

Gammapy v2.0: New features and plans

Gammapy dev team: gammapy-ld-1@in2p3.fr

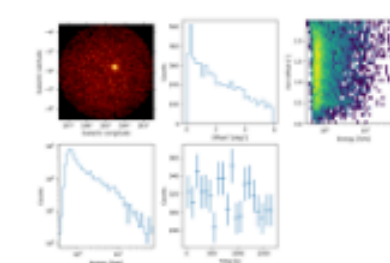
Development cycle since 1.0

- ❖ Gammapy v2.0 released on 26/08/2025
 - ❖ Will be maintained for the next development cycle (~ 2yrs)
 - ❖ *No more bug fixes for 1.x versions*
- ❖ Almost 3 yrs since the last LTS
 - ❖ Gammapy v1.0: 10/11/2022
 - ❖ 3 major releases, and 2 bug fixes
 - ❖ ~ 31 new contributors (~ 10 very regular developers)
 - ❖ Presented at ~11 major conferences, ~ 15 schools/hand-on sessions
- ❖ Many new features:
 - ❖ Analysis validated for most gamma-ray telescopes: CTAO, VERITAS, H.E.S.S., MAGIC, LST1, HAWC, SWGO, Fermi-LAT
 - ❖ Separation of Internal Data Model from the GADF
 - ❖ Performance improvements with multi-processing
 - ❖ Meta-data handling
 - ❖ Enhanced multi-instrument joint likelihood
 - ❖ Improved timing analysis
 - ❖ Addition of parameter priors
 - ❖ Estimator for energy dependent morphology

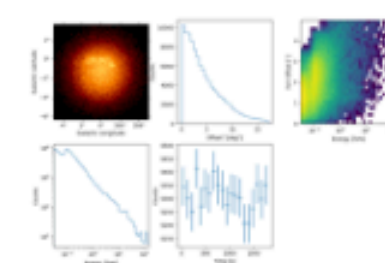
Project Manager: B. Khelifi
Lead dev: A. Donath, R. Terrier, Q. Remy, A. Sinha
Documentation in-charge: K. Feijen
CI and DevOps: N. Pigoux and D. Morcuende
General website : H. Stapel

Data exploration

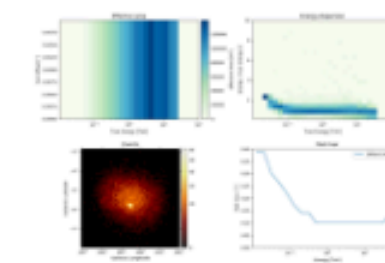
These tutorials show how to perform data exploration with Gammapy, providing an introduction to H.E.S.S., MAGIC, VERITAS, CTAO, HAWC and Fermi-LAT data and their instrument response functions (IRFs). You will learn to explore and filter event lists according to different criteria, view multidimensional IRF files and perform instrument-specific analysis.



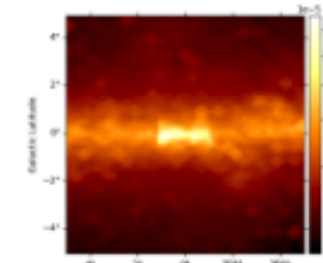
H.E.S.S. with
Gammapy



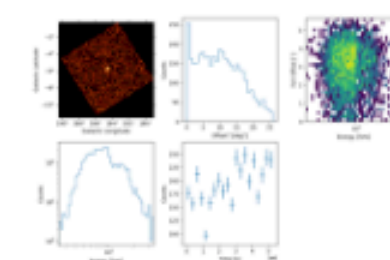
CTAO with
Gammapy



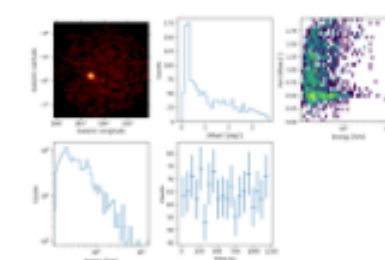
MAGIC with
Gammapy



Fermi-LAT with
Gammapy



HAWC with
Gammapy

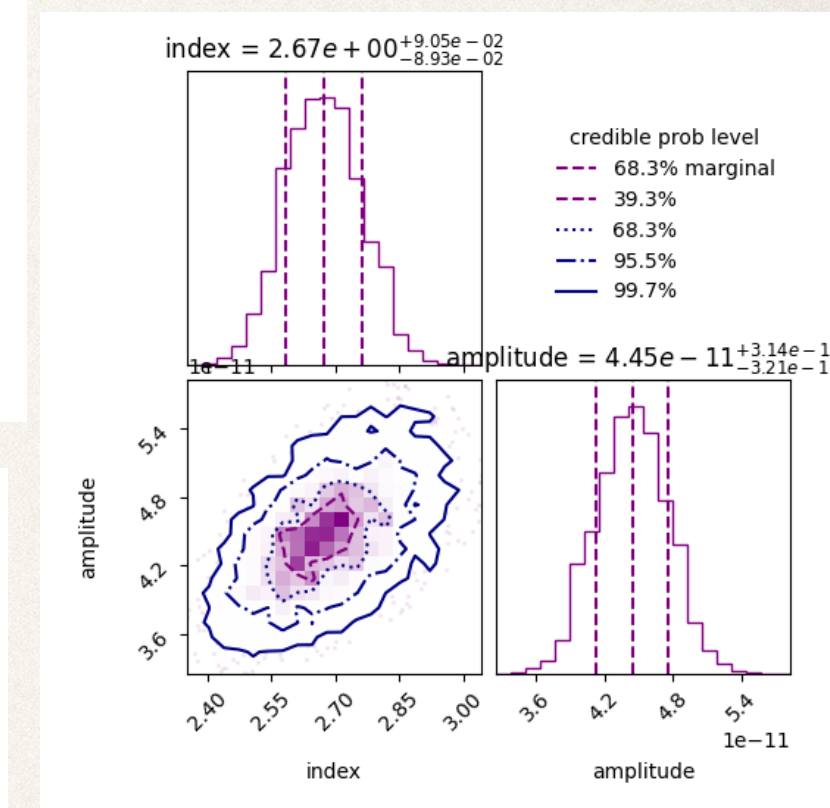
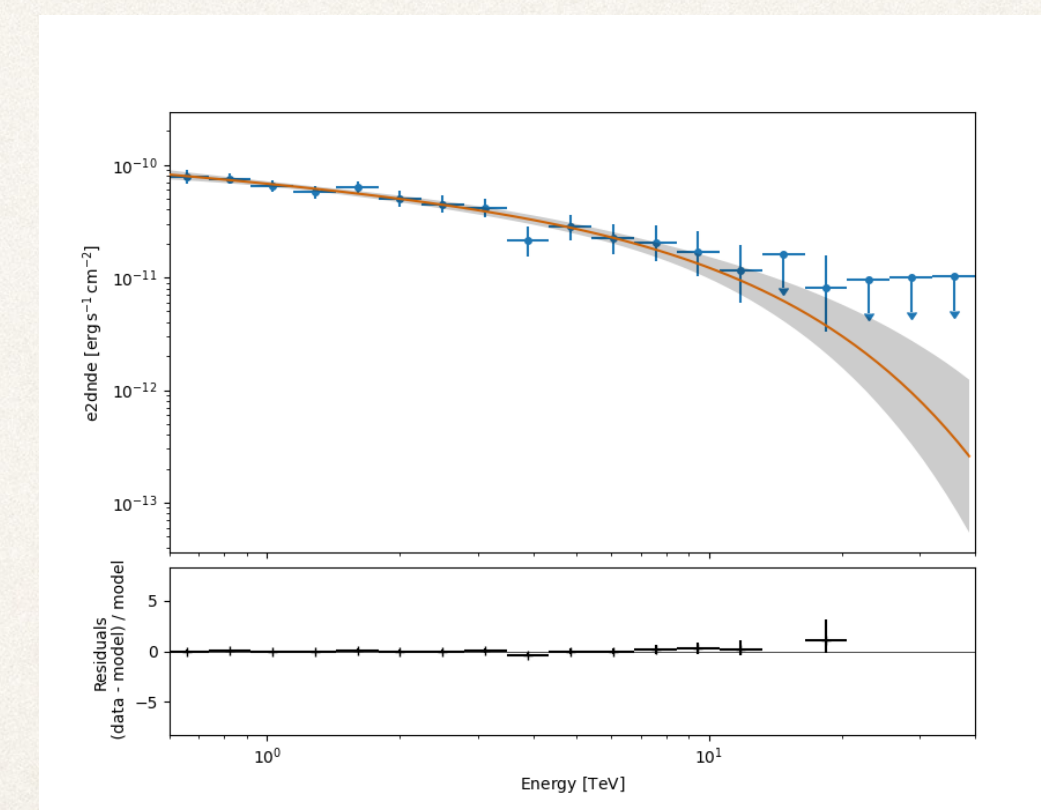
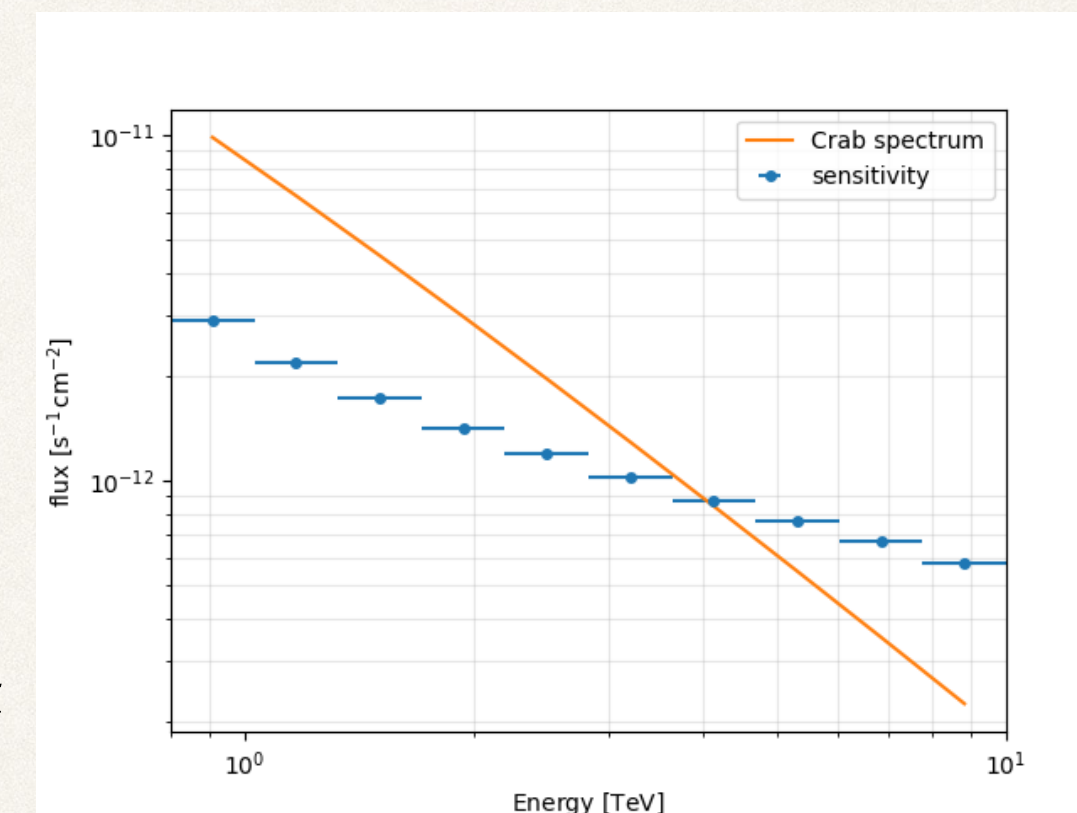


VERITAS with
Gammapy

See full changelog at: <https://docs.gammapy.org/2.0/release-notes/v2.0.html>

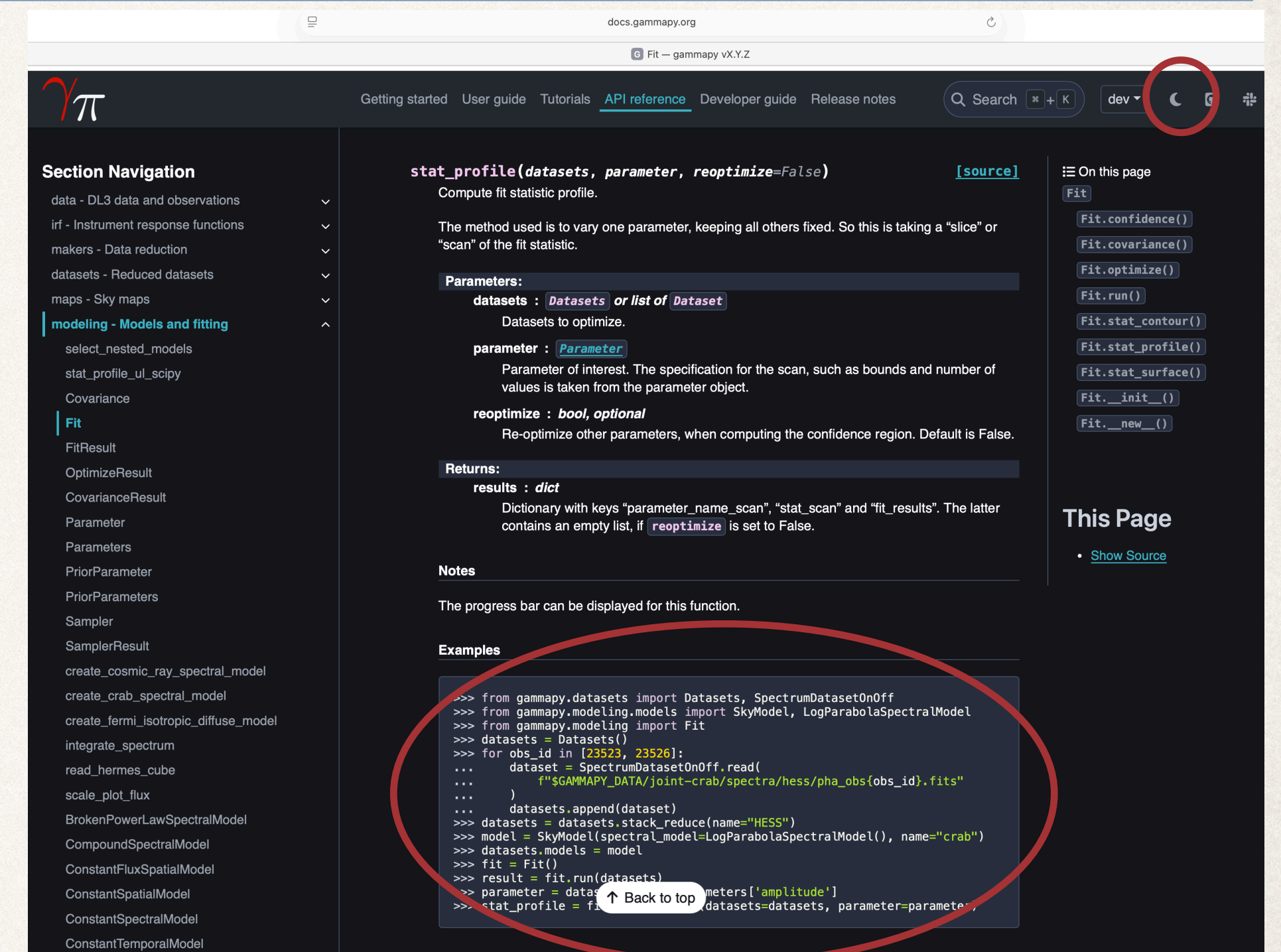
Major new features since 1.3

- ❖ All estimators can create a *flux_sensitivity* now
 - ❖ Minimum detectable flux for the given data and analysis
 - ❖ Calculation based on Asimov dataset following Cowan et al 2011 <https://arxiv.org/abs/1007.1727>
 - ❖ Joint sensitivity can be computed if all data from same instrument (eg: different event type)
- ❖ Error bands computed using multi-normal sampling of spectral parameters
 - ❖ Takes into account all correlations between parameters
 - ❖ Can be slow for models where vectorization is not possible (eg: Naima models)
- ❖ Support for nested sampling with Bayesian analysis
- ❖ Added pulse profiles for the 3rd Fermi Pulsar Catalog
- ❖ Allow configuration of statistic on a dataset using `FitStatistic`` class



Improved documentation

- ❖ Fresh look and feel
 - ❖ Updated appearance with cleaner CSS and **dark theme support** 🌙
- ❖ Easier learning path
 - ❖ Tutorials reordered for better flow & clarity based on user feedback
 - ❖ Code examples now embedded directly in doc-strings
- ❖ New tutorials explaining specific science cases:
 - ❖ How to compute flux upper limit for non-detected sources
 - ❖ Constraining parameter limits from an analysis
 - ❖ End-to-end VERITAS analysis
 - ❖ Fermi-LAT analysis and integration with Fermi-py
 - ❖ Bayesian inference with nested sampling



Upcoming development plans

- ❖ We are currently planning the next release cycle:
 - ❖ Priorities: Requirements of collaborations, maintenance and new features
 - ❖ User inputs on suggested features welcome:
 - ❖ Please consider contributing (get in touch before opening a PR)
 - ❖ All contributors will be listed as authors on Zenodo releases
- ❖ Upcoming user calls will be focussed on specific topics:
 - ❖ **Background modeling:** Understand approaches and experiences from different experiments
 - ❖ **Dark matter analysis:** Need to build a WG on this, collect feedback on available functionalities
 - ❖ Please get in touch if you want to attend/comment

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