The Barnes Hut Algorithm: Implementation and Simulation in Julia

20602 - Computer Science (Algorithms)

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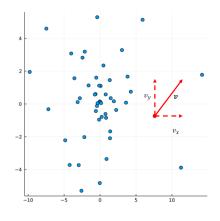
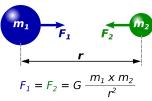
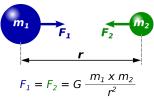


Figure: A cloud of 50 bodies

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Then, the point is moved according to Newton's second law of mechanics:

$$\vec{a}_i = \frac{\vec{F}_i^{net}}{m_i}$$

Therefore, to compute the net force acting on a body one needs to compute all the pairwise gravitational interactions between points.

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- Such time complexity is feasible only for small N
- ullet For example: Earth–Sun simulation, Solar System (N \sim 20)

```
unction onestepBrute(time::Float64,stars::Array{Star,1},spaceScale::Int64)
new stars = copy(stars)
F mat = zeros(length(stars),length(stars),2)
for i in 1:length(stars)
    for j in i:length(stars)
        if i != i
             F = newton(stars[i], stars[i])
            d j = stars[j]-stars[i]
            cos 0j, sin 0j = get cos sin(d j)
                                                                               O(N^2)
            f j = [F * cos 0j, F * sin 0j]
            F_{mat[i,j,:]} = f_{j} # * 10^{9}
            F mat[j,i,:] = -f j # * 10^9
    net force = [sum(F mat[i,:,1]),sum(F mat[i,:,2])]
    new stars[i] = moveStar(net force, stars[i], time, spaceScale)
end
return new stars
```

Figure: Naive brute force approach

Earth-Sun Simulation

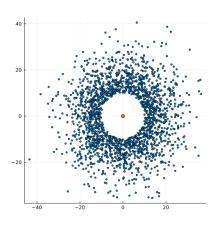
Why is there a spaceScale?

$$m_{sun} \sim$$
 1e30 $m_{earth} \sim$ 1e24 $d_{e,s}^2 \sim$ (1e11) 2 $G \sim$ 1e-11

$$\Rightarrow F_{e,s} \sim 1$$
e21

Divide the masses by 1e20 and the distance by 1e10, so that

$$F_{e.s}^{scaled} \sim 10^1$$



Earth Sun Simulation

MAIN IDEA:

Approximate the net force on a body in a clever way.

When a point is "far enough" from a cluster, the gravitational field generated by the cloud of points is *roughly* the same as if it were generated by a **single point** having as mass the sum of the masses and located in the **center of mass**¹.

¹The center of mass is the average of the coordinates weighted by the mass of the points, $(x_{CM}, y_{CM}) = (\sum m_i)^{-1} (\sum m_i x_i; \sum m_i y_i)$

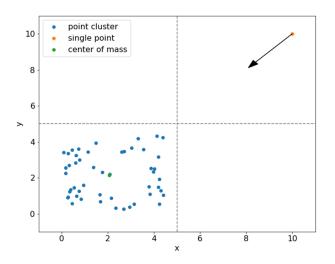


Figure: Approximating the net force with Barnes Hut

Problems:

- How do I define a cluster of points?
- What classifies as "far enough"?

Problem 1: how to define a cluster of points

- find a way to characterize the space of particles and organize their positions;
- optimize the necessary retrieval operations and computations.

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IDEA: divide the space using a tree structure called **Quadtree** (or Octree in 3D)

A **Quadtree** is a "map" of space that helps us model groups of points as a single center of mass. Formally, it is a tree structure with the following properties:

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- Each child node is recursively defined in the same way.

An example:²

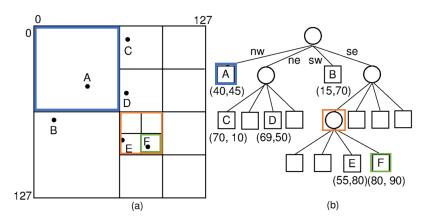


Figure: To the left the space carved by the tree, and to the right the tree structure of the points in the space

²source: website

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• We start at the root of the tree, which represents the entire space, but it is empty at the beginning.

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- 3 At the end, return the root.

Quadtree Data Structure

Once we constructed the tree, we will only need to traverse it from the root to the leaves, in order to compute the force acting on a point.

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Hence, we can store the quadtree through Node2D objects which contain four pointers to the children nodes and also information about the quadrant.

Quadtree Data Structure

Therefore, it is enough to keep track of the root of the tree, and from there we can explore all the other nodes.

This data structure has the following operation times:

	Insert	Get
Average	O(logN)	O(logN)
Worst	O(N)	O(N)

Table: Time complexity of operations

As for space complexity, the number of nodes in the tree is

$$\frac{4N-1}{3}=O(N)$$

Build Quadtree in Julia

```
function buildOTree(root::Union{Nothing,Node2D},stars::Union{Nothing,Array{Star,1}},
                x lim::Array(Float64.1), v lim::Array(Float64.1))
if isnothing(root)
    root = createRoot(stars,x lim,y lim)
for s in stars
    0 = getQuadrant(root.regCenter,s.s)
    if isnothing(root.children[Q])
        center, x lim, y lim = getCenterLimits(root,Q)
        root.children[Q] = Node2D(root,center,x lim,y lim)
        root.children[0] = updateNode!(root.children[0].s)
        root.children[0] = updateNode!(root.children[0].s) 	
        if root.children[Q].n == 2
            stars = root.children[0].stars
             stars = [s]
        root.children[Q] = buildQTree(root.children[Q], stars ,
                                      root.children[Q].x lim, root.children[Q].y lim)
return root
```

O(NlogN)

Problem 2: What classifies as "far enough"?

Once we have a way to divide the space into regions and the points into clusters, we need a rule to define what can be considered "far enough".

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The Multipole Acceptance Criterion (MAC)

If the size of the cluster s divided by the distance of the cluster's CoM d from the point is lower than a parameter $\theta>0$, then consider the cluster as a whole, i.e. approximate if:

$$\theta > \frac{s}{d}$$

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• If $\theta = 0$, then BH will never approximate, and it will produce a result equivalent to brute force.

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A usual good starting point is $\theta=1$, which can be then adjusted according to the interest of the simulation.

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- **1** (MAC) If $\theta > s/d$ then compute the force between the point i and the CoM (having as mass the sum of all points in the quadrant);
- If not, repeat the same procedure for each non-empty children of the node considered.

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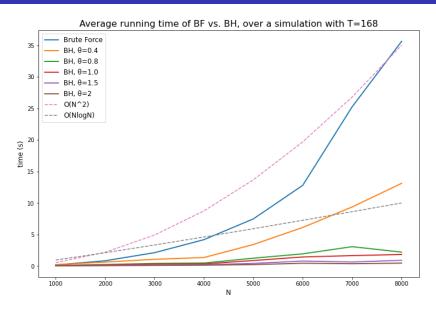
For each point i (starting at the root node) do:

- Compute the node's size s and the distance d(i, CoM) of i from the CoM;
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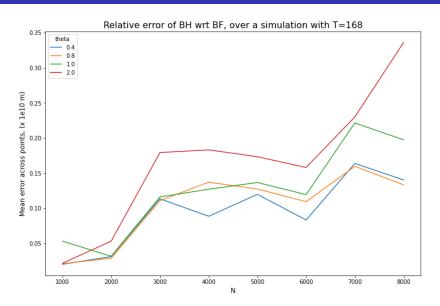
At the end of each iteration, we'll have \vec{F}_i^{net} and we are able to move the point accordingly.

```
function computeForceTree(node::Node2D,star::Star,theta::Float64)
F_s = [0.,0.]
 s = abs(node.x lim[2]-node.x lim[1])
d = distance(star.s.node.centerOfMass)
 if node.n == 1
     if star ∉ node.stars
        f = newton(star.node.stars[1])
        dir = node.stars[1]-star
        cos0, sin0 = get cos sin(dir)
        F s += [f * cos0, f * sin0] # * 10^9
        F s += [0.,0.]
 elseif s/d < theta
                                                                                                   O(log N)
    cm = Star(node.centerOfMass,[0.,0.],node.totalMass)
    f = newton(star.cm)
    dir = cm-star
    cos0, sin0 = get cos sin(dir)
    F_s += [f * cos0, f * sin0] # * 10^9
                                                                O(logN)
    for k in keys(node.children)
         if ! isnothing(node.children[k])
             F s += computeForceTree(node.children[k],star,theta)
return F s
```

Simulations and Benchmarking



Simulations and Benchmarking



Galaxy clash simulation

Galaxy clash simulation

Code available at github.com/PietroDomi/BarnesHut.jl

References

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