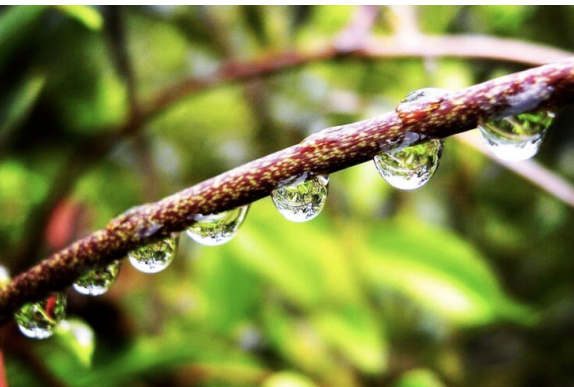


# Imaging through turbulence (2)

EC 522 Computational Optical Imaging

Lei Tian



# Admins

- » HWs

- » HW 6: posted, due 4/17

- » **Group lecture**

- » Reach out to me & your student mentor for discussions if needed

- » All questions are good questions

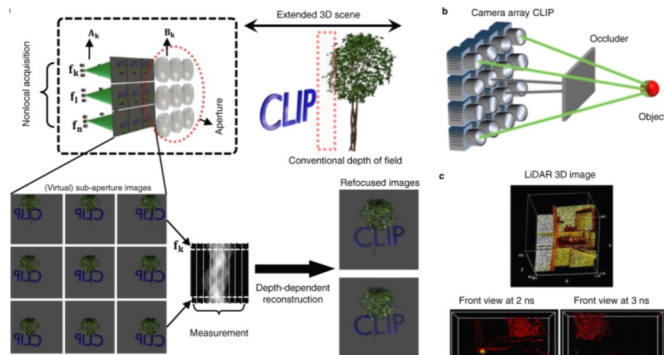
- » Grades are based on the demonstration of technical comprehension during the presentation



| Group lectures begin |   |  |           |                         |      |
|----------------------|---|--|-----------|-------------------------|------|
| 3/18<br>(M)          | Selected Topic in computational imaging |  | Group 1   | HW 5: inverse problem 2 | HW 4 |
| 3/20<br>(W)          | Selected Topic in computational imaging |  | Group 2   |                         |      |
| 3/25<br>(M)          | Selected Topic in computational imaging |  | Group 3   |                         |      |
| 3/27<br>(W)          | Selected Topic in computational imaging |  | Group 4   |                         |      |
| 4/1<br>(M)           | Selected Topic in computational imaging |  | Group 5   | HW 6: inverse problem 3 | HW 5 |
| 4/3<br>(W)           | Selected Topic in computational imaging |  | Group 6   |                         |      |
| 4/8<br>(M)           | Selected Topic in computational imaging |  | Group 7   |                         |      |
| 4/10<br>(W)          | Selected Topic in computational imaging |  | Group 8   |                         |      |
| 4/15<br>(M)          | Patriots' Day Holiday                   |  |           |                         |      |
| 4/17<br>(W)          | Selected Topic in computational imaging |  | Group 9   |                         | HW 6 |
| 4/22<br>(M)          | Cancelled                               |  |           |                         |      |
| 4/24<br>(W)          | <b>Final Projects</b>                   |  | Group 1-3 |                         |      |
| 4/29<br>(M)          | <b>Final Projects</b>                   |  | Group 4-6 |                         |      |
| 5/1<br>(W)           | <b>Final Projects</b>                   |  | Group 7-9 |                         |      |

# Theme: Computational photography

## Group Lecture / Class Project 3: Lightfield imaging



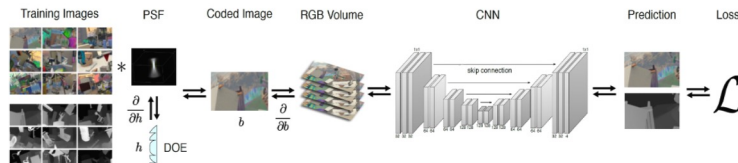
Team 4 - 3/27 (W) / / 4/29 (M)

**Paper: CLIP**

- Saini Ye
- Chen Qian
- Yuxiang Su

Student mentor:  
**Qianwan Yang**

## Group Lecture / Class Project 5: Computational 3D photography



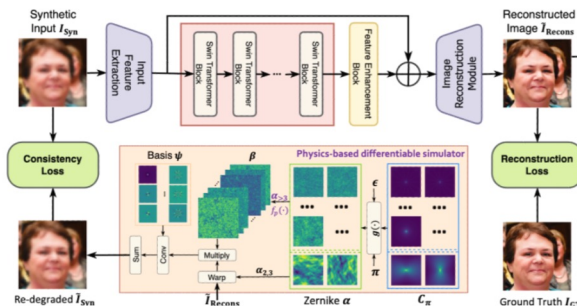
Team 5 - 4/1 (M) / 4/29 (M)

**Paper: Depth from defocus**

- Ian Lee
- Susan Zhang
- Wangyi Chen

Student Mentor:  
**Qianwan Yang**

## Group Lecture / Class Project 9: Computational imaging in complex media



Team 6 - 4/3 (W) / 4/29 (M)

**Paper: Turbulence**

- Shuyue Jia
- Hua Tong
- Yujie Zheng

Student mentor:  
**Tongyu Li**

Team 7 - 4/8 (M) / 5/1 (W)

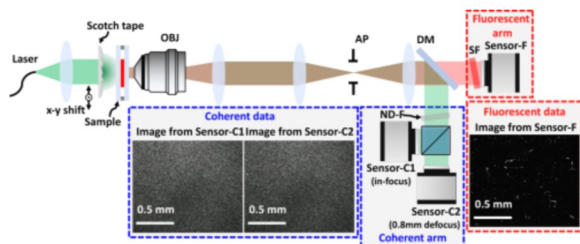
**Paper: NeuWS**

- Noa Margolin
- Nolan Vild
- Adhithi Ramasubramanian

Student Mentor:  
**Hao Wang**

# Theme: computational microscopy

## Group Lecture / Class Project 8: Computational imaging with structured illumination

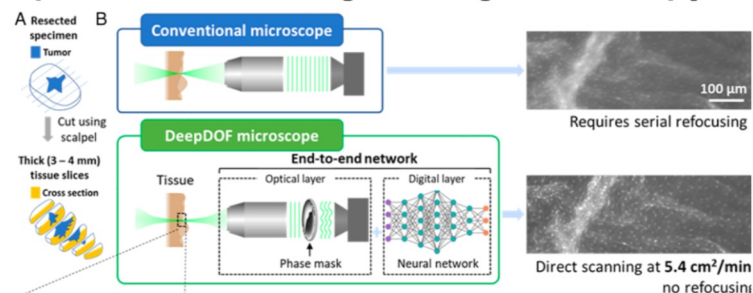


Team 8 - 4/10 (W) / 5/1 (W)

- Yi Shen
- Deming Li
- Kara Stratton

Student mentor:  
**Tongyu Li**

## Group Lecture / Class Project 7: Point spread function engineering microscopy

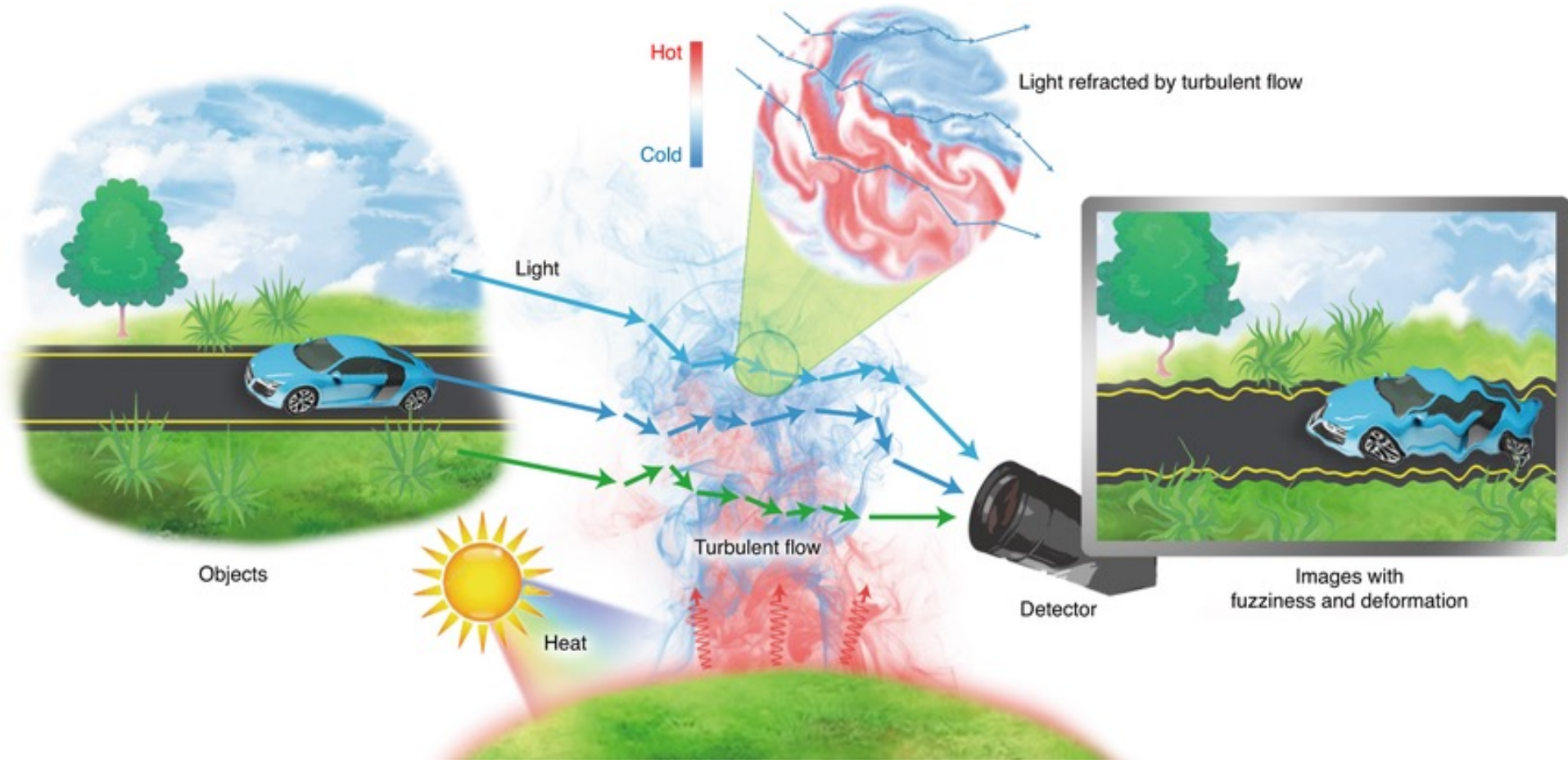


Team 9 - 4/17 (W) / 5/1 (W)

- Rachel Chan
- Qilin Deng

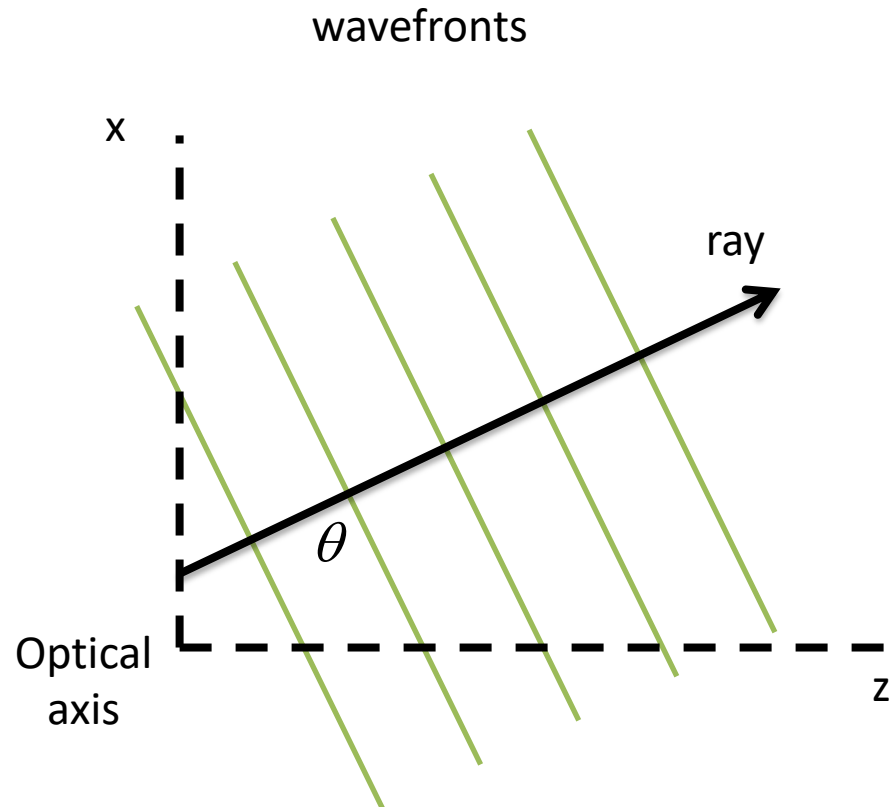
Student Mentor:  
**Joe Greene**

# Why imaging through turbulence



# Introduction to Adaptive Optics

# Precursor: rays & wavefront

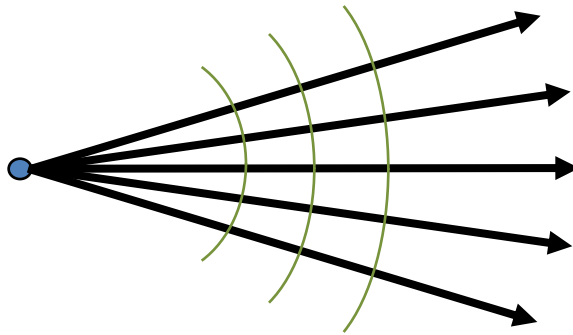




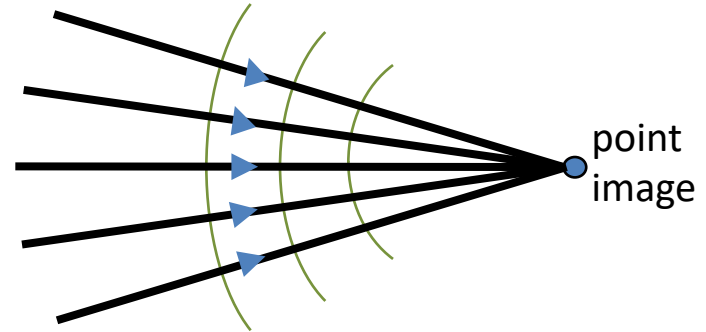
# rays & wavefront

diverging spherical wave

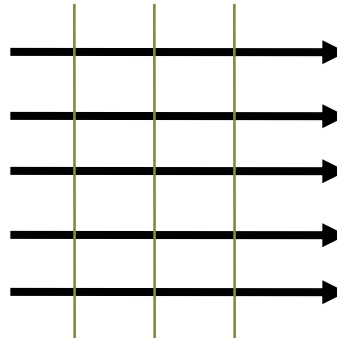
point source  
point object



converging spherical wave



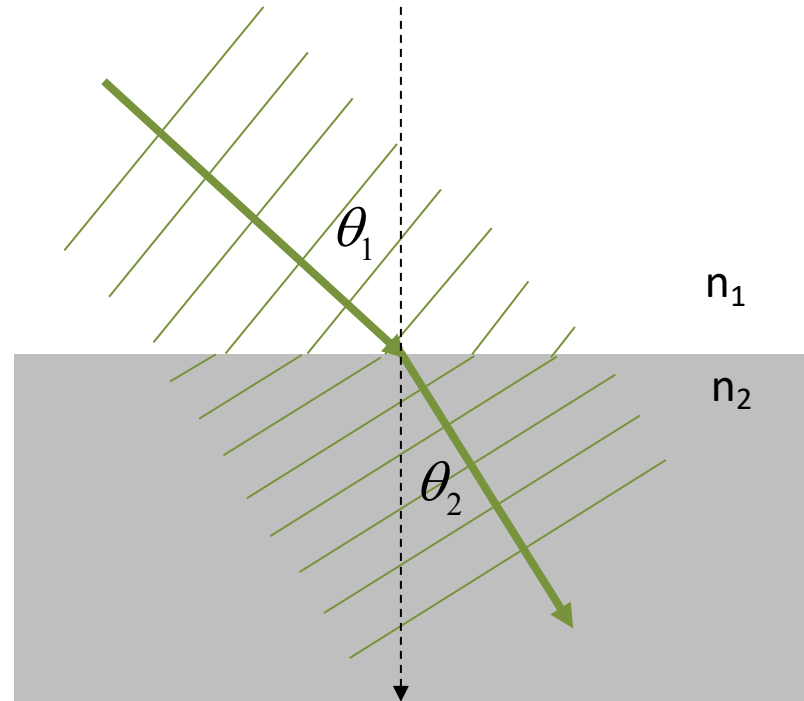
**Remember: wave-fronts  $\perp$  to rays**  
Rays travel left  $\rightarrow$  right



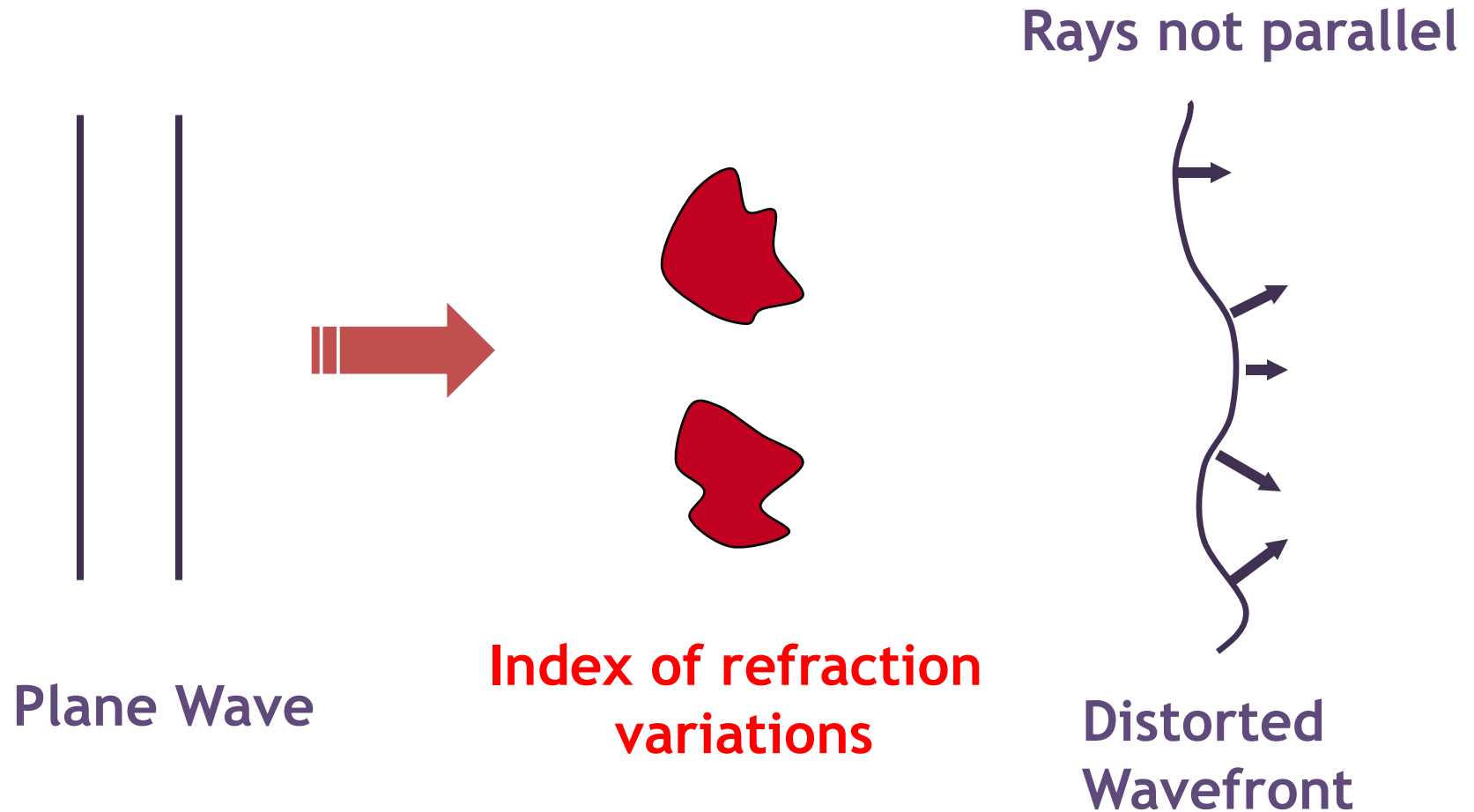
plane wave

# Rays of light bend when index of refraction Wavefront is also bent accordingly

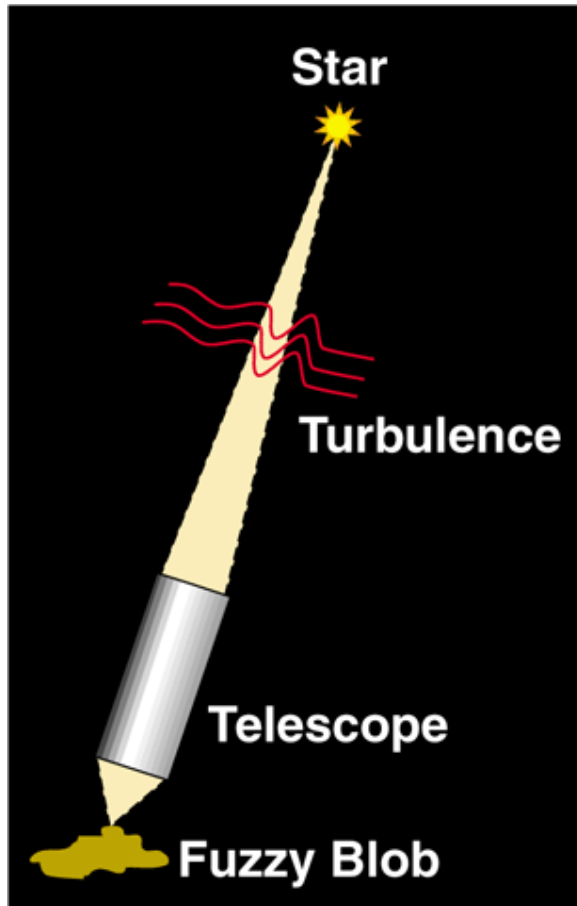
- » Light travels slower in more optically dense (higher index of refraction  $n$ ) materials
- » Causes rays to bend at the interface of two materials



# Atmospheric perturbations cause distorted wavefronts



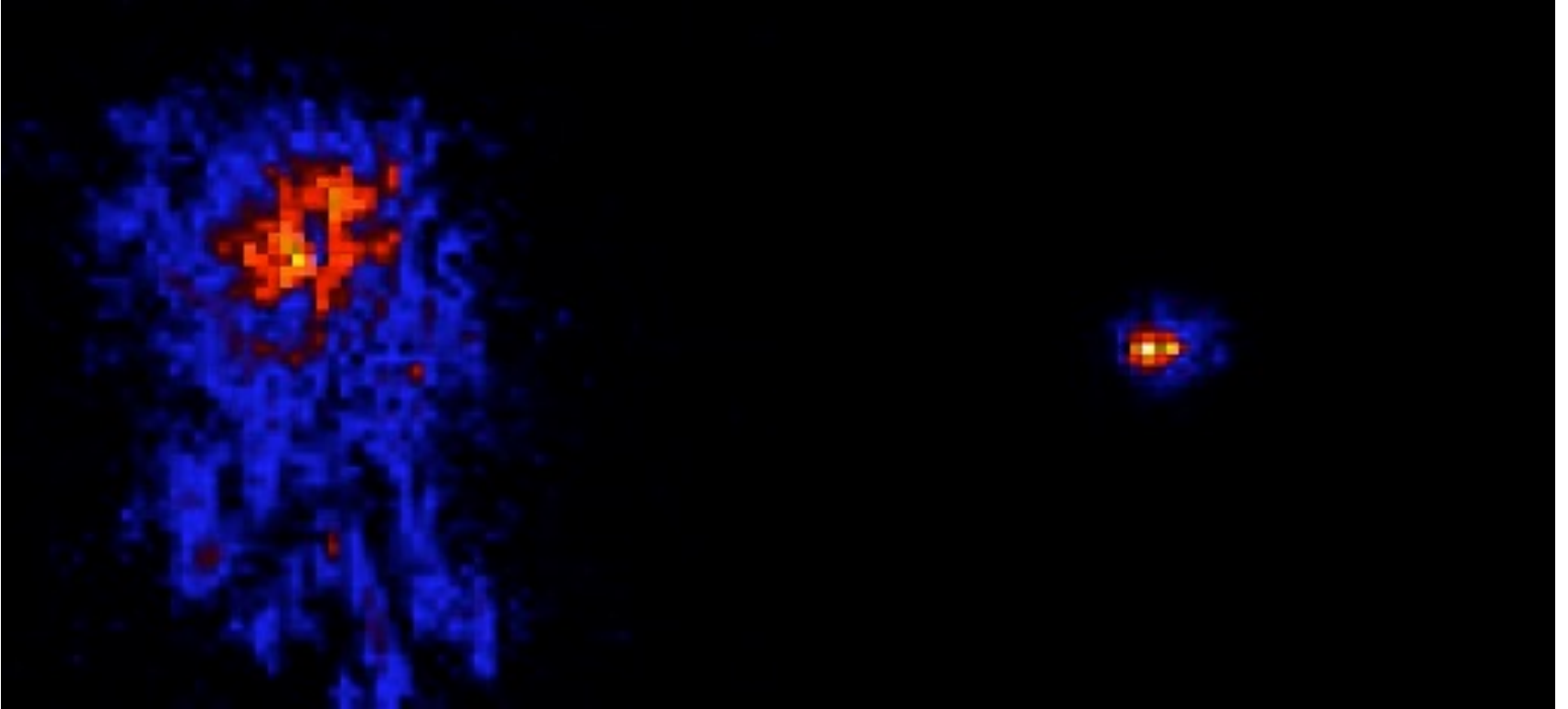
# Turbulence leads to fuzzy images



Turbulence in earth's atmosphere makes stars twinkle

More importantly, turbulence spreads out light; makes it a blob rather than a point

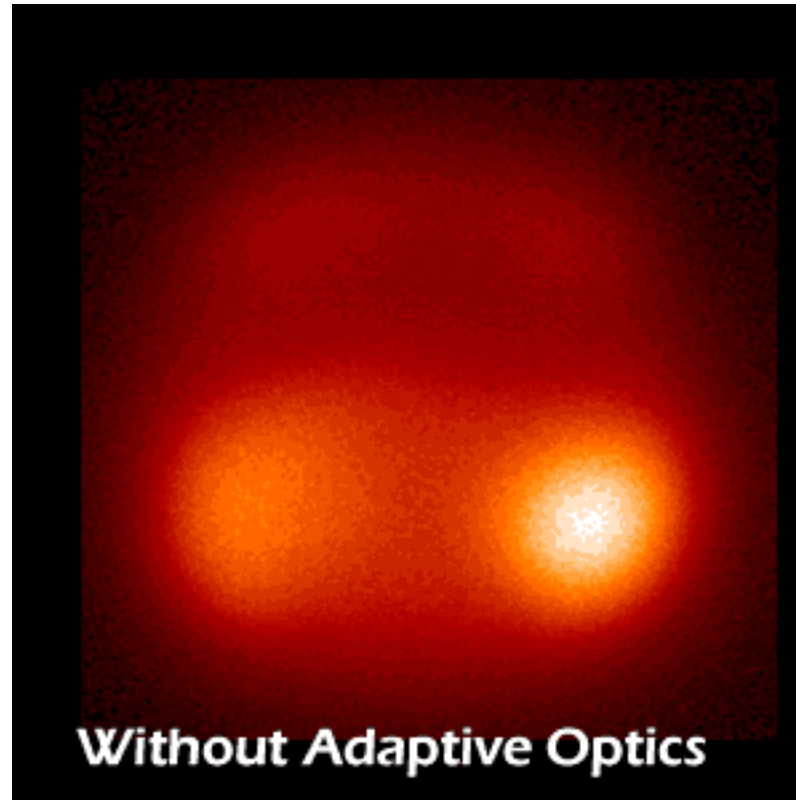
# Solution? → Adaptive Optics



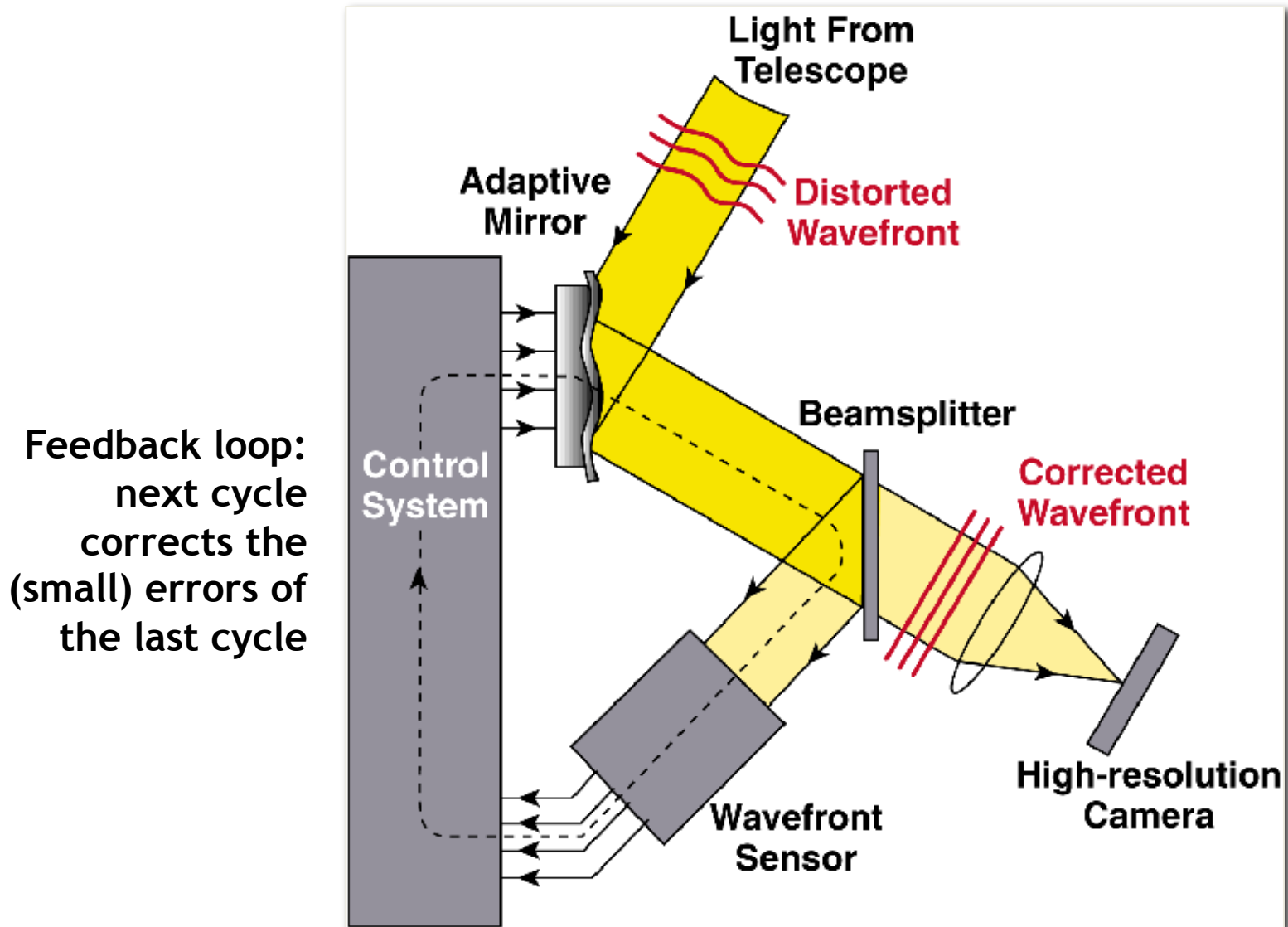
Infra-red images of a star, from Lick Observatory adaptive optics system

<https://www.ucolick.org/~max/289/>

# Solution? → Adaptive Optics



# How Adaptive Optics works?

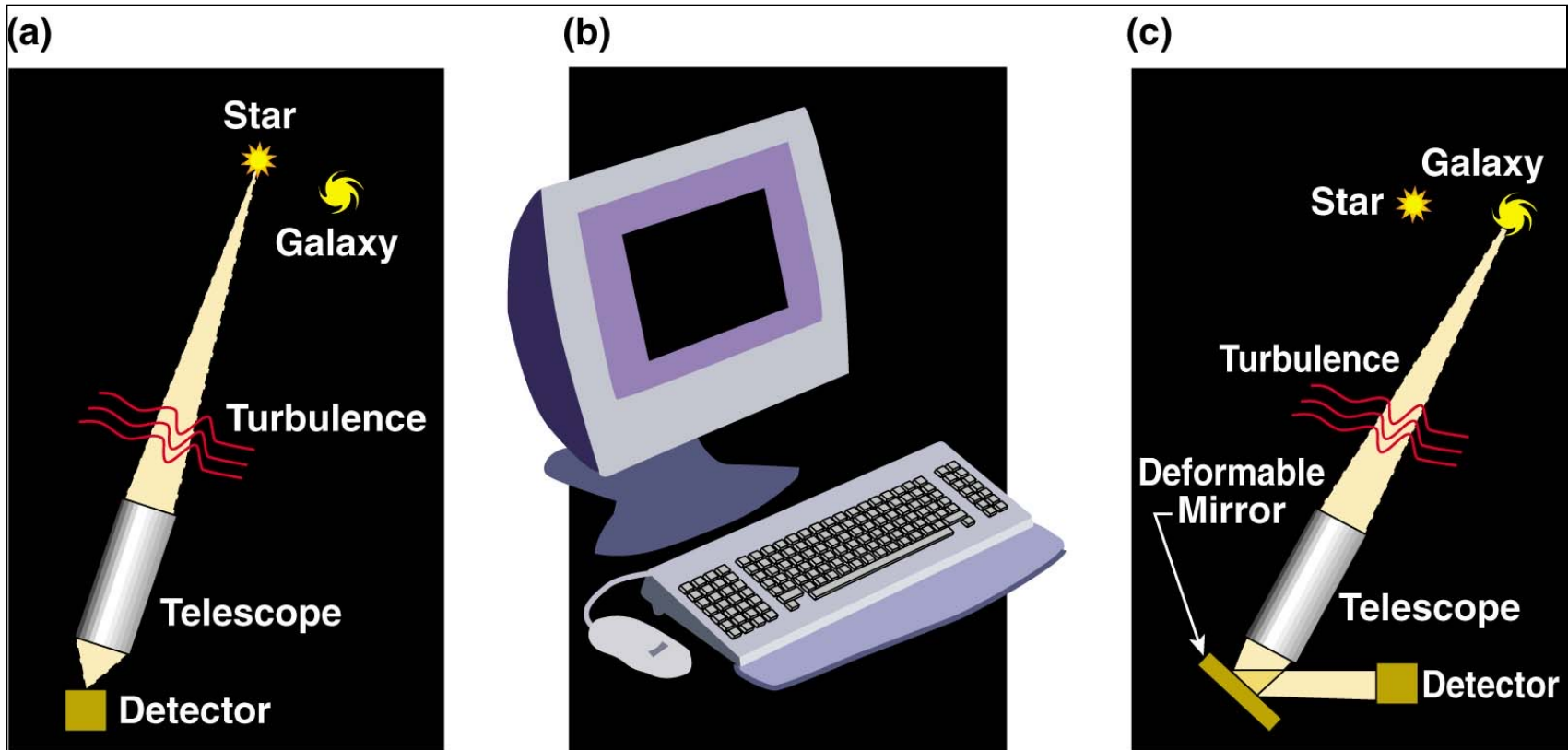


# How Adaptive Optics works?

*Measure details of blurring from “guide star” near the object you want to observe*

*Calculate (on a computer) the shape to apply to deformable mirror to correct blurring*

*Light from both guide star and astronomical object is reflected from deformable mirror; distortions are removed*



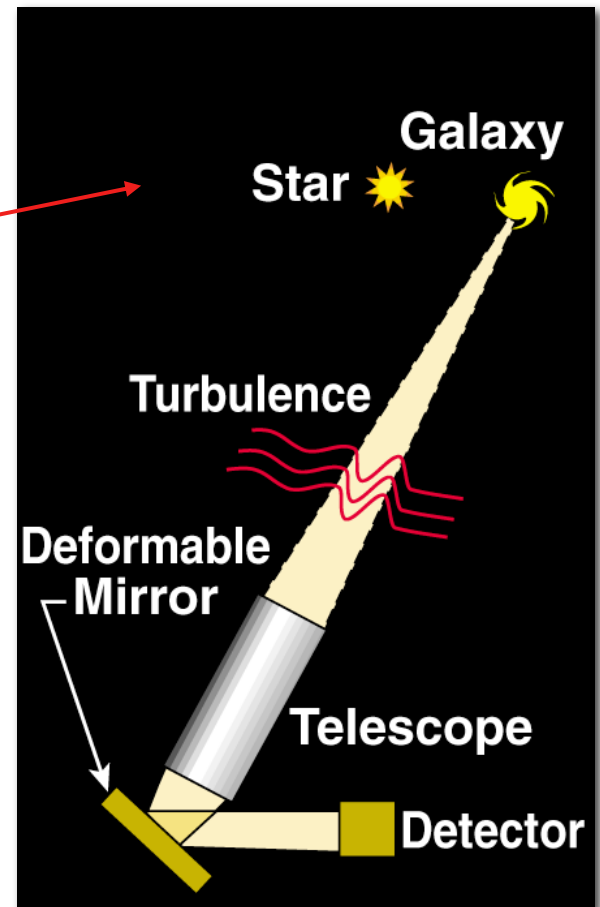


# How Adaptive Optics works?

## Need for “guide star”

If there's no close-by “real” star, create one with a laser

- » Use a laser beam to create artificial “star” at altitude of 100 km in atmosphere



# How Adaptive Optics works?

## Need for “guide star”



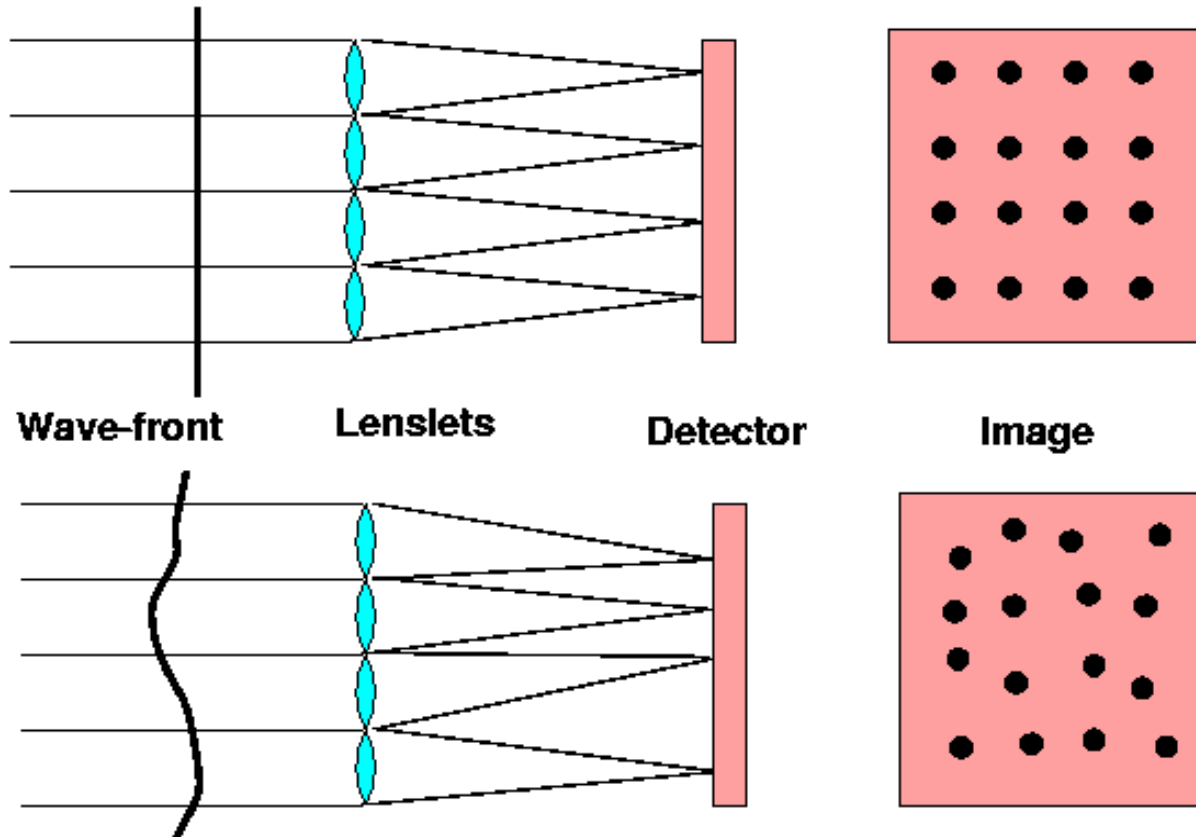
Four lasers on Mauna Kea: Keck 1 and 2, Gemini, Subaru telescopes

<https://www.ucolick.org/~max/289/>

# How Adaptive Optics works?

## How to measure the wavefront?

→ wavefront sensing

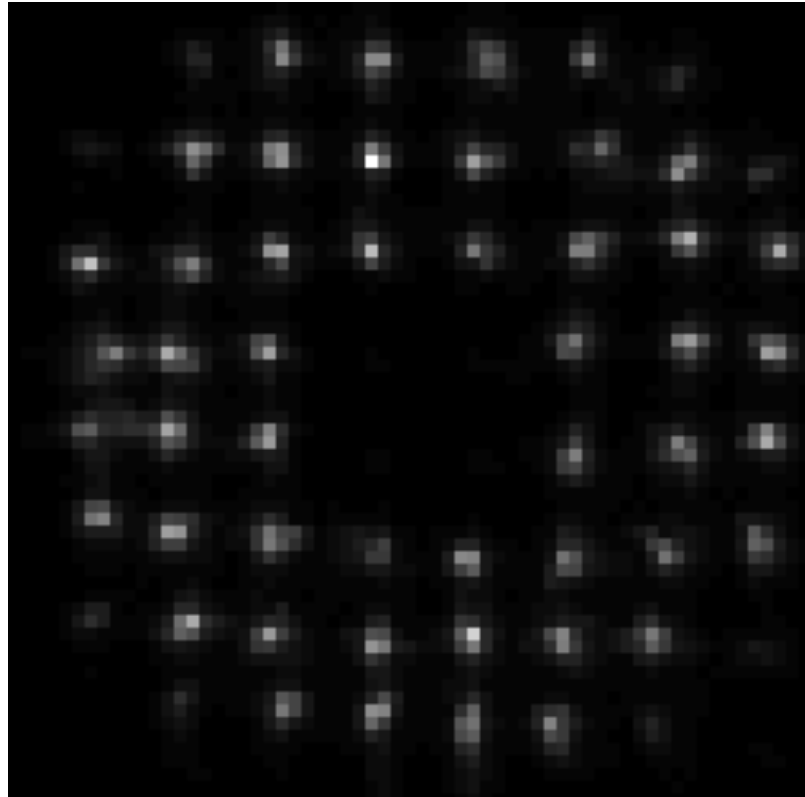


**Shack-Hartmann wavefront sensor**

# How Adaptive Optics works?

## How to measure the wavefront?

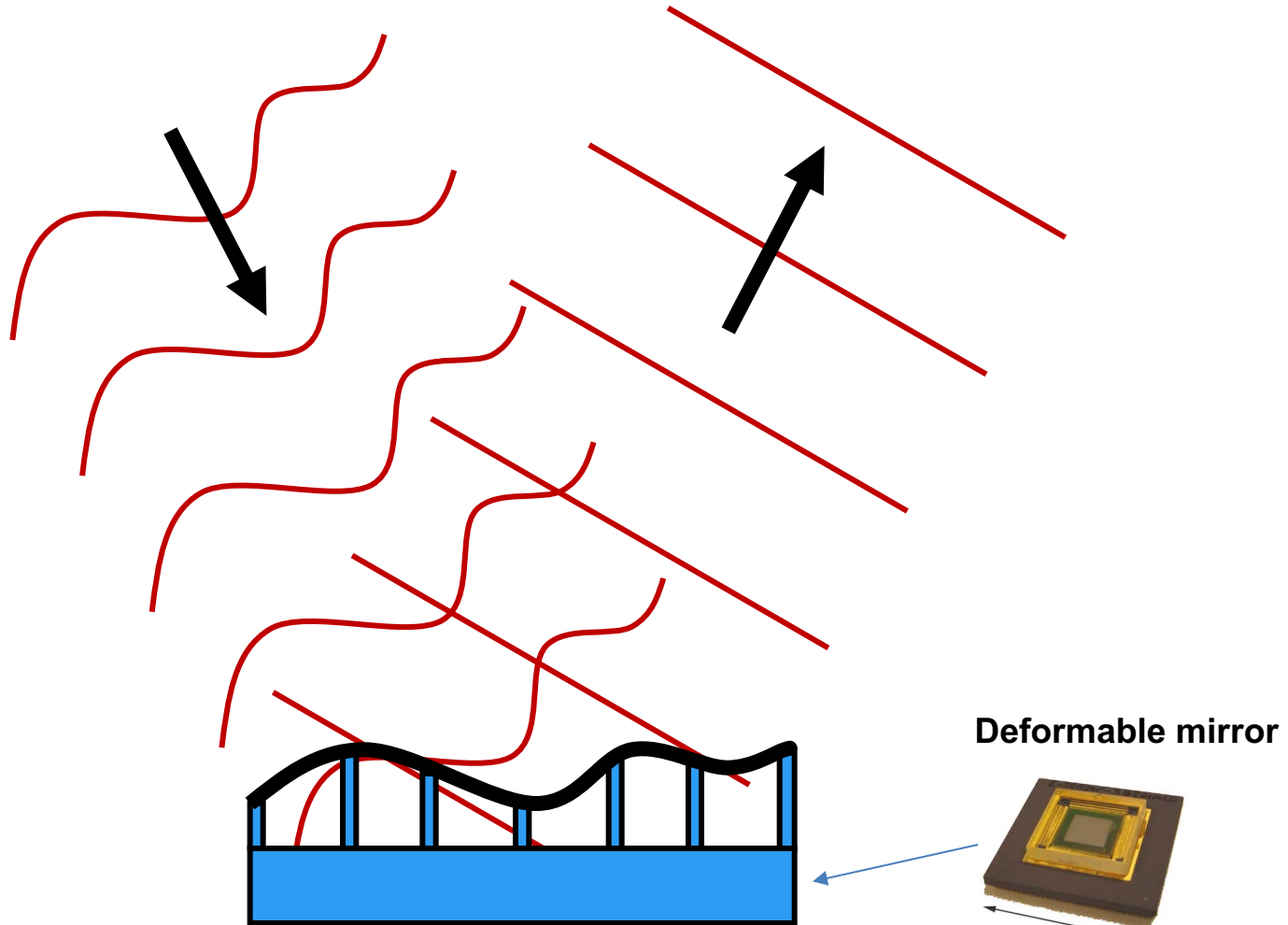
→ wavefront sensing



**Shack-Hartmann wavefront sensor**

# How Adaptive Optics works?

## How to **physically** correct for the wavefront?



**Emerging idea: Computational Adaptive Optics**