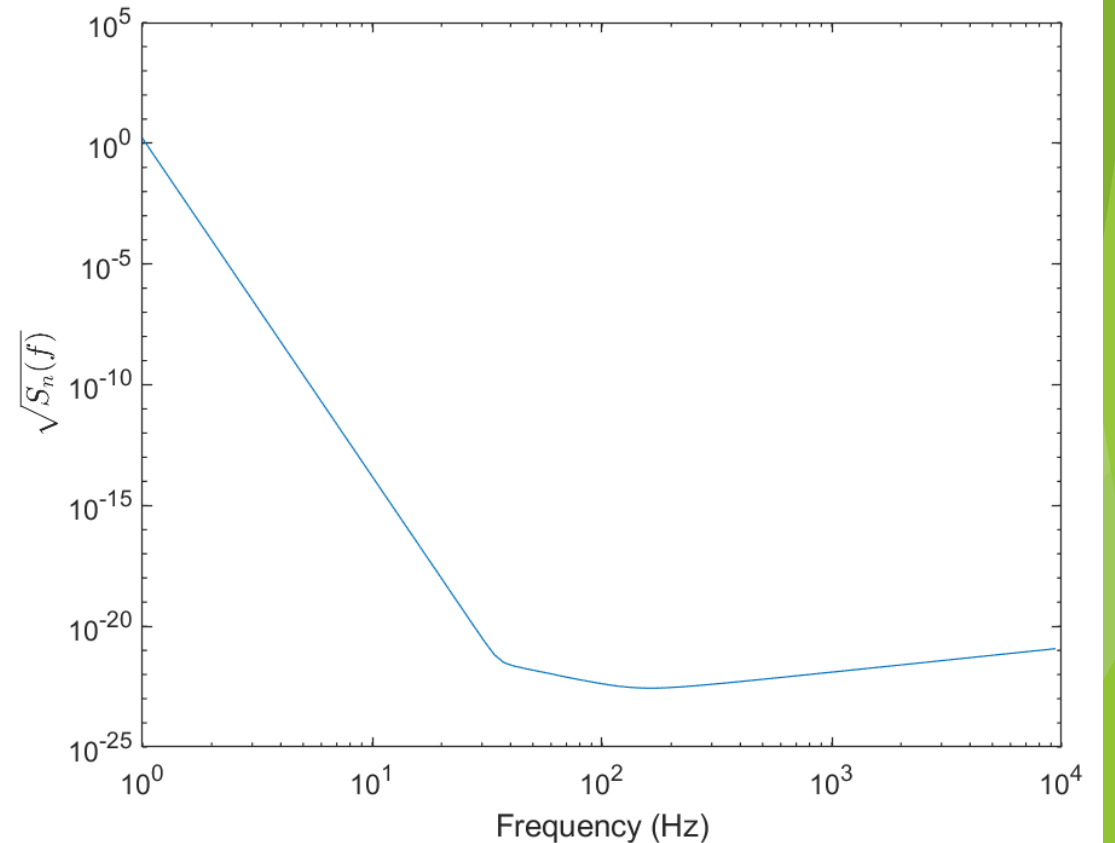


Simulating LIGO noise

Objective

- ▶ Simulate the noise of an interferometric detector
- ▶ We will pick the sensitivity curve of the initial LIGO detector as an example for the target PSD
- ▶ The same steps can be used for any other design sensitivity curve (e.g., advanced LIGO, LISA etc)



Initial LIGO design sensitivity

- ▶ The PSD is provided in the file **NOISE/iLIGOSensitivity.txt**
- ▶ It is a plain text file which can be read into Matlab
- ▶ First column is Frequency f (Hz) and second column is $\sqrt{S_n(f)}$

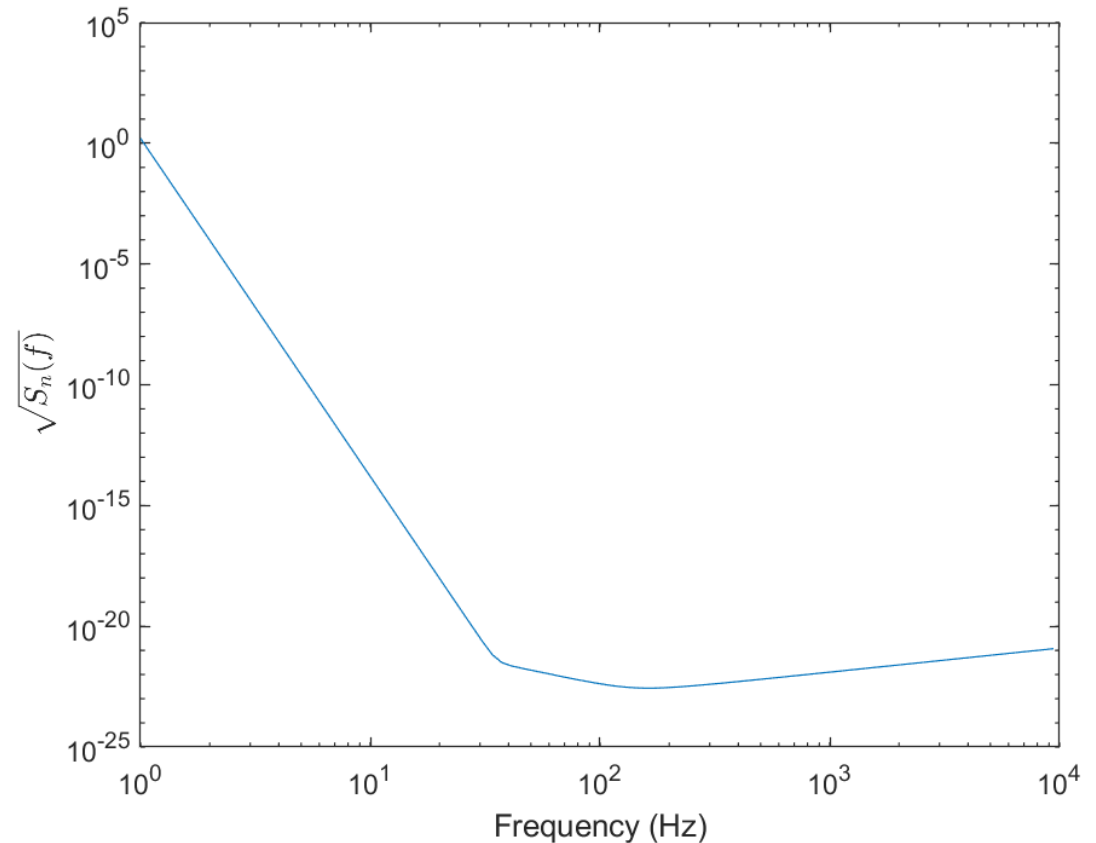
```
>> y = load('iLIGOSensitivity.txt', '-ascii');
```

```
>> whos y
```

Name	Size	Bytes	Class
------	------	-------	-------

y	97x2	1552	double
---	------	------	--------

```
>> loglog(y(:,1), y(:,2))
```



Initial LIGO design sensitivity

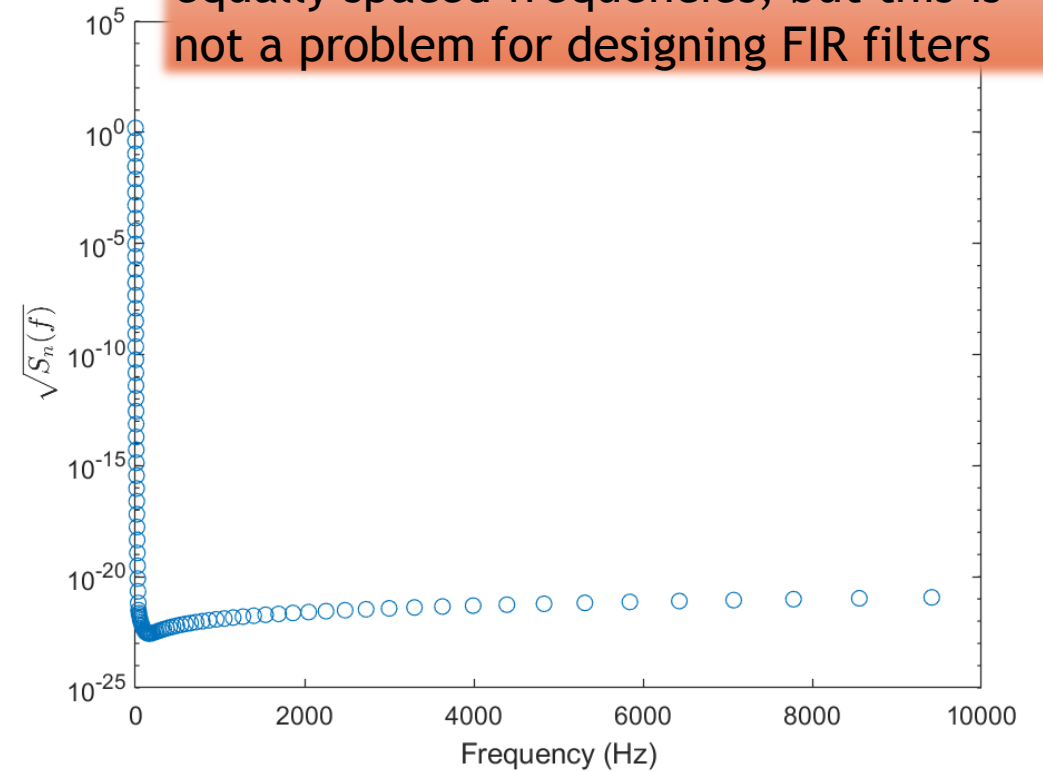
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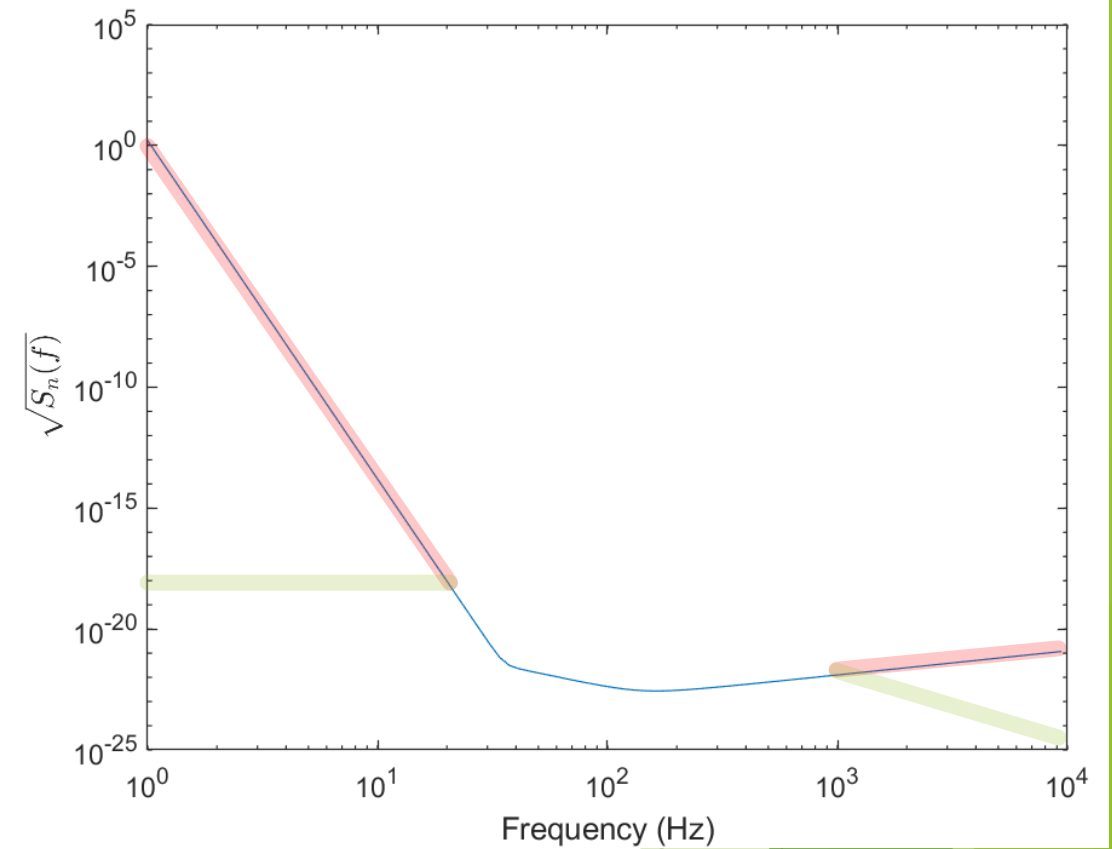
Name	Size	Bytes	Class
y	97x2	1552	double

```
>> loglog(y(:,1), y(:,2))
```



Modifications

- ▶ In any data analysis method, the low and high frequency parts will be filtered out \Rightarrow the PSD of the simulated noise need not match the design PSD in those parts
- ▶ The order of an FIR filter that can reproduce the steeply rising seismic part in its transfer function will be very high \Rightarrow Making the PSD goes smoothly to zero or just be a constant in these parts will help Matlab in designing better filters



Modifications

- ▶ For $f \leq 50$ Hz: $S_n(f) \rightarrow S_n(f = 50)$
- ▶ For $f \geq 700$ Hz: $S_n(f) \rightarrow S_n(f = 700)$
 - ▶ 700 Hz is where the inspiral phase of a binary of double neutron star will terminate
 - ▶ No point in keeping noise above this frequency in the data
- ▶ Remember that you need normalized frequencies of 0 and 1 for input to FIR1
 - ▶ Add $f = 0, S_n(f = 0)$ and $\frac{f_s}{2}, S_n\left(f = \frac{f_s}{2}\right)$ to the list if these are missing
- ▶ **Task:** Use the supplied codes (**NOISE/colGaussNoiseDemo.m**, **statgaussnoisegen.m**) to produce a simulated LIGO noise realization and estimate the PSD

Example of simulated LIGO noise PSD

- Note that the LIGO design sensitivity plots are logarithmic in frequency while the plot here is on a linear scale
 - Also, different truncation and sampling frequencies were used here
- The “bumpiness” in the PSD near the minimum is an artifact of the approximation inherent in filter design
 - You should try to minimize such artifacts by choosing the design parameters appropriately.

