Simulation of a gas-driven rocket engine

Oskar Idland & Jannik Eschler Institute of Theoretical Astrophysics, University of Oslo (Dated: September 13, 2022)

This is an abstract Complete this summary at the end of the paper

I. INTRODUCTION

Humanity is the only known species to use and shape everything from elements to their entire planet to their benefit. Now, the next logical step is to look further and find out what there can be found beyond our home planet. There are different ways to get to, and explore space, but the most established way to do this is by using

Rocket engines, while highly advanced, still simply utilize Newton's laws of motion to create thrust. More specifically the third law, which states "For every action, there is an equal and opposite reaction". When wanting to propel our rocket engine as fast as possible upwards into space and being able to maneuver, we are therefore required to be able to expel the right amount of matter at the right time to create the force needed to propel the rocket as wanted.

The first step to creating such an engine is simulating both its inner workings and the engine as part of a rocket during a launch to determine engine-, rocketand launch-parameters as well as gaining a better understanding of the complex system, which makes up a functional rocket. This will happen in a simulated solar system created by the ast2000tools package The rocket in this paper will be using hot H_2 gas under high pressure which will be expelled out the end of the rocket engine to create a force. This is due to H_2 being an ideal gas, which simplifies calculations.

THEORY II.

To complete the calculations for the engine we are going to use a lot of statistics to simplify the behavior of the gas particles.

III. METHOD

As we are trying to simulate a rocket engine, using gas being expelled at high speed, it is important to be able to simulate and understand the gas. A gas is defined as "a fluid (such as air) that has neither independent shape nor volume but tends to expand indefinitely" and can be comprised of one or multiple individual atoms, or in the case of Hydrogen gas molecules. Hence, if we want to simulate a gas, we need to be able to simulate individual

molecules for themselves. To not make calculations excessively difficult we assume to have an ideal gas. This means it's particles can be looked at as point particles without any spatial extension and it's density and temperature being uniform throughout the gas. The particles will therefore be distributed at random positions. However, on the molecular scale some particles will have more energy than others, and since a molecules speed is tightly correlated to its energy, statistical physics need to be used to find their velocity. The most important function we will need is the gaussian probability distribution function, also called normal distribution function and given by

$$f(\mu, \sigma; x) = \frac{1}{\sqrt{2\pi}\sigma} exp \left[-\frac{1}{2} \left(\frac{x - \mu}{\sigma} \right)^2 \right]$$
 (1)

The mean value μ and standard deviation σ are parameters which help defining the distribution, and x is a free variable. The shape of the gaussian distribution is a very distinct bell curve. The position of the curve is given by the mean value μ , whereas the width of the curve is controlled by the standard deviation σ given by

$$\sigma = \sqrt{\frac{1}{N} \sum_{i=1}^{N} (x_i - \mu)^2}$$

This value can be very hard to visualise, as it is given by such an abstract formula. To make visualisation easier, the width of the curve can also be given by another unit called "Full width at half maximum" or in short FWHM. The definition of this unit is already in the name. The width of the curve at half of the maximum value. Since the maximum value is attained at the mean value where $x = \mu$ and therefore equal to $f(\mu, \sigma; x = \mu)$, are we looking for values for which

$$f(\mu, \sigma; x_1) = \frac{1}{2} f(\mu, \sigma; \mu)$$
 (2)

Half the width of the curve at half maximum will then be given by $W_{Half} = \sqrt{(\mu - x_1)^2}$. Due to the symmetry of the curve around $x = \mu$, the full width of the curve can be calculated by multiplying W_{Half} by 2. More specifically

$$FWHM = 2\sqrt{(\mu - x_1)^2}$$

Where x1 satisfies 2. This unit for the shape of the curve is much easier to visulatize for a given curve. Since both σ and FWHM are controlling the width of the curve, they can be put in relation to one another. The relationship between σ and FWHM is equal to

$$FWHM = 2\sqrt{2\ln 2}\sigma$$

The mathematical derivation of this can be found in

the appendix However, the gaussian probability distribution function alone is not sufficient to calculate the probability of a value as it only returns the probability density. To calculate the probability of the value x being in the interval [a, b], we need to integrate the probability distribution function over the given interval.

$$P(a \le x \le b) = \int_a^b f(\mu, \sigma; x) dx$$

A handy rule of thumb for a value being inside an interval expressed by σ is

$$P(-1\sigma \le x - \mu \le 1\sigma) \approx 0.68$$

$$P(-2\sigma \le x - \mu \le 2\sigma) \approx 0.95$$

$$P(-3\sigma \le x - \mu \le 3\sigma) \approx 0.997$$

To simulate our gas later used in the rocket engine, we will use the Maxwell-Boltzmann probability distribution. This distribution is based on the gaussian probability distribution, and uses a particles absolute temperature (T) and it's mass (m) as arguments as well as the Boltzmann constant $k = 1.380649 \cdot 10^{-23} J/K$.

When integrating the Maxwell-Boltzmann distribution over a given interval, it can be used to calculate the probability of a particle's speed being in the given interval.

$$P(v) = \int_{a}^{b} \left(\frac{m}{2\pi kT}\right)^{\frac{3}{2}} e^{-\frac{1}{2}\frac{mv^{2}}{kT}} 4\pi v^{2}$$
 (3)

$$P(v_x) = \int_a^b \sqrt{\frac{m}{2\pi kT}} e^{-\frac{1}{2}\frac{mv_x^2}{kT}}$$
 (4)

The Maxwell-Boltzmann distribution for the absolute speed in three dimensions 3 and for one dimension 4.

It is important to not confuse these two distributions as one is only for one-dimensional velocities, whereas the other is for the absolute velocity calculated from the velocity in all three spatial dimensions and will yield different results. Using $10^5\,H_2$ molecules with a temperature of $T=3000\,K$ each of the distributions can now be plotted to verify the formula and determine approximate values for both the absolute velocity and one-dimensional velocity.

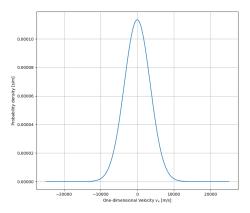


Figure 1. Maxwell-Boltzmann probability distribution for one-dimensional velocities of H_2 particles

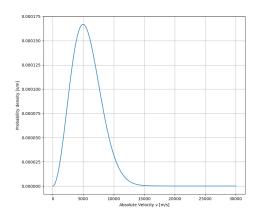


Figure 2. Maxwell-Boltzmann probability distribution for absolute velocity of \mathcal{H}_2 particles

IV. RESULTS

V. DISCUSSION

VI. CONCLUSION

VII. REFERENCES

1: https://www.merriam-webster.com/dictionary/gas

VIII. APPENDIX: MATHEMATICAL DERIVATIONS

Challenge A.3.1

$$\langle v \rangle = \int_0^\infty v P(v) \, \mathrm{d}v$$

$$\int_0^\infty x^{\frac{3}{2}} e^{-x} \mathrm{d}x = \frac{3}{4} \sqrt{\pi}$$

We try to rewrite the first integral to fit the form of the second

$$\int_0^\infty v P(v) \, dv = \int_0^\infty \left(\frac{m}{2\pi kT}\right)^{\frac{3}{2}} e^{-\frac{1}{2} \frac{mv^2}{kT}} 4\pi v^3$$

$$\frac{4}{\sqrt{\pi}} \int_0^\infty \left(\frac{mv^2}{2\pi kT}\right)^{\frac{3}{2}} e^{-\frac{1}{2}\frac{mv^2}{kT}}$$

We use substitution to get the aforementioned integral

$$\frac{4}{\sqrt{\pi}} \frac{3}{4} \sqrt{\pi} = 3$$