

GAMMAPY

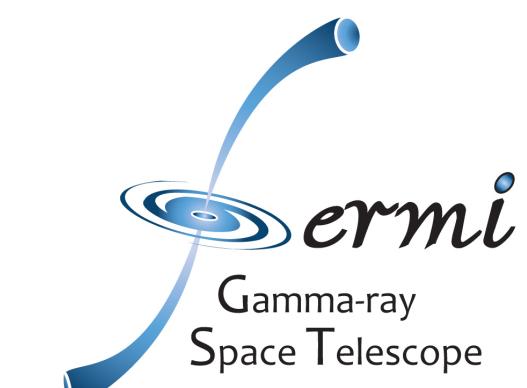
A Python Package for (not only) Gamma-Ray Astronomy

Axel Donath for the Gammapy developers

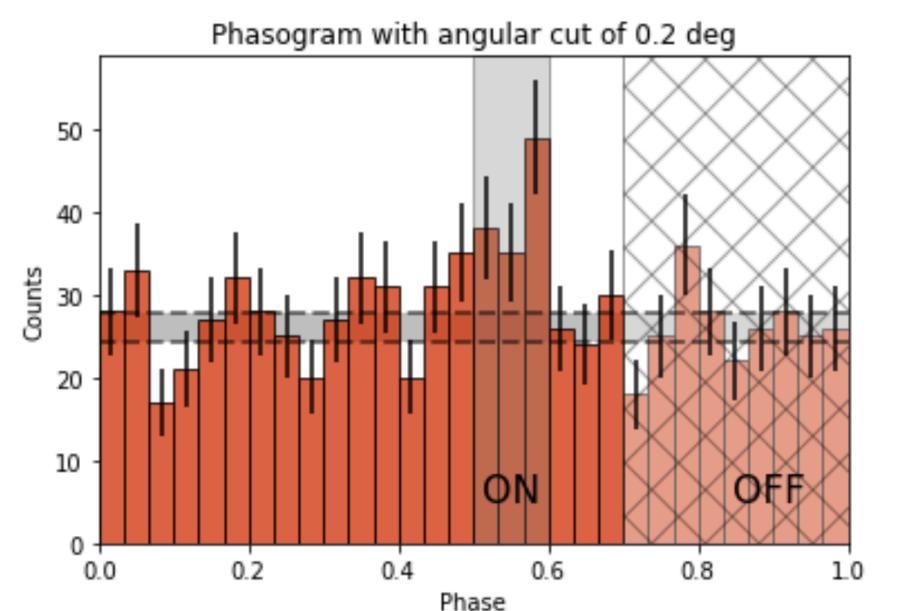
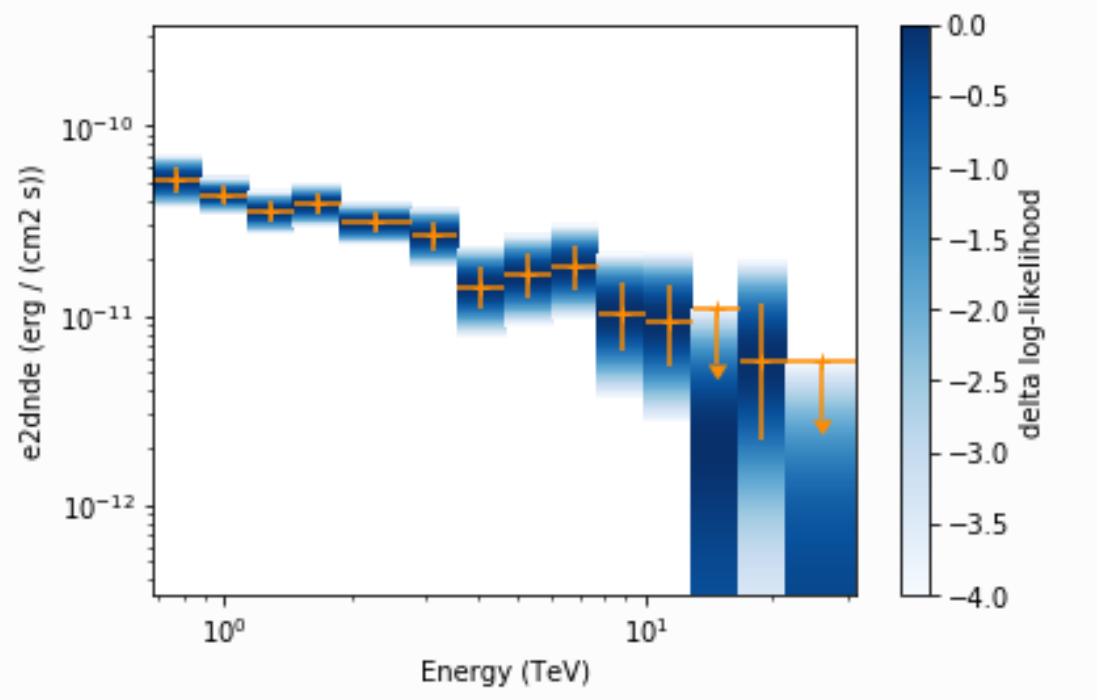
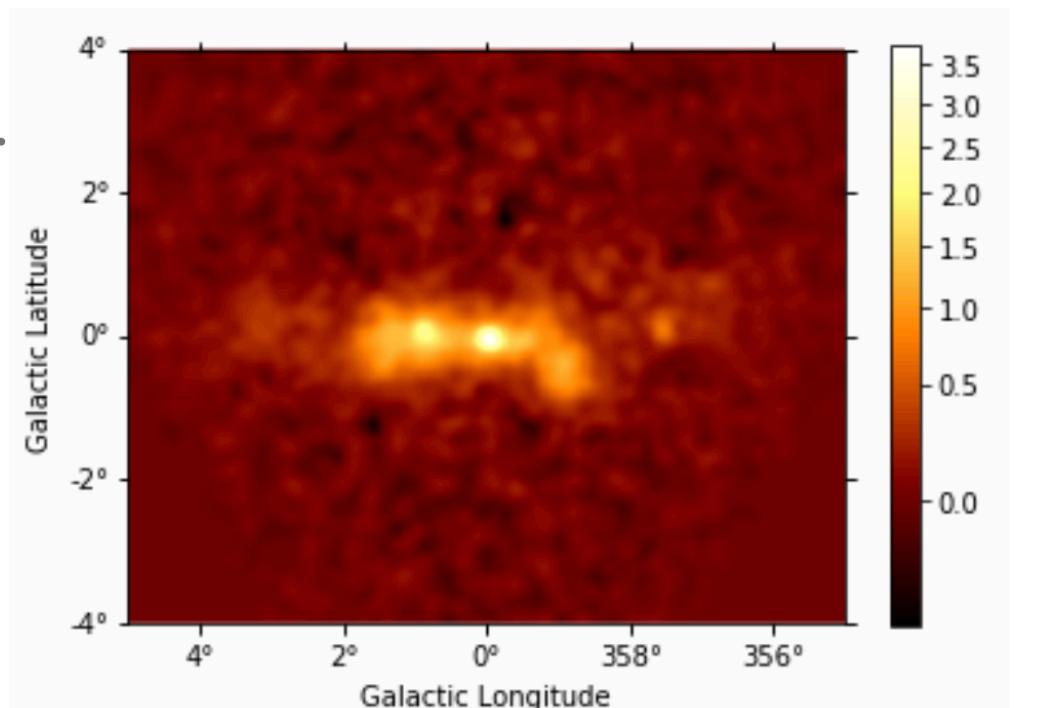
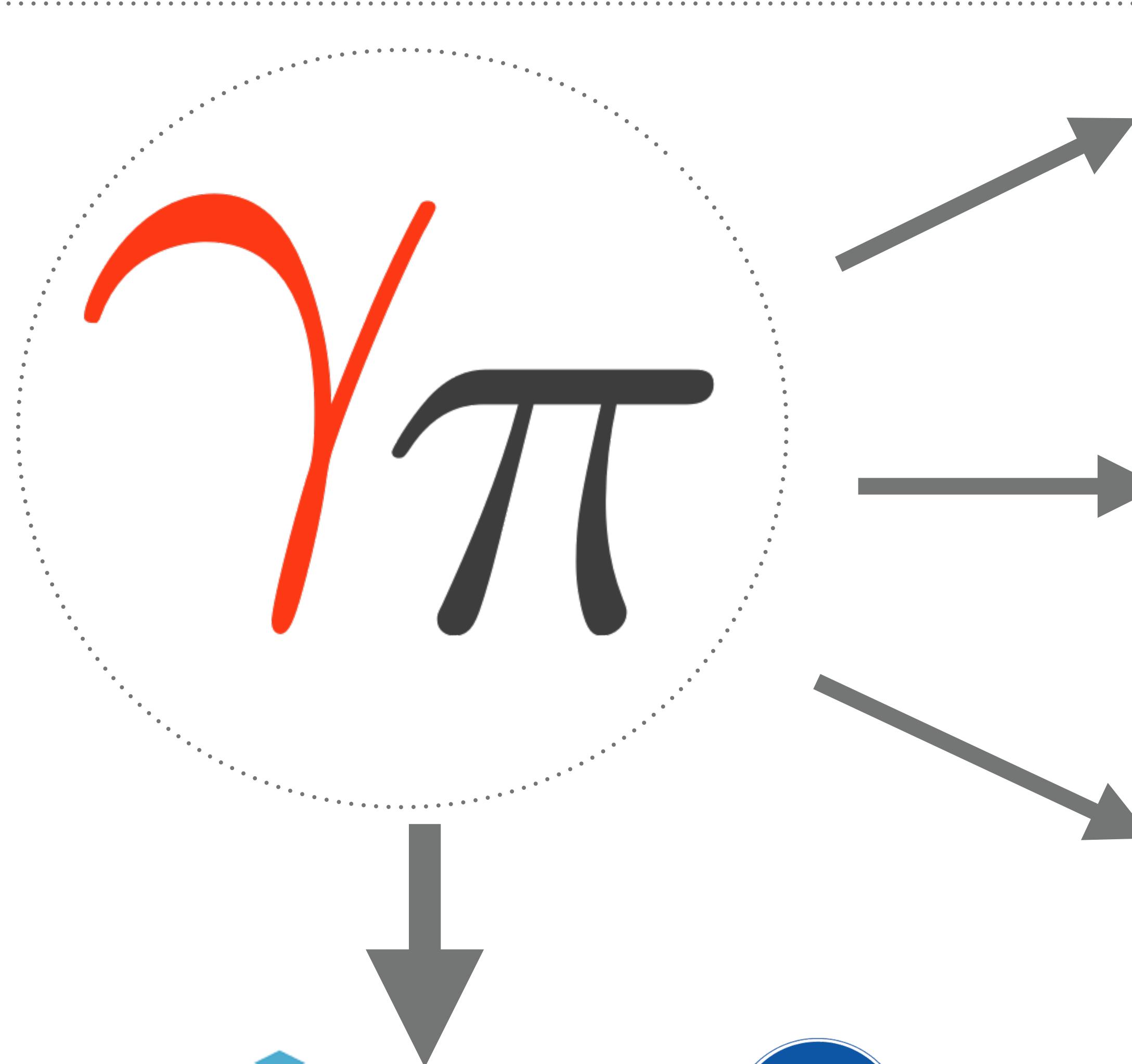
Gammapy User Call, Oct. 26th

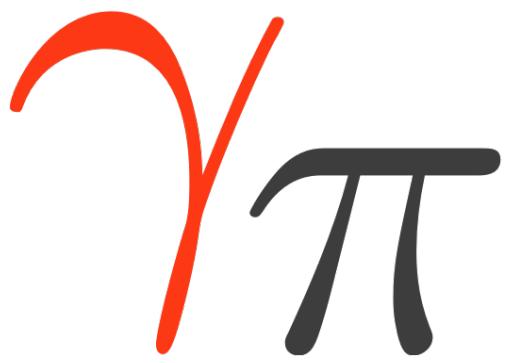
INTRODUCTION

WHAT IS GAMMAPY?



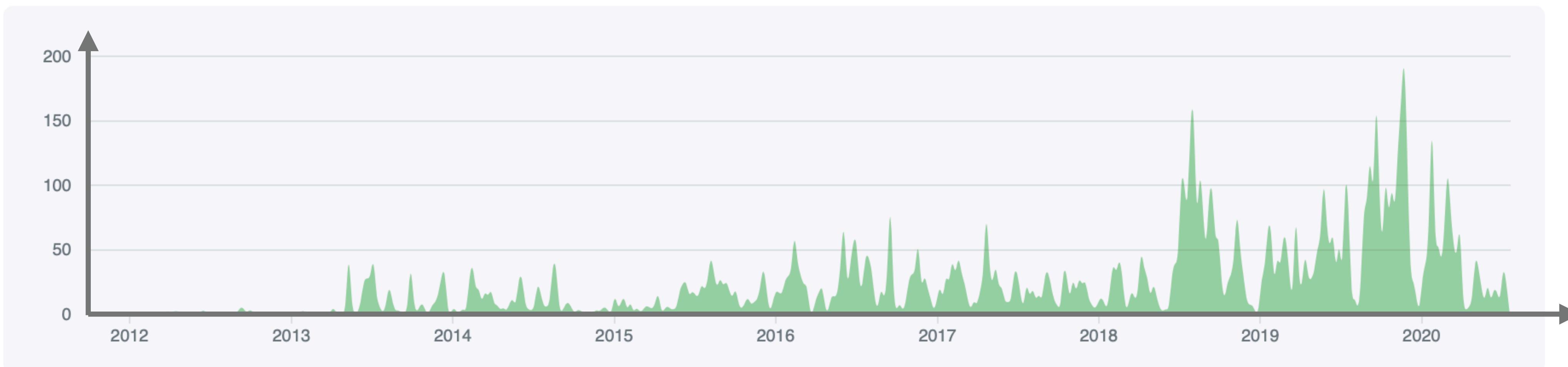
etc.



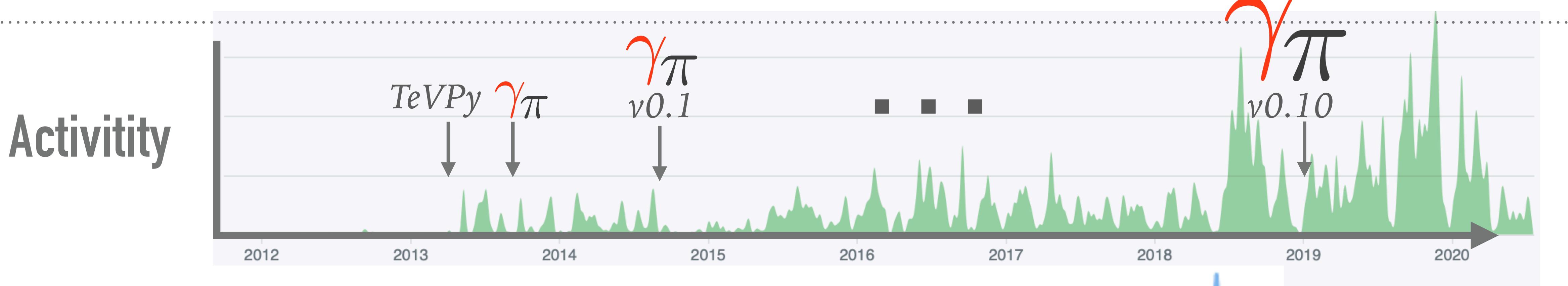


WHAT IS GAMMAPY?

- An openly developed Python package for Gamma-Ray astronomy
- Started in ~2012 with a set of Python scripts developed for the HESS Galactic Plane Survey analysis by Christoph Deil and myself
- Approximately 8 years of experience in developing and maintaining an open source Python package and have probably re-written the package already 3 times...:-)
- **~120 forks, ~60 contributors**
- **~13.000 commits**



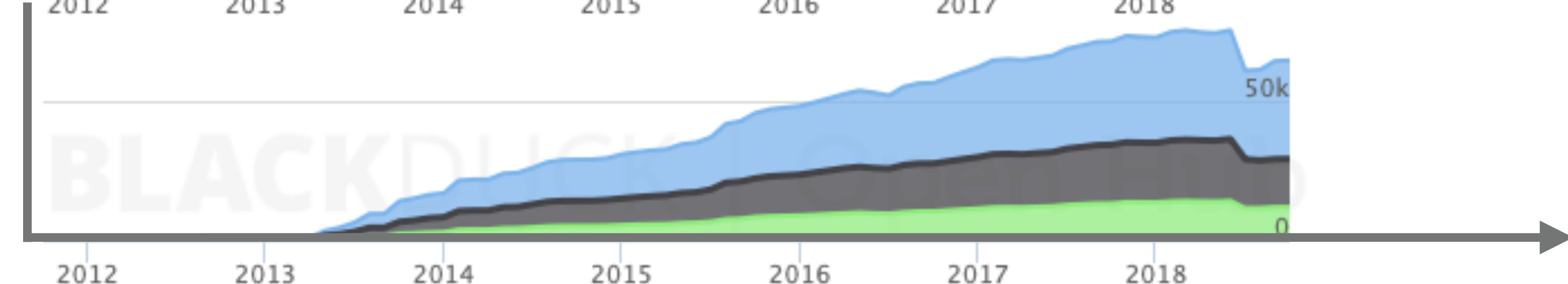
WHAT IS GAMMAPY?



Contributors



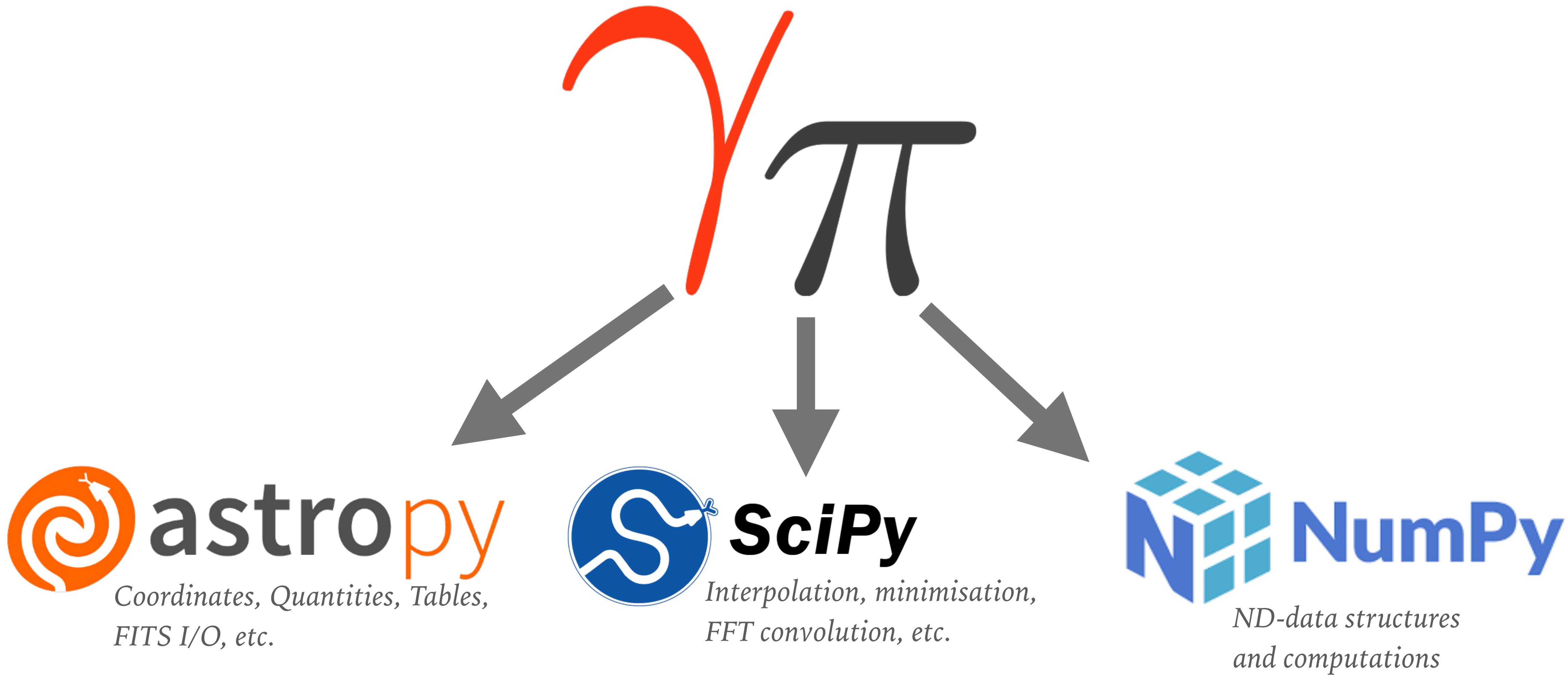
Lines of code



DEVELOPMENT & SETUP

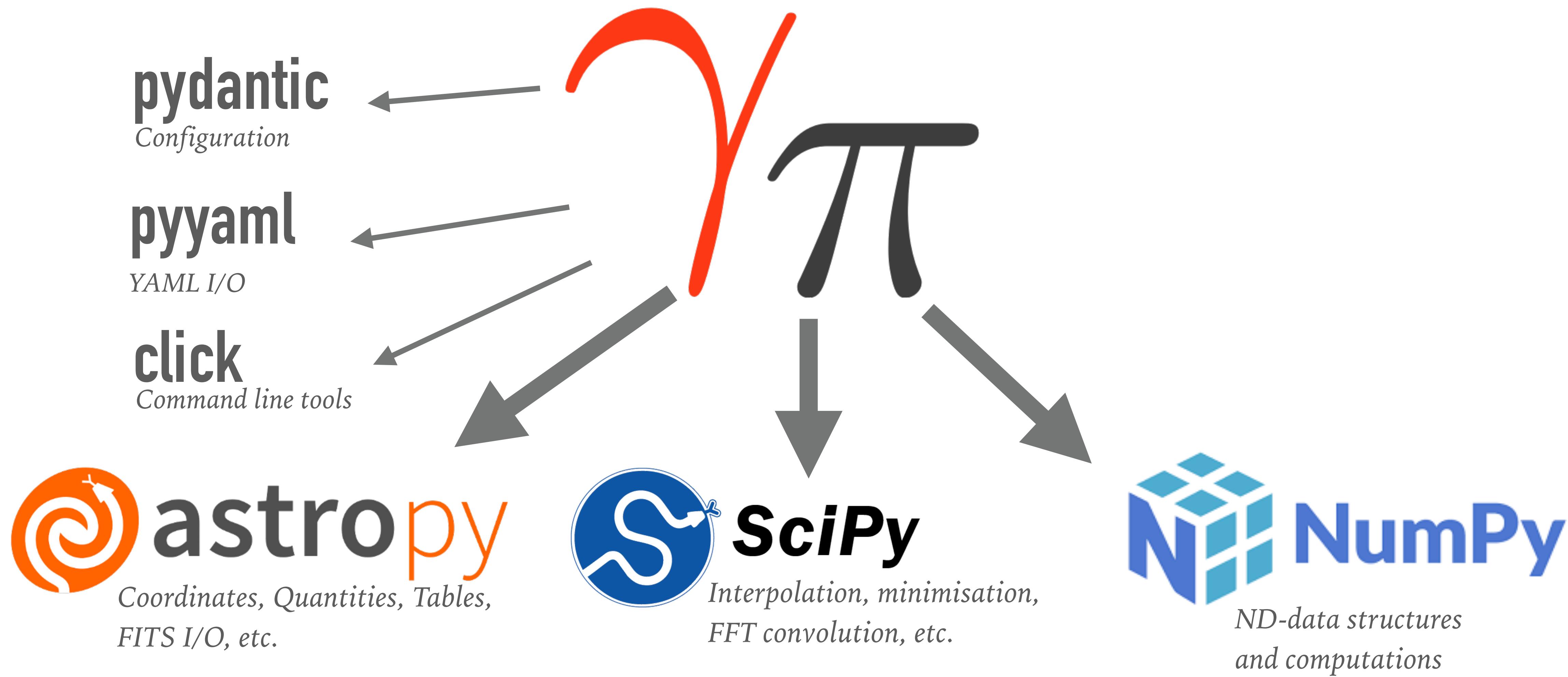
DEPENDENCIES

Required dependencies
→

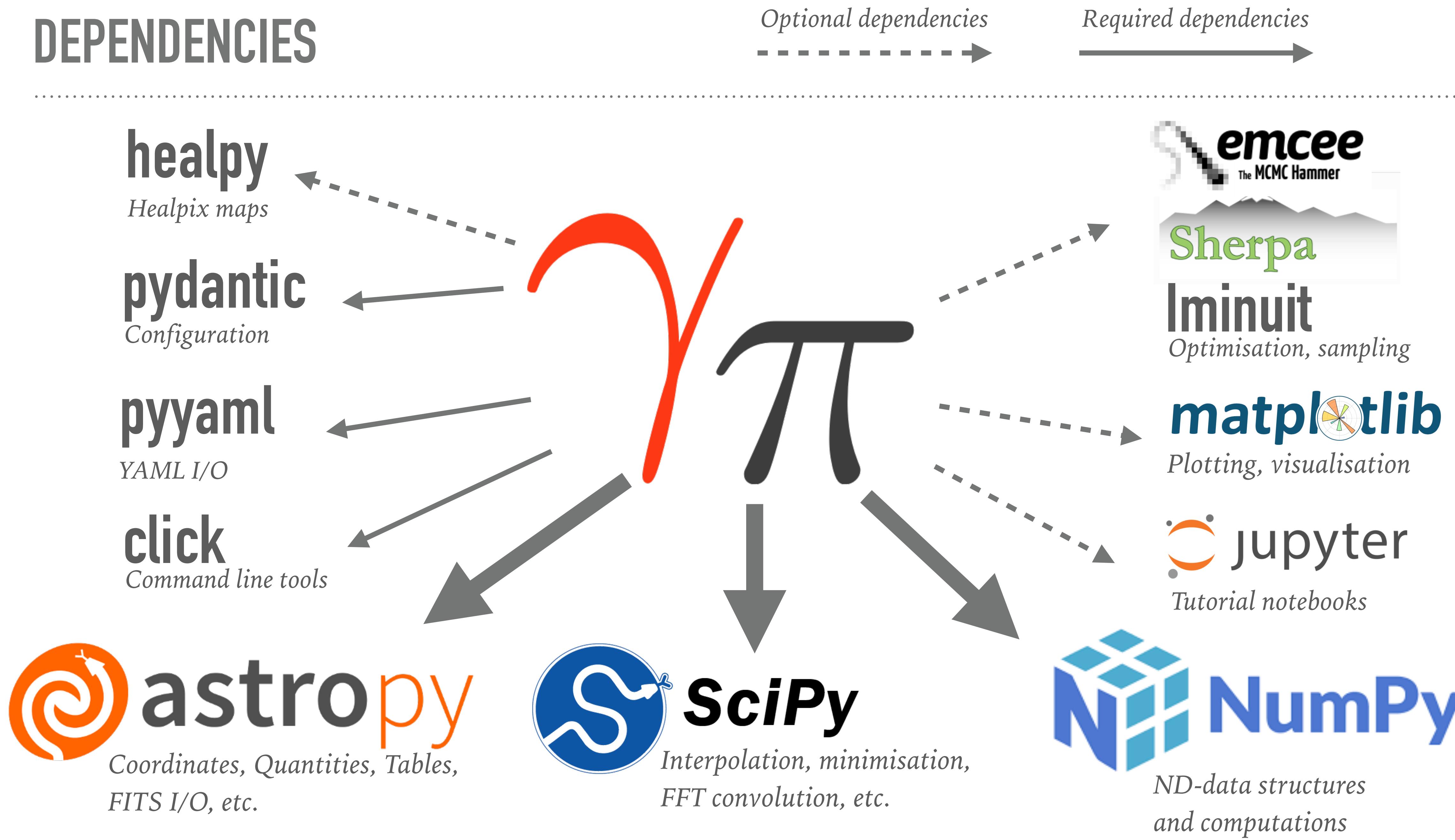


DEPENDENCIES

Required dependencies
→



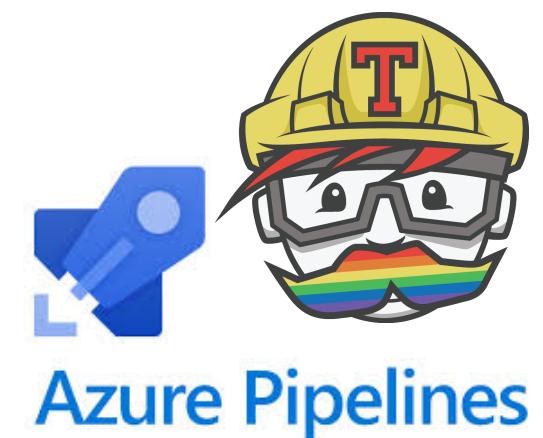
DEPENDENCIES



DEVELOPMENT AND CI SETUP



Hosted and openly developed on Github:
<https://github.com/gammipy/gammipy>



Travis-CI and Azure Pipelines used
for continuous integration



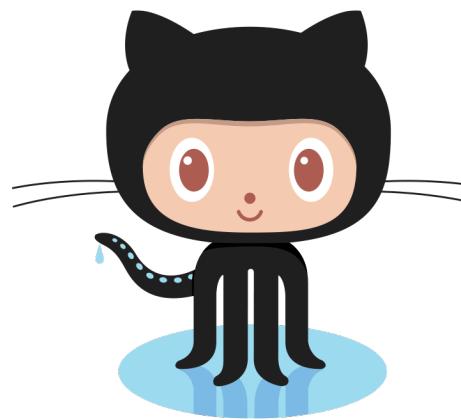
COVERALLS

Coveralls used for monitoring
of code test coverage



Docs are build and deployed
manually: <https://docs.gammipy.org/>

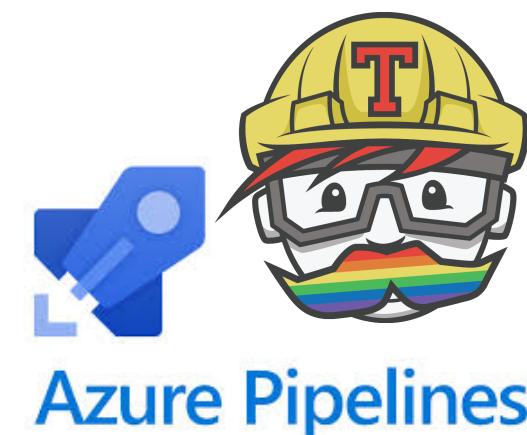
DEVELOPMENT AND CI SETUP



Hosted and openly developed on Github:
<https://github.com/gammipy/gammipy>



Black code formatting used
for a consistent code format.



Travis-CI and Azure Pipelines used
for continuous integration



Sphinx used to build the
documentation



COVERALLS

Coveralls used for monitoring
of code test coverage



pytest

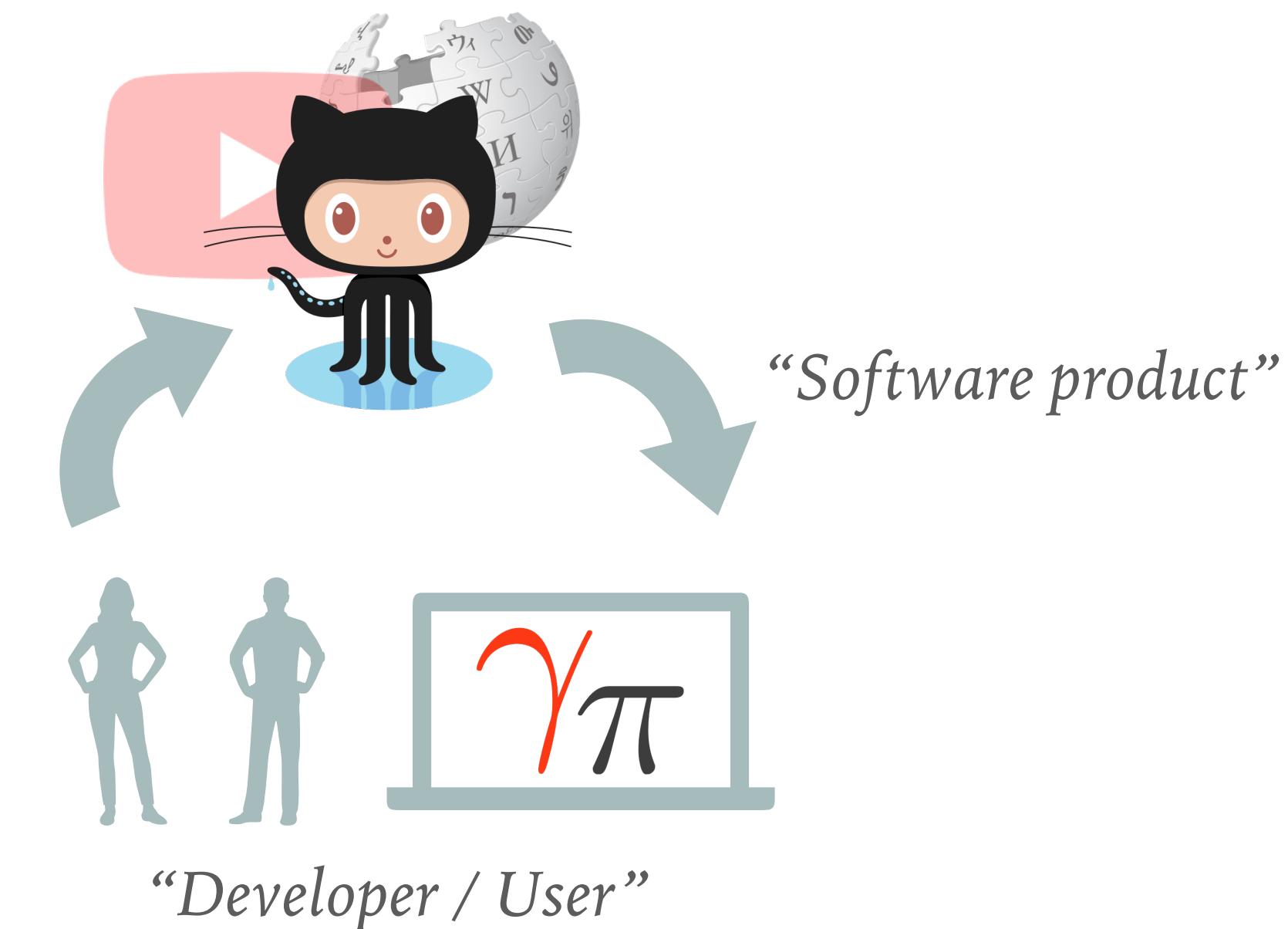
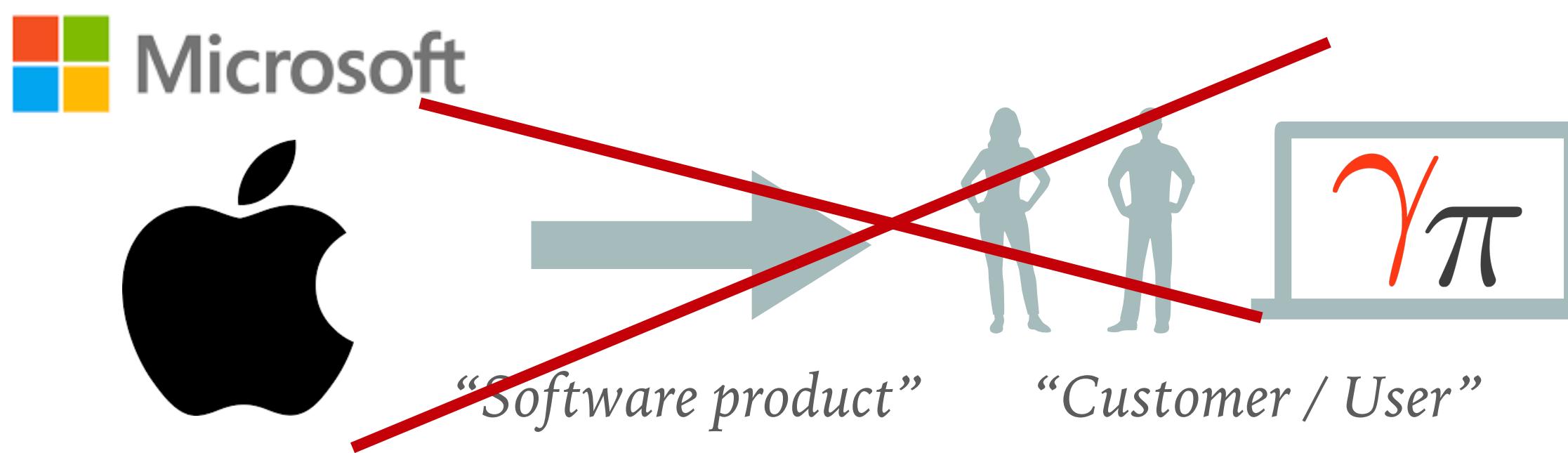
Pytest used for testing



Docs are build and deployed
manually: <https://docs.gammipy.org/>

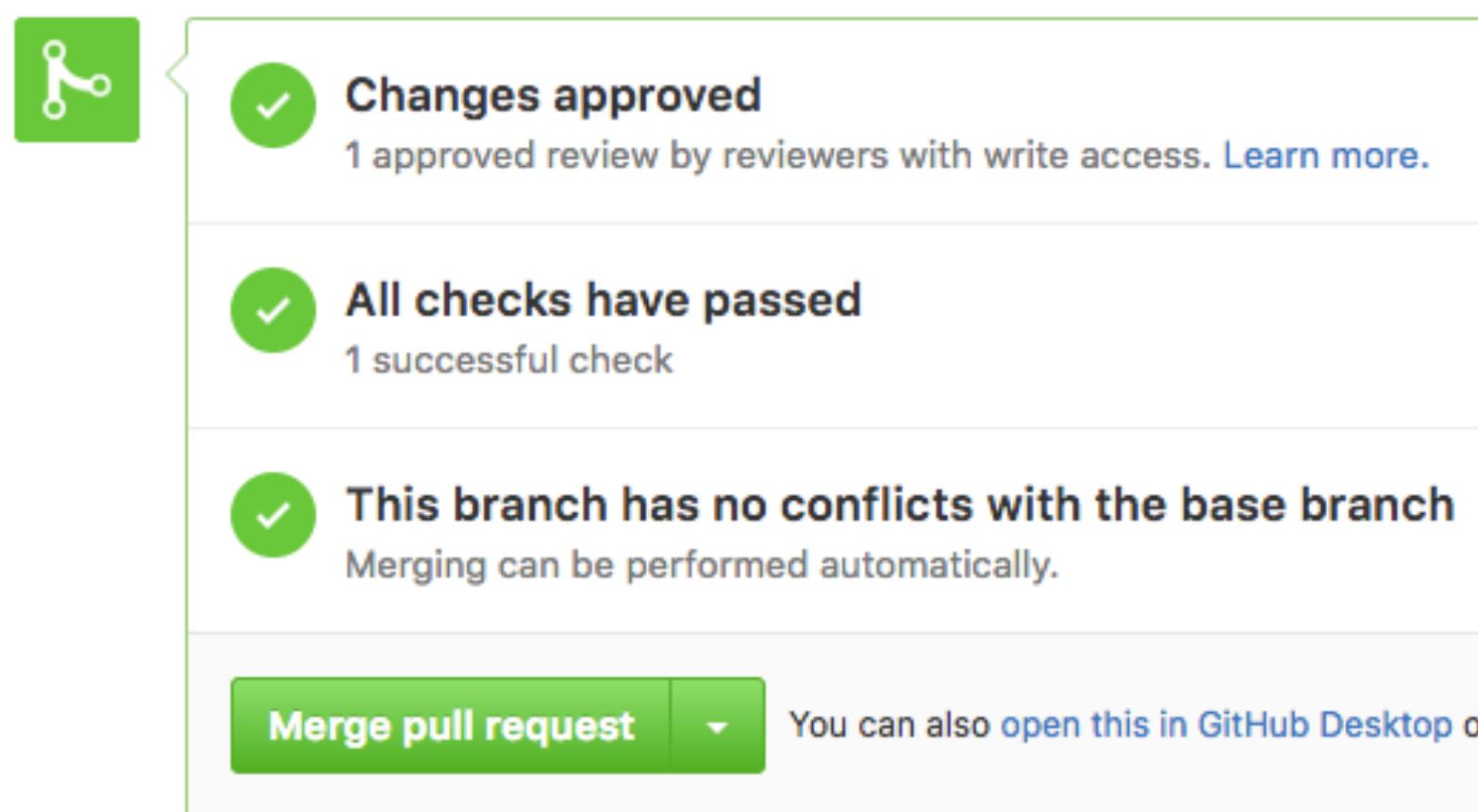
WHAT IS “COMMUNITY DRIVEN” / OPEN DEVELOPMENT?

- Not a classical linear “developer -> user” or “company -> customer” relationship
- **No strict, structural boundary between developers and users**
 - Users know best what they need and deliver it for the benefit of the community
- Maybe **best understood in a Web 2.0 context?** There is no strict boundary between content creator and consumer... (Youtube / Wikipedia etc.)



“GITHUB” WORKFLOW

- Standard multi-branch git workflow:
 - Contributors fork a repository
 - Features are developed in new a branch “on the side”
 - A pull requests is opened
 - Every pull requests (PR) is reviewed at least once by more experienced developers (lead developers). Sometimes “all fine”, sometime “Here is a number of substantial comments”
 - Once review comments are implemented and the CI builds pass a PR gets merged



A screenshot of a GitHub pull request merge screen. At the top left is a green icon with a gear and wrench. To its right, a green circle with a checkmark contains the text "Changes approved". Below it, another green circle with a checkmark contains "All checks have passed". Further down, a third green circle with a checkmark contains "This branch has no conflicts with the base branch". At the bottom left is a green button labeled "Merge pull request". To its right, a message says "You can also open this in GitHub Desktop or".

INSTALLATION & SETUP

<https://docs.gammapy.org/0.17/install/index.html>

Install Gammapy:

- 1.** `$ curl -O https://gammapy.org/download/install/gammapy-0.17-environment.yml`
- 2.** `$ conda env create -f gammapy-0.17-environment.yml`
- 3.** `$ conda activate gammapy-0.17`



ANACONDA®



INSTALLATION & SETUP

<https://docs.gammapy.org/0.17/install/index.html>

Install Gammapy:

1. `$ curl -O https://gammapy.org/download/install/gammapy-0.17-environment.yml`
2. `$ conda env create -f gammapy-0.17-environment.yml`
3. `$ conda activate gammapy-0.17`



Download tutorials:

1. `$ gammapy download tutorials`
2. `$ cd gammapy-tutorials`
3. `$ export GAMMAPY_DATA=$PWD/datasets`



TUTORIALS

- Best starting point are the tutorials notebooks: <https://docs.gammipy.org/0.17/tutorials/index.html>
- CTA data access: <https://docs.gammipy.org/0.17/notebooks/cta.html>
- CTA data analysis: https://docs.gammipy.org/0.17/notebooks/cta_data_analysis.html
- And many, many others...
- Tutorials can be executed online as well (using Binder):

Click here!

This is a fixed-text-formatted version of a Jupyter notebook

Try online [launch binder](#)

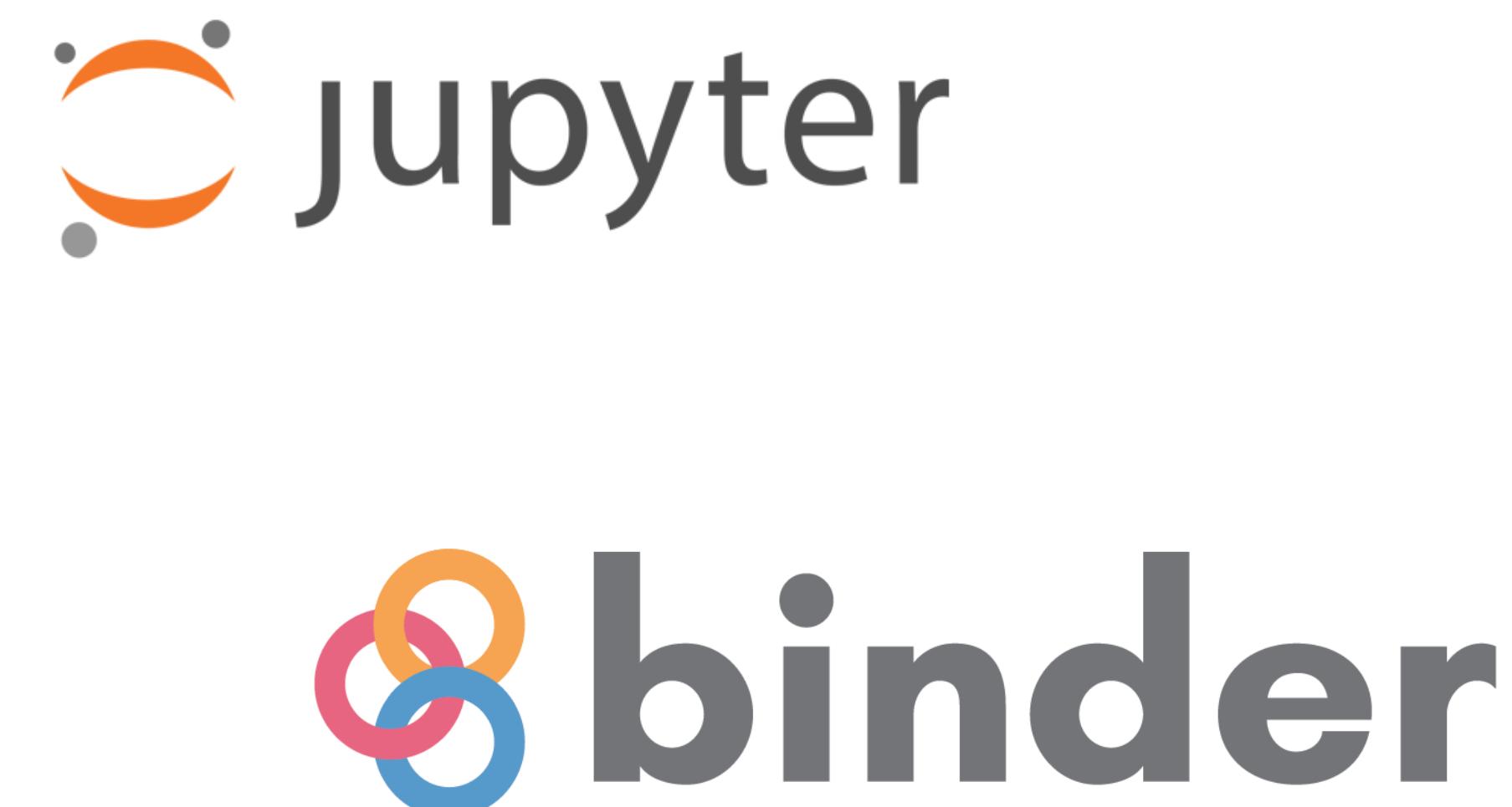
You can contribute with your own notebooks in this [GitHub repository](#).

Source files: [cta.ipynb](#) | [cta.py](#)

CTA with Gammipy

Introduction

The Cherenkov Telescope Array (CTA) is the next generation ground-based observatory for gamma-ray astronomy. Gammipy is a prototype for the Cherenkov Telescope Array (CTA) science tools (2017ICRC...35..766D).



RELATED PROJECTS



“An open data collection
and source catalog for VHE
gamma-ray astronomy”

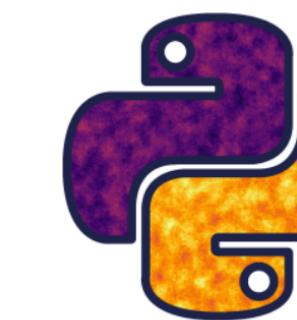
<https://github.com/gammipy/gamma-cat>



gammasky.net

“A portal to the gamma-ray sky”

<https://github.com/gammipy/gamma-sky>



PyGamma19

“Python and open data for
Gamma-Ray Astronomy
Workshop”, March 18th - 22nd

<https://indico.cern.ch/event/783425/>



Gamma-astro-data-formats

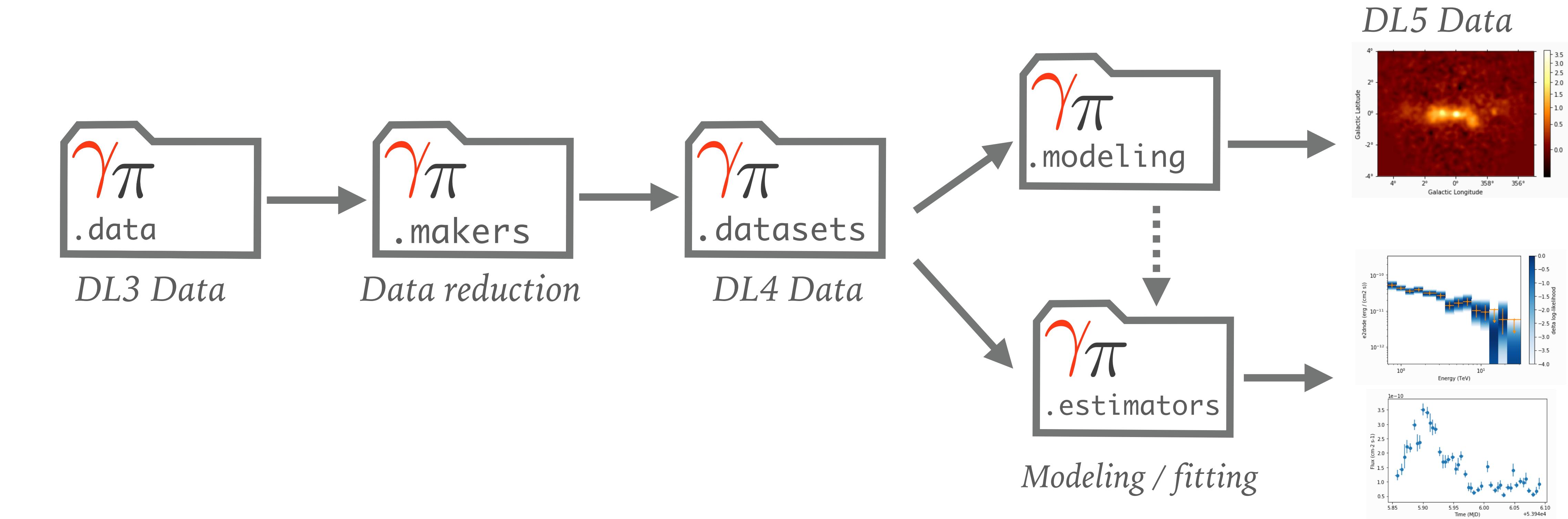
“A place to propose and share data format
descriptions for gamma-ray astronomy.”

<https://github.com/open-gamma-ray-astro>

FEATURES

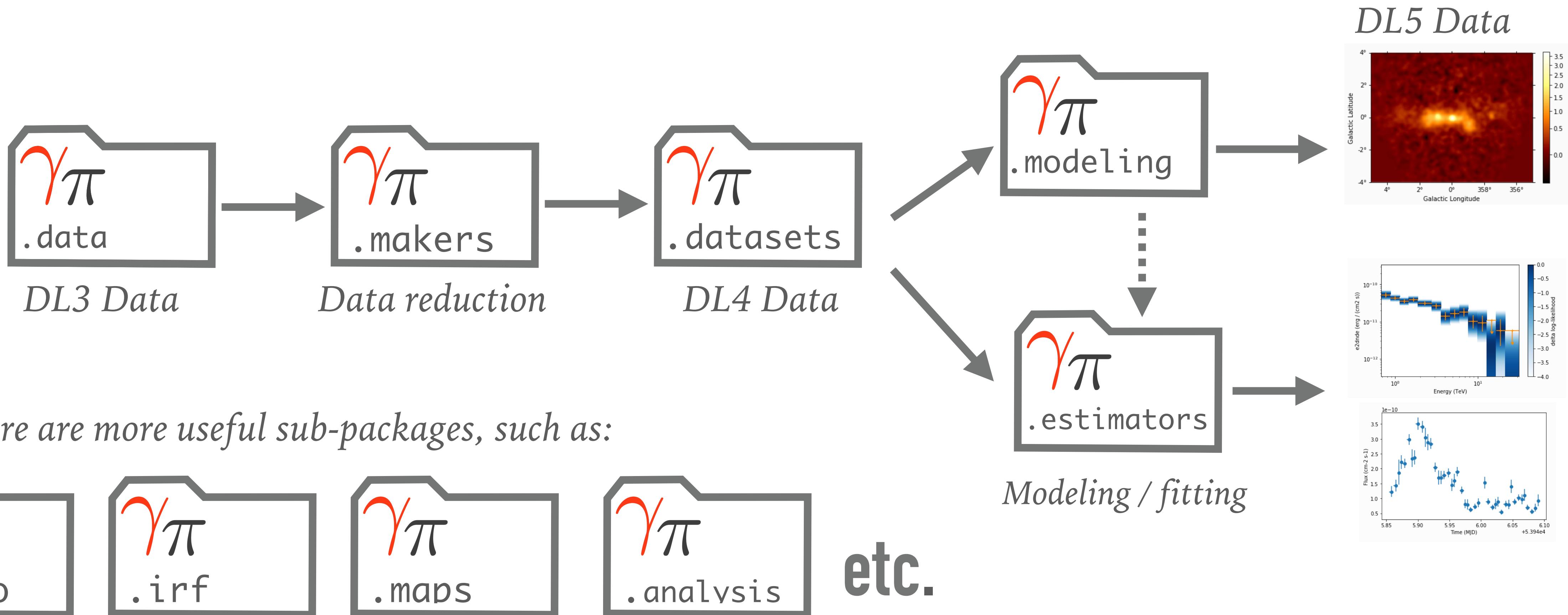
PACKAGE STRUCTURE

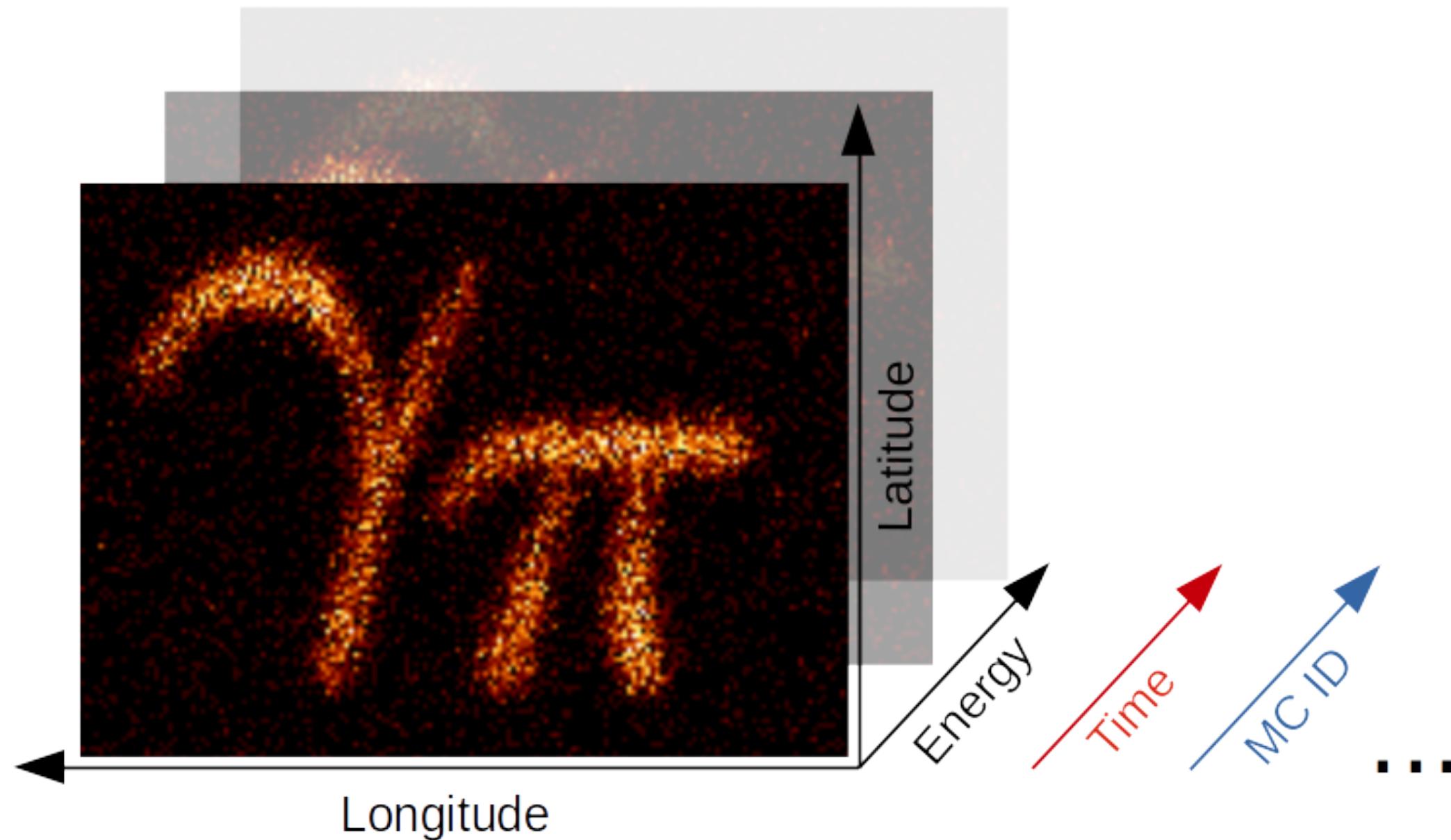
- Gammapy is structured into sub-package, mostly based by API and data level. The core functionality for a typical analysis workflow is distributed as following:



PACKAGE STRUCTURE

- Gammapy is structured into sub-package, mostly based by API and data level. The core functionality for a typical analysis workflow is distributed as following:





Uniform API for WCS and HEALPix...

```
from gammapy.maps import Map
from astropy.coordinates import SkyCoord

position = SkyCoord(0.0, 5.0, frame='galactic', unit='deg')
# Create a WCS Map
m_wcs = Map.create(binsz=0.1, map_type='wcs', skydir=position, width=10.0)
# Create a HPX Map
m_hpx = Map.create(binsz=0.1, map_type='hpx', skydir=position, width=10.0)
```

- WCS and HEALPix image based data structures with arbitrary number of non-spatial extra axes such as energy or time
- Uniform API for WCS and HEALPix based maps
- Interpolation, re-projection, smoothing, convolution, FITS serialisation, (interactive) plotting etc. all generalised to N-dimensions

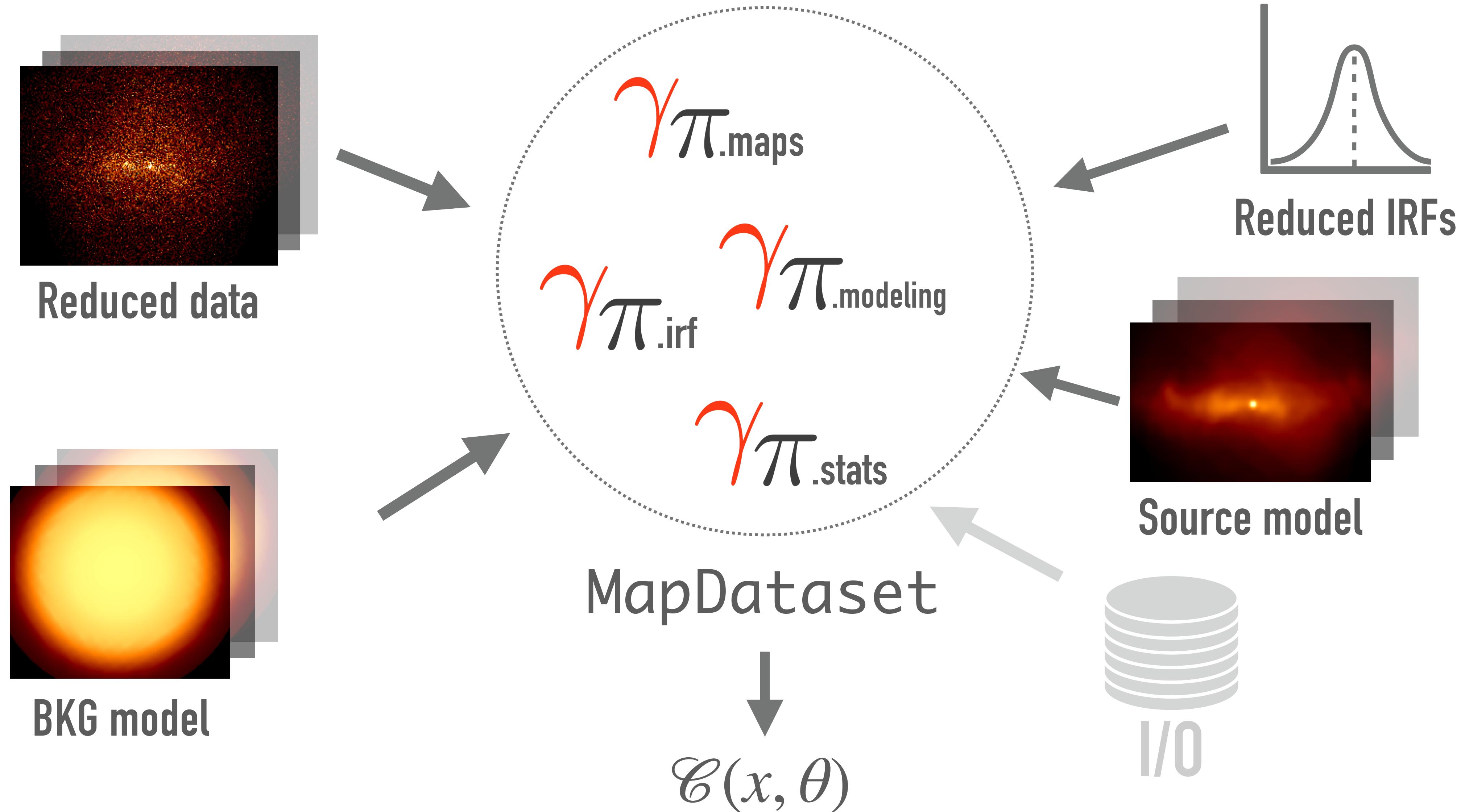
Handling of arbitrary non-spatial extra axes...

```
from gammapy.maps import Map, MapAxis
from astropy.coordinates import SkyCoord

position = SkyCoord(0.0, 5.0, frame='galactic', unit='deg')
energy_axis = MapAxis.from_bounds(100., 1E5, 12, interp='log', unit='GeV', name='energy')
time_axis = MapAxis.from_bounds(0., 12, 12, interp='lin', unit='h', name='time')

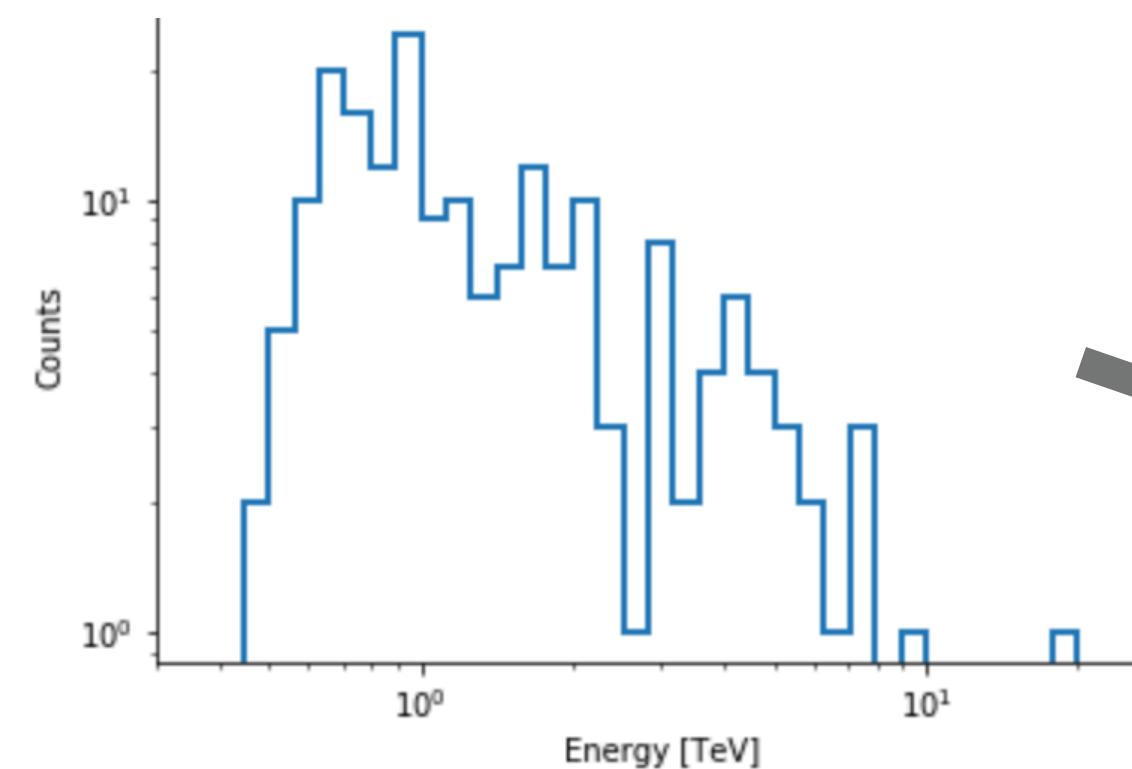
# Create a WCS Map
m_wcs = Map.create(binsz=0.02, map_type='wcs', skydir=position, width=10.0,
                   axes=[energy_axis, time_axis])
```

GAMMAPY.DATASETS

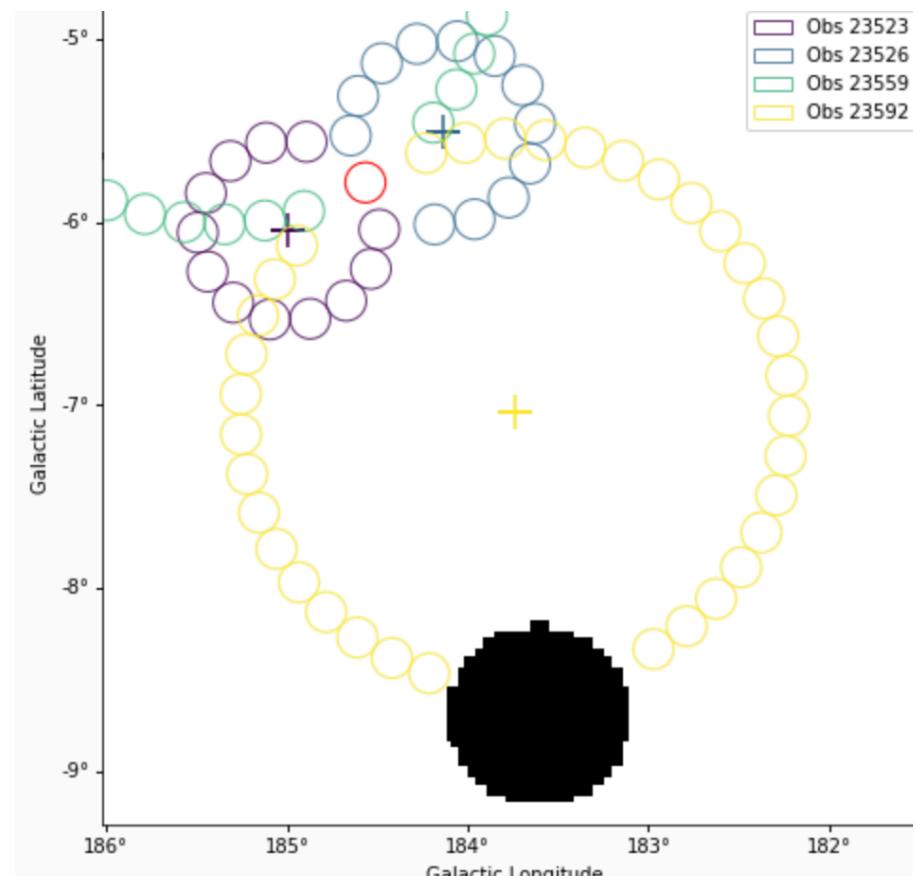


GAMMAPY.DATASETS

OGIP Format



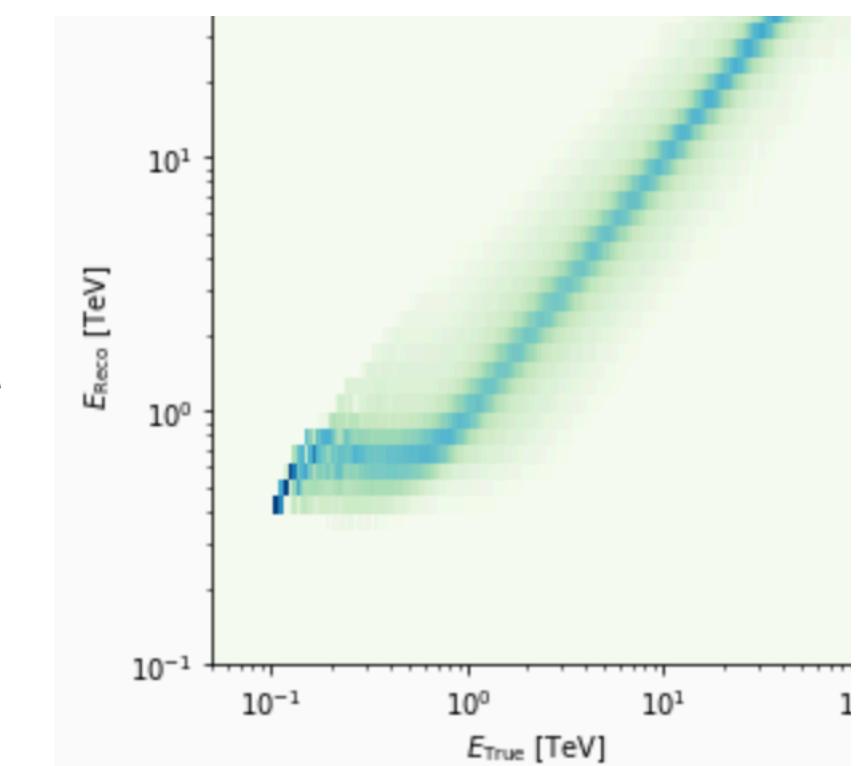
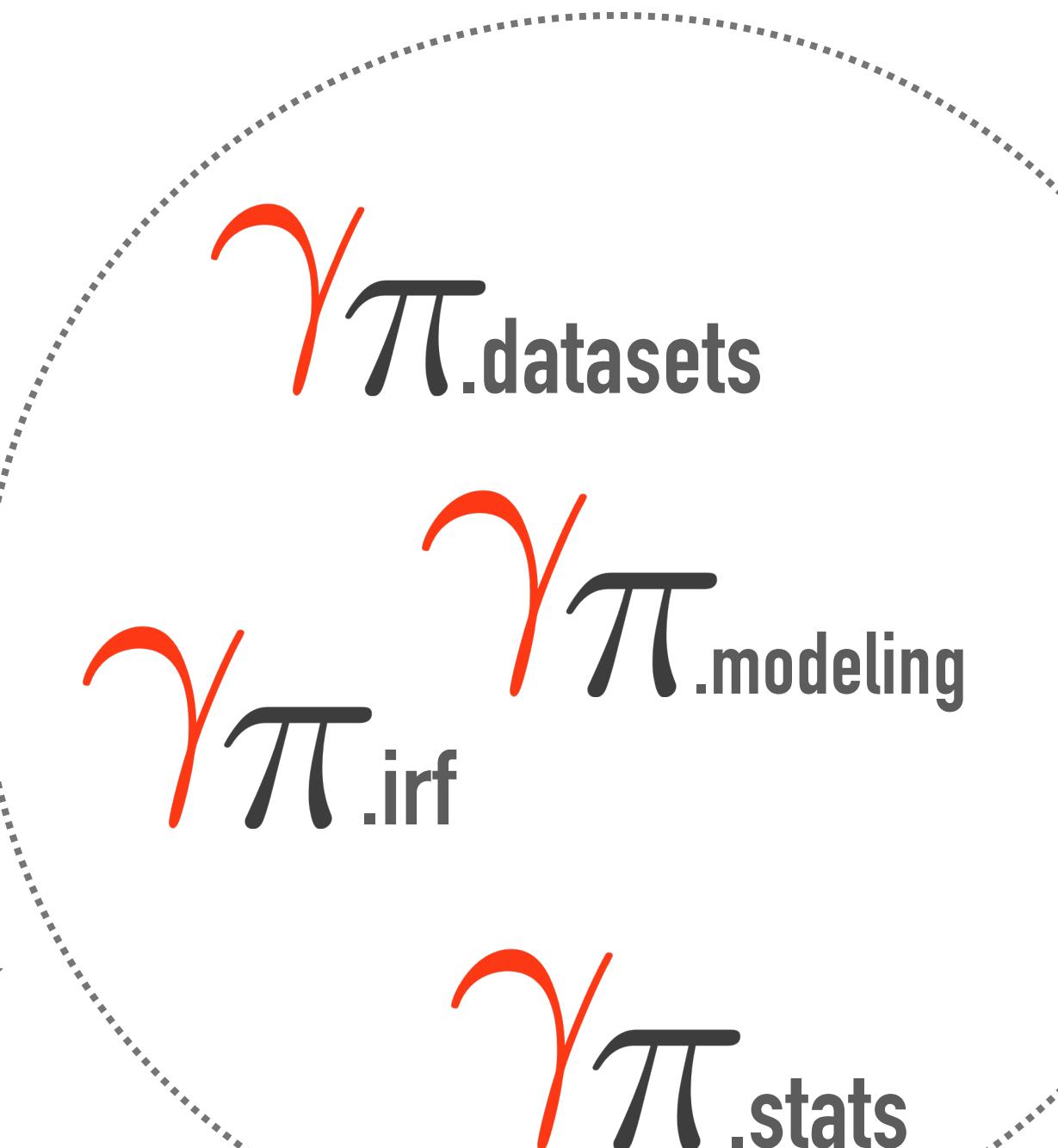
Reduced data



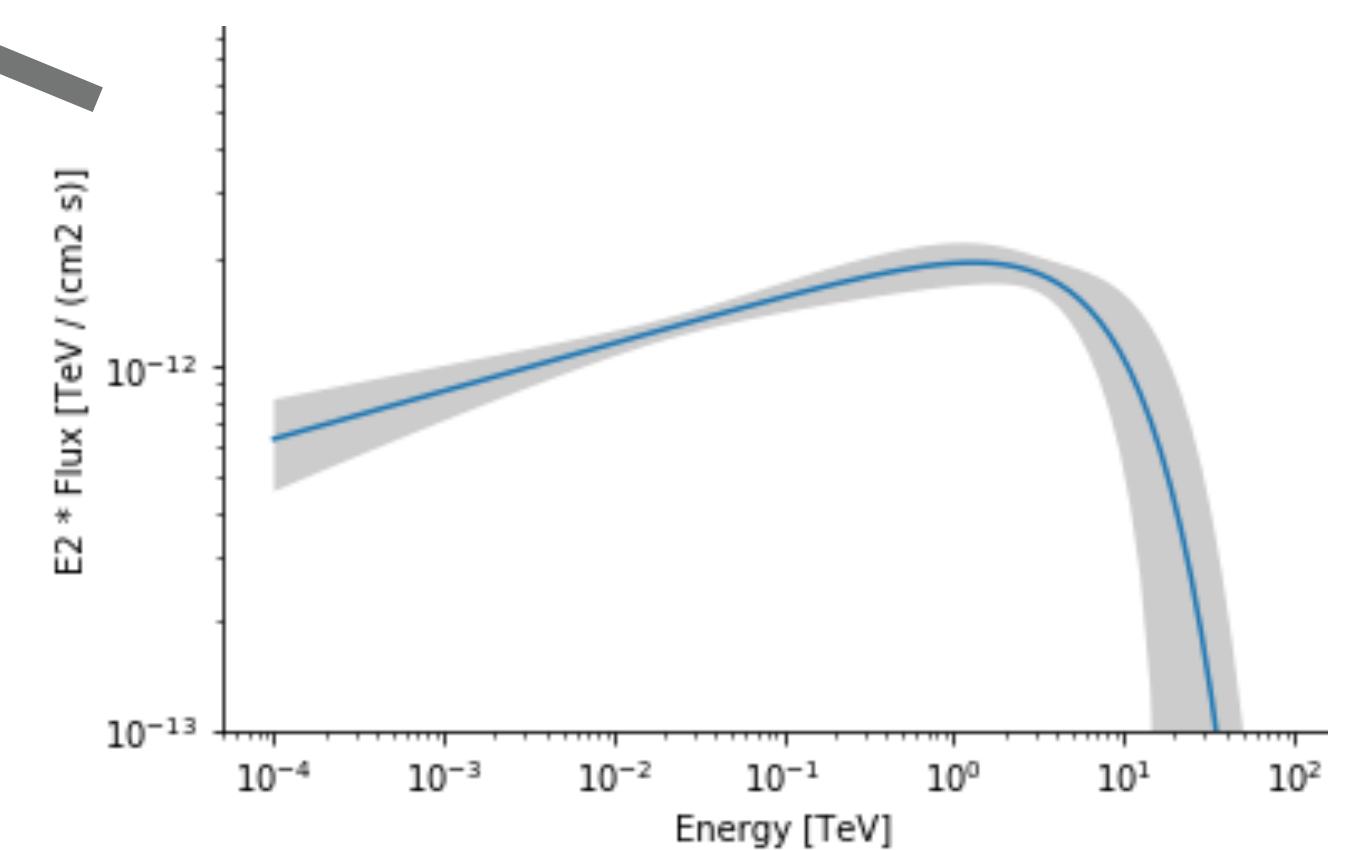
BKG data / model

SpectrumDatasetOnOff

$$\mathcal{W}(x, \theta)$$

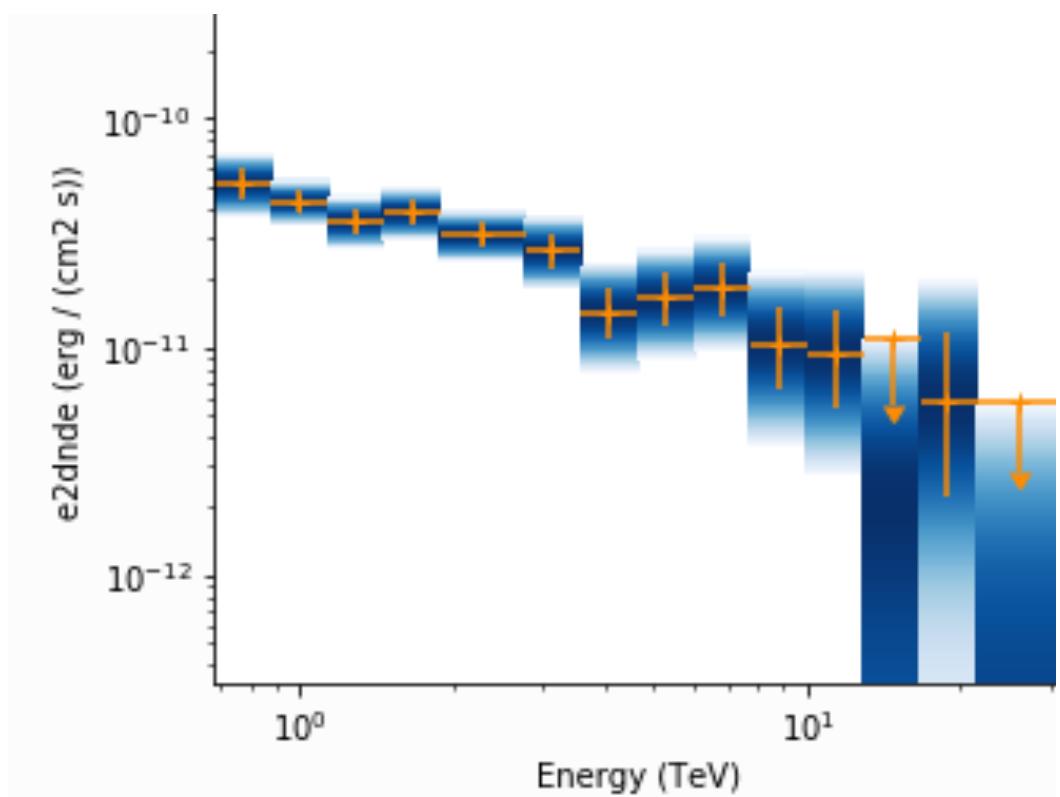


Reduced IRFs

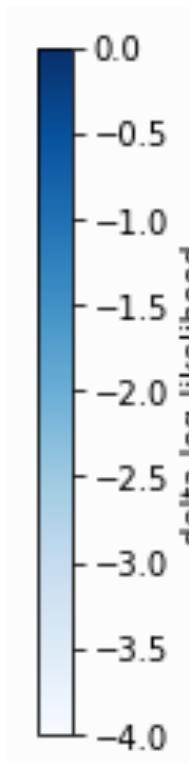


Spectral model

GAMMAPY.DATASETS



Reduced data

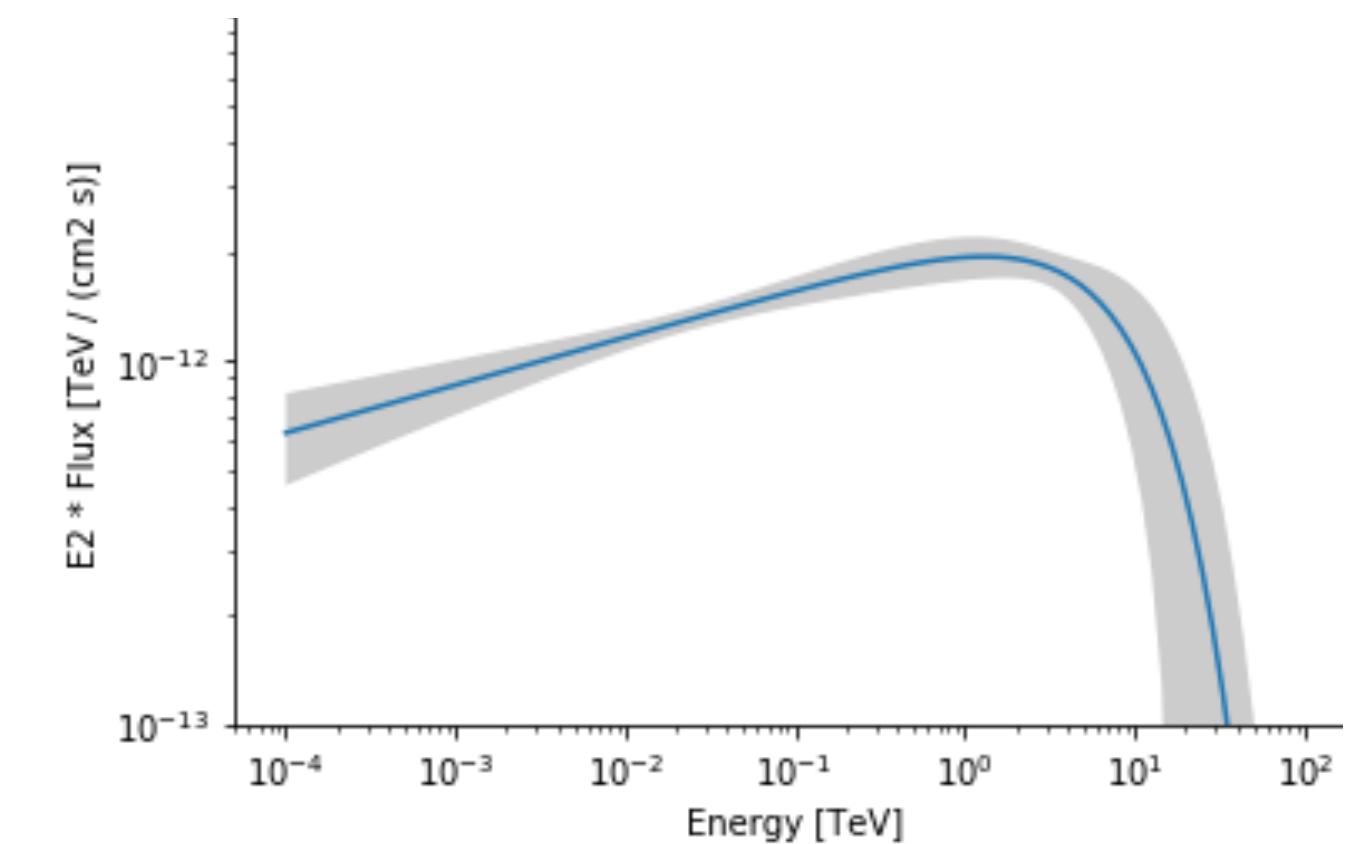


FluxPointsDataset

$\gamma \pi_{\text{estimators.flux_points}}$

$\gamma \pi_{\text{modeling.models}}$

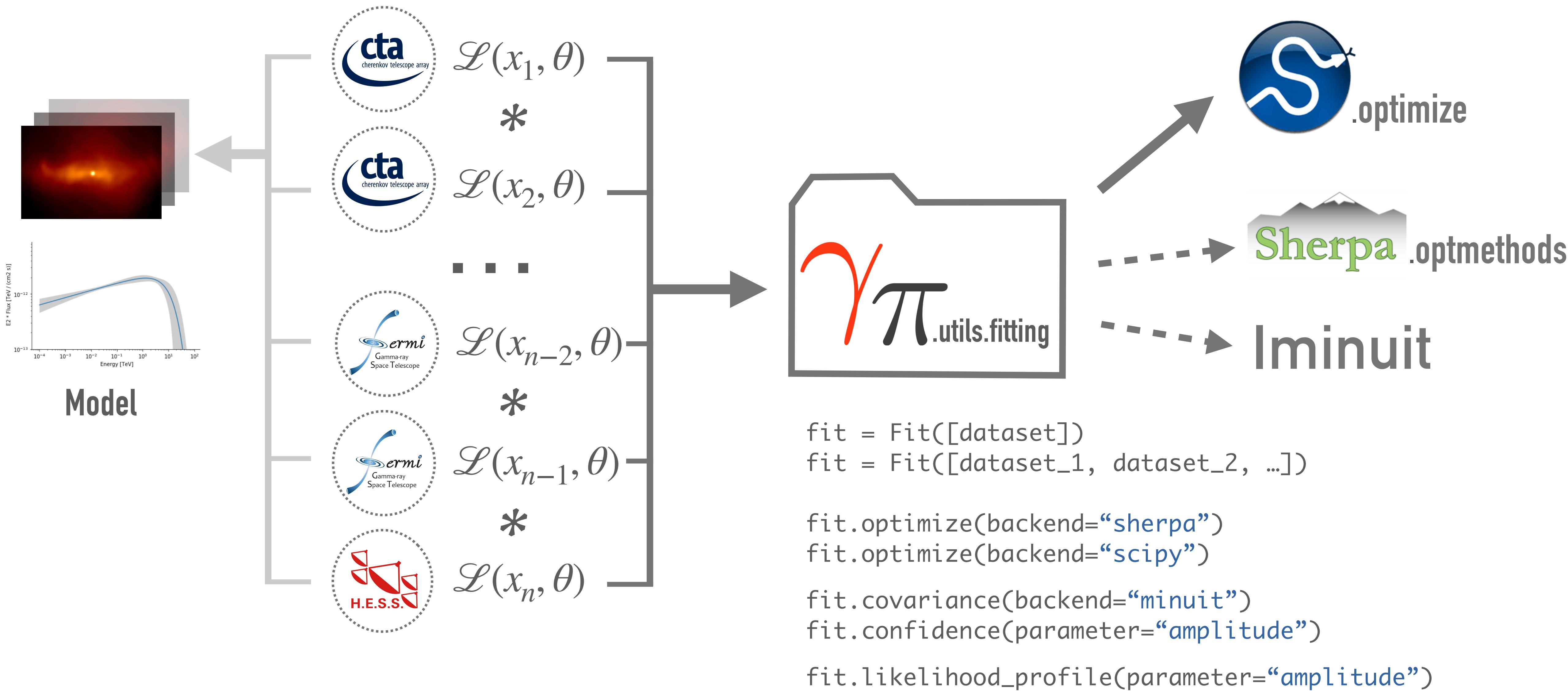
$\gamma \pi_{\text{stats}}$



Spectral model

$$\chi^2(x, \theta)$$

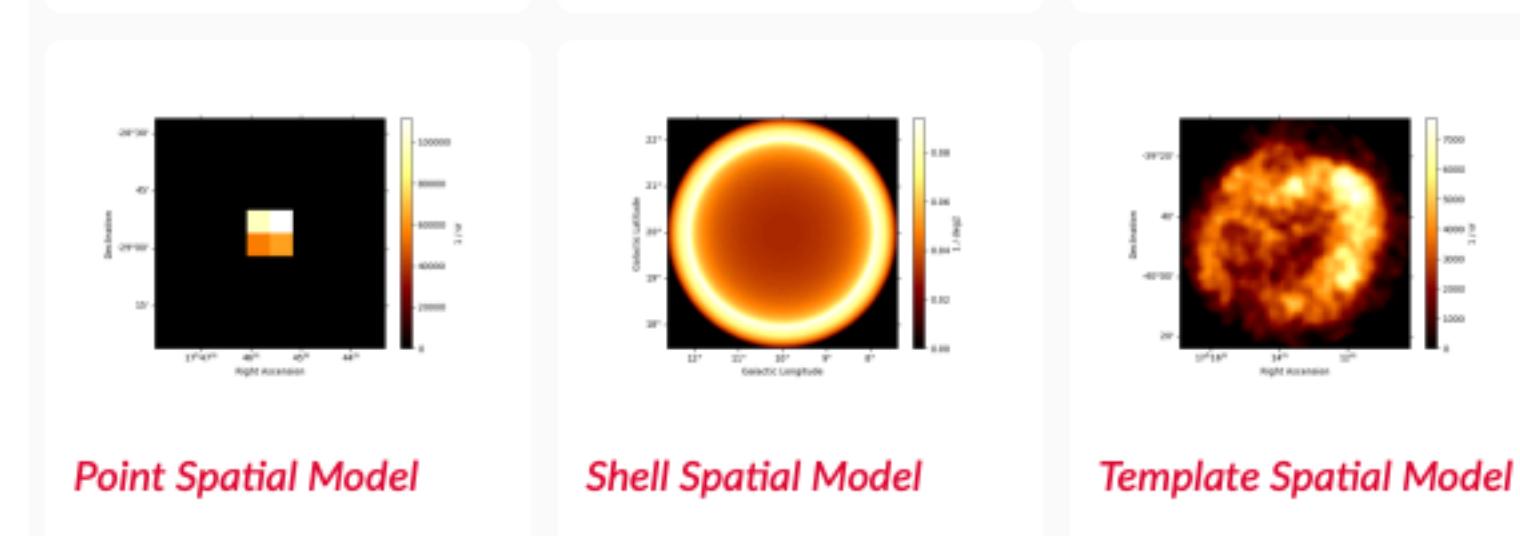
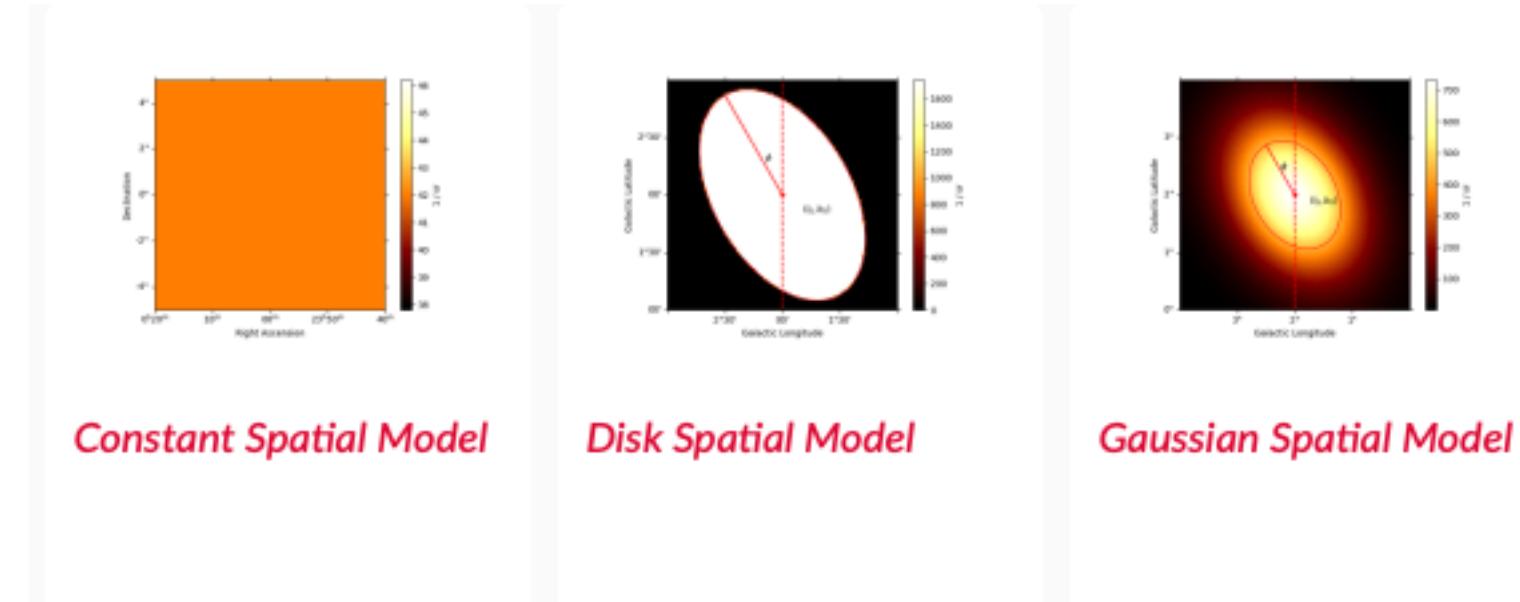
JOINT LIKELIHOOD FITTING



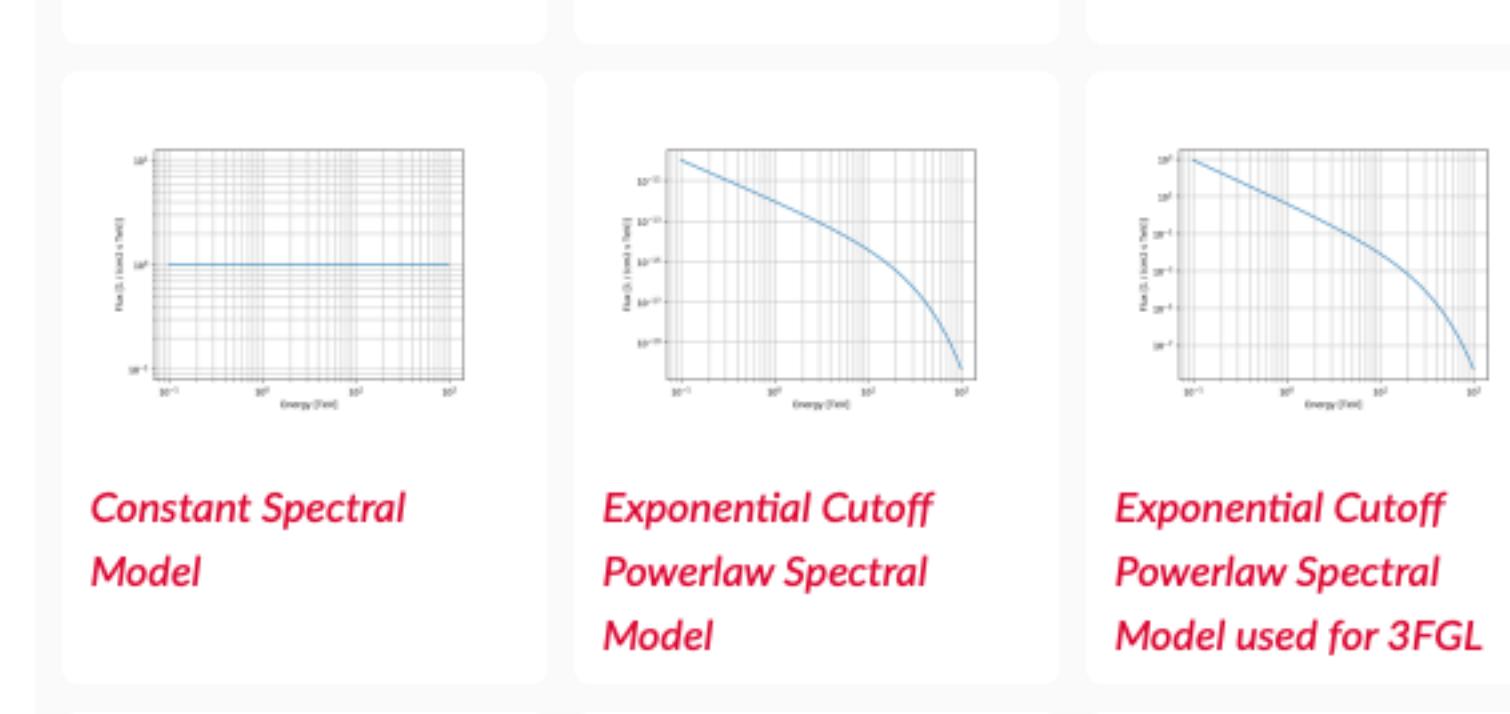
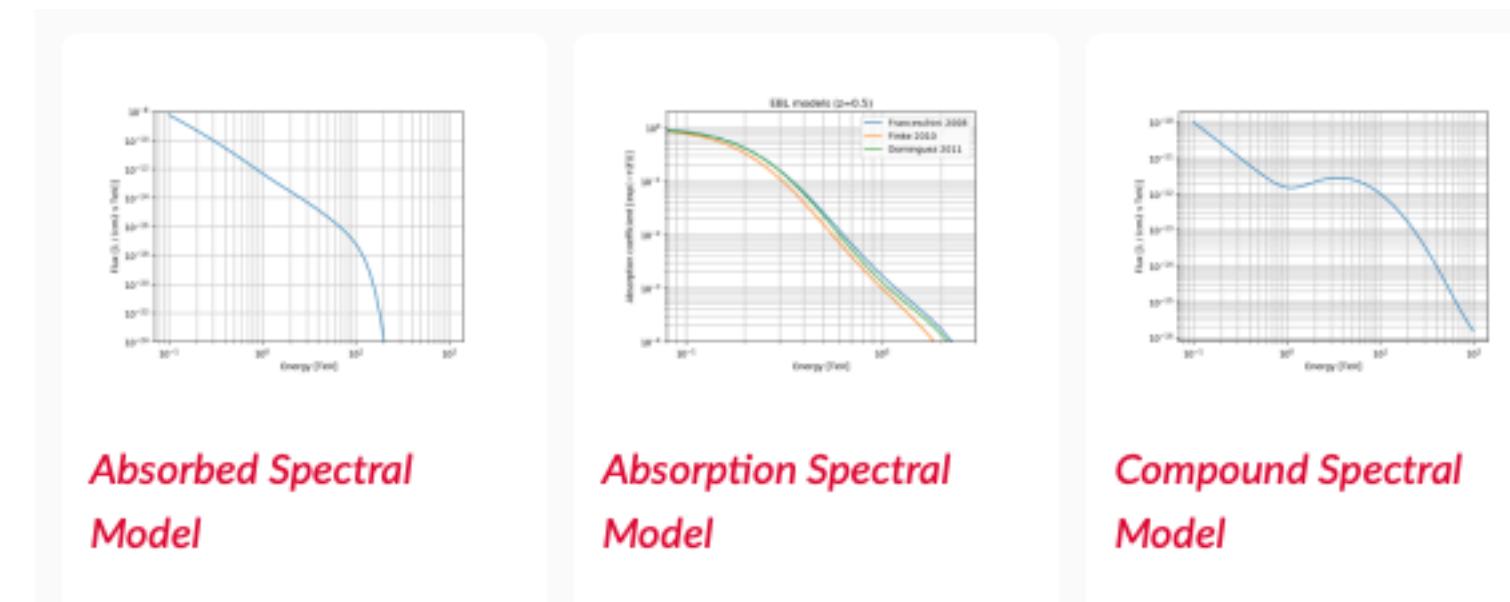
GAMMAPY.MODELING.MODELS

- Model gallery: <https://docs.gammapy.org/0.17/modeling/gallery/index.html#model-gallery>

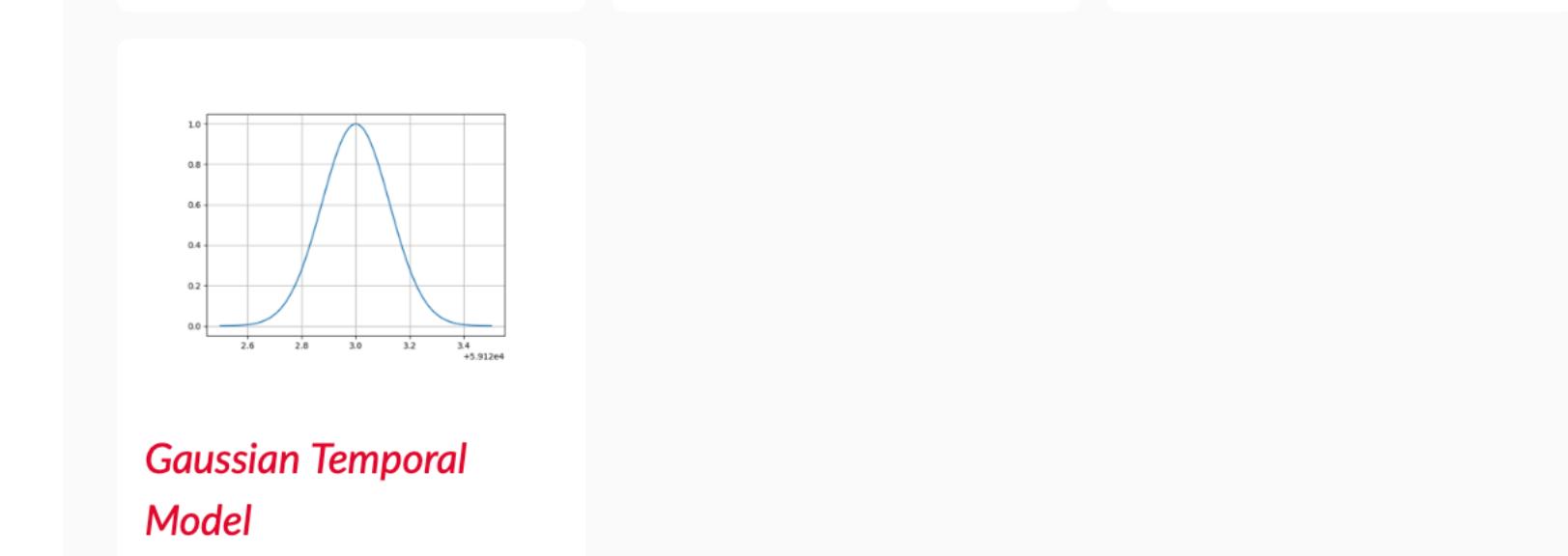
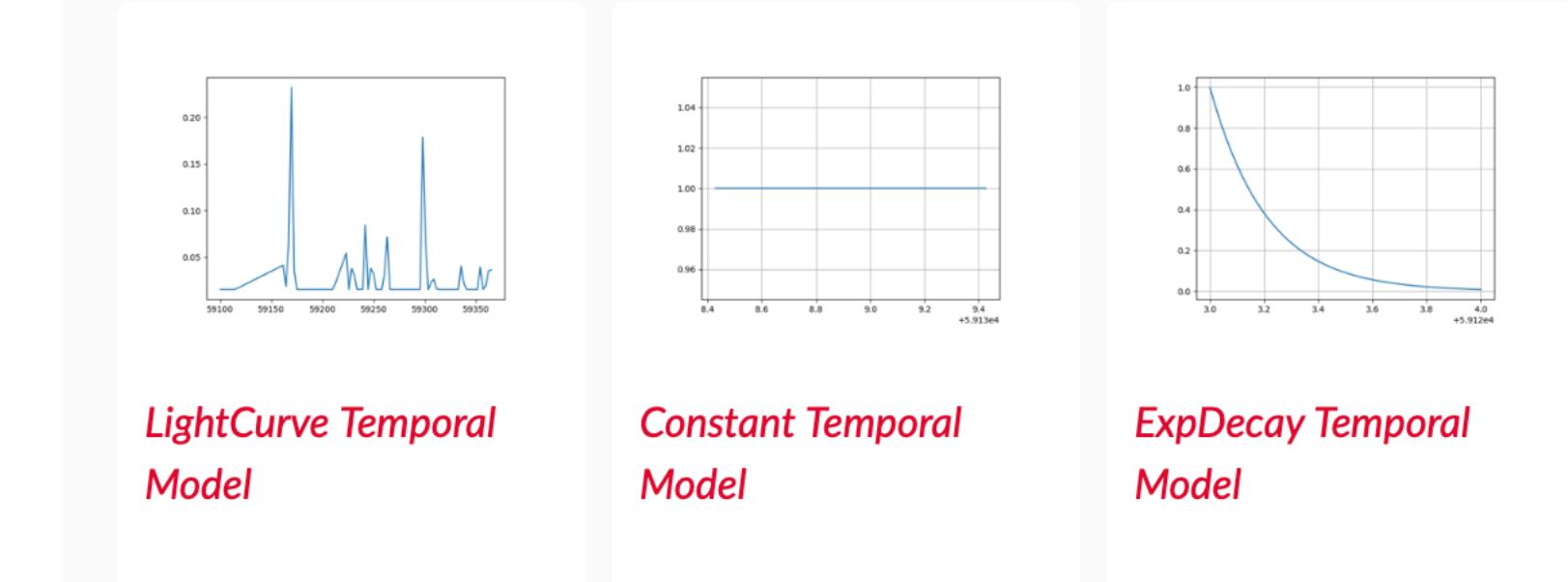
Spatial models



Spectral models...



Temporal models...



- Models tutorial: <https://docs.gammapy.org/0.17/notebooks/models.html>, includes example how to implement a custom model and register it for I/O
- Configure IRF application to model components using: `SkyModel(..., apply_irf={"psf": False})`

FITTING OF TEMPORAL MODELS

https://docs.gammapy.org/0.17/notebooks/light_curve_simulation.html

- In Gammapy v0.17 we added support to evaluate and fit temporal models to a list of datasets

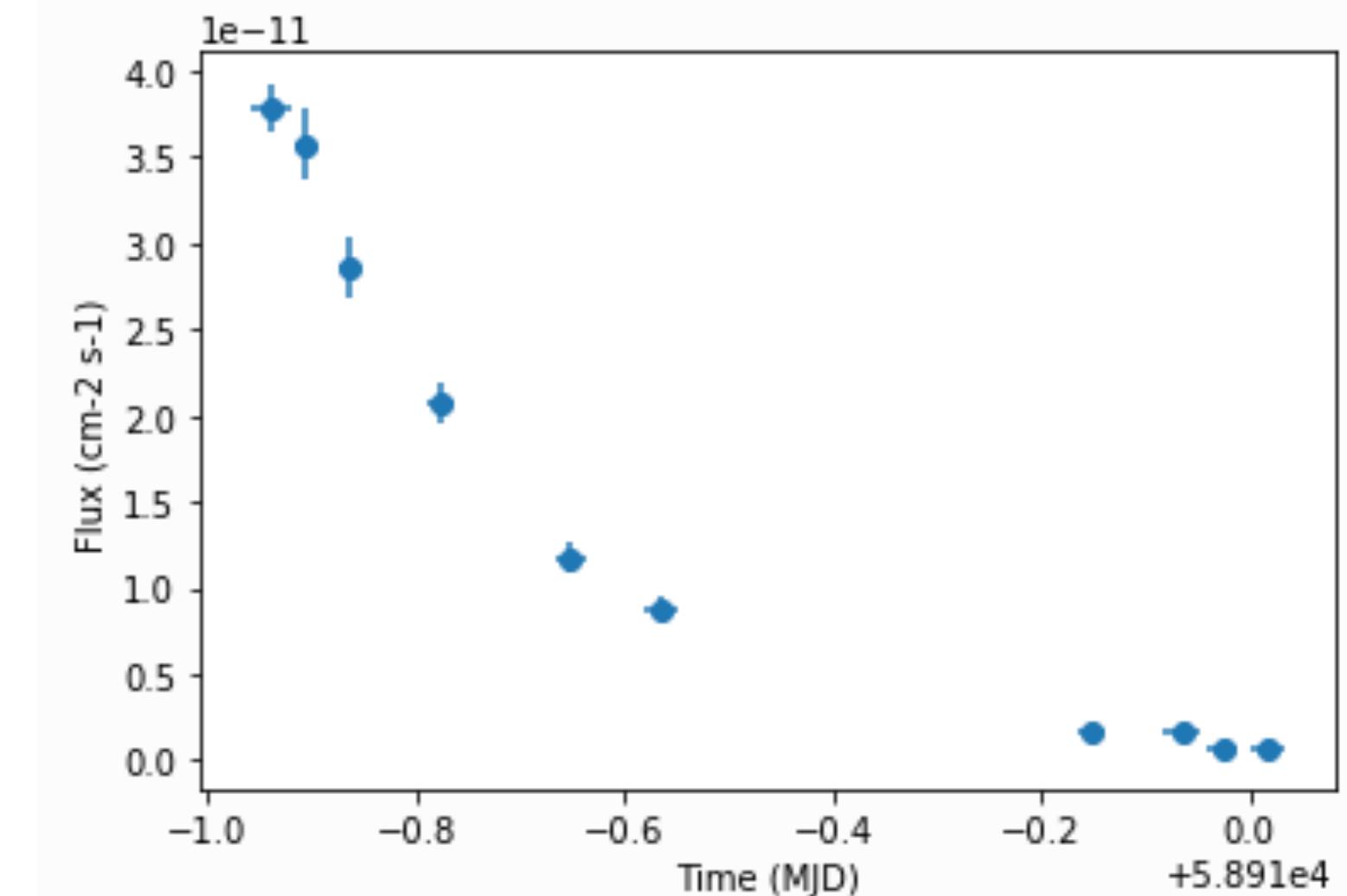
```
# Define the model:  
  
spectral_model1 = PowerLawSpectralModel(  
    index=2.0, amplitude="1e-12 cm-2 s-1 TeV-1", reference="1 TeV"  
)  
temporal_model1 = ExpDecayTemporalModel(t0="10 h", t_ref=gti_t0.mjd * u.d)  
model = SkyModel(  
    spectral_model=spectral_model1,  
    temporal_model=temporal_model1,  
    name="model-test",  
)
```

Define an additional temporal model component...

```
for i in range(n_obs):  
    obs = Observation.create(  
        pointing=pointing,  
        livetime=lvtm[i],  
        tstart=tstart[i],  
        irfs=irfs,  
        reference_time=gti_t0,  
)  
    empty_i = empty.copy(name=f"dataset_{i}")  
    maker = SpectrumDatasetMaker(selection=["aeff", "background", "edisp"])  
    dataset = maker.run(empty_i, obs)  
    dataset.models = model_simu  
    dataset.fake()  
    datasets.append(dataset)
```

Simulate datasets in a loop...

Compute lightcurve...



Available temporal models

Registry

`ConstantTemporalModel`

`LightCurveTemplateTemporalModel`

`ExpDecayTemporalModel`

`GaussianTemporalModel`

CONVENIENCE COVARIANCE HANDLING

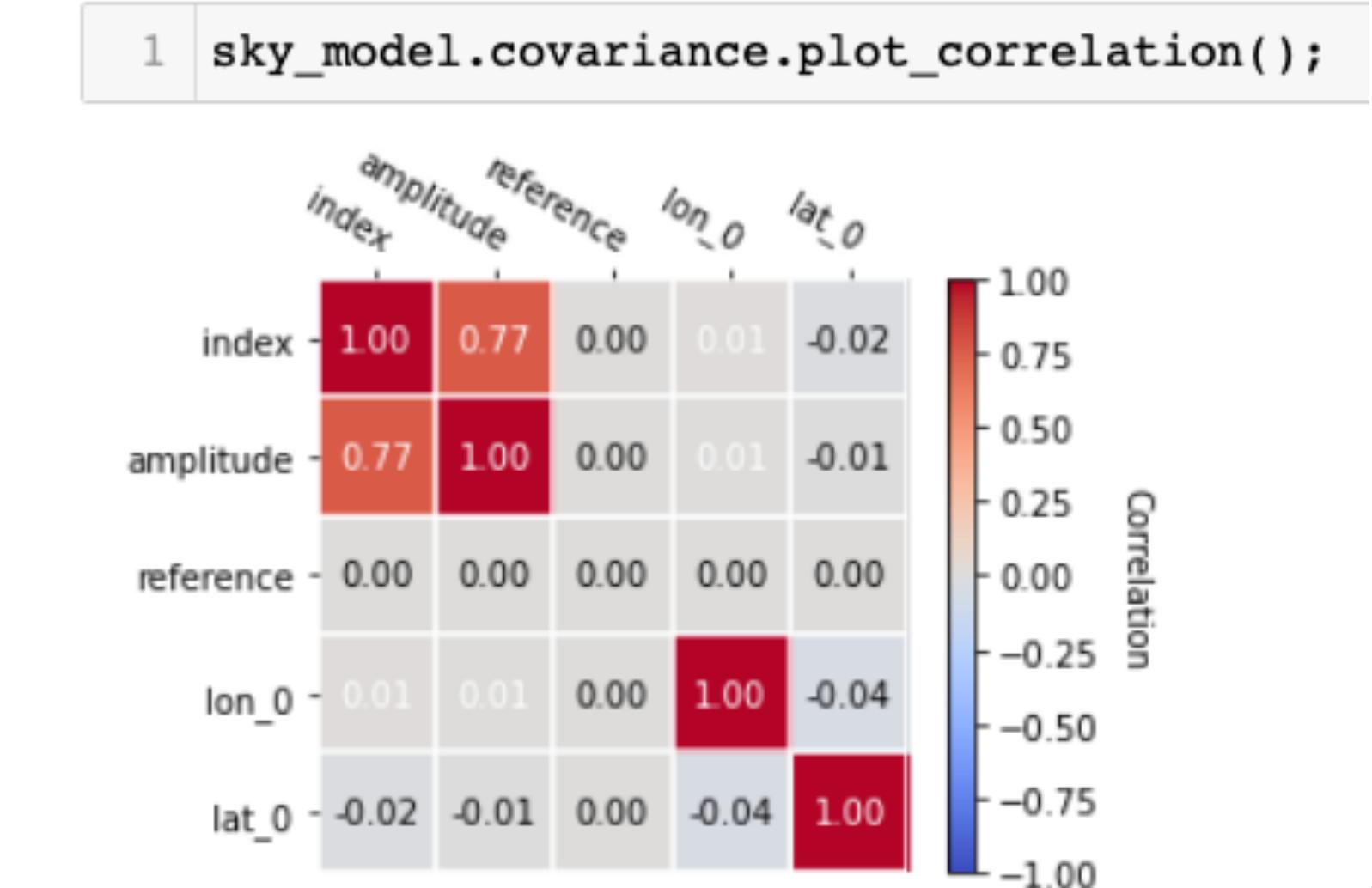
- After running `Fit.run()` or `Fit.covariance()` the covariance is now available on the fitted model
- Sub-covariance are automatically tracked to sub-model components and can be accessed like:

```
print(models.covariance)
print(models["my-model"].covariance)
print(models["my-model"].spectral_model.covariance)
```

- Write / read covariance to file:

```
models.write_covariance("covariance.dat")
models.read_covariance("covariance.dat")
```

Plot correlation between parameters

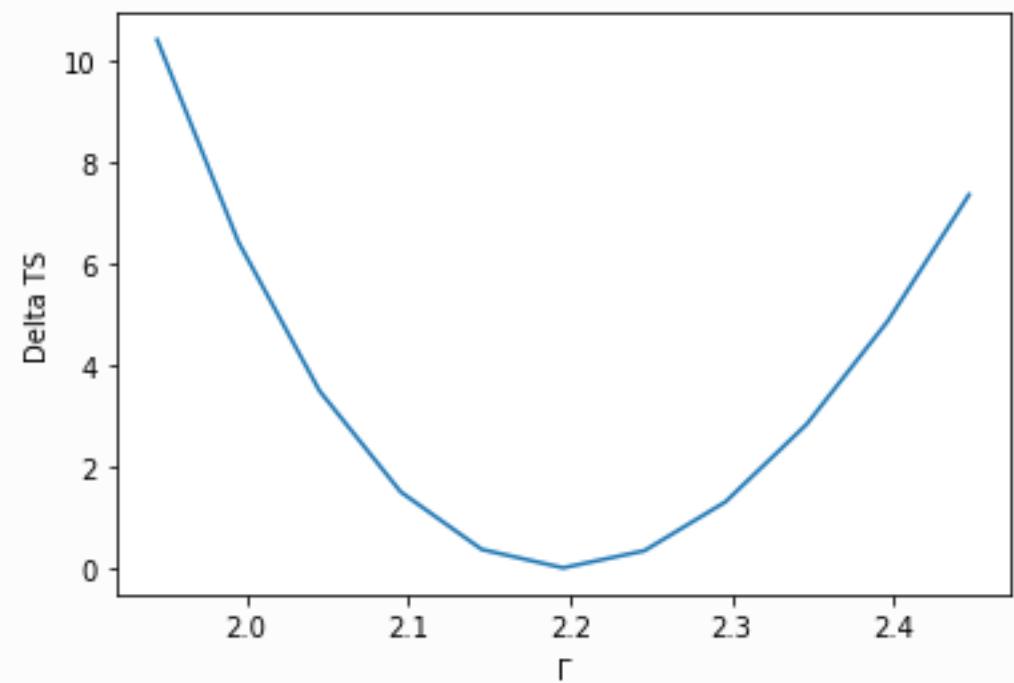


MODELLING

<https://docs.gammapy.org/0.17/notebooks/modeling.html>

- Any fitting / modelling of data requires **careful checking of fit convergence and results!**
- Demonstrates how to use different optimisers & fitting backends, such as iminuit, scipy and sherpa
- Shows how to compute likelihood profiles and contours for fit diagnostics...

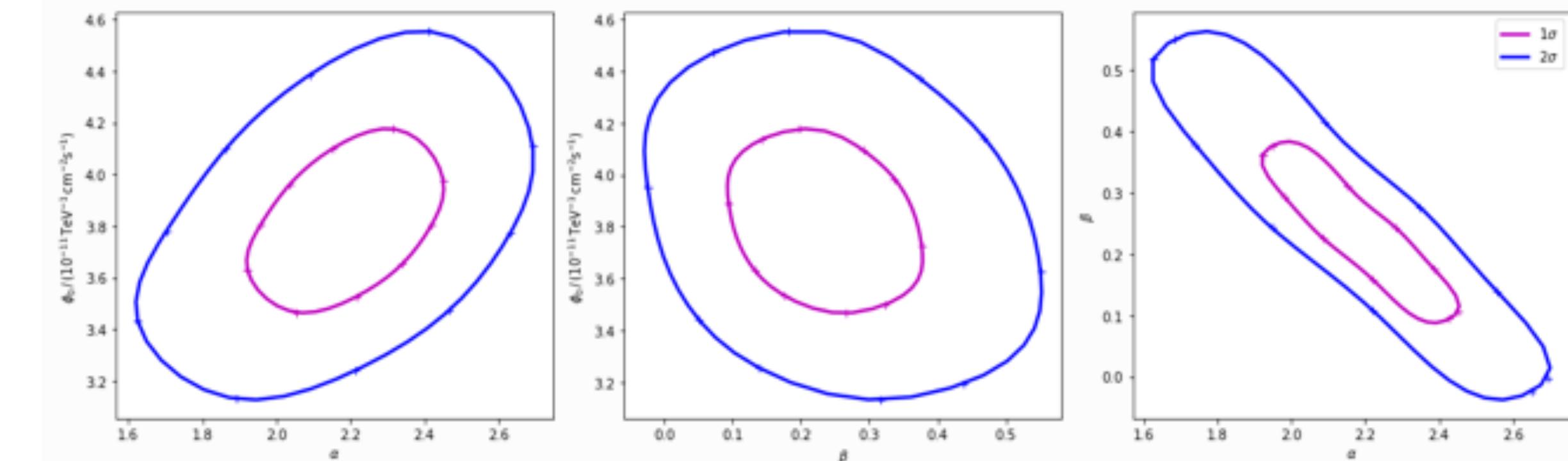
Create likelihood profiles



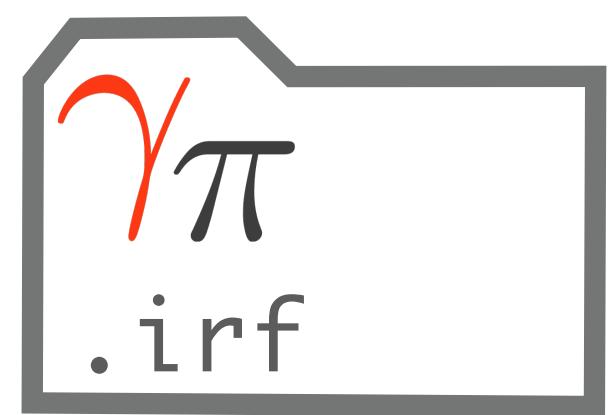
```
profile = fit.stat_profile(parameter="alpha")

total_stat = result_minuit.total_stat
plt.plot(profile["values"], profile["stat"] - total_stat)
plt.xlabel(r"$\Gamma$")
plt.ylabel("Delta TS");
```

Create likelihood contours



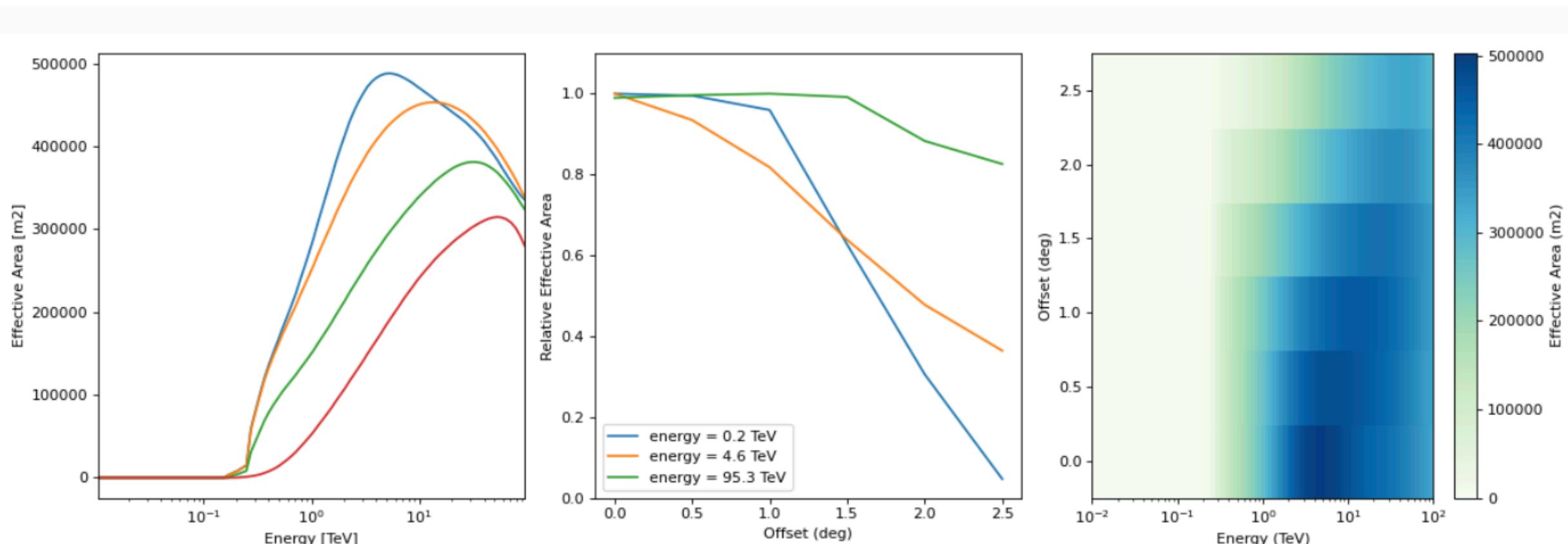
```
for par_1, par_2 in combinations(["alpha", "beta", "amplitude"], r=2):
    contour = fit.minos_contour(
        result.parameters[par_1],
        result.parameters[par_2],
        numpoints=npoints,
        sigma=sigma,
    )
```



Read and plot effective area...

```
"""Plot an effective area from the HESS DL3 data release 1."""
import matplotlib.pyplot as plt
from gammapy.irf import EffectiveAreaTable2D

filename = "$GAMMAPY_DATA/hess-dl3-dr1/data/hess_dl3_dr1_obs_id_020136.fits.gz"
aeff = EffectiveAreaTable2D.read(filename, hdu="AEFF")
aeff.peek()
plt.show()
```

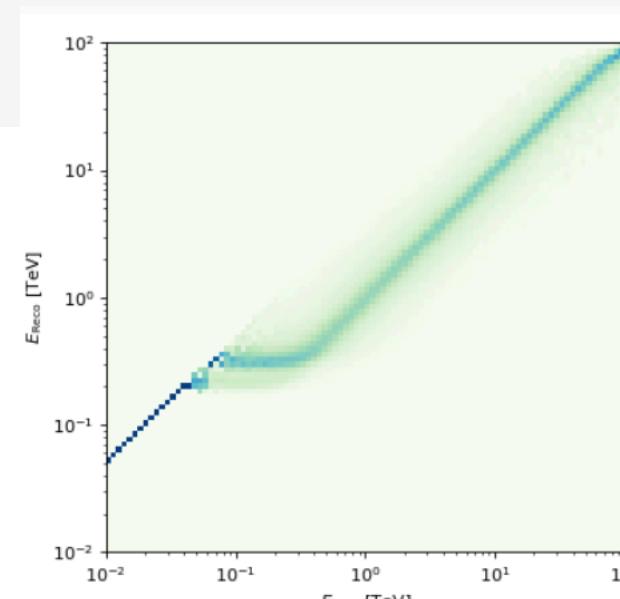


- Handling of instrument response functions such as: effective area, point spread function (PSF), energy resolution (EDisp) and instrumental background....
- Handling of position dependent IRFs via `PSFMap` and `EDispKernelMap`
- Mostly visualisation, interpolation and serialisation from and to FITS

Read and plot energy resolution....

```
"""Plot an energy dispersion from the HESS DL3 data release 1."""
import matplotlib.pyplot as plt
from gammapy.irf import EnergyDispersion2D

filename = "$GAMMAPY_DATA/hess-dl3-dr1/data/hess_dl3_dr1_obs_id_020136.fits.gz"
edisp = EnergyDispersion2D.read(filename, hdu="EDISP")
edisp.peek()
plt.show()
```



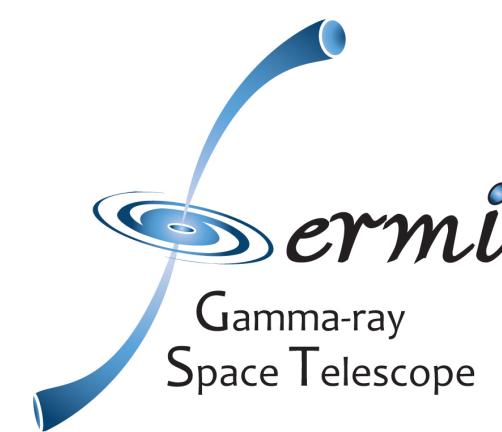
GAMMAPY.CATALOG

<https://docs.gammapy.org/0.17/tutorials/catalog.html>

```
1 from gammapy.catalog import SourceCatalog3FHL
2
3 fermi_3fhl = SourceCatalog3FHL()
4 crab_3fhl = fermi_3fhl["Crab Nebula"]
5
6 crab_3fhl.spectral_model.energy_flux("1 TeV", "10 TeV").to("erg cm-2 s-1")
```

$$1.112798 \times 10^{-10} \frac{\text{erg}}{\text{s cm}^2}$$

Source_Name	RAJ2000	DEJ2000	GLON	GLAT	Conf_95_SemiMajor	Conf_95_SemiMinor	Conf_95_PosAng	ROI_num	Signif_Avg	Pivot_Energy	Flux_Density
	deg	deg	deg	deg	deg	deg	deg			GeV	1 / (cm ² GeV s)
bytes18	float32	float32	float32	float32	float32	float32	float32	int16	float32	float32	float32
3FHL J0001.2-0748	0.3107	-7.8075	89.0094	-67.3118	0.0424	0.0424	nan	64	5.362	23.73	5.3174e-13
3FHL J0001.9-4155	0.4849	-41.9303	334.1216	-72.0697	0.1018	0.1018	nan	429	5.638	28.42	5.4253e-13
3FHL J0002.1-6728	0.5283	-67.4825	310.0868	-48.9549	0.0357	0.0357	nan	386	8.470	20.82	1.2062e-12



2FHL
3FHL
3FGL
4FGL

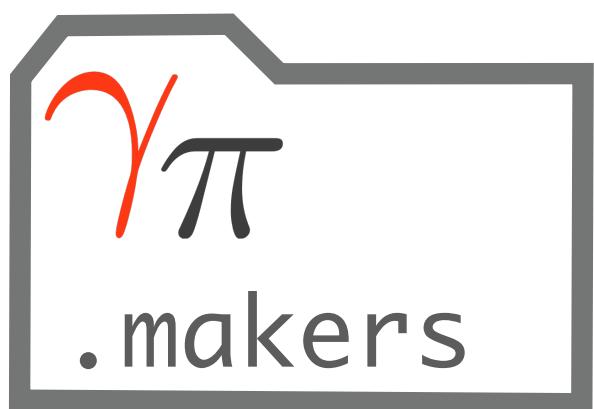


“An open data collection
and source catalog for VHE
gamma-ray astronomy”

- Convenience access to existing gamma-ray catalogs, from Fermi-LAT, HESS, HAWC, etc.
- Uniform access API and translation of the info to Gammapy objects, such as models, flux points, light-curves

GAMMAPY.MAKERS

<https://docs.gammapy.org/0.17/makers/index.html>



- A uniform “Maker” API for data reduction:

```
from gammapy.datasets import MapDataset
from gammapy.makers import MapDatasetMaker, SafeMaskMaker
from gammapy.data import DataStore
from gammapy.maps import MapAxis, WcsGeom

data_store = DataStore.from_dir("$GAMMAPY_DATA/hess-dl3-dr1")
observations = data_store.get_observations([23523, 23592, 23526, 23559])

energy_axis = MapAxis.from_bounds(1, 10, nbins=11, name="energy", unit="TeV", interp="log")
geom = WcsGeom.create(skydir=(83.63, 22.01), axes=[energy_axis], width=5, binsz=0.02)

maker = MapDatasetMaker()
safe_mask_maker = SafeMaskMaker(methods=["aeff-default", "offset-max"], offset_max="3 deg")

stacked = MapDataset.create(geom)

for obs in observations:
    cutout = stacked.cutout(obs.pointing_radec, width="6 deg")
    dataset = maker.run(cutout, obs)
    dataset = safe_mask_maker.run(dataset, obs)
    stacked.stack(dataset)

print(stacked)
```

Makers are created and configured first...

...and then executed using .run(dataset, obs) in a chain in the data reduction loop.

Available makers:

ReflectedRegionsBackgroundMaker [...]
AdaptiveRingBackgroundMaker (r_in, r_out_max, ...)
FoVBackgroundMaker ([method, exclusion_mask])
PhaseBackgroundMaker (on_phase, off_phase)
RingBackgroundMaker (r_in, width[, ...])
SpectrumDatasetMaker [...]
MapDatasetMaker ([background_oversampling, ...])
SafeMaskMaker ([methods, aeff_percent, ...])

<https://docs.gammapy.org/dev/makers/index.html#classes>

- Example for reflected regions: <https://docs.gammapy.org/dev/makers/reflected.html#using-the-reflected-background-estimator>
- Also check tutorials...

EVENT SAMPLING

https://docs.gammify.org/0.17/tutorials/event_sampling.html

- In Gammapy v0.17 we introduced event sampling via the `MapDatasetEventSampler` object, which simulates a list of events for a given dataset comprising a sky model and reduced IRFs

Simulate events for multiple observations

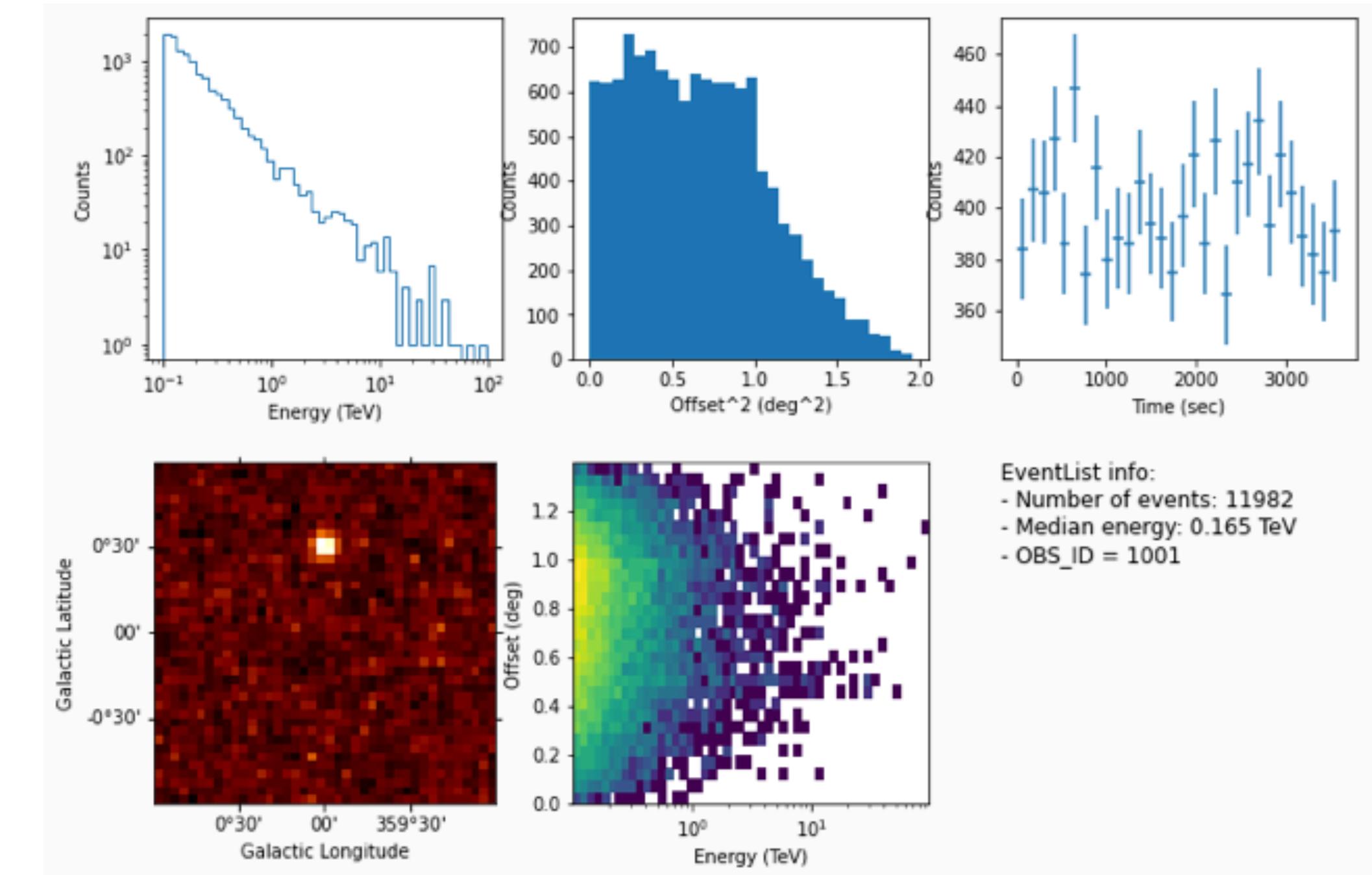
```
for idx, tstart in enumerate(tstarts):

    observation = Observation.create(
        obs_id=idx,
        pointing=POINTING,
        tstart=tstart,
        livetime=livetimes[idx],
        irfs=irfs,
    )

    dataset = maker.run(empty, observation)
    dataset.models.extend(models_pntpwl)

    sampler = MapDatasetEventSampler(random_state=idx)
    events = sampler.run(dataset, observation)
    events.table.write(
        f"./event_sampling/events_{idx:04d}.fits", overwrite=True
    )
```

Statistics of simulated events



- Event sampling is validated here: <https://github.com/gammify/gammify-benchmarks/tree/master/validation/event-sampling>

MISCELLANEOUS

- With `MapDataset.to_spectrum_dataset(region=)` there is the possibility to extract a spectrum dataset from a 3D datasets for further analysis, see https://docs.gammapy.org/0.17/notebooks/extended_source_spectral_analysis.html
- To unify the API between spectral and “cube” analysis we introduced `RegionGeom` / `RegionNDMap` classes, which bundle the region information with the spectral information and are compatible in API with the `Map` classes. So far not fully documented, but used in `SpectrumDataset` and `SpectrumDatasetOnOff`
- In Gammapy v0.16 we started to adapt the `Estimator` API to work with `Dataset` objects as well
- An example for a multi-instrument analysis tutorial combining HESS / Fermi-LAT and HAWC can be found here: https://docs.gammapy.org/0.17/notebooks/analysis_mwl.html
- Introduced unified statistics API with `CashCountsStatistics` and `WStatCountsStatistics` (represent standard Li & Ma solutions using cash and wstat fit statistics)

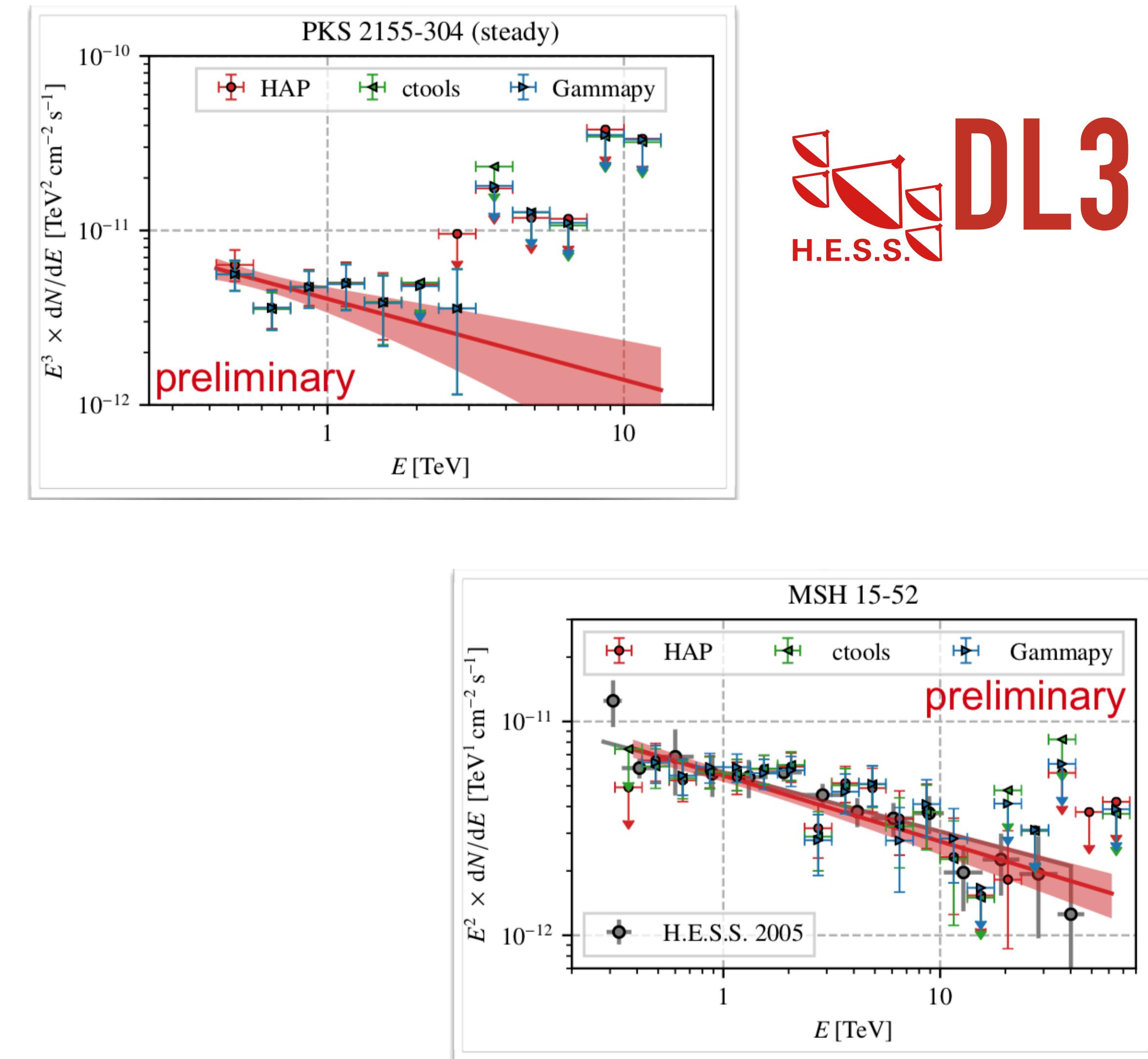
SCIENCE VERIFICATION

H.E.S.S. DL3 VALIDATION PAPER

- H.E.S.S public data released last year (<https://www.mpi-hd.mpg.de/hfm/HESS/pages/dl3-dr1/>)
- Validation of FITS data and open source tools Gammapy and ctools (lead by Lars Mohrmann from ECAP)
- Perfect reference analysis for a CI / benchmarking setup for the science tools
- Setup a science verification CI system: <https://github.com/gammapy/gammapy-benchmarks/tree/master/validation>

Source	Type	Size	Livetime	Period
Crab nebula	PWN	Point-like	1.9 h	2004
PKS 2155-304	AGN	Point-like	9.8 h	2006/2008
MSH 15-52	PWN	Small	9.1 h	2004
RX J1713.7-3946	SNR	Large	7.0 h	2004
Off data	-	-	20.7 h	2004/2005

For joint crab analysis see Cosimo's presentation...



SUMMARY

- Gammipy is a Python package for Gamma-Ray astronomy, based on Numpy, Astropy and Scipy
- It implements standard high level analysis use-cases for event based data, such as model fitting, flux points estimation, light-curve estimation, etc.
- It implements a joint-joint-likelihood API, that allows to combine data of multiple instruments
- **Gammipy v0.18 will be released on Oct 30th**, which will include a 98% final API. It will be the “release candidate” for v1.0. After v0.18 only bug fixes, docs improvements and minor clean up...
- Questions?

CONTACT AND SUPPORT

- Slack: gammapy.slack.com (quick questions, immediate help)
- Github issues: <https://github.com/gammipy/gammipy/issues>
(feature requests & bug reports)
- Gammipy mailing list (not used much...):
 - gammipy@googlegroups.com
 - Archive: <https://groups.google.com/forum/#!forum/gammipy>
- Gammipy dev meetings (technical meetings in case you'd like to contribute):
 - Every Friday 10am on Vibe
 - Regular Gammipy user calls, will be announced on the CTA mailing list

