

Preparing the basic ingredients for trajectory design

# 2. ORBIT REPRESENTATION



### Lab 2.1: Cartesian to Keplerian

- Write a function car2kep.m that takes as input [state = (x,y,z,vx,vy,vz)] and returns as output: [kep=(a,e,i,Om,om,theta)]
- Write a function kep2car.m that takes as input [kep=(a,e,i,Om,om,theta)] and returns as output: [state = (x,y,z,vx,vy,vz)]
- 3. Write a function plotOrbit.m that plot an orbit given 5 orbital elements: kep=(a,e,i,Om,om,theta) and returns the handle to the figure

NB: Also mu is an input to the function. For testing the function write a script so that all the input you use are stored.



### Lab 2.2: Keplerian orbit

- 1. Numerically integrate with ode45 the equation  $\ddot{\mathbf{r}} + \frac{\mu}{r^3}\mathbf{r} = 0$  function dyn\_2BP.m
  - Initial conditions: positions and velocity in LEO at altitude of 500 km, inclination = 30 degrees, RAAN = 40 degrees, omega = 0 degrees, e = 0.
- 2. Plot r(t), v(t)
- 3. Demonstrate with a graph that h(t) and e(t) are constant in magnitude and direction and that e(t) is perpendicular to h(t)
- 4. Convert r(t) and v(t) to kep(t)
- Plot a(t), e(t), i(t), Om(t), om(t), theta(t)
- Plot vr(theta) and vtheta(theta)

NB: declare all initial parameters and numerical constants at the beginning of the script, use scripts!



integration time  $(n_T \times 1 \text{ array})$ 

[T,X]=ode45(@dynamics,tspan,X0)

initial state (1 x n<sub>x</sub> array)

state evolution in time  $(n_T \times n_X)$ 

time span for integration

- $[T_0 T_f]$
- $[T_0 T_1 ... T_i ... T_f]$  (1 x n<sub>T</sub> array)

function [X<sub>d</sub>]=dynamics(t,X)

•••

< dynamics written as system of first order ode >

•••

Xd=[...]'; % column vector

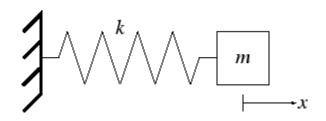
n<sub>X</sub> = length of state vector

 $n_T$  = length of integration time array



#### **Example:** harmonic oscillator

**Equation of motion** 



$$m\ddot{x} + kx = 0$$

$$\ddot{x} + \omega_0^2 x = 0$$

$$\omega_0 = \sqrt{\frac{k}{m}}$$

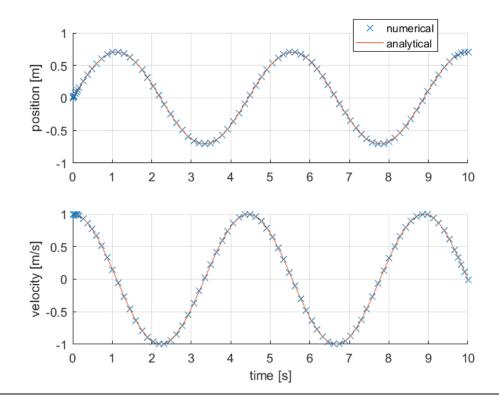
Analytical solution

$$x(t) = \frac{\dot{x}_0}{\omega_0} \sin(\omega_0 t) + x_0 \cos(\omega_0 t)$$

Compare analytical solution with numerical integrated one



- Write the equation of motion (second order differential equation) into a system of ode into a function
- 2. Call the function using ode45
- 3. Plot the results and compare with analytical solution





#### <u>Astrodynamics problems</u>

Require high accuracy and more stringent tolerances

Default values:

Abstol = 1e-6

[T,X]=ode45(@dynamics,tspan,X0,options)

[T,X]=ode113(@dynamics,tspan,X0,options)

NB: Sometimes ode113 works better than ode45 for astrodynamics problems



### Important notes

- In dyn\_2BP.m use mu as an external parameter
   [T,X]=ode45(@dynamics,tspan,X0,options,parameters)
   [X<sub>d</sub>]=dynamics(t,X,parameters)
- Use vector notation in Matlab r=X(1:3); v=X(4:6)
- At the beginning of the main script use clc; clear; (clear allows using debug mode)
- Use debug mode to validate your code
- help function to read the help of a Matlab function

