

Area and error calculations for the HEK

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1 Event_C1Error, Event_C2Error (in helio-projective cartesian coordinates)

$$\text{Event_C1Error} = W \cdot \frac{180^\circ}{\pi D_\odot} \cdot \left(1 + \frac{\Delta D_\odot}{D_\odot} |x - x_{\text{center}}|\right),$$

$$\text{Event_C2Error} = W \cdot \frac{180^\circ}{\pi D_\odot} \cdot \left(1 + \frac{\Delta D_\odot}{D_\odot} |y - y_{\text{center}}|\right),$$

where (x, y) are pixel coordinates, D_\odot is the distance observer-Sun in Mm, ΔD_\odot is its error estimate in Mm, and W is the *number of Mm in 1 pixel length*. From http://en.wikipedia.org/wiki/Astronomical_unit, we find

$$D_\odot \simeq 149\,597.871 \text{ Mm}.$$

Using the Earth orbit's eccentricity of $e \simeq 0.0167$, the yearly maximal error $\Delta D_\odot \simeq 149\,597.871 \times 0.0167 \text{ Mm}$. Note that $\frac{\Delta D_\odot}{D_\odot} \simeq 0.0167$.

2 Area_Raw (in Mm²)

This is the area as it is seen in the image, hence projected from a 3-D ball onto a 2-D disk.

$$\text{Area_Raw} = \text{Event_Npixels} \cdot \left(\frac{R_0}{R}\right)^2,$$

where R_0 and R are the radius of the Sun expressed in Mm respectively pixels. In [1], we find $R_0 = 695.508 \text{ Mm}$ and $\Delta R_0 = 0.026 \text{ Mm}$. Note that R can be found in the FITS header of the image, and $\Delta R = 2 \text{ (pixels)}$.

3 Area_Uncert (in Mm²)

$$\text{Area_Uncert} = \left(\frac{R_0}{R}\right)^2 \cdot \left((\text{number of pixels in contour}) + 2 \cdot \text{Event_Npixels} \cdot \left(\frac{\Delta R_0}{R_0} + \frac{\Delta R}{R} \right) \right).$$

4 Area_AtDiskCenter (in Mm²)

This is the deprojected area, i.e., the area on the 3-D ball.

$$\text{Area_AtDiskCenter} = \left(\frac{R_0}{R}\right)^2 \cdot \sum_{\text{pixels } (x_i, y_i)} \frac{R}{\sqrt{R^2 - (x_i - x_{\text{center}})^2 - (y_i - y_{\text{center}})^2}}.$$

Note that Area_AtDiskCenter will go to infinity as the pixel goes toward the eastern or western limb. In order to avoid this, I propose to adopt Paul Higgins's approach to limit the correction factor (i.e., the terms of the sum) to 16, which corresponds to a LOS angle of $\simeq 86^\circ$. For AR containing a pixel with a corresponding term exceeding 16, we don't provide Area_AtDiskCenter.

5 Area_AtDiskCenterUncert (in Mm²)

$$\begin{aligned} \text{Area_AtDiskCenterUncert} = & \left(\frac{R_0}{R}\right)^2 \cdot \sum_{\text{pixels } (x_i, y_i) \in \text{boundary AR}} \frac{R}{\sqrt{R^2 - (x_i - x_{\text{center}})^2 - (y_i - y_{\text{center}})^2}} \\ & + \left(\frac{R_0^2}{R}\right) \cdot \sum_{\text{pixels } (x_i, y_i) \in \text{AR}} \\ & \left(\frac{|x_i - x_{\text{center}}| + |y_i - y_{\text{center}}| + 2 \frac{\Delta R_0}{R_0} \cdot (R^2 - (x_i - x_{\text{center}})^2 - (y_i - y_{\text{center}})^2) + \frac{\Delta R}{R} \cdot (2R^2 - (x_i - x_{\text{center}})^2 - (y_i - y_{\text{center}})^2)}{(R^2 - (x_i - x_{\text{center}})^2 - (y_i - y_{\text{center}})^2)^{1.5}} \right). \end{aligned}$$

References

- [1] T.M. Brown, J. Christensen-Dalsgaard, "Accurate determination of the solar photosphere radius", *ApJ*, 500:L195–L198, 1998.