

**Radiative Processes in Astrophysics / Problem Set #1**  
**Due February 15, 2021**

Pick **two** of the problems below to answer. Submit as a PDF by email!

1. Show that for an optically thin cloud around a source with mass  $M$  and luminosity  $L$  (integrated over frequency), the condition that the luminosity drives away the cloud through radiation pressure is  $L > 4\pi GMc/\kappa$ , where  $\kappa$  is the opacity (i.e. absorption coefficient per unit mass, assumed constant with frequency).
2. Sunspots on the Sun have a temperature of  $\sim 4000$  K, relative to the typical location on the Sun, which has a temperature of  $\sim 5500$  K. What is the ratio of the specific intensity integrated over frequency ( $\int d\nu I_\nu$ ) that we should observe in the location of a Sunspot relative to a typical location on the solar disk?
3. Light that reaches us from the Sun's "limb" (the area near its apparent angular edge) has a lower specific intensity than light that reaches us from its angular center. Explain why. Do you predict that the spectrum as a function of frequency differs in shape from center to edge, and if so how?
4. Given that for redshifted light  $\nu_o = \nu_e/(1+z)$  show that  $I_\nu \nu^{-3}$  is a constant.
5. In cosmology the angular diameter distance is defined as  $D_A = s/\theta$ , where  $s$  is the physical size of an object and  $\theta$  is its angular size, and the luminosity distance is defined as  $D_L = \sqrt{L/4\pi f}$ , where  $L$  is the total luminosity of an object (integrated over all frequencies) and  $f$  is its measured flux. Using only the fact that  $I_\nu \nu^{-3}$  is a constant show that  $D_L/D_A = (1+z)^2$ .