

Neutron transport theory and discrete ordinates method

For structural grid systems in Cartesian coordinate system

Xingguang Zhou¹

¹Nuclear Thermal-hydraulics Laboratory
Xi'an Jiaotong University

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Physical quantities and their relationships

Neutron angular flux (unit: $\text{cm}^{-2} \cdot \text{s}^{-1}$)

$$\phi(\mathbf{r}, \boldsymbol{\Omega}, E, t) = V \cdot N(\mathbf{r}, \boldsymbol{\Omega}, E, t).$$

Neutron scalar flux (unit: $\text{cm}^{-2} \cdot \text{s}^{-1}$)

$$\Phi(\mathbf{r}, E, t) = \int_0^{4\pi} \phi(\mathbf{r}, \boldsymbol{\Omega}, E, t) d\boldsymbol{\Omega}.$$

Neutron angular current density (unit: $\text{cm}^{-2} \cdot \text{s}^{-1}$)

$$\mathbf{J}(\mathbf{r}, \boldsymbol{\Omega}, E, t) = \boldsymbol{\Omega} \cdot \phi(\mathbf{r}, \boldsymbol{\Omega}, E, t).$$

Neutron current density (unit: $\text{cm}^{-2} \cdot \text{s}^{-1}$)

$$\mathbf{J}(\mathbf{r}, E, t) = \int_0^{4\pi} \mathbf{J}(\mathbf{r}, \boldsymbol{\Omega}, E, t) d\boldsymbol{\Omega}.$$

Generalized derivation of neutron transport equation

According to the neutron balance and Reynolds transport theorem, we can get the generalized form of neutron transport equation,

$$\begin{aligned} \frac{D}{Dt} \int_{\mathbb{V}} N(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^3r &= \frac{\partial}{\partial t} \int_{\mathbb{V}} N(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^3r \\ &+ \oint_{\mathbb{S}} \mathbf{n} \cdot \mathbf{V} \otimes N(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^2r, \end{aligned}$$

where \mathbf{n} is the unit normal vector of the surface, and $\mathbf{V} = \boldsymbol{\Omega}V$. Hence the equation can be rewritten as

$$\begin{aligned} \frac{D}{Dt} \int_{\mathbb{V}} N(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^3r &= \frac{\partial}{\partial t} \int_{\mathbb{V}} N(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^3r \\ &+ \oint_{\mathbb{S}} \mathbf{n} \cdot \boldsymbol{\Omega} \otimes \phi(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^2r. \end{aligned}$$

Generalized derivation of neutron transport equation

Considering the source terms and using Gauss theorem, we can get

$$\frac{\partial}{\partial t} \int_{\mathbb{V}} N(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^3r + \int_{\mathbb{V}} \nabla \cdot \boldsymbol{\Omega} \phi(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^3r = \int_{\mathbb{V}} Q(\mathbf{r}, \boldsymbol{\Omega}, E, t) d^3r.$$

The conservative neutron transport equation can be obtained by removing the volume integration,

$$\frac{\partial}{\partial t} N(\mathbf{r}, \boldsymbol{\Omega}, E, t) + \nabla \cdot \boldsymbol{\Omega} \phi(\mathbf{r}, \boldsymbol{\Omega}, E, t) = Q(\mathbf{r}, \boldsymbol{\Omega}, E, t).$$

The above equation still be rewritten into a form with more explicit physical meaning,

$$\frac{\partial}{\partial t} N(\mathbf{r}, \boldsymbol{\Omega}, E, t) + \nabla \cdot \mathbf{J}(\mathbf{r}, \boldsymbol{\Omega}, E, t) = Q(\mathbf{r}, \boldsymbol{\Omega}, E, t).$$

Generalized derivation of neutron transport equation

Hence, the conservative form of neutron transport equation with isotropic fission source in critical problem is

$$\begin{aligned}\nabla \cdot \mathbf{J}(\mathbf{r}, \boldsymbol{\Omega}, E) &= -\Sigma_t(\mathbf{r}, E) \phi(\mathbf{r}, \boldsymbol{\Omega}, E) \\ &+ \int_0^{4\pi} d\boldsymbol{\Omega}' \int_0^\infty dE' \Sigma_s(\mathbf{r}, \boldsymbol{\Omega} \rightarrow \boldsymbol{\Omega}', E' \rightarrow E) \phi(\mathbf{r}, \boldsymbol{\Omega}', E') \\ &+ \frac{\chi(E)}{4\pi k_{eff}} \int_0^\infty dE' \nu \Sigma_f(\mathbf{r}, E') \int_0^{4\pi} \phi(\mathbf{r}, \boldsymbol{\Omega}', E') d\boldsymbol{\Omega}',\end{aligned}$$

where

$$\mathbf{J}(\mathbf{r}, \boldsymbol{\Omega}, E) = \boldsymbol{\Omega} \phi(\mathbf{r}, \boldsymbol{\Omega}, E)$$

Discrete ordinates method - S_N

In neutron transport equation, there are four types arguments which are need to be solved, angular argument Ω , spatial argument \mathbf{r} , neutron energy E , and time t .

The most important argument for neutron transport equation is angular argument Ω , and actually Ω comes from the **Boltzmann equation**, which also can be used in radiative heat transfer. Discrete ordinates Method (S_N) is a popular method for solve Ω , that is first proposed by **Chandrasekhar to calculate the radiation and heat transfer of celestial bodies in the universe.**

Since S_N has the clear physical meaning, we use S_N to discretize Ω in neutron transport equation.

Discrete ordinates method - S_N

With the help of **Gaussian integration**, the integration of neutron angular flux to Ω can be derived as

$$\psi(\mathbf{r}, E, t) = \int_0^{4\pi} \phi(\mathbf{r}, \Omega, E, t) d\Omega = \sum_{m=1}^M \omega_m \phi(\mathbf{r}, \Omega_m, E, t),$$

where ω_m is the weight in **Gaussian integration**. Hence, if we want to discretize the Ω in **Cartesian coordinate system** from $\Omega_{m-\frac{1}{2}}$ to $\Omega_{m+\frac{1}{2}}$, we can get

$$\int_{\Omega_{m-\frac{1}{2}}}^{\Omega_{m+\frac{1}{2}}} \phi(\mathbf{r}, \Omega, E, t) d\Omega = \omega_m \phi(\mathbf{r}, \Omega_m, E, t).$$

In **Cartesian coordinate system**, the convective term can be written as

$$\therefore \nabla \cdot \Omega = 0,$$

$$\therefore \nabla \cdot (\Omega \phi(\mathbf{r}, \Omega, E, t)) = \phi \nabla \cdot \Omega + \Omega \cdot \nabla \phi \Rightarrow \Omega \cdot \nabla \phi.$$

Discrete ordinates method - S_N

For $\Omega \cdot \nabla \phi$, we also can derive the discretization of Ω from $\Omega_{m-\frac{1}{2}}$ to $\Omega_{m+\frac{1}{2}}$ in **Cartesian coordinate system**, as

$$\int_{\Omega_{m-\frac{1}{2}}}^{\Omega_{m+\frac{1}{2}}} \Omega \cdot \nabla \phi(\mathbf{r}, \Omega, E, t) d\Omega = \omega_m \Omega_m \cdot \nabla \phi(\mathbf{r}, \Omega_m, E, t).$$

Why we always alert **"Cartesian coordinate system"**? This is because only in Cartesian coordinate system, which does not contain the **leakage term** (part of neutron flux) due to coordinate change, above operation of discretization of Ω is established in such a way like "integral mean value theorem".

$$\begin{aligned} \int_{\Omega_{m-\frac{1}{2}}}^{\Omega_{m+\frac{1}{2}}} \Omega \cdot \nabla \phi(\Omega) d\Omega &= \omega_m \Omega_m \cdot \nabla \phi(\Omega_m) \\ &= \Omega_{m'} \cdot \nabla \phi(\Omega_{m'}) \left| \Omega_{m+\frac{1}{2}} - \Omega_{m-\frac{1}{2}} \right|. \end{aligned}$$

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Basic principle of symmetry and normalization

Actually, in neutron transport equation, one-dimensional condition and two-dimensional condition are both the **special cases of three-dimensional condition**. The integrals of the independent angle variables are

1). one-dimensional ("only polar angle $\theta \in [0, \pi]$ ")

$$\frac{1}{2\pi} \int_0^{2\pi} d\varphi \int_0^\pi \sin \theta d\theta = 2 \Rightarrow \int_0^\pi \sin \theta d\theta,$$

2). two-dimensional (azimuthal angle $\varphi \in [0, 2\pi]$, polar angle $\theta \in [0, \pi/2]$)

$$\frac{1}{2} \int_0^{2\pi} d\varphi \int_0^\pi \sin \theta d\theta = 2\pi \Rightarrow \int_0^{2\pi} d\varphi \int_0^{\frac{\pi}{2}} \sin \theta d\theta,$$

3). three-dimensional (azimuthal angle $\varphi \in [0, 2\pi]$, polar angle $\theta \in [0, \pi]$)

$$\int_0^{2\pi} d\varphi \int_0^\pi \sin \theta d\theta = 4\pi.$$

Basic principle of symmetry and normalization

However, the integrals of the independent angle variables are not the total integrals. In fact, we still get

$$\sum_{m=1}^{M_{(1)}} \omega_{m_{(1)}} = \frac{1}{2\pi} \sum_{m=1}^{M_{(3)}} \omega_m = 2, \quad \sum_{m=1}^{M_{(2)}} \omega_{m_{(2)}} = \frac{1}{2} \sum_{m=1}^{M_{(3)}} \omega_m = 2\pi,$$

for 1-D, 2-D, and 3-D, respectively. In most context of neutron transport equation, the increment solid angle $d\Omega$ is usually defined as

$$d\Omega = \frac{d\varphi \sin \theta d\theta}{2\pi \cdot 2},$$

to yield the convenient normalization over all angles

$$\int_0^{4\pi} d\Omega = \int_0^{2\pi} \frac{d\varphi}{2\pi} \int_0^\pi \frac{\sin \theta d\theta}{2} = 1.$$

Basic principle of symmetry and normalization

After the **normalization of $d\Omega$** , for 1-D, 2-D, and 3-D, that

$$\int_0^{4\pi} d\Omega = \sum_{m=1}^M \omega_m = 1.$$

The reason we consider that normalization of $d\Omega$ is that we need to use **Gaussian quadrature set** in S_N method.

From the **Numerical Analysis**, we can know that usually $\sum \omega_m = 1$ or 2 in Gaussian quadrature set. It is **rarely and hard** to make $\sum \omega_m = 4\pi$, which means that **Gaussian integral points also need to be re-designed**. In the other words, the quadrature sets with $\sum \omega_m = 1$ or 2 are very mature. Hence, we make the normalization of $d\Omega$ to adapt Gaussian quadrature set.

In addition, there are some other physical considerations, which will make the computational load much lesser, according to "old" computers.

Basic principle of symmetry and normalization

For isotropic fission source in steady-state critical problems, we always get

$$Q_f(\mathbf{r}, \boldsymbol{\Omega}, E) = \frac{\chi(E)}{4\pi k_{eff}} \int_0^\infty dE' \int_0^{4\pi} d\boldsymbol{\Omega}' \nu \Sigma_f(E') \phi(\mathbf{r}, \boldsymbol{\Omega}', E'),$$

if we use the normalization of $d\boldsymbol{\Omega}$, Q_f can be written as

$$\begin{aligned} Q_f(\mathbf{r}, \boldsymbol{\Omega}, E) &= \frac{\chi(E)}{k_{eff}} \int_0^\infty dE' \int_0^{4\pi} \frac{d\boldsymbol{\Omega}'}{4\pi} \nu \Sigma_f(E') \phi(\mathbf{r}, \boldsymbol{\Omega}', E') \\ &= \frac{\chi(E)}{k_{eff}} \int_0^\infty dE' \nu \Sigma_f(E') \sum_{m=1}^M \frac{\omega_m}{4\pi} \phi(\mathbf{r}, \boldsymbol{\Omega}'_m, E'). \end{aligned}$$

We can see that 4π is integrated in the quadrature weight ω_m , and then the identity of the Q_f is unchanged. This identity transformation also meets the requirements of existing Gaussian quadrature sets.

Basic principle of symmetry and normalization

For scattering source term in steady-state critical problems, we usually use "isotropic" scattering, as

$$Q_s(\mathbf{r}, \boldsymbol{\Omega}, E) = \frac{1}{2\pi} \int_0^\infty dE' \int_0^{4\pi} d\boldsymbol{\Omega}' \Sigma_s \left(\mathbf{r}, \underbrace{\boldsymbol{\Omega}' \cdot \boldsymbol{\Omega}}_{\mu_0}, E' \rightarrow E \right) \phi(\mathbf{r}, \boldsymbol{\Omega}', E').$$

We can expand the scattering macroscopic with Legendre expansions, as

$$\Sigma_s(\mathbf{r}, \mu_0, E' \rightarrow E) = \sum_{l=0}^{\infty} \frac{2l+1}{2} \Sigma_{s,l}(\mathbf{r}, E' \rightarrow E) P_l(\mu_0),$$

if $l = 0$, then we get the "isotropic" scattering source term,

$$Q_s(\mathbf{r}, \boldsymbol{\Omega}, E) = \int_0^\infty \frac{1}{4\pi} \Sigma_s(\mathbf{r}, E' \rightarrow E) \int_0^{4\pi} \phi(\mathbf{r}, \boldsymbol{\Omega}', E') d\boldsymbol{\Omega}' dE',$$

Basic principle of symmetry and normalization

$$\begin{aligned} Q_s(\mathbf{r}, \boldsymbol{\Omega}, E) &= \int_0^\infty \frac{1}{4\pi} \Sigma_s(\mathbf{r}, E' \rightarrow E) \sum_{m=1}^M \omega_m \phi(\mathbf{r}, \boldsymbol{\Omega}_m, E') dE' \\ &= \int_0^\infty \Sigma_s(\mathbf{r}, E' \rightarrow E) \sum_{m=1}^M \frac{\omega_m}{4\pi} \phi(\mathbf{r}, \boldsymbol{\Omega}_m, E') dE'. \end{aligned}$$

From the above discussions, we can clearly know that if we use S_N method with Gaussian integrations, multiplication factors (such as 4π) can be introduced to maintain the $\sum \omega_m = 1$. And the Gaussian integral points $\boldsymbol{\Omega}_m$ do not need to be changed. Hence the source terms as Q_s and Q_f are need to be re-written into the identity form. After the above identity transformation, the "real" neutron (angular) scalar flux can be calculated.

Basic principle of symmetry and normalization

For the isotropic source terms in 1-D problem with single group, we can get

$$\mu \frac{\partial \phi(z, \boldsymbol{\Omega})}{\partial z} + \Sigma_t(z) \phi(z, \boldsymbol{\Omega}) = \frac{1}{2\pi} \int_0^{2\pi} d\varphi' \int_1^{-1} \Sigma_s(z, \mu_0) \phi(z, \boldsymbol{\Omega}') d\mu' \\ + \frac{1}{4\pi k_{eff}} \nu \Sigma_f \Phi(z),$$

where $\mu_0 = \boldsymbol{\Omega}' \cdot \boldsymbol{\Omega}$. With the basic principle of symmetry, the equation can be re-written into

$$\mu \frac{\partial \phi(z, \mu)}{\partial z} + \Sigma_t(z) \phi(z, \mu) = \int_1^{-1} \Sigma_s(z, \mu_0) \phi(z, \mu') d\mu' \\ + \frac{1}{2 \cdot k_{eff}} \nu \Sigma_f \Phi(z),$$

where $\phi(z, \boldsymbol{\Omega}') = \frac{1}{2\pi} \phi(z, \mu')$, and $\Sigma_s(z, \boldsymbol{\Omega}' \rightarrow \boldsymbol{\Omega}) = \frac{1}{2\pi} \Sigma_s(z, \mu_0)$

Basic principle of symmetry and normalization

With the isotropic source terms, the final 1-D problem with single group can be written as

$$\mu \frac{\partial \phi(z, \mu)}{\partial z} + \Sigma_t(z) \phi(z, \mu) = \frac{1}{2} \Sigma_s(z) \Phi(z) + \frac{1}{2} \frac{1}{k_{eff}} \nu \Sigma_f \Phi(z).$$

Hence, $\Phi(z) = \sum_{m=1}^M \omega_m \phi(z, \mu_m)$. We suppose that move $\frac{1}{2}$ into ω_m to ensure that $\sum_{m=1}^M \omega_m = 1$. Therefore, the equation in the code for 1-D critical problem can be written as

$$\mu \frac{\partial \phi(z, \mu)}{\partial z} + \Sigma_t \phi(z, \mu) = \Sigma_s \sum_{m=1}^M \omega_m \phi(z, \mu_m) + \frac{\nu \Sigma_f}{k_{eff}} \sum_{m=1}^M \omega_m \phi(z, \mu_m).$$

$$\sum_{m=1}^M \omega_m = 1$$

Basic principle of symmetry and normalization

Discretization of the above equation with S_N method, we can get

$$\mu_m \frac{\partial \phi_m(z)}{\partial z} + \Sigma_t \phi_m(z) = \Sigma_s \sum_{m=1}^M \omega_m \phi(z, \mu'_m) + \frac{\nu \Sigma_f}{k_{eff}} \sum_{m=1}^M \omega_m \phi_m(z),$$

and also use FVM to discretize the spatial parameters, as

$$\mu_m \left(\phi_{i+\frac{1}{2},m} - \phi_{i-\frac{1}{2},m} \right) + \Sigma_{t,i} \phi_{i,m} \Delta z = \left(\Sigma_{s,i} + \frac{\nu \Sigma_{f,i}}{k_{eff}} \right) \sum_{m=1}^M \omega_m \phi_{i,m} \Delta z.$$

Let

$$Q_{i,m} = \left(\Sigma_{s,i} + \frac{\nu \Sigma_{f,i}}{k_{eff}} \right) \sum_{m=1}^M \omega_m \phi_{i,m}$$

Basic principle of symmetry and normalization

Hence

$$\mu_m \left(\phi_{i+\frac{1}{2},m} - \phi_{i-\frac{1}{2},m} \right) + \Sigma_{t,i} \phi_{i,m} \Delta z = Q_{i,m} \Delta z.$$

With the central difference, we can get

$$\begin{cases} \phi_{i+\frac{1}{2},m} = 2\phi_{i,m} - \phi_{i-\frac{1}{2},m}, & \mu_m > 0 \\ \phi_{i-\frac{1}{2},m} = 2\phi_{i,m} - \phi_{i+\frac{1}{2},m}, & \mu_m < 0, \end{cases}$$

and first-order upwind scheme

$$\begin{cases} \phi_{i+\frac{1}{2},m} = \phi_{i,m}, & \mu_m > 0 \\ \phi_{i-\frac{1}{2},m} = \phi_{i,m}, & \mu_m < 0, \end{cases}$$

substituting the scheme into the equation, we can get the solution of the iteration.

Basic principle of symmetry and normalization

Solution of central difference and first-order upwind scheme

$$\phi_{i,m} = \frac{Q_{i,m} \Delta z + C \mu_m \phi_{i-\frac{1}{2},m}}{C \mu_m + \Sigma_{t,i} \Delta z}, \quad \mu_m > 0,$$
$$\phi_{i,m} = \frac{Q_{i,m} \Delta z - C \mu_m \phi_{i+\frac{1}{2},m}}{-C \mu_m + \Sigma_{t,i} \Delta z}, \quad \mu_m < 0,$$

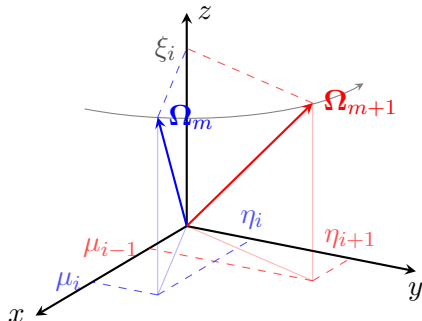
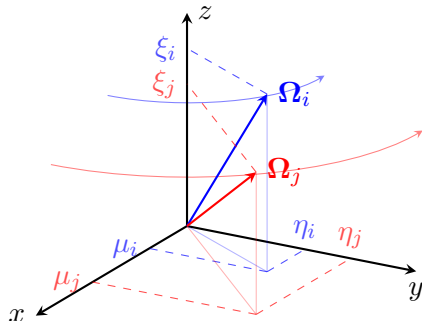
where, $C = 2$ in central difference scheme, and $C = 1$ in first-order upwind scheme.

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Ω and its properties

$\mu = \Omega \cdot \mathbf{e}_x$, $\eta = \Omega \cdot \mathbf{e}_y$, $\xi = \Omega \cdot \mathbf{e}_z$ are the projections of angular parameter Ω . It means that Ω can be written as $\Omega = (\mu, \eta, \xi)$.



$$\mu^2 + \eta^2 + \xi^2 = 1, \quad \mu_i^2 + \eta_i^2 + \xi_i^2 = 1,$$

$$\Rightarrow \mu_{i-1}^2 + \eta_{i+1}^2 + \xi_i^2 = 1.$$

Ω and its properties

Ω and Gaussian quadrature set also have the following properties. For $\phi(\mathbf{r}, \Omega) = 1$, that

$$\Phi(\mathbf{r}) = \int_0^{4\pi} \phi(\mathbf{r}, \Omega) d\Omega = \sum_{m=1}^M \omega_m \phi(\mathbf{r}, \Omega_m) = 4\pi.$$

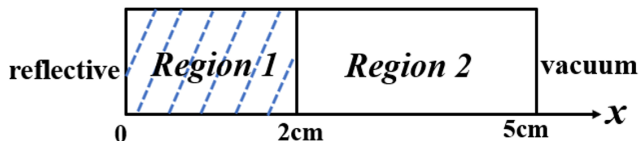
With the isotropic neutron angular flux, that

$$\mathbf{J}(\mathbf{r}) = \int_0^{4\pi} \mathbf{J}(\mathbf{r}, \Omega) = \sum_{m=1}^M \omega_m \Omega_m \phi_C(\mathbf{r}, \Omega_m) = \mathbf{0} \Rightarrow \sum_{m=1}^M \omega_m \Omega_m = \mathbf{0}.$$

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Problem description



Group constants of Region 1

$$\Sigma_t = 1.0 \text{ cm}^{-1}, \quad \Sigma_s = 0.5 \text{ cm}^{-1}, \quad \nu\Sigma_f = 1.0 \text{ cm}^{-1}.$$

Group constants of Region 2

$$\Sigma_t = 0.8 \text{ cm}^{-1}, \quad \Sigma_s = 0.4 \text{ cm}^{-1}, \quad \nu\Sigma_f = 0.0 \text{ cm}^{-1}.$$

Use S8 method to solve this problem with central difference scheme and first-order upwind scheme.

Algorithm 1 solve $\phi_{i,m}, \mu_m > 0$

Require: $\phi_{i,m}$, boundary, $\phi_{i,m}^{User}$

```
1: while  $m \in [1, M], \mu_m > 0$  do
2:   while  $i \in [1, I]$  do
3:     if Reflective then
4:        $\phi_{i,m} \Rightarrow \phi_{i,M-m+1}$ 
5:     else
6:        $\phi_{i,m} \Rightarrow \phi_{i,m}^{User}$ 
7:     end if
8:      $\phi_{i,m} = \frac{Q_{i,m}, \Delta z + C\mu_m \phi_{i-\frac{1}{2},m}}{C\mu_m + \Sigma_{t,i}\Delta z}$ 
9:      $\phi_{i,m+\frac{1}{2}} \Leftarrow C\phi_{i,m} - (C-1)\phi_{i,m-\frac{1}{2}}$ 
10:    call NegativeFluxCorrection
11:  end while
12: end while
```

Algorithm 2 solve $\phi_{i,m}, \mu_m < 0$

Require: $\phi_{i,m}$, boundary, $\phi_{i,m}^{User}$

```
1: while  $m \in [1, M], \mu_m < 0$  do
2:   while  $i \in [1, I]$  do
3:     if Reflective then
4:        $\phi_{i,m} \Rightarrow \phi_{i,M-m+1}$ 
5:     else
6:        $\phi_{i,m} \Rightarrow \phi_{i,m}^{User}$ 
7:     end if
8:      $\phi_{i,m} = \frac{Q_{i,m}, \Delta z - C\mu_m \phi_{i+\frac{1}{2},m}}{-C\mu_m + \Sigma_{t,i}\Delta z}$ 
9:      $\phi_{i,m-\frac{1}{2}} \Leftarrow C\phi_{i,m} - (C-1)\phi_{i,m+\frac{1}{2}}$ 
10:    call NegativeFluxCorrection
11:  end while
12: end while
```

Algorithm 3 source term

Require: $\phi_{i,m}, k_{eff}$

```
1: while  $m \in [1, M]$  do
2:   while  $i \in [1, I]$  do
3:      $\Phi_i \leftarrow 0$ 
4:     while  $m' \in [1, M]$  do
5:        $\Phi_i \leftarrow \sum_{m'}^M \omega_{m'} \phi_{i,m'}$ 
6:     end while
7:      $Q_{i,m} \leftarrow \left( \Sigma_{s,i} + \frac{\nu \Sigma_{f,i}}{k_{eff}} \right) \Phi_i$ 
8:   end while
9: end while
```

Algorithm 4 eigenvalue

Require: $\Phi_i, \Phi_i^{Old}, k_{eff}^{Old}$

```
1:  $Q_f \leftarrow \sum_{i=1}^I \nu \Sigma_i \Phi_i$ 
2:  $Q_f^{Old} \leftarrow \sum_{i=1}^I \nu \Sigma_i \Phi_i^{Old}$ 
3:  $k_{eff} = \frac{Q_f \cdot k_{eff}^{Old}}{Q_f^{Old}}$ 
```

Algorithm 5 power iteration

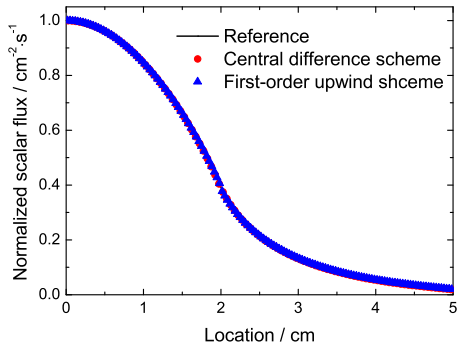
Require: $\phi_{i,m}$, boundary, k_{eff}

```
1: call MeshGeneration
2: call Initialization
3: while  $iter \in [1, iMAX]$  do
4:   call SourceTerm
5:   call Solve  $\phi_{i,m}, \mu_m < 0$ 
6:   call Solve  $\phi_{i,m}, \mu_m > 0$ 
7:   call EigenValue
8:   if  $|k_{eff} - k_{eff}^{Old}| < 1.e-6$  then
9:     return
10:  end if
11:   $\phi_{i,m}^{Old} \Leftarrow \phi_{i,m}$ 
12:   $k_{eff}^{Old} \Leftarrow k_{eff}$ 
13: end while
```

Result and discussion

Table 1: Comparison of k_{eff}

Method	k_{eff}	Deviation
MCNP	1.67869	-
S8-CD	1.68169	+300 pcm
S8-FUD	1.67837	-32 pcm



The results show that S8 method with central difference and first-order upwind schemes are both can calculation the ISSA problem correctly. We use 100 nodes in CD scheme and 150 nodes in FUD scheme, that in second order accuracy and first order accuracy, respectively.

Conclusion

- ① Based on the conservation of neutron number and the basic physical meaning of each physical quantity, we deduced the neutron transport equation.
- ② Symmetry and discrete ordinate method in neutron transport equation are introduced.
- ③ Finally, we code the S8 to solve ISSA benchmark with Fortran 90. The results show that in the deterministic neutronics calculation, grid system, discretized angle and scheme accuracy are all related, and the characteristics such as **error cancellation** need to be considered.

Thanks for your attention!

zxg818@stu.xjtu.edu.cn