

# All-Optical Radiation Reaction at $10^{21} \text{ W/cm}^2$

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Notebook: Óscar Amaro, November 2022 @ [GoLP-EPP](#)

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## Introduction

In this notebook we reproduce figure 3.

## Figure 3: Electron beam energy loss

(In the units chosen in the paper, the classical electron radius, present in the Thompson cross section, is  $r_e = \frac{e^2}{m c^2}$  while in SI  $r_e = \frac{1}{4 \pi \epsilon_0} \frac{e^2}{m c^2}$ )

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In[ ]:= (* LP PW laser colliding against a relativistic electron *)
Clear[γ, γf, γ0, k, k3, η, EGeV, τ0, τ0fs, e, m, c, λ, a0, ω0, ΔEE, χ, ΔEEχ, ε0]

(* equation 3 *)
γf = γ0 / (1 + k γ0); (*[] final electron energy *)

$$k3 = \frac{1}{4 \pi \epsilon_0} (1 - \cos[\pi])^2 \frac{\eta}{3} \frac{e^2 \omega_0^2}{m c^3} a_0^2 \tau_0; (*[] \text{CRR } k \text{ factor from equation 3} *)$$


$$k = (1 - \cos[\pi])^2 3.2 \times 10^{-5} (10^{-22}) \tau_0 \text{fs}; (*[] \text{CRR } k \text{ factor Engineering formula} *)$$

ΔEE = (γ0 - γf) / γ0; (*[] 0 ≤ energy loss ≤ 1 *)

(* χ=1: χ~2 γ0 a0/aS *)
(*Solve[γf==γ0/(1+k γ0), γ0][[1,1,2]]*)
γ = Solve[1 == 2 γ a0 / (411823), γ][[1, 1, 2]];
ΔEEχ = ΔEE /. {γ0 → γ};

(* physical constants *)
e = 1.6 × 10-19; (*[C]*)
m = 9.1 × 10-31; (*[Kg]*)
c = 3 × 108; (*[m/s]*)
ε0 = 8.854 × 10-12; (*[F/m] vacuum permittivity*)

(* parameters *)
η = 0.4; (*[] temporal profiles *)
τ0 = 26.5 × 10-15; (*[s]*)
τ0fs = τ0 1015; (*[fs]*)
λ = 1; (*[μm] laser central wavelength *)
ω0 = 2 π c / (λ 10-6) // N; (*[1/s] laser frequency *)
a0 = 0.855 Sqrt[10-18] λ; (*[] laser a0 *)
γ0 = EGeV / (0.511 × 10-3); (*[] electron γ from GeV to boost *)

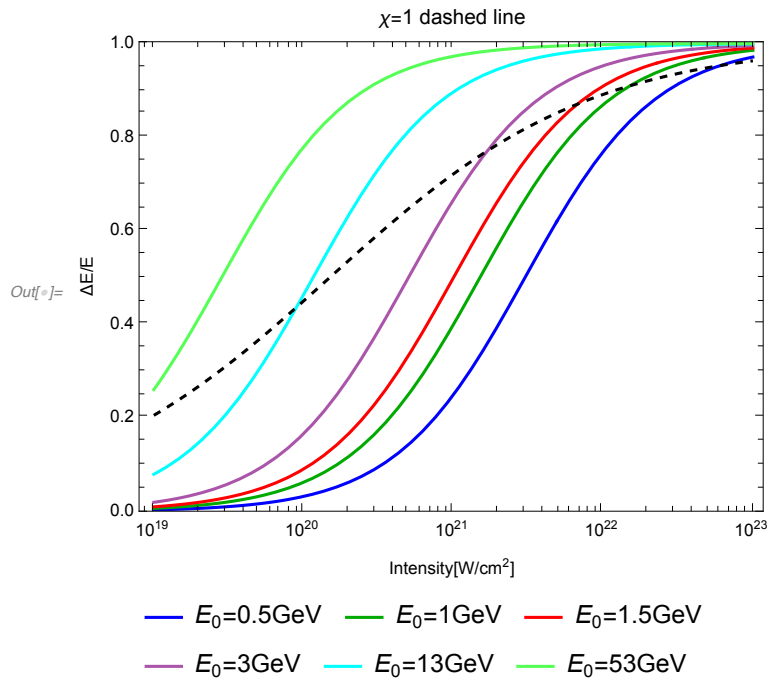
(* compare the CRR k factors of equation 3 and the "engineering formula" *)
k3
k
(*k ~ 4.64 10-7 a02 for these parameters *)

(* plot *)
LogLinearPlot[{ΔEE /. {EGeV → 0.5}, ΔEE /. {EGeV → 1}, ΔEE /. {EGeV → 1.5},
  ΔEE /. {EGeV → 3}, ΔEE /. {EGeV → 13}, ΔEE /. {EGeV → 53}, ΔEEχ},
{1, 1019, 1023}, Frame → True, FrameLabel → {"Intensity[W/cm2]", "ΔE/E"},
PlotRange → {0, 1}, PlotStyle → {Blue, Darker[Green], Red,
  Lighter[Purple], Cyan, Lighter[Green], {Dashed, Black}}, PlotLegends →
{"E0=0.5GeV", "E0=1GeV", "E0=1.5GeV", "E0=3GeV", "E0=13GeV", "E0=53GeV"},
AspectRatio → 3 / 4, PlotLabel → "χ=1 dashed line"]

```

Out[ ]:= 3.43767 × 10<sup>-25</sup> I

Out[ ]:= 3.392 × 10<sup>-25</sup> I



# $\eta$ factor for the two envelopes

```

In[242]:= Clear[t,  $\tau$ ,  $\tau_0$ , aPoly, aSin, intPoly, intSin]
 $\tau_0 = 1$ ; (*[] normalize all times to  $\tau_0$  *)
 $\tau = t / \tau_0$ ;

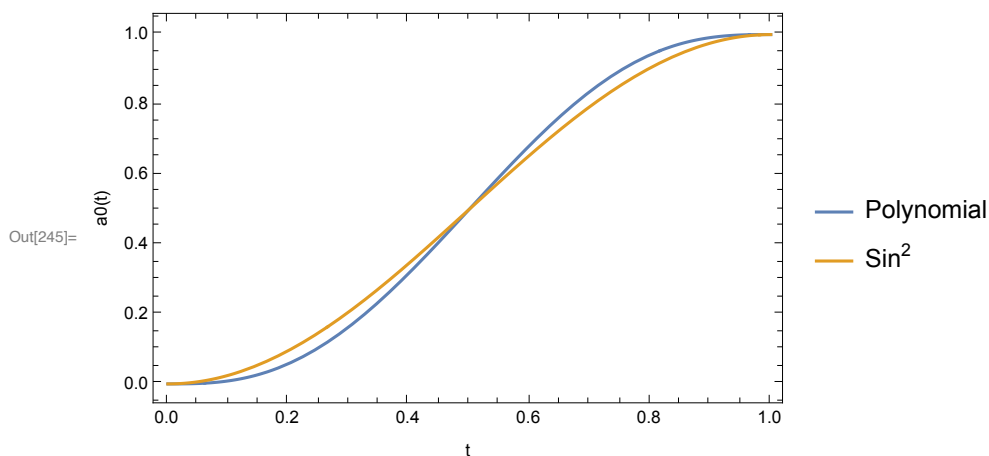
aPoly =  $10 \tau^3 - 15 \tau^4 + 6 \tau^5$ ; (* polynomial envelope. instead of using  $\tau =$ 
 $\sqrt{2} t / \tau_0$  as in the paper (which would give incorrect integrals) we use
simply  $\tau = t / \tau_0$  to be able to compare between the two envelopes *)
aSin =  $\text{Sin}\left[\frac{\pi t}{2 \tau_0}\right]^2$ ; (*  $\text{sin}^2$  envelope *)

(* plot the two envelopes *)
Plot[{aPoly, aSin}, {t, 0,  $\tau_0$ }, Frame  $\rightarrow$  True,
FrameLabel  $\rightarrow$  {"t", " $a_0(t)$ "}, PlotLegends  $\rightarrow$  {"Polynomial", "Sin2"}]

(* the k factor will be proportional to  $a(t)^2$ . since
the envelopes are symmetric, we only integrate up to  $t = \tau_0$  *)
intPoly = Integrate[aPoly2, {t, 0,  $\tau_0$ }]
intSin = Integrate[aSin2, {t, 0,  $\tau_0$ }]

(* the ratio of the integrals from this notebook *)
intSin / intPoly // N
(* the ratio of the integrals from the paper *)
0.375 / 0.392

```



Out[248]= 0.957182

Out[249]= 0.956633