### I. SKYRME INTERACTION - FINITE TEMPERATURE

#### A. Potential Matrix Element

Interaction Matrix:

$$V_{ij} = t_0 (1 + x_0 P_\sigma) \delta(\mathbf{r}_{ij}) + \frac{1}{2} t_1 (1 + x_1 P_\sigma) \frac{1}{\hbar^2} [\overleftarrow{p}^{\dagger}_{ij} \delta(\mathbf{r}_{ij}) + \delta(\mathbf{r}_{ij}) \overrightarrow{p}_{ij}^2]$$

$$+ t_2 (1 + x_2 P_\sigma) \frac{1}{\hbar^2} \overleftarrow{\mathbf{p}}^{\dagger}_{ij} \cdot \delta(\mathbf{r}_{ij}) \overrightarrow{\mathbf{p}}_{ij} + \frac{1}{6} t_3 (1 + x_3 P_\sigma) \rho^\alpha(\mathbf{r}) \delta(\mathbf{r}_{ij})$$

$$+ \frac{i}{\hbar^2} W_0(\sigma_i + \sigma_j) \cdot \overleftarrow{\mathbf{p}}^{\dagger}_{ij} \times \delta(\mathbf{r}_{ij}) \overrightarrow{\mathbf{p}}_{ij}$$

$$+ \frac{1}{4} t_4 (1 + x_4 P_\sigma) \frac{1}{\hbar^2} [\overleftarrow{\mathbf{p}}^{\dagger}_{ij} \rho^\beta(\mathbf{r}) \delta(\mathbf{r}_{ij}) + \delta(\mathbf{r}_{ij}) \rho^\beta(\mathbf{r}) \overrightarrow{p}_{ij}^2]$$

$$+ t_5 (1 + x_5 P_\sigma) \frac{1}{\hbar^2} \overleftarrow{\mathbf{p}}_{ij} \cdot \rho^\gamma(\mathbf{r}) \delta(\mathbf{r}_{ij}) \overrightarrow{\mathbf{p}}_{ij}$$

$$(1)$$

where,  $\mathbf{r}_{ij} = \frac{\mathbf{r}_i - \mathbf{r}_j}{2}$ ,  $\mathbf{r} = \frac{\mathbf{r}_i + \mathbf{r}_j}{2}$ ,  $P_{\sigma} = \frac{1}{2}(1 + \sigma_1 \cdot \sigma_2)$ ,  $\mathbf{p}_{ij} = -i\hbar \frac{\nabla_i - \nabla_j}{2}$ ,  $\rho(\mathbf{r}) = \rho_p(\mathbf{r}) + \rho_n(\mathbf{r})$  and the arrows show the direction on which the momentum operators act.

# B. Energy Density

Assume time reversal invariance. The Energy of the ground state can be written as integration over an energy density functional,  $E_{HFB} = \int d^r \varepsilon_{HFB}$ 

$$\varepsilon_{HFB} = \sum_{t=n,p} \frac{\hbar^{2}}{2M_{q}} \tau_{q} + \frac{1}{2} t_{0} \left[ (1 + \frac{1}{2} x_{0}) \rho^{2} - (\frac{1}{2} + x_{0}) \sum_{q=n,p} \rho_{q}^{2} \right] \\
+ \frac{1}{4} t_{1} \left[ (1 + \frac{1}{2} x_{1}) (\rho \tau + \frac{3}{4} (\nabla \rho)^{2}) - (\frac{1}{2} + x_{1}) \sum_{q=n,p} (\rho_{q} \tau_{q} + \frac{3}{4} (\nabla \rho_{q})^{2}) \right] \\
+ \frac{1}{4} t_{2} \left[ (1 + \frac{1}{2} x_{2}) (\rho \tau + \frac{3}{4} (\nabla \rho)^{2}) + (\frac{1}{2} + x_{2}) \sum_{q=n,p} (\rho_{q} \tau_{q} - \frac{1}{4} (\nabla \rho_{q})^{2}) \right] \\
+ \frac{1}{12} t_{3} \rho^{\alpha} \left[ (1 + \frac{1}{2} x_{3}) \rho^{2} - (\frac{1}{2} + x_{3}) \sum_{q=n,p} \rho_{q}^{2} \right] \\
+ \frac{1}{4} t_{4} \left[ (1 + \frac{1}{2} x_{4}) (\rho \tau + \frac{3}{4} (\nabla \rho)^{2}) - (\frac{1}{2} + x_{4}) \sum_{q=n,p} (\rho_{q} \tau_{q} + \frac{3}{4} (\nabla \rho_{q})^{2}) \right] \rho^{\beta} \\
+ \frac{\beta}{8} t_{4} \left[ (1 + \frac{1}{2} x_{4}) \rho (\nabla \rho)^{2} - (\frac{1}{2} + x_{4}) \nabla \rho \cdot \sum_{q=n,p} \rho_{q} \nabla \rho_{q} \right] \rho^{\beta-1} \\
+ \frac{1}{4} t_{5} \left[ (1 + \frac{1}{2} x_{5}) (\rho \tau + \frac{3}{4} (\nabla \rho)^{2}) + (\frac{1}{2} + x_{5}) \sum_{q=n,p} (\rho_{q} \tau_{q} - \frac{1}{4} (\nabla \rho_{q})^{2}) \right] \rho^{\gamma} \\
- \frac{1}{16} (t_{1} x_{1} + t_{2} x_{2}) J^{2} + \frac{1}{16} (t_{1} - t_{2}) \sum_{q=n,p} J_{q}^{2} \\
- \frac{1}{16} (t_{4} x_{4} \rho^{\beta} + t_{5} x_{5} \rho^{\gamma}) J^{2} + \frac{1}{16} (t_{4} \rho^{\beta} - t_{5} \rho^{\beta}) \sum_{q=n,p} J_{q}^{2} \\
+ \frac{1}{2} W_{0} (\mathbf{J} \cdot \nabla \rho + \sum_{q=n,p} \mathbf{J}_{q} \cdot \nabla \rho_{q})$$

where,

$$\rho = 2 \int \frac{d^3k}{(2\pi\hbar)^3} n(k)$$

$$\tau = 2 \int \frac{d^3k}{(2\pi\hbar)^3} k^2 n(k)$$

$$\mathbf{J} = \int \frac{d^3k}{(2\pi\hbar)^3} \mathbf{k} \times \sum_{s,s'} \langle s|\boldsymbol{\sigma}|s'\rangle \ n(k)$$
(3)

The different terms can be grouped together in simpler notation:

$$\varepsilon = \sum_{t=n,p} \frac{\hbar^{2}}{2M_{q}} \tau_{q} + \frac{1}{4} t_{0} \Big[ (2 + x_{0})\rho^{2} - (1 + 2x_{0}) \sum_{q=n,p} \rho_{q}^{2} \Big] \\
+ \frac{1}{8} \Big[ a(\rho\tau + \frac{3}{4}(\nabla\rho)^{2}) + 2b \sum_{q=n,p} (\rho_{q}\tau_{q} + \frac{3}{4}(\nabla\rho_{q})^{2}) \Big] \\
+ \frac{1}{24} t_{3} \rho^{\alpha} \Big[ (2 + x_{3})\rho^{2} - (1 + 2x_{3}) \sum_{q=n,p} \rho_{q}^{2} \Big] \\
+ \frac{1}{8} t_{4} \Big[ (2 + x_{4})(\rho\tau + \frac{3}{4}(\nabla\rho)^{2}) - (1 + 2x_{4}) \sum_{q=n,p} (\rho_{q}\tau_{q} + \frac{3}{4}(\nabla\rho_{q})^{2}) \Big] \rho^{\beta} \\
+ \frac{\beta}{16} t_{4} \Big[ (2 + x_{4})\rho(\nabla\rho)^{2} - (1 + 2x_{4})\nabla\rho \cdot \sum_{q=n,p} \rho_{q}\nabla\rho_{q} \Big] \rho^{\beta-1} \\
+ \frac{1}{8} t_{5} \Big[ (2 + x_{5})(\rho\tau + \frac{3}{4}(\nabla\rho)^{2}) + (1 + 2x_{5}) \sum_{q=n,p} (\rho_{q}\tau_{q} - \frac{1}{4}(\nabla\rho_{q})^{2}) \Big] \rho^{\gamma} \\
- \frac{1}{16} (t_{1}x_{1} + t_{2}x_{2})J^{2} + \frac{1}{16} (t_{1} - t_{2}) \sum_{q=n,p} J_{q}^{2} \\
- \frac{1}{16} (t_{4}x_{4}\rho^{\beta} + t_{5}x_{5}\rho^{\gamma})J^{2} + \frac{1}{16} (t_{4}\rho^{\beta} - t_{5}\rho^{\beta}) \sum_{q=n,p} J_{q}^{2} \\
+ \frac{1}{2} W_{0}(\mathbf{J} \cdot \nabla\rho + \sum_{q=n,p} \mathbf{J}_{q} \cdot \nabla\rho_{q})$$

where,  $a = t_1(x_1 + 2) + t_2(x_2 + 2)$ ,  $b = \frac{1}{2}[t_2(2x_2 + 1) - t_1(2x_1 + 1)]$ .

In uniform matter  $\nabla \rho = 0$ :

$$\varepsilon = \sum_{t=n,p} \frac{\hbar^2}{2M_q} \tau_q + \frac{1}{4} t_0 \Big[ (2+x_0)\rho^2 - (1+2x_0) \sum_{q=n,p} \rho_q^2 \Big] 
+ \frac{1}{8} \Big[ a\rho\tau + 2b \sum_{q=n,p} \rho_q \tau_q \Big] 
+ \frac{1}{24} t_3 \rho^\alpha \Big[ (2+x_3)\rho^2 - (1+2x_3) \sum_{q=n,p} \rho_q^2 \Big] 
+ \frac{1}{8} t_4 \Big[ (2+x_4)\rho\tau - (1+2x_4) \sum_{q=n,p} \rho_q \tau_q \Big] \rho^\beta 
+ \frac{1}{8} t_5 \Big[ (2+x_5)\rho\tau + (1+2x_5) \sum_{q=n,p} \rho_q \tau_q \Big] \rho^\gamma 
- \frac{1}{16} (t_1 x_1 + t_2 x_2) J^2 + \frac{1}{16} (t_1 - t_2) \sum_{q=n,p} J_q^2 
- \frac{1}{16} (t_4 x_4 \rho^\beta + t_5 x_5 \rho^\gamma) J^2 + \frac{1}{16} (t_4 \rho^\beta - t_5 \rho^\beta) \sum_{q=n,p} J_q^2$$

In unpolarized matter,  $\mathbf{J} = 0$ :

$$\varepsilon = \sum_{t=n,p} \frac{\hbar^2}{2M_q} \tau_q + \frac{1}{4} t_0 \Big[ (2+x_0)\rho^2 - (1+2x_0) \sum_{q=n,p} \rho_q^2 \Big]$$

$$+ \frac{1}{8} \Big[ a\rho\tau + 2b \sum_{q=n,p} \rho_q \tau_q \Big]$$

$$+ \frac{1}{24} t_3 \rho^\alpha \Big[ (2+x_3)\rho^2 - (1+2x_3) \sum_{q=n,p} \rho_q^2 \Big]$$

$$+ \frac{1}{8} t_4 \Big[ (2+x_4)\rho\tau - (1+2x_4) \sum_{q=n,p} \rho_q \tau_q \Big] \rho^\beta$$

$$+ \frac{1}{8} t_5 \Big[ (2+x_5)\rho\tau + (1+2x_5) \sum_{q=n,p} \rho_q \tau_q \Big] \rho^\gamma$$
(6)

Energy per bayon,  $\mathcal{E} \equiv \varepsilon/rho$ :

$$\mathcal{E} = \sum_{t=n,p} \frac{\hbar^2}{2M_q} \frac{\tau_q}{\rho} + \frac{1}{4} t_0 \Big[ (2+x_0)\rho - (1+2x_0) \sum_{q=n,p} \frac{\rho_q^2}{\rho} \Big]$$

$$+ \frac{1}{8} \Big[ a\tau + 2b \sum_{q=n,p} \frac{\rho_q \tau_q}{\rho} \Big]$$

$$+ \frac{1}{24} t_3 \rho^{\alpha} \Big[ (2+x_3)\rho - (1+2x_3) \sum_{q=n,p} \frac{\rho_q^2}{\rho} \Big]$$

$$+ \frac{1}{8} t_4 \Big[ (2+x_4)\tau - (1+2x_4) \sum_{q=n,p} \frac{\rho_q \tau_q}{\rho} \Big] \rho^{\beta}$$

$$+ \frac{1}{8} t_5 \Big[ (2+x_5)\tau + (1+2x_5) \sum_{q=n,p} \frac{\rho_q \tau_q}{\rho} \Big] \rho^{\gamma}$$

$$(7)$$

In terms of proton fraction,  $y = \frac{\rho_p}{\rho_p + \rho_n}$ :

$$\mathcal{E} = \sum_{t=n,p} \frac{\hbar^2}{2M_q} \frac{\tau_q}{\rho} + \frac{1}{4} t_0 \Big[ (2+x_0) - (1+2x_0) [y^2 + (1-y)^2] \Big] \rho$$

$$+ \frac{1}{8} \Big[ a\tau + 2b [y\tau_p + (1-y)\tau_n] \Big]$$

$$+ \frac{1}{24} t_3 \rho^{\alpha+1} \Big[ (2+x_3) - (1+2x_3) [y^2 + (1-y)^2] \Big]$$

$$+ \frac{1}{8} t_4 \Big[ (2+x_4)\tau - (1+2x_4) [y\tau_p + (1-y)\tau_n] \Big] \rho^{\beta}$$

$$+ \frac{1}{8} t_5 \Big[ (2+x_5)\tau + (1+2x_5) [y\tau_p + (1-y)\tau_n] \Big] \rho^{\gamma}$$
(8)

## II. SINGLE PARTICLE PROPERTIES

From the energy density the single particle spectrum can be derived. By performing functional variation of the energy density with repsect to the single particle wavefunction, a modified Schrodinger equation can be derived:

$$\delta \varepsilon_i = \left[ \frac{\delta \varepsilon_i}{\delta \tau_i} + \frac{\delta \varepsilon_i}{\delta \rho_i} \right] \delta \phi_i = \epsilon_i \delta \phi_i \tag{9}$$

Since the Skyrme potential is at mostly quadratic in momenta with nonlinear density dependence, its effect is exactly included by effective mass and mean field shift(residual interaction), both density dependent:

$$\epsilon_{i}(k) = \frac{\hbar^{2}k^{2}}{2M_{i}^{*}} + U_{i}$$

$$\frac{\hbar^{2}}{2M_{q}^{*}} \equiv \frac{\partial \varepsilon}{\partial \tau_{q}}$$

$$U_{i} \equiv \frac{\partial \varepsilon}{\partial \rho_{i}}$$
(10)

From eq. 6 the effective baryon masses:

$$M_{p}^{*}/M = \left\{1 + \frac{M \rho}{4\hbar^{2}} \left[a + 2 y b + t_{4}[(2 + x_{4}) - (1 + 2x_{4}) y]\rho^{\beta} + t_{5}[2 + x_{5} + (1 + 2x_{5}) y]\rho^{\gamma}\right]\right\}^{-1}$$

$$M_{n}^{*}/M = \left\{1 + \frac{M \rho}{4\hbar^{2}} \left[a + 2 (1 - y) b + t_{4}[(2 + x_{4}) - (1 + 2x_{4}) (1 - y)]\rho^{\beta} + t_{5}[2 + x_{5} + (1 + 2x_{5}) (1 - y)]\rho^{\gamma}\right]\right\}^{-1}$$

$$(11)$$

and the residual potentials:

$$U_{p} = \frac{1}{8} (2b \tau_{p} + a \tau) + \frac{1}{2} t_{0} [(2 + x_{0}) - (1 + 2x_{0}) y] \rho$$

$$+ \frac{1}{24} t_{3} \Big[ 4 + \alpha - 2y(1 - (1 - y)\alpha) + x_{3}(1 - 2y)[2 - (1 - 2y)\alpha] \Big] \rho^{\alpha+1}$$

$$U_{n} = \frac{1}{8} (2b \tau_{n} + a \tau) + \frac{1}{2} t_{0} [(1 - x_{0}) + (1 + 2x_{0})y] \rho$$

$$+ \frac{1}{24} t_{3} \Big[ 2 + \alpha + 2y(1 + \alpha - y\alpha) - x_{3}(1 - 2y)[2 + (1 - 2y)\alpha] \Big] \rho^{\alpha+1}$$

$$(12)$$

#### III. T=0 DFT

At T=0, there are simple relation that can be drawn between the 2 integrations in Fourier space since the occuppation number is a step function:

$$\rho_{q} = 2 \int \frac{d^{3}k}{(2\pi\hbar)^{3}} \theta(k_{F,q} - k) = \frac{k_{F,q}^{3}}{3\pi^{2}\hbar^{3}}$$

$$\tau_{q} = 2 \int \frac{d^{3}k}{(2\pi\hbar)^{3}} (k/\hbar)^{2} \theta(k_{F,q} - k) = \frac{k_{F,q}^{5}}{5\pi^{2}\hbar^{5}} \to$$

$$\tau = \frac{3}{5} (3\pi^{2})^{2/3} \rho_{q}^{5/3}, H_{n}(y) = 2^{n-1} [y^{n} + (1-y)^{n}], y = \rho_{p}/\rho$$

$$\tau = \tau_{p} + \tau_{n} = \frac{3}{5} (\frac{3\pi^{2}}{2})^{2/3} H_{5/3}(y) \rho^{5/3}$$
(13)

So,

$$\mathcal{E}_{0} = \frac{3\hbar^{2}}{10M} \left(\frac{3\pi^{2}}{2}\right)^{2/3} \rho^{2/3} H_{5/3}(y) + \frac{1}{8} t_{0} \rho \left[2(2+x_{0}) - (1+2x_{0})H_{2}(y)\right]$$

$$+ \frac{3}{40} \left(\frac{3\pi^{2}}{2}\right)^{2/3} \left[aH_{5/3}(y) + bH_{8/5}(y)\right] \rho^{5/3}$$

$$+ \frac{1}{48} t_{3} \rho^{\alpha+1} \left[2(2+x_{3}) - (1+2x_{3})H_{2}(y)\right]$$

$$+ \frac{3}{40} \left(\frac{3\pi^{2}}{2}\right)^{2/3} t_{4} \rho^{\beta+5/3} \left[(2+x_{4})H_{5/3}(y) - (\frac{1}{2}+x_{4})H_{8/3}(y)\right]$$

$$+ \frac{3}{40} \left(\frac{3\pi^{2}}{2}\right)^{2/3} t_{5} \rho^{\gamma+5/3} \left[(2+x_{5})H_{5/3}(y) + (\frac{1}{2}+x_{5})H_{8/3}(y)\right]$$

$$+ \frac{3}{40} \left(\frac{3\pi^{2}}{2}\right)^{2/3} t_{5} \rho^{\gamma+5/3} \left[(2+x_{5})H_{5/3}(y) + (\frac{1}{2}+x_{5})H_{8/3}(y)\right]$$

with  $H_n(y) = 2^{n-1}[y^n + (1-y)^n]$ . A common choice is to set  $M = 1/2(M_n + M_p)$ , or use the individual value for each species. In compact notation,

$$\mathcal{E}_0 = C(y)\rho^{2/3} + A(y)\rho + B(y)\rho^{\alpha+1} + D(y)\rho^{5/3} + G(y)\rho^{\beta+5/3} + K(y)\rho^{\gamma+5/3}$$
(15)

By comparing the 2 expressions, the following relations can be easily deduced:

$$C(y) = \frac{3\hbar^2}{10M_y} \left(\frac{3\pi^2}{2}\right)^{2/3} H_{5/3}(y)$$

$$A(y) = \frac{1}{8} t_0 \left[ 2(2+x_0) - (1+2x_0)H_2(y) \right]$$

$$B(y) = \frac{1}{48} t_3 \left[ 2(2+x_3) - (1+2x_3)H_2(y) \right]$$

$$D(y) = \frac{3}{40} \left(\frac{3\pi^2}{2}\right)^{2/3} \left[ aH_{5/3}(y) + bH_{8/5}(y) \right]$$

$$G(y) = \frac{3}{40} \left(\frac{3\pi^2}{2}\right)^{2/3} t_4 \left[ (2+x_4)H_{5/3}(y) - (\frac{1}{2}+x_4)H_{8/3}(y) \right]$$

$$K(y) = \frac{3}{40} \left(\frac{3\pi^2}{2}\right)^{2/3} t_5 \left[ (2+x_5)H_{5/3}(y) + (\frac{1}{2}+x_5)H_{8/3}(y) \right]$$

For pure neutron matter and symmetric matter the kinetic coefficient is

$$C_n = C(0) = \frac{3\hbar^2}{10M_n} (\frac{3\pi^2}{2})^{2/3} H_{5/3}(0) = 118.995 \text{ Mev fm}^2 \approx 119 \text{ Mev fm}^2$$

$$C_{sym} = C(1/2) = \frac{3\hbar^2}{5(M_n + M_p)} (\frac{3\pi^2}{2})^{2/3} H_{5/3}(1/2) = 75.0139 \text{ Mev fm}^2 \approx 75 \text{ Mev fm}^2$$
(17)

The effective mass is due to the terms dependent on kinetic energy:

$$\tau^{T=0}(\rho,y) \equiv \frac{3\hbar^{2}}{10M^{*}} \left(\frac{3\pi^{2}}{2}\right)^{2/3} \rho^{2/3} H_{5/3}(y)$$

$$= \frac{3\hbar^{2}}{10M} \left(\frac{3\pi^{2}}{2}\right)^{2/3} \rho^{2/3} H_{5/3}(y) + \frac{3}{40} \left(\frac{3\pi^{2}}{2}\right)^{2/3} \left[aH_{5/3}(y) + bH_{8/5}(y)\right] \rho^{5/3}$$

$$+ \frac{3}{40} \left(\frac{3\pi^{2}}{2}\right)^{2/3} t_{4} \rho^{\beta+5/3} \left[(2+x_{4})H_{5/3}(y) - \left(\frac{1}{2}+x_{4}\right)H_{8/3}(y)\right]$$

$$+ \frac{3}{40} \left(\frac{3\pi^{2}}{2}\right)^{2/3} t_{5} \rho^{\gamma+5/3} \left[(2+x_{5})H_{5/3}(y) + \left(\frac{1}{2}+x_{5}\right)H_{8/3}(y)\right]$$

$$\equiv C(y) \rho^{2/3} \left[1 + (D(y)\rho + G(y)\rho^{\beta+1} + K(y)\rho^{\gamma+1})/C(y)\right]$$
(18)

Thus,

$$M^*/M = \left\{1 + \frac{M}{4H_{5/3}(y)\hbar^2} \left[\rho[aH_{5/3}(y) + bH_{8/5}(y)] + \rho^{\beta+1}t_4[(2+x_4)H_{5/3}(y) - (\frac{1}{2}+x_4)H_{8/3}(y)]\right] + \rho^{\gamma+1}t_5[(2+x_5)H_{5/3}(y) - (\frac{1}{2}+x_5)H_{8/3}(y)]\right]\right\}^{-1}$$

$$= \frac{C(y)}{C(y) + D(y)\rho + G(y)\rho^{\beta+1} + K(y)\rho^{\gamma+1}}$$
(19)

Also, the thermodynamic pressure:

$$\mathcal{P}^{T=0}(\rho, y) = \rho^2 \frac{\partial \mathcal{E}^{T=0}}{\partial \rho}$$

$$= \frac{2}{3} C(y) \rho^{5/3} + A(y) \rho^2 + (\alpha + 1) B(y) \rho^{\alpha+1} + \frac{5}{3} D(y) \rho^{8/3} + \beta G(y) \rho^{\beta+8/3} + \gamma K(y) \rho^{\gamma+8/3}$$
(20)

## IV. LATTIMER - SCHWESTY NOTATION

The standard parametrization:

$$\varepsilon = \frac{\hbar^2 \tau_n}{2M_n} + \frac{\hbar^2 \tau_p}{2t M_p} + \frac{1}{8} [a(\rho_n + \rho_p)(\tau_n + \tau_p) + 2b(\rho_n \tau_n + \rho_p \tau_p)] + \frac{1}{4} t_0 [(2 + x_0)(\rho_n + \rho_p)^2 - (1 + 2x_0)(\rho_n^2 + \rho_p^2)] + \frac{1}{24} t_3 (\rho_n + \rho_p)^{\alpha} [(2 + x_3)(\rho_n + \rho_p)^2 - (1 + 2x_3)(\rho_n^2 + \rho_p^2)]$$
(21)

$$a = \frac{4\hbar^{2}}{M}(F+G)$$

$$b = -\frac{4\hbar^{2}}{M}G$$

$$t_{0} = \frac{8}{3}(A+B)$$

$$x_{0} = -\frac{1}{2}\frac{A-2B}{A+B}$$

$$t_{3} = 16(C+D)$$

$$x_{3} = -\frac{1}{2}\frac{C-D}{C+D}$$

$$\alpha = \delta - 1$$

$$a = \frac{4\hbar^{2}}{M}(F+G)$$

$$b = -\frac{4\hbar^{2}}{M}G$$

$$t_{0} = \frac{8}{3}(A+B)$$

$$x_{0} = -\frac{1}{2}\frac{A-2B}{A+B}$$

$$(23)$$

$$x_{1} = -\frac{1}{2}\frac{C-D}{C+D}$$

$$\alpha = \delta - 1$$

$$\varepsilon = 4B\rho_{n}\rho_{p} + A(\rho_{n} + \rho_{p})^{2} + (\rho_{n} + \rho_{p})^{\delta-1}[4D\rho_{n}\rho_{p} + C(\rho_{n} + \rho_{p})^{2}] + \frac{5(\frac{2}{3})^{2/3}}{3\pi^{4/3}}\alpha_{S}[M(\frac{\tau_{n}}{M_{n}} + \frac{\tau_{p}}{M_{p}}) + F(\rho_{n} + \rho_{p})(\tau_{n} + \tau_{p}) - G(\rho_{n} - \rho_{p})(\tau_{n} - \tau_{p})]$$
(24)

The standard parametrization for  $T=0, (M_n, M_p) \to M = \frac{1}{2}(M_n + M_p)$ :

$$\varepsilon = \frac{3\hbar^2}{10M} \left(\frac{3\pi^2}{2}\right)^{2/3} \rho^{5/3} H_{5/3}(y) + \frac{1}{8} t_0 \rho^2 \left[ 2(2+x_0) - (1+2x_0) H_2(y) \right]$$

$$+ \frac{3}{40} \left(\frac{3\pi^2}{2}\right)^{2/3} \left[ a H_{5/3}(y) + b H_{8/5}(y) \right] \rho^{8/3}$$

$$+ \frac{1}{48} t_3 \rho^{\alpha+2} \left[ 2(2+x_3) - (1+2x_3) H_2(y) \right]$$
(25)

By comparing the expression in these notes with the ones from Lattimer parametrization, the energy density is:

$$\varepsilon = \alpha_S \rho^{5/3} H_{5/3}(y) + [A + B(2 - H_2(y))] \rho^2 + [C + D(2 - H_2(y))] \rho \delta + \alpha_S \rho^{5/3} [(F + G)H_{5/3} - GH_{8/3}(y)]$$
(26)

where, 
$$\alpha_S = \frac{3\hbar^2}{10M} (\frac{3}{2}\pi^2)^{2/3}$$

#### A. Skyrme parametrization from Saturation Observables

Given the following set of physical observables,

$$E_{0} = \varepsilon \Big|_{\rho_{0},y=1/2} = (A+B)\rho_{0}^{2} + (C+D)\rho_{0}^{\delta+1} + \alpha_{S}\rho_{0}^{5/3}(1+F\rho_{0})$$

$$P = \rho^{2} \frac{d(\varepsilon/\rho)}{d\rho} \Big|_{\rho_{0},y=1/2} = \frac{2}{3}\alpha_{S}\rho_{0}^{5/3} + (A+B)\rho_{0}^{2} + \frac{5}{3}F\alpha_{S}\rho_{0}^{8/3} + (C+D)\delta\rho_{0}^{1+\delta} = 0$$

$$(M^{*}/M) = \frac{d\varepsilon}{d\tau} \Big|_{\rho_{0},y=1/2} = (1+F\rho_{0})^{-1}$$

$$K_{m} = 9\rho^{2} \frac{d^{2}(\varepsilon/\rho)}{d\rho^{2}} \Big|_{\rho_{0},y=1/2} = -2\alpha_{S}\rho_{0}^{2/3} + 10F\alpha_{S}\rho_{0}^{5/3} + 9(C+D)(\delta-1)\delta\rho_{0}^{\delta}$$

$$S = \frac{1}{8} \frac{d^{2}(\varepsilon/\rho)}{dy^{2}} \Big|_{\rho_{0},y=1/2} = \frac{5}{9}\alpha_{S}\rho_{0}^{2/3} - B\rho_{0} + \frac{5}{9}(F-3G)\alpha_{S}\rho_{0}^{5/3} - D\rho_{0}^{\delta}$$

$$L = 3\rho \frac{dS}{\rho} \Big|_{\rho_{0},y=1/2} = \frac{10}{9}\alpha_{S}\rho_{0}^{2/3} - 3B\rho_{0} + \frac{25}{9}(F-3G)\alpha_{S}\rho_{0}^{5/3} - 3D\delta\rho_{0}^{\delta}$$

$$K_{s} = 9\rho^{2} \frac{d^{2}S}{d\rho^{2}} \Big|_{\rho_{0},y=1/2} = -\frac{10}{9}\alpha_{S}\rho_{0}^{2/3} + \frac{50}{9}(F-3G)\alpha_{S}\rho_{0}^{5/3} - 9D(\delta-1)\delta\rho_{0}^{\delta}$$

the skyrme parameters can be found as follows,

$$F = \frac{(M^*/M)^{-1} - 1}{\rho_0}$$

$$\delta = \frac{K_m + 2\rho_0^{2/3}(1 - 5F\rho_0)\alpha_S}{3\rho_0^{2/3}(1 - 2F\rho_0)\alpha_S - 9E_0}$$

$$G = \frac{9K_S - 27(L - 3S)\delta + 5\rho_0^{2/3}\alpha_S[2 - 3\delta + 2F\rho_0(3\delta - 5)]}{30\rho_0^{5/3}\alpha_S(3\delta - 5)}$$

$$D = \frac{5(3L - 9S + \rho_0^{2/3}\alpha_S) - 3K_S}{9(5 - 8\delta + 3\delta^2)\rho_0^{\delta}}$$

$$C = \frac{\rho_0^{2/3}(1 - 2F\rho_0)\alpha_S - 3E_0}{3(\delta - 1)\rho_0^{\delta}} - D$$

$$B = \frac{L(6 + 9\delta) + 5(\rho_0^{2/3}\alpha_S(3\delta - 2) - 9S\delta) - 3K_S}{18\rho_0(\delta - 1)}$$

$$A = -\left[\frac{2}{3}\alpha_S\rho^{-1/3} + B + \frac{5}{3}F\alpha_S\rho^{2/3} + (C + D)\delta\rho^{\delta - 1}\right]$$

#### V. FINITE TEMPERATURE DFT

In order to obtain the relationship between density and chemical potential, the following set of coupled equations need to be solved self-consistenly:

$$f_{k} = \left[1 + e^{\left(\frac{k^{2}}{2M^{*}} + U - \mu\right)/T}\right]^{-1}$$

$$\rho = \sum_{s,is} \int \frac{d^{3}k}{(2\pi)^{3}} f_{k}$$

$$\tau = \sum_{s,is} \int \frac{d^{3}k}{(2\pi)^{3}} k^{2} f_{k}$$

$$E \equiv E(\rho, \tau)$$

$$(29)$$

where, a sum over all discrete quantum numbers is performed (spin and isospin). And from the energy density functional, the mean field parameters can be derived,

$$M^* = \frac{1}{2} (\frac{\delta E}{\delta \tau})^{-1}$$

$$U = \frac{\delta E}{\delta \rho}$$
(30)

The chemical potential can be found by inverting the expression for the density.

### VI. THERMODYNAMIC POTENTIALS

Since the effect of phenomenological mean field models can be incorporated into  $(M^*, U)$  which are density dependent for Skyrme, and also temperature dependent for RMF, the thermodynamic properties of assymetric matter at finite temperature can be expressed by fermi integrals of 'modified' non-interacting fermi gases. The single particle spectrum and dustribution function:

$$\xi_{i} = \frac{k^{2}}{2M_{i}^{*}} + U_{i}$$

$$F_{i} = \left\{ \exp\left[\frac{\xi_{i} - \mu_{i}}{T}\right] + 1 \right\}^{-1} = \left\{ \exp\left[\frac{\frac{k^{2}}{2M_{i}^{*}} - \eta_{i}}{T}\right] + 1 \right\}^{-1}, \ \overline{F} = 1 - F$$
(31)

The density and kinetic density:

$$\rho_i = \int_0^\infty \frac{dk}{\pi^2} k^2 F_i$$

$$\tau_i = \int_0^\infty \frac{dk}{\pi^2} k^4 F_i$$
(32)

The entropy density can be calculated from the distribution function:

$$S/V = -\int_{0}^{\infty} \frac{dk}{\pi^{2}} k^{2} \left[ F_{i} \ln(F_{i}) + (1 - F_{i}) \ln(1 - F_{i}) \right]$$

$$= -\int_{0}^{\infty} \frac{dk}{\pi^{2}} k^{2} F_{i} \ln\left(\frac{F_{i}}{1 - F_{i}}\right) - \int_{0}^{\infty} \frac{dk}{\pi^{2}} k^{2} \ln(1 - F_{i})$$

$$= -\int_{0}^{\infty} \frac{dk}{\pi^{2}} k^{2} F_{i} \ln\left(\exp\left[\frac{-\xi_{i} + \mu_{i}}{T}\right]\right) - \left[\frac{k^{3}}{3\pi^{2}} \ln(1 - F_{i})\right] \Big|_{0}^{\infty} + \int_{0}^{\infty} \frac{dk}{\pi^{2}} \frac{k^{3}}{3} \frac{k}{M_{i}^{*}} \frac{\exp\left[\frac{\xi_{i} - \mu_{i}}{T}\right]}{1 - F_{i}}$$

$$= \left[\frac{1}{2M_{i}^{*}T} + \frac{1}{3M_{i}^{*}T}\right] \int_{0}^{\infty} \frac{dk}{\pi^{2}} k^{4} F_{i} + \frac{U_{i} - \mu_{i}}{T} \int_{0}^{\infty} \frac{dk}{\pi^{2}} k^{2} F_{i}$$

$$= \frac{1}{T} \left[\frac{5\tau_{i}}{6M_{i}^{*}} + (U_{i} - \mu_{i})\rho\right]$$
(33)

From the first law of thermodynamics:

$$E_{i} = TS_{i} + \mu_{i}N_{i} - P_{i}V$$

$$P_{i} = \frac{TS_{i} + \mu_{i}N_{i} - E_{i}}{V}$$

$$= \frac{5\tau_{i}}{6M_{i}^{*}} + U_{i}\rho_{i} - \frac{E_{i}}{V}$$

$$(34)$$

## VII. THERMODYNAMIC DERIVATIVES

Let,

$$\alpha_1 = \frac{\hbar^2}{2M}(F - G), \alpha_2 = \frac{\hbar^2}{2M}(G + F)$$
 (35)

Then,

$$\frac{\hbar^2}{2M_i^*} = \frac{\hbar^2}{2M} + \alpha_1 n_i + \alpha_2 n_{-i} \tag{36}$$

where, i denotes the isospin value. From epxression above, the derivative of the effective mass in terms of the density can be found:

$$\partial_{n_i} M_r^* = -2 \frac{M_r^{*2}}{\hbar^2} \left[ \alpha_2 + (\alpha_1 - \alpha_2) \delta_{ir} \right]$$
(37)

The density and kinetic density can be expressed in terms of the general fermi integration:

$$F_{n}(\eta) = \int_{0}^{\infty} \frac{u^{n}}{e^{u-\eta} + 1} du, \quad \eta_{i} = (\mu_{i} - U_{i})/T, \quad U_{i} = \frac{\delta \mathcal{E}}{\delta n_{i}}$$

$$n_{i} = \frac{1}{2\pi^{2}} \left(\frac{2M_{r}^{*}T}{\hbar^{2}}\right)^{3/2} F_{1/2}(\eta_{i}) \leftrightarrow \eta_{i} = F_{1/2}^{-1} \left[2\pi^{2} n_{i} \left(\frac{\hbar^{2}}{2M_{r}^{*}T}\right)^{3/2}\right] = F_{1/2}^{-1}(u_{i})$$
(38)

And,

$$\partial_{n_{i}} u_{r} = 2\pi^{2} \left(\frac{\hbar^{2}}{2M_{r}^{*}T}\right)^{3/2} \left[\delta_{ir} + \frac{3n_{r}M_{r}^{*}}{\hbar^{2}} \left[\alpha_{2} + (\alpha_{1} - \alpha_{2})\delta_{ir}\right]\right]$$

$$= \frac{F_{1/2}(\eta_{r})}{n_{r}} \left[\delta_{ir} + \frac{3n_{r}M_{r}^{*}}{\hbar^{2}} \left[\alpha_{2} + (\alpha_{1} - \alpha_{2})\delta_{ir}\right]\right]$$
(39)

Also,  $\partial_{\eta} F_n(\eta) = n F_{n-1}(\eta) \leftrightarrow \partial_u F_n^{-1}(u) = \frac{1}{n F_{n-1}(u)}$ :

$$\partial_{n_{i}} \eta_{r} = d_{u} F_{1/2}^{-1}(u_{r}) \times \partial_{n_{i}} u_{r} 
= \frac{\partial u F_{1/2}(u_{r})}{n_{r} F_{-1/2}(\eta_{r})} \left[ \delta_{ir} + \frac{3n_{r} M_{r}^{*}}{\hbar^{2}} \left[ \alpha_{2} + (\alpha_{1} - \alpha_{2}) \delta_{ir} \right] \right] 
= \frac{2F_{1/2}(u_{r})}{n_{r} F_{-1/2}(\eta_{r})} \left[ \delta_{ir} + \frac{3n_{r} M_{r}^{*}}{\hbar^{2}} \left[ \alpha_{2} + (\alpha_{1} - \alpha_{2}) \delta_{ir} \right] \right] 
\equiv \frac{G}{n_{r}} \left[ \delta_{ir} + \frac{3n_{r} M_{r}^{*}}{\hbar^{2}} \left[ \alpha_{2} + (\alpha_{1} - \alpha_{2}) \delta_{ir} \right] \right]$$
(40)

Thus,

$$\partial_{n_i} \mu_r = T \partial_{n_i} \eta_r + \partial_{n_i} U_r$$

$$= \frac{TG}{n_r} \left[ \delta_{ir} + \frac{3n_r M_r^*}{\hbar^2} \left[ \alpha_2 + (\alpha_1 - \alpha_2) \delta_{ir} \right] \right] + \partial_{n_i} U_r$$
(41)