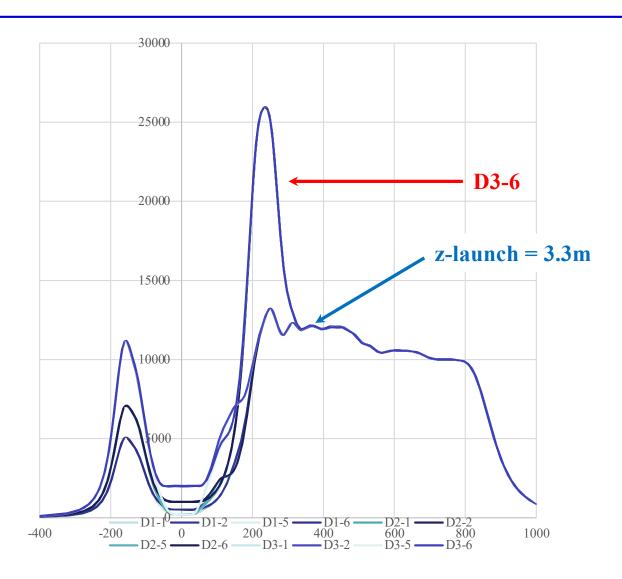
Preliminary RAYS results for MPEX, OXB

August, 2025 D. B. Batchelor

- RAYS added multiple mirror equilibria, post-processing and graphics, Ordinary to Extraordinary mode conversion coefficients
- ⇒ Can now trace cold-plasma rays in MPEX and calculate OX mode conversion
- Ran several cases for 2nd harmonic ECH OX conversion
 - Tried to extract parameters from various MPEX design documents
 - Several unknowns, uncertainties, and inconsistencies
 - Put together parameter sets representing what I think, that you think are representative of MPEX
 - Looked at rays, and conversion coefficients for off-axis launch haven't seen any such in MPEX documentation
- Need more information before doing more → especially a good characterization of rays launched from the antenna

Magnetics information from Earle Burkhardt (spread sheet)



- Case D3-6
- Field in RAYS is calculated using elliptic integrals (not paraxial approximation), spline fit

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Hyperbolic density model based on Marin, MPEX-03-003

$$n_e(r) = \left[n_e(0) - n_e(r_{edge})\right] \frac{\tanh\left(\frac{\psi + \psi_P}{\Delta}\right) - \tanh\left(\frac{\psi - \psi_P}{\Delta}\right)}{2\tanh\left(\frac{\psi_P}{\Delta}\right)} + n_e(r_{edge})$$

• He takes $\psi_P = 0.6$, $\Delta = 0.19$, which seems to fit Proto MPEX

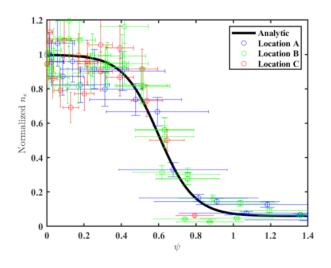
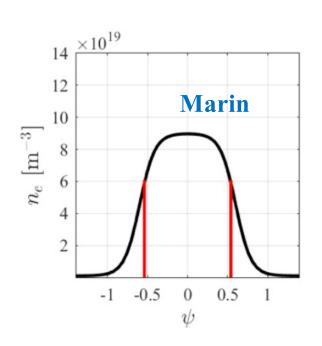


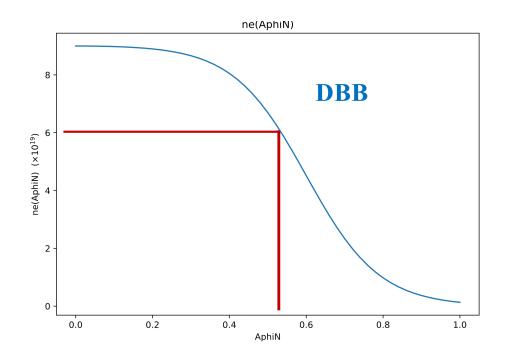
Figure 1, Experimentally measured radial electron density profiles normalized to the peak density at different axial locations along Proto-MPEX. Analytical fit to the data uses Eq. 9 with $\psi_p=0.6$ and $\Delta\psi=0.19$.

• He apparently uses the same parameters to model MPEX

However, I get a different radial profile from him for MPEX using these parameters

Density versus ψ are similar, $\psi_P = 0.6, \Delta = 0.19$

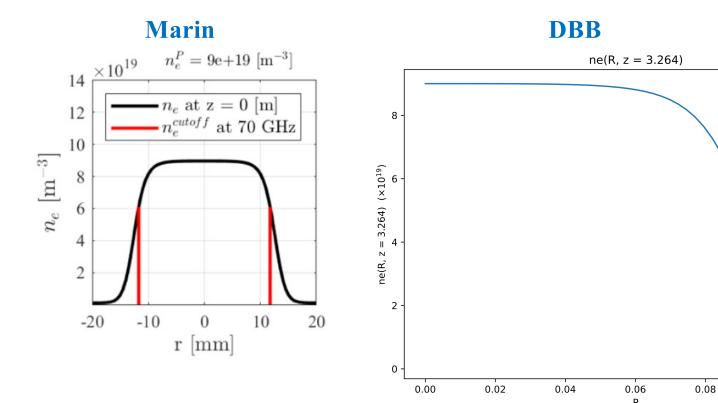




Cutoff density is at same location in ψ

Profiles in radius are very different

• Again $\psi_P = 0.6, \Delta = 0.19$

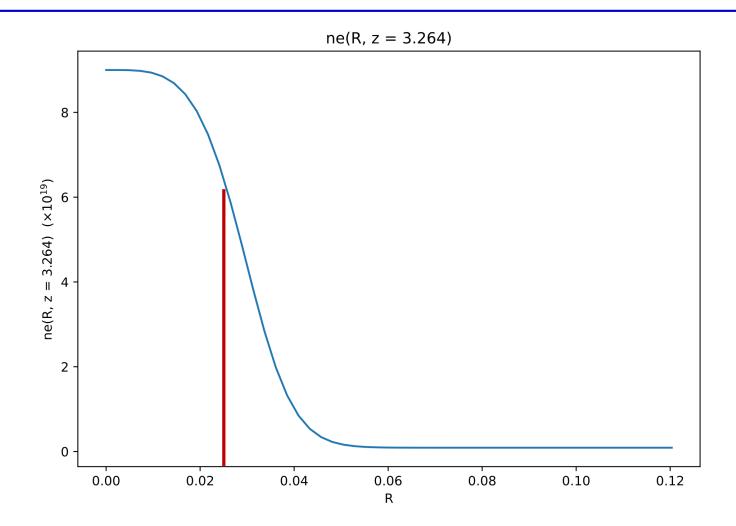


- His plasma is about 2 cm wide, mine is about 17 cm at z = 3.3m with these profile parameters
- Maybe the discrepancy is the definition of LUFS. I haven't been able to find out the limiter strike point for the D3-6 case.

0.10

0.12

Cases were run with narrower profile $\psi_P = 0.05$, $\Delta = 0.05$



• Plasma is about 4 cm wide at O-mode cutoff, z = 3.3m

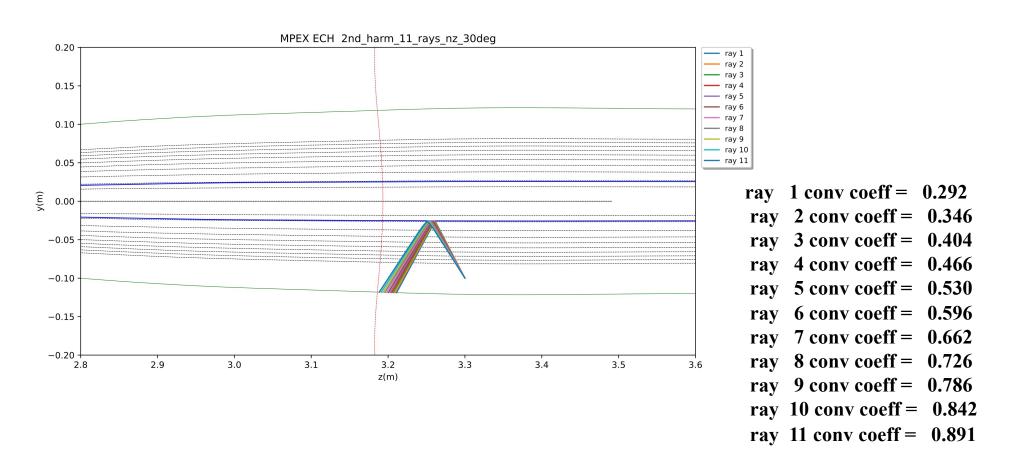
There appears to be a discrepancy in the optimum launch angle

• From Preinhalter and Kopecky (1973) the optimum value of $n_{||}$ is given by

$$n_{z\,opt} = \sqrt{\frac{\beta}{1+\beta}}$$
, where $\beta \equiv \frac{\Omega_{ce}}{\omega_{rf}}$

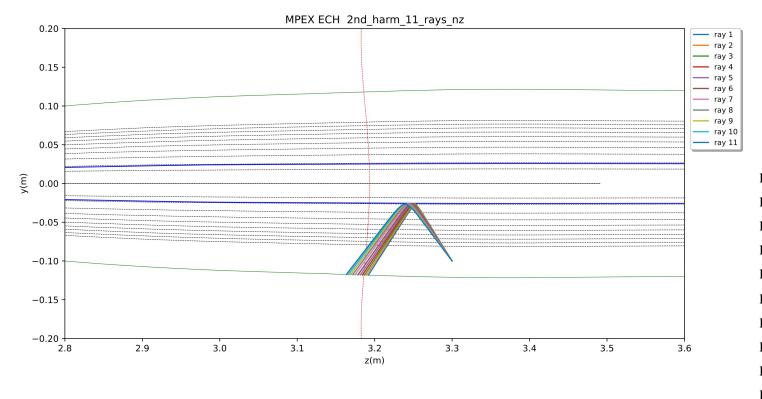
- For 2nd harmonic i.e. $\beta = 0.5$, $n_{zopt} = 0.577 \Rightarrow \theta_{opt} = 35.3^{\circ}$.
- Here θ is the angle between the density gradient (essentially Y here) and the incident wave vector. The magnetic field is taken to be perpendicular to the density gradient (essentially Z here)
- The MPEX design seems to shoot for a launch angle of 30°
- It makes a difference

Rays launched around 30° to Y, in Y-Z plane



- Launch angle varies from $30^{\circ} \pm 2.5^{\circ}$ in 0.5° steps
- Ray 6 is the 30° case, conversion is about 60%

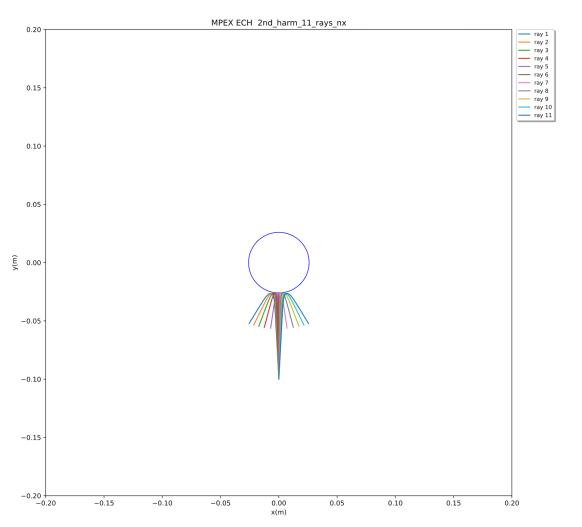
Rays launched around 35.3° to Y, in X-Z plane



ray 1 conv coeff = 0.842
ray 2 conv coeff = 0.891
ray 3 conv coeff = 0.932
ray 4 conv coeff = 0.964
ray 5 conv coeff = 0.986
ray 6 conv coeff = 0.998
ray 7 conv coeff = 0.999
ray 8 conv coeff = 0.989
ray 9 conv coeff = 0.971
ray 10 conv coeff = 0.945
ray 11 conv coeff = 0.913

- Launch angle varies from 35.3° \pm 2.5° in 0.5° steps
- Ray 6 is the 35.3° case, conversion is about 100%

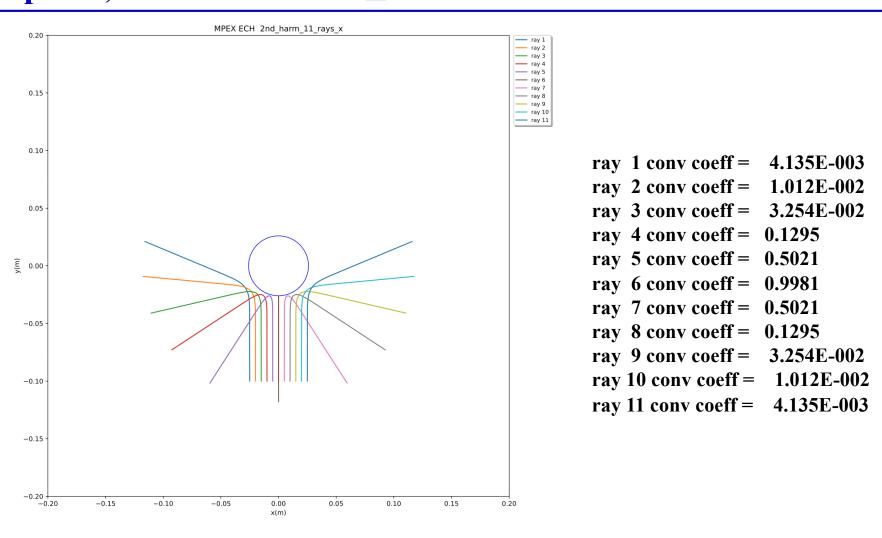
Rays launched not directly at axis, at 35.3° to Y in Y-Z plane, and around 0° in X-Y plane, $\pm 2.5^{\circ}$



ray 1 conv coeff = 0.468
ray 2 conv coeff = 0.596
ray 3 conv coeff = 0.733
ray 4 conv coeff = 0.863
ray 5 conv coeff = 0.960
ray 6 conv coeff = 0.998
ray 7 conv coeff = 0.960
ray 8 conv coeff = 0.863
ray 9 conv coeff = 0.733
ray 10 conv coeff = 0.596
ray 11 conv coeff = 0.468

- Launch angle varies from $0^{\circ} \pm 2.5^{\circ}$ in 0.5° steps
- Ray 6 is the 0° case, conversion is about 100%
- But drops off pretty fast if not aimed directly at axis

Rays launched not directly at axis – at 35.3° to Y in Y-Z plane, X-launch varies $\pm 2.5cm$ about X = 0



- Ray 6 is the X = 0 case, conversion is about 100%
- But drops off really fast if not aimed directly at axis

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Final points

- These results are preliminary. There are a number of inputs that need to be nailed down better before anything else is done. (Questions for MPEX team next slide)
- The calculation of O-X conversion gives an upper limit to the amount of O-X-B heating that can be obtained. Converted X-mode may not go and be absorbed where you want it, but first it has to be converted.
- This situation looks similar to me to what we had in Elmo Bumpy Torus (EBT). For that we developed a "power balance" model, which was essentially a three region zero-D model, supplemented with ray tracing. ("Detailed modeling of microwave energy deposition in EBT", Batchelor, Goldfinger, Rasmussen, Phys. Fluids 27, 1984). Something similar might be useful for MPEX.
- What I would really like to do is to turn this analysis over to an ORNL person so I can focus on adding the X-mode propagation and Bernstein mode propagation and absorption to the code.

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Questions for MPEX team

- What are the relevant coil current cases to examine? D3-6? Are there others? What is the operational rationale for the various cases?
- What is the location and geometry of the limiter? Or better yet:
- What are the strike points on the limiter for the various coil current cases? What is the LUFS for the various cases?
- What is the geometry of the first wall all along the device?
- Is there more recent thought on expected plasma profiles than is contained in Marin's 2021 report, MPEX-03-RTP-003?
- What is the geometry of the ECH launcher? The design is sketched out in MPEX-03-DAC-007, 2022, But I can't find that the actual location and orientation as built are specified there.
- The biggest uncertainty is how to represent the launched ECH power as an initial distribution of rays. I think these RAYS calculations do show how sensitive the O-X-B performance is to details of the launched antenna spectrum. I think it would be advisable to put some effort into refining such a model, maybe to the point of revisiting the COMSOL work.

Reality check – Are there important things that I have wrong or left out?

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