





A frequency domain approach to datadriven control





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Abstract

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Keywords: data-driven controller tuning; model reference control; Frequency domain approach; Nonparametric; Frequency response function

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List of Abbreviations

CT Continuous-time

DFT Discrete Fourier transform

DT Discrete-time FD Frequency domain

FRF Frequency response function IDFT Inverse discrete Fourier transform

LPM Local Polynomial Method LTI Linear time-invariant MSE Mean squared error NSR Noise-to-signal ratio RMS Root mean square SNR Signal-to-noise ratio

TD Time domain

WNLS Weighted nonlinear least squares

ZOH Zero-order hold

Chapter 1

Introduction

Systems that react to stimuli are all around us. However, many systems don't behave as we want them to. Control theory is a framework that allows engineers to find ways to manipulate these stimuli in order to get these systems to behave as they want them to. Traditionally, the first step to controlling a system is to model the underlying system. If the system is a linear time-invariant (LTI) system, it can be modelled by a transfer function or state space equations. This transfer function or state space model is a parametric representation of the system. When a parametric model of the system is obtained, a controller can be designed for it.

A model of an LTI system can be obtained by applying a well-designed input to the system and measuring the corresponding output.

Data-driven control is an attempt to skip the modelling part of this process. Instead of going from data to a parametric model to a controller, a controller is found directly from data. This is illustrated in figure 1.1.

One such data-driven approach is model reference control [1]. The idea is to define a reference model and to tune a controller such that the closed loop system is as close as possible to the reference model. However, as will be shown in chapter 3, the name "data-driven" can be a bit misleading. In fact, it will be shown that a nonparametric model is hidden in the maths. A nonparametric model differs from a parametric model in the sense that the parametric model tries to describe the data using a relatively small amount of parameters. A nonparametric model still has parameters, but the amount of parameters can be equal to the number of data points. Concretely, applied to LTI systems, a parametric model could be a transfer function where the parameters are the coefficients of the numerator and denominator. A nonparametric model of an LTI system could be the frequency response function at different frequencies.

Chapter 2 gives an introduction to nonparametric models of LTI systems. Chapter 3 explains model reference control and shows that there is an underlying nonparametric model in the maths. Finally, chapter 4 shows how stability can be guaranteed in model reference control in the absence of a parametric model.

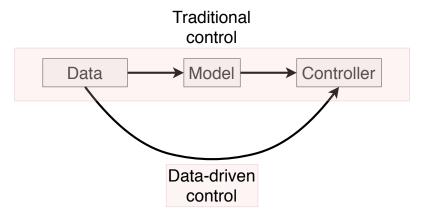


Figure 1.1 – Traditional control vs. data-driven control.

Chapter 2

Preliminaries

2.1 Introduction

A refresher about linear time-invariant systems is given here. Afterwards, this section focuses on frequency domain (FD) methods for estimating nonparametrically the frequency response function (FRF). White noise and system transients are analysed in the FD. It is important to understand how white noise manifests itself in the frequency domain as it will increase the variability of the nonparametric estimate of the FRF. System transients not only increase the variability of the FRF estimates but also introduce a bias error. Therefore, it is also important to study its properties so that it can be accounted for during nonparametric estimation of the FRF. Finally, this chapter ends with some advanced methods to suppress transients and noise in FRF estimation.

2.2 LTI systems

Linear time-invariant (LTI) systems can be represented in different ways. In this work, we will use the transfer function representation for single-input single-output (SISO) systems.

$$G(\Omega) = \frac{B(\Omega)}{A(\Omega)}$$

with

$$\Omega = \begin{cases} s & \text{if working in continuous-time (CT)} \\ z^{-1} & \text{if working in discrete-time (DT)} \end{cases}$$

and $B(\Omega)$ and $A(\Omega)$ being polynomials of Ω . For now, we will keep our focus on CT systems. In CT, the output of this system is given by

$$y(t) = \mathcal{F}^{-1}\{Y(j\omega)\} = \mathcal{F}^{-1}\{G(j\omega)U(j\omega)\} = \mathcal{F}^{-1}\{G(j\omega)\mathcal{F}\{u(t)\}\}$$

With \mathcal{F} and \mathcal{F}^{-1} denoting the Fourier transform and inverse Fourier transform respectively.

$$\mathcal{F}\{x(t)\}\}(j\omega) = \int_{-\infty}^{+\infty} x(t)e^{-j\omega t}dt$$

$$\mathcal{F}^{-1}\{X(j\omega)\}(t) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} X(j\omega)e^{j\omega t}d\omega$$

Because a convolution in the time domain (TD) becomes a multiplication in the frequency domain, the output in the FD is simply found my performing a multiplication.

$$Y(j\omega) = G(j\omega)U(j\omega)$$
(2.1)

2.3 Single sine excitations

Assume that the input u(t) of the system is a complex exponential with a single frequency.

$$u(t) = e^{j(\omega_u t + \phi)}$$

In the FD this becomes

$$U(j\omega) = 2\pi e^{j\phi} \delta(\omega - \omega_u)$$

With δ denoting the Dirac delta function. Using (2.1) gives the output in the FD is

$$Y(j\omega) = 2\pi |G(j\omega_u)| e^{j(\phi + \angle G(j\omega_u))} \delta(\omega - \omega_u)$$

Transforming back to the TD we find

$$y(t) = |G(j\omega_u)|e^{j\angle G(j\omega_u)}e^{j(\omega_u t + \phi)} = |G(j\omega_u)|e^{j\angle G(j\omega_u)}u(t)$$

Since the system is linear and real, this result can be used to calculate the output of the system if the input is a cosine wave.

$$u(t) = \cos(\omega_u t + \phi) = \frac{e^{j(\omega_u t + \phi)} + e^{-j(\omega_u t + \phi)}}{2}$$

After a brief calculation, the output is found to be given by

$$y(t) = |G(j\omega_u)| \cos(\omega_u t + \angle G(j\omega_u) + \phi)$$

To summarize, if the input of a system is a sine wave, then the output will be a sine wave with the same frequency, but with a different phase and amplitude determined by the value of the transfer function at that frequency.

2.4 Discrete Fourier transform

Using the Fourier transform is not practical as it requires an infinite amount of data and an infinite time resolution to compute. The discrete Fourier transform (DFT) solves both of these problems.

The DFT of a sequence x(n), n = 0, 1, ..., N-1 is defined as

$$X(k) = \mathrm{DFT}\{x(n)\} = \frac{1}{N} \sum_{n=0}^{N-1} x(n) e^{-j2\pi k n/N} \,, \quad k \in \mathbb{Z}$$

X(k) is periodic with period N, so k is usually confined to k = 0, 1, ..., N-1. Usually, the DFT is defined without the factor $\frac{1}{N}$. However, this will simplify the results in this work. The inverse DFT (IDFT) is defined in a similar way.

$$x(n) = \text{IDFT}\{X(k)\} = \sum_{k=0}^{N-1} X(k)e^{j2\pi kn/N}, \quad n \in \mathbb{Z}$$

2.5 Perfect reconstruction

The question now is: under which conditions can we perfectly reconstruct a CT signal x(t) from a DT measurement?

$$x_d(n) = x(nT_s), \quad n = 0, \dots, N-1$$

It is simpler to see when this is the case by working in the FD. The question then becomes: under which conditions does the DFT of $x_d(n)$ contain the same information as the Fourier transform of x(t)?

Periodicity First, as the DFT has a limited spectral resolution, the Fourier spectrum of the continuous signal must also be discrete. This is the case when the CT signal is periodic. If the CT signal has a period T

$$x(t) = x(t+T) \quad \forall t$$

then the Fourier spectrum of x(t) will consist of Dirac delta functions with a fixed spacing between the Dirac pulses.

$$X(\omega) = \mathcal{F}\{x(t)\} = \sum_{n=-\infty}^{+\infty} c_n \delta(\omega - n\omega_0), \quad \omega_0 = \frac{2\pi}{T}$$

Leakage Next, the DFT frequencies should coincide with the Dirac pulses to avoid leakage. The DFT bins k correspond to certain frequencies depending on the sampling frequency f_s and the number of samples N.

$$\omega_k = 2\pi k \frac{f_s}{N} = 2\pi k \frac{1}{NT_s} = 2\pi k \frac{1}{T_{\rm meas}}$$

 $T_{\rm meas} = NT_s$ is the measurement time. The lowest frequency of the CT signal $\omega_{\rm signal}$ needs to correspond to one of the DFT frequencies in order to avoid leakage.

$$\omega_{\text{signal}} = \omega_k \Rightarrow \boxed{T_{\text{meas}} = kT_{\text{signal}}}$$

In simple words: the measurement time must contain an integer number of periods of the CT signal.

Aliasing Finally, the famous Nyquist-Shannon sampling theorem states that the CT signal x(t) must not contain frequency components higher than half the sampling frequency.

$$f_{\rm max} < \frac{f_s}{2}$$

To summarize, a CT signal is perfectly reconstructable from a sampled version of it in a limited time window if the DFT contains the same information as the CT Fourier transform. This is the case when the following conditions are met:

- The CT signal is periodic.
- The measurement time contains an integer number of periods.
- The bandwidth of the signal does not exceed half the sampling frequency.

2.6 Frequency response function estimation

If the input and output of an LTI system satisfy the conditions of perfect reconstructability, then the transfer function can be calculated at the excited frequencies. Taking the input to be a periodic single sine as before results in a sine in the output with the same frequency.

$$u(t) = \cos(\omega_u t + \phi) \Rightarrow y(t) = |G(j\omega_u)| \cos(\omega_u t + \angle G(j\omega_u) + \phi)$$

In the FD this becomes

$$Y(j\omega_u) = G(j\omega_u)U(j\omega_u)$$

Assuming that the excited frequency corresponds to one of the DFT frequencies $\omega_u = \omega_k = 2\pi k f_s/N$ and $|\omega_k| < \pi f_s$ fulfils the last 2 conditions of perfect reconstructability respectively.

$$U(j\omega_k) = 2\pi U(k)\delta(\omega - \omega_k), \quad Y(j\omega_k) = 2\pi Y(k)\delta(\omega - \omega_k)$$

With U(k) and Y(k) being the k-th DFT bin of $u_d(n)$ and $y_d(n)$ respectively. Note that the factor 2π is just a consequence of how the Fourier transform and the DFT are defined. In the end, we are interested in ratios, so this won't matter. This means that

$$Y(k) = G(j\omega_k)U(k) \tag{2.2}$$

Finally, the value of the frequency response function at $s = j\omega_k$ can be calculated as

$$G(j\omega_k) = \frac{Y(k)}{U(k)}$$

2.7 Multisine excitations

Using a single sine excitations allows one to calculate the value of the FRF at one single frequency. We can make use of the linearity of LTI systems to calculate the FRF at multiple frequencies at once. This is where a multisine excitation can be useful.

$$u(t) = \sum_{k \in K_{avc}} A_k \sin(\omega_k t + \phi_k), \quad \omega_k = 2\pi k \frac{f_s}{N}, \quad K_{exc} \subseteq \{n \in \mathbb{Z} | 0 \le n < N/2\}$$

In the FD this gives

$$U(k) = \frac{A_k}{2j} e^{j\phi_k}, \quad k \in \mathcal{K}_{\text{exc}}$$

By using (2.2) the output is

$$Y(k) = G(j\omega_k) \frac{A_k}{2j} e^{j\phi_k}, \quad k \in K_{\text{exc}}$$

Thus, the FRF can be calculated for all $k \in K_{exc}$.

$$G(j\omega_k) = \frac{Y(k)}{U(k)}, \quad k \in K_{\text{exc}}$$
 (2.3)

Magnitude choice The magnitude of the sine components A_k can be chosen by the user. This can be useful when noise occurs in a certain frequency band. More power can be attributed to this frequency band to get a better signal-to-noise ratio at those frequencies.

Phase choice The phases of the sine components ϕ_k can also be chosen by the user. A concise example is given to see what the consequences are of choosing a different phase. Two possibilities are shown here.

A linear phase multisine is the simplest case: all the sine components have the same phase

$$\phi_k = 0$$

In a random phase multisine the phases of the sine components have random phases drawn from a uniform distribution between 0 and 2π .

$$\phi_k \sim U(0, 2\pi)$$

Two multisines with the same RMS values are shown in figure 2.1. One has a linear phase, while the other has a random phase. The magnitude spectrum of both are the same. However, the linear phase multisine peaks at t=0, while the power of the random phase multisine is more evenly distributed over time. This can be useful when the generator has a limited range due to saturation. Even though both signals contain the same power, the random phase multisine is less likely to saturate.

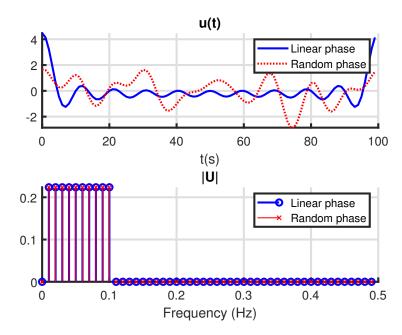


Figure 2.1 – Comparison between a linear phase multisine and a random phase multisine with the same RMS value.

Example Consider a DT system.

$$G(z^{-1}) = \frac{0.4097z^{-1} + 0.407z^{-2}}{1 - 1.165z^{-1} + 0.9813z^{-2}}$$
(2.4)

This system is excited by the random phase multisine shown in figure 2.2. The normalized frequency is the frequency divided by the sampling frequency. 0.5 represents the Nyquist frequency $f_s/2$. The steady state output of this system is shown in figure 2.3. Then the FRF can be calculated by using (2.3) at the excited frequencies. The actual FRF and calculated FRF are plotted in figure 2.4.

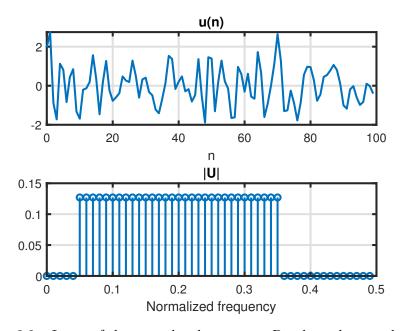


Figure 2.2 – Input of the second-order system. Random phase multisine.

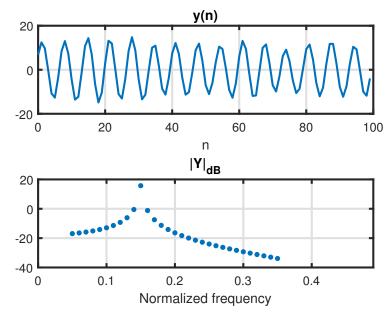


Figure 2.3 – Output of the second-order system.

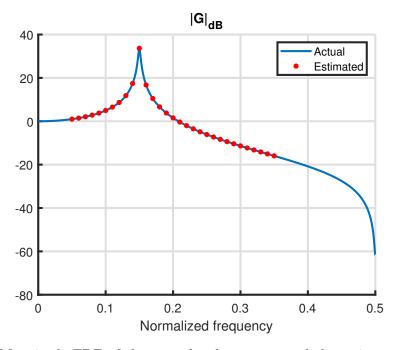


Figure 2.4 – Magnitude FRF of the second-order system and the estimated FRF at the excited frequencies using (2.3).

2.8 Measurement set-ups

2.8.1 Zero-order hold set-up

In a lot of measurement set-ups, the input signal is generated digitally with a Zero-order hold (ZOH). This means that the signal is kept constant for a whole sampling period. At the sampling instants, a CT system G(s) preceded by a ZOH can be modelled exactly as a DT system. [3, eq. (34)]

$$G_{\text{ZOH}}(z) = (1 - z^{-1}) \mathcal{Z} \left\{ \mathcal{L}^{-1} \left\{ \frac{G(s)}{s} \right\} |_{t=nT_s} \right\}$$
 (2.5)

The ZOH measurement set-up is shown in figure 2.5. In this case, the actuator $G_{\text{act}}(s)$ is part of the FRF that is measured. The dynamics of the measurement device $G_y(s)$ must be calibrated perfectly $(G_y(s) = 1)$ if one wants to measure the FRF from generator to output. Note that, in theory, the ZOH set-up is not allowed to contain anti-alias filters.

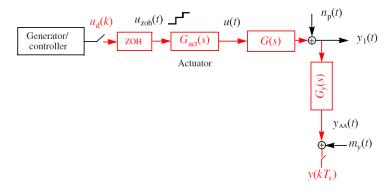


Figure 2.5 – Zero-order hold measurement set-up. Red indicates which parts are modelled. Taken from [2].

Thus, if no anti-alias filter is used at the output and assuming that the actuator is perfect $G_{\text{act}}(s) = 1$, when a CT system is excited with a ZOH and is sampled, we are actually measuring the ZOH version of the FRF and not the CT FRF directly. This is bad news if we want to measure the CT FRF G(s). However, this can be circumvented. When $|f| \ll f_s/2$, there isn't a big difference between the CT FRF G(s) and the ZOH version of G(s).

$$G(j2\pi f) \approx G_{\rm ZOH}(e^{j2\pi f T_s})$$
 (2.6)

So, when the goal is to measure a CT FRF with a ZOH set-up, care must be taken to stay within the region where this approximation holds. This can be accomplished by making sure that the highest excited frequency in the input signal is well below the sampling frequency of the measurement set-up.

$$f_{\rm max} \ll f_s/2$$

Example Consider a second order system.

$$G(s) = \frac{\omega_0^2}{s^2 + 2\zeta\omega_0 s + \omega_0^2}$$
 with $\omega_0 = 2\pi 0.3 [\mathrm{rad}/s]$ and $\zeta = 0.01$

Applying the ZOH transformation (2.5) to this system with $f_s = 2$ Hz results in the DT system that was used in previous examples (2.4). The magnitude FRF of both the CT system and the ZOH version of it are plotted in figure 2.6. Notice that the DT system has a periodic FRF while the CT system does not. Moreover, at the low frequencies, both FRFs overlap. The approximation (2.6) is worse once the frequency gets close to $f_s/2$.

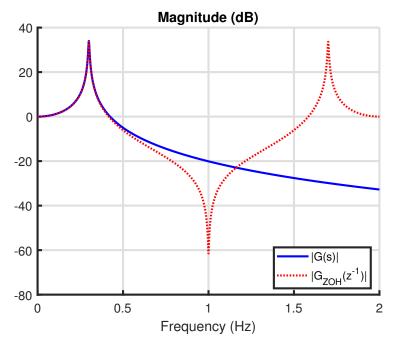


Figure 2.6 – Magnitude FRF of the second-order system and the magnitude FRF of the ZOH version of it with $f_s = 2$ Hz.

2.8.2 Band-limited set-up

Even if one uses a high sampling frequency f_s , the actuator dynamics will still influence the measurement of the FRF when using a ZOH set-up. This is not an issue when a band-limited set-up is used. The band-limited set-up is shown in figure 2.7. In a bandlimited set-up, the input to the system can still be generated by a ZOH. But the input to the system must be measured and an anti-alias filter must be used to avoid aliasing. By measuring the input to the system, the dynamics of the actuator won't be part of the FRF estimate. The output of the system must also go through an anti-alias filter before being measured. A relative calibration is needed if one wants to measure the FRF from the input to the output. Concretely this means that the ratio of the anti-alias filters $G_y(s)/G_u(s)$ must be taken into account when processing the measurements.

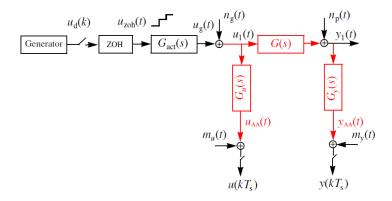


Figure 2.7 – Band-limited measurement set-up. Red indicates which parts are modelled. Taken from [2].

2.9 System transients

The first condition of perfect reconstructability (periodicity) implies that the signal has been repeating forever until now and will repeat forever in the future. This is not realistic; when a system is excited, the excitation must have started at some time in the past and must end at some point in the future. In the example of the previous section we had briefly mentioned that the measurements were taken in steady state. Concretely, the input was applied to the system for 20 periods and only the last period was used for the analysis, thereby ensuring that the system transients have faded away. This can be quantified by calculating the RMS of the difference between the output of the p-th period and the output of the last period. This is plotted in figure 2.8. The RMS of the difference decreases exponentially. If the measurements were noisy, the RMS of the difference would decrease until the transients are below noise level, at which point the transients can be said to have faded away. In this case, the RMS of the difference reaches approximately 10^{-7} by the 20-th period, which is negligible for our purposes.

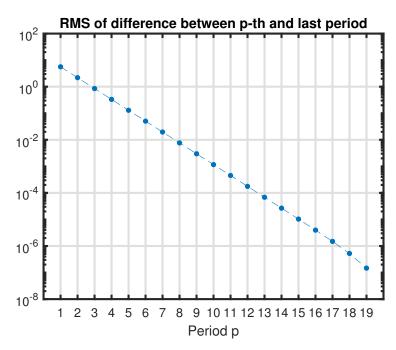


Figure 2.8 - RMS of the difference between the output of every period and the output of the last period.

It turns out that the transient term is just a rational function added on top of the steady state output spectrum.

$$Y(k) = G(e^{-j2\pi k/N})U(k) + T(k)$$
 (2.7)

with

$$T(k) = \frac{I(e^{-j2\pi k/N})}{A(e^{-j2\pi k/N})}$$

This is proven for DT systems in appendix 2.A. The transient term for CT systems will be discussed afterwards.

If u(n) and y(n) are periodic, then T(k) = 0, because $I(e^{-j2\pi k/N})$ only depends on the difference between the in- and output at the end of the current period and the in- and output at the end of the previous period. In this case (2.7) simplifies to what we had before.

$$Y(k) = G(e^{-j2\pi k/N})U(k)$$

A key property of the transient term T(k) is that it is a "smooth" function of the frequency k. This is due to the fact that it is a rational form of $e^{-j2\pi k/N}$. This property can be used to suppress the transient, as will be explained later. Another property of T(k) is that its denominator is the same as the denominator of $G(e^{-j2\pi k/N})$. Therefore, the transient term will "resemble" the shape of the FRF.

CT transients Similar results can be derived for CT systems.

$$Y(k) = G(j\omega)U(k) + T(k) + \delta(k)$$

In this case, the numerator of T(k) depends on the difference in initial conditions at t = 0 and $t = T = nT_s$:

$$\left[\frac{d^py}{dt^p}(T)-\frac{d^py}{dt^p}(0)\right] \text{ and } \left[\frac{d^pu}{dt^p}(T)-\frac{d^pu}{dt^p}(0)\right]$$

with $p \in \mathbb{N}$ and $p < \max(n_a, n_b)$, n_a and n_b being the order of the denominator and numerator of G(s) respectively. An additional term $\delta(k)$ pops up. This is the alias error and it can be generated when the transient $T(j\omega)$ overextends into the aliasing frequencies $f > f_s/2$. It is present even if the signals have been low-pass filtered. Because the alias error $\delta(k)$ is also smooth, it can be grouped together with the transient term T(k).

Example Again, the same second-order system (2.4) of the previous section is used. This time I'll take a look at the estimation error.

$$\begin{split} G_{\rm est}(e^{-j2\pi k/N}) - G(e^{-j2\pi k/N}) &= \frac{Y(k)}{U(k)} - G(e^{-j2\pi k/N}) \\ &= \frac{G(e^{-j2\pi k/N})U(k) + T(k)}{U(k)} - G(e^{-j2\pi k/N}) \\ &= \frac{T(k)}{U(k)} \end{split}$$

As the magnitude spectrum of u(n) is flat, the magnitude of the estimation error will be proportional to the magnitude of the transient term.

The estimation error for the FRF estimated from the first and 20th period is shown in figure 2.9. The error of the FRF calculated from the last period is around -150dB, which is negligible. In other words, there is no transient term as the data is in steady state. The error for the first period is a smooth function of the frequency and "resembles" the shape of the transfer function as expected.

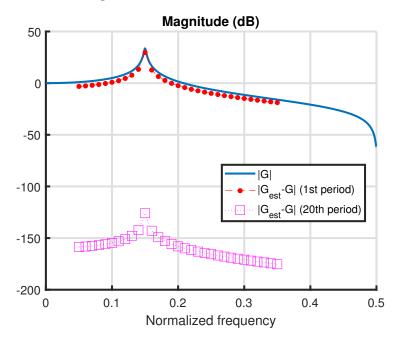


Figure 2.9 – Magnitude FRF of the second-order system and the estimation errors of the FRF at the excited frequencies using (2.3) for the data from the first period and the 20th period.

2.10 Frequency domain methods

Till now we have been treating CT and DT systems as separate cases. However, FD methods can be generalized to both CT and DT systems. There is no reason why a FD method should only work for one and not the other. To generalize these methods it is useful to define Ω to encompass both CT and DT models. Ω can either be s or z^{-1} .

$$\Omega = \begin{cases} s & \text{if working in continuous-time (CT)} \\ z^{-1} & \text{if working in discrete-time (DT)} \end{cases}$$

Thus, denoting an LTI system as $G(\Omega)$ is not specific to CT or DT systems. When the FRF is evaluated in the DFT frequencies $\omega_k = 2\pi k f_s/N$ it, is denoted Ω_k .

$$\Omega_k = \begin{cases} j\omega_k = j2\pi k f_s/N & \text{if working in continuous-time (CT)} \\ e^{-j\omega_k T_s} = e^{-j2\pi k/N} & \text{if working in discrete-time (DT)} \end{cases}$$

2.11 White noise

White noise is a random signal with a flat power spectrum. A zero-mean Gaussian sequence is white noise.

$$v(n) \sim \mathcal{N}(0, \sigma^2)$$

Let's assume that a measurement is perturbed by this white noise.

$$x(n) = x_0(n) + v(n)$$

Due to the linearity of the DFT one obtains

$$X(k) = X_0(k) + V(k)$$

Thus it is interesting to understand the properties of the DFT of this white noise sequence.

$$V(k) = \frac{1}{N} \sum_{n=0}^{N-1} v(n) e^{-j2\pi kn/N}$$

It turns out that (see appendix 2.B)

$$\begin{split} \mathbb{E}\{V(k)\} &= 0\\ \mathbb{E}\{|V(k)|^2\} &= \frac{1}{N}\sigma^2\\ \mathbb{E}\{V^2(k)\} &= 0 \text{ if } \mathrm{mod}(2k,N) \neq 0 \end{split}$$

2.12 Periodic signals

What happens to the DFT spectrum when we apply a signal periodically? Let's assume that we have a sequence $\tilde{x}(n)$ of length N with corresponding DFT

$$\tilde{X}(k) = \sum_{n=0}^{N-1} \tilde{x}(n)e^{-j2\pi kn/N}$$

Now, the sequence $\tilde{x}(n)$ will be repeated one more time to obtain a sequence of length 2N.

$$x(n) = \begin{cases} \tilde{x}(n) & \text{if } 0 \le n < N - 1\\ \tilde{x}(n - N) & \text{if } N \le n < 2N - 1 \end{cases}$$

The DFT of x(n) is then given by

$$\begin{split} X(k) &= \frac{1}{2N} \sum_{n=0}^{2N-1} \tilde{x}(n) e^{-j2\pi k n/(2N)} = \frac{1}{2N} \sum_{n=0}^{N-1} \tilde{x}(n) e^{-j2\pi k n/(2N)} + \frac{1}{2N} \sum_{n=N}^{2N-1} \tilde{x}(n-N) e^{-j2\pi k n/(2N)} \\ &= \frac{1}{2N} \sum_{n=0}^{N-1} \tilde{x}(n) e^{-j2\pi k n/(2N)} + \frac{1}{2N} \sum_{n=0}^{N-1} \tilde{x}(n) e^{-j2\pi k n/(2N)} e^{-j2\pi k n/(2N)} e^{-j2\pi k n/(2N)} \\ &= \frac{1}{N} \sum_{n=0}^{N-1} \tilde{x}(n) e^{-j2\pi k n/(2N)} \frac{1 + e^{-j2\pi k/2}}{2} \end{split}$$

Evaluating this in 2k and 2k + 1 gives

$$X(2k) = \tilde{X}(k)$$
$$X(2k+1) = 0$$

Thus, the even DFT lines will contain the information of X(k) and the uneven DFT lines will be zero.

This result can be generalized to a signal that is repeated for P times.

$$x(n) = \tilde{x}(\text{mod}(n, P)) \Rightarrow X(kP + r) = \begin{cases} \tilde{X}(k) & \text{if } r = 0\\ 0 & \text{if } r = 1, \dots, P - 1 \end{cases}$$

Example A random phase multisine (N = 40) is created where the 3 lowest DFT bins are excited with an RMS value of 1. This signal is repeated P = 3 times. The repeated signal is plotted in the TD and the FD in figure 2.10. It can be seen that in the periodic signal there are P - 1 = 2 non-excited lines in between the excited lines.

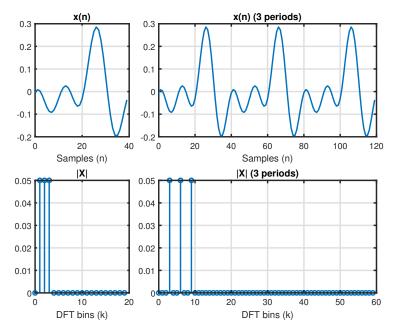


Figure 2.10 – A multisine is repeated 3 times.

2.13 Transient suppression

In section 2.9 it was shown that transients in the output will negatively impact the quality of the FRF estimate. However, there are multiple ways to suppress these transients. The most obvious way is to wait for the system to enter steady state and to only start measuring once the transients have faded away. This is exactly what was done in the example of section 2.9 (see figure 2.9). There are also ways to suppress the transient without throwing away the measurements that are not in steady state. 3 of them will be discussed:

- Windowing
- Parametric estimation
- The local polynomial method

2.13.1 Windowing

There are many possible windows that one can choose from. The Hann window is a popular one.

$$w(n) = \frac{1}{2} \left[1 - \cos(\frac{2\pi n}{N}) \right], \quad n = 0, \dots, N - 1$$
 (2.8)

The windowed input and output are then given by

$$u_w(n) = w(n)u(n)$$
 and $y_w(n) = w(n)y(n)$

A multiplication in the TD becomes a convolution in the FD.

$$U_W(k) = W(k) * U(k)$$
 and $Y_W(k) = W(k) * Y(k)$

with $W(k) = \frac{1}{2}\delta(k) - \frac{1}{4}\delta(k+1) - \frac{1}{4}\delta(k-1)$. The windowed estimate of the FRF is then given by

$$G_W(\Omega_k) = \frac{Y_W(k)}{U_W(k)} \tag{2.9}$$

Let's now see why windowing will reduce the effect of the transient.

$$Y_W(k) = W(k) * (G(\Omega_k)U(k) + T(k))$$

For the first term the following approximation can be made:

$$W(k) * [G(\Omega_k)U(k)] \approx G(\omega_k)[W(k) * U(k)] = G(\omega_k)U_W(k)$$

Hereby it is assumed that $G(\Omega_{k-1}) \approx G(\Omega_k) \approx G(\Omega_{k+1})$. In other words, $G(\Omega)$ is flat around Ω_k . (2.9) then becomes

$$G_W(\Omega_k) \approx G(\Omega_k) + \frac{T_W(k)}{U_W(k)}$$

The trick now is to use the property that T(k) is a smooth function of the frequency. Let's assume that T(k) can be captured locally by a second order polynomial.

$$T(k) = ak^2 + bk + c$$

Windowing T(k) gives

$$T_W(k) \propto 2T(k) - T(k+1) - T(k-1) = -2a$$

This is a bit like taking the second order derivative, which suppresses the transient.

Example The DT system (2.4) is taken again. The same input as in figure 2.2 is used. The estimation error of the FRF with and without windowing is shown in figure 2.11. Windowing gives better results, except near the resonance frequency. This is because the approximation that $G(\Omega)$ is flat does not hold in this region.

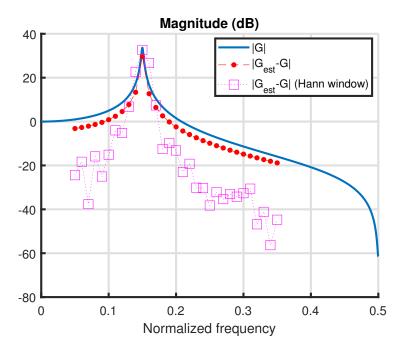


Figure 2.11 – Comparison of FRF estimation error without and with Hann windowing.

2.13.2 Parametric estimation

Till now, only nonparametric estimates of the FRF were discussed. Unlike line-fitting, where one is interested in the slope and offset of the line, the nonparametric estimate does no such thing. The result of the nonparametric estimate is the FRF estimate at every excited frequency. However, in a sense there are still parameters: these are the FRF estimates at every frequency bin $k \in K_{exc}$.

The TF of the system is however still given by a rational function. The coefficients of this rational function can be estimated. The transient term is also a rational function and the parameters describing it can also be included in the estimation.

$$Y(k) = G(\Omega_k)U(k) + T(k)$$

Using $G(\Omega_k) = \frac{B(\Omega_k)}{A(\Omega_k)}$ and $T(k) = \frac{I(\Omega_k)}{A(\Omega_k)}$ the expression above can be rewritten as

$$A(\Omega_k)Y(k) - B(\Omega_k)U(k) - I(\Omega_k) = 0$$
(2.10)

 $A(\Omega_k)$ is a polynomial of order n_a and $B(\Omega_k)$ is a polynomial of order n_b . (2.18) in appendix 2.A is an explicit formula for $I(\Omega_k)$. It is easy to see that $I(\Omega_k)$ is a polynomial of order n_I with

$$n_I = \max\left(n_a, n_b\right) - 1$$

For CT systems, there is also an alias error on top of the transient term. These errors can also be captured well by a polynomial. Grouping the alias error together with the $I(\Omega_k)$ results in [4, Section 6.3.2.3]

$$n_I > \max(n_a, n_b)$$

After a bit of calculations, (2.10) can be turned into

$$\begin{bmatrix} \vdots & & \vdots & & \vdots & & \vdots & & \vdots \\ Y(k) & \dots & Y(k)\Omega_k^{n_a} & U(k) & \dots & U(k)\Omega_k^{n_b} & 1 & \dots & \Omega_k^{n_I} \\ \vdots & & \vdots & & \vdots & & \vdots & & \vdots \end{bmatrix} \begin{bmatrix} a_0 \\ \vdots \\ a_{n_a} \\ -b_0 \\ \vdots \\ -b_{n_b} \\ -i_0 \\ \vdots \\ -i_{n_I} \end{bmatrix} = 0$$

This can be solved by calculating the right null-space of the first matrix. The coefficients i_p will capture the transient. Note that the null space is empty if the measurements are noisy. We won't go into these details in this work, so interested reader is referred to [5].

Example Again, the DT system (2.4) is used with the same input shown in figure 2.2. The parameters are determined as described above. The resulting FRF estimation error is plotted in figure 2.12. Note that as a parametric representation is obtained, the FRF can be calculated at all frequencies. The error is around -300dB, which is MATLAB's precision. This means that the error is negligible and that the transient has been fully suppressed.

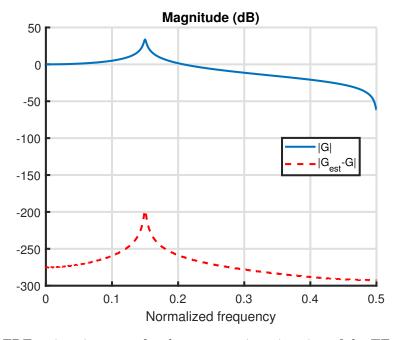


Figure 2.12 – FRF estimation error for the parametric estimation of the TF with transient terms.

2.14 Local polynomial method

The idea of the local polynomial method (LPM) is quite simple. There are multiple variants of the LPM, but here we will focus on the robust LPM for periodic excitations. More information about all the variants of the LPM can be found in [4, Chapter 7].

2.14.1 Response of a system excited by a periodic input

For the robust LPM for periodic signals to work, at least P=2 periods must be measured. Let's assume for the sake of simplicity that the input $\tilde{u}(n)$ is noiseless and in steady state. $\tilde{u}(n)$ is applied P consecutive times to the system.

$$u(n) = \tilde{u}(\operatorname{mod}(n, P)), \quad n = 0, 1, \dots, NP - 1$$

The spectrum of u(n) is

$$U(kP+r) = \begin{cases} \tilde{U}(k) & \text{if } r = 0\\ 0 & \text{if } r = 1, \dots, P-1 \end{cases}$$
 (2.11)

It will resemble the spectrum shown in figure 2.10. Assuming that the output is perturbed by additive noise, the output spectrum is given by

$$Y(kP+r) = G(kP+r)U(kP+r) + T(kP+r) + N_y(kP+r)$$

T(kP+r) is the transient term that arises from the fact that the system might not be in steady state. $N_y(kP+r)$ is additive noise that is added to the output. It can either be white noise (see section 2.11) or filtered white noise. Using (2.11), we can be more specific about the output spectrum.

$$Y(kP+r) = \begin{cases} G(kP)U(kP) + T(kP) + N_Y(kP) & \text{if } r = 0\\ T(kP+r) + N_Y(kP+r) & \text{if } r = 1, \dots, P-1 \end{cases}$$

Example The system (2.4) is excited with the signal shown in figure 2.10. The output is perturbed by Gaussian white noise with a standard deviation of 0.005. The input and output spectra are plotted in figure 2.13. The output spectrum at the DFT lines 3, 6 and 9 are dominated by the G(kP)U(kP) term. There is a peak around the 18-th DFT line that corresponds to the transient term T. The 18-th DFT line corresponds to the normalized frequency 18/(NP) = 18/120 = 0.15, which corresponds to the normalized resonance frequency of the system (see figure 2.12 for example). This is to be expected as the transient resembles the shape of the transfer function. Finally, after the 30-th DFT bin, the noise terms N_Y dominate the output.

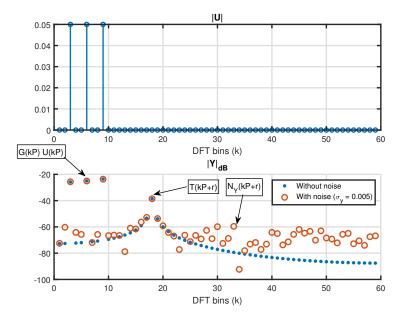


Figure 2.13 – Input and output of the DT system (2.4)

2.14.2 Algorithm

Now we can finally discuss the algorithm of the robust LPM. The idea is to estimate the contribution of the transient term at the excited frequencies using the data at the non-excited frequencies. To this end, we will work in a window around every excited DFT bin. In the case of the previous example, the excited bins are k = 1, 2 and 3. Two parameters must be chosen by the user. The first one is the window size 2n. We must use n unexcited bins before and after kP.

Example k = 3, P = 3, n = 4. We are working in a window around Y(kP) = Y(9). We must use 4 unexcited bins before and after 9.

$$Y(9+r_i)$$
 with $r_i = -5, -4, -2, -1, 1, 2, 4, 5$

For every one of the unexcited lines the following holds

$$Y(kP + r_i) = T(kP + r_i) + N_Y(kP + r_i)$$

The transient term can be modelled as a polynomial of r. This is because the transient is a smooth function of the frequency. The order of the polynomial R is the second parameter that the user can choose.

$$T(kP+r) \approx T(kP) + \sum_{s=1}^{R} t_s(k)r^s$$

We can then find the polynomial of best fit through the points $Y(kP + r_i)$

Example continued Putting all the $Y(9+r_i)$ into a vector and choosing R=2 gives

$$\begin{bmatrix} Y(4) \\ Y(5) \\ Y(7) \\ Y(8) \\ Y(10) \\ Y(11) \\ Y(13) \\ Y(14) \end{bmatrix} = \begin{bmatrix} 1 & (-5) & (-5)^2 \\ 1 & (-4) & (-4)^2 \\ 1 & (-2) & (-2)^2 \\ 1 & (-1) & (-1)^2 \\ 1 & 1 & 1^2 \\ 1 & 2 & 2^2 \\ 1 & 4 & 4^2 \\ 1 & 5 & 5^2 \end{bmatrix} \begin{bmatrix} T(9) \\ t_1(3) \\ t_2(3) \end{bmatrix} + \begin{bmatrix} N_Y(4) \\ N_Y(5) \\ N_Y(7) \\ N_Y(8) \\ N_Y(10) \\ N_Y(11) \\ N_Y(13) \\ N_Y(14) \end{bmatrix} \longrightarrow Y_n = K_n\Theta + V_n$$

The least squares solution is given by¹

$$\hat{\Theta} = (K_n^H K_n)^{-1} K_n^H Y_n$$

This results in an estimation of the transient term at the excited DFT bin $\hat{T}(9)$. The other terms $t_1(3)$ and $t_2(3)$ are not important.

Finally, the estimated transient term can be removed from the output spectrum at the excited DFT line.

$$\hat{Y}(kP) = Y(kP) - \hat{T}(kP)$$

Thus, the transient has been suppressed. The FRF can then be calculated simply as

$$\hat{G}(\Omega_k) = \frac{\hat{Y}(kP)}{U(kP)}$$

The entire procedure outlined here must be repeated for all $k \in K_{exc}$.

¹Note that calculating the solution like this results in an ill-conditioned problem. The \ operator in MATLAB solves this problem using QR-factorisation, which is better conditioned. $(\theta = K_n \setminus Y_n)$

2.14.3 Variance estimate

Having a variance estimate of the output spectrum is useful for providing uncertainty bounds. It can also be used as a nonparametric weighting in parametric identification of the system transfer function.

The residual is defined as the difference between the measured output spectrum and the predicted output spectrum.

$$\hat{V}_n = Y_n - K_n \hat{\Theta}$$

Assuming that the variance of the noise is white (flat) in the window 2n, it can be used to estimate the variance at every $k \in K_{exc}$.

$$\hat{\sigma}_{\hat{Y}}^2(kP) = \frac{1}{q^{\text{noise}}} V_n^H V_n \,, \quad \text{ with } q^{\text{noise}} = 2n - (R+1) \text{ (degrees of freedom)}$$

A proof of this is given in appendix 2.C and is based on [4, Appendix 7.B]. The reason why we must divide by the degrees of freedom q^{noise} and not by 2n is because R+1 parameters are estimated. This is analogous to the reason why the unbiased sample variance is calculated by dividing by the number of observations - 1 when the population mean is also estimated.

As $\hat{G}(\Omega_k) = \hat{Y}(kP)/U(kP)$, the variance of the FRF can be calculated as

$$\hat{\sigma}_{\hat{G}}^{2}(\Omega_{k}) = \frac{\hat{\sigma}_{Y}^{2}(kP)}{|U(kP)|^{2}} \tag{2.12}$$

2.14.4 Example

The robust LPM is applied to measurements from the system (2.4). This time N = 16384 and P = 2. The first F = 5000 frequencies are excited with a random phase multisine with an RMS value 1. White Gaussian noise is added to the output with $\sigma_y = 0.2$. This corresponds to a signal-to-noise ratio (SNR) of 29.5dB. This number is calculated as follows.

$$\text{SNR}_{\text{dB}} = 10 \log_{10} \left(\frac{\frac{1}{NP} \sum_{t=0}^{NP-1} y_0(t)^2}{\sigma_y^2} \right) = 29.5 \text{dB}$$

The numerator is the mean of the squares of the noiseless output y_0 and the denominator is the power of the noise. The SNR is quite high as a consequence of the resonance that is present in the transfer function. 100 noise realizations are simulated while keeping the same random phase multisine as an input and the root-mean square (RMS) error is calculated to get an idea of the effectiveness of the estimator.

$$RMS[|\hat{G} - G|](\Omega_k) = \sqrt{\frac{1}{100} \sum_{i=1}^{100} |\hat{G}^{(i)}(\Omega_k) - G(\Omega_k)|^2}$$

with $\hat{G}^{(i)}(\Omega_k)$ being the nonparametric estimate of $G(\Omega_k)$ for the *i*-th noise realization. The parameters used for the LPM are R=2, n=6 which results in $q^{\text{noise}}=9$ degrees of freedom. To make the results more presentable, the data points are taken together in windows of size 50 and are averaged. The results are shown in figure 2.14. Not taking the transient into account is significantly worse around the resonance frequency of the system. The transient resembles the FRF of the system, which is why the error is most pronounced around the resonance frequency. However, far away from the resonance frequency, not taking the transient into account is approximately 1 dB better than using the Robust LPM. This is because the transient term is below the random noise contribution. LPM uses a noisy estimate of the transient and this estimate is subtracted from the output spectrum, leading to an increased variance. Finally, the robust LPM seems to be slightly better than windowing in this simulation.

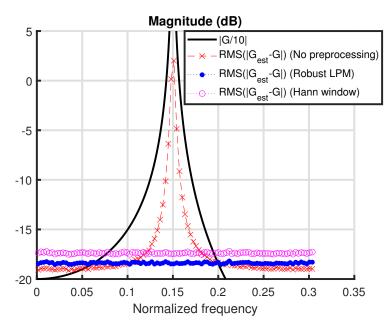


Figure 2.14 – Comparison of the RMS error for the nonparametric estimate acquired without preprocessing, with windowing and with the robust LPM. R = 2, n = 6, $q^{\text{noise}} = 9$.

2.14.5 Choice of the order and degrees of freedom

Two parameters can be chosen by the user when performing the robust LPM analysis: the order of the polynomial approximation and the degrees of freedom used to estimate the noise variance.

Order A good way to choose the order is to start at R = 2 and increment it in steps of two until the estimate of the variance $\hat{\sigma}_{\hat{C}}^2$ stops decreasing.

Degrees of freedom Increasing the degrees of freedom q^{noise} will increase the window size 2n. This gives a better estimation of the variance. There is a trade-off however: it was assumed that the noise variance is white (flat) in the window 2n. Thus, making the window size too big will result in a loss of frequency resolution. Additionally, making the window size 2n bigger means that the transient is approximated by a polynomial over a larger window. At that point it might be necessary to increase the order R of the polynomial.

2.15 Generalization

Many things were simplified until now. A few of the assumptions that were made are:

- The input is noiseless.
- White Gaussian noise noise was simulated in the example, but what if the noise is filtered white noise?
- There is no feedback from output to input.
- The input is a multisine. What about random excitations?

Each of these points will be discussed briefly.

2.15.1 Noisy input

Of course, the measurement of the input can be noisy. Thankfully, the robust LPM is also able to estimate the input noise variance for periodic excitations. Given that the input is not known perfectly, the estimate of the FRF variance (2.12) is not correct any more. In general, the variance of an FRF estimate $\hat{G}(\Omega_k)$ can be approximated by

$$\hat{\sigma}_{\hat{G}}^2(k) = |\hat{G}(\Omega_k)|^2 \left(\frac{\hat{\sigma}_Y^2(k)}{|\hat{Y}(k)|} + \frac{\hat{\sigma}_U^2(k)}{|\hat{U}(k)|} - 2\operatorname{Re}\left(\frac{\hat{\sigma}_{YU}^2(k)}{Y(k)\overline{U(k)}} \right) \right)$$

The equation above is only applicable when the excitation is periodic and when the FRF is estimated by dividing the output spectrum Y(k) by the input spectrum U(k). It is also possible to estimate the variance of the noise for arbitrary excitations, but different formulas must be used. [4, Section 2.6]

2.15.2 Filtered white noise

For simplicity sake, filtered white noise at the sampling instants is defined as

$$v(n) = S(z^{-1})e(n)$$
, with $e(n) \sim \mathcal{N}(0, \sigma^2)$

The power spectrum of this noise is not flat. This also means that the noise samples can be correlated over time.

$$\exists n, m \text{ with } n \neq m \text{ such that } \mathbb{E}\{v(n)v(m)\} \neq 0$$

An important consequence of filtered white noise is that there will also be noise transients in the measurements.

$$V(k) = S(\Omega_k)E(k) + T_S(\Omega_k)$$

As the input of the noise filter is random, the noise transients will never fade away. This means that when one waits long enough for the system $G(\Omega)$ to enter steady state, there will still be noise transients in the measurements. This is where the LPM can also be useful.

2.15.3 Feedback

Consider the measurement set-up shown in figure 2.15. An LTI system $G(z^{-1})$ is in negative feedback and the output is perturbed by process noise.

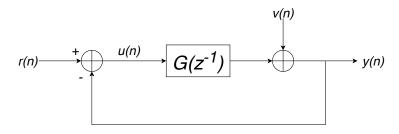


Figure 2.15 – LTI system in negative feedback with additive process noise.

The Z-transform of the output and input are given by

$$Y(z) = \frac{G}{1+G}R(z) + \frac{1}{1+G}V(z)$$
$$U(z) = \frac{1}{1+G}R(z) - \frac{1}{1+G}V(z)$$

Noise that affects the output also affects the input due to the feedback. Two possibilities will be considered: r(n) is a multisine and r(n) is a random signal. Assuming that the process noise v(n) is a Gaussian white noise sequence with variance σ_v^2 we get

$$\mathbb{E}\{V(k)\}=0$$
 and $\mathbb{E}\{|V(k)|^2\}=\sigma_v^2(k)/N$

r(n) is a multisine In this case we can apply multiple periods P to the system. It is assumed that enough time has passed for the system transient to fade away and the noise transients will be neglected.

$$Y^{(p)}(k) = \frac{G(\Omega_k)}{1 + G(\Omega_k)} R^{(p)}(k) + \frac{1}{1 + G(\Omega_k)} V^{(p)}(k)$$
$$U^{(p)}(k) = \frac{1}{1 + G(\Omega_k)} R^{(p)}(k) - \frac{1}{1 + G(\Omega_k)} V^{(p)}(k)$$

The superscript ^(p) denotes the period of the measurement. The nonparametric FRF can then be estimated with

$$\hat{G}(\Omega_k) = \frac{\frac{1}{P} \sum_{p=0}^{P} Y^{(p)}(k)}{\frac{1}{P} \sum_{p=0}^{P} U^{(p)}(k)}$$
(2.13)

Taking the limit for $P \to \infty$ will allow us to establish whether this estimator is consistent.

$$\lim_{P \to \infty} \hat{G}(\Omega_k) = \frac{\mathbb{E}\{Y^{(p)}(k)\}}{\mathbb{E}\{U^{(p)}(k)\}} = G(\Omega_k)$$

The first equality applies the law of large numbers for independent experiments. And so this estimator is indeed consistent.

r(n) is an arbitrary signal For arbitrary signals it is not advised to use the estimator (2.13) because a division by a small number is possible, resulting in an estimator that won't converge. Another thing that must be considered when working with arbitrary signals is that arbitrary signals are aperiodic. If we want to get a better estimate of the FRF we will need to cut the measurement into M pieces of N samples. If the total number of samples NM is constant, this will results in a trade-off between the frequency resolution f_s/N and noise suppression. An estimator that is used for arbitrary signals is

$$\hat{G}(\Omega_k) = \frac{\frac{1}{M} \sum_{m=0}^{M} Y^{(m)}(k) \overline{U^{(m)}(k)}}{\frac{1}{M} \sum_{m=0}^{M} U^{(m)}(k) \overline{U^{(m)}(k)}}$$
(2.14)

Here there is no danger for the denominator to become small. This estimator is consistent when only the output is perturbed by noise. However, in our case the input is also perturbed. Taking the limit as $M \to \infty$ gives

$$\lim_{M \to \infty} \hat{G}(\Omega_k) = \frac{\mathbb{E}\{Y^{(m)}(k)\overline{U^{(m)}(k)}\}}{\mathbb{E}\{U^{(m)}(k)\overline{U^{(m)}(k)}\}} = \frac{S_{YU}(k)}{S_{UU}(k)}$$

with S_{YU} being the cross-power spectrum between the output and input and S_{UU} being the auto-power spectrum of the input. Assuming that the reference is zero-mean Gaussian white noise with variance σ_r^2 we get

$$\mathbb{E}\{R(k)\} = 0 \text{ and } \mathbb{E}\{|R(k)|^2\} = \sigma_r^2(k)/N$$

and the fact that R(k) and V(k) are uncorrelated we get the following result

$$\lim_{P \to \infty} \hat{G}(\Omega_k) = \frac{G(\Omega_k) \sigma_r^2(k) - \sigma_v^2(k)}{\sigma_r^2(k) + \sigma_v^2(k)} \neq G(\Omega_k)$$

Thus, the estimator (2.14) is inconsistent.

Indirect method It is possible to get around the problem of the estimator (2.14) by also using the reference signal r(n). Instead of modelling the transfer function from input (u) to output (y), we will model the transfer function from reference (r) to output (y) and from reference (r) to input (u).

$$\hat{G}(\Omega_k) = \frac{\frac{1}{P} \sum_{m=0}^{P} Y^{(m)}(k) \overline{R^{(m)}(k)}}{\frac{1}{P} \sum_{m=0}^{P} U^{(m)}(k) \overline{R^{(m)}(k)}}$$
(2.15)

Doing some calculations brings us to

$$\lim_{M \to \infty} \hat{G}(\Omega_k) = \frac{\mathbb{E}\{Y^{(m)}(k)\overline{R^{(m)}(k)}\}}{\mathbb{E}\{U^{(m)}(k)\overline{R^{(m)}(k)}\}} = \frac{S_{YR}(k)}{S_{UR}(k)} = G(\Omega_k)$$

Thus, by keeping the reference signal and using it, it is possible to get a consistent estimate when using arbitrary excitations.

2.16 Conclusion

Multisines can be used to identify nonparametric estimates of the FRF. The robust LPM allows for the transients to be suppressed. Even if the system is in steady state, the noise transients can still deteriorate the quality of the FRF estimate. This is a reason why the robust LPM can be quite effective.

It is useful to keep the reference signal that is applied to the system. When the system is excited by an arbitrary excitation, this information can be used to get a consistent estimate of the FRF even if both the input and the output are perturbed by noise.

Finally, the authors of [4] offer MATLAB code that can perform the robust LPM for periodic and arbitrary excitations. These MATLAB function can also identify the best linear approximation of a nonlinear system. These functions can also handle MIMO systems.

Appendix

2.A Transient term

Suppose that a DT LTI system is of the following form.

$$a_0y(n) + a_1y(n-1) + a_2y(n-2) = b_0u(n) + b_1u(n-1)$$
(2.16)

The DFT of a sequence is actually just a windowed version of the Z-transform. To see this, the window is defined as

$$w(n) = \begin{cases} 1, & 0 \le n < N \\ 0, & \text{otherwise} \end{cases}$$

The windowed Z-transform of a sequence x(n) is then given by

$$\mathcal{Z}\{w(n)x(n)\} = \sum_{n=-\infty}^{+\infty} w(n)x(n)z^{-n} = \sum_{n=0}^{N-1} x(n)z^{-n}$$

Evaluating the expression above in $z = e^{j2\pi k/N}$ is exactly the DFT of x(n).

$$\mathcal{Z}\{w(n)x(n)\}|_{z=e^{j2\pi k/N}} = \sum_{n=0}^{N-1} x(n)e^{-j2\pi kn/N} = DFT\{x(n)\}$$

Thus, taking the windowed Z-transform of both sides of (2.16) and evaluating it in $z = e^{j2\pi k/N}$ is the same as taking the DFT of both sides.

For simplicity, let's only consider one of the terms.

$$a_p w(n) y(n-p)$$

 a_p is just a constant, so that can also be left out in the analysis.

$$\mathcal{Z}\{w(n)y(n-p)\} = \sum_{n=0}^{N-1} y(n-p)z^{-n}$$

After some manipulations:

$$\mathcal{Z}\{w(n)y(n-p)\} = z^{-p} \sum_{n=0}^{N-1} y(n)z^{-n} + \sum_{n=0}^{p-1} [y(n-p) - y(n-p+N)z^{-N}]z^{-n}$$

Evaluating this expression in $z = e^{j2\pi k/N}$ gives

$$\mathcal{Z}\{w(n)y(n-p)\}|_{z=e^{j2\pi k/N}} = (e^{-j2\pi k/N})^p \sum_{n=0}^{N-1} y(n)e^{-j2\pi kn/N} + \sum_{n=0}^{p-1} [y(n-p) - y(n-p+N)](e^{-j2\pi k/N})^n$$

Note that z^{-N} disappears because $e^{-j2\pi Nn/N}=1$. The first term contains the DFT of y(n). The second term is a polynomial in $e^{-j2\pi k/N}$ that depends on y(n-p)-y(n-p+N). In other words, it depends on the difference between samples from the previous period and samples from the current period.

$$\frac{1}{N} \mathcal{Z}\{w(n)y(n-p)\}|_{z=e^{j2\pi k/N}} = (e^{-j2\pi k/N})^p Y(k) + I_{y,p}(e^{-j2\pi k/N})$$

Applying this to (2.16) gives

$$Y(k)\left(\sum_{p=0}^{2} a_{p} (e^{-j2\pi k/N})^{p}\right) = U(k)\left(\sum_{p=0}^{1} b_{p} (e^{-j2\pi k/N})^{p}\right) + I(e^{-j2\pi k/N})$$
(2.17)

with

$$I(e^{-j2\pi k/N}) = \sum_{p=0}^{1} b_p I_{u,p}(e^{-j2\pi k/N}) - \sum_{p=0}^{2} a_p I_{y,p}(e^{-j2\pi k/N})$$
(2.18)

Notice that

$$G(z^{-1})|_{z=e^{j2\pi k/N}} = G(e^{-j2\pi k/N}) = \frac{B(e^{-j2\pi k/N})}{A(e^{-j2\pi k/N})} = \frac{\sum_{p=0}^{1} b_p (e^{-j2\pi k/N})^p}{\sum_{p=0}^{2} a_p (e^{-j2\pi k/N})^p}$$

Dividing (2.17) by $A(e^{-j2\pi k/N})$ then gives the final form

$$Y(k) = G(e^{-j2\pi k/N})U(k) + T(k)$$
 (2.19)

with

$$T(k) = \frac{I(e^{-j2\pi k/N})}{A(e^{-j2\pi k/N})}$$

2.B DFT of white noise

A white noise sequence v(n) has the following properties

$$\mathbb{E}\{v(n)\} = 0$$

$$\mathbb{E}\{v(n)v(m)\} = \sigma^2\delta(n-m)$$

The DFT of this white noise sequence is

$$V(k) = \frac{1}{N} \sum_{n=0}^{N-1} v(n)e^{-j2\pi kn/N}$$

 $\mathbb{E}\{V(k)\}$ The expected value of V(k) is

$$\mathbb{E}\{V(k)\} = \frac{1}{N} \sum_{n=0}^{N-1} \mathbb{E}\{v(n)\} e^{-j2\pi kn/N} = \frac{1}{N} \sum_{n=0}^{N-1} 0 e^{-j2\pi kn/N} = 0$$

 $\mathbb{E}\{V(k)\overline{V(l)}\}$ The expected value of $V(k)\overline{V(l)}$ is

$$\begin{split} \mathbb{E}\{V(k)\overline{V(l)}\} &= \frac{1}{N^2} \mathbb{E}\{\sum_{n=0}^{N-1} v(n)e^{-j2\pi kn/N} \sum_{m=0}^{N-1} v(m)e^{j2\pi lm/N} \} \\ &= \frac{1}{N^2} \sum_{n=0}^{N-1} \sum_{m=0}^{N-1} \mathbb{E}\{v(n)v(m)\} e^{-j2\pi kn/N} e^{j2\pi lm/N} \\ &= \frac{\sigma^2}{N^2} \sum_{n=0}^{N-1} \sum_{m=0}^{N-1} \delta(n-m)e^{-j2\pi kn/N} e^{j2\pi lm/N} \\ &= \frac{\sigma^2}{N^2} \sum_{n=0}^{N-1} e^{-j2\pi (k-l)n/N} = \begin{cases} \frac{\sigma^2}{N} \text{ if } \operatorname{mod}(k-l,N) = 0 \\ 0 \text{ otherwise} \end{cases} \end{split}$$

When k = l, this result becomes

$$\mathbb{E}\{|V(k)|^2\} = \frac{\sigma^2}{N}$$

 $\mathbb{E}\{V(k)V(l)\}$ The expected value of V(k)V(l) is

$$\begin{split} \mathbb{E}\{V(k)V(l)\} &= \frac{1}{N^2} \mathbb{E}\{\sum_{n=0}^{N-1} v(n) e^{-j2\pi k n/N} \sum_{m=0}^{N-1} v(m) e^{-j2\pi l m/N} \} \\ &= \frac{1}{N^2} \sum_{n=0}^{N-1} \sum_{m=0}^{N-1} \mathbb{E}\{v(n)v(m)\} e^{-j2\pi k n/N} e^{-j2\pi l m/N} \\ &= \frac{\sigma^2}{N^2} \sum_{n=0}^{N-1} e^{-j2\pi (k+l)n/N} = \begin{cases} \frac{\sigma^2}{N} & \text{if } \operatorname{mod}(k+l,N) = 0 \\ 0 & \text{otherwise} \end{cases} \end{split}$$

When k = l and $mod(2k, N) \neq 0$, this result becomes

$$\mathbb{E}\{V^2(k)\} = 0$$

2.C Covariance estimation

The output spectrum in a window of size 2n around an excited frequency line is given by

$$Y_n = K_n \Theta + V_n$$

The least squares solution is given by

$$\hat{\Theta} = (K_n^H K_n)^{-1} K_n^H Y_n$$

A key assumption used to estimate the covariance $C_Y(kP)$ is that V_n is assumed to have a flat power spectrum in the window 2n.

The residual is the difference between the measured spectrum and the predicted spectrum.

$$\hat{V}_n = Y_n - K_n \hat{\Theta} = (I_{2n} - K_n (K_n^H K_n)^{-1} K_n^H) Y_n = P_n Y_n$$

Using the fact that $P_nK_n = K_n - K_n = 0$ we get

$$\hat{V}_n = P_n V_r$$

Next up, we want to to see how $\hat{V}_n^H \hat{V}_n$ relates to $V_n^H V_n$.

$$\hat{V}_n^H \hat{V}_n = V_n^H P_n^H P_n V_n$$

First, it is easy to see that $P_n^H = P_n$. Next, it turns out that P_n is an idempotent matrix.

$$P_n P_n = (I_{2n} - K_n (K_n^H K_n)^{-1} K_n^H) (I_{2n} - K_n (K_n^H K_n)^{-1} K_n^H)$$

= $I_{2n} - 2K_n (K_n^H K_n)^{-1} K_n^H + K_n (K_n^H K_n)^{-1} K_n^H = P_n$

Thus, we get

$$\hat{V}_n^H \hat{V}_n = V_n^H P_n V_n = \operatorname{trace}(V_n^H P_n V_n) = \operatorname{trace}(P_n V_n V_n^H)$$
(2.20)

In this step we used the fact that the trace of a scalar is a scalar and the fact that matrices in a trace can be circularly permuted (trace(ABC) = trace(BCA)). It is assumed that V_n has a flat power spectrum in the window 2n, i.e. $\sigma_V^2(kP + r_i) = \sigma_V^2(kP)$. This means that

$$\mathbb{E}\{V_n V_n^H\} = \sigma_V^2(kP)I_{2n}$$

Plugging this into (2.20) and taking the expected value gives

$$\mathbb{E}\{\hat{V}_n^H \hat{V}_n\} = \sigma_V^2(kP) \operatorname{trace}(P_n)$$
(2.21)

 P_n is an idempotent matrix, meaning that its eigenvalues can only be 0 or 1. Additionally, because P_n and K_n are each other's orthogonal complement, the rank of P_n is related to the rank of K_n .

$$K_n P_n = 0 \Rightarrow \operatorname{rank}(P_n) = 2n - \operatorname{rank}(K_n)$$

If K_n is full column rank and if K_n has more rows than columns, then the rank of K_n is equal to the number of columns in K_n .

$$rank(K_n) = R + 1$$

The trace of a matrix is the sum of the eigenvalues of a matrix. The sum of the eigenvalues of P_n is exactly equal to the rank of P_n , because the rank of P_n is equal to the number of nonzero eigenvalues and because the eigenvalues can only be 0 or 1.

$$\operatorname{trace}(P_n) = \operatorname{rank}(P_n) = 2n - (R+1) = q^{\text{noise}}$$

Finally, this explains why $\hat{V}_n^H \hat{V}_n$ must be divided by q^{noise} to get an unbiased estimate of the covariance.

$$\sigma_Y^2(kP) = \frac{\mathbb{E}\{\hat{V}_n^H \hat{V}_n\}}{q^{\text{noise}}}$$

Chapter 3

Model reference control

3.1 Introduction

The first part of this chapter is a summary of the work presented in [1]. From section 3.6 and on, improvements are made to the existing methods.

3.2 Problem statement

The goal of model reference control is to design introller for a single-input single-output system $G(\Omega)$. Traditionally, the first step to finding a controller is to estimate a parametric representation of $G(\Omega)$. In this controller is the modelling step will be skipped. We will go directly from input-output data to a controller. assumed that input-output measurements (u(n)) and u(n) respectively) of the system in open loop are available to the user. It is also assumed that u(n) is stable and minimum-phase. It is also possible to extend this theory to unstable nonminimum-phase systems. This is done in appendix 3.C.

The system is controlled by an unknown controller $K(\Omega, \rho)$ in closed loop (CL). This is shown graphically in figure 3.1.

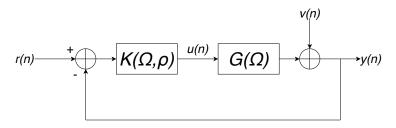


Figure 3.1 – Closed loop system.

The transfer function from the reference r to the output y is given by

$$CL(\Omega) = \frac{K(\Omega, \rho)G(\Omega)}{1 + K(\Omega, \rho)G(\Omega)}$$

 $\rho = \begin{bmatrix} \rho_1 & \dots & \rho_{n_\rho} \end{bmatrix}^T$ is a vector containing the controller parameters that should be optimized. In this work $K(\Omega, \rho)$ is linear in the parameters.

$$K(\Omega, \rho) = \beta(\Omega)\rho \tag{3.1}$$

with $\beta(\Omega)$ being a row vector with n_{ρ} elements. The idea of model reference control, is to get the closed loop system "as close" as possible to a user-defined reference system $M(\Omega)$.

$$M(\Omega) \approx \frac{K(\Omega, \rho)G(\Omega)}{1 + K(\Omega, \rho)G(\Omega)}$$

 $M(\Omega)$ needs to be chosen such that it is a stable causal LTI system. This "closeness" criterion can be quantified by using the 2-norm of a transfer function.

$$J_{mr}(\rho) = \left| \left| F(\Omega) \left[M(\Omega) - \frac{K(\Omega, \rho)G(\Omega)}{1 + K(\Omega, \rho)G(\Omega)} \right] \right| \right|_{2}^{2}$$
(3.2)

 $F(\Omega)$ is a user-defined weighing filter that can be chosen to highlight specific frequencies. The 2-norm of a SISO system is defined differently for CT and DT systems. For CT systems it is

$$||H(s)||_2^2 = \frac{1}{2\pi} \int_{-\infty}^{+\infty} |H(j\omega)|^2 d\omega$$

and for DT systems it is

$$||H(z^{-1})||_2^2 = \frac{1}{2\pi} \int_{-\pi}^{\pi} |H(e^{j\omega})|^2 d\omega$$

3.3 Convex cost

A key problem with the use of (3.2) as a cost function is that it is not convex. In order to solve this, we must first define the ideal controller $K^*(\Omega)$ as

$$K^*(\Omega) = \frac{M(\Omega)}{G(\Omega)(1 - M(\Omega))}$$
(3.3)

This definition ensures that the closed loop system is equal to the reference system by construction.

$$\frac{K^*(\Omega)G(\Omega)}{1+K^*(\Omega)G(\Omega)}=M(\Omega)$$

Note that it is possible that K^* is not realizable i.e.

$$\not\equiv \rho$$
 such that $K(\Omega, \rho) = K^*(\Omega)$

For example, if $K^*(\Omega)$ is a polynomial of Ω of degree 4, then $K(\Omega, \rho) = \rho_0 + \rho_1 \Omega + \rho_2 \Omega^2$ will not be able to realize $K^*(\Omega)$ perfectly.

Next, both terms in (3.2) can be put on the same denominator.

$$M(\Omega) - \frac{K(\Omega, \rho)G(\Omega)}{1 + K(\Omega, \rho)G(\Omega)} = \frac{M(\Omega) - (1 - M(\Omega))K(\Omega, \rho)G(\Omega)}{1 + K(\Omega, \rho)G(\Omega)}$$

The sensitivity function is approximated by the ideal sensitivity function.

$$\frac{1}{1+K(\Omega,\rho)G(\Omega)}\approx\frac{1}{1+K^*(\Omega)G(\Omega)}=\frac{1}{1+\frac{M(\Omega)}{1-M(\Omega)}}=1-M(\Omega)$$

The validity of this approximation should be verified afterwards. The sensitivity function is the transfer function from the disturbance v(n) to the output y(n). It quantifies how sensitive the output is to disturbances.

This approximation leads to the definition of the convex cost function.

$$J(\rho) = \left| \left| F(\Omega)(1 - M(\Omega)) \left[M(\Omega) - (1 - M(\Omega)) K(\Omega, \rho) G(\Omega) \right] \right| \right|_{2}^{2}$$
 (3.4)

Of course, not all forms of $K(\Omega, \rho)$ will make this cost function convex. However, it is convex when $K(\Omega, \rho)$ is linear in the parameters (3.1). Note that the cost is minimized for the ideal controller if the ideal controller is realizable.

$$K(\rho^*, \Omega) = K^*(\Omega) \Longrightarrow J(\rho^*) = 0$$

3.4 Other cost functions

We can devise many other cost functions that solve this problem. An example is the following.

$$J_K(\rho) = ||K^*(\Omega) - K(\Omega, \rho)||_2^2$$
(3.5)

This cost function minimizes the square of the difference between the actual controller $K(\Omega, \rho)$ and the ideal controller $K^*(\Omega)$. If $K(\Omega, \rho)$ is linear in the parameters, then this is just a simple linear least squares regression. However, this is not the same as optimizing ρ to the cost function (3.4) and will result in a different outcome. In fact, it can be shown that (3.4) is a weighted linear least squares regression version of (3.5). By using (3.3) we get

$$M(\Omega) = G(\Omega)(1 - M(\Omega))K^*(\Omega)$$

Using this expression in (3.4) results in

$$J(\rho) = \left| \left| F(\Omega)(1 - M(\Omega)) \left[G(\Omega)(1 - M(\Omega))K^*(\Omega) - (1 - M(\Omega))K(\Omega, \rho)G(\Omega) \right] \right| \right|_2^2$$
$$= \left| \left| F(\Omega)(1 - M(\Omega))^2 G(\Omega) \left[K^*(\Omega) - K(\Omega, \rho) \right] \right| \right|_2^2$$

Note that $J(\rho)$ depends on $G(\Omega)$ both directly and indirectly as K^* also depends on $G(\Omega)$. No expression for $G(\Omega)$ is known in the data-driven approach and it will therefore have to be estimated. This shows that the choice of the cost function is an important part of thee optimization and must be done with care.

3.5 Correlation-based approach

As $G(\Omega)$ is not known in (3.4), the authors of [1] propose the correlation-based approach. They give algorithms for periodic and arbitrary excitations. Here we will focus on their equations concerning periodic excitations. We will also restrict this section to DT systems $(\Omega = z^{-1})$, as they have done in their paper.

One thing to note, is that all the equations in [1] are written in the TD. These will be translated to the FD in section 3.6. The reader is recommended to not spend too much time trying to understand the definitions in this section as they will make more sense in the next section.

Model First it is assumed that the output of the DT system is perturbed by some noise v(n).

$$y(n) = G(q^{-1})u(n) + v(n)$$
(3.6)

with q^{-1} being the backshift operator: $q^{-1}x(n)=x(n-1)$. N denotes the period of the input.

$$u(n) = u(n+N) \tag{3.7}$$

The additive noise is modelled as DT filtered white noise.

$$v(n) = S_v(q^{-1})e_v(n)$$
, with $\mathbb{E}\{e_v(n)\} = 0$ and $\mathbb{E}\{e_v(n)^2\} = \sigma^2$

Error signal Then a new quantity $\epsilon(n, \rho)$ is defined [1, eq. (15)].

$$\epsilon(n,\rho) = M(q^{-1})u(n) - K(q^{-1},\rho)(1 - M(q^{-1}))y(n)$$
 (computational form)
= $[M - K(\rho)(1 - M)G]u(n) - K(\rho)(1 - M)v(n)$ (analytic form)

In the second equality operators q^{-1} are left out for clarity. The quantitational form is used in practice—as the user has access to measurements u(n) and v(n). The analytic form is used to develop the theory further.

Notice that when $K(\rho) = K^*$, the first term in the analytic form is zero.

$$\epsilon^*(n) = -K^*(1 - M)v(n)$$

Thus $\epsilon(n,\rho)$ is uncorrelated with u(n) when $\rho = \rho^*$. The main idea behind the correlation-based approach is to tune the parameters ρ such that $\epsilon(n,\rho)$ becomes uncorrelated with u(n). However, note that this reasoning only holds if K^* is realizable with the proposed controller scheme $K(\rho)$. If K^* is not realizable, then the first term $[M - K(\rho)(1 - M)G]$ will not be zero when replacing $K(\rho)$ with the optimal controller and $\epsilon(n,\rho)$ will still be correlated with u(n).

Input filtering Additionally, the filter W is defined [1, eq. (41)]

$$W(e^{-j\omega_k}) = \frac{F(e^{-j\omega_k})(1 - M(e^{-j\omega_k}))}{S_{UU}(e^{-j\omega_k})}$$
$$\omega_k = \frac{2\pi k}{N}, \quad k = 0, \dots, N-1$$

 $S_{UU}(e^{-j\omega_k})$ is the DFT of the autocorrelation of u(n) [1, eqs. (38) and (39)].

$$R_{uu}(\tau) = \frac{1}{N} \sum_{n=0}^{N-1} u(n-\tau)u(n)$$
(3.8)

$$S_{UU}(e^{-j\omega_k}) = \sum_{\tau=0}^{N-1} R_{uu}(\tau)e^{-j\tau\omega_k}$$
(3.9)

If the system is in steady state, $u(n-\tau)$ is known for negative time indices by using (3.7). If the experiment starts with zero initial conditions, then $u(n-\tau) = 0$ for $n-\tau < 0$.

The filter W is applied to the input u(n). As is remarked at the end of [1, Sec. 4.4], if a parametric representation of $S_{UU}(q^{-1})$ is known, this filter can be applied in the TD.

$$u_W(n) = W(q^{-1})u(n)$$

Note that this can be problematic if $S_{UU}(q^{-1})$ has zeros that are not on the unit circle, as the filter $W(q^{-1})$ will then be unstable. More information concerning this is given in appendix 3.A.

Correlation criterion Then, the cross-correlation between between $u_W(n)$ and $\epsilon(n,\rho)$ is calculated.

$$R_{u_W\epsilon}(\tau,\rho) = \frac{1}{NP} \sum_{n=0}^{NP-1} u_W(n-\tau)\epsilon(n,\rho)$$
(3.10)

with P being the number of periods measured. Note that as in (3.8), $u_W(n-\tau)$ can be found by using (3.7) if the system is in steady state. If the system starts off in zero initial conditions, $u_W(n-\tau) = 0$ for $n-\tau < 0$.

The correlation criterion $J_{NP,l_1}(\rho)$ can then be defined.

$$J_{NP,l_1}(\rho) = \sum_{\tau = -l_1}^{l_1} R_{u_W \epsilon}^2(\tau, \rho)$$
(3.11)

with $l_1 \leq N/2$ being a parameter that can be chosen by the user. The idea behind this parameter l_1 is discussed in detail in section 3.7. It is then proven in [1, Appendix II] that (3.11) converges to (3.4) for $N, P \to \infty$ with probability 1 under certain conditions.

$$\lim_{N \to \infty} J_{NP,l_1}(\rho) = J(\rho), \text{ w.p. } 1$$

However, for finite data, it is proven that the estimator is biased [1, eq. (37)] The bias is discussed further in section 3.8.

3.6 Translation to the frequency domain

In this section the results from the previous section will be translated to the FD. Here we will show that a nonparametric estimate of the system $G(\Omega)$ is actually hidden in the mathematics. Thus, even though the parametric modelling step is skipped, there is still a nonparametric model.

3.6.1 Disadvantages of TD

There are some disadvantages of working in the TD. Here are some that are relevant to this subject.

- TD filtering can only be done easily if the underlying system is a DT system.
- The filtering of u(n) with $W(q^{-1})$ can only be performed in the TD if a parametric representation of $S_{UU}(q^{-1})$ is known and if $S_{UU}(q^{-1})$ only has zeros on the unit circle. (see appendix 3.A).
- If the system starts with zero initial conditions or in steady state, then the transient response can be taken into account, as knowledge of the input and output signals are known before they are applied. This makes it possible to calculate (3.8) and (3.10) for $n \tau < 0$. However, if the system does not start with zero initial conditions or in steady state, this approach will not work and the transient term will not be suppressed.

These problems can be solved by working in the FD. To address the first point: FD methods can also handle CT systems. For the second point: a convolution in the TD will explode when the system is unstable. However, this is not a problem in the FD as a convolution in the TD becomes a simple multiplication in the FD. Finally, for the third point: the robust LPM (see section 2.14) is able to estimate nonparametrically the frequency response function (FRF) from noisy input-output data, while suppressing the transient term.

3.6.2 Nonparametric estimate

By translating the formulas of section 3.5 to the FD, it will become immediately apparent that a nonparametric estimate of the FRF is being calculated. The first step in translating the problem from the TD to the FD is to use Parseval's theorem. According to Parseval's theorem, the sum of squares in the TD is equivalent to the sum of the norms squared in the FD. The cost function (3.11) represents a sum of squares in the TD. Parseval's theorem will be discussed in greater detail in section 3.7.

Now, instead of calculating $\epsilon(n,\rho)$ in the TD, it can be calculated in the FD. Note that this also makes the computations less intensive, as a convolution in the TD becomes a multiplication in the FD. As the input to the system is assumed to be periodic with period P, the frequencies Ω_k will correspond to the DFT bins kP.

$$E(kP, \rho) = M(\Omega_k)U(kP) - K(\Omega_k, \rho)(1 - M(\Omega_k))Y(kP)$$

Applying the filter W to u(n) can also be done in the FD.

$$U_W(kP) = W(\Omega_k)U(kP)$$

with

$$W(\Omega_k) = \frac{F(\Omega_k)(1 - M(\Omega_k))}{S_{UU}(e^{-j\omega_k})}$$

The denominator $S_{UU}(e^{-j\omega_k})$ is proportional to $U(kP)\overline{U(kP)}$.

$$S_{UU}(e^{-j\omega_k}) = NU(kP)\overline{U(kP)}$$
(3.12)

This equivalence is proven in appendix 3.B. Then, the cross-power spectrum between $\epsilon(n, \rho)$ and $u_W(n)$ is also calculated in the FD by doing a simple multiplication.

$$S_{U_WE}(\Omega_k, \rho) = NU_W(kP)\overline{E(kP, \rho)}$$
(3.13)

Expanding (3.13) makes the link to the cost function (3.4) immediately apparent.

$$S_{U_WE}(\Omega_k, \rho) = U(kP) \frac{F(1-M)}{U(kP)\overline{U(kP)}} \overline{(MU(kP) - K(\rho)(1-M)Y(kP))}$$
(3.14)

$$= F(1-M)\overline{\left(M - K(\rho)(1-M)\hat{G}(\Omega_k)\right)}$$
(3.15)

with

$$\hat{G}(\Omega_k) = \frac{Y(kP)\overline{U(kP)}}{U(kP)\overline{U(kP)}} = \frac{Y(kP)}{U(kP)}$$
(3.16)

The index Ω_k was left out in F, M and K for clarity. $\hat{G}(\Omega_k)$ is a nonparametric estimate of the system. If \hat{G} is replaced by the actual system G, then $S_{UwE}(\Omega_k, \rho)$ is exactly the quantity being integrated over in (3.4).

Thus, a nonparametric estimate of the system \hat{G} can be found, followed by calculating the cost function in the FD.

$$J_N(\rho) = \frac{1}{|\mathcal{K}_{\text{exc}}|} \sum_{k \in \mathcal{K}_{\text{exc}}} \left| H(\Omega_k, \rho) \right|^2$$
(3.17)

with $K_{\rm exc}$ being the set of excited DFT bins, $|K_{\rm exc}|$ being the cardinality of this set and

$$H(\Omega_k, \rho) = F(\Omega_k)(1 - M(\Omega_k)) \left[M(\Omega_k) - (1 - M(\Omega_k)) K(\Omega_k, \rho) \hat{G}(\Omega_k) \right]$$
(3.18)

Proceeding in this way, we can be much more flexible with the manner in which we estimate the FRF nonparametrically. Let's take a closer look at the nonparametric estimator that is hidden inside the formulas of the TD.

$$\hat{G}(\Omega_k) = \frac{Y(kP)}{U(kP)}$$

Now, let's separate y(n) into its P periods $y^{(p)}(n)$.

$$y^{(p)}(n) = y(n+pP), \quad p = 0, \dots, P-1$$

We can also define the DFT for each of the periods.

$$Y^{(p)}(k) = \sum_{n=0}^{N-1} y^{(p)}(n)e^{-j2\pi kn/N}, \quad p = 0, \dots, P-1$$

With these definitions we can get a better understanding of what Y(kP) represents.

$$\begin{split} Y(kP) &= \sum_{n=0}^{NP-1} y(n) e^{-j2\pi(kP)n/(NP)} \\ &= \sum_{p=0}^{P-1} \sum_{n=0}^{N-1} y(n+pP) e^{-j2\pi kn/N} = \sum_{p=0}^{P-1} Y^{(p)}(k) \end{split}$$

Thus, the nonparametric estimate becomes

$$\hat{G}(\Omega_k) = \frac{\frac{1}{P} \sum_{p=0}^{P-1} Y^{(p)}(k)}{\frac{1}{P} \sum_{p=0}^{P-1} U^{(p)}(k)}$$
(3.19)

This is a consistent estimator for the model (3.6). Indeed, if we translate (3.6) to the FD we obtain

$$Y^{(p)}(k) = G(\Omega_k)U^{(p)}(k) + V^{(p)}(k), \quad p = 0, \dots, P-1$$

Given that the input is periodic $U^{(p)}(k) = U_0(k)$, we get

$$\hat{G}(\Omega_k) = \frac{\frac{1}{P} \sum_{p=0}^{P-1} \left[G(\Omega_k) U_0(k) + V^{(p)}(k) \right]}{U_0(k)} = G(\Omega_k) + \frac{1}{U_0(k)} \frac{1}{P} \sum_{p=0}^{P-1} V^{(p)}(k)$$

The output noise is assumed to be Gaussian white noise, which means that $\mathbb{E}\{V(k)\}=0$. Taking the limit for $P\to\infty$ gives

$$\lim_{P\to\infty} \hat{G}(\Omega_k) = G(\Omega_k) + \frac{1}{U_0(k)} \mathbb{E}\{V(k)\} = G(\Omega_k) \,, \text{ w.p. 1}$$

The law of large numbers for independent experiments is used in the first equation. Following the same lines, it can easily be verified that the estimator is also consistent if the measurement of the input is noisy.

$$U^{(p)} = U_0(k) + N_U^{(p)}(k)$$

3.7 Parseval's theorem

Parseval's theorem is very simple: the energy of the signal in the TD is the same as the energy of the signal in the FD.

$$\sum_{n=0}^{N-1} x(n)^2 = N \sum_{k=0}^{N-1} |X(k)|^2$$
(3.20)

with $X(k) = DFT\{x(n)\}$. The factor N is specific to the definition of the DFT that is used in this work and will be different if another definition is used.

The cost function in the FD (3.17) is also the sum of squares.

$$J_N(\rho) = \frac{1}{N} \sum_{k=0}^{N-1} |H(\Omega_k, \rho)|^2$$

According to Parseval's theorem, the above sum is equivalent to a sum of squares in the TD.

$$J_N(\rho) = \frac{1}{N^2} \sum_{n=0}^{N-1} h(n, \rho)^2$$
 (3.21)

with $h(n,\rho)$ being the IDFT of $H(\Omega_k,\rho)$. If $N\to\infty$, $h(n,\rho)$ is the impulse response of the system $H(\Omega,\rho)$. The trick that the authors use in [1] is the fact that the impulse response of a stable system will fade away after some time.

$$\exists l_1 \text{ such that } h(n) \approx 0 \text{ for } n > l_1$$

Applying this approximation to the cost function (3.21), gives

$$J_N(\rho) \approx \frac{1}{N^2} \sum_{n=0}^{l_1} h(n,\rho)^2$$
 (3.22)

Example Let's take a very simple first order system.

$$H(z^{-1}) = \frac{0.2z^{-1}}{1 - 0.8z^{-1}} \tag{3.23}$$

The first 51 samples of the impulse response of (3.23) is plotted in figure 3.2. Additionally, the red dotted line shows the impulse response with a bit of additive Gaussian white noise.

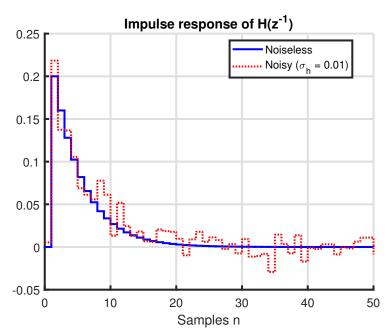


Figure 3.2 – Impulse response of (3.23) with and without additive Gaussian white noise.

The approximate cost function of the signal in the TD $\frac{1}{51^2}\sum_{n=0}^{l_1}h(n,\rho)^2$ is plotted in figure 3.3 (blue full line). It quickly converges to the cost function of the signal in the FD $\frac{1}{51}\sum_{k=0}^{50}|H(k)|^2$. This is also the case for the noisy signal (dotted red line); it converges to its cost in the FD. Ideally however, we would like the cost in the noisy case to be as close to the actual, noiseless cost. This is not the case; there is a certain bias. In this case, taking $l_1=10$, would result in less bias, while still keeping the information that is needed. This can also be seen in figure 3.2: the impulse response after n=10 is at or below noise level. Thus the impulse response after n=10 does not contain much useful information and will only contribute to a biased cost.

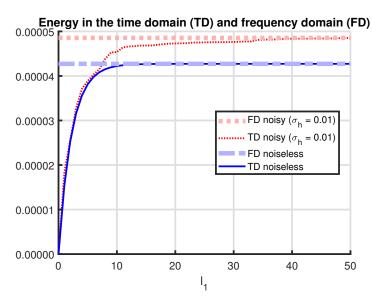


Figure 3.3 – Energy of the signal in the TD and the FD (divided by N).

In [1], l_1 has a slightly different interpretation. As it can be seen in (3.10), the sum goes from $-l_1$ to l_1 . This is because the sum is taken over the cross-correlation of u_W and ϵ . The cross-correlation is meaningful for both positive and negative indices, which is why the sum also extends into negative values of τ . The impulse response for negative indices is zero for causal systems, which is why the sum starts at n = 0 in (3.22).

3.8 Bias

Let's quantify the bias that was mentioned in the previous section. Let's assume that the input is a periodic signal that excites all the DFT frequencies. The output is perturbed by filtered white noise.

$$Y^{(p)}(k) = G(\Omega_k)U_0(k) + S_v(\Omega_k)V^{(p)}(k)$$

with $\mathbb{E}\{V^{(p)}(k)\}=0$ and $\mathbb{E}\{|V^{(p)}(k)|^2\}=\sigma^2/N$. The nonparametric estimate of the FRF is

$$\hat{G}(\Omega_k) = \frac{\frac{1}{P} \sum_{p=0}^{P-1} Y^{(p)}(k)}{\frac{1}{P} \sum_{p=0}^{P-1} U^{(p)}(k)} = G(\Omega_k) + \frac{1}{U_0(k)} \frac{1}{P} \sum_{p=0}^{P-1} S_v(\Omega_k) V^{(p)}(k)$$

The statistical properties of this estimator are

$$\mathbb{E}\{\hat{G}(\Omega_k)\} = G(\Omega_k)$$
$$\operatorname{Var}\{|\hat{G}(\Omega_k)|\} = \frac{1}{NP} \frac{|S_v(\Omega_k)|^2 \sigma^2}{|U_0(k)|^2}$$

 $|S_v(\Omega_k)|^2\sigma^2$ quantifies the power of the noise at the k-th DFT bin and $|U_0(k)|^2$ quantifies the energy of the signal at the k-th bin. This becomes evident when the RMS of the signals is calculated.

$$RMS^{2} = \frac{1}{N} \sum_{n=0}^{N-1} x(n)^{2} = \sum_{k=0}^{N-1} |X(k)|^{2}$$

Thus, $\frac{|S_v(\Omega_k)|^2\sigma^2}{|U_0(k)|^2}$ is the noise-to-signal ratio at the k-th DFT bin. Now we have the information we need to calculate the expected value of the cost function.

$$J_N(\rho) = \frac{1}{N} \sum_{k=0}^{N-1} |H(\Omega_k, \rho)|^2 = \frac{1}{N} \sum_{k=0}^{N-1} \left| F(\Omega_k) (1 - M(\Omega_k)) \left[M(\Omega_k) - (1 - M(\Omega_k)) K(\Omega_k, \rho) \hat{G}(\Omega_k) \right] \right|^2$$

Taking the expected value gives

$$\mathbb{E}\{J_{N}(\rho)\} = \frac{1}{N} \sum_{k=0}^{N-1} \left| F(\Omega_{k})(1 - M(\Omega_{k})) \left[M(\Omega_{k}) - (1 - M(\Omega_{k}))K(\Omega_{k}, \rho)G(\Omega_{k}) \right] \right|^{2}$$

$$+ \frac{1}{N} \sum_{k=0}^{N-1} \frac{1}{NP} \frac{|S_{v}(\Omega_{k})|^{2} \sigma^{2}}{|U_{0}(k)|^{2}} \left| F(\Omega_{k})(1 - M(\Omega_{k}))^{2}K(\Omega_{k}, \rho) \right|^{2}$$

$$= \tilde{J}_{N}(\rho) + \frac{\sigma^{2}}{N^{2}P} \sum_{k=0}^{N-1} \frac{|F(\Omega_{k})|^{2}|1 - M(\Omega_{k})|^{4}|K(\Omega_{k}, \rho)|^{2}}{\mathrm{SNR}(k)}$$

with $\tilde{J}_N(\rho)$ being the community contains the noiseless case and $SNR(k) = \frac{|U_0(k)|^2}{|S_v(\Omega_k)|^2\sigma^2}$. So, it has been shown that the cost function in the noisy case is not equal to the noiseless cost function. This is not a problem if the second term does not depend on ρ , as the optimization parameters that minimize $\mathbb{E}\{J_N(\rho)\}$ would also number $\tilde{J}_N(\rho)$. However, in this case the second term does depend on ρ , which is what causes this estimator to be biased in the presence of noise.

What if we now transform $H(\Omega_k, \rho)$ to the TD using the IDFT and approximate the cost function by only summing till l_1 ?

$$J_N(\rho) = \frac{1}{N} \sum_{k=0}^{N-1} |H(\Omega_k, \rho)|^2 = \frac{1}{N^2} \sum_{n=0}^{N-1} h(n, \rho)^2 \approx \frac{1}{N^2} \sum_{n=0}^{l_1} h(n, \rho)^2 = J_{N, l_1}(\rho)$$

Because the sum only contains (l_1+1) noisy terms, the bias will be smaller compared with the full sum. To be more specific, the bias will be a factor $(l_1+1)/N$ smaller. Thus the expected value of the approximate cost function becomes

$$\mathbb{E}\{J_{N,l_1}(\rho)\} \approx \tilde{J}_N(\rho) + \frac{\sigma^2}{N^2 P} \frac{l_1 + 1}{N} \sum_{k=0}^{N-1} \frac{|F(\Omega_k)|^2 |1 - M(\Omega_k)|^4 |K(\Omega_k, \rho)|^2}{\text{SNR}(k)}$$
(3.24)

Thus, the bias can be decreased by only taking the first $l_1 + 1$ terms in the TD. However, making l_1 too small will invalidate the approximation $J_N(\rho) \approx J_{N,l_1}(\rho)$. Additionally, the bias can also be decreased by increasing the number of periods P, increasing the number of samples per period N or by increasing the signal-to-noise ratio SNR(k).

3.9 Weighted nonlinear least squares

It is possible to reduce the variability by using a weighted cost function. The weight will be based on the variance of the FRF estimate.

$$\sigma_{\hat{G}}^2(\Omega_k) = \mathbb{E}\{|\hat{G}(\Omega_k) - \mathbb{E}\{\hat{G}(\Omega_k)\}|^2\}$$

The robust LPM can estimate the variance of the FRF estimate $\hat{\sigma}_{\hat{G}}^2$ for systems excited by periodic inputs. The weighted nonlinear least squares cost function is

$$J_{\text{WNLS}}(\rho) = \frac{1}{|\mathbf{K}_{\text{exc}}|} \sum_{k \in \mathbf{K}_{\text{exc}}} \frac{|H(\Omega_k, \rho)|^2}{\hat{\sigma}_H^2(\Omega_k, \rho)}$$
(3.25)

with K_{exc} being the set of excited DFT bins and

$$H(\Omega_k, \rho) = F(\Omega_k)(1 - M(\Omega_k)) \Big[M(\Omega_k) - (1 - M(\Omega_k)) K(\Omega_k, \rho) \hat{G}(\Omega_k) \Big]$$

and

$$\hat{\sigma}_H^2(\Omega_k, \rho) = \hat{\sigma}_{\hat{G}}^2(\Omega_k) \Big| F(\Omega_k) (1 - M(\Omega_k))^2 K(\Omega_k, \rho) \Big|^2$$

The expected value of (3.25) is

$$\begin{split} \mathbb{E}\{J_{\text{WNLS}}(\rho)\} &= \frac{1}{|\mathbf{K}_{\text{exc}}|} \sum_{k \in \mathbf{K}_{\text{exc}}} \frac{|F(1-M)|^2 |M - (1-M)K(\rho)G|^2}{|F(1-M)^2 K(\rho)|^2 \hat{\sigma}_{\hat{G}}^2} \\ &+ \frac{1}{|\mathbf{K}_{\text{exc}}|} \sum_{k \in \mathbf{K}_{\text{exc}}} \frac{|F(1-M)^2 K(\rho)|^2 \sigma_{\hat{G}}^2}{|F(1-M)^2 K(\rho)|^2 \hat{\sigma}_{\hat{G}}^2} \\ &= \tilde{J}_{\text{WNLS}}(\rho) + \frac{1}{|\mathbf{K}_{\text{exc}}|} \sum_{k \in \mathbf{K}} \frac{\sigma_{\hat{G}}^2(\Omega_k)}{\hat{\sigma}_{\hat{G}}^2(\Omega_k)} \end{split}$$

with $\tilde{J}_{\text{WNLS}}(\rho)$ being the cost function when $G(\Omega_k)$ is known perfectly. The frequencies Ω_k are left out in the first equation for simplicity. The second term does not depend on ρ , which means that the optimization parameters ρ that minimize $\tilde{J}_{\text{WNLS}}(\rho)$ will also minimize $\tilde{J}_{\text{WNLS}}(\rho)$. After all, the second term is just a constant which ρ .

Realizable controller If the controller is realizable, then

$$\tilde{J}_{\text{WNLS}}(\rho^*) = 0$$

The convex cost function (3.17) is also zero in $\rho = \rho^*$ in the noiseless case. This means that the original cost function (3.2), the convex cost function (3.4) and the WNLS cost function (3.25) are all minimal in $\rho = \rho^*$. Therefore, if the controller is realizable, then the WNLS optimization has the potential to give a better estimate of the optimal controller as the WNLS cost function also takes the noise variance into account.

Non realizable controller On the other, what happens if the controller is not realizable? In this case, the minimum of any of the costs will be greater than 0, even in the noiseless case. Thus, the optimization parameters ρ that minimize the original cost function, the convex cost function and the WNLS cost function will be different in general.

Optimization strategy The cost function (3.25) is not convex any more as the denominator also depends on the optimization parameters ρ . Thus, the minimization of this cost function cannot be solved with convex optimization. It can be solved with the Gauss-Newton algorithm. An initial estimate of ρ can be found by minimizing the convex cost function (3.17). The danger however, is that this optimization might not converge to the global minimum.

3.10 Discrete time simulations

First, an important approximation was made in the previous sections. Initially, it was assumed that the sensitivity function can be approximated by the ideal sensitivity function.

$$\frac{1}{1+K(\Omega,\rho)G(\Omega)}\approx\frac{1}{1+K^*(\Omega)G(\Omega)}$$

This was done in order to make the cost function convex. Thus we should verify after optimization that the original cost function (3.2) can be approximated by the convex cost function (3.4).

$$J_{mr}(\rho) \approx J(\rho)$$

System models In order to compare the FD methods to the TD methods, we will initially work with DT systems. 3 DT systems $G(q^{-1})$ will be simulated. The first 2 systems will have a reference model $M(q^{-1})$ that will be realizable by the proposed control structure $K(q^{-1}, \rho)$. The third system will have a reference model that cannot be realized by the proposed control structure. The systems, the reference model and the proposed control structure are given in table 3.1. These are also shown in figure 3.4. The systems will start with zero initial conditions.

Name	Simple system	Long transient system	
$G(q^{-1})$	$0.25q^{-1}$	$1.785q^{-1} + 1.701q^{-2}$	
	$\overline{1 - 0.75q^{-1}}$	$\overline{1 + 1.558q^{-1} + 0.9274q^{-2}}$	
$M(q^{-1})$	$0.15q^{-1} - 0.075q^{-2}$	$0.1785q^{-1} + 0.3486q^{-2} + 0.1701q^{-3}$	
	$\overline{1 - 1.6q^{-1} + 0.675q^{-2}}$	$1 + 0.7369q^{-1} - 0.2824q^{-2} - 0.7573q^{-3}$	
$K(\rho, q^{-1})$	$\rho_0 + \rho_1 q^{-1}$	$\rho_0 + \rho_1 q^{-1}$	
	$1 - q^{-1}$	$\frac{1 - q^{-1}}{1 - q^{-1}}$	
Realizable	Yes	Yes	
Name	Non realizable controller		
$G(q^{-1})$	$0.7893q^{-3}$		
	$\overline{1 - 1.418q^{-1} + 1.59q^{-2} - 1.316q^{-3} + 0.886q^{-4}}$		
$M(q^{-1})$	$0.1552q^{-3}$		
	$1 - 1.212q^{-1} + 0.3672q^{-2}$		
$K(\rho, q^{-1})$	$\rho_0 + \rho_1 q^{-1} + \rho_2 q^{-2} + \rho_3 q^{-3}$		
	$\frac{1 - q^{-1}}{1 - q^{-1}}$		
Realizable		No	

Table 3.1 – Simulated system and the reference models.

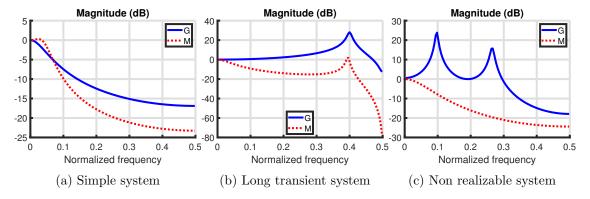


Figure 3.4 – Magnitude bode plots of the system and the reference models.

Noise models The output of the systems will be perturbed by DT filtered white noise.

$$y(n) = G(q^{-1})u(n) + S_u(q^{-1})e(n)$$
 with $\mathbb{E}\{e(n)\} = 0$ and $\mathbb{E}\{e(n)^2\} = \sigma^2$

2 different noise models will be used:

• $S_y(q^{-1}) = 1$, i.e white noise

•
$$S_y(q^{-1}) = \frac{0.1105q^{-1} - 0.06831q^{-2} + 0.04222q^{-3} + 0.04222q^{-3} - 0.06831q^{-4} + 0.1105q^{-5}}{1 - 0.3337q^{-1} - 0.3872q^{-2} - 0.1103q^{-3}}$$

The magnitude bode plots of the noise models are shown in figure 3.5. For every simulation the noise standard deviation is set to $\sigma = 0.2$.

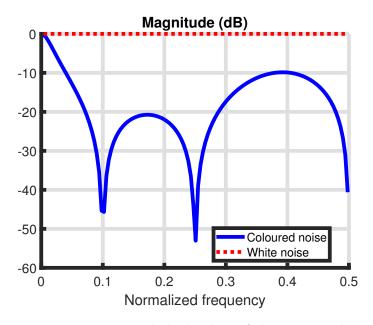


Figure 3.5 – Magnitude bode plots of the noise models.

Optimization strategies As explained in the previous sections, there are multiple ways to find the optimal controller.

First, the optimal ρ for a certain frequency resolution F_s/N can be found by minimizing (3.17) and by replacing $\hat{G}(\Omega_k)$ with the actual system $G(\Omega_k)$. This can be done because we are working in a simulation. Note however, that we are optimizing the convex cost function. If the reference model M is realizable, then $J_{mr}(\rho^*) = J(\rho^*) = 0$, which means that, in the noiseless case, the parameters ρ that minimize the convex cost function will also minimize the original cost function. This is not the case any more if the reference model is not realizable.

The TD method consists of minimizing (3.11) w.r.t. ρ . However, the parameter l_1 must be chosen by the user. The maximum value that can be chosen for l_1 is N/2.

The FD method consists of minimizing (3.17) w.r.t. ρ . The l_1 -trick can also be used here by minimizing $J_{N,l_1}(\rho) = \frac{1}{N^2} \sum_{n=0}^{l_1} h(n,\rho)^2$ with $h(n,\rho)$ being the IDFT of $H(\Omega_k,\rho)$. Notice that when $l_1 = N - 1$, $J_{N,l_1}(\rho) = J_N(\rho)$ in the FD.

Finally, the WNLS method will also be used by minimizing (3.25). This cost function will be minimized by using the Gauss-Newton method. The maximum number of iterations is set to 100 and the algorithm is stopped if the relative difference between the previous and the next estimate is smaller than 10^{-10} for all ρ_i in ρ . An initial estimate for ρ is found by using the FD method with $l_1 = N - 1$. As there is no guarantee that this optimization will converge to the global minimum, the optimization will also be started at the optimal ρ for comparison. Of course, this can only be done because we are working in a simulation.

Input The input will be a random phase multisine where all harmonics are excited. The RMS of this multisine is 1 and it is repeated for P = 4 periods. None of the periods are discarded. The number of samples per period is N = 255.

Nonparametric estimate The FD methods require a nonparametric estimate. Two different nonparametric estimates are used. The first one is without any removal of system and noise transients by using (3.19). The second way of obtaining a nonparametric estimate is by using the robust LPM with order R = 2 and degrees of freedom $q^{\text{noise}} = 1$. These values were obtained by using the heuristics mentioned in section 2.14.5. Finally, the WNLS method needs the variance of the FRF estimate. This is also obtained via the robust LPM.

Evaluation 100 noise realizations are simulated. The resulting controllers are compared by evaluating the cost function with the real system $G(\Omega)$. Both the original cost function and the convex cost function are calculated.

$$J_{mr}(\rho) = \frac{1}{127} \sum_{k=1}^{127} \left| M(\Omega_k) - \frac{K(\Omega_k, \rho)G(\Omega_k)}{1 + K(\Omega_k, \rho)G(\Omega_k)} \right|^2$$
$$J(\rho) = \frac{1}{127} \sum_{k=1}^{127} \left| (1 - M(\Omega_k)) \left[M(\Omega_k) - (1 - M(\Omega_k))K(\Omega_k, \rho)G(\Omega_k) \right] \right|^2$$

where ρ is obtained by optimization for one noise realization with any of the methods. Note that $F(\Omega_k) = 1$ in the above equations. The sum starts at 1 because DC is not excited by the input. The sum ends at 127 because there are 127 excited harmonics in the input. Then the cost of the controllers obtained from the different noise realizations can be averaged.

$$\bar{J}_{mr} = \frac{1}{100} \sum_{i=1}^{100} J_{mr}(\rho^{(i)})$$
$$\bar{J} = \frac{1}{100} \sum_{i=1}^{100} J(\rho^{(i)})$$

where $\rho^{(i)}$ are the parameters obtained by optimizing the cost corresponding to the *i*-th noise realization with one of the optimization methods. Note that TD methods with different values of l_1 are considered different and shouldn't be mixed up when taking the mean. Finally, the mean cost function can be expressed in decibels.

$$\bar{J}_{mr}|_{dB} = 10 \log_{10}(\bar{J}_{mr})$$

 $\bar{J}|_{dB} = 10 \log_{10}(\bar{J})$

Influence of l_1 Let's start by seeing how the choice of l_1 influences the performance of the resulting controller. The simple system is simulated with disturbing white output noise as explained above. For the TD method, all values of l_1 between 1 and $\lfloor N/2 \rfloor = 127$ are tried for each noise realization. For the FD method, all values of l_1 between 1 and N-1=254 are tried for each noise realization. The resulting closed loop system that is obtained by applying the FD method with $l_1=1,11$ and 254 is shown in figure 3.6. The closed loop system resulting from $l_1=1$ performs poorly and doesn't get close to the reference model M. The result is better when $l_1=254$, but is best when $l_1=17$.

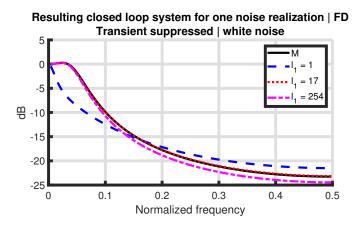
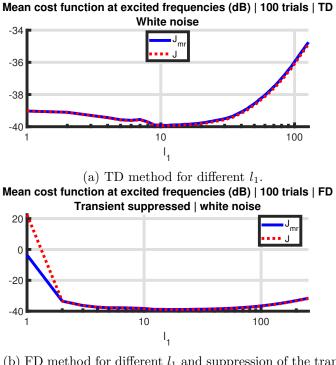


Figure 3.6 – Resulting CL system for one white noise realization by applying the FD method with different values of l_1 .

The mean cost function for the TD and FD method for different values of l_1 is shown in figure 3.7. The mean cost function has a bowl shape as a function of l_1 . This is because too much information is thrown away when taking a too small l_1 . However, taking a big l_1 just increases the bias. This was also observed in the figure 3.6. One more thing to note is that the original cost function J_{mr} coincides with the convex cost function J most of the time. This means that the sensitivity function can be approximated quite well by the ideal sensitivity function. The cost functions don't coincide for the FD method when $l_1 = 1$. This is because the resulting closed loop system doesn't come very close to the reference model M, as can be seen figure 3.6.



(b) FD method for different l_1 and suppression of the transient.

Figure 3.7 – Mean cost functions for 100 noise realizations, applied to the simple system with a white noise model.

Transient suppression In the previous experiment, the transient was suppressed by using the robust LPM. What will happen if the transient is not suppressed? The mean cost function obtained when applying the FD methods with $l_1 = 254$ with and without transient suppression are given in table 3.2. Not suppressing the transient gives better results. This is to be expected as the simple system has a very low transient. Moreover, there are no noise transients because the noise is white.

Method	$\bar{J} _{dB}$	$J_{mr}^- _{dB}$
Transient suppressed	-31.74	-31.57
Transient not suppressed	-34.86	-34.73

Table 3.2 – FD method with $l_1 = 254$ applied to simple system with white noise. With and without transient suppression.

Let's see what happens when we try coloured noise. The results of the same experiment, but with coloured noise are shown in table 3.3. This time, not suppressing the transient is still better.

Method	$\bar{J} _{dB}$	$\bar{J_{mr}} _{dB}$
Transient suppressed	-50.99	-50.98
Transient not suppressed	-52.51	-52.52

Table 3.3 – FD method with $l_1 = 254$ applied to simple system with coloured noise. With and without transient suppression.

To see why this is the case, we can look at the MSE error of the nonparametric estimate of the transfer function over all the noise realizations. This is plotted in figure 3.8. The

estimate is better around the transmission zeros of the noise model. To understand why this is so, let's assume that the noise model is given by

$$S_y(q^{-1}) = \frac{C(q^{-1})}{D(q^{-1})}$$

In that case, the contribution of the noise on the output at the k-th bin is

$$\frac{C(\Omega_k)}{D(\Omega_k)}E(k) + \frac{I_E(\Omega_k)}{D(\Omega_k)} = \frac{C(\Omega_k)E(k) + I_E(\Omega_k)}{D(\Omega_k)}$$

with E(k) being the DFT of e(n) and $I_E(\Omega_k)$ being a polynomial in Ω_k that depends on the initial and end conditions of e(n). When $C(\Omega_k)$ is small, then the contribution of the noise is mainly attributed to the transient term $\frac{I_E(\Omega_k)}{D(\Omega_k)}$. Thus, the transient term is dominant at transmission zeros of the noise model. However, at the other frequencies where the random term $\frac{C(\Omega_k)}{D(\Omega_k)}E(k)$ is dominant, the estimate of the FRF is slightly worse (around 1 dB) when taking the transient into account. As the optimization takes all the frequency bins into account, an improvement around the transmission zeros is not enough to improve the estimate of the optimal controller in this case.

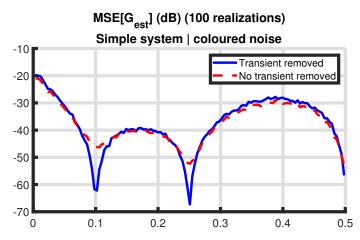


Figure 3.8 – MSE error of the nonparametric estimate of the simple system in the presence of coloured noise with and without transient suppression.

TD vs. FD vs WNLS

Long transient system

Non realizable system

3.11 Continuous time simulations

3.12 Conclusion

Appendix

3.A Division by auto-power spectrum

In section 3.5, the following filter is defined.

$$W(q^{-1}) = \frac{F(q^{-1})(1 - M(q^{-1}))}{S_{UU}(q^{-1})}$$

with $S_{UU}(q^{-1})$ being a parametric representation of the auto-power spectrum of u(n).

If u(n) is DT filtered white noise $u(n) = S_u(q^{-1})e_u(n)$ and a parametric representation of $S_u(q^{-1})$ is known, then the auto-power spectrum of u(n) becomes

$$S_{UU}(q^{-1}) = S_u(q^{-1})S_u(q)$$

The zeros of $S_u(q^{-1})$ and $S_u(q)$ have an inverse relationship. If a is a zero of $S_u(q^{-1})$, then a^{-1} is a zero of $S_u(q)$. These zeros then become poles of $W(q^{-1})$. Thus, if $S_u(q^{-1})$ contains zeros that are not on the unit circle, the filter $W(q^{-1})$ will be unstable. And so, applying the filter in the time domain is in general not possible.

3.B DFT of cross-correlation

The circular cross-correlation of two discrete signals x and y is defined by [6, eq. (2.22)]

$$R_{xy}(\tau) = \sum_{n=0}^{N-1} \overline{x(n)} y(\text{mod}(n+\tau, N))$$
 (3.26)

Taking the DFT of (3.26) gives

$$\begin{split} S_{XY}(k) &= \frac{1}{N} \sum_{\tau=0}^{N-1} R_{xy}(\tau) e^{-j2\pi k\tau/N} = \frac{1}{N} \sum_{\tau=0}^{N-1} \sum_{n=0}^{N-1} \overline{x(n)} y(\operatorname{mod}(n+\tau,N)) e^{-j2\pi k\tau/N} \\ &= \frac{1}{N} \sum_{n=0}^{N-1} \overline{x(n)} \sum_{\tau=0}^{N-1} y(\operatorname{mod}(n+\tau,N)) e^{-j2\pi k\tau/N} \\ &= \frac{1}{N} \sum_{n=0}^{N-1} \overline{x(n)} e^{j2\pi kt/N} \sum_{\tau=0}^{N-1} y(\operatorname{mod}(n+\tau,N)) e^{-j2\pi k(n+\tau)/N} \end{split}$$

The second sum is actually independent of n.

$$\begin{split} \sum_{\tau=0}^{N-1} y(\text{mod}(n+\tau,N)) e^{-j2\pi k(n+\tau)/N} &= \sum_{\tau=0}^{N-t-1} y(n+\tau) e^{-j2\pi k(n+\tau)/N} \\ &+ \sum_{\tau=N-t}^{N-1} y(n+\tau-N) e^{-j2\pi k(n+\tau)/N} \\ &= \sum_{\tau=0}^{N-t-1} y(n+\tau) e^{-j2\pi k(n+\tau)/N} \\ &+ \sum_{\tau=-t}^{-1} y(n+\tau) e^{-j2\pi k(n+\tau)/N} \\ &= \sum_{\tau=-t}^{N-t-1} y(n+\tau) e^{-j2\pi k(n+\tau)/N} \end{split}$$

In the last step we used the fact that $e^{-j2\pi k(n+\tau+N)/N} = e^{-j2\pi k(n+\tau)/N}$. Now by doing one last substitution we get

$$\sum_{\tau=0}^{N-1} y(\text{mod}(n+\tau,N)) e^{-j2\pi k(n+\tau)/N} = \sum_{\tau=0}^{N-1} y(\tau) e^{-j2\pi k\tau/N}$$

This equation is valid because the sum runs over all the samples $0, \ldots, N-1$. Finally, the DFT of the cross-correlation is

$$S_{XY}(k) = \frac{1}{N} \sum_{n=0}^{N-1} x(n) e^{-j2\pi kt/N} \sum_{\tau=0}^{N-1} y(\tau) e^{-j2\pi k\tau/N}$$

$$\Rightarrow S_{XY}(k) = N \overline{X(k)} Y(k)$$

3.C Unstable systems

Open loop experiments are not practical on unstable systems. If we want to measure the FRF of an unstable system, it will have to be done in closed loop with a stabilizing controller. As discussed in section 2.15.3, the danger here is that process noise on the output will be fed back to the input. Not taking the right precautions will lead to an inconsistent nonparametric estimate of the FRF in some cases.

First, a quick summary of the solution given in [1] will be discussed. Then it will be shown that a nonparametric estimate of the FRF is again hidden in the maths.

3.C.1 Correlation-based approach

Model The set-up shown in figure 3.C.1 is considered. An unstable system $G(q^{-1})$ is stabilized by a controller $K_s(q^{-1})$ in negative feedback. The reference signal r(n) is entered between the controller and the system. The reference signal is assumed to be periodic. The input to the system is u(n) and the output of the system y(n) is perturbed by process noise v(n). The signal x(n) is the reference signal that would be used in the actual closed loop system. However, to keep the notation similar to [1], x(n) is set to 0. The closed loop TF from x(n) to y(n) is

$$M_s(q^{-1}) = \frac{K_s(q^{-1})G(q^{-1})}{1 + K_s(q^{-1})G(q^{-1})}$$

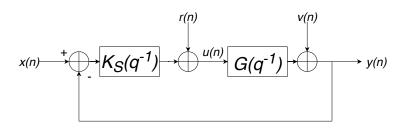


Figure 3.C.1 – Unstable LTI system with a stabilizing controller in feedback.

Error signal The error signal is the same as in section 3.5.

$$\epsilon(n, \rho) = Mu(n) - K(\rho)(1 - M)y(n)$$

Reference filtering This time, the reference signal r(n) is filtered instead of the input to the system u(n). The filter $W(q^{-1})$ which is used to obtain $r_W(n)$ is defined differently this time.

$$W(e^{-j\omega}) = \frac{F(e^{-j\omega})(1 - M(e^{-j\omega}))}{(1 - M_s(e^{-j\omega}))S_{RR}(e^{-j\omega})}$$
(3.27)

with $S_{RR}(e^{-j\omega})$ being the auto-power spectrum of the reference signal. It is pointed out in [1] that $W(e^{-j\omega})$ cannot be implemented in this way, because M_s is not known as we don't have access to a parametric representation of G. This is solved in the following way. First, u(n) is written as a function of r(n) and v(n).

$$u(n) = \frac{1}{1 + K_s G} r(n) - \frac{K_s}{1 + K_s G} v(n)$$

Assuming that there is no noise v(n) = 0 and rewriting $1 + K_sG$ in terms of M_s , gives

$$u(n) = (1 - M_s)r(n)$$

Hence, there is a relation between the auto-power spectrum of r(n) and the cross-power spectrum between r(n) and u(n).

$$S_{RU}(e^{-j\omega}) = (1 - M_s(e^{-j\omega}))S_{RR}(e^{-j\omega})$$

And so (3.27) becomes

$$W(e^{-j\omega}) = \frac{F(e^{-j\omega})(1 - M(e^{-j\omega}))}{S_{RU}(e^{-j\omega})}$$

 $S_{RU}(e^{-j\omega})$ can be estimated from the data and the filter can be applied to r(n) in the FD to obtain $r_W(n)$.

Correlation criterion Then, the cross-correlation between $r_W(n)$ and $\epsilon(n,\rho)$ is calculated.

$$R_{r_W \epsilon}(\tau, \rho) = \frac{1}{NP} \sum_{n=0}^{NP-1} r_W(n-\tau) \epsilon(n, \rho)$$

with P being the number of periods measured. The cost function can then be calculated by using (3.11).

3.C.2 Nonparametric estimate

The error signal in the FD becomes

$$E(kP, \rho) = M(\Omega_k)U(kP) - K(\Omega_k, \rho)(1 - M(\Omega_k))Y(kP)$$

Here, because the reference signal is periodic, the frequencies Ω_k correspond to the DFT bins kP. For the rest of this section, the frequencies Ω_k will be left out from the equations for clarity. The filtered reference signal becomes

$$R_W(kP) = \frac{F(1-M)R(kP)}{R(kP)\overline{U(kP)}}$$

Then the cross-power spectrum between $\epsilon(n,\rho)$ and $r_W(n)$ is

$$S_{R_WE}(\Omega_k, \rho) = R_W(kP)\overline{E(kP, \rho)} = F(1 - M)\overline{\left[M\frac{U(kP)\overline{R(kP)}}{U(kP)\overline{R(kP)}} - K(\rho)(1 - M)\frac{Y(kP)\overline{R(kP)}}{U(kP)\overline{R(kP)}}\right]}$$
$$= F(1 - M)\overline{[M - K(\rho)(1 - M)\hat{G}(\Omega_k)]}$$

with

$$\hat{G}(\Omega_k) = \frac{Y(kP)\overline{R(kP)}}{U(kP)\overline{R(kP)}} = \frac{Y(kP)}{U(kP)}$$
(3.28)

If \hat{G} is replaced by the actual system G, $S_{R_WE}(\Omega_k, \rho)$ is exactly the quantity being integrated over in (3.4). As was discussed in section 3.5, (3.28) is equivalent to taking the DFT of every period and taking the mean of the spectra over the periods.

$$\hat{G}(\Omega_k) = \frac{\frac{1}{P} \sum_{p=0}^{P-1} Y^{(p)}(k)}{\frac{1}{P} \sum_{p=0}^{P-1} U^{(p)}(k)}$$

Note that this estimate is consistent when the excitation is periodic. For arbitrary excitations it is inconsistent. It is interesting to see that the unsimplified fraction in (3.28) looks like the indirect method for estimating the FRF (see section 2.15.3).

$$\hat{G}(\Omega_k) = \frac{\frac{1}{P} \sum_{m=0}^{P} Y^{(m)}(k) \overline{R^{(m)}(k)}}{\frac{1}{P} \sum_{m=0}^{P} U^{(m)}(k) \overline{R^{(m)}(k)}}$$

This FRF estimate is consistent when using arbitrary excitations.

Chapter 4 Guaranteed stability

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