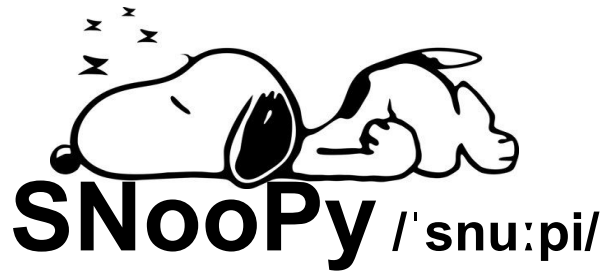


SN-Ia LC-fitting

SN parameters correlations

SN and host galaxy correlations

Nataliya Ramos Chernenko



def., **SuperNovae in Object Oriented Python**

(python package) a collection of tools useful to build your own fitter
in order to analyze the Type Ia supernovae.

1. EBV_model. Only work with Δm_{15} - based template.

2. EBV_model2. Δm_{15} & s_{BV} - based templates

3. Max_model. Δm_{15} & s_{BV} - based templates

4. Color_model. Only work with 'color-stretch' s_{BV} parameter

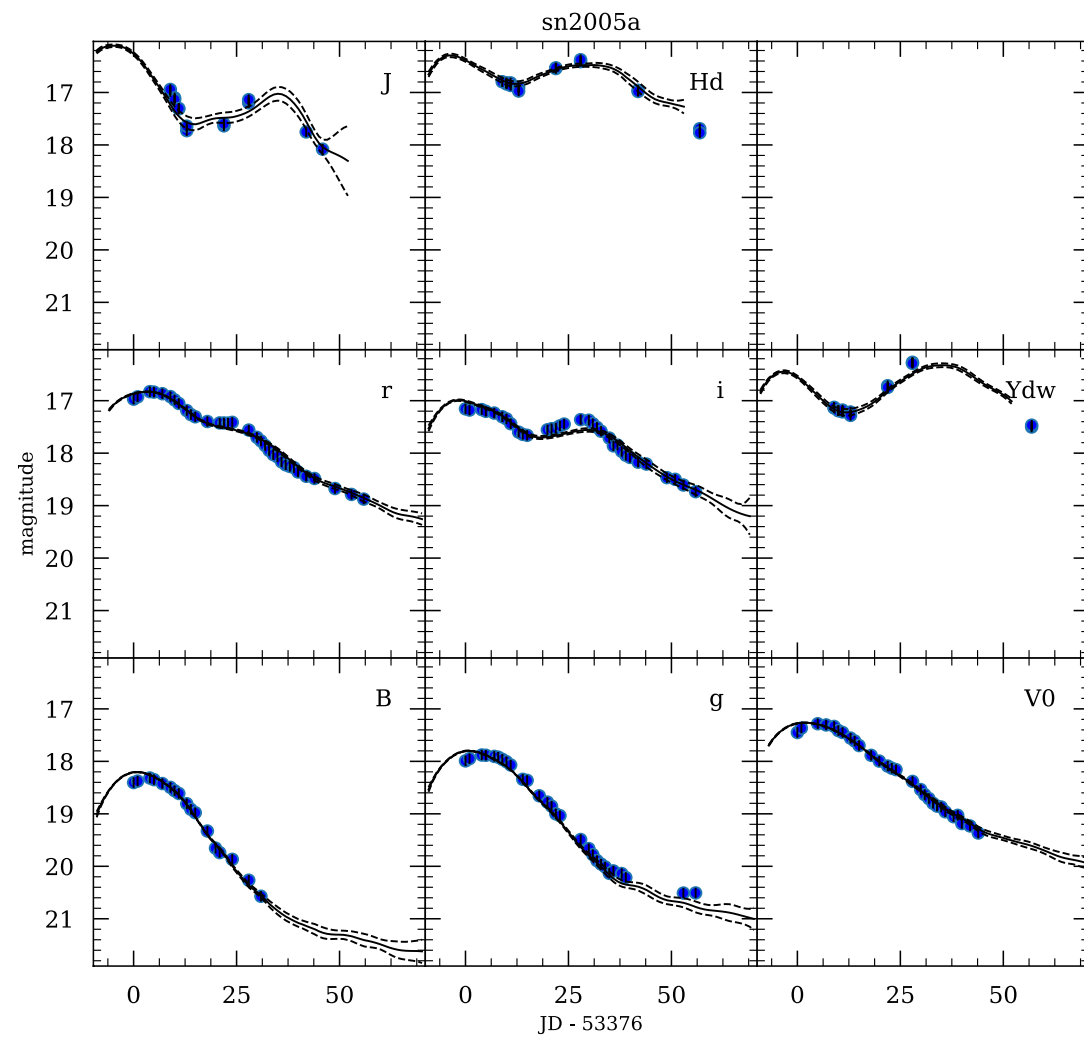
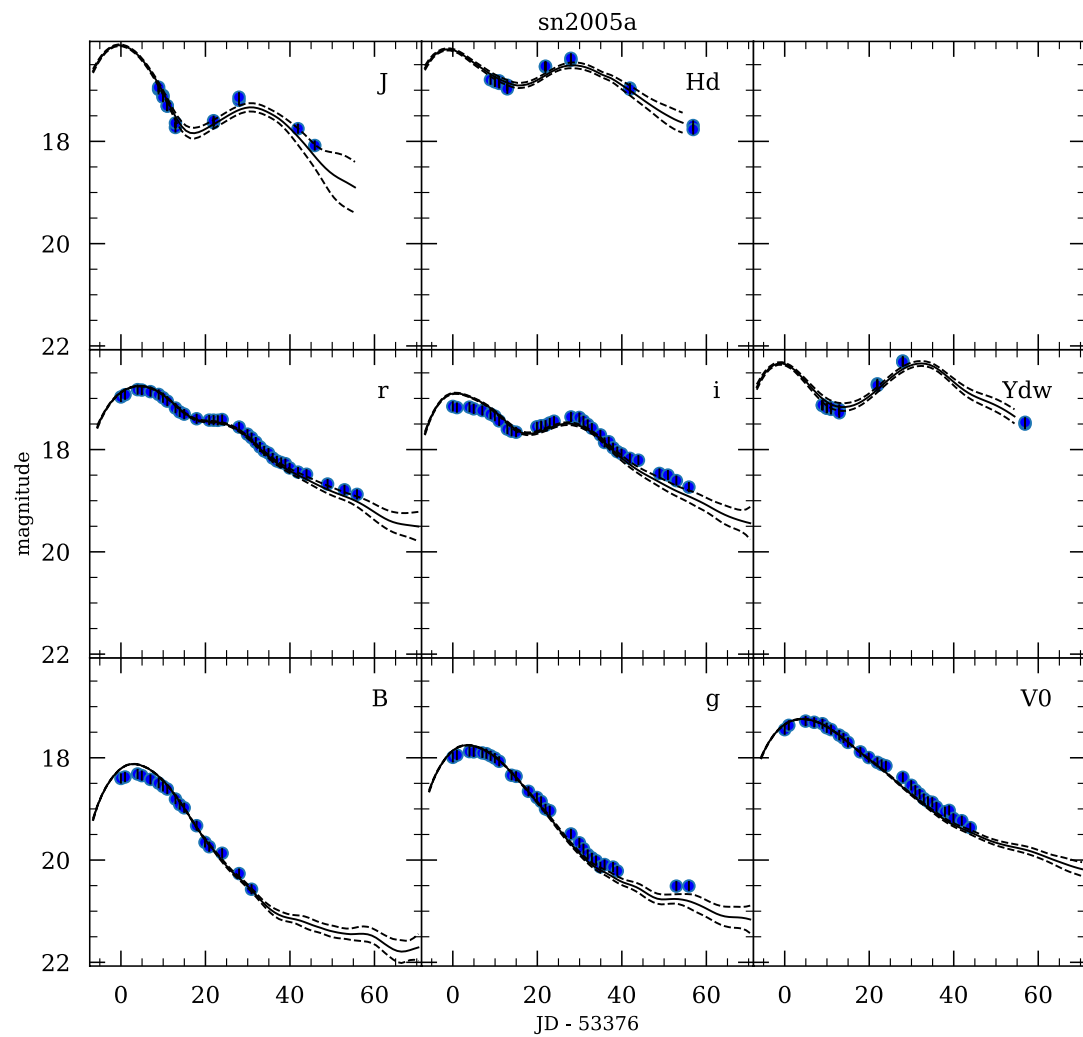
Two fitters: **non-linear least-square Levenberg-Marquart** algorithm
and **Markov Chain Monte Carlo** fitter.

1. **EBV_model2.** Δm_{15} & s_{BV} - based templates

$$\begin{aligned} m_X(t) &= T_Y(t', stype) + M_Y(stype) + \mu \\ &+ R_X E(B - V)_{gal} + R_Y E(B - V)_{host} + K_{XY} \end{aligned}$$

- $m_X(t)$: observed magnitude in t , time relative to B_{\max} .
- $E(B - V)_{gal}$ and $E(B - V)_{host}$: reddening due to galactic foreground and host galaxy, respectively.
- R_X and R_Y : total-to-selective absorptions for filters X and Y, respectively.
- t' is the de-redshifted time relative to B_{\max} .

This model fits **4 parameters**: T_{\max} , Δm_{15} or s_{BV} , EBV_{host} , and DM .

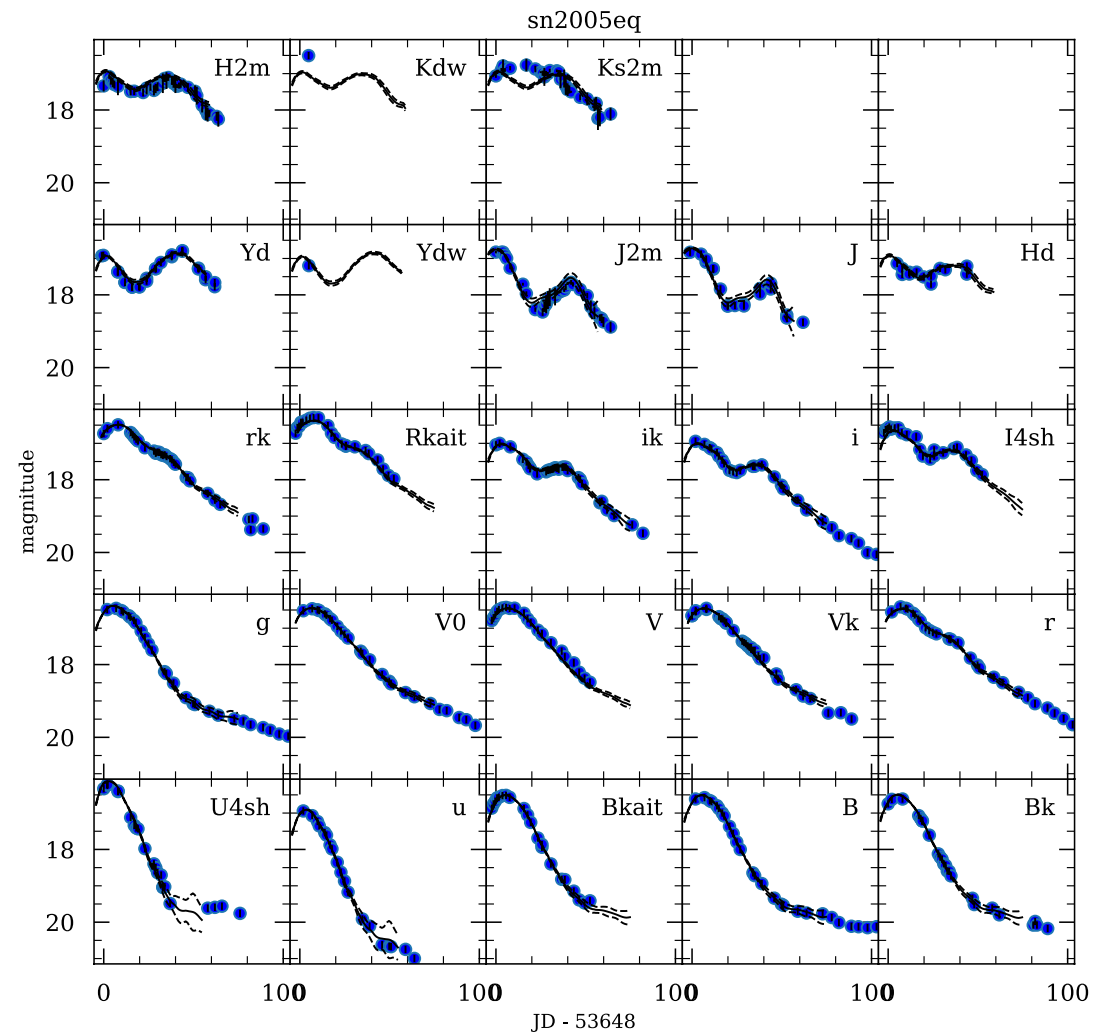
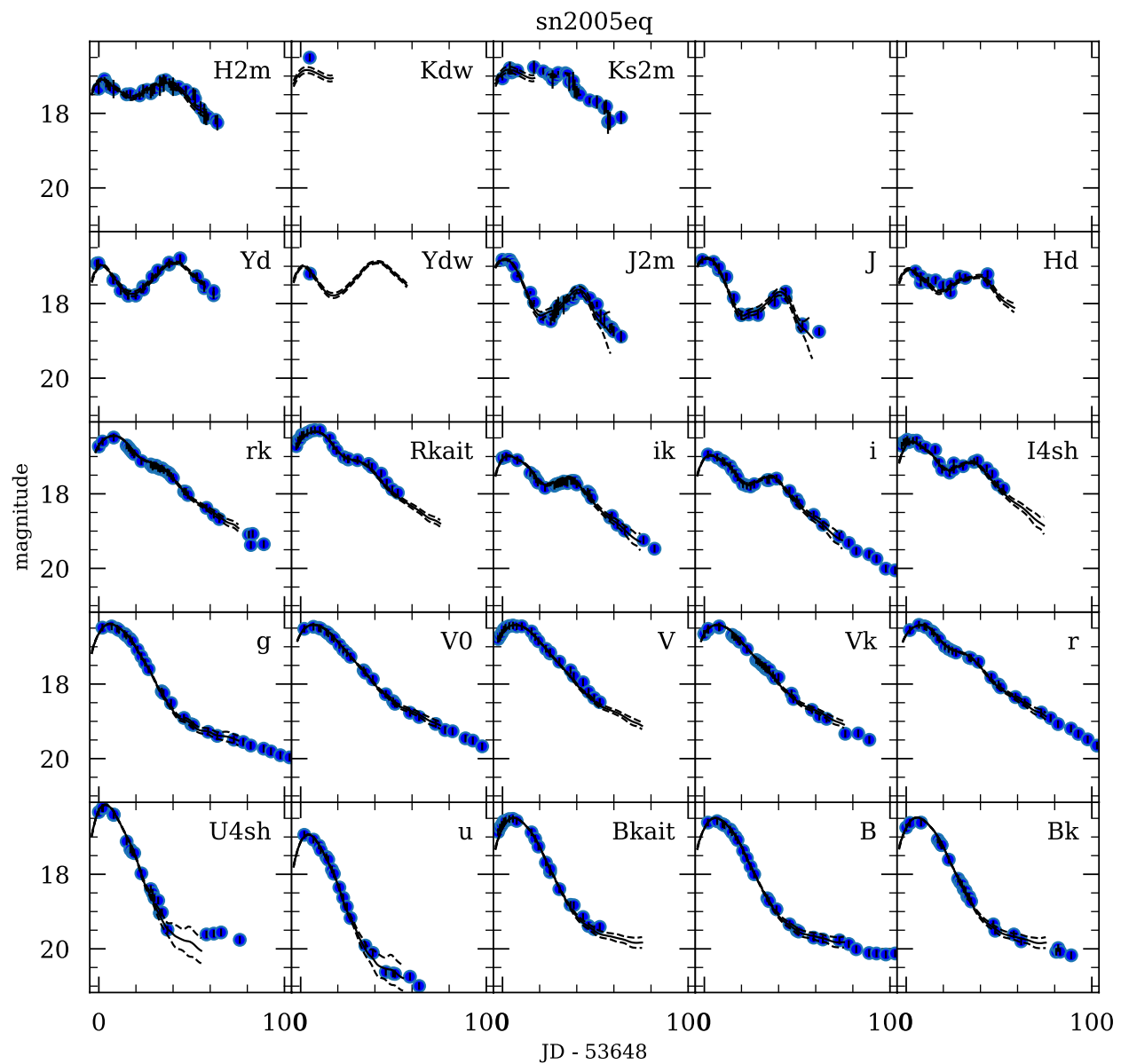


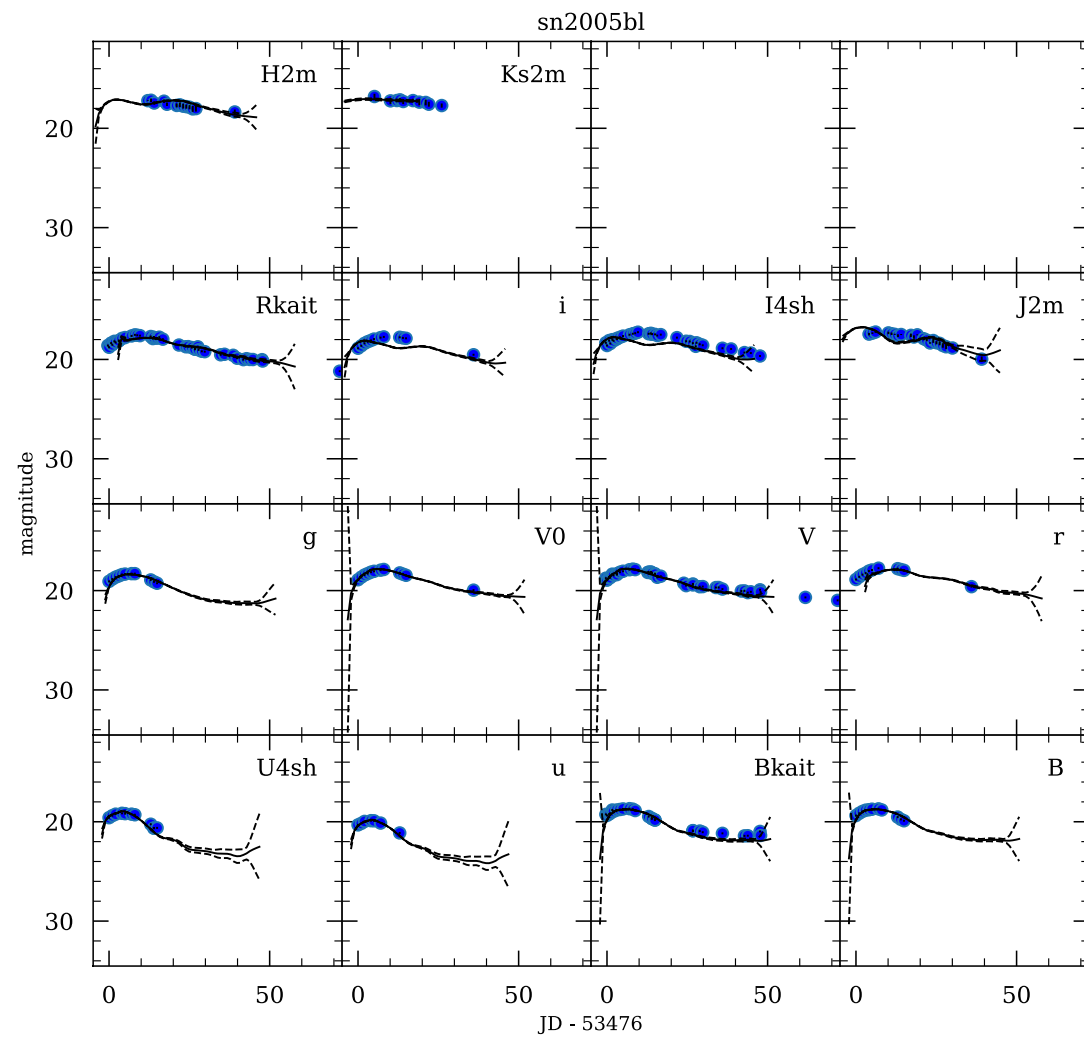
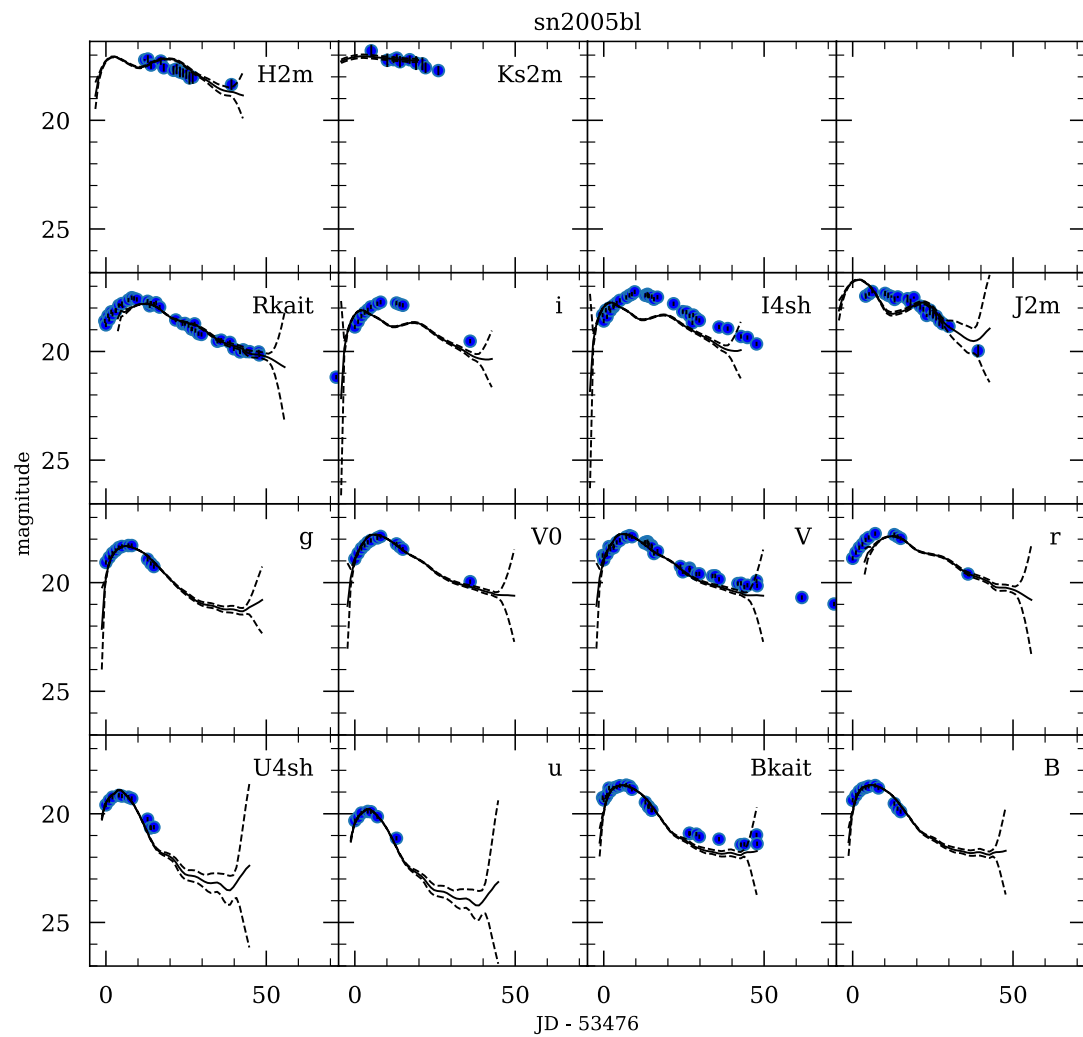
2. Max_model. Δm_{15} & s_{BV} - based templates.

$$m_X(t) \\ = T_Y(t', stype) + m_Y + R_X E(B - V)_{gal} + K_{XY}$$

- m_Y is the peak magnitude in filter Y
- K_{XY} the cross-band k-correction from rest-frame X to observed filter Y.
- $stype = \Delta m_{15}$ or s_{BV} .

For **N filters**, this model fit **N+2 parameters**: T_{\max} , Δm_{15} or s_{BV} and $N f_{\max}$.



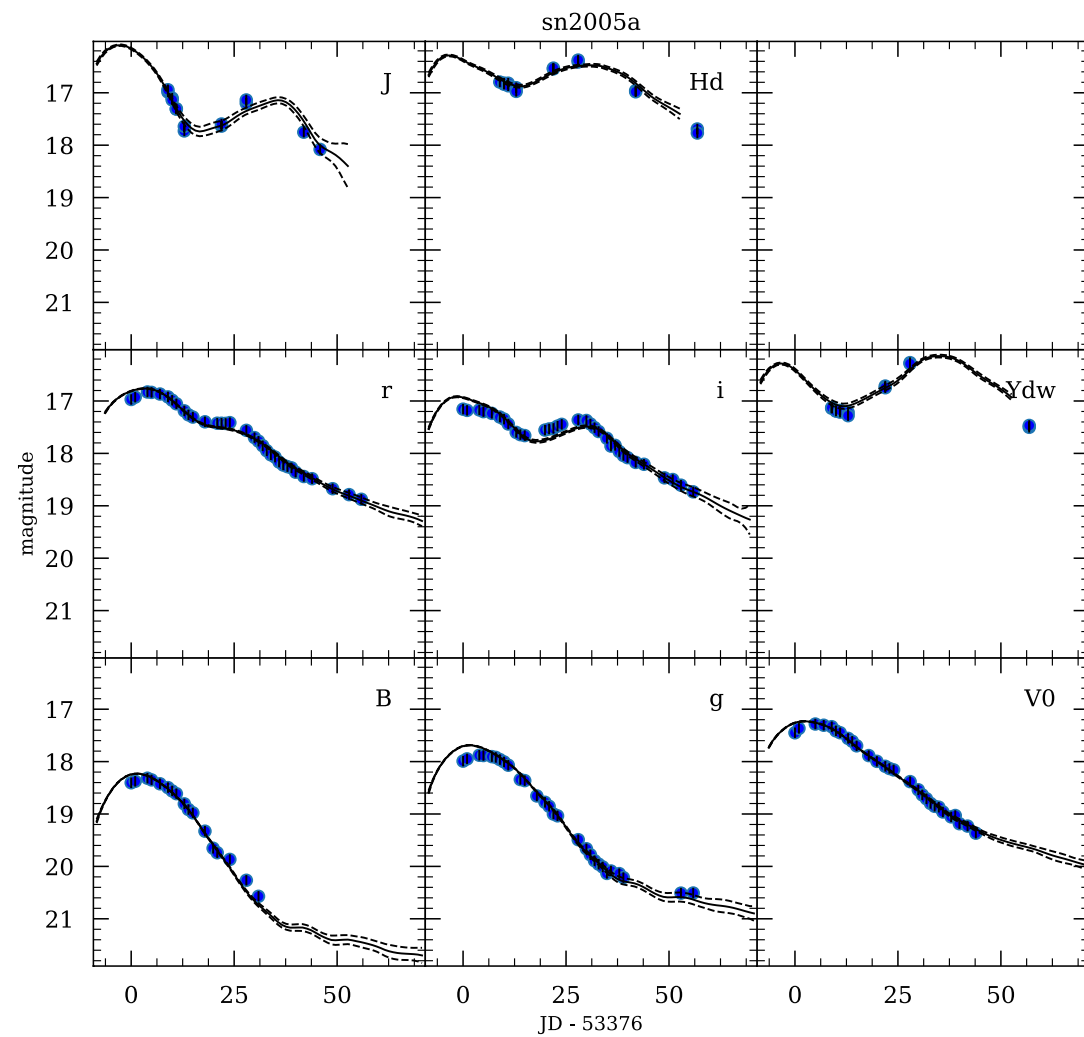
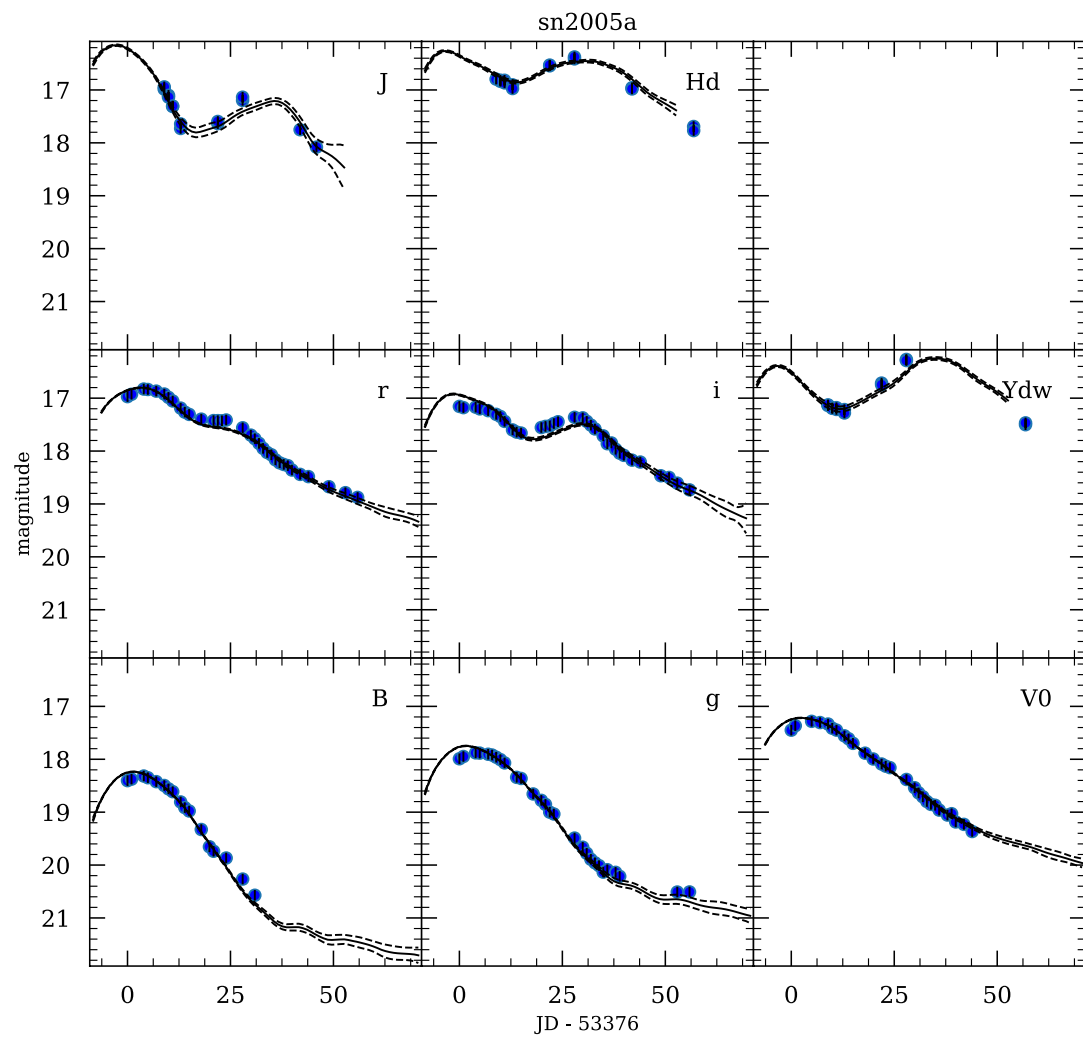


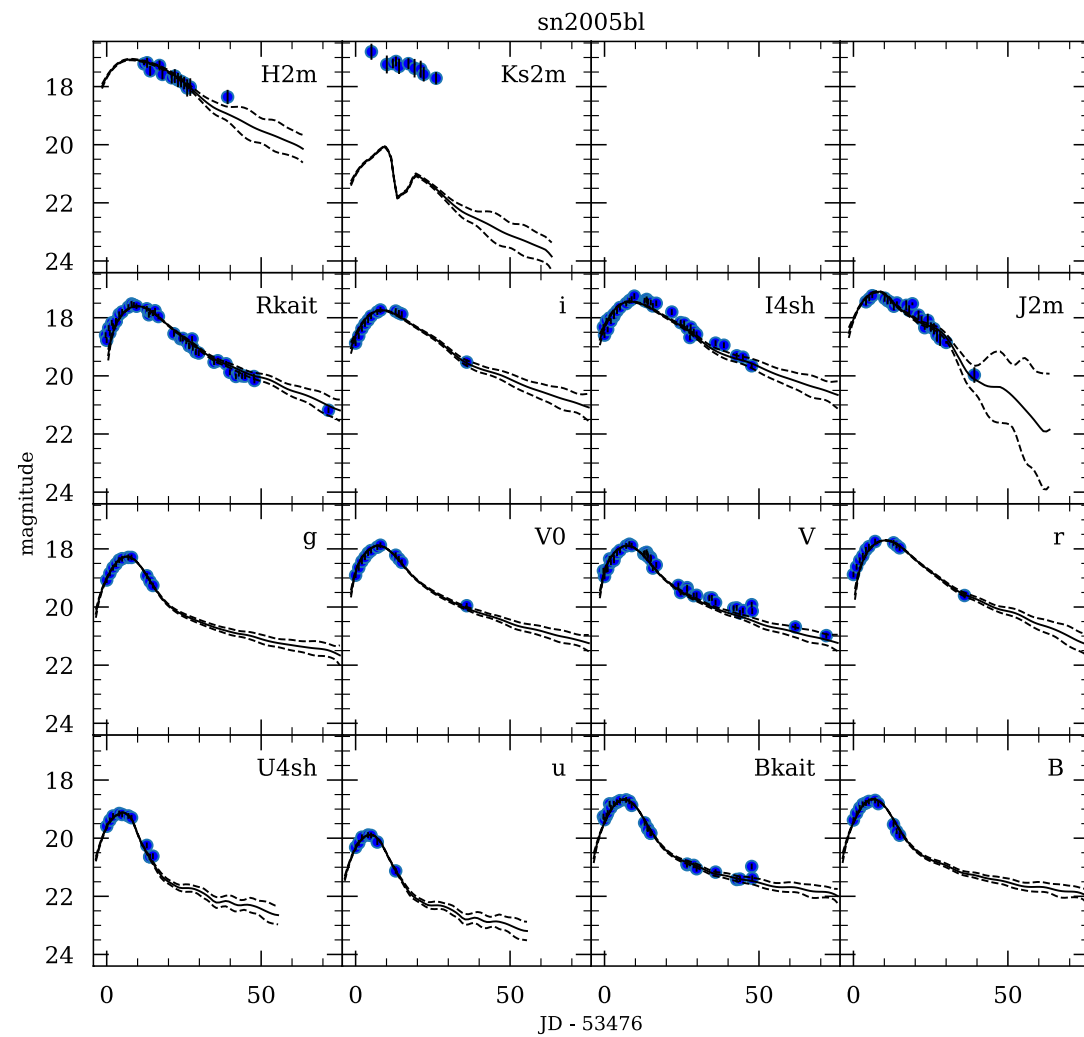
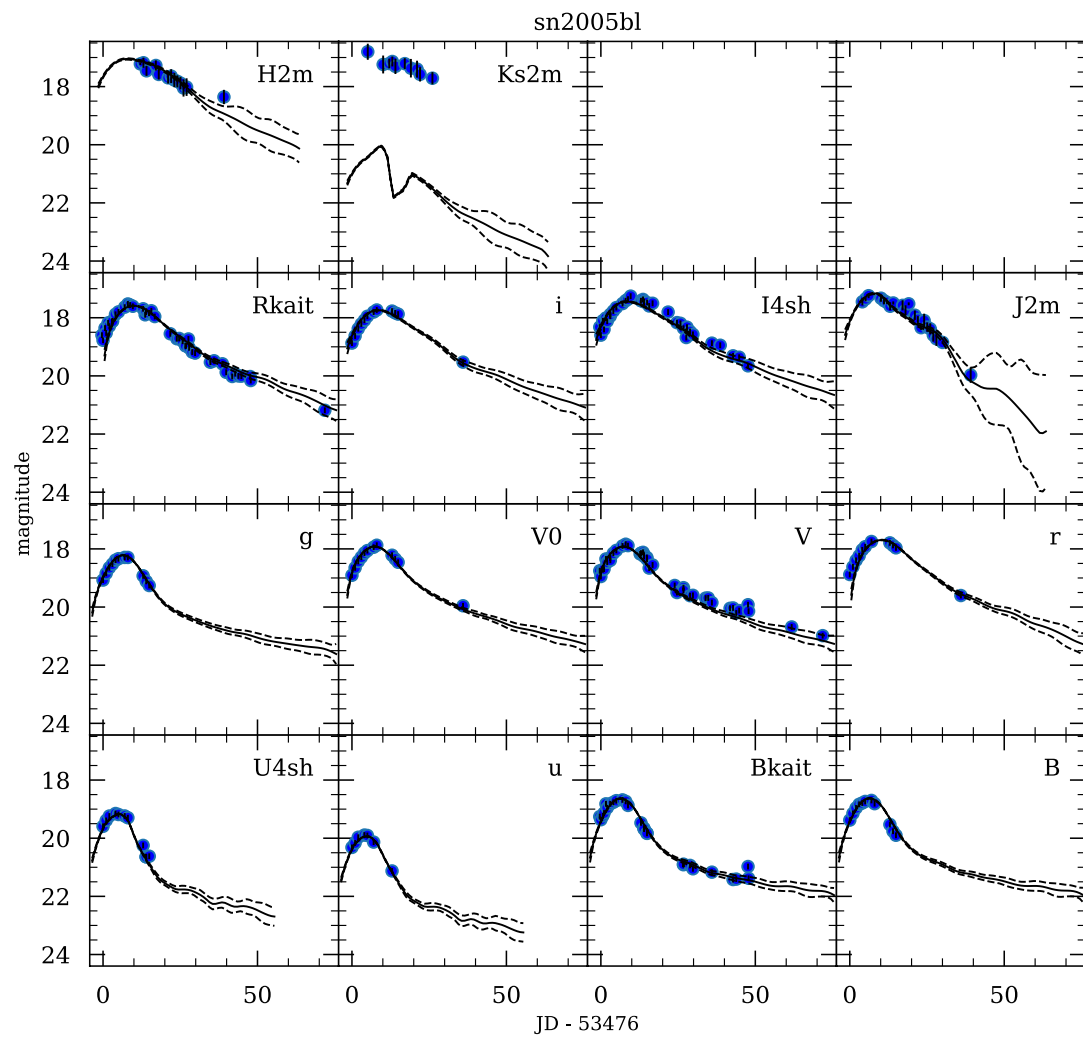
3. Color_model. *Only work with ‘color-stretch’ s_{BV} parameter*

$$\begin{aligned} m_X(t) \\ = T_Y(t', s_{BV}) + B_{max} + (X - B)(s_{BV}) + R_X \cdot E(B - V)_{gal} \\ + R_Y(R_V) \cdot E(B - V)_{host} + K_{XY} \end{aligned}$$

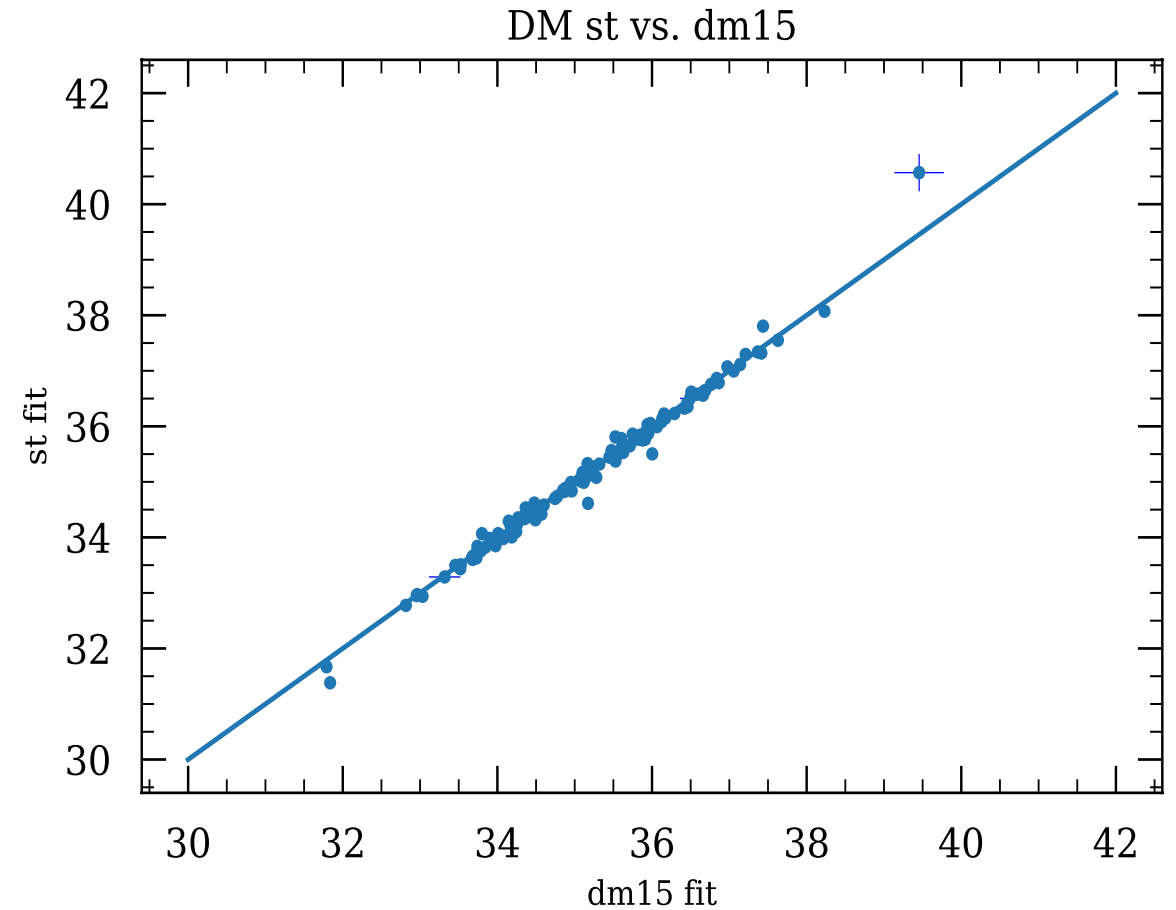
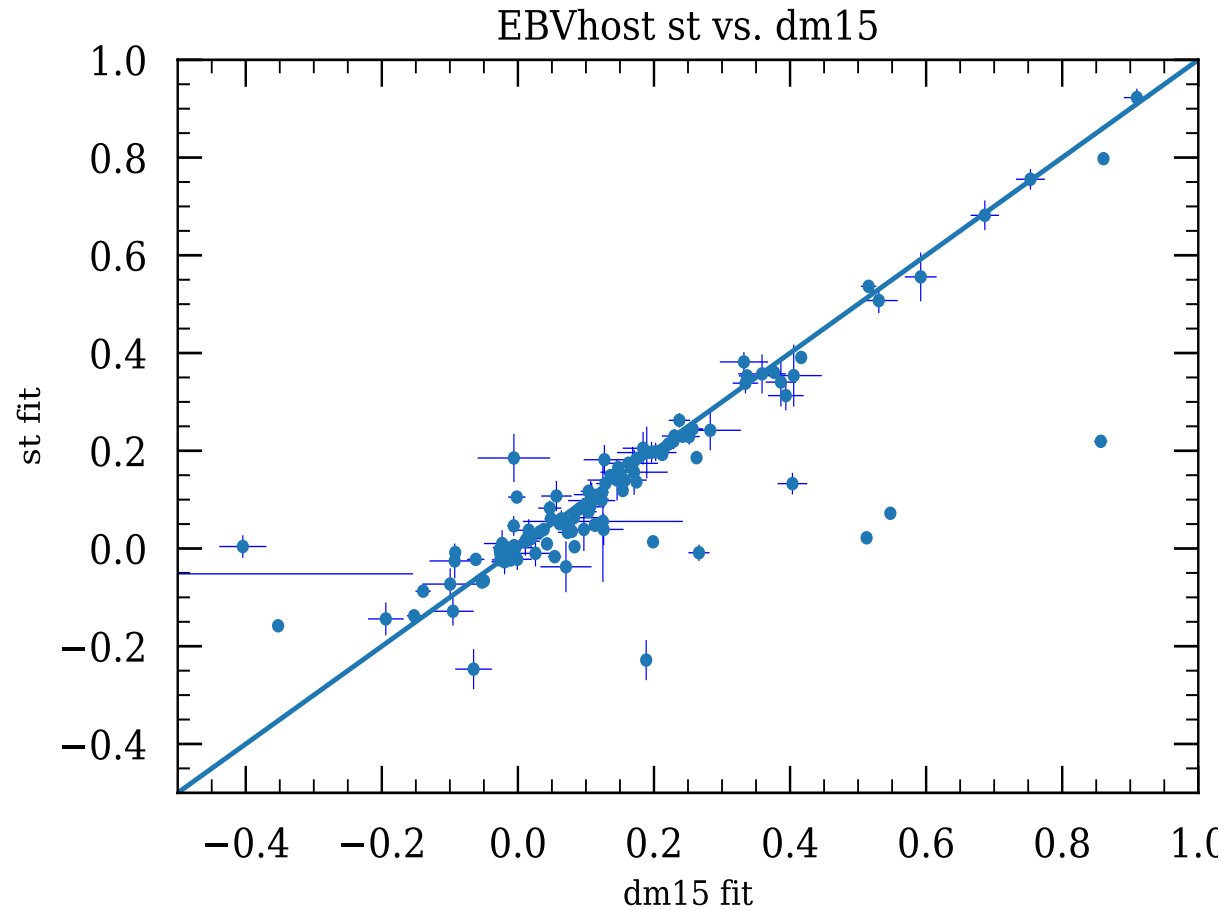
- B_{max} is the de-reddened and K-corrected B maximum
- $(X - B)(s_{BV})$ is the intrinsic X-B color, which is a function of s_{BV} .

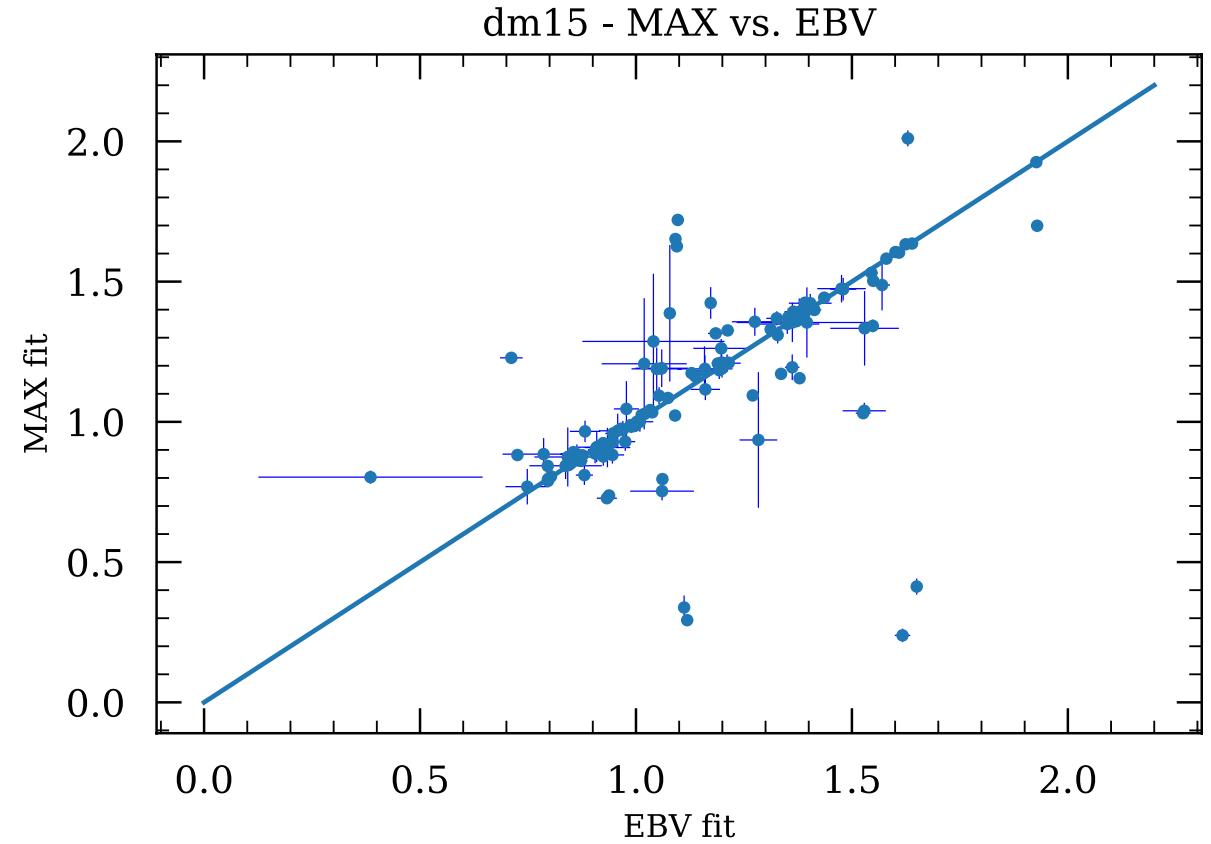
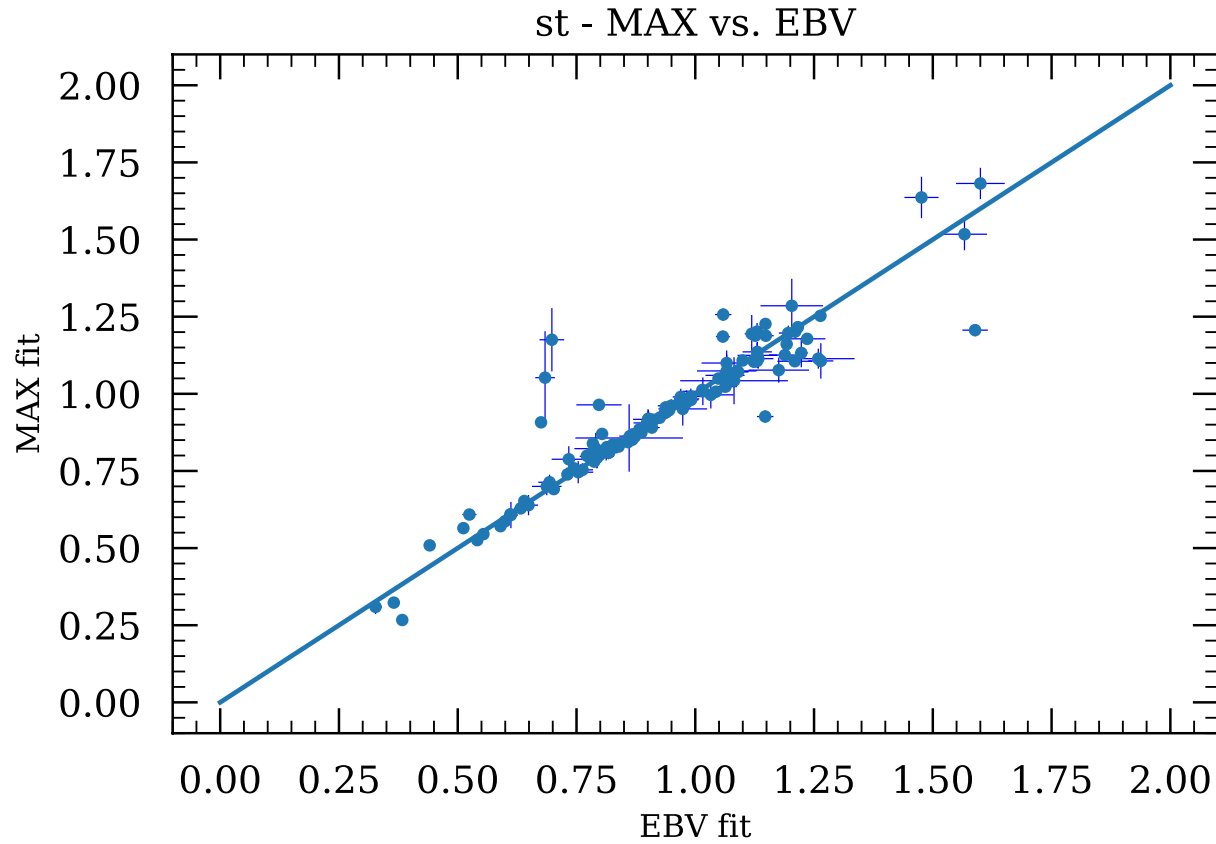
The model has **5 free parameters**: T_{max} , Δm_{15} or s_{BV} , EBV_{host} and R_V .





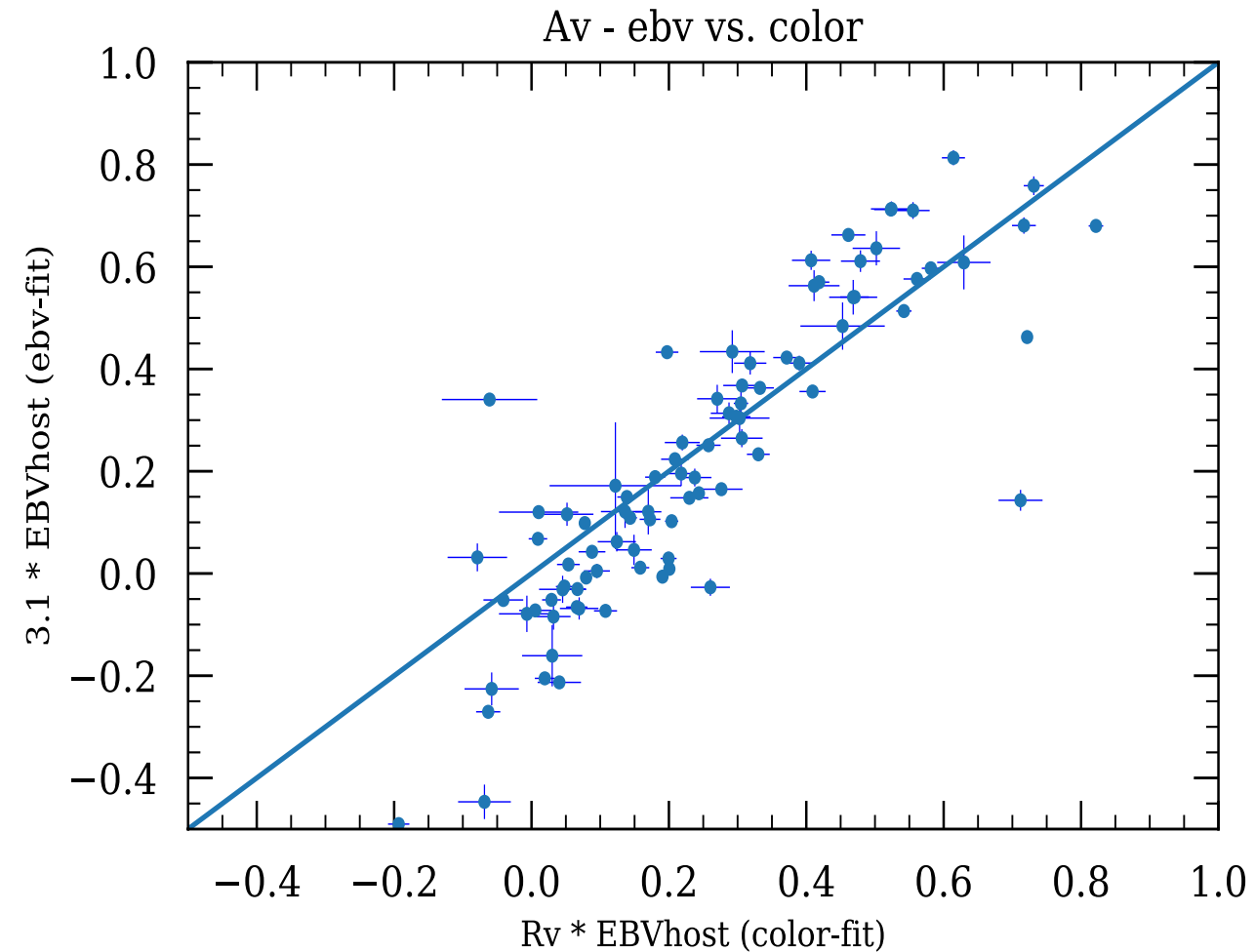
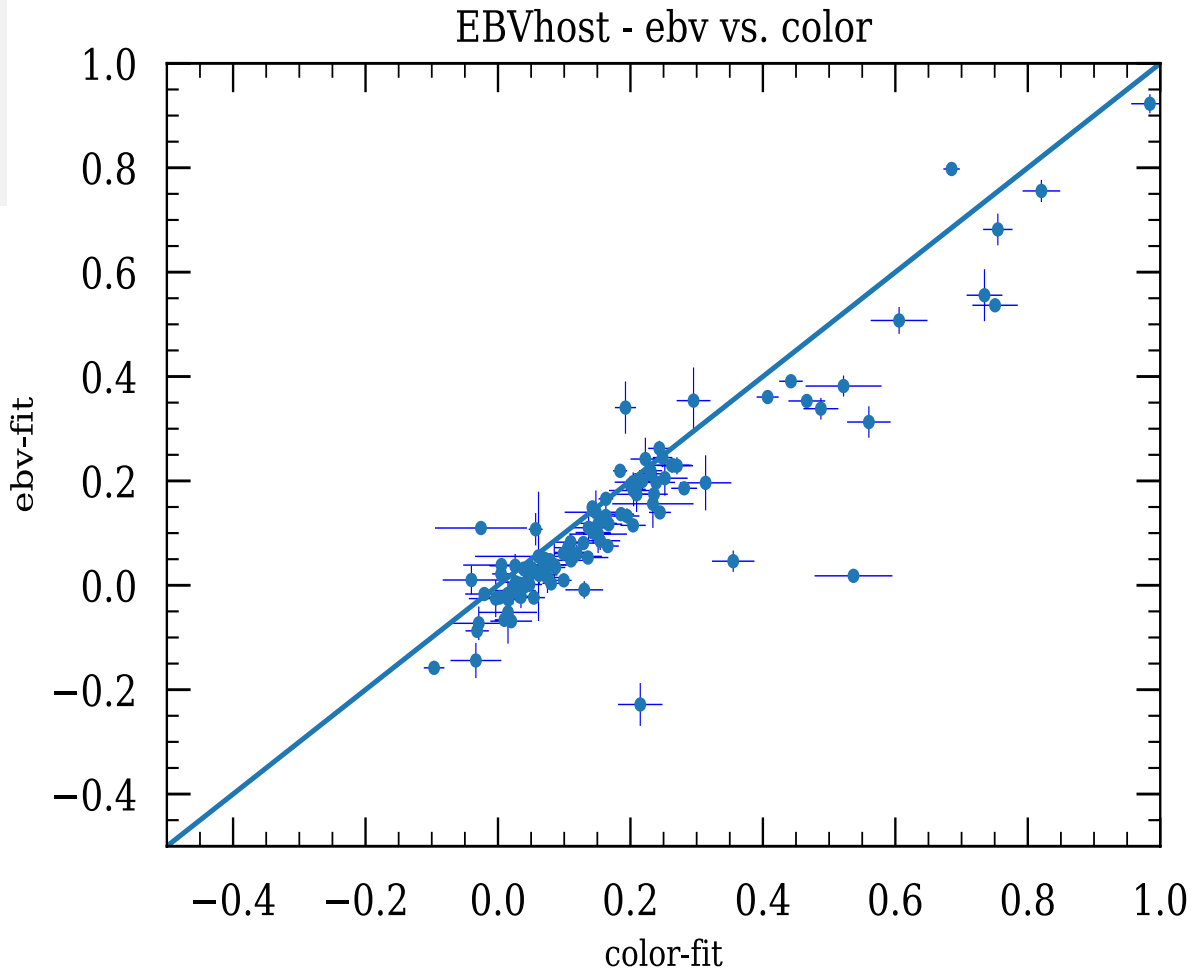
Are the estimated values independent of the selected LC-shape-parameter template within the same model?



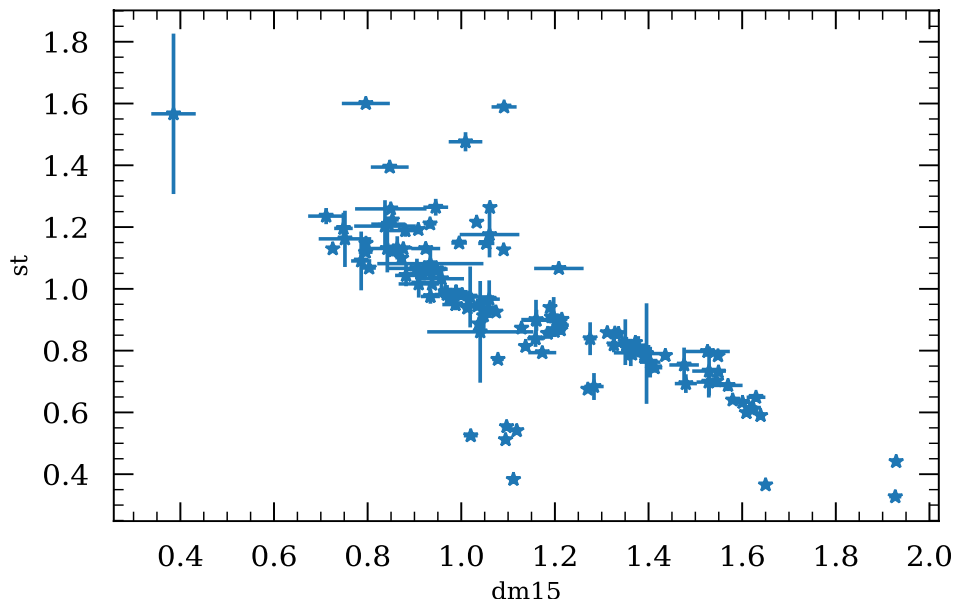


Is the estimated LC-shape-parameter independent of the selected model?

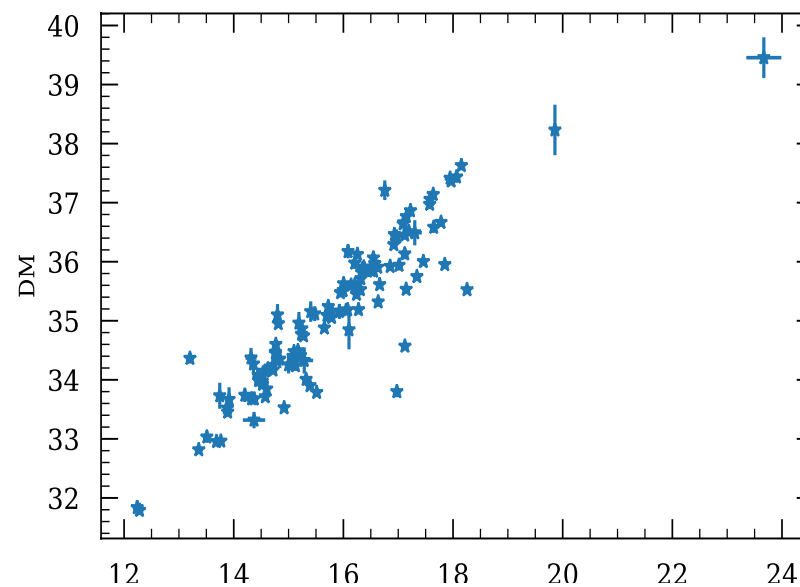
And if some, what kind of dependance ?



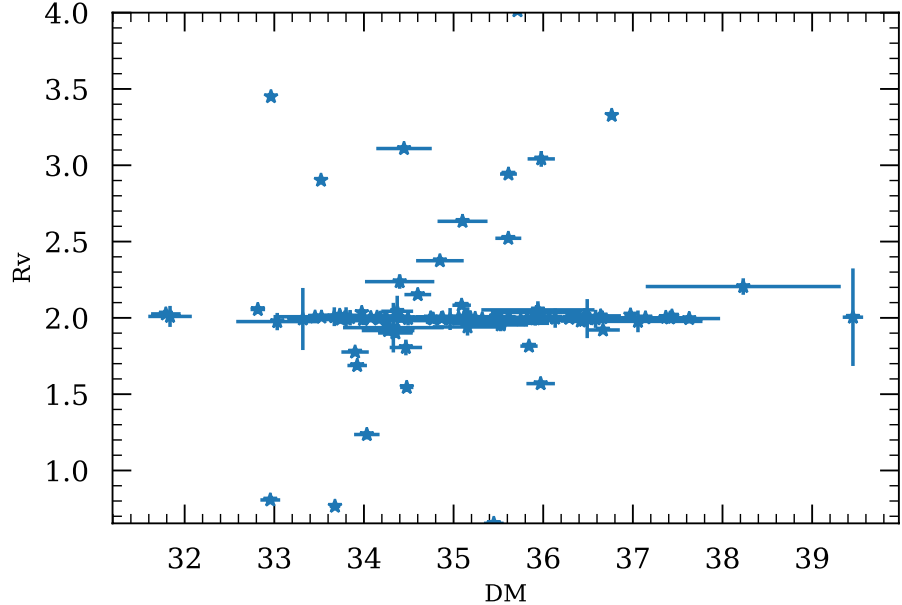
st vs. dm15



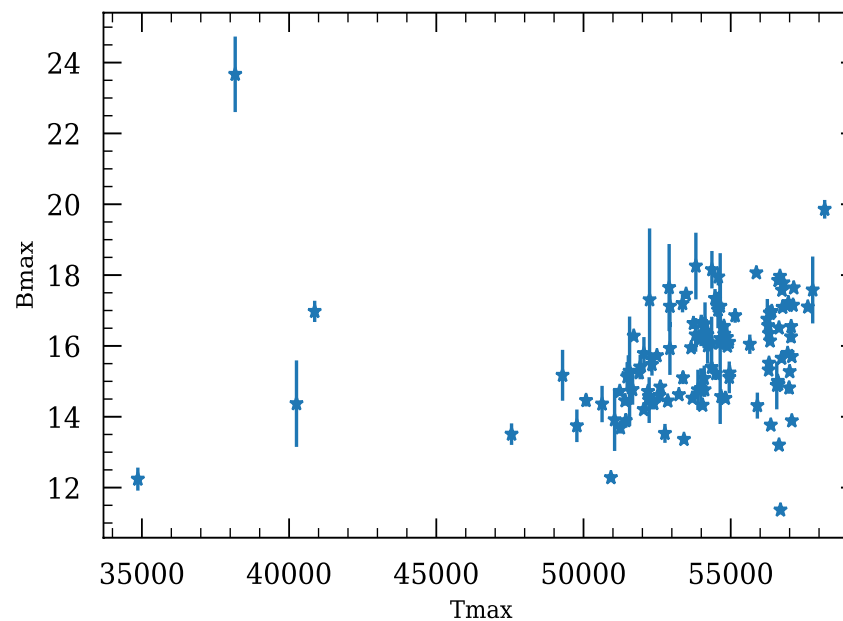
DM vs. Bmax



Rv vs. DM



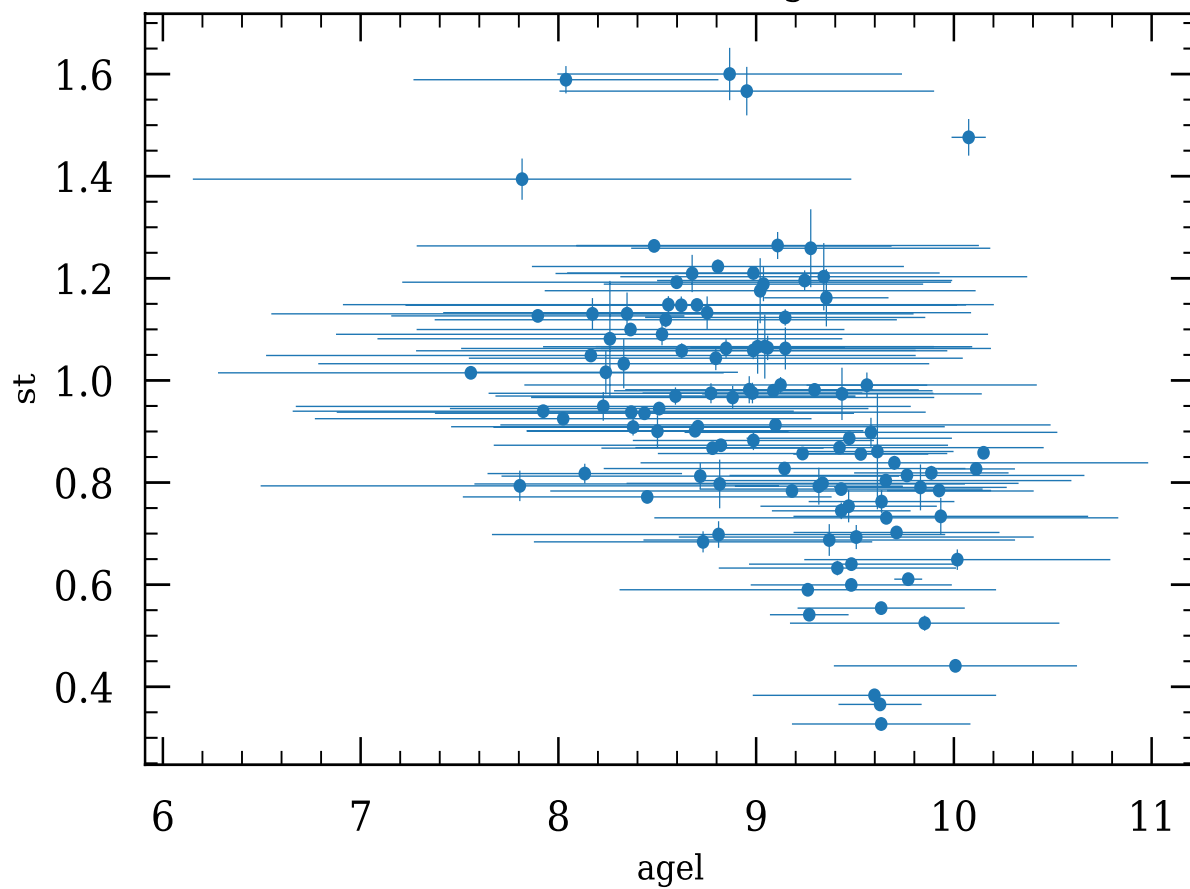
Bmax vs. Tmax



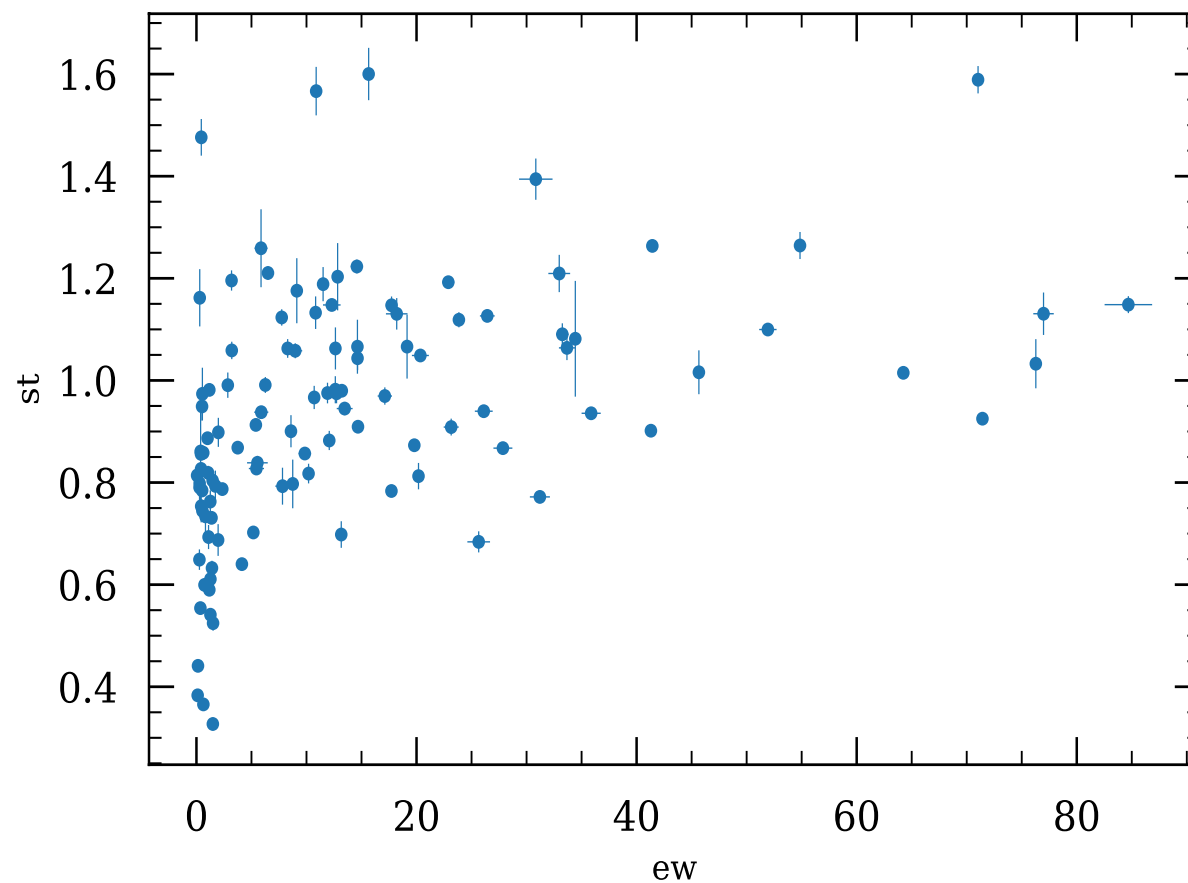
**SN
parameters
correlations**

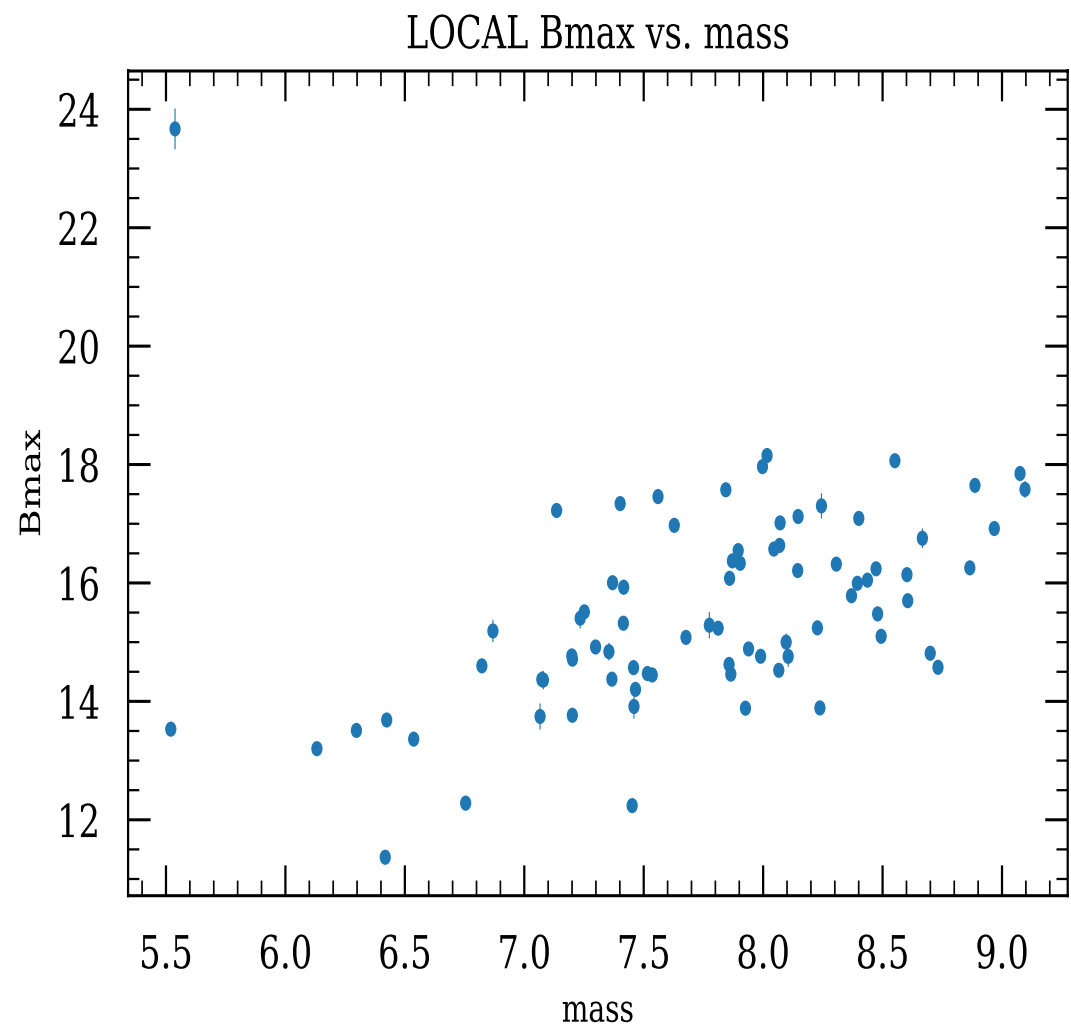
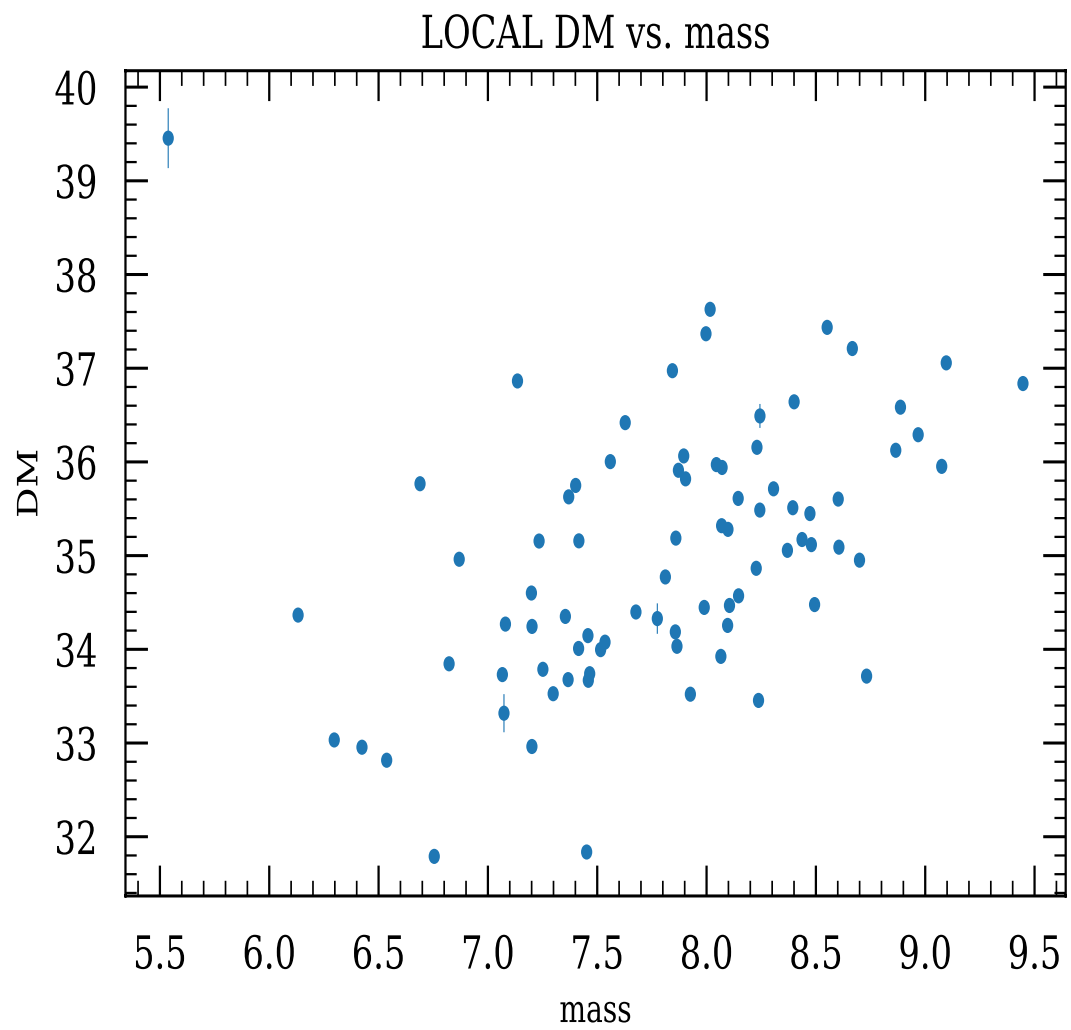
SN parameters – host galaxy local environment correlations

LOCAL st vs. agel



LOCAL st vs. ew





SN parameters – host galaxy TOTAL environment correlations

