD2:
  - 1 extension **SPRE\_2\_E**: The serial pre-scan for the E-side of CCD2 (10 frames)



- dimensions: 25 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SPRE\_2\_F**: The serial pre-scan for the F-side of CCD2 (10 frames)  
→ dimensions: 25 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SOVER\_2\_E**: The serial over-scan for the E-side of CCD2 (10 frames)  
→ dimensions: 15 columns x (4510 + 30) rows x 10 layers
  - 1 in extension **SOVER\_2\_F**: The serial over-scan for the F-side of CCD2 (10 frames)  
→ dimensions: 15 columns x (4510 + 30) rows x 10 layers
  - 1 extension **POVER\_2\_E**: The parallel over-scan for the E-side of CCD2 (10 frames)  
→ dimensions: 2255 columns x 30 rows x 10 layers
  - 1 extension **POVER\_2\_F**: The parallel over-scan for the F-side of CCD2 (10 frames)  
→ dimensions: 2255 columns x 30 rows x 10 layers
  - 1 extension **IMAGE\_2\_E**: The image data for the E-side of CCD2 (10 frames)  
→ dimensions: 2255 columns x 4510 rows x 10 layers
  - 1 extension **IMAGE\_2\_F**: The image data of CCD2 (10 frames)  
→ dimensions: 2255 columns x 4510 rows x 10 layers
- For CCD3:
- 1 extension **SPRE\_3\_E**: The serial pre-scan for the E-side of CCD3 (10 frames)  
→ dimensions: 25 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SPRE\_3\_F**: The serial pre-scan for the F-side of CCD3 (10 frames)  
→ dimensions: 25 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SOVER\_3\_E**: The serial over-scan for the E-side of CCD3 (10 frames)  
→ dimensions: 15 columns x (4510 + 30) rows x 10 layers
  - 1 in extension **SOVER\_3\_F**: The serial over-scan for the F-side of CCD3 (10 frames)  
→ dimensions: 15 columns x (4510 + 30) rows x 10 layers
  - 1 extension **POVER\_3\_E**: The parallel over-scan for the E-side of CCD3 (10 frames)  
→ dimensions: 2255 columns x 30 rows x 10 layers
  - 1 extension **POVER\_3\_F**: The parallel over-scan for the F-side of CCD3 (10 frames)  
→ dimensions: 2255 columns x 30 rows x 10 layers
  - 1 extension **IMAGE\_3\_E**: The image data for the E-side of CCD3 (10 frames)  
→ dimensions: 2255 columns x 4510 rows x 10 layers
  - 1 extension **IMAGE\_3\_F**: The image data of CCD3 (10 frames)  
→ dimensions: 2255 columns x 4510 rows x 10 layers
- For CCD4:
- 1 extension **SPRE\_4\_E**: The serial pre-scan for the E-side of CCD4 (10 frames)  
→ dimensions: 25 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SPRE\_4\_F**: The serial pre-scan for the F-side of CCD4 (10 frames)  
→ dimensions: 25 columns x (4510 + 30) rows x 10 layers
  - 1 extension **SOVER\_4\_E**: The serial over-scan for the E-side of CCD4 (10 frames)  
→ dimensions: 15 columns x (4510 + 30) rows x 10 layers
  - 1 in extension **SOVER\_4\_F**: The serial over-scan for the F-side of CCD4 (10 frames)



- dimensions: 15 columns x (4510 + 30) rows x 10 layers
- 1 extension **POVER\_4\_E**: The parallel over-scan for the E-side of CCD4 (10 frames)
  - dimensions: 2255 columns x 30 rows x 10 layers
- 1 extension **POVER\_4\_F**: The parallel over-scan for the F-side of CCD4 (10 frames)
  - dimensions: 2255 columns x 30 rows x 10 layers
- 1 extension **IMAGE\_4\_E**: The image data for the E-side of CCD4 (10 frames)
  - dimensions: 2255 columns x 4510 rows x 10 layers
- 1 extension **IMAGE\_4\_F**: The image data of CCD4 (10 frames)
  - dimensions: 2255 columns x 4510 rows x 10 layers

**Example 2:** E-side of CCD2, 100 lines in partial-readout mode, 25 cycles

Relevant FEE parameters (see [Section 12.4.2](#)):

- `ccd_readout = [2, 2, 2, 2];`
- `ccd_side = E;`
- `num_cycles = 25;`
- `row_end = row_start - 100 - 1`

Note that this means that the E-side of CCD2 will be transmitted 4x25 times.

Structure/extensions of the FITS file:

- **PRIMARY**: The primary extension (PrimaryHDU object). This does not contain any transmitted data.
- For CCD2:
  - 1 extension **SPRE\_2\_E**: The serial pre-scan for the E-side of CCD2 (4x25 frames)
    - dimensions: 25 columns x (4510 + 30) rows x 100 layers
  - 1 extension **SOVER\_2\_E**: The serial over-scan for the E-side of CCD2 (4x25 frames)
    - dimensions: 15 columns x (4510 + 30) rows x 100 layers
  - 1 extension **POVER\_2**: The parallel over-scan for the E-side of CCD2 (4x25 frames)
    - dimensions: 4510 columns x 30 rows x 100 layers
  - 1 extension **IMAGE\_2**: The image data of CCD2 (4x25 frames)
    - dimensions: 4510 columns x 4510 rows x 100 layers

### 5.3.3. Inspecting the content

In the test scripts analysis package, i.e. the `fitsfiles.py` in `camtest.analysis.functions`, a number of convenience functions have been implemented to access information in the FITS files without detailed knowledge of the file structure. The remainder of this section will explain how to use these functions to access the different data parts in the FITS files.

#### 5.3.3.1. Overview

The structure of the FITS files can be inspected with the following commands:



```
>>> get_overview(filename)
```

The output looks like to this (here only the E-side of CCD1 has been transmitted, without parallel over-scan):

No.	Name	Ver	Type	Cards	Dimensions	Format
0	PRIMARY	1	PrimaryHDU	5	()	
1	SPRE_1_E	0	ImageHDU	12	(25, 100)	float64
2	SOVER_1_E	0	ImageHDU	12	(15, 100)	float64
3	IMAGE_1	0	ImageHDU	28	(2255, 100)	float64
4	SPRE_1_E	0	ImageHDU	12	(25, 100)	float64
5	SOVER_1_E	0	ImageHDU	12	(15, 100)	float64
6	IMAGE_1	0	ImageHDU	28	(2255, 100)	float64
7	SPRE_1_E	0	ImageHDU	12	(25, 100)	float64
8	SOVER_1_E	0	ImageHDU	12	(15, 100)	float64
9	IMAGE_1	0	ImageHDU	28	(2255, 100)	float64
10	SPRE_1_E	0	ImageHDU	12	(25, 100)	float64
11	SOVER_1_E	0	ImageHDU	12	(15, 100)	float64
12	IMAGE_1	0	ImageHDU	28	(2255, 100)	float64

Alternatively, you can get hold of a list with the extension names with

```
>>> get_ext_names(filename)
```

### 5.3.3.2. Primary header

The primary header can be retrieved as

```
>>> primary_header = get_primary_header(filename)
```

Note that all headers (for the primary header as well as for the other datasets) are returned as `astropy.io.fits.header.Header` objects. To get hold of the value of a specific header key, use:

```
>>> header[key]
```

The relevant keywords in the primary header are:

Keyword	Description
LEVEL	This is present for historical reason and is fixed at 2, which indicates that it is a FITS file in which the data has been arranged into cubes.
V_START	The index of the first row that is transmitted.
V_END	The index of the last row that is transmitted.
H_END	The index of the last column that is transmitted.
ROWS_FINAL_DUMP	The number of rows that are dumped after the requested number of rows is transmitted.



Keyword	Description
TELESCOP	Set to "PLATO".
INSTRUME	The camera ID (as taken from the setup).
SITENAME	The name of the test site at which the data was acquired.
SETUP	The setup ID.
CCD_READOUT_ORDER	String representation of an array with the order inwhich the CCDs will be read out.
CYCLETIME	The image cycle time [s].
READTIME	The time needed to read out the requested part for a single CCD side [s].
OBSID	The observation identifier.
DATE-OBS	The timestamp of the first exposure (of any side of any CCD in the file).
CGSE	Version of the Common EGSE with which the FITS file was produced.

### 5.3.3.3. Images

The image data of a specific exposure (counting starts at zero) of a specific side of a specific CCD can be retrieved as a numpy array, as follows:

```
>>> image_data = get_image_data(filename, ccd_number, ccd_side, exposure_number)
```

You can retrieve the image cubes and header for a specific side of a specific CCD as follows:

```
>>> image_cube_data = get_image_cube_data(filename, ccd_number, ccd_side)
>>> image_cube_header = get_image_cube_header(filename, ccd_number, ccd_side)
```

The relevant keywords in the image header are:

Keyword	Description
NAXIS1	The number of columns in the image area of the CCD, that are transmitted (max 2255).
NAXIS2	The number of rows in the image area of the CCD, that are transmitted (max 4510).
NAXIS3	The number of exposures.
FOCALLEN	The focal length of the telescope [mm].
CTYPE1/CTYPE2	Set to "LINEAR" to indicate that a linear coordinate transformation is used (between pixels and mm), both in the column and row direction.
CTYPE3	Set to "TIMETAB" to indicate that the 3rd axis will be characterised by tabulated (time) values. In the older FITS files, this was set to "TIME-TAB" (consistent with the Greisen & Calabretta papers), but the C code underlying astropy and the traditional FITS viewer need it to be "TIMETAB" instead.



Keyword	Description
CUNIT1/CUNIT2	Set to "MM" to indicate that the focal-plane coordinates are expressed in mm, both in the column and row direction.
CUNIT3	Set to "s" to indicate that the time is expressed in s.
CDELT1/CDELT2	The pixels size [mm], both in the column and row direction.
PS3_0	Set to "WCS-TAB_<1/2/3/4>_<E/F>" to indicate the the 3rd axis is characterised by the values in in the table, that is stored in a extension with that name.
PS3_1	Set to "TIME" to indicate that the 3rd axis is characterised by the values in the table (see PS3_0) in the column with that name.
SITENAME	The name of the test site at which the data was acquired.
EXTNAME	The extension name, following the convention:  IMAGE_<CCD number (1/2/3/4)>_<CCD side (E/F)>
CCD_ID	The CCD number (1/2/3/4).
CROTA2	The orientation angle of the CCD [degrees]. This indicates over which angle the CCD reference frame is rotated w.r.t. the focal-plane reference plane, in counter-clockwise direction.
CD1_1	The product of the pixel size and the cosine of the CCD orientation angle.
CD1_2	The negative product of the pixel size and the sine of the CCD orientation angle.
CD2_1	The product of the pixel size and the sine of the CCD orientation angle.
CD2_2	The product of the pixel size and the cosine of the CCD orientation angle.
CRVAL1	The focal-plane x-coordinate of the CCD origin [mm].
CRVAL2	The focal-plane y-coordinate of the CCD origin [mm].
CRPIX1	The column coordinate of the CCD origin w.r.t. the first transmitted column of the image area.
CRPIX2	The row coordinate of the CCD origin w.r.t. the first transmitted row of the image area.
OBSID	The observation identifier.
DATE-OBS	The timestamp of the start of the data acquisition of the first exposure.

#### 5.3.3.4. Parallel over-scan

To get hold of the data (as a numpy array) of a parallel over-scan (if present) for a specific exposure (counting starts at 0) of a specific side of a specific CCD, execute the following command:

```
>>> parallel_overscan_data = get_parallel_overscan_data(filename, ccd_number, ccd_SIDE, exposure_number)
```

You can retrieve the parallel over-scan data and header for a specific side of a specific CCD as follows:



```
>>> parallel_overscan_cube_data = get_parallel_overscan_cube_data(filename, ccd_number, ccd_side)
>>> parallel_overscan_cube_header = get_parallel_overscan_cube_header(filename, ccd_number, ccd_side)
```

The relevant keywords in the parallel over-scan header are:

Keyword	Description
NAXIS1	The number of columns in the parallel over-scan, that are transmitted.
NAXIS2	The number of rows in the parallel over-scan of the CCD, that are transmitted.
NAXIS3	The number of exposures.
CTYPE3	Set to "TIMETAB" to indicate that the 3rd axis will be characterised by tabulated (time) values. In the older FITS files, this was set to "TIME-TAB" (consistent with the Greisen & Calabretta papers), but the C code underlying astropy and the traditional FITS viewer need it to be "TIMETAB" instead.
CUNIT3	Set to "s" to indicate that the time is expressed in s.
PS3_0	Set to "WCS-TAB_<1/2/3/4>_<E/F>" to indicate the the 3rd axis is characterised by the values in in the table, that is stored in a extension with that name.
PS3_1	Set to "TIME" to indicate that the 3rd axis is characterised by the values in the table (see PS3_0) in the column with that name.
FOCALLEN	The focal length of the telescope [mm].
SITENAME	The name of the test site at which the data was acquired.
EXTNAME	The extension name, following the convention:  IMAGE_<CCD number (1/2/3/4)>_<CCD side (E/F)>
CCD_ID	The CCD number (1/2/3/4).
OBSID	The observation identifier.
DATE-OBS	The timestamp of the start of the data acquisition of the first exposure.

### 5.3.3.5. Serial pre-scan

To get hold of the data (as a numpy array) of a serial pre-scan for a specific exposure (counting starts at 0) of a specific side of a specific CCD, execute the following command:

```
>>> serial_prescan_data = get_serial_prescan_data(filename, ccd_number, ccd_side, exposure_number)
```

You can retrieve the serial pre-scan data and header for a specific side of a specific CCD as follows:

```
>>> serial_prescan_cube_data = get_serial_prescan_cube_data(filename, ccd_number, ccd_side)
>>> serial_prescan_cube_header = get_serial_prescan_cube_header(filename, ccd_number, ccd_side)
```

The relevant keywords in the serial pre-scan header are:



Keyword	Description
NAXIS1	The number of columns in the serial pre-scan (fixed at 25).
NAXIS2	The number of rows in the serial pre-scan, that are transmitted.
NAXIS3	The number of exposures.
CTYPE3	Set to "TIMETAB" to indicate that the 3rd axis will be characterised by tabulated (time) values. In the older FITS files, this was set to "TIME-TAB" (consistent with the Greisen & Calabretta papers), but the C code underlying astropy and the traditional FITS viewer need it to be "TIMETAB" instead.
CUNIT3	Set to "s" to indicate that the time is expressed in s.
PS3_0	Set to "WCS-TAB_<1/2/3/4>_<E/F>" to indicate the the 3rd axis is characterised by the values in in the table, that is stored in a extension with that name.
PS3_1	Set to "TIME" to indicate that the 3rd axis is characterised by the values in the table (see PS3_0) in the column with that name.
SITENAME	The name of the test site at which the data was acquired.
EXTNAME	The extension name, following the convention:  SPRE_<CCD number (1/2/3/4)>_<CCD side (E/F)>
CCD_ID	The CCD number (1/2/3/4).
OBSID	The observation identifier.
DATE-OBS	The timestamp of the start of the data acquisition of the first exposure.

### 5.3.3.6. Serial over-scan

To get hold of the data (as a numpy array) of a serial over-scan for a specific exposure (counting starts at 0) of a specific side of a specific CCD, execute the following command

```
>>> serial_overscan_data = get_serial_overscan_data(filename, ccd_number, ccd_side, exposure_number)
```

You can retrieve the serial over-scan data and header for a specific side of a specific CCD as follows:

```
>>> serial_overscan_cube_data = get_serial_overscan_cube_data(filename, ccd_number, ccd_side)
>>> serial_overscan_cube_header = get_serial_overscan_cube_header(filename, ccd_number, ccd_side)
```

The relevant keywords in the serial over-scan header are:

Keyword	Description
NAXIS1	The number of columns in the serial over-scan.
NAXIS2	The number of rows in the serial over-scan, that are transmitted.
NAXIS3	The number of exposures.

Keyword	Description
CTYPE3	Set to "TIMETAB" to indicate that the 3rd axis will be characterised by tabulated (time) values. In the older FITS files, this was set to "TIME-TAB" (consistent with the Greisen & Calabretta papers), but the C code underlying astropy and the traditional FITS viewer need it to be "TIMETAB" instead.
CUNIT3	Set to "s" to indicate that the time is expressed in s.
PS3_0	Set to "WCS-TAB_<1/2/3/4>_<E/F>" to indicate the the 3rd axis is characterised by the values in in the table, that is stored in an extension with that name.
PS3_1	Set to "TIME" to indicate that the 3rd axis is characterised by the values in the table (see PS3_0) in the column with that name.
SITENAME	The name of the test site at which the data was acquired.
EXTNAME	The extension name, following the convention:  SOVER_<CCD number (1/2/3/4)>_<CCD side (E/F)>
CCD_ID	The CCD number (1/2/3/4).
OBSID	The observation identifier.
DATE-OBS	The timestamp of the start of the data acquisition of the first exposure.

### 5.3.3.7. Time

To get hold of the **relative** time (as a numpy array) in seconds w.r.t. the first exposure (of any side of any CCD in the file) for the exposures for a specific side of a specific CCD, execute the following command:

```
>>> from camtest.analysis.functions.fitsfiles import get_relative_time
>>> relative_time = get_relative_time(filename, ccd_number, ccd_side)
>>> relative_time = get_relative_time(hdu_list_object, ccd_number, ccd_side)
```

The **absolute** time (as a numpy array) in seconds since epoch 1958, for the exposures for a specific side of a specific CCD, can be retrieved as follows:

```
>>> from camtest.analysis.functions.fitsfiles import get_absolute_time
>>> absolute_time = get_absolute_time(filename, ccd_number, ccd_side)
>>> absolute_time = get_absolute_time(hdu_list_object, ccd_number, ccd_side)
```

### 5.3.3.8. Focal-plane coordinates

To convert image pixel coordinates (row, column) for a given side of a given CCD to focal-plane coordinates (x, y), in mm, execute:

```
>>> from camtest.analysis.functions.fitsfiles import get_fp_coordinates
```



```
>>> x, y = get_fp_coordinates(filename, ccd_number, ccd_side, row, column)
>>> x, y = get_fp_coordinates(hdu_list_object, ccd_number, ccd_side, row, column)
```

### 5.3.3.9. Astropy WCS objects

It is possible to load the headers into an astropy WCS object, which can be used for coordinate conversions:

```
>>> from astropy.io import fits
>>> from astropy.wcs import WCS
>>> with fits.open(<filename>) as hdul:
...     header = hdul[<extension name>].header
>>> wcs = WCS(header)
```

For the older FITS files, you may get this error message (see #1714):

```
ERROR: ValueError: ERROR 5 in wcsset() at line 2352 of file cextern/wcslib/C/wcs.c: Invalid parameter value.
ERROR 3 in tabset() at line 747 of file cextern/wcslib/C/tab.c: Invalid tabular parameters: Each element of K must
be positive, got 0.
```

This can be circumvented by correcting the value for the CTYPE3 keyword (before creating the WCS object):

```
>>> header["CTYPE3"] = "TIMETAB"
```

## 5.4. Telecommand history

The complete telecommand history is not yet saved but can be reconstructed from the setup\_id (see below) and the obsid-table. The [obsid-table.txt](#) is a text file located at the root of the data storage location. The file contains one entry per observation, associating

- the main parameters of the observation, obsid and site
- the time of execution
- the setup\_id active at execution time (contains the version number of the plato-test-script on the operational server)
- The building block name
- All parameter names and values passed to the execute command

The following lines are examples taken from the obsid-table.txt file at CSL:

```
01108 CSL1 00084 2023-06-07T08:56:22.832+0000 cam_single_cube_int_sync(theta="8.3", phi="12.0", num_cycles="5",
exposure_time="0.2", n_rows="1000", attenuation="0.00413")
01109 CSL1 00084 2023-06-07T08:57:30.003+0000 cam_single_cube_int_sync(theta="8.3", phi="12.0", num_cycles="5",
exposure_time="0.2", n_rows="1000", attenuation="0.00413")
01110 CSL1 00084 2023-06-07T09:02:16.140+0000
check_and_move_relative_user(csamodel="egse.coordinates.csamodel.CSLReferenceFrameModel", translation="[0, 0,
0.01]", rotation="[0, 0, 0]", setup="egse.setup.Setup", verbose="True")
01111 CSL1 00084 2023-06-07T09:03:09.426+0000 cam_single_cube_int_sync(theta="8.3", phi="12.0", num_cycles="5",
```



```
exposure_time="0.2", n_rows="1000", attenuation="0.00413")
01112 CSL1 00084 2023-06-07T09:06:40.668+0000
check_and_move_relative_user(cslmodel="egse.coordinates.cslmodel.CSLReferenceFrameModel", translation="[0, 0, 0.005]", rotation="[0, 0, 0]", setup="egse.setup.Setup", verbose="True")
01113 CSL1 00084 2023-06-07T09:08:28.446+0000 cam_single_cube_int_sync(theta="8.3", phi="12.0", num_cycles="5",
exposure_time="0.2", n_rows="1000", attenuation="0.00413")
01114 CSL1 00084 2023-06-07T09:29:58.225+0000 cam_single_cube_int_sync(theta="8.3", phi="12.0", num_cycles="5",
exposure_time="0.2", n_rows="1000", attenuation="0.00413")
01115 CSL1 00084 2023-06-07T09:37:08.043+0000 cam_single_cube_int_sync(theta="8.3", phi="12.0", num_cycles="5",
exposure_time="0.2", n_rows="1000", attenuation="0.00413")
01116 CSL1 00084 2023-06-07T09:40:39.185+0000 cam_single_cube_int_sync(theta="8.3", phi="12.0", num_cycles="5",
exposure_time="0.2", n_rows="1000", attenuation="0.00413")
```

The **obsid-table.txt** file is explained in more detail in the [Interface Control Document \(ICD\) \[RD-04\]](#).



# 6. Configuration and Setups

The complete documentation on the EGSE configuration and on the concept of Setup can be found in the [developer manual](#).

Setup is the concept we attach to the entity encapsulating the entire set of

- Identifiers and configuration items
- Calibration values and calibration files

Necessary to

- Describe the test-environment and the item under test
- operate the test

## 6.1. Example Setup file

A snippet of a setup file is shown below. It shows the tree structure of the YAML file. At the top level are the main components: gse (devices), camera (fpa, tou, fee), telemetry, etc. The snippet only shows part of the Setup. Under the **gse** branch we have **hexapod** and **stages** and many more that are not shown. All ground equipment that is part of the test setup shall have an entry under the **gse** branch. The information that shall go into these entries is device identification, calibration information, specific device settings, etc.

Everything that is connected to the Camera, i.e. SUT, shall go under the **camera** branch. This is e.g. identifiers for the TOU, FEE, FPA, DPU, CCD, ..., calibration information, defaults, numbering, avoidance information, etc.

**NOTE**

The gse branch is also used by the process manager to determine which devices are part of the Setup and for which devices it should present a status LED and a start/stop button.

The users are allowed to modify or add items and branches to the setup and save a new version of it. As an example, the stages branch in this example contains calibration values defining metrology of the rotation and translation stages used to position the light beam at CSL.

*Snippet of a Setup file from CSL*

```

NavigableDict
├── site_id: CSL2
├── position: 2
└── gse
    ├── hexapod
    │   ├── device: class//egse.hexapod.symetrie.puna.PunaProxy
    │   ├── device_name: Symetrie Puna Hexapod
    │   ├── ID: 2B
    │   ├── time_request_granularity: 0.1
    │   ├── CID: 603382
    │   └── label: 172543 - PUNA
    └── stages
        ├── ID: 1
        └── BIG_ROTATION_STAGE_ID: 420-20913

```



```

    |   |
    |   |   └── SMALL_ROTATION_STAGE_ID: 409-10661
    |   |   └── TRANSLATION_STAGE_ID: 5101.30-943
    |   |   └── device: class//egse.stages.huber.smc9300.HuberSMC9300Proxy
    |   |   └── device_name: Huber SMC9300 Stages
    |   |   └── calibration
    |   |       ├── height_collimated_beam: 513.9
    |   |       ├── offset_phi: 0.4965
    |   |       ├── offset_alpha: 0.0
    |   |       ├── offset_delta_x: 96.884
    |   |       ├── phi_correction_coefficients: [-0.0049, 0.0003]
    |   |       ├── alpha_correction_coefficients: [0.0856, -0.5]
    |   |       └── delta_x_correction_coefficients: [-0.1078, 0.2674, -0.0059]
    |   |   └── big_rotation_stage
    |   |       ├── avoidance: 3.0
    |   |       ├── hardstop: 179.316
    |   |       └── default_speed: 15000
    |   |   └── small_rotation_stage
    |   |       └── default_speed: 15000
    |   |   └── translation_stage
    |   |       └── default_speed: 15000
    |
    ...  

    |
    |   └── camera
    |       ├── TOU
    |       |   └── ID: BA-N1-11130000-FM-01
    |       └── fpa
    |           ├── avoidance
    |           |   ├── clearance_xy: 2.0
    |           |   ├── clearance_z: 2.0
    |           |   ├── vertices_nb: 60
    |           |   └── vertices_radius: 100.0
    |           |   └── ID: N-FPA-11200000-FM-01
    |           |   └── max_offset: 20
    |           └── dpu
    |               ├── device: class//egse.dpu.DPUProxy
    |               └── device_name: DPU
    |
    ...  

    |
    |   └── telemetry
    |       ├── dictionary: pandas//.../common/telemetry/tm-dictionary-brigand.csv
    |       └── separator: ;
    |   └── sensor_calibration
    |       └── callendar_van_dusen
    |           └── EN60751
    |               ├── A: 0.0039083
    |               ├── B: -5.775e-07
    |               └── C: -4.183e-12
    |   └── history
    |       ├── 0: Initial zero Setup for CSL2
    |       ├── 1: Copy of CSL setup 97 (last EM setup)
    |       └── 2: Removed TCS block

```

You can check the version of the Setup with the following command:

```
>>> print(setup.get_id())
00068
```

The number printed on your system will be different.



## 6.2. Available Setups

Setups will be available in the form of YAML files that are stored in the `plato-cgse-conf` repository and are located (probably) at `~/git` where the repos are kept.

**WARNING**

Do never edit these YAML files directly since they are maintained through changes on the operational machine. The configuration manager on the egse-server machine manages these Setups and brings them under configuration control in GitHub automatically upon submitting a new Setup.

Browsing through the available setups can either be done in Python or via a GUI.

### 6.2.1. Browsing the Setups in Python

To get a list of the setups that are available in the system, execute the following command (to be imported from `camtest`):

```
>>> list_setups()
```

This will return a list of (setup identifier, site identifier) pairs, e.g.

```
('00037', 'CSL2', 'Use v2 of N-FEE sensor calibration for Chimay (#293)', 'brigand')
('00038', 'CSL2', 'Incl. nominal ranges for power consumption checks (#312)', 'brigand')
('00039', 'CSL2', 'Updated N-cam voltages for the AEU PSU (#315)', 'brigand')
('00040', 'CSL2', 'New CSLReferenceFrameModel [csl_model_from_file]', 'brigand')
('00041', 'CSL2', 'Putting back N-cam voltages for the AEU PSU', 'brigand')
('00042', 'CSL2', 'updated stages calibration (beam height and phi correction)', 'brigand')
('00043', 'CSL2', 'Updated AEU voltages + voltage/current protection values (#324)', 'brigand')
('00044', 'CSL2', 'updated translation stage zero position', 'brigand')
('00045', 'CSL2', 'Update AEU configuration according to - NRB NCR-CSL-0036 disposition, email by Yves on
24/05/2023', 'brigand')
```

The `list_setups()` command also allows you to filter the results, by using any of the keywords inside the setups. For instance, to list all the setups related to the STM version of the TOU, tested with the hexapod No 1 at CSL, you would type

```
>>> list_setups(camera__ID="achel") ①
```

① Note that the double underscore “`__`” is used to navigate the Setup. All parameters passed will be joined with a logical AND.

```
list_setups(camera__ID="achel")
('00004', 'CSL2', 'Copy camera and telemetry info for achel from CSL1 setup 38', 'achel')
('00005', 'CSL2', 'Incl. sensor calibration', 'achel')
('00006', 'CSL2', 'Updated hexapod ID', 'achel')
('00007', 'CSL2', 'Updated device name for DAQ6510 (#235)', 'achel')
('00008', 'CSL2', 'Using short sync pulses of 200ms (instead of 150ms)', 'achel')
('00009', 'CSL2', 'Copy camera and telemetry info for achel from CSL1 setup 45', 'achel')
('00010', 'CSL2', 'Incl. MGSE calibration coefficients (#255)', 'achel')
('00011', 'CSL2', 'New CSLReferenceFrameModel [csl_model_from_file]', 'achel')
('00012', 'CSL2', 'Changed offset_phi for validation purposes', 'achel')
```



```
('00013', 'CSL2', 'Recalibration of the SMA (#258)', 'achel')
('00014', 'CSL2', 'Updated reference Hartmann positions (#254)', 'achel')
('00015', 'CSL2', 'fixed alpha correction coefficients', 'achel')
('00016', 'CSL2', 'Incl. reference_full_76 (taken from CSL1 setup 46)', 'achel')
('00017', 'CSL2', 'New CSLReferenceFrameModel [csl_model_from_file]', 'achel')
('00018', 'CSL2', 'Copy camera and telemetry info for achel from CSL1 setup 47', 'achel')
('00019', 'CSL2', 'Updated x, y measured positions w.r.t. LDO input (#266)', 'achel')
```

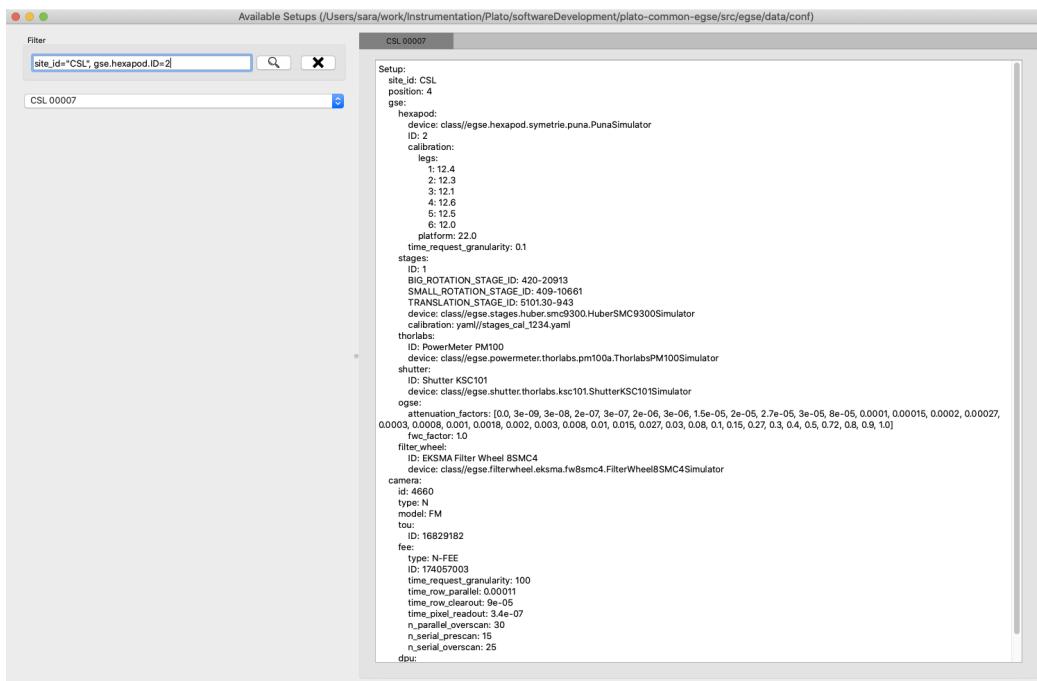
## 6.2.2. Using the Setup GUI

To open the GUI to inspect all available setups, type the following command:

```
$ setup_ui
```

This will fire up a window as shown in [Figure 3](#). The directory that is mentioned in the window title, is the one where the available setups are located.

A text field on the left-hand side allows you to filter the setups, similar to the arguments of the `list_setups` command from [Section 6.2.1](#). You can navigate through the tree both with the '`_`' and the dot notation. For the available setups that pass the filtering, the site and setup identifier will appear in the drop-down menu, after either hitting the return key in the filter text field or by pressing the search button next to it.



*Figure 3. Screenshot of the GUI to inspect all available setups.*

## 6.3. Loading a Setup

After inspection of the available setups, a specific setup can be loaded, based on the identifier.

### IMPORTANT

*loading* a Setup means to load it in the system such that it then becomes the reference for the system configuration. This means **it impacts the GlobalState and the ConfigurationManager!** It is different from getting a copy of a Setup



as a variable in a python script (see below).

Ideally, the Setup will be loaded **one single time** at the start of a test phase, with a setup reflecting the HW present in the test environment. **The preferred way to do so is via the setup GUI.** That can be launched via

```
$ setup_ui
```

In python:

```
>>> from camtest import load_setup
>>> setup = load_setup(7)
```

The above command will load the Setup in the configuration manager and also in the Python console you used to execute the `load_setup()` command.

#### **WARNING**

Loading or submitting a Setup will have an effect on the configuration manager. The Setup is however NOT automatically propagated to other components in the system like the device drivers or the synoptics manager. Therefore, always check if you have the right Setup available, especially in the QtConsole.

You can check which Setup is loaded in the configuration manager with

```
$ cm_cs status
Configuration manager:
  Status: active
  Site ID: CSL2
  No observation running
  Setup loaded: 00068
  Hostname: 192.168.0.163
  Monitoring port: 6001
  Commanding port: 6000
  Service port: 6002
  Listeners: Storage CS
```

Also the process manager (`pm_ui`) and the Operator GUI (e.g. `csl_ui`) indicate which Setup is loaded, but keep in mind that this is only true for that particular process and is not necessarily propagated to all processes running in this distributed environment.

## 6.4. Inspecting, accessing, and modifying a Setup

First, make sure a Setup is already loaded in the system, and that you have a variable attached to a setup in your Python session. Here we call it `setup`.

You can get a Setup with the following command

```
>>> from camtest import get_setup
>>> setup = get_setup(7)
```



This will read the content of Setup "00007" for the site you are currently at and assign it to a variable called `setup`.

**NOTE**

This has no effect on the system configuration (the ConfigurationManager will not know about it, and the GlobalState won't be affected).

### 6.4.1. Content of the setup

To print the entire content of the setup:

```
>>> print(setup)
```

The Setups, as well as all of their branches are “navigable dictionaries”. In practice that means that they have a tree structure, and every part of the tree can be accessed with a simple syntax, using dot notation (in contrast to using a double underscore (`_`) when filtering the available Setups).

### 6.4.2. Inspect a given branch or leaf

You can inspect any branch or leaf of the Setup by navigating the Setup and printing the result:

```
>>> print(setup.branch.subbranch.leaf)
```

For instance, printing the hexapod configuration at CSL:

```
>>> print(setup.gse.hexapod)
NavigableDict
└── device: class//egse.hexapod.symetrie.puna.PunaProxy
    ├── device_name: Symetrie Puna Hexapod
    ├── ID: 2B
    ├── time_request_granularity: 0.1
    ├── CID: 603382
    └── label: 172543 - PUNA
```

### 6.4.3. Modify an entry

Any Setup entry can be assigned to with a simple assignment statement.

```
>>> setup.branch.subbranch.leaf = object
```

For instance:

```
>>> setup.camera.fpa.avoidance.clearance_xy = 3
```



## 6.4.4. Add a new entry

When you want to replace a complete sub-branch in the Setup, use a dictionary.

```
>>> setup.branch.subbranch = []
>>> setup.branch.subbranch.leaf = object
```

for instance, to introduce the fpa subbranch in the example file above, one would write:

```
>>> setup.camera.fpa = {}
>>> setup.camera.fpa.ID = "STM"
>>> setup.camera.fpa.avoidance = {}
>>> setup.camera.fpa.avoidance.clearance_xy = 3
```

The above can be simplified by adding a predefined dictionary to the Setup:

```
>>> setup.camera.fpa = {"ID": "STM", "avoidance": {"clearance_xy": 3}}
```

## 6.5. Saving a new setup

The setups are stored under configuration control in the plato-cgse-conf repository. The EGSE is taking care of the interface with that repository when a user is submitting a new setup, with no additional action necessary than:

```
>>> response = submit_setup(setup, description="A sensible description of this setup")
>>> if isinstance(response, Setup):
...     setup = response
... else:
...     print(response)
```

The new setup receives a unique `setup_id`, and a new entry is created in the list of setups. The new `Setup` is then loaded and made active in the configuration manager. When the `Setup` is processed by the system, brought under configuration control and no errors occurred, the new `Setup` is returned and will be assigned to the `setup` variable. In case of an error, the response contains information on the cause of the Failure and is printed.

**NOTE**

Do not directly catch the returned `Setup` in the `setup` variable, because you will lose your modified `Setup` in case of an error.

**NOTE**

the `description` is mandatory. Setups keep track of their history, so we strongly encourage to provide concise but accurate descriptions each time this command is used.

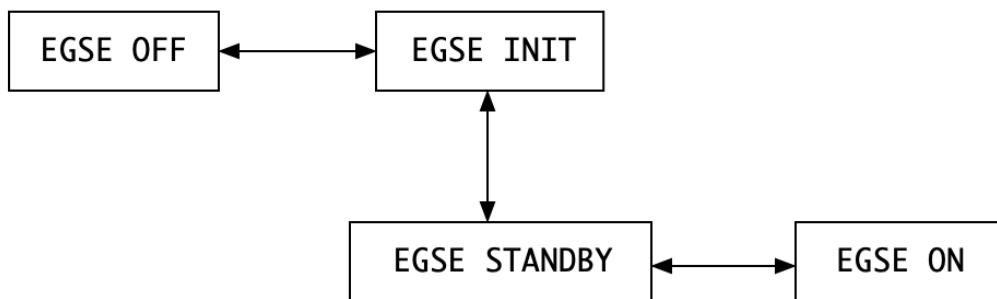
# 7. Common-EGSE startup, shutdown, sleep

Intended readers: site-operator, test-operator

## 7.1. EGSE States

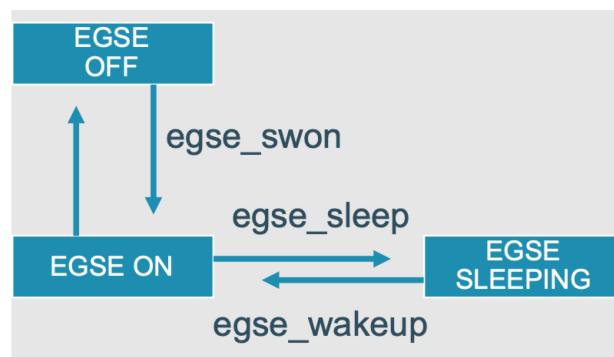
The following states have been defined for the EGSE system, as depicted in [Figure 4](#):

- EGSE OFF: all systems are powered off;
- EGSE INIT: egse-server is booted, user logged in as plato-user;
- EGSE STANDBY: critical housekeeping is being acquired, a sub-set of critical functions is available (e.g. temperature and safety monitoring of the test-environment and test-article);
- EGSE ON: the system and hardware are ready to receive commands.



*Figure 4. Transitions between the EGSE states.*

The transitions between these states are all handled by the Process Manager GUI. Under the hood, the Process Manager queries the setup (from the Configuration Manager) for information on the relevant processes. A distinction is made between core and devices processes.



## 7.2. Core & Device Processes

The following processes are considered as Common-EGSE **core processes** (depicted in the left panel of [Figure 5](#)) and should always be running:

- the **Storage**, which is responsible for archiving of housekeeping and image data (see [Chapter 5](#));
- the **Configuration Manager**, which manages of the configuration and setup of the system (see



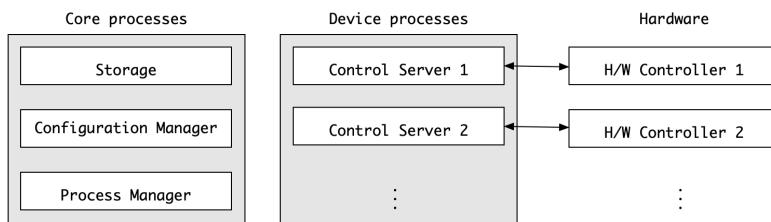
[Chapter 6](#));

- the **Process Manager**, which can be used to start device processes and monitor their status (typically via the corresponding GUI; see further),
- the **Logger**, which collects all log messages from the different components of the Common-EGSE and test scripts. The log messages are saved in a log file at a location denoted by the environment variable PLATO\_LOG\_FILE\_LOCATION.
- the **Synoptics Manager**, which handles all generic and device independent housekeeping.

These processes are started automatically when the egse-server is booted. They cannot be re-started nor shut down via the Process Manager GUI; they can only be monitored there.

The **device processes** (depicted in the middle panel of [Figure 5](#)) are the so-called Control Servers that talk to the hardware Controllers (depicted in the right panel of [Figure 5](#)), for commanding and monitoring the devices.

These processes can be (re-)started and shut down individually from the Process Manager GUI. They will be running on the same machine as the Process Manager itself, which is the egse-server during normal operations.



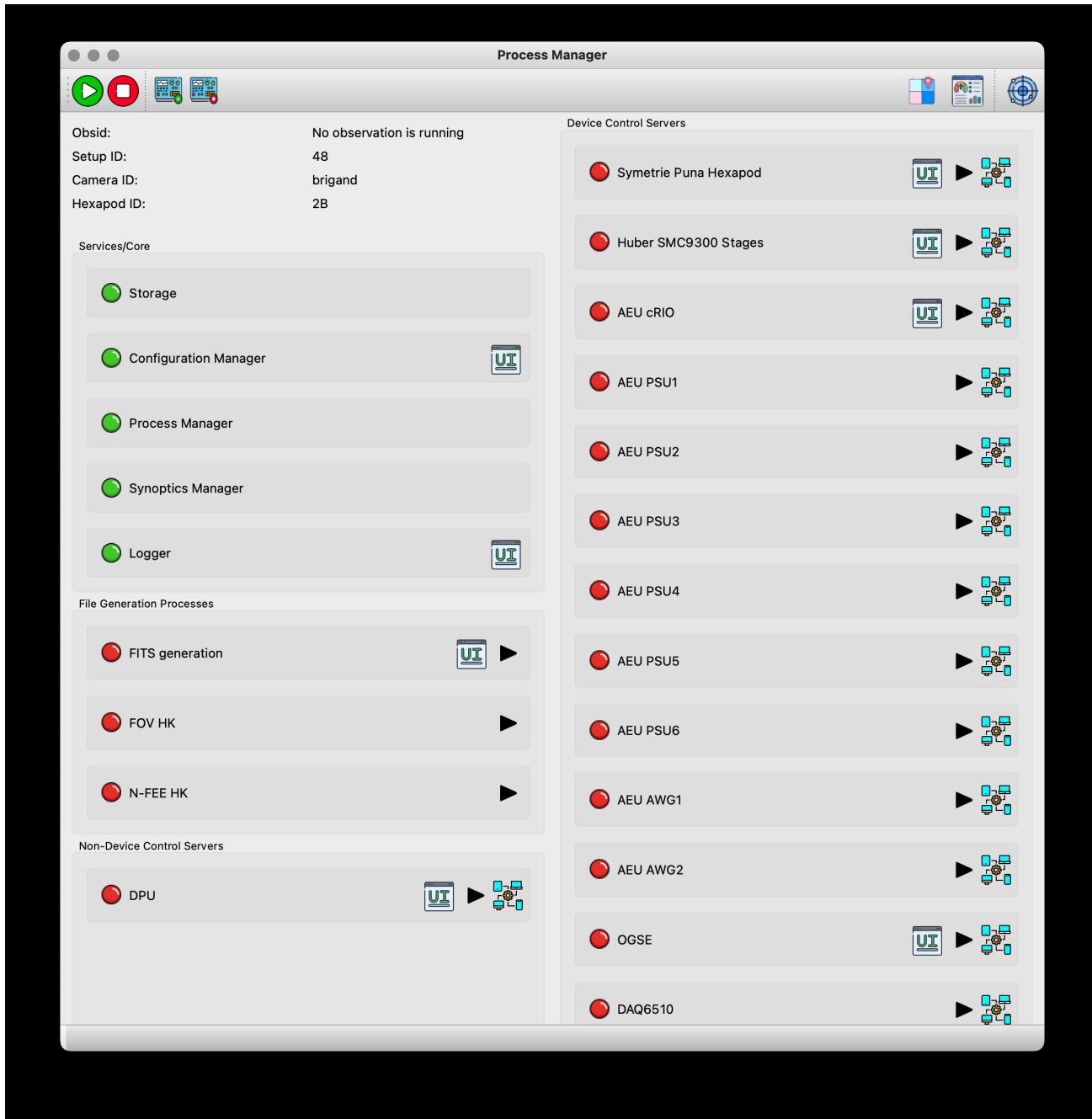
*Figure 5. Core and devices processes. The latter are Control Servers that talk to the hardware Controllers, for commanding and monitoring the devices.*

## 7.3. Process Manager GUI

A desktop icon will be provided to start the Process Manager GUI. Alternatively, it can be started from the terminal command line with the following command:

```
$ pm_ui
```

This will fire up the GUI shown in [Figure 6](#).



*Figure 6. The screenshot of the Process Manager GUI. The colour of the leds (green/orange/red) indicates what the status is of the core processes and the device processes in the current Setup. Devices processes can be started (either in operational or in simulator mode) or shut down by pressing the corresponding play/stop buttons. If a GUI has been implemented for a process, it can be opened (in a separate window) by pressing the corresponding "UI" button.*

On the top left, an overview of the core services will be given. When a (new) Setup is loaded (see Section 6.3]), the content of the right part (with the device control servers) will be updated. Only the processes for the devices that are included in the current Setup will be shown.

A led in front of the process name will give you a quick impression of the state of the process. It can have the following colour for device processes:

- **green:** the process is running and a connection with the hardware Controller has been established;



- **orange**: the process is running, but the connection with the hardware Controller could not be established or was lost;
- **red**: the process is not yet or no longer running or the process hangs for some undefined reason.

For the *core processes*, the led should always be green (indicating the process is running). In case one of them turns red, the corresponding core process is no longer alive, and the system may have to be re-started. Consult your site-operator in this case.

For the processes for which a "GUI" button is present, a GUI for the corresponding device can be opened in a separate window by pressing this button. A GUI for any device can be opened only once.

In contrast to the core processes, the device processes can be started from the Process Manager GUI, either in operational mode (when a hardware Controller is available) or in simulator mode (when no hardware Controller is available; for testing purposes).

### 7.3.1. EGSE OFF □ EGSE INIT

When booting the egse-server, the core processes (i.e. Storage, Configuration Manager, Process Manager, Synoptics Manager, and Logger) will be started automatically. If the Process Manager GUI is started at this point and no Setup is loaded, the area with the device processes will be empty, and the five core processes will show a green LED in the GUI.

Note that the Process Manager GUI can only be used to monitor the core processes, not to (re-)start them. Should any of the LEDs for the core processes turn red (at any point), more detailed inspection of the system will be needed (and the system may even have to be re-booted).

Any GUI for these processes can be started by pressing the GUI button. Note that although the core processes are running on the egse-server, GUIs that are launched from the process manager will be started on your local machine, desktop client.

### 7.3.2. EGSE INIT □ EGSE STANDBY

After having brought the Common-EGSE into its INIT state, a Setup must be loaded (see [Section 6.3](#)). This comprises (amongst others) all devices that are currently relevant to you. Once a Setup is loaded, the area with the device processes will be updated in the process manager GUI.

To start the critical device processes, click on the play/stop button for these devices. If all goes well, the LEDs for these processes will turn green in the GUI.

To shut down these critical processes (to return to INIT mode), the play/stop button of these devices must be pressed a second time.

Any GUI for these processes can be started by pressing the GUI button. The GUI will start on your local machine and can only be started once. Starting and stopping a GUI for a device has no effect on the state of that device.

### 7.3.3. EGSE STANDBY □ EGSE ON

In STANDBY mode, only the critical processes will be running. The other device processes can be started by pressing their play/stop button, bringing the Common-EGSE into ON mode. To return to



STANDBY mode, press the buttons again.

Any GUI for these processes can be started by pressing the GUI button.



# 8. Utility functions

## 8.1. Logging

Errors, warnings, general debugging or any useful information can be logged using the logging infrastructure. The logging infrastructure will take care that the messages you generate are stored in the log files associated with the measurement. It also takes care of visualising the messages to the operator.

Log levels: CRITICAL, ERROR, WARNING, INFO, DEBUG

To log a message in your script:

```
>>> from camtest import camtest_logger
>>> camtest_logger.info("Starting performance verification test")
2023-06-15 17:10:30,857:           IPython:  INFO:  1:camtest      :Starting performance verification
test
```

This will print a log message in the REPL and send a log record with the message to the Logger. Messages for any of the above levels will end up in the Common-EGSE log file, but only messages of level INFO and above will also be printed in the terminal. So, if you need to check any debugging messages, make sure you check the log file at [\\$PLATO\\_LOG\\_FILE\\_LOCATION](#).

## 8.2. Handling Errors

Inevitably, the code other developers write will generate error and exceptions. Whenever you expect an exception and you know what to do with it or how to handle it, catch it with the following construct:

```
try:
    call_a_function_from_another_dev(...)
except ZeroDivisionError:
    # do something to recover here
```

If you cannot handle the error, just let it pass to the next level in your script and eventually to the top building\_block.

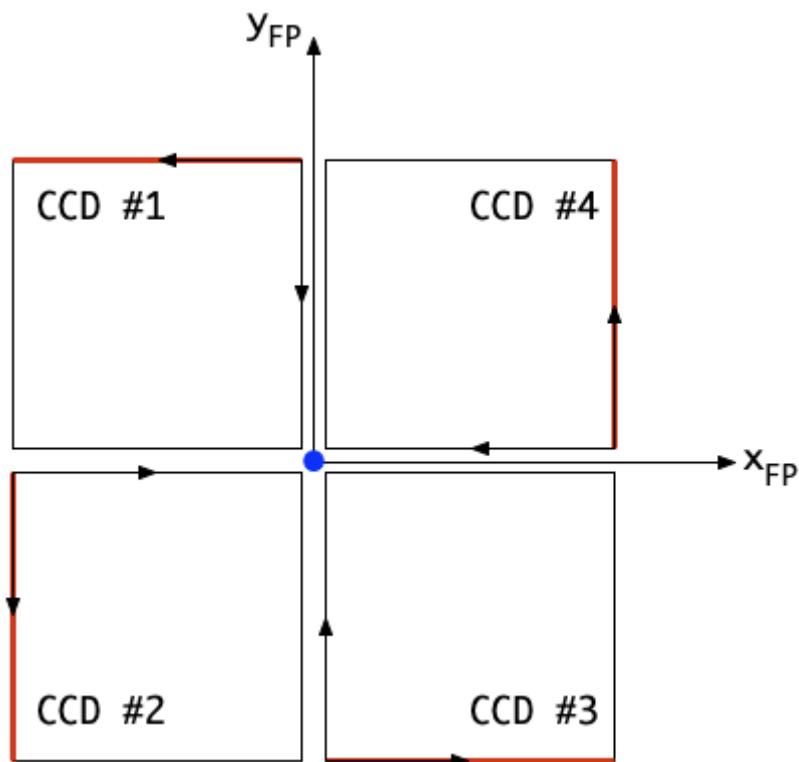
There is a fully detailed description of exception handling in the on-line Common-EGSE documentation in the section [Exception Handling](#) of the developer manual.

## 8.3. Coordinate transformations

[Figure 7](#) below shows the schematic overview of the focal plane with the four CCDs. The focal-plane reference frame ( $x_{FP}$ ,  $y_{FP}$ ,  $z_{FP}$ ) is associated with the focal plane of one camera. Its origin is in the middle of the four CCDs (indicated by the blue dot in the figure) and the  $z_{FP}$ -axis coincides with the optical axis (pointing towards the reader). The coordinates in the ( $x_{FP}$ ,  $y_{FP}$ ) reference frame are typically expressed in mm.

Each of the four CCDs has its own CCD reference frame ( $x_{\text{CCD}}$ ,  $y_{\text{CCD}}$ ) assigned to it (indicated with the small arrows in the figure). When keeping the readout register of a CCD at the bottom, the origin of the associated CCD reference frame is at the lower left corner of the CCD. The parallel charge transfer happens in the negative  $y_{\text{CCD}}$  direction; the serial charge transfer in the readout register happens in the negative  $x_{\text{CCD}}$  direction for the left detector half and in the positive  $x_{\text{CCD}}$  direction for the right detector half. The coordinates in the  $(x_{\text{CCD}}, y_{\text{CCD}})$  reference frame are typically expressed in pixels (where a pixel measures  $18\mu\text{m}$  in both directions)

- Definition of important coordinate systems
- Field of view position (degrees) to CCD number and pixel coordinates
- CCD number and pixel coordinates to field of view position (degrees)



*Figure 7. Schematic overview of the focal plane of one camera, with the four CCDs. The blue dot in the middle of the CCDs is the origin of the focal-plane reference frame ( $x_{\text{FP}}$ ,  $y_{\text{FP}}$ ) and denotes the location where the optical axis intersects with the focal-plane. The readout registers of the CCDs are marked in red and the associated CCD reference frames are indicated with the small arrows.*

Alternatively, the position in the focal plane can be expressed in field angles ( $\theta, \phi$ ), as shown in [Figure 8](#). The angular distance to the optical axis (marked with a blue dot in the figure) is denoted by angle  $\theta$ , whereas  $\phi$  represents the in-field angle, measured in counter-clockwise direction from the  $x_{\text{FP}}$  axis.

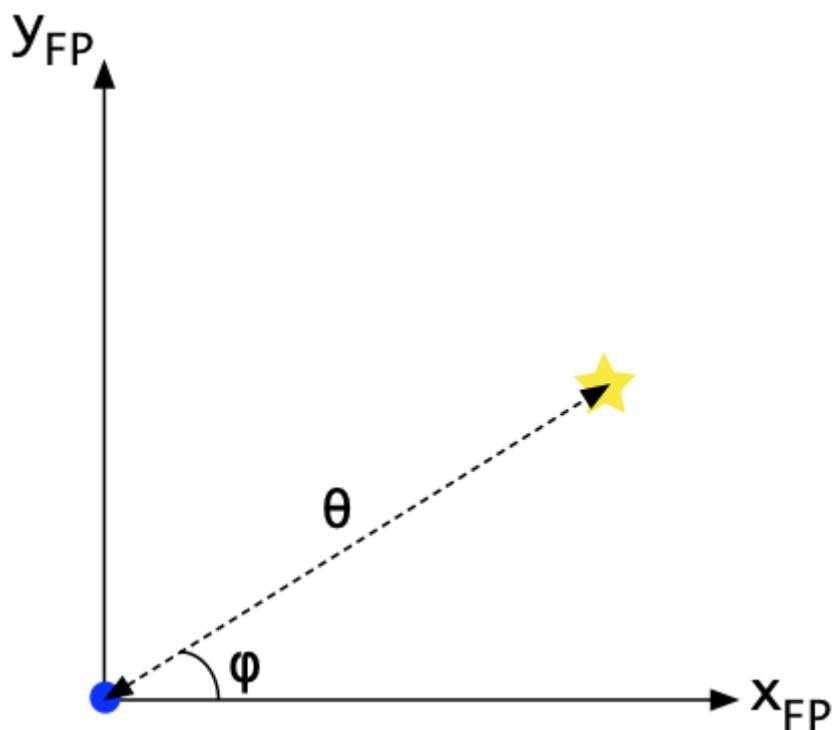


Figure 8. Definition of the field angles  $\theta$  and  $\varphi$ . The former is the angular distance to the optical axis (marked by means of the blue dot), the latter is the in-field angle, measured in the counter-clockwise direction from the  $x_{FP}$  axis.

The following coordinate conversions can be imported from `egse.coordinates`:

```
row, column, ccd_code = focal_plane_to_ccd_coordinates(x_fp, y_fp)
```

- Identifies the CCD on which the given focal-plane coordinates ( $x_{fp}, y_{fp}$ ) are located, and returns the pixel coordinates (`row, column`) in the corresponding CCD reference frame and the CCD code (`ccd_code`).
  - The input focal-plane coordinates ( $x_{fp}, y_{fp}$ ) should be specified in mm.
  - The output pixels coordinates (`row, column`) are given in pixels.
  - The output CCD code `ccd_code` should be 1/2/3/4 (as in Figure 7).
  - If the given focal-plane coordinates do not fall on any of the CCDs, (None, None, None) is returned.

```
theta, phi = focal_plane_coordinates_to_angles(x_fp, y_fp)
```

- Converts the given focal-plane coordinates ( $x_{fp}, y_{fp}$ ) to field angles (`theta, phi`).
  - The input focal-plane coordinates ( $x_{fp}, y_{fp}$ ) should be specified in mm.
  - The output field angles (`theta, phi`) are given in degrees.

```
x_fp, y_fp = ccd_to_focal_plane_coordinates(row, column, ccd_code)
```



- Converts the given pixel coordinates (`row`, `column`) in the reference frame of the CCD with the given code `ccd_code` to focal-plane coordinates (`x_fp`, `y_fp`).
  - The input pixel coordinates (`row`, `column`) should be given in pixels.
  - The input CCD code `ccd_code` should be 1/2/3/4 (as in [Figure 7](#)).
  - The output focal-plane coordinates (`theta`, `phi`) are given in mm.

```
x_fp, y_fp = angles_to_focal_plane_coordinates(theta, phi)
```

- Converts the given field angles (`theta`, `phi`) to focal-plane coordinates (`x_fp`, `y_fp`).
  - The input field angles (`theta`, `phi`) should be given in degrees.
  - The output focal-plane coordinates (`x_fp`, `y_fp`) are given in mm.

```
x_distorted, y_distorted = undistorted_to_distorted_focal_plane_coordinates(  
    x_undistorted, y_undistorted, distortion_coefficients, focal_length)
```

- Converts the given undistorted focal-plane coordinates (`x_undistorted`, `y_undistorted`) to distorted focal-plane coordinates (`x_distorted`, `y_distorted`), based on the given distortion coefficients(`distortion_coefficients`) and focal length (`focal_length`).
  - The input undistorted focal-plane coordinates (`x_undistorted`, `y_undistorted`) should be given in mm.
  - The input distortion coefficients (`distortion_coefficients`) should be an array [k1, k2, k3] with the coefficients of the distortion polynomial.
  - The input focal length (`focal_length`) should be given in mm.
  - The output distorted focal-plane coordinates (`x_distorted`, `y_distorted`) should be given in mm.



# 9. When a new camera arrives

This chapter describes the different actions to be taken when a new camera arrives at either CSL or one of the test houses. The actions are different for CSL with respect to the test houses, so this will be described in separated sections.

## 9.1. Setup

The setup file is a configuration file, part of the CGSE, describing all components and calibrations of the actual test article, test equipment, and test environment. In addition to holding some information like the ensemble of reference frame definitions (in CSL) or the history, it consists of two main components: a first describing the GSE and gathering the information over the test environment (devices control and calibration, temperature sensors, ...), and a second describing the camera, i.e. gathering all the information over the description and calibration of the test articles (TOU, FPA, FEE...).

The setup is a YAML file, i.e. a hierarchical “dictionary” (in the python sense of the word). Items can be referred to by a [key, value] pair, where the key refers to their location in the hierarchy and consists of the full path to that location, e.g. `setup.gse.stages.calibration.phi_correction_coefficients: [-0.0014, -0.003]`. The value can be a built-in type (int, float, str, ...), a serialized object, a list, or the path to another calibration file. A more detailed explanation of the Setup is in [Chapter 6](#).

The setup is under configuration control in the GitHub repository [plato-cgse-conf](#) and new setup versions are created exclusively on the server or client machine of the cleanroom for which the setup is intended.

## 9.2. Purpose

This chapter describes the values susceptible to change when the components of a new camera are integrated on a test setup in CSL. The goal of this chapter is to provide a step-by-step procedure for the creation of the new setup, including the source of information for every item to update.

At the time of writing, several setups do already exist that include the calibration of the GSE components on both test setups in CSL (“CSL1” and “CSL2”). So this document assumes their pre-existence rather than creating them from scratch every time again.

Setups exist and are maintained in different configuration directories (and in GitHub) for every test environment, so every test chamber has its own numbering for the setups `SETUP_<SITE>_<NUMBER>_<date>_<time>`, e.g. `SETUP_CS1_00065_230407_091711.yaml` (the date and time are automatically included at creation time). The `<NUMBER>` is sequentially incremented for each new version of a setup and is not reset for each new camera.

**NOTE**

In this chapter, the `setup.components.to_update` will be marked in bold, the `setup.components` that should not need any update are written in italics.



## 9.3. Useful Information

- Test status of each of the cameras and/or its components: <https://s2e2.cosmos.esa.int/confluence/display/PCOT/Test+Status>
- CAM Configuration connects the camera id with the identifiers of its components: <https://s2e2.cosmos.esa.int/confluence/pages/viewpage.action?pageId=171639593>
- CAM Logistics Control keeps track of the shipments from CSL to the THs: <https://s2e2.cosmos.esa.int/confluence/pages/viewpage.action?spaceKey=SCIPLM&title=CAM+Logistics+Control>
- TOU Logistics Control keeps track of the shipments from TAS to LDO: <https://s2e2.cosmos.esa.int/confluence/pages/viewpage.action?spaceKey=SCIPLM&title=TOU+Logistics+Control>

## 9.4. A new camera arrives at CSL

### 9.4.1. Necessary Inputs

This is a list of the input data and files necessary anytime during the process described below, so maybe first check all of them are at hand:

- The latest version of AD01, freshly pulled from Eclipse
- The metrology excel file from the FPA
- The metrology excel file from the TOU
- The raw data from the TOU
- The laser tracker with the input `CSLReferenceFrameModel` from CSL
- The default FPGA parameters file (`FPGA_Configuration_file.ini`)
- The sensor calibration parameters file (`HK_Calibrations_File.ini`)
- The FEE register map calibration file (the same for all (P)FM cameras, but different for EM)
- The HK map calibration file (the same for all (P)FM cameras, but different for EM)
- Telemetry dictionary file (CSV file in plato-cgse-conf)
- The acceptable limits for the CCD offsets & readout noise (`analog_chain_limits_tvac` and `_ambient.yaml`)
- The acceptable limits for the FEE currents and powers (`sft_limits_fee_hk_<camera>.xlsx`)

The references for some of those documents can be found in [Section 9.7](#) or in the latest weekly minutes from CSL.

### 9.4.2. Baseline

I don't understand what you mean here with baseline? I think this should be more elaborate like what information goes into the GitHub ticket, the ticket basically describes the steps to take, and how to do that by providing the Python code that needs to be executed to ingest the proper information into the Setup. There is also a lot of other information that needs to go in, like N-FEE sensor calibration, the



new Register Map definition, the HK packet definition, etc.

1. Open a GitHub ticket on the plato-cgse-conf repository, using the "Configuration for new camera template".
2. Create a new setup, starting from the last setup on the given test-chamber (e.g. CSL1).
3. Create a GitHub ticket for the inclusion of the CSL metrology (xls) file on the CSL server (cgse-conf)
4. Reduce the raw data from LDO
5. Create a GitHub ticket for the inclusion of the LDO-data-reduction (fits) file on the CSL server (data/reduced)

### 9.4.3. Updating the GSE section

There are only a few items to update in this section, as the information doesn't change with the camera.

#### **setup.gse.hexapod**

There are 2 hexapods per test room in CSL → check the proper hexapod ID & information with CSL (source TBD).

There are 4 hexapods : 1A, 1B, 2A, 2B, with 1, 2 referring to the test room CSL1 or CSL2.

The `setup.gse.hexapod.ID`, `CID` and `label` are updated together (TBC).

#### **setup.gse.stages**



Bound to the room. Should not need any update.

The section `setup.gse.stages.calibration` comes from

- CSL1 : PLATO-PL-CSL-RP-0031\_3.0\_R1\_CalibrationReport
- CSL2 : PLATO-PL-CSL-RP-0032\_2.1\_R2\_CalibrationReport

The offsets are simple scalar values describing the value at field angles [boresight, azimuth] = [0,0]

The correction coefficients are polynomials, expressed as [order\_0\_par, order\_1\_par, ...]. The parameters are expressed in the equations presented for each of the stages:

## The origin of the values

Correction of the azimuth  $\varphi$

`setup.gse.stages.phi_correction_coefficients = [order0, order1]` → from equation 4-4 in section "calculation of correction angle  $\varphi$  versus FOV  $\theta$ "

Mirror orientation  $\alpha$

`setup.gse.stages.alpha_correction_coefficients = [order0, order1]` → from equation 4-1 in section "Definition of Mirror orientation"

`setup.gse.stages.calibration.offset_alpha = value` → this value is 0. = the difference between the "calculated mirror orientation" for FoV  $\theta$  in Table 4-8 and `alpha_correction_coefficients[0]`

Mirror position  $x$

`setup.gse.stages.delta_x_correction_coefficients = [order0, order1, order2]` → from equation 4-3 in section "Definition of Mirror position"

`setup.gse.stages.calibration.offset_delta_x = value` → This value is listed under equation 4-3

### `setup.gse.stages.calibration.offset_phi`

This must be updated from the z-component of the rotation vector expressing TOU\_MEC in GL\_ISO.

This is only made available to us after alignment of the camera on the CSL setup.

It can then be found in the xls sheet gathering the laser-tracker metrology from CSL, provided as input to define the CSLReferenceFrame (see [Section 9.6.2](#)).

### `setup.gse.aeu`



Bound to the room. Should not need any update

Original info from TBD

#### **setup.gse.ogse**

Bound to the room. Should not need any update

Original info from TBD

#### **setup.gse.DAQ6510**

Bound to the room. Should not need any update (TBC)

Original info from TBD

This still needs to be checked as the sensors need to be connected in the same way for all cameras. Check the sensor / channel connection.

### **9.4.4. Updating the Camera section**

#### **Camera identifiers**

CSL UI → "Configuration" tab → Sect. "2 - Include information in setup" → "Camera identifiers"

- Set the camera\_id (choose from drop-down menu)
- Set the camera\_serial\_number (from AD01)
- Set the tou\_id (from AD01; e.g. TOU-11100000-FM-01)
- Set the fee\_id (from AD01)
- Set the fpa\_id (from AD01; e.g. FPA-11200000-FM-SN103)
- Complete the description (should contain camera name + issue number)
- Press "run"
- Check the proposed changes → Agree to submit setup

#### **setup.camera.TOU.ID**

See "Camera identifiers" above

#### **setup.camera.fpa**

- **setup.fpa.ID**: See "Camera identifiers" above
- **setup.fpa.avoidance** parameters don't need any update

#### **setup.camera.dpu**

No update needed.

#### **setup.camera.fee**



**setup.camera.fee.type:** N-FEE or F-FEE

**setup.camera.fee.ID:** See "Camera identifiers" above

**setup.camera.fee.calibration:**

1. Pull the “\*HK\_Calibrations\_File.ini” from the FEE / Analogue chain EIDP
2. CSL UI: "Configuration" tab → Sect. "1 - Create configuration / calibration files" → "N-FEE sensor calibration":
  - a. Set the camera\_id (choose from drop-down menu)
  - b. Load the HK calibration file from step 1
  - c. Fill out the handling jigs (taken from AD01)
  - d. Press "run"
3. Submit the file to the plato-cgse-conf GitHub repository (in common/n-fee)
4. CSL UI: "Configuration" tab → Sect. "2 - Include information in setup" → "N-FEE sensor calibration":
  - a. Set the camera\_id (choose from drop-down menu)
  - b. Load the N-FEE calibration file created in steps 2 - 3
  - c. Complete the description (should contain camera name + issue number)
  - d. Press "run"
  - e. Check the proposed changes → Agree to submit setup

**setup.camera.fee.register\_map & setup.camera.fee.hk\_map**

1. Make a copy of the register map and HK map YAML files (in common/n-fee) of a previous (P)FM
2. Replace the beer name in the filenames
3. Submit the files to the plato-cgse-conf GitHub repository (in common/n-fee)
4. CSL UI: "Configuration" tab → Sect. "2 - Include information in setup" → "N-FEE register & HK map":
  - a. Set the camera\_id (choose from drop-down menu)
  - b. Load the register and HK maps from steps 1 - 3
  - c. Complete the description (should contain camera name + issue number)
  - d. Press "run"
  - e. Check the proposed changes → Agree to submit setup

**setup.camera.fee.fpga\_defaults**



1. Pull the “\*FPGA\_Configuration file.ini” from the FEE / Analogue chain EIDP
2. CSL UI: "Configuration" tab → Sect. "1 - Create configuration / calibration files" → "N-FEE FPGA defaults":
  - a. Set the camera\_id (choose from drop-down menu)
  - b. Load the FPGA defaults file from step 1
  - c. Press "run"
3. Submit the file to the plato-cgse-conf GitHub repository (in common/n-fee)
4. CSL UI: "Configuration" tab → Sect. "2 - Include information in setup" → "FEE FPGA defaults":
  - a. Set the camera\_id (choose from drop-down menu)
  - b. Load the FPGA defaults file created in steps 2 - 3
  - c. Complete the description (should contain camera name + issue number)
  - d. Press "run"
  - e. Check the proposed changes → Agree to submit setup

`setup.camera.fee.ccd_numbering, ccd_sides, sensor_sel`

no update

`setup.camera.fee.power_consumption`

#### A. Prepare the calibration file

1. Identify the appropriate versions of the MSSL-0058 short functional test reports for both the ambient and tvac cases. For instance, the ambient case might be labelled “before shipping”. Many versions of MSSL-0058 exist for every camera, so have your selection validated by ESA before anything else.
2. Find an existing calibration file in the cgse-conf repo, in common/n-fee. For instance sft\_limits\_fee\_hk\_duvel\_v01.xlsx
3. Copy it to a new file bearing the name of the new camera, to make sure to respect its format in the new version
4. (Manually) edit the cover page and the values in the ‘ambient’ and ‘tvac’ sheets according to the MSSL-0058 reports (table numbers are indicative; in case of deviation, apply common-sense; powers are computed from voltages and currents)
  - i. Voltages (table 14-8). The tolerances on the voltages are set to 100%.
  - ii. Currents ON mode (table 14-9)
  - iii. Currents STANDBY mode (table 14-14)
  - iv. Currents in FULL\_IMAGE mode READOUT (table 14-16)
  - v. Currents in FULL\_IMAGE mode INTEGRATION (table 14-17)

#### B. Update the setup accordingly

1. CSL UI: “Configuration” tab → Sect. 2 (SFT) I and P limits”
2. version (mandatory) : the version of your calibration file (integer)
3. confdir (optional) : can be left blank : the directory where the calibration files will be found. The default is taken from the env. variable PLATO\_CONF\_DATA\_LOCATION
4. description (optional) : if left blank, replaced with “Incl. power consumption limits for <camera>. Ref. file version <version>”

The calibration file is selected based on the camera.ID, which is determined from the current setup (`setup.camera.ID`), so prior to this, make sure to start after loading a preliminary setup for the new camera.

**setup.camera.ccd.limits**



- A. Calibration file received by email from Sami Niemi, and placed on `$PLATO_CONF_DATA_LOCATION/common/ccd`. Two independent files exist for ambient and tvac. Both have the same format and contain the information for all cameras at once. The filenames are “analogue\_chain\_limits\_<temp>\_<version>.yaml”, where <temp> is either ‘ambient’ or ‘tvac’ and <version> is ‘v??’ with ?? = the version number, e.g. 04.

Verify that the file received from Sami has the right version number (= the ‘next one’) and that the temperatures are well formulated as “ambient” and “tvac” (e.g. not “cold”)

- B. Update the setup accordingly:

1. CSL UI: “Configuration” tab → Sect. 2 (SFT) CCD offsets and readout noise”
2. version (mandatory) : the version of your calibration file (integer)
3. temp (mandatory) : either ‘ambient’ or ‘tvac’
4. confdir (optional) : can be left blank : the directory where the calibration files will be found. The default is taken from the env. variable `PLATO_CONF_DATA_LOCATION`
5. description (optional) : if left blank, replaced with “Incl. power consumption limits for <camera>. Ref. file version <version>”

The calibration file is selected based on the camera.ID, which is determined from the current setup (`setup.camera.ID`), so prior to this, make sure to start after loading a preliminary setup for the new camera.

#### **setup.camera.ID, setup.camera.serial\_number**

See "Camera identifiers" above

#### **setup.camera.ccd**

`setup.camera.orgin_offset_x`, `origin_offset_y` and `orientation`: must be updated from the FPA metrology excel sheet.

To compute the new values from plato-test-scripts, run the ipython notebook `scripts/ccd_info.ipynb`.

Alternative:

1. Input data: FPA metrology excel sheet. PL-ALN-CSL-0050 "coordinates of the corner pixels")
2. Script `camtest.analysis.analysis_fpa_eidp_corners.py`
  - a. Ingest the coordinates according to the format existing for previous cameras
  - b. Run the code snippet for the new camera.
3. CSL UI: "Configuration" tab → Sect. "2 - Include information in setup" → "CCD positions":
  - a. Set the `camera_id` (choose from drop-down menu)
  - b. Load the FPA metrology spreadsheet from step 1
  - c. Complete the description (should contain camera name + issue number)
  - d. Press "run"
  - e. Check the proposed changes to the setup → Agree to submit the setup

`setup.camera.fov`: TOU metrology excel sheet

`setup.camera.fov.focal_length`: update from LDO TOU metrology excel sheet.

`setup.camera.fov.distortion_coefficients`: update from LDO TOU metrology excel sheet.

`setup.camera.fov.inverse_distortion_coeffcients`: computed from the distortion coefficients

1. Input data: TOU metrology excel sheet
2. To compute : from the plato-test-scripts, run the ipython notebook `scripts/field_distortion.ipynb`
3. CSL UI: "Configuration" tab → Sect. "2 - Include information in setup" → "Field distortion & Focal length":
  - a. Set the `camera_id` (choose from drop-down menu)
  - b. Load the TOU metrology spreadsheet from step 1
  - c. Complete the description (should contain camera name + issue number)
  - d. Press "run"
  - e. Check the proposed changes to the setup → Agree to submit the setup

`setup.camera.fov.tau`: no update. If needed, it will be updated after recalibration.



## 9.4.5. Remaining Setup sections

### `setup.cs1_model`

is automatically updated during operations. This demands a laser-tracker metrology input file from CSL.

See [Section 9.6.1](#) for details.

### `setup.fov_positions.reference_full_40.x & y`

To be updated from the LDO metrology, for use with ‘use\_angles = False’ in the verification of the LDO metrology.

They can be found in the columns presenting the TOU\_MEC / MRF X & Y coordinates of the LDO measurements at ambient, in their metrology deliverables document (columns x and y in PL-ALN-CSL-0520 in sheet “TOU” of the TOU/LDO “RP” document; see [Section 9.7](#)).

### `setup.telemetry`

1. Make a copy of the telemetry dictionary of the previous camera (CSV file)
2. Replace the beer name in the filename
3. Submit the files to the plato-cgse-conf GitHub repository (in common/telemetry)
4. CSL UI: "Configuration" tab → Sect. "2 - Include information in setup" → "Telemetry"
  - a. Set the camera\_id (choose from drop-down menu)
  - b. Load the file from step 1
  - c. Complete the description (should contain camera name + issue number)
  - d. Press "run"
  - e. Check the proposed changes → Agree to submit the setup

### `setup.sensor_calibration`

No update needed.

## 9.5. A new camera arrives at the test houses

TBW

## 9.6. Additional preparation steps

### 9.6.1. Reduce the raw LDO data

The raw data from LDO’s measurements at ambient must be reduced and the results formatted as if acquired in CSL to allow for easy comparisons by the CSL operators.

The raw data is attached to a document, whose reference can be found in the CSL weekly meeting (updated there). See also [Section 9.7](#).



1. Verify the next obsid that is free for LDO (see table in Section XXXXX).
2. Reduce the LDO data
3. Compare the results (ellipse-sizes) with those from LDO (presented in their metrology excel sheet)
4. Copy the resulting fits file on the CSL server, in the /data/reduced/ directory (TBC)

So far the raw data is reduced manually, due to various issues and changes in the measurements and in the formats of the LDO deliveries.

### 9.6.2. CSLReferenceFrame

The CSL operator GUI needs a laser-tracker metrology file from CSL. That is an excel file with a “data” sheet that is read by the task in the GUI.

1. The file is communicated to KUL after alignment of the camera on the setup. It doesn’t need any modification.
2. It must be pushed to the plato-cgse-conf repository in the 'data/CSL[12]/conf' directory and after that the repository must be pulled on both the server and client machines.

NB: The name of the file will be PLATO-CSL-PL-RP\_00YY vX.3\_CSL\_RFModel\_FMX\_CAMID\_ROT?\_ISO???.xlsx where YY is a document number, X is the FM-TOU-number (e.g. FM2 is Chimay), and the single and double question marks point to the rotation and isostatic tables used. The filename is not critical, it is a free parameter in the arguments panel of the CSL operator GUI.

## 9.7. Reference Information

Camera configuration & ABCL:

- Most information can be found in AD01: PLATO-INAF-PL-LI-0071 : PFM\_FM\_FS\_CAM\_ABCL
- Links between CAM ID, TOU, FPA & metrology reports are in the minutes of the CSL weekly (the table is permanent and updated in there)

CAM ID : from PLATO-CSL-PL-PR-0019 version 1.1

TOU is the usual reference FM#n.

- The serial number can be found in AD01 or in Confluence: <https://s2e2.cosmos.esa.int/confluence/pages/viewpage.action?pageId=171639593>
- It has the following structure: N?-CAM-1?100000-FM-## where ? runs from 1 to 4 and denotes the camera-group, whereas ## runs from 01 to 06 and marks the camera number within the group.

DEAD : from AD01 (original reference: PTO-EST-PL-TN-1369 version 1.0)

Metrology

- For the TOU, metrology reports start with PLATO-INAF-PL-RP and can be found in the TOU EIDP, in Eclipse or [/STER/platodata/EIDP/{camera}/TOU/… /12-Test Reports/](#)



- For the TOU, the raw data can be found in documents starting with PLATO-INAF-PL-TR, and are in the EIDP or delta\_EIDP
- For the FPA, metrology reports start with PTO-EST-PL-RP and can be found directly in Eclipse

TODO: link the TOU & FPA Ref. to their EIDP in Eclipse ?

To be allocated to the right camera:

- PTO-EST-PL-REP-1551 - PLATO FM7 FPA Metrology Summary
- PTO-EST-PL-REP-1552 - PLATO FM8 FPA Metrology Summary
- PTO-EST-PL-REP-1553 - PLATO FM10 FPA Metrology Summary

# 10. Switching ON/OFF the Camera

Before you can start any operations or tests, you will need to switch on the Camera N-CAM following a dedicated Camera Switch ON procedure. Also for switching off the Camera you will need to follow a series of steps to be executed in the correct order.

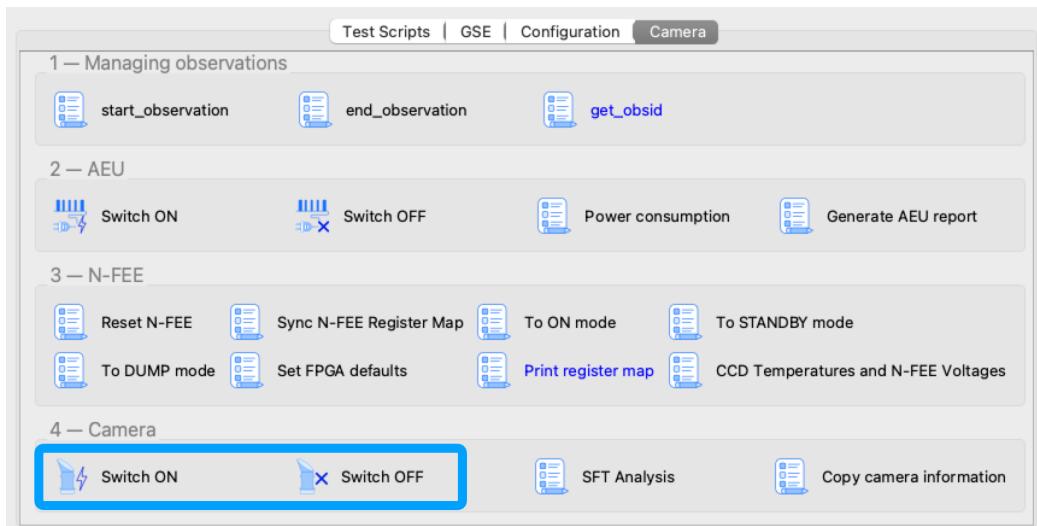
The steps for Camera Switch ON are the following:

1. Send commands to the AEU to power on the camera and enable the sync signals → Camera is now in ON mode
2. Send the FPGA defaults to the N-FEE
3. Go to STANDBY mode
4. Go to DUMP mode, external sync

The steps for the Camera Switch OFF:

1. Go to STANDBY mode
2. Go to ON mode
3. Send commands to the AEU to disable the sync signals and power off the camera

Note that, although the individual steps all exist as tasks in the Operator GUI, the procedure to switch OFF and ON the Camera should be used from the Operator GUI: [Camera TAB > 4 – Camera > Switch ON / Switch OFF](#). These tasks step through the different steps in the procedure and perform additional checks like voltages/currents/power, N-FEE mode etc. prompting the user for confirmation at each step.



*Figure 9. Screenshot of the Operator Task GUI where the Camera Switch On and Switch OFF have been framed.*

## 10.1. Detailed description of Camera Switch ON

This section describes what is done during the Camera Switch ON procedure, i.e. by running the task 'Camera Switch ON' in the Operator GUI. Although the individual steps are explained in detail, you



should still run the **Camera Switch ON** task instead of each step separately. **Only use the individual commanding when in a contingency.**

The AEU Test EGSE shall be on StandBy mode before we start. You can check that in the AEU GUI where the *Stand-by* LED in the left panel (EGSE mode) shall be green.

**WARNING**

Before powering on the AEU, stop all AEU related control servers. When you power on the AEU, the LEDs on the front panel start blinking (you won't see anything in the AEU GUI since the control servers are not running). This blinking takes about three minutes and you shall NOT start the AEU control servers before the LEDs stop blinking.

### Step 1. Send commands to the AEU to power on the camera and enable the sync signals

The AEU Test EGSE is now ready to power the camera and enable the synchronisation. This is the first step in the procedure. The individual task is: **Task GUI > Camera TAB > 2 – AEU > Switch ON**. From the Python REPL<sup>[1]</sup>, you can execute these commands:

```
>>> from camtest import start_observation, end_observation ①
>>> from camtest.commanding import aeu

>>> start_observation("AEU N-CAM Switch ON")
>>> aeu.n_cam_swon()
>>> aeu.n_cam_sync_enable(image_cycle_time=25, svm_nom=1, svm_red=0)
>>> end_observation()
```

① import statements will be given only once, they are normally automatically loaded from the **startup.py** file.

The camera shall now be in ON mode (check this in the DPU GUI) and the AEU GUI should have the following LEDs turned green: 'Functional check & TVAC', 'N-CAM', all power lines

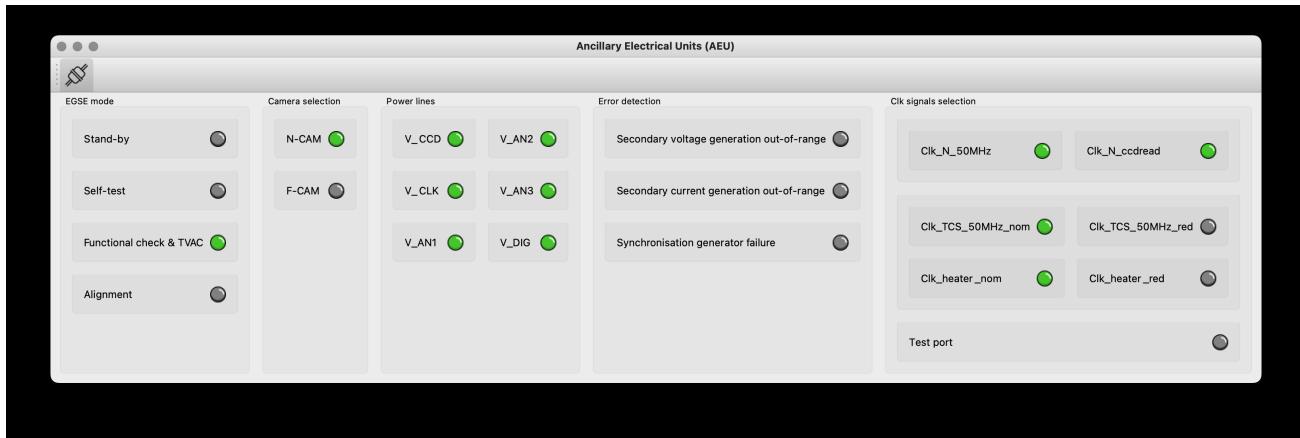


Figure 10. The state of the AEU Test EGSE after a AEU N-CAM Switch ON

At this point, the Camera Switch ON procedure will perform a few additional tasks that are not done when you execute the individual steps:

1. The AEU (cRIO & PSU) HK sampling frequency will be increased to 4Hz for the time of the Camera Switch ON procedure. The frequency can be changed in the arguments panel of the task GUI.



2. The N-FEE register map is loaded by the DPU Processor. This will synchronise the DPU Processor internal state with the N-FEE state. This step requires that the DPU CS was started and is running.
3. The AEU cRIO voltages and currents will be printed in an overview table and checked against their limits. The operator is asked to confirm or abort.

## Step 2. Send the FPGA defaults to the N-FEE

### NOTE

This step shall be skipped for the EM camera as there are not updated FPGA defaults defined for EM.

Each camera N-FEE requires a update of a number of FPGA parameters in the register map. That is what we call the *FPGA defaults*. Please note that these settings are camera specific and influence the proper readout of the CCDs, so it's important that this step is executed at the right time in the procedure. The individual task is: [Task GUI > Camera TAB > 3 – N-FEE > Set FPGA defaults](#). From the REPL you can execute these commands:

```
>>> from camtest.commanding import dpu  
>>> execute(dpu.n_fee_set_fpga_defaults)
```

The user will be prompted to confirm the values have been correctly applied in the N-FEE FPGA.

The FPGA defaults are read from the current Setup. In the Setup there is an entry `setup.camera.fee.fpga_defaults` that loads the YAML file with the correct values for the camera. The values are expressed as registers (32bit values) and not as individual parameters. You can inspect the values as follows:

```
>>> setup.camera.fee.fpga_defaults  
  
NavigableDict  
|   reg_0_config: 119D0000  
|   reg_1_config: 0  
|   reg_2_config: E40FA36B  
|   reg_3_config: 8F60000  
|   reg_4_config: 186A7D8C  
|   reg_5_config: 3EA030D4  
|   reg_6_config: 0  
|   reg_7_config: 0  
|   reg_8_config: 0  
|   reg_9_config: 0  
|   reg_10_config: 0  
|   reg_11_config: 0  
|   reg_12_config: 0  
|   reg_13_config: 0  
|   reg_14_config: 0  
|   reg_15_config: 0  
|   reg_16_config: 0  
|   reg_17_config: 0  
|   reg_18_config: 7FE7EF16  
|   reg_19_config: FE7EE7FE  
|   reg_20_config: 198CD  
|   reg_21_config: 5E5000  
|   reg_22_config: 4241AE9  
|   reg_23_config: 0  
|   reg_24_config: 0  
|   reg_25_config: 6400000  
|   reg_26_config: 3E807D0
```



```
>>> setup.camera.fee.get_raw_value('fpga_defaults')
'yaml://.../common/n-fee/nfee_fpga_defaults Brigand.yaml'
```

### Step 3. Go to STANDBY mode

Now, the camera will be brought into STANDBY mode. This means the CCDs will be powered and start accumulating flux. The individual task is: [Task GUI > Camera TAB > 3 – N-FEE > To STANDBY mode](#). From the REPL, the command is:

```
>>> execute(dpu.n_cam_to_standby_mode)
```

In the Camera Switch ON procedure, the user will be prompted to confirm the N-FEE is actually in STANDBY mode.

Again, the procedure will check the AEU cRIO voltages and currents and print an overview table. Please note that the configured limits are different for ON mode, STANDBY mode and also DUMP mode. At each step the procedure will perform this check. The operator is asked to confirm or abort.

### Step 4. Go to DUMP mode, external sync

The last step in the switch-on procedure is to bring the camera in DUMP mode (external sync). As explained in [Chapter 12](#), DUMP mode is not a genuine FEE operation mode, but is defined in the CGSE as a state in which the N-FEE is in FULL IMAGE mode in which the dump gate is kept high, continuously resetting the readout register. No data is acquired or send to the DPU. The individual task is: [Task GUI > Camera TAB > 3 – N-FEE > To DUMP mode](#). The command you can use in the REPL is:

```
>>> execute(dpu.n_cam_to_dump_mode)
```

The user will then be prompted to confirm that the N-FEE is actually in DUMP mode. After confirmation, the AEU cRIO voltages and currents will be checked against their limits and printed in an overview table asking the user for another confirmation.

One last step that is part of the Camera Switch ON procedure is to take a series of single images without the light source on. That is, it will be full frame darks with the dump gate enabled.

```
>>> dpu.n_cam_acquire_and_dump(num_cycles=5, row_start=0, row_end=4539, rows_final_dump=0, ccd_order=[1, 2, 3, 4], ccd_side="BOTH")
```

Finally, the AEU cRIO and PSU HK frequency is reset to their default values from the Settings, i.e. [HK\\_DELAY](#).

This concludes the Camera Switch ON procedure as it is currently implemented and required to execute. The camera will be in DUMP mode after this switch-on which is the starting state for all the TVAC tests that will be executed.

## 10.2. Detailed description of Camera Switch OFF

Switching OFF the camera is much simpler and doesn't require specific checks. The N-FEE should normally be in DUMP mode when you start this switch-off and is brought first to STANDBY mode, then to ON mode. Each of these steps require confirmation from the operator. Finally, the AEU Switch OFF is executed which disables the sync signals and powers off the camera. The individual tasks are [Camera TAB > 3 – N-FEE > To STANDBY mode](#), [Camera TAB > 3 – N-FEE > To ON mode](#), and [Camera TAB > 2 – AEU > Switch OFF](#). From the REPL, the following commands accomplish the same result:

```
>>> start_observation("AEU N-CAM Switch OFF")
>>> dpu.n_cam_to_standby_mode()
>>> dpu.n_cam_to_on_mode()
>>> aeu.n_cam_sync_disable()
>>> aeu.n_cam_swoff()
>>> end_observation()
```

**IMPORTANT**

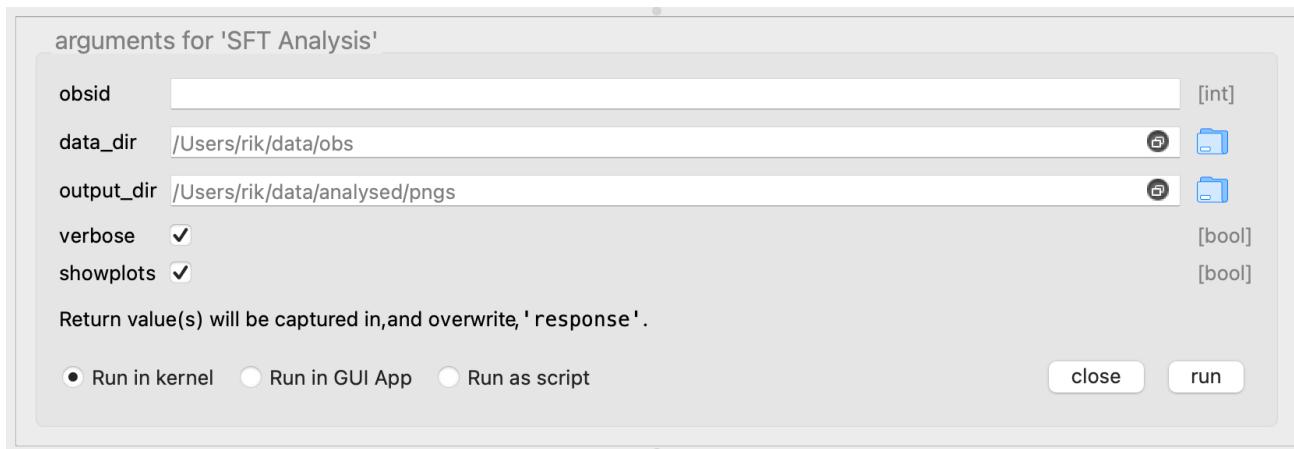
As with the Camera Switch ON procedure, also for the switch-off, you shall use the Camera Switch OFF task from the [Task GUI > Camera TAB > 4 – Camera > Switch OFF](#) instead of using the individual tasks or commands.

## 10.3. Analysis of the Short Function Test (SFT)

Every time you execute the Camera Switch ON procedure, the task will perform a short functional test using the [acquire\\_and\\_dump](#) command as shown below. This is a full-frame measurement of all four CCDs and both sides.

```
>>> dpu.n_cam_acquire_and_dump(num_cycles=5, row_start=0, row_end=4539,
                                rows_final_dump=0, ccd_order=[1, 2, 3, 4],
                                ccd_side="BOTH")
```

In the Operator Task GUI, Camera TAB (see [Figure 9](#)), there is a task button to execute the SFT analysis. In the arguments panel, you should provide the [obsid](#) of the Camera Switch ON as an integer, i.e. only the test id, not the site id. The [data\\_dir](#) is the parent folder where the observations are stored and the [output\\_dir](#) is where the result of the analysis will be saved.



*Figure 11. Input arguments panel that will appear when selecting the SFT Analysis task button in the Operator Task GUI.*

The SFT analysis runs two different functions:

- The first one verifies the currents and power limits on the N-FEE with the HK of the AEU voltages and currents.
- The second analyses the image data and verifies that the electronic offset and readout noise are within the acceptable boundaries for the given camera.

These functions produce several standard output files, consisting in

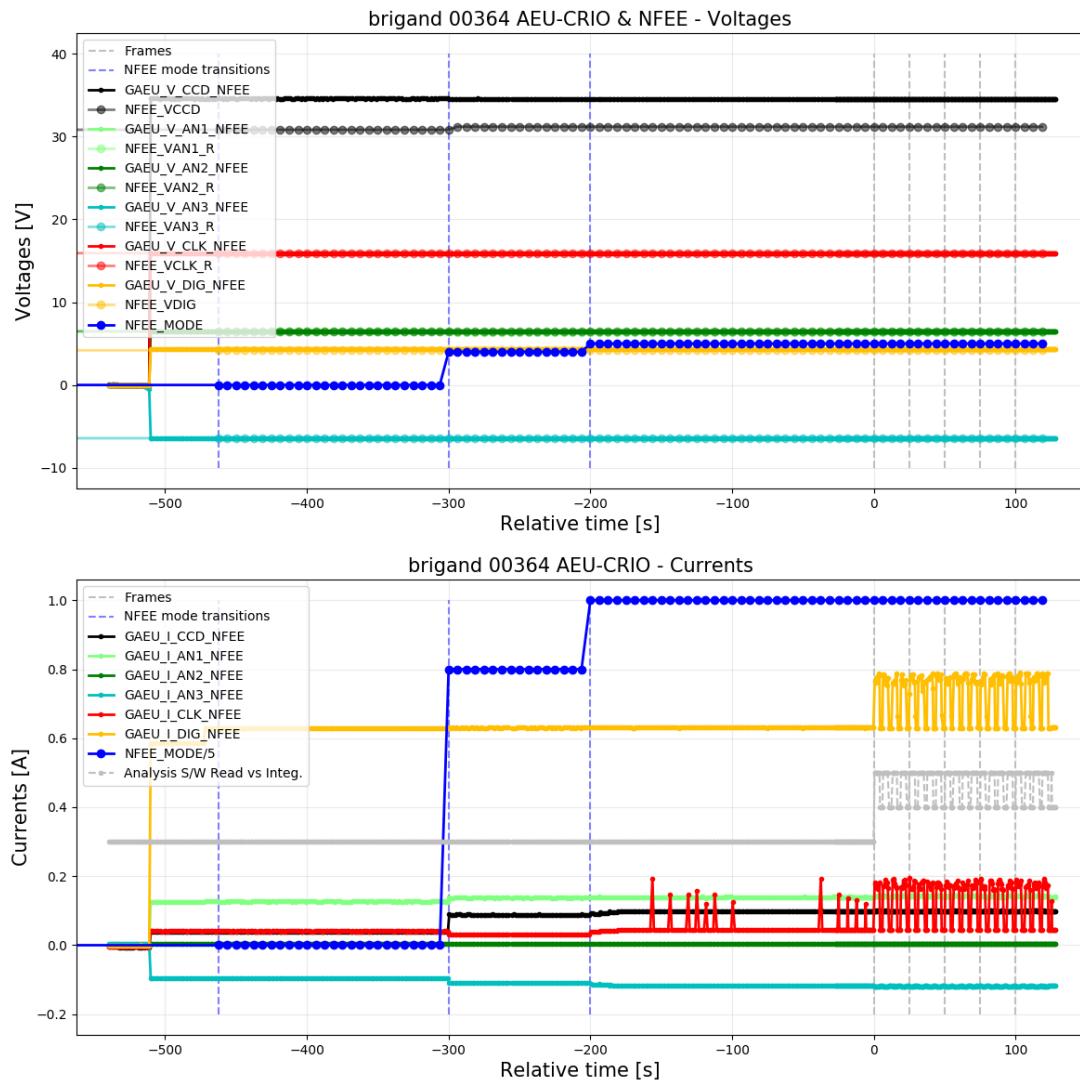
- plots of the data
- text files with the limit-checks and success-criteria.

The HK checks produce the following files:

- `sft_analysis_<camera>_<obsid>_HK_AEU-CRIO_VI.png`
- `sft_analysis_<camera>_<obsid>_HK_checks.txt`

where `<camera>` is the camera.ID, e.g. 'duvel', and `<obsid>` is the observation id, e.g. '01234'.

The former is a plot of the HK, looking like this for an SFT measurement on Brigand:



*Figure 12. Standard output of the HK checks in the SFT analysis. The top panel shows the voltages, while the currents are on the bottom. The blue dashed lines mark the FEE-mode transitions (starting with OFF, then ON, STANDBY and FULL\_IMAGE\_MODE). The dashed gray lines mark the frame acquisition times during the execution of `n_cam_acquire_and_dump`.*

The latter file is a text file and contains the limit-checks for currents and power consumption, and nominally ends with a message formatted like this:

```

All V-AEU OK: True
All V-FEE OK: True
All I OK:      True
All P OK:      True

All values OK.

IF offsets & readout noise OK, you can PROCEED.

```



If any of the checks are invalid it recommends to “STOP for analysis”.

The image analysis part produces the following files:

- `sft_analysis_<camera>_<obsid>_img_raw.png`
- `sft_analysis_<camera>_<obsid>_img_offsetSubtracted.png`
- `sft_analysis_<camera>_<obsid>_img_avg_rows_and_columns.png`
- `sft_analysis_<camera>_<obsid>_img_checks_offsets_ron_<func>.txt`

where `<func>` is the statistical function used to combine the results obtained from individual columns (default is ‘mean’).

The ‘raw’ image displays one full-frame image of the 4 CCDs (the layer is chosen at run-time of the analysis, and is reported in the image itself).

The `offsetSubtracted` image corresponds to the same data, after subtraction from the offset computed from the serial overscan columns.

The `avg_rows_and_columns` presents the average row and column profiles of the image area (including parallel overscan) of the 8 half-CCDs independently.

Examples of such images can be found below.

The text file contains the limit-checks for electronic offsets and readout noise, and nominally ends with a message formatted like this:

```

Summary
off_spre      OK
off_sover     OK
off_pover     OK
off_img       OK
noise_sover   OK
noise_img_col OK

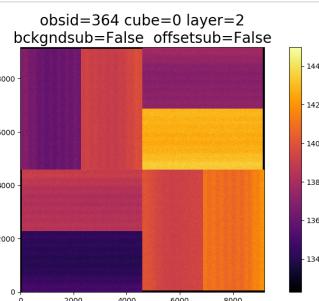
Checks of CCD offsets & r.o.n. OK.

IF voltage checks OK, you can PROCEED.

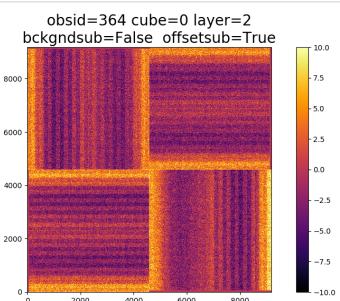
```

Here as well, if any of the checks are invalid it recommends to “STOP for analysis”.

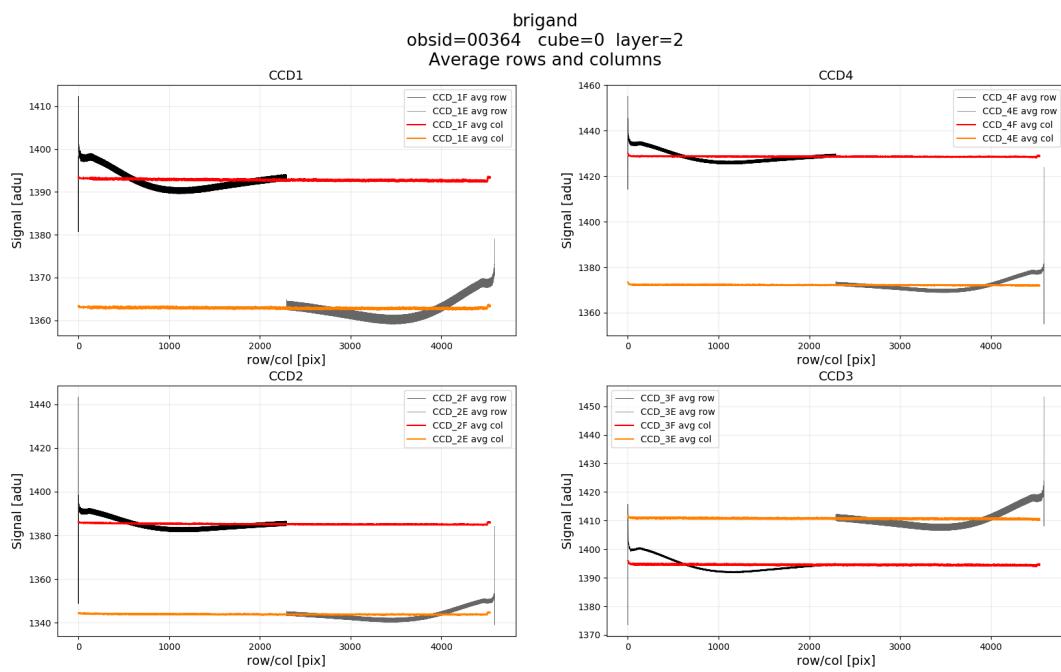
*Table 2. Standard images output by the SFT analysis.*



*Figure 13. The RAW image.*



*Figure 14. Image after offset-subtraction.*



*Figure 15. Standard image output of the SFT analysis, presenting the average row and column profiles for the image-area in all half-CCDs.*

[1] A REPL (Read-Eval-Print Loop) is an interactive programming environment that allows users to enter Python code, which is then executed, and the results are immediately displayed, aka the Python command prompt.



# 11. Operating the AEU EGSE

## 11.1. Introduction

On the spacecraft, the Ancillary Electronics Unit (AEU) is providing the FEE with the necessary stabilised secondary voltages and is also providing synchronisation signals to allow the different FEEs to synchronise their CCD readout cycle to the temperature control system (TCS) heater power switching.

On CAM level, we do not use a flight-like AEU. A dedicated AEU EGSE is providing the power supplies and synchronisation signals needed to operate a single N-type or single F-type camera.

The AEU EGSE supplies 6 voltages to the FEE:

- V\_CCD
- V\_CLK
- V\_AN1
- V\_AN2
- V\_AN3
- V\_DIG

[Figure 16](#), [Figure 17](#), [Figure 18](#), and [Figure 19](#) list the default voltage and current values for the N-AEU and F-AEU.

N-AEU Power Line	Min. output voltage (V)	Nom. Output voltage (V)	Max. Output voltage (V)	Max. Output fault voltage considering a single failure (V)
V_CCD	34.00	34.70	35.40	39.00
V_CLK	15.70	16.05	16.40	18.00
V_AN1	6.50	6.65	6.80	7.50
V_AN2	6.50	6.65	6.80	7.50
V_AN3	-6.50	-6.65	-6.80	-7.50
V_DIG	4.45	4.55	4.65	5.20

*Figure 16. Default voltage values for the N-AEU.*



N-AEU Power Line	Peak supply current (A)	Max. steady state current (A)	Max. average current (A)	Min. current (A)
V_CCD	0.211	0.105	0.105	0.000
V_CLK	0.295	0.208	0.148	0.000
V_AN1	0.380	0.190	0.190	0.000
V_AN2	0.116	0.058	0.058	0.000
V_AN3	-0.449	-0.224	-0.224	-0.000
V_DIG	1.023	0.628	0.549	0.100

Figure 17. Default current values for the N-AEU.

F-AEU Power Line	Min. output voltage (V)	Nom. Output voltage (V)	Max. Output voltage (V)	Max. Output fault voltage considering a single failure (V)
V_CCD	30.8	31.6	32.4	38.0
V_CLK	15.6	16.2	16.8	21.5
V_AN1	7.3	7.7	8.1	9.0
V_AN2	5.0	5.4	5.8	6.4
V_AN3	-7.1	-7.6	-8.1	-9.0
V_DIG	4.9	5.3	5.7	6.5

Figure 18. Default voltage values for the F-AEU.

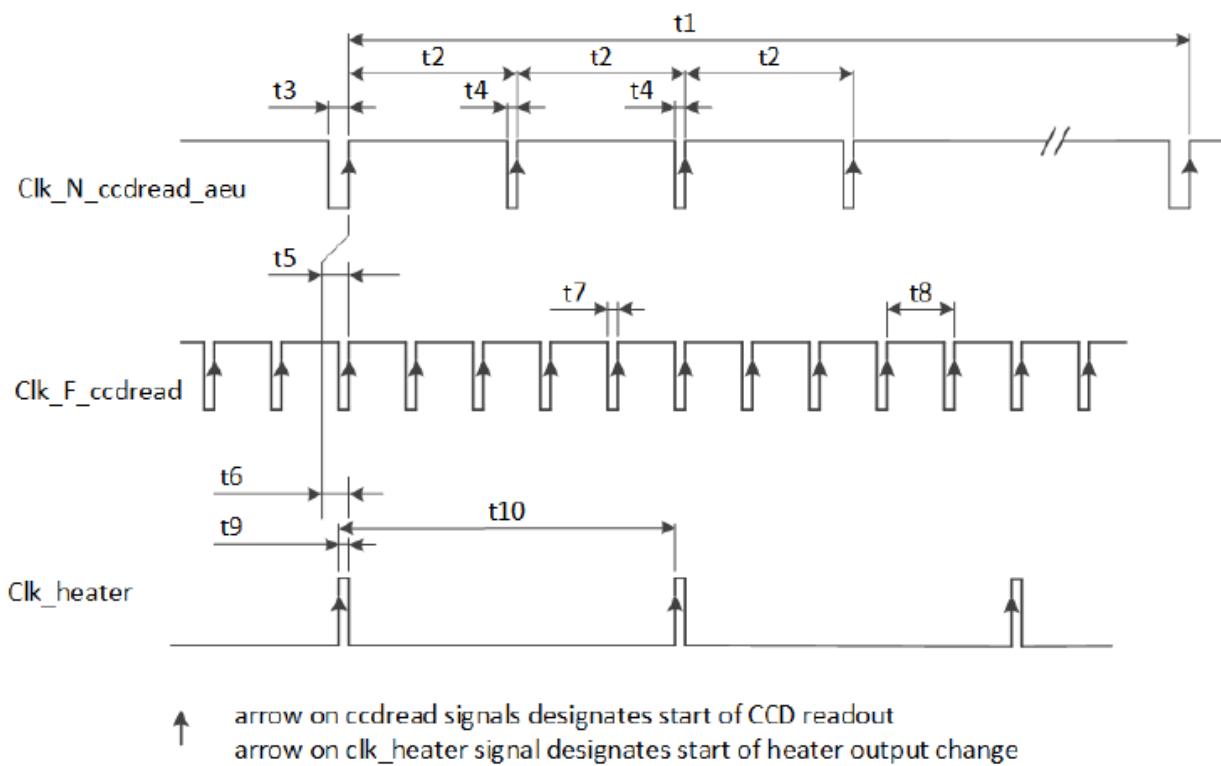
F-AEU Power Line	Peak supply current (A)	Max. steady state current (A)	Max. average current (A)	Min. current (A)
V_CCD	0.213	0.164	0.162	0.000
V_CLK	0.920	0.708	0.656	0.000
V_AN1	0.665	0.511	0.511	0.000
V_AN2	1.820	1.400	1.400	0.000
V_AN3	-0.222	-0.171	-0.171	0.000
V_DIG	3.093	2.379	2.379	0.728

Figure 19. Default current values for the F-AEU.

The AEU EGSE provides sync pulses to the N-FEE, F-FEE, TCS EGSE, and the EGSE (e.g. shutter controller of the OGSE). The following synchronisation signals can be configured:

- `Clk_N_ccdread_AEU / Clk_F_ccdread_AEU`;
- `Clk_heater` (nominal/redundant in case of operating an F-CAM);
- `Clk_50MHz` (nominal redundant in case of operating an N-CAM)

The synchronisation output signal timings are shown in [Figure 20](#).



ID	Description	Min	Nom	Max	Note
t0	50 MHz cycle time = clk		20 ns		
t1	N-AEU Image cycle time (programmable)		1.2500E9 clk 1.5625E9 clk 1.8750E9 clk 2.1875E9 clk 2.5000E9 clk		25.00 s 31.25 s 37.50 s 43.75 s 50.00 s
t2	N-AEU single CCD readout period		3.125E8 clk		6.25 s
t3	CCD 1 start readout pulse width		2E7 clk		400 ms
t4	CCD 2 to 4 start readout pulse width		1E7 clk		200 ms
t5	Skew Clk_F_ccdread to Clk_N_ccdread AEU	-50 ns		50 ns	
t6	Skew Clk_heater to Clk_N_ccdread AEU	-50 ns		50 ns	
t7	Clk_F_ccdread pulse width		1E7 clk		200 ms
t8	Clk_F_ccdread cycle time		1.25E8 clk		2.5 s
t9	Clk_heater pulse width		1E7 clk		200 ms
t10	Clk_heater cycle time		6.25E8 clk		12.5 s

Figure 20. Synchronisation output signal timings: Clk\_ccdread & Clk\_heater, as taken from RD-05.

## 11.2. AEU switch on and off

The AEU is switched on by the operator by powering up the unit. During the first 3 minutes after powering up no AEU commanding shall take place. The front-panel leds stand-by, self-test, functional check & TVAC, and alignment will be blinking until the system is ready.

## 11.3. Changing between AEU EGSE operation modes

The AEU EGSE has 4 functional modes:

**Stand-by:**



- AEU EGSE has power and is ready to receive commands;
- AEU EGSE distributes telemetry;
- No sync signal is generated;
- All FEE secondary voltages are down and cannot be activated.

### Functional check & TVAC mode

- AEU EGSE has power and is ready to receive commands;
- AEU EGSE distributes telemetry;
- All sync signals can be commanded on/off;
- Secondary voltage generation can be commanded on and off.

### (ambient) Alignment operating mode

- AEU EGSE has power and is ready to receive commands;
- AEU EGSE distributes telemetry;
- Only following synchronisation signals can be commanded on/off:
  - `Clk_50MHz_N-AEU_N-FEE`;
  - `Clk_50MHz_F-AEUnom_F-FEE`;
  - `Clk_50MHz_F-AEUred_F-FEE`.
- The secondary voltage generation can be commanded on/off;

### Self-test

- EGSE runs a self-diagnostic to check the status of its sync and secondary voltage outputs, and reports the results in its TM;
- The GSE autonomously reverts to stand-by mode once the self-diagnostic is complete;
- In this mode, the FEE and TCS EGSE have to be physically disconnected from the AEU EGSE, and their connectors on the AEU EGSE have to be connected to the test port connector on the AEU EGSE.

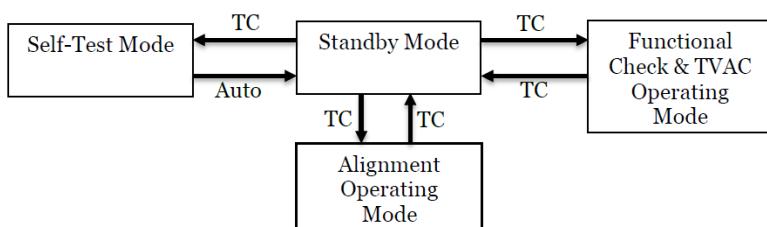


Figure 21. AEU EGSE mode transitions.

Switching modes:

```

>>> from egse.aeu.aeu import OperatingMode
>>> aeu.set_operating_mode(OperatingMode.STANDBY)
>>> aeu.set_operating_mode(OperatingMode.FC_TVAC)
>>> aeu.set_operating_mode(OperatingMode.ALIGNMENT)
    
```



```
>>> aeu.set_operating_mode(OperatingMode.SELFTEST)
```

Checking in which mode the AEU is:

```
>>> mode = aeu.get_operating_mode()
```

## 11.4. Power supply Unit: Setting and checking Current and voltage protections

To read the voltage and current setpoints, and the corresponding over-protection values (OVP and OCP) from the power supply units:

```
>>> v_ccd, v_clk, v_an1, v_an2, v_an3, v_dig = aeu.get_psu_voltage_setpoints()
>>> ovp_ccd, ovp_clk, ovp_an1, ovp_an2, ovp_an3, ovp_dig = aeu.get_psu_ovp()
>>> i_ccd, i_clk, i_an1, i_an2, i_an3, i_dig = aeu.get_psu_current_setpoints()
>>> ocp_ccd, ocp_clk, ocp_an1, ocp_an2, ocp_an3, ocp_dig = aeu.get_psu_ocp()
```

To read the measured voltages and currents from the power supply units:

```
>>> v_ccd, v_clk, v_an1, v_an2, v_an3, v_dig = aeu.get_psu_voltages()
>>> i_ccd, i_clk, i_an1, i_an2, i_an3, i_dig = aeu.get_psu_currents()
```

## 11.5. FEE voltages and currents

To read the measured values for the voltages and currents, and the corresponding protection values (UVP, OVP, and OCP):

For the N-CAM:

```
>>> v_ccd, v_clk, v_an1, v_an2, v_an3, v_dig = aeu.get_n_cam_voltages()
>>> uvp_ccd, uvp_clk, uvp_an1, uvp_an2, uvp_an3, uvp_dig = aeu.get_n_cam_uvp()
>>> ovp_ccd, ovp_clk, ovp_an1, ovp_an2, ovp_an3, ovp_dig = aeu.get_n_cam_ovp()
>>> i_ccd, i_clk, i_an1, i_an2, i_an3, i_dig = aeu.get_n_cam_currents()
>>> ocp_ccd, ocp_clk, ocp_an1, ocp_an2, ocp_an3, ocp_dig = aeu.get_n_cam_ocp()
```

For the F-CAM:

```
>>> v_ccd, v_clk, v_an1, v_an2, v_an3, v_dig = aeu.get_f_cam_voltages()
>>> uvp_ccd, uvp_clk, uvp_an1, uvp_an2, uvp_an3, uvp_dig = aeu.get_f_cam_uvp()
>>> ovp_ccd, ovp_clk, ovp_an1, ovp_an2, ovp_an3, ovp_dig = aeu.get_f_cam_ovp()
>>> i_ccd, i_clk, i_an1, i_an2, i_an3, i_dig = aeu.get_f_cam_currents()
>>> ocp_ccd, ocp_clk, ocp_an1, ocp_an2, ocp_an3, ocp_dig = aeu.get_f_cam_ocp()
```



## 11.6. FEE voltage memories

The AEU EGSE provides three memory positions to store default values for the FEE voltages: position A, B, C.

TBD: We store the nominal N-FEE values in memory position A, the nominal values for F-FEE in memory position B. + commands to store voltages, currents, protection values in the memory positions

## 11.7. AEU powering up and down FEE

To power on and off the N- or F-CAM, the following building blocks can be used:

```
>>> aeu.n_cam_swon()
>>> aeu.n_cam_swoff()

>>> aeu.f_cam_swon()
>>> aeu.f_cam_swoff()
```

Note that switching on the camera, will put the AEU in functional check and TVAC mode. Switching off, will put it back to stand-by mode.

## 11.8. AEU configuring synchronisation signals

For the N-CAM, the synchronisation signals must be configured according to the desired image cycle time. Allowed values for the image cycle time are: 25, 31.25, 37.50, 43.75, and 50s. Note that - for image cycle times longer than 25s - not all heater sync pulses are synchronised with a Clk\_ccdread sync pulse.

The following building blocks enable and disable the clock sync signals that are sent to the N-FEE (i.e. Clk\_50MHz, Clk\_ccdread (here with an image cycle time of 25s), and Clk\_heater (synchronised with Clk\_ccdread)):

```
>>> aeu.n_cam_sync_enable(image_cycle_time=25)
>>> aeu.n_cam_sync_disable()
```

For the F\_CAM, the following building blocks enable and disable the clock sync signals that are sent to the F-FEE (i.e. Clk\_50MHz, Clk\_F\_ccdread, and Clk\_heater (synchronised with Clk\_F\_ccdread); nominal clocks only):

```
>>> aeu.f_cam_sync_enable()
>>> aeu.f_cam_sync_disable()
```

To check the sync status of the clocks (i.e. whether or not they are enabled) and whether or not a synchronisation failure has been detected:

For the N-CAM:



```
>>> clk_50mhz, clk_ccdread = aeu.get_n_cam_sync_status()
>>> clk_50mhz, clk_ccdread = aeu.get_n_cam_sync_quality()
```

For the F-CAM:

```
>>> clk_50mhz_nom, clk_50_mhz_red, clk_ccdread_nom, clk_ccdread_red = aeu.get_f_cam_sync_status()
>>> clk_50mhz_nom, clk_50_mhz_red, clk_ccdread_nom, clk_ccdread_red = aeu.get_f_cam_sync_quality()
```

For the SVM/heater:

```
>>> clk_50mhz_nom, clk_50mhz_red, clk_heater_nom, clk_heater_red = aeu.get_svm_sync_status()
>>> clk_50mhz_nom, clk_50mhz_red, clk_heater_nom, clk_heater_red = aeu.get_svm_sync_quality()
```

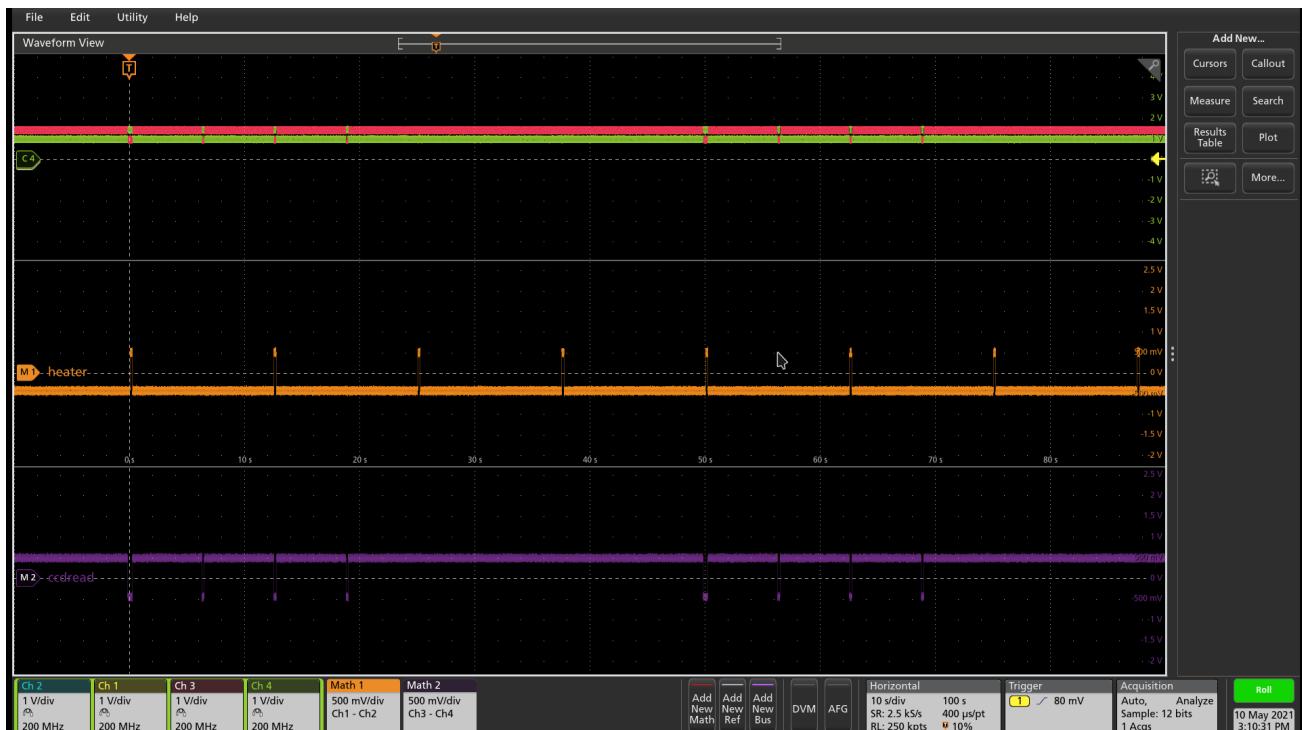


Figure 22. *Clk\_N\_ccdread* (purple) sync pulses and *Clk\_heater* pulses (orange) with image cycle time set to 50s (instead of the nominal 25s). Note that the pulse to read the 1st CCD is wider. The *Clk\_heater* pulses continue with a period of 12.5s, independently of the image cycle time so if the image cycle time is not 25s, the heater will not always be synchronised with a CCD readout.

## 11.9. AEU self test

When the AEU is put in self-test mode (see [Section 11.3](#)), the loopback option can be set as follows:

```
>>> aeu.selftest(LoopBack.NO_LOOPBACK)
>>> aeu.selftest(LoopBack.F_CAM_NOM)
>>> aeu.selftest(LoopBack.F_CAM_RED)
>>> aeu.selftest(LoopBack.N_CAM)
>>> aeu.selftest(LoopBack.SVM_NOM)
>>> aeu.selftest(LoopBack.SVM_RED)
```



## 11.10. AEU Telemetry parameters

```

I_N_V_CCD<NR2>,
I_N_V_CLK<NR2>,
I_N_V_AN1<NR2>,
I_N_V_AN2<NR2>,
I_N_V_AN3<NR2>,
I_N_V_DIG<NR2>,
I_F_V_CCD<NR2>,
I_F_V_CLK<NR2>,
I_F_V_AN1<NR2>,
I_F_V_AN2<NR2>,
I_F_V_AN3<NR2>,
I_F_V_DIG<NR2>,
V_N_V_CCD<NR2>,
V_N_V_CLK<NR2>,
V_N_V_AN1<NR2>,
V_N_V_AN2<NR2>,
V_N_V_AN3<NR2>,
V_N_V_DIG<NR2>,
V_F_V_CCD<NR2>,
V_F_V_CLK<NR2>,
V_F_V_AN1<NR2>,
V_F_V_AN2<NR2>,
V_F_V_AN3<NR2>,
V_F_V_DIG<NR2>,
Standby<boolean>,
Selftest<boolean>,
FC_TVAC<boolean>,
Alignment<boolean>, N-
CAM<boolean>, F-
CAM<boolean>,
V_CCD<boolean>,
V_CLK<boolean>,
V_AN1<boolean>,
V_AN2<boolean>,
V_AN3<boolean>,
V_DIG<boolean>,
S_voltage_oor<boolean>,
S_current_oor<boolean>,
Sync_gf<boolean>,
Clk_50MHZ<boolean>,
Clk_ccdread<boolean>,
Clk_heater<boolean>,
Clk_F_FEE_N<boolean>,
Clk_F_FEE_R<boolean>,
TestPort<boolean>

```

## 11.11. Functional summary

At the start of a test day, the following two AEU building blocks must be executed (either independently or in another building block):



For N-CAM testing:	For F-CAM testing:
>>> aeu.n_cam_swon() >>> aeu.n_cam_sync_enable(image_cycle_time)	>>> aeu.f_cam_swon() >>> aeu.f_cam_sync_enable()

At the end of a test day, the following two AEU building blocks must be executed (either independently or in another building block):

For N-CAM testing:	For F-CAM testing:
>>> aeu.n_cam_sync_disable() >>> aeu.n_cam_swoff()	>>> aeu.f_can_sync_disable() >>> aeu.f_cam_swoff()



# 12. Operating the N-FEE

This section describes some basic principles on the N-FEE design and their practical impact on the N-FEE, CCD and camera operations.

## 12.1. Glossary

- **Parallel:** direction parallel to the columns of the CCD (sometimes referred to as “vertical”).
- **Serial:** direction parallel to the rows of the CCD (sometimes referred to as “horizontal”).
- **Readout register:** single row of pixels below the active region of the CCD, used to transfer the charges in the serial direction towards the readout amplifier.
- **Partial readout:** describes a CCD readout process in which only a given range of CCD rows are digitized. The rows between the region and the readout register are dumped during the readout process, i.e., the recorded signal is not digitized, it is not transferred to the FEE and will not appear in the data.
- **Windowing:** refers to an FEE operating mode in which a pre-defined collection of windows on the CCDs is transferred to the DPU (during CAM-tests: the EGSE). The entire CCDs are readout and digitized, but only the pre-defined windows are transferred. The windowing mode is not used during alignment nor TVAC Camera testing and will not be discussed in this document.
- **Dumping:** a row during readout means to continuously reset the readout register. The charges transferred from the active region of the CCD are lost. The absence of serial transfer makes it much faster to dump a line (90 µs) than to read it out normally (parallel transfer 110µs + serial transfer ~800µs).
- **Clearout:** dumping all or part of a CCD.
- **E and F:** the Plato CCDs dispose of 2 readout amplifiers. The left and right halves of the CCD are transferred to the corresponding amplifier (they behave almost like independent CCDs). For some reason that thy shall not ask, E and F refer to the right and left halves of a CCD (mnemonic: same order as in “FEE”) respectively.
- **Cycle:** period between two long-synchronisation pulses (see below)

## 12.2. N-FEE operating modes

A complete list of the N-FEE operating modes can be found in [RD-06](#). The main modes for camera testing are:

- **ON:** FEE powered on; CCDs powered off. This mode allows transitions to the test-pattern acquisition modes and to the STANDBY mode.
- **STANDBY:** FEE powered on; CCDs powered on. The CCDs are properly biased but not readout, i.e., they are “indefinitely integrating”.
- **FULL\_IMAGE:**
  - allowing for full-CCD acquisition
  - highly configurable mode, also allows partial readout, reverse clocking, charge injection etc.



- This is the workhorse for the ground-based tests.

- **WINDOWING:**

- allowing for the acquisition of multiple windows on the CCDs
- Highly configurable mode, also allows partial readout, reverse clocking, charge injection etc.
- This mode is the baseline for the in-flight operations.

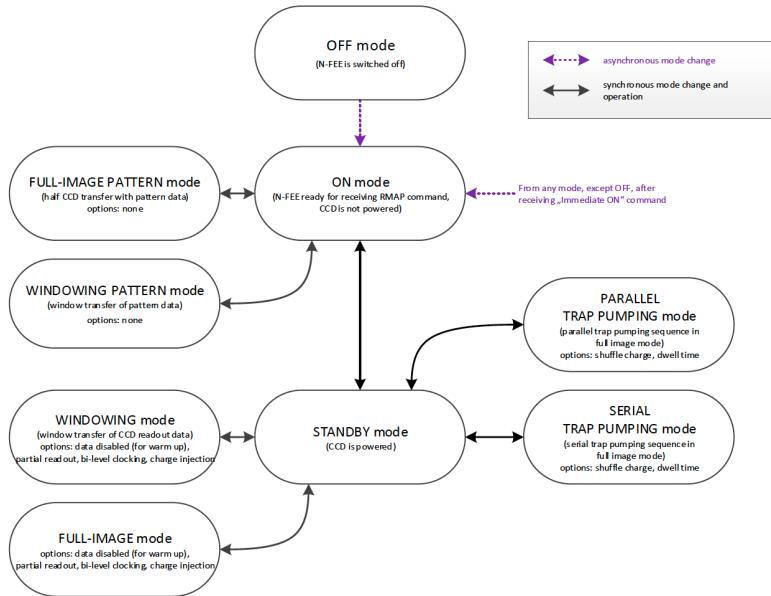


Figure 23. FEE Operating modes (RD-06)

The N-FEE Operating modes are described in more detail in the appendix attached to [RD-14](#). The mode identifiers that you will probably see in Grafana screens or in housekeeping entries, are defined in the N-FEE Register Map and listed here for your convenience.

*Table 3. The N-FEE modes as defined in the Register Map for the PFM [RD-15]. The N-FEE mode is defined by the `ccd_mode_config` parameter in register 21.*

Mode	ID (hex)
On-Mode	0x0
Full Image Pattern Mode	0x1
Windowing-Pattern Mode	0x2
Stand-By-Mode	0x4
Full Image Mode	0x5
Windowing-Mode	0x6
Soft asynchronous Reset	0x7
Immediate On-Mode (Command)	0x8
Parallel trap pumping mode 1 – Full-Image	0x9
Parallel trap pumping mode 2 – Full-Image	0xA
Serial trap pumping mode 1 – Full Image	0xB
Serial trap pumping mode 2 – Full Image	0xC



Mode	ID (hex)
Reserved	0xD
Reserved	0xE
Reserved	0xF

## 12.3. Cycle, Timing and Synchronisation

**Readout timing:** The AEU sends synchronization pulses to the FEE every 6.25 seconds. Every pulse triggers a CCD-readout. In nominal operations, the 4 CCDs in one camera are addressed sequentially, i.e. readout one at a time, delayed by one pulse period, i.e. 6.25 seconds.

**Cycle time and FEE configuration:** all sync-pulses trigger a CCD readout. During nominal operations, every fourth pulse is “long” (it lasts 400ms instead of 200ms). **We define the long-pulse period as the “cycle-time”.** The cycle-time is important in two respects. First, in nominal operations, it takes 4 pulses to cycle over the 4 CCDs, i.e. each CCD is readout every cycle-time seconds. Second, the FEEs, i.e. the operating mode of the CCDs can be reconfigured whenever, but only when the FEE receives a long pulse. The FEE-register (containing the configuration parameters) is read during the pulse and the **new configuration is immediately applied** to the subsequent readouts, i.e. *to integrations that were already on-going*. This is important to keep in mind for the timing of your tests (see the timing examples in Appendix).

**Exposure time:** the PLATO cameras have no shutter. Consequently, the CCDs integration never stops. In practice, the sync-pulses trigger the readout process, and the exposure time effectively corresponds to the cycle-time minus the readout time. That means for instance that for a given cycle-time, the effective exposure time will be longer when performing partial readout than when acquiring full-CCD images.

**Modifying the exposure time:** the exposure time itself cannot be commanded at the level of the FEE. There are nevertheless various ways to modify the exposure time:

- Increase it by changing the cycle time (see building block [n\\_cam\\_partial\\_cycle\\_config](#), [Section 12.4.8](#))
- Shorten it by changing the order in which we address the CCDs in the course of a cycle: e.g. readout the same CCD at every pulse instead of cycling through the 4 CCDs (see parameter [ccd\\_order](#) in [Section 12.4.2](#)).
- Increase it by not addressing a given CCD. If some given CCDs are not addressed for readout, they continue to integrate. The next time they are addressed (after reconfiguring the FEE), their effective exposure time will have been much longer than the nominal exposure time (e.g. for dark current or faint ghosts).
- Disregard the AEU sync pulses and use FEE internal sync-pulses instead. This allows for exposure times shorter than 6.25 seconds (e.g. ambient).

**N-FEE internal sync-pulses:** to accommodate short exposure times, the FEE can generate its own sync-pulses. The source of the sync pulses and period of the internal pulses can be configured with the following EGSE commands to the DPU:



```
>>> n_fee_set_internal_sync(int_sync_period) # in milliseconds
>>> n_fee_set_external_sync()
```

You shouldn't use these commands directly but rather call the dedicated building block:

```
>>> n_cam_partial_int_sync(...)
```

**IMPORTANT**

all N-FEE-generated pulses are long pulses. The cycle-time is hence identical to the CCD readout period, and there is no “natural” cycling through the 4 CCDs. Only when in DUMP mode, the 4 CCDs will be cycled also in internal sync. This is a feature of the DPU Processor implemented in the CGSE and is not available in flight. This CCD cycling in DUMP mode internal sync was implemented as of release 2023.19.0+CGSE.

## 12.4. Commanding the N-FEEs

The following sections describe a collection of building blocks designed to configure and operate the FEEs and the CCDs. A list of the building blocks can also be found on the PLATO Confluence, in the PCOT space, by following links to the “On-Ground Testing”.

Examples of time-sequencing for some operational approaches are presented in [Appendix 17.A](#).

In this section, for the sake of simplicity, the names of the building blocks directly appear at the python prompt (`>>>`), but remember that a commanding building block will exclusively be accepted either within another building block or function, or (hence generating an observation) by the execute command (see [Section 4.1](#)).

### 12.4.1. CCD and pixel references

[Figure 24](#) (for FM) and [Figure 25](#) (for EM) show the CCD numbering adopted for the commanding. We will further refer to these coordinate systems as `CCDn_PIX`  $n = 1,2,3,4$ . Note that these

- each cover an entire CCD, without interruption at the “border” between E & F (columns 2254 to 2255), and
- differ from the CCD coordinate systems adopted in RD-10 (pix [1,1] close to the optical axis), as well as of those adopted at FEE-level (2 coord. systems/CCD, with the x-axes in opposite directions on E & F).

**FM**

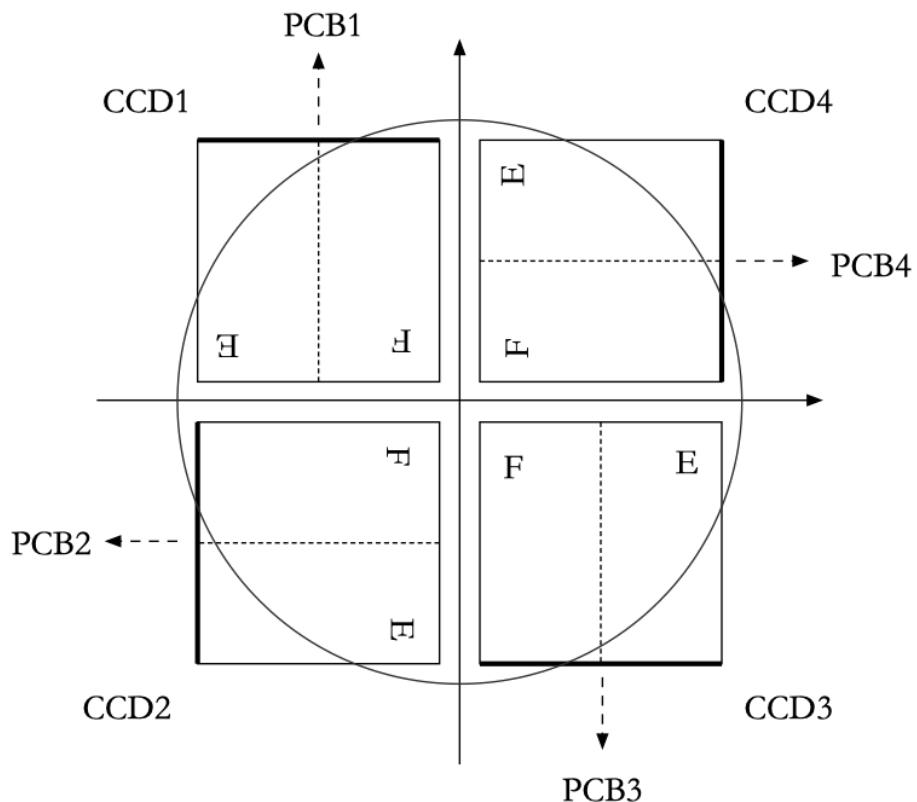
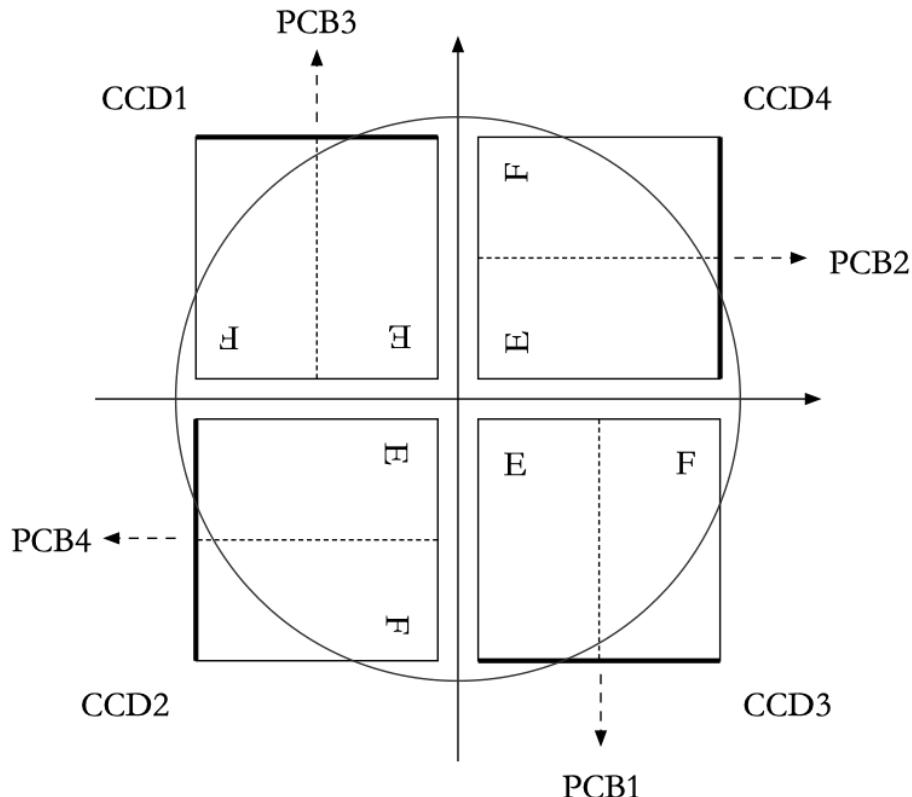


Figure 24. CCD numbering and pixel coordinates for the FM camera on every CCD (CCD\_PIXn reference frames). The thicker black border lines represent the readout registers.

EM



*Figure 25. CCD numbering and pixel coordinates for the EM camera on every CCD (CCD\_PIXn reference frames). The thicker black border lines represent the readout registers.*

### 12.4.2. Standard building block parameters

Some of the input parameters are common to several building blocks. We list some below, to avoid repeating them for every building block.

- **num\_cycles**
  - num\_cycles = 0 sets the FEEs in the required configuration until commanded otherwise
  - num\_cycles > 0 indicates a finite number of cycles after which the N-FEE will automatically be (re)set to dump mode (see `dump_mode` below).
- **row\_start, row\_end**
  - These parameters configure the region of the CCD that will be readout, resulting in partial readout mode.
  - First, the rows < `row_start` are transferred and dumped.
  - Then (`row_end` – `row_start` + 1) rows are readout and digitized.
  - If `rows_final_dump` = 0, nothing else happens
- **rows\_final\_dump**
  - If `rows_final_dump` > 0, after the requested number of lines have been read, `rows_final_dump`



rows to transfer and dumped.

- This allows e.g. for a clearout of the CCD from all dark-current charges accumulated during the readout process before starting a new integration (important at ambient temperature)
- **ccd\_order**
  - During nominal operations, the four CCDs are sequentially addressed during every cycle. This parameter allows to specify and alter that sequence.
  - Examples: [1,2,3,4], [1,3,1,3], [2,2,2,2]

- **ccd\_side**

- This parameter indicates which side(s) of the CCD will be recorded. With the readout register at the bottom, the E-side is the right half and the F-side is the left half.
- In full-image mode, the SpaceWire link to the DPU cannot cope with transferring full frames. Consequently, a choice must be made, either E or F.

The entire CCDs (E and F) is readout and transmitted to the FEE, but only one side is transmitted to the DPU (or EGSE) every cycle. Consequently, it takes a minimum of 2 cycles to obtain full-CCD images, while the exposure time nevertheless still corresponds to one cycle

- This parameter accepts the following values:
  - “E”, “F”, or “BOTH”
  - ~~A string of 4 characters, being either “E” or “F”, e.g. [“EEFF”]~~
  - ~~A string of 8 characters, being either “E” or “F”, e.g. [“EEEEFFFF”]~~
  - ~~If a 4 values are given, the ccd\_side will be changed at every sync pulse, long or short. Four values will hence cover one cycle\_time, but it will take two full cycles to iterate over 8 values.~~
  - “BOTH” means both E and F sides are recorded every cycle. While standard in windowing mode, this is not possible in full-image mode when the camera is connected to an actual DPU, e.g. at integrated system level (spacecraft). The N-FEEs were also neither designed nor extensively tested for this (i.e. full image) by MSSL, but they can do it, and this mode was shortly tested at EGSE level (with a real FEE). It shall be used with caution but remains a possibility to speed up full-CCD image acquisition if needed due to scheduling constraints.

### 12.4.3. N-FEE mode transitions

Two FEE-specific building blocks currently exist to put them into specific operational “modes”:

**ON mode** ([Section 12.2](#)).

```
>>> n_fee_to_on_mode() -- building block
>>> n_fee_is_on_mode() -- function
```

NB: ON mode is the default mode after FEE switch on, but this building block cannot be used to power on the FEE. That is handled by the AEU ([Chapter 11](#)).

**STANDBY mode** ([Section 12.2](#)).



```
>>> n_fee_to_standby_mode() -- building block
>>> n_fee_is_standby_mode() -- function
```

## 12.4.4. DUMP mode

DUMP is not a genuine FEE operation mode. We defined it as a full-image mode in which the dump-gate is maintained high, i.e. the readout register is continuously reset. That is a convenient way to avoid saturation between tests, or building blocks of a given test. In addition, the N-FEE is configured with digitisation disabled which means the image data is not transmitted to the DPU.

### External sync

The CCD operation proceeds over the 4 CCDs with a nominal cycle-time of 25 seconds, but the data are dumped. The default is to read out 0 lines, then do a full-frame clearout (rows\_final\_dump=4510).

```
>>> n_fee_to_dump_mode() # building block
>>> n_fee_is_dump_mode() # function
```

### Internal sync

The CCD operation proceeds over the 4 CCDs. In this mode, we read out 10 lines normally, then perform a full-frame clearout (rows\_final\_dump = 4510). The cycle time is a free parameter, but it must by all means not be chosen shorter than the readout+clearout time. We therefore recommend cycle\_time >= 1 second.

```
>>> n_fee_to_dump_mode_int_sync(cycle_time, ccd_order) # building block
>>> n_fee_is_dump_mode() # function
```

## 12.4.5. N-CAM full-image, basic

Standard full-image acquisition, with a nominal cycle time of 25 seconds, cycling over the 4 CCDs, and 30 rows of over-scan. Only the duration and ccd\_side must be specified. The simplest mode to acquire full, or half-CCD images.

```
>>> n_cam_full_standard(num_cycles, ccd_side)
```

## 12.4.6. N-CAM full-image

Identical to n\_cam\_full\_standard, but allows to configure the ccd\_order and number of over-scan rows as well.

```
>>> n_cam_full_ccd(num_cycles, ccd_order, ccd_side, rows_overscan)
```



## 12.4.7. N-CAM full image, partial readout and final clearout

Identical to n\_cam\_full\_ccd, but allows for partial readout & clearout after readout. The over-scan rows are commanded via the partial-readout parameters: over-scan is only acquired if row\_end > 4509.

```
>>> n_cam_partial_ccd(num_cycles, row_start, row_end, rows_final_dump, ccd_order, ccd_side)
```

## 12.4.8. N-CAM full image, with configurable cycle-time

Identical to n\_cam\_partial\_ccd, including the possibility to configure longer cycle times (from 25 to 50 sec, by steps of 6.25 seconds). The readout process is not affected by the cycle\_time, so the additional time directly corresponds to an increase in exposure time.

```
>>> n_cam_partial_cycle_config (num_cycles, row_start, row_end, rows_final_dump, ccd_order, ccd_side, cycle_time)
```

## 12.4.9. N-CAM full image, with internal sync-pulses

Identical to n\_cam\_partial\_ccd, with configurable exposure time. As explained in [Section 12.3](#), the exposure time cannot be commanded directly at CCD level but results indirectly from the long-pulse period(cycle-time). In this mode, the input parameters are used to compute the duration of a CCD readout, and that in turn is used to compute the cycle time allowing for the desired exposure time.

```
>>> n_cam_partial_int_sync(num_cycles, row_start, row_end, rows_final_dump, ccd_order, ccd_side, exposure_time)
```

In this mode, all sync-pulse are long pulses, i.e. the FEEs can be reconfigured before any readout.

## 12.4.10. N-FEE reverse clocking

Reverse clocking consists in clocking the CCD transfer voltages so that the charges are moved away from the readout register and readout amplifier rather than towards it. It is described in [RD-07](#), and exists in two flavors, depending on the operation of the readout register:

- 1: serial REV
- 2: serial FWD

Both modes provide a reliable measure of the readout noise, but only the second one guarantees a reliable measure of the digital offset. In both cases, the parallel clocks are REV.

It can be operated via the following building block:

```
>>> n_cam_reverse_clocking(num_cycles, clock_dir_serial, ccd_order, ccd_side)
```

**clock\_dir\_serial** must be either "FWD" (standard readout, representative digital offset), or "REV", for reverse clocking in the serial direction as well.



## 12.4.11. Charge injection

Charge injection is described in [RD-08](#). It is envisaged as a means to reduce the negative effects of an increasing CTI towards EOL.

```
>>> n_cam_charge_injection_full(num_cycles, row_start, row_end, rows_final_dump, ccd_order, ccd_side, ci_width,
ci_gap)
```

- `ci_width` expresses the number of rows covered by charge-injection in each block
- `ci_gap` expresses the number of rows between two blocks of charge-injection.

## 12.5. Synchronization with CCD-readouts

It may be beneficial to synchronize some commands with the CCD readouts. For instance small movements of the source on the detector (dithering) may be fast enough to occur entirely during the CCD readout. Synchronizing the movements on the readout hence alleviates the need to lose one image cycle or more to “let things happen”. This can be achieved with the following approach (e.g. in standard mode)

```
from camtest.commanding.dpu import on_long_pulse_do, wait_cycles, n_cam_full_standard

n_cam_full_standard(num_cycles=0, ccd_side=0E0)
for i in range(n_dithers):
    wait_cycles(num_images-1)
    on_long_pulse_do(point_source_to_fov(theta[i],phi[i],wait=False))
```

The first command sets the FEE into an infinite image-acquisition loop, and returns immediately. In the loop, the `wait` command “counts” the right number of cycles (images) before returning. Finally, the command `on_long_pulse_do` will wait until the next long synchronisation pulse and then trigger the embedded command, making sure it starts simultaneously with the CCD readout.

The `on_long_pulse_do` command will only execute the function passed as the argument to `on_long_pulse_do`. If you want to execute several commands, or include a delay: use `wait_cycles(1)` to hold till the system sees the long pulse. The subsequent commands in your script will be executed after the long pulse.

Notes:

- `on_long_pulse_do(command)` triggers the command on the long pulse, i.e. it synchronize it with the readout of one particular CCD (the first one appearing in the parameter `ccd_order`), not all four.
- A mechanical movement + its stabilisation may take a significant amount of time, even for small movements. If this turns out to be longer than the readout time, it will spill over the next integration time, which is undesirable. In order to avoid losing a 25 sec cycle for just a few hundred milliseconds, one can artificially increase the readout duration thanks to the parameter `n_rows_final_dump`, available in all partial readout observing modes ([Section 12.4.2](#)).



# 13. Operating the TCS EGSE

TCS EGSE is providing monitoring and control of the TRP1 and TRP22 temperatures on the camera (TOU and FEE).

TRP (Thermal Reference Points) that are monitored:

- TRP1 (3 sensors, 1nominal/1redundant heater): the TRP controlling the overall TOU temperature, used for thermal focusing
- TRP 22 (1 sensor, 1 nominal / 1 redundant heater): the survival heater on the FEE. This is not used when the FEE is on, it is only used to bring the FEE to its operational temperature if it is too cold to be switched on.

The TCS EGSE has different operating modes:

- **Normal** operating mode of the camera: all the temperatures are acquired and nominally only the TRP1 heater is controlled but TRP22 heater could be added (change of configuration in configuration mode required)
- **Safe** operating mode: all the temperatures are acquired and TRP1 heater is controlled as well as FEE heater
- **Decontamination** mode or de-icing mode: all the temperatures are acquired and nominally only TRP1 heater is controlled but TRP22 heater could be added (change of configuration in configuration mode required)
- Camera **calibration** mode: only thermistor reading is performed and heaters are not activated
- **EMC** mode: replay of a heater control sequence, sensor used only for monitoring and switch off in case of temperature limit is exceeded
- **Extended** mode: all parameters are fully configurable.

The TCS EGSE is changing the TRP1 heater power with PWM (Pulse Width Modulation): the current is varied by changing the width of square full-power pulses. The width of the pulses can only be changed at the time of an AEU clk\_heater sync pulse, between CCD readouts.

The PWM pulses come at frequencies between 30 and 50 Hz - this frequency is configurable.

Clk\_heater sync pulse, to avoid switching the power during a CCD readout. The TCS EGSE is updating the heater power within 200ms after the leading edge (up-stroke) of the Clk\_heater pulse.

The TCS EGSE can be operated in a *local* mode using the man-machine-interface (MMI). During camera tests, for monitoring and commanding the TCS EGSE shall be in *Remote Control* mode.

## 13.1. Switching between operating modes

Switching between modes is done with just one command, set\_operating\_mode. The command takes one parameter that defines the mode.

1. normal



2. safe
3. decontamination | de-icing
4. calibration
5. EMC
6. self-test
7. extended

You can use either the number or the name of the mode as the argument, but it is preferred to use the constants defined in the egse.tcs module.

```
>>> from egse.tcs import OperatingMode

>>> tcs.set_operating_mode(OperatingMode.NORMAL)
>>> tcs.set_operating_mode(OperatingMode.SAFE)
>>> tcs.set_operating_mode(OperatingMode.DECONTAMINATION)
>>> tcs.set_operating_mode(OperatingMode.CALIBRATION)
>>> tcs.set_operating_mode(OperatingMode.EMC)
>>> tcs.set_operating_mode(OperatingMode.SELF_TEST)
>>> tcs.set_operating_mode(OperatingMode.EXTENDED)
```

The mode that will be used most during camera testing is the *extended* mode, which allows all parameters to be configurable.

### 13.1.1. Decontamination

The Decontamination heating is performed by turning ON the TRP1 heaters at full power.

### 13.1.2. Replay

## 13.2. Remote Commanding

The TCS EGSE has a set of commands at your disposal for remote commanding the device. These commands can only be used in *Remote Control* mode, i.e., when the button shows orange in the MMI. We can categorise these commands as follows:

**Change the operating mode:** as seen above, the command for changing the operating mode is `tcs.set_operating_mode(...)` which takes one parameters that defines the mode.

```
>>> tcs.set_operation_mode(OperatingMode.EXTENDED)
```

**Set configuration parameter:** to change configuration parameters, use the command `tcs.set_parameter(...)` which takes two arguments, the parameter name and its value. The following command changes the temperature set point for the PI controller on channel 1 to -40°C.

```
>>> tcs.set_parameter("ch1_tset", -40)
```



A full list of parameters that can be configured is available in table 6.8 on page 61-68 of [RD-09](#).

Please note that several parameters can form a group, e.g., to configure the PI controller for a certain channel. In this case the best approach is to set all the parameters needed for that configuration and switch to the proper mode, i.e., closed loop or open loop.

After your configuration is finished, the parameter set needs to be committed on the device. This is done by sending a `tcs.commit()` command.

**Run the task:** When you have properly configured the TCS EGSE, you need to run the task. This is needed to drive the controllers for the heaters and to execute the PID controller to reach and maintain the intended temperatures.

A task is run with the `tcs.run_task()` command and stopped with `tcs.stop_task()`.

**WARNING**

Only when a task is running, the housekeeping parameters will be updated continuously.

**Retrieve Information:** several commands are available to retrieve information from the TCS EGSE.

- `get_data()`: this command works in both local mode and remote-control mode. It retrieves telemetry from the TCS EGSE. When a task is running, the full housekeeping is available through this command, if no task is running, only load time and system resources are available.
- `get_housekeeping_value(<parameter name>)`: a list of all housekeeping parameters is available in table 6.6 on page 50-58 of [RD-09](#).
- `get_all_housekeeping()`: this command only work in remote-control mode and retrieves all the housekeeping values from the TCS EGSE. Note however, that, when no task is running, the housekeeping will be outdated. To follow up-to-date telemetry with this command, make sure the task is running.
- `get_error()`: this command returns any error that occurred and showed up in the MMI display. Note that there is no way to clear the errors, so the output of this command can be outdated also.
- `get_configuration()`: this command retrieves the full set of configuration parameters with their current value.

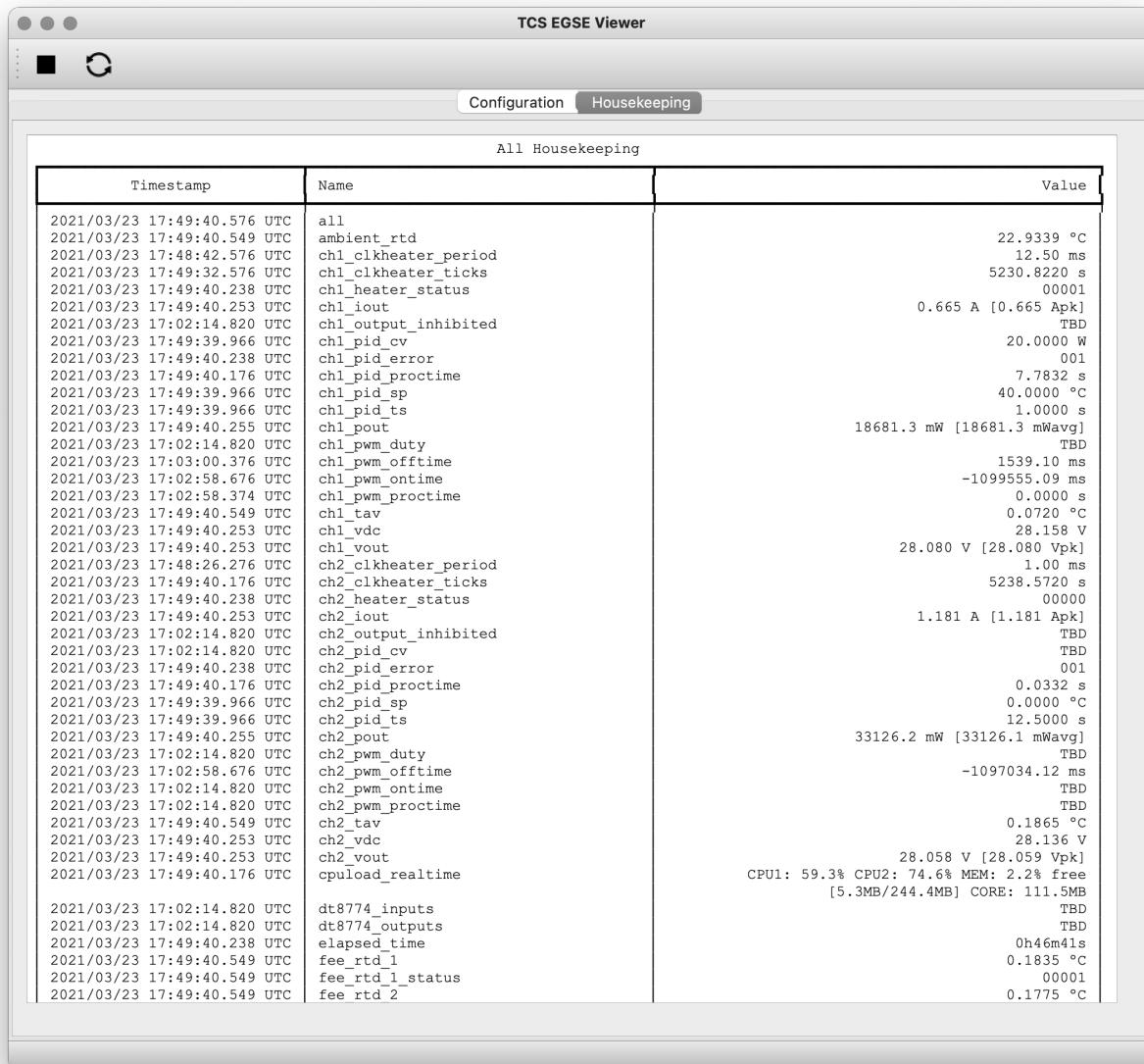


Figure 26. The TCS EGSE Viewer gives an overview of the TCS Housekeeping parameters with their latest engineering values. There are two TABs (1) the Configuration tab which shows the current configuration parameters and their values, and (2) the Housekeeping tab which updates every 10s with the latest housekeeping parameters.

**Convenience Functions:** two convenience functions are available to show information in a nice colour table. If you need a quick overview of the configuration or the housekeeping in the REPL or in a Jupyter Notebook, use the following functions:

```
>>> from egse.tcs.tcs import print_all_housekeeping
>>> from egse.tcs.tcs import print_configuration

>>> print_all_housekeeping()
>>> print_configuration()
```



## 13.3. The TCS Data Acquisition System —DAS

Monitoring TCS EGSE telemetry can best be done with the TCS data acquisition das. The das command is used from the terminal as follows:

```
$ das tcs --interval 10
```

The above command will retrieve TCS housekeeping telemetry every 10 seconds. The data is sent to the Storage Manager that saves the telemetry in a CSV file on the egse-server. The das also makes the temperatures of the TOU, FEE, and the ambient and internal temperatures available for Prometheus to be monitored by Grafana.

Please note that this command should preferably run on the egse-server and be started by the site-operator (see [Section 1.3](#)). The test-operator can inspect the metrics from the TCS EGSE data acquisition in the Grafana display.

## 13.4. Setting the temperature setpoints

TBW

## 13.5. Enabling / disabling temperature control

TBW

## 13.6. Temperature sensor configuration

TBW

### 13.6.1. Disabling TRP1 sensors to verify redundancy function

### 13.6.2. Switching from TRP1 mean to median

## 13.7. Changing temperature sensor calibration curves

TBW

## 13.8. Changing PI control parameters

TBW

## 13.9. Changing the PWM frequency

TBW



# 14. Operating the TEB, shroud and MARI thermal control

## 14.1. Context

The test houses implement the temperature control of 6 Temperature Reference points (TRP) of temperature interfaces to the camera: three temperatures of the Thermal Environment Box (TEB): the TEB\_SKY, a shroud offering a cold radiative cooling sink to the TOU baffle, an upper segment TEB\_TOU, offering the radiative environment seen by the TOU tube (covered in MLI) above the optical bench on the spacecraft, a lower segment TEB\_FEE offering the radiative environment in the optical bench cavity seen by the FEE onboard the spacecraft, and the mounting points of the 3 TOU bipods on the manipulation ring (TRP2,3,4) which are representing the conductive interface of the optical bench as seen by the TOU bipods on the spacecraft.

All six TRPs are controlled independently, and we can check and change the temperature setpoints, and switch the control on / off for each TRP.

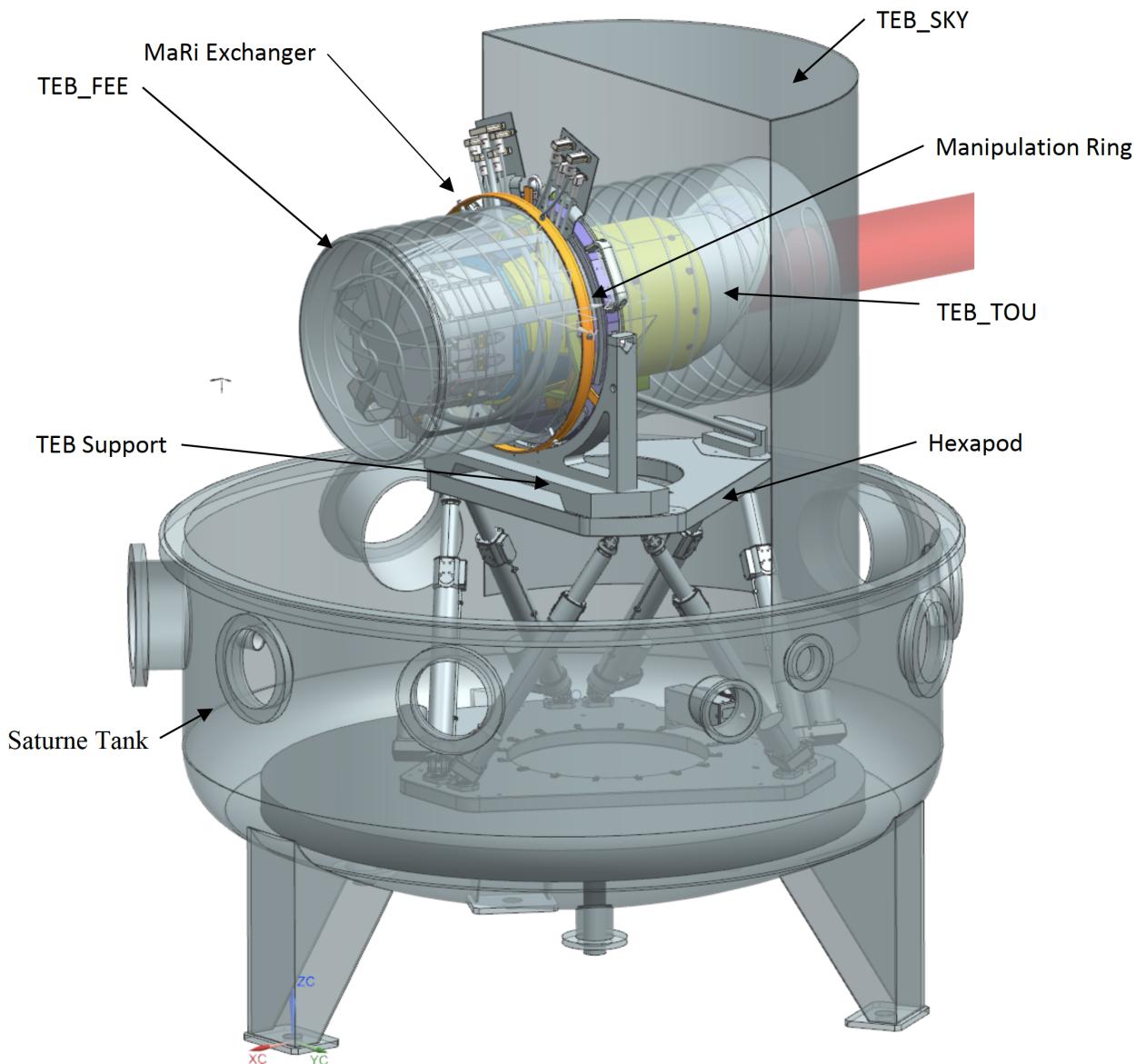


Figure 27. Test setup at IAS, with TEB\_SKY, TEB\_FEE, TEB\_TOU and MaRi indicated

## 14.2. Checking and setting the temperature setpoints

Table 4. Heater patches at the bottom of the MaRi for control of TRP 2, 3,4 (left) and corresponding temperature sensors on the top, right next to the interface holes for the TOU bipods (right)

Heater patches	Temperature Sensors



Changing the setpoints:

```
>>> execute(tgse.set_temp_setpoint(trp=tgse.TRP.TEB_SKY, temperature=<temperatureCelsius>))
>>> execute(tgse.set_temp_setpoint(trp=tgse.TRP.TEB_TOU, temperature=<temperatureCelsius>))
>>> execute(tgse.set_temp_setpoint(trp=tgse.TRP.TEB_FEE, temperature=<temperatureCelsius>))
>>> execute(tgse.set_temp_setpoint(trp=tgse.TRP.TRP2, temperature=<temperatureCelsius>))
>>> execute(tgse.set_temp_setpoint(trp=tgse.TRP.TRP3, temperature=<temperatureCelsius>))
>>> execute(tgse.set_temp_setpoint(trp=tgse.TRP.TRP4, temperature=<temperatureCelsius>))
>>> execute(tgse.set_temp_setpoint(trp=tgse.TRP.TRP234, temperature=<temperatureCelsius>))
```

The effect of the last command (changing TRP234) is the same as setting TRP2, 3, 4 to the same setpoint.

Checking the setpoints:

```
>>> setpoint = tgse.get_temp_setpoint(trp=tgse.TRP.TEB_SKY)
>>> setpoint = tgse.get_temp_setpoint(trp=tgse.TRP.TEB_TOU)
>>> setpoint = tgse.get_temp_setpoint(trp=tgse.TRP.TEB_FEE)
>>> setpoint = tgse.get_temp_setpoint(trp=tgse.TRP.TRP2)
>>> setpoint = tgse.get_temp_setpoint(trp=tgse.TRP.TRP3)
>>> setpoint = tgse.get_temp_setpoint(trp=tgse.TRP.TRP4)
```

## 14.3. Starting / stopping the temperature control loop

```
>>> execute(tgse.start_control(trp=tgse.TRP.TEB_SKY))
>>> execute(tgse.start_control(trp=tgse.TRP.TEB_TOU))
...
>>> execute(tgse.stop_control(trp=tgse.TRP.TEB_SKY))
>>> execute(tgse.stop_control(trp=tgse.TRP.TEB_TOU))
...
```



# 15. Operating the OGSE

The OGSE (optical Ground Support Equipment) consists of a collimator with a pattern mask in the focus. This pattern mask is illuminated by a fiber. The fiber is fed with a laser-driven light source that produces a flat spectrum. Between the light source and the collimator, there is a control box that has:

- Two filter wheels with neutral density filters, allowing to attenuate the light injected into the collimator to the different flux levels required
- A slow asynchronous shutter (not implemented at all test-houses)
- A fast shutter that is synchronised to the AEU synchronisation signal (not implemented at all test-houses)
- A power meter monitoring the light intensity from the light source (power meter channel 1)
- A power meter monitoring the light injected into the collimator (power meter channel 2)

## 15.1. Switching entire OGSE on

### Utility building block

The OGSE can be switched on with one simple command:

```
>>> execute(ogse.ogse_swon)
```

or the command `ogse.ogse_swon()` can be used in a test script.

### Effect

- Switches on the OGSE controllers
- Closes the shutter (if present)
- Puts the filter wheels in a position with 100% attenuation (blank position)
- Switches on the power meter
- Switches on the light source.
- OGSE telemetry is recording.

**NOTE**

the light source needs warming up and stabilisation; the OGSE start up script does not wait for that.

## 15.2. Switching entire OGSE off

### Utility building block

```
>>> execute(ogse.ogse_swoff)
```



## Effect

- Switches off the power meter
- Switches off the light source.
- Switches off the OGSE controllers
- OGSE telemetry is recording.

## 15.3. Attenuation with Neutral density filters

The attenuation is set by a double filter wheel where specific combinations of the wheel positions define a predefined attenuation. The following commands can be used to set the attenuation:

```
>>> ogse.att_level_up()
```

Effect: changes filter wheel combination one attenuation step higher.

```
>>> ogse.att_level_down()
```

Effect: changes filter wheel combination one attenuation step lower.

```
>>> ogse.set_fwc_fraction(<attenuation factor>)
```

Effect: will choose combination of ND filters that matches commanded factor of the full-well capacity as close as possible.

### *Attenuation examples*

- Attenuation factor 1.00 : both filter wheels are in the position with no Neutral density filter, no attenuation.
- Attenuation factor 0.01: filter wheels are in a combination of Neutral density filters delivering approximately 99% attenuation - the signal will be 100 times smaller than with no attenuation.
- Attenuation factor 0 will put a filter wheel in the opaque plate position.

```
>>> ogse.set_fw_position((<wheel_a_pos>, <wheel_b_pos>))
```

Effect: choose manual ND filters in the two wheels, note the argument is a tuple.

You can also read this back through the housekeeping:

```
>>> get_housekeeping("GOGSE_ATT_LVL")
```

Returns the actual attenuation factor realised with the ND filter combination



The CSL ambient alignment collimator provides the following attenuation factors:

Att Selection	Transmittance	Att Selection	Transmittance
0	0,0E+00	19	1,8E-03
1	3,0E-09	20	2,0E-03
2	3,0E-08	21	3,0E-03
3	2,0E-07	22	8,0E-03
4	3,0E-07	23	1,0E-02
5	2,0E-06	24	1,5E-02
6	3,0E-06	25	2,7E-02
7	1,5E-05	26	3,0E-02
8	2,0E-05	27	8,0E-02
9	2,7E-05	28	1,0E-01
10	3,0E-05	29	1,5E-01
11	8,0E-05	30	2,7E-01
12	1,0E-04	31	3,0E-01
13	1,5E-04	32	4,0E-01
14	2,0E-04	33	5,0E-01
15	2,7E-04	34	7,2E-01
16	3,0E-04	35	8,0E-01
17	8,0E-04	36	9,0E-01
18	1,0E-03	37	1,0E+00

Figure 28. CSL ambient alignment collimator attenuation factors (from )

## 15.4. Attenuation specifying the full well fraction

For every facility, a calibration will be done to relate the attenuation factor (0...1) to the fraction of the full well that will be filled in a nominal (25sec) integration near the center of the field.

```
>>> ogse.set_fwc_fraction(0.5)
```

Will set the OGSE attenuation factor that results in the brightest pixel in the PSF near the center of the field to be about 50% of the full well (which is roughly 1E6 for PLATO CCDs)

## 15.5. Switching on/off light intensity stabilisation loop

Only implemented at IAS (TBD)

The N-camera reads out a different CCD every 6.25 seconds. You want to synchronise the shutter open to the start of the new integration on the CCD you are interested in (where the collimator image is seen). Setting the exposure time allows to avoid exposing the CCD during readout (avoiding e.g. smearing) and/or attenuate the light source in finer steps than allowed by the neutral density filter wheels.

```
>>> ogse.shutter_startloop(<ccd number to synchronise to>, <exposure time>)
>>> ogse.shutter_stoploop()
```



## 15.6. Power meter

The readings of the power meters are stored in OGSE housekeeping telemetry.

If you need access to the power, read the housekeeping parameters GOGSE\_PM\_CH1\_PWR (light source monitor) or GOGSE\_PM\_CH2\_PWR (light injected into the collimator)

**NOTE**

CSL collimator power meter 2 will only provide readings between attenuation levels 2E-3 and 1.

## 15.7. OGSE housekeeping parameters

Parameter name	Description	Grafana screen	type	unit
GOGSE_LDLS_INTERLOCK	Laser Driven Light Source Power Interlock on	GOGSE_MON	bool	
GOGSE_LDLS_POWER	Laser Driven Light Source power on	GOGSE_MON	bool	
GOGSE_LDLS_LAMP	Laser Driven Light Source lamp on	GOGSE_MON	bool	
GOGSE_LDLS_LASER	Laser Driven Light Source laser on	GOGSE_MON	bool	
GOGSE_LDLS_LAMP_FAULT	Laser Driven Light Source lamp fault	GOGSE_MON	bool	
GOGSE_LDLS_CTRLFAULT	Laser Driven Light Source controller fault	GOGSE_MON	bool	
GOGSE_LDLS_PSU	Power Supply Unit on	GOGSE_MON	bool	
GOGSE_LDLS_OPERATE	Laser Driven Light Source operate status on	GOGSE_MON	bool	
GOGSE_PM_CH1_PWR	Power meter channel 1 power	GOGSE_MON	Uint16	
GOGSE_PM_CH1_TEMP	Power meter channel 1 temperature	GOGSE_MON	Uint16	DegCelsius
GOGSE_PM_CH1_STATUS	Power meter channel 1 on	GOGSE_MON	bool	
GOGSE_PM_CH2_PWR	Power meter channel 2 power	GOGSE_MON	Uint16	
GOGSE_PM_CH2_TEMP	Power meter channel 2 temperature	GOGSE_MON	Uint16	DegCelsius
GOGSE_PM_CH2_STATUS	Power meter channel 2 on	GOGSE_MON	bool	
GOGSE_ATT_LVL	Attenuation factor (0..1)	GOGSE_MON	Uint16	
GOGSE_ATT_FWELL	Attenuation factor (approximate fraction of Full well)	GOGSE_MON	Uint16	

# 16. Operating the tests, system states

We distinguish between two different entities, corresponding to different timescales for the tests:

- Test phase : consists in one or several days or uninterrupted tests, i.e. without switch off (intentional or not).
- Test : the execution of a single test-script.

We also define the following “system states”:

- INITIALIZED: all subsystems in the test environment and test article (hereafter the ‘system’) are switched on, set to predefined conditions and ready to accept commands.
  - FEEs in “STANDBY” (CCDs powered).
  - OGSE shutter closed
  - No requirement on the OGSE filters (attenuation level unknown).
  - AEU : switched on and syncing
  - TCS: powered on, no task running, configured in remote operational mode, i.e. accepting commands.
  - MGSE mechanisms: controllers on, mechanisms homed if relevant, ready to accept commands

This is the initial state at the start of a test-phase.

- IDLE: All subsystems are in nominal conditions, as for INITIALIZED, but the FEEs are in “DUMP\_MODE” (full\_image mode, nominal clocking, dump gate high), preventing the accumulation of charges between tests.

This state is indeed aimed to serve an “inter-test-known-condition”.

- SAFE: as INITIALIZED, with the FEEs in “ON\_MODE” (CCDs not powered)
- RUNNING: test running.

A standard test procedure will describe all steps to be followed at the start of a test phase, to bring the system from “switched off” to INITIALIZED. It mainly consists in three blocks

- Power on all subsystems (manual hardware switch on)
- Switch on the EGSE components, launch the GUIs and commanding prompt (software switch on)
- Bring the system to the INITIALIZED state.

A dedicated commanding script is provided to operate this latter step:

```
>>> system_to_initialized()
```

You can test if the system is in this state with



```
>>> system_test_if_initialized()
```

This function will abort if the system is not in INITIALIZED state, and do nothing otherwise.

As said above, by convention, every test script should be able to assume that the system is in IDLE state before it starts, and it should return into that state before ending. A dedicated commanding script is provided to this aim as well:

```
>>> system_to_idle()
```

Finally, it shall be possible at any moment to test if the system is in IDLE mode. The command therefore is

```
>>> system_test_if_idle()
```

This function will abort if the system is not in IDLE state, and do nothing otherwise.

The test procedure should explicitly bring back the system to IDLE after every test, but whenever possible, it is nevertheless recommended to start every test with a check via `system_test_if_idle`, and call `system_to_idle` as a last command, to return the system to that known state.



# 17. Appendices

## Appendix A: Examples of CCD acquisition timing sequence

**Commanding:** To keep notations light, the building blocks in the examples below are presented directly at the python prompt, but in operations they will be refuted outside of either another function or building-block, or (if they by themselves constitute the entire test) an execute() command (see [Section 4.1](#)).

**Timing:** In all cases, our time reference starts at the pulse triggering

- a. the commanded configuration
- b. the readout of CCD 1 (integrating previously, but readout under the new configuration)

In most cases, as long as the cycle-time and the ccd\_order remain unchanged by the configuration occurring at  $t = 0$ , the images read out during the first cycle can probably be used directly since, e.g. the just commanded partial readout would be applied. This must be assessed on a case-by-case basis. It will make a difference of one cycle wrt the first moment where a representative set of images corresponding to the new configuration has been recorded. This explains why an apparently redundant cycle is represented in all examples below.

### 17.A.1. 4 CCDs nominal cycle time, acquiring E & F sides simultaneously (full-image or windowing mode)

Examples of building block parametrisation to achieve this:

```
>>> n_cam_full_standard(num_cycles=0, ccd_side="BOTH")
>>> n_cam_full_ccd (num_cycles=0, ccd_order=[1,2,3,4], ccd_side="BOTH", rows_overscan=0)
```

```
start readout CCD1: 0.00 long pulse, FEE config, start cycle 1, readout E & F-sides of previously ongoing exposures
start integr. CCD1: 4.00
start readout CCD2: 6.25 short pulse
start integr. CCD2: 10.25
start readout CCD3: 12.50 short pulse
start integr. CCD3: 16.50
start readout CCD4: 18.75 short pulse
start integr. CCD4: 22.75
start readout CCD1: 25.00 long pulse, start cycle 2, readout *E & F-sides*
start integr. CCD1: 29.00
start readout CCD2: 31.25
start integr. CCD2: 35.25
start readout CCD3: 37.50
start integr. CCD3: 41.50
start readout CCD4: 43.75
start integr. CCD4: 47.75
start readout CCD1: 50.00 long pulse, start cycle 3
start integr. CCD1: 54.00
```



## 17.A.2. Partial readout, external sync, single side (full-image mode)

```
>>> n_cam_partial_ccd (num_cycles=0, row_start=500, row_end=1000, rows_final_dump=4510, ccd_order=[1,1,1,1], ccd_side=EE)
```

For the same of simplicity, the timing below assumes the readout & clearout take exactly one second in total.

```
start readout CCD1_E: 0.00 long pulse, FEE config, start cycle 1
start integr. CCD1: 1.00
start readout CCD1_E: 6.25 short pulse
start integr. CCD1: 7.25
start readout CCD1_E: 12.50 short pulse
start integr. CCD1: 13.50
start readout CCD1_E: 18.75 short pulse
start integr. CCD1: 19.75
start readout CCD1_E: 25.00 long pulse, FEE config, start cycle 2
start integr. CCD1: 26.00
```

## 17.A.3. Partial readout, internal sync, single side (full-image mode)

```
>>> n_cam_partial_ccd_int_sync (num_cycles=0, row_start=500, row_end=1000, rows_final_dump=4510, ccd_order=[1,1,1,1], ccd_side=EE, exposure_time=3)
```

For the same of simplicity, the timing below assumes the readout & clearout take exactly one second in total. In fact they are estimated from row\_start, row\_end and rows\_final\_dump, and the cycle time is then commanded to (readout\_time + exposure\_time)

```
start readout CCD1_E: 0.00 long pulse, FEE config, start cycle 1
start integr. CCD1_E: 1.00
start readout CCD1_E: 4.00 long pulse, start cycle 2
start integr. CCD1_E: 5.00
start readout CCD1_E: 8.00 long pulse, start cycle 3
start integr. CCD1_E: 9.00
```

## Appendix B: Field of view representation with visited positions in CSL

Starting a GUI on the operator screen showing the FOV:

```
$ visited_positions_ui
```

Then you can add the visited locations as follows:

- in focal-plane coordinates (x\_fp, y\_fp) [mm]:

```
>>> from egse.visitedpositions import visit_focal_plane_position +
```



```
>>> visit_focal_plane_position(x_fp, y_fp)
```

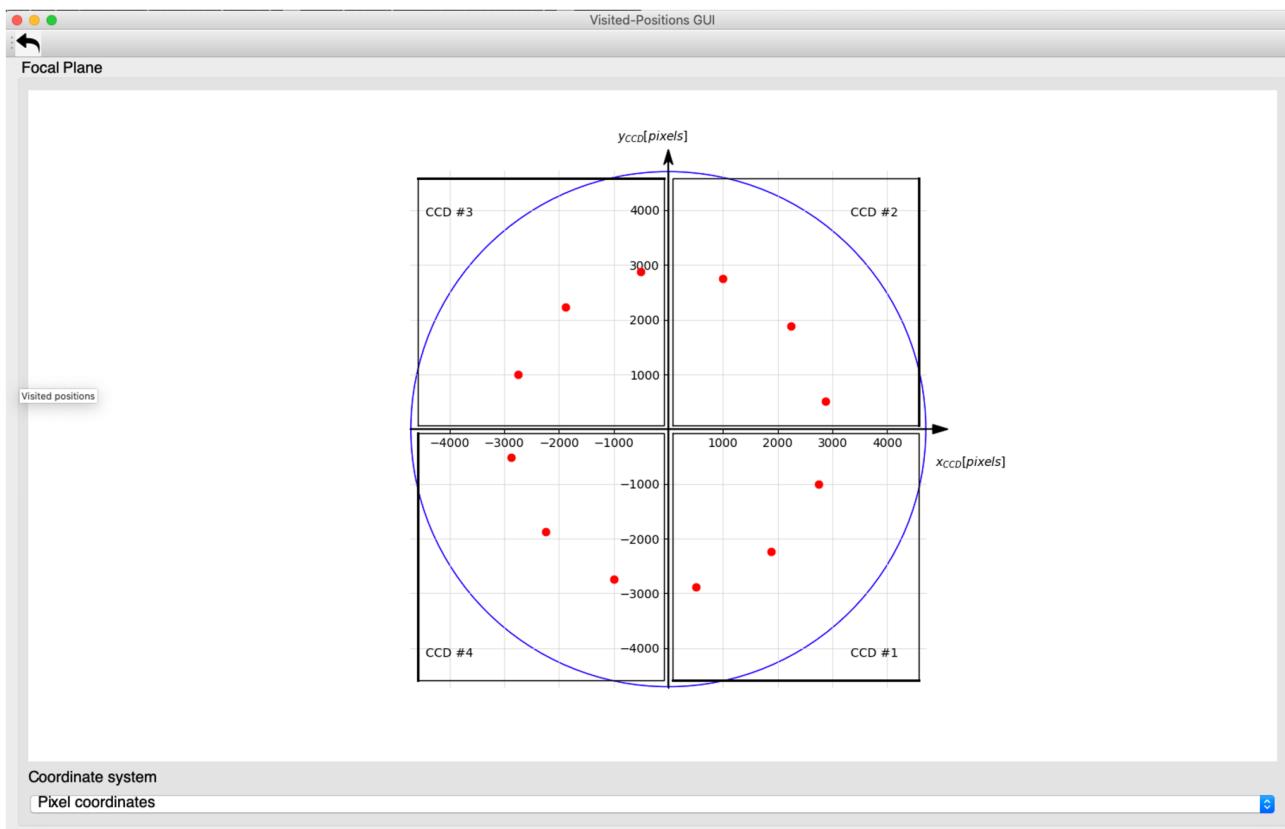
- in CCD coordinates (row, column) [pixel] on a given CCD:

```
>>> from egse.visitedpositions import visit_ccd_position +
>>> visit_ccd_position(row, column, ccd_code)
```

- in field angles (theta, phi) [degrees]:

```
>>> from egse.visitedpositions import visit_field_angles +
>>> visit_field_angles(theta, phi)
```

At each of these locations, a red dot will appear on the plot. You can switch between coordinate system (used in the plot) with the combobox below the plot window (focal-plane coordinates, pixel coordinates, and field angles).



*Figure 29. Visited positions gui. From the commanding script you can launch the gui visible to the operator and add positions (red dots) to visualise the progress of a measurement.*

Alternatively, you can fire up this GUI with

```
$ visited_positions_ui
```

and add the visited locations with the following commands:

- in focal-plane coordinates (x\_fp, y\_fp) [mm]:



```
>>> visit_focal_plane_position(x_fp, y_fp)
```

- in CCD coordinates (row, column) [pixel] on a given CCD:

```
>>> visit_ccd_position(row, column, ccd_code)
```

- in field angles (theta, phi) [degrees]:

```
>>> visit_field_angles(theta, phi)
```

These commands can be used in test scripts, to visualise visited positions, without having to pass on the GUI object (or checking for its existence). In case the GUI has been fired up, the positions will be marked in the GUI. If not, nothing will happen (no error will occur).

In [camtest.commanding.cs1\\_gse](#), there are a couple of building blocks to move the mechanisms (i.e. hexapod and stages) such that the point sources falls on a specified position:

```
>>> point_source_to_fov(theta, phi)
>>> point_source_to_fp(x, y)
```

When executing these building blocks, a red dot will be added to the GUI, marking that position (in case the GUI was fired up).

## Appendix C: What should be started where?

On the EGSE server:

- The core services (Process Manager, Configuration Manager, Storage Manager, and Logging) should be running at all times. Ideally the EGSE server is configured in a way that these services will be re-started automatically in case they would go down.
- If you want to start the Control Servers for the devices on the command line: do it here (but probably you'll want to launch them from the PM UI (see below)).
- The FITS generator.
- The housekeeping generator for the N-FEE.

On the EGSE client:

- All GUIs should be started here (Process Manager UI, Configuration Manager UI, Setup UI, DPU UI, FOV UI, etc.).
- You can start the Control Server for the devices from the PM UI. By doing so, these processes will be started on the EGSE server.
- Executing scripts, commands,...



## Appendix D: Generating FITS files off-line

It sometimes happens that the FITS files need to be re-processed off-line for some observations. This can be done, based on the relevant HDF5 files that are stored in the daily folders for the OD(s) during which the observation took place. Preferably, this is done on the EGSE server of the TH, with the following command:

```
$ fitsgen for-obsid <obsid>
```

This will use the default data location from the `PLATO_DATA_STORAGE_LOCATION` environment variable. It is possible to use a different data location, as follows:

```
$ fitsgen for-obsid <obsid> --location <full path to the data folder (in which obs and daily can be found)>
```

Note that it can be dangerous to run the off-line FITS generation on the client or server during testing, as it could overload the Storage Manager. In case you would do this on a dedicated machine (with the same commands), preferably over VNC (as some observation take a long time to be processed), you need:

- a full installation of plato-common-egse (make sure to use a release);
- a checkout of plato-cgse-conf;
- a running Storage Manager;
- access to the data folder (with sub-folders `/daily` (in which the HDF5 files reside) and `/obs` (in which the FITS files will be stored)) from the machine on which you will launch the FITS generation.

Before you can go ahead, make sure that there are no FITS files in the folder for this observation (remove them if there would be any)!

## Appendix E: GitHub issues, NCR, and PVS

In this section we explain what the difference is between a GitHub issue, a non-conformance report (NCR) and a procedure variation sheet (PVS). All three have their value and life cycle, but its important to understand when and where they have to be used, and how they are tracked.

### 17.E.1. GitHub issues

A GitHub issue is a software issue that is raised in the GitHub repository of that part of the software. Issues on the common software, device drivers, storage etc. should be raised on the `plato-common-egse` repository, while issues with test scripts or, e.g., the Operator Task GUI shall be raised on the `plato-test-scripts` repo. We have also the `plato-cgse-conf` repo which is used for the configuration and calibration of the camera and test equipment, and where the Setups are stored/tracked. Finally, there is the `plato-cgse-doc` repo which contains the documentation suite (of which you are reading one part right now).



An issue doesn't necessarily mean there is a problem or an error. GitHub issues can be raised for various reasons: feature requests, software improvements, documentation update, bug, clarification request, system crash, dependency problem, etc. Whenever you want to raise an issue, use the button at the top of the GitHub repo page and fill in a clear title and a proper description. Some issues have a template that help you provide all information that a developer would need to analyse and reproduce the problem or to proceed with an implementation.

at the top-left of the repo Project page, click the **Issues** button to navigate to the overview page of open issues...

then at the top-right of the Issues page, click the **New issue** button to create a new issue...

When you raise an issue, make sure you associate as many labels to it as applicable. Labels identify the issue and provide a powerful way to filter and search issues. If an issue is urgent and needs to go into a release, attach it to the milestone for that release.

Anyone who has access to the repository can inspect and create issues. If you encounter a problem, contact the maintainer of the GitHub repos.

- Examples:

- a process doesn't start: processes are all part of the CGSE and you will notify when their LED is red in the process manager GUI. Try to investigate, but if you can not solve it, raise an issue in plato-common-egse.
- an ERROR or Exception is raised in the console area of the operator Task GUI. Raise an issue in the test scripts and copy the error message from the console area into the issue description. If needed, the GitHub maintainer will migrate the issue to the CGSE repo.
- the test script is aborted: This can be due to incorrect input values, or a system component that is not connected etc. Most of the time such an issue can be solved, but if the problem remains, raise an issue on the test scripts.
- data is missing: This can be HK, metrics, FITS or CSV files, etc. This is usually a problem on the CGSE since all data files are handled by the Storage component. Only reduction and analysis scripts that are started from the Operator Task GUI generate intermediate results or PNGs.

GitHub issues is an issue tracking system, not a software configuration control system.

How is the software configuration tracked? Where is that described?

- milestones → CCB → releases

## 17.E.2. NCR

A NCR is a *Non-Conformance Report* and this type of report is used when there is a problem with either the Camera or one of the GSE that are under the responsibility of an ESA contract. As a consequence, NCRs are tracked and handled by ESA. An NCR can however be a result of an analysis on an issue that was previously raised in GitHub. When that is the case, add a link to or the ID of the



GitHub issue to the NCR description. You should also add the link to the NCR (or the ID) to the GitHub issue.

The NCRs are tracked from the [Eclipse system](#) used at ESA. Select the **PLATO PA** scope and then **eNCTS**:

The screenshot shows a software interface for managing Non-Conformities (NCRs). At the top, there's a navigation bar with the PLATO PA scope selected. Below it, the eNCTS tab is active. The main area displays a list of NCRs, with a search bar and various filtering and sorting options available.

This page is the starting point for creating and viewing all NCRs.

### Examples

- the N-FEE reports errors on the SRAM Buffer Full continuously, we could not identify a cause in the CGSE or site hardware/network setup.
- CCD 2 of FM2 is not working properly anymore after Camera vibration. (PLATO-CSL-PL-NCR-0032)
- after a camera switch ON, we measure a background of 650 ADU instead of 1300 ADU (PLATO-IAS-PL-NCR-0029)

### 17.E.3. PVS

A PVS is a *Procedure Variation Sheet* and is used to describe any changes or deviations from a test procedure like TP-11 or TP-12. We make a distinction between changes in the procedure due to (1) an error in the procedure, lacking or incomplete information and (2) a small deviation or command that is used which is basically not part of the procedure. The former changes should actually go into a PVS since they introduce a change in the procedure itself and need to be tracked in a proper way. The latter change is usually done by red-lining the as-run copy of the procedure. This way, when the steps that were executed are checked, it is clear some extra commands have been executed or some commands have been skipped. That is not necessarily a permanent change in the procedure.

The PVS is a sheet, so where does this reside? Is the PVS attached to the procedure itself, to the as-run, or is it a separate document?