# MEK4320 Energy in surface waves.

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## Two approaches

- Averaged fluxes and densities are calculated directly.
   This approach is given, albeit imprecisely, in leaflet sec. 2.4.
  - Important for physical insight. Should be done first.
- An energy equation is derived from the basic equations for momentum and mass conservation, together with the boundary conditions.
  - Gain: correct, general expressions for densities and fluxes are found, also including nonlinearity
  - Gain: existence and expressions for capillary energy follows from the basic equation.
  - Disadvantage: much algebra. In particular related to the surface tension.

We will only do one horizontal dimension.



## The role surface pressure, $p_a$

Application of a uniform surface pressure will lead to a constant addition of  $p_a$  to the pressure throughout the fluid, while the velocities and the surface elevation are unaffected.

In the energy balance  $p_a$  will enter only through the pressure work at surfaces. Looking at any volume,  $\Omega$ , with surface  $\Sigma$ , the effect of the work exerted by  $p_a$  is:

$$\int_{\Sigma} -p_{a}\vec{v}\cdot\vec{n}dA = \int_{\Omega} -p_{a}\nabla\cdot\vec{v}d\Omega = 0,$$

where Gauss' theorem is employed and we have invoked  $\nabla \cdot \vec{v} = 0$ . We exempt  $p_a$  in the subsequent calculations by putting it equal to zero

# Direct evaluation of densities and fluxes

#### The wave solution

The harmonic solution reads

$$\phi = -\frac{a\omega \cosh k(z+H)}{k \sinh kH} \cos(kx - \omega t), \quad \eta = a \sin(kx - \omega t),$$

where  $\omega$  and k fulfill the dispersion relation.

In the integration we mainly need simple integrals akin to

$$\int\limits_{x_1}^{x_1+\lambda}\sin^2(kx-\omega t)dx=\frac{1}{2}\lambda,$$

and

$$\int_{H}^{0} \cosh^{2}(k(z+H))dz = \frac{1}{2}(\frac{1}{k}\cosh kH\sinh kH + H).$$

### Calculation of energy densities.

#### Accuracy of basic solution

Our wave mode solution is linear in the amplitude (a).  $\Rightarrow$  Only approximate energy densities may be calculated. Linear modes yield energy densities of order  $a^2$  correctly to leading order.

#### **Energy densities**

Densities is averaged over a wavelength and measured per crest length.

- (i) Kinetic energy,  $E_k$ .
- (ii) Potential energy due to gravity,  $E_p^t$ .
- (iii) Potential energy due to capillary effects,  $E_p^k$ .

Notation: (K2.1), say, is equation 2.1 in leaflet. Indices denote partial derivatives.



### Averaged kinetic energy

The full expression for averaged kinetic energy:

$$E_k = \frac{1}{\lambda} \int_{x_1}^{x_1 + \lambda} \int_{-H}^{\eta} \frac{1}{2} \rho \vec{\mathbf{v}}^2 dz dx \tag{1}$$

where the result is independent of  $x_1$  due to periodicity. The integral in z can be split in two. One from the bottom to z=0 and anther from z=0 to the free surface  $(z=\eta)$ . Since the integrand is quadratic in a the latter will be of order  $O(a^3)$  which may be neglected in comparison with the first. Insertion of the linear wave mode followed by straightforward integration yields

$$E_{k} = \frac{1}{4}\rho g a^{2} (1 + \frac{\sigma k^{2}}{\rho g}) + O(a^{3})$$
 (2)

## Averaged potential energy due to gravity

In contrast to the leaflet<sup>(a)</sup> we define the zero level at z = 0.

$$\int_{x_1}^{x_1+\lambda} \int_{-H}^{\eta} \rho g z dz dx = -\frac{1}{2} \rho g \lambda H^2 + \int_{x_1}^{x_1+\lambda} \frac{1}{2} \rho g \eta^2 dx$$
 (3)

The first term in the rightmost expression is the equilibrium energy, while the second is the increase due to the wave. Hence

$$E_p^t = \frac{1}{\lambda} \int_{x_1}^{x_1 + \lambda} \frac{1}{2} \rho g \eta^2 dx = \frac{1}{4} \rho g a^2 + O(a^3)$$
 (4)

<sup>a</sup> The odd definition below (K2.26) is necessary to mask the erroneous omission of advection of potential energy in the leaflet.



## Potential energy due to surface tension

Stretching of surface: work is done against the surface tension.

Shrinking of surface: energy may be retrieved.

Simple flat surface  $\Rightarrow$ 

potential energy= $\sigma A + \mathrm{const.}$ , where A is surface area.

Assume this holds for a general surface (c) We study energy per width; Area per width = arc-length  $\Rightarrow$ 

$$\int_{x_1}^{x_1+\lambda} \sigma \sqrt{1+\eta_x^2} dx = \sigma \lambda + \int_{x_1}^{x_1+\lambda} \frac{1}{2} \sigma \eta_x^2 dx + O(a^4)$$
 (5)

First term is the equilibrium part; the wave part becomes

$$E_p^k = \frac{1}{\lambda} \int_{x_1}^{x_1 + \lambda} \frac{1}{2} \sigma \eta_x^2 dx + O(a^4) = \frac{1}{4} \sigma k^2 a^2 + O(a^3)$$
 (6)

<sup>(</sup>c) Shown later and in leaflet through derivation of energy equation.



### Equipartition of energy

The total potential energy

$$E_p = E_p^t + E_p^k = \frac{1}{4} \rho g a^2 (1 + \frac{\sigma k^2}{\rho g}) + O(a^3),$$
 (7)

which is the same as the average kinetic energy

$$E_k = E_p$$

In average potential and kinetic energies are equal.

This is a general property for linear waves.

It is also true for linear oscillators, such as a pendulum in the gravity field.

## Energy transport through a vertical cross-section

#### Advection of kinetic and gravity-potential energy

Transport through interface element  $d\sigma$ :

 $\gamma \vec{v} \cdot \vec{n} d\sigma$ ,  $\underline{\gamma}$  is density (i.e.  $\rho gz$ ) per volume,  $\vec{n}$  unit normal.

Here  $\vec{n} = \vec{i}$ , transport per width:  $d\sigma \Rightarrow dz$ .

#### Advection of capillary-potential energy

Due to transport of material surface with u at surface.

#### Effect of pressure work

Pressure times normal velocity, here: pudz

#### Effect of surface tension work

Force due to surface tension  $\cdot \vec{v}$  at the surface.



## Three contribution to the energy flux

Transport of potential energy from gravitation, kinetic energy and the effect of pressure work is combined.

$$f_1 = \int_{-H}^{\eta} (\frac{1}{2} \rho \vec{v}^2 + \rho gz + p) u dz$$
 (8)

The expression simplifies by means of Euler's pressure equation<sup>(e)</sup> (below (K2.7) with  $p_a = 0$ )

$$f_1 = -\int_{-H}^{\eta} (\rho \phi_t - p_a) \phi_x dz = -\int_{-H}^{0} \rho \phi_t \phi_x dz + O(a^3)$$
 (9)

The averaged flux becomes

$$F_1 = -\frac{1}{T} \int_{t}^{t+T} \int_{-H}^{0} \rho \phi_t \phi_x \mathrm{d}z \mathrm{d}t + O(a^3)$$
 (10)

## The energy flux due to capillary effects

#### Energy transport due to surface tension

 $f_2 = advection + effect of surface tension$ 

- advection =  $\sigma \frac{ds}{dx} u = \sigma \times \text{transport of surface length.}$
- effect of surface tension =  $\sigma \vec{s} \cdot \vec{v}$ .

Here s is arc-length of surface and  $\vec{s}$  is unit tangent. Combined

$$f_2 = \sigma[(1+\eta_x^2)^{\frac{1}{2}}u - (1+\eta_x^2)^{-\frac{1}{2}}(u+\eta_x w)]_{z=\eta} = -\sigma\eta_x\phi_z|_{z=0} + O(a^3)$$
(11)

Averaged:

$$F_2 = -\frac{1}{T} \int_{t}^{t+T} f_2 \mathrm{d}t$$

### The group velocity

We find  $F_1$  and  $F_2$  by inserting the wave mode solution and do the straightforward integration.

#### Recognition of group velocity

With  $F = F_1 + F_2$  we may write

$$F = c_g E \tag{12}$$

where  $c_{g} = d\omega/dk$  is the group velocity.

The relation (12) is a general property for linear wave modes.

 $c_g$  is interpreted as the celerity of wave energy



# Conservation laws derived from the PDEs

### Conservation laws, general.

Conservation of energy for a volume  $\Omega$  may be expressed as

$$\frac{\partial}{\partial t} \int_{\Omega} \varepsilon d\Omega = -\int_{\Sigma} \vec{q} \cdot \vec{n} dA$$
 (13)

where  $\Sigma$  is the surface that envelopes  $\Omega$ ,  $\epsilon$  is the energy density per volume and  $\vec{q}$  is the energy flux density.

#### Side track

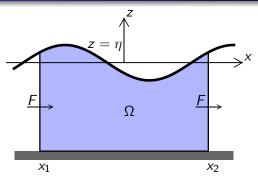
 $\Omega$  fixed in time (not true in our case!): Division by the volume, use of Gauss' theorem on the right hand side, followed by  $\Omega \to 0$ 

$$\varepsilon_t = -\nabla \cdot \vec{q},\tag{14}$$

which is the differential form.



## Energy conservation for waves; special case.



The integrated energy equation for  $\Omega$ 

$$\frac{d}{dt}(e_p^{(t)} + e_k + e_p^{(k)}) = F|_{x=x_1} - F|_{x=x_2}$$
 (15)

 $e_p^{(t)}$ ,  $e_k$ ,  $e_p^{(k)}$  are obtained by integrating the energy densities in  $\Omega$ . Objective: Produce (15) from basic equations, manipulations and interpretations. Recognize the expressions from the direct

Euler's equation of motion:

$$\rho \vec{\mathbf{v}}_t = -\nabla \mathbf{p} - \rho \nabla (\frac{1}{2} \vec{\mathbf{v}}^2) - \rho \nabla \Phi \tag{16}$$

where the convective term is simplified due to  $\nabla \times \vec{v} = 0$ . We multiply (16) by  $\vec{v}$  and integrate in  $\Omega$ . Gauss' theorem and  $\rho = \text{const.} \Rightarrow$ 

$$\int_{\Omega} \left(\frac{1}{2}\rho \vec{v}^{2}\right)_{t} d\Omega = -\int_{r} \left(-p + \frac{1}{2}\rho \vec{v}^{2} + \rho \Phi\right) \vec{v} \cdot \vec{n} ds \qquad (17)$$

$$(i) \quad (ii) \quad (iii)$$

where r is the union of the free surface and the vertical cross sections at  $x_1$  and  $x_2$  (the contribution from the bottom is zero). On the left hand side of (17) we recognize the kinetic energy.

Temporal derivation may be moved outside integral by Leibniz' rule (only the surface is time dependent):

$$\frac{d}{dt} \int_{\Omega} G d\Omega = \int_{\Omega} \frac{\partial G}{\partial t} d\Omega + \int_{x_1}^{x_2} G|_{z=\eta} \vec{v} \cdot \vec{n} ds$$
 (19)

Here

$$\int_{\Omega} (\frac{1}{2}\rho \vec{v}^2)_t d\Omega = \frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} (\frac{1}{2}\rho \vec{v}^2) d\Omega - \int_{x_1}^{\infty} (\frac{1}{2}\rho \vec{v}^2)|_{z=\eta} \vec{v} \cdot \vec{n} ds \quad (20)$$

and

$$\frac{d}{dt} \int_{\Omega} \Phi d\Omega = \int_{x_1}^{x_2} \Phi|_{z=\eta} \vec{v} \cdot \vec{n} ds, \qquad (21)$$

since 
$$\frac{\partial \Phi}{\partial t} = 0$$
.



#### Equation (17) reorganized

Last term (20) and (ii) cancel out  $(21) \Rightarrow$  (iii) = time derivative of total potential energy in the gravity field.

$$\frac{d}{dt}(e_p^{(t)} + e_k) = -\int_{x_1}^{x_2} p\vec{v}|_{z=\eta} \cdot \vec{n} ds + I_1 - I_2$$
 (22)

where p is the pressure under the surface and

$$e_k = \int_{\Omega} \frac{1}{2} \rho \vec{v}^2 d\Omega \qquad e_p^{(t)} = \int_{\Omega} \rho \Phi d\Omega$$
 (23)

$$I_{i} = \int_{-H}^{\eta} \{ (p + \frac{1}{2}\rho\vec{v}^{2} + \rho\Phi)u \}|_{x=x_{i}} dz$$
 (24)

**ENERGY** 

All in place, relative to direct approach, save the capillary contributions (from first term on r.h.s of (22))

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### Pure gravity waves, results

Expression for  $I_i$  contains, in sequence:

- Effect of pressure in vertical transect
- Advection of kinetic energy in vertical transect
- Advection of potential energy in gravity field in vertical transect

Hence, when  $\sigma = 0$  we have I = F.

Moreover, when  $\sigma = 0$  the first term on r.h.s. of (22) vanishes.

Hence, we have shown the simplified version of (15)

$$\frac{d}{dt}(e_p^{(t)} + e_k) = F|_{x=x_1} - F|_{x=x_2}$$

First term on r.h.s. of (22) can be rewritten to yield  $e_p^{(k)}$  term as well as contribution to F.

Details follow after "EXTRA" flag, but will not be required.



Definition of vertically integrated energy densities,  $E_k^{(i)}$  and  $E_p^{(i)}$ 

$$e_k = \int_{\Omega} \frac{1}{2} \rho \vec{v}^2 d\Omega = \int_{x_1}^{x_2} E_k^{(i)} dx, \text{ where } E_k^{(i)} = \int_{-H}^{\eta} \frac{1}{2} \rho \vec{v}^2 dz$$
 (25)

$$e_{\rho}^{(t)} = \int_{\Omega} \rho \Phi d\Omega = \int_{x_1}^{x_2} E_{\rho}^{(i)} dx, \quad \text{where} \quad E_{\rho}^{(i)} = \int_{-H}^{\eta} \rho \Phi dz \quad (26)$$

F (I) on preceding slides are vertically integrated fluxes, now renamed  $F^{(i)}$ . Dividing (15) by  $x_2-x_1$  and forming the limit  $x_2-x_1\to 0$  we obtain the depth integrated energy equation

$$\frac{\partial}{\partial t} (E_p^{(i)} + E_k^{(i)}) = -\frac{\partial F^{(i)}}{\partial x}$$
 (27)

The averaged quantities,  $E_k$  etc, may then be obtained by inserting the wave mode solution in  $E_k^{(i)}$  etc., removing the equilibrium contribution, make the appropriate approximations and average over t or x.

## **EXTRAS**

### Capillary effects

Last step: recasting the surface integral that includes p. Dynamic boundary condition at  $z=\eta$ 

$$p\vec{n} = -\sigma \frac{\partial \vec{s}}{\partial s} \Rightarrow p = -\frac{\sigma \eta_{xx}}{(1 + \eta_x^2)^{\frac{3}{2}}}$$
 (28)

Integration by parts

$$-\int_{x_{1}}^{x_{2}} p\vec{v} \cdot \vec{n} ds = \int_{x_{1}}^{x_{2}} \sigma \vec{v} \cdot d\vec{s} = (\sigma \vec{v} \cdot \vec{s})|_{x_{1}}^{x_{2}} - \int_{x_{1}}^{x_{2}} \sigma \vec{s} \cdot d\vec{v}$$
 (29)

First term on r.h.s. is the effect the surface tension.

Last term must correspond to the combination of temporal change and advection of potential capillary energy. We must unravel them.

We must identify the following as a part of the last term in (29)

#### Capillary energy as defined previously

$$e_p^{(k)} \equiv \sigma L = \int_{x_1}^{x_2} \sigma ds = \int_{x_1}^{x_2} \sigma s_x dx, \text{ where } s_x = \sqrt{1 + \eta_x^2}$$
 (30)

Time differentiation

$$\frac{\mathrm{d}e_{p}^{(k)}}{\mathrm{d}t} = \sigma \frac{\mathrm{d}L}{\mathrm{d}t} = \int_{x_{1}}^{x_{2}} \frac{\sigma \eta_{x} \eta_{xt}}{s_{x}} dx \tag{31}$$

#### Manipulations on last term of (29)

Definitions:  $\hat{u} = u(x, \eta, t)$ ,  $\hat{w} = w(x, \eta, t)$ 

Kinematic boundary condition:  $\hat{w} = \eta_t + \hat{u}\eta_x \Rightarrow \hat{w}_x = \eta_{xt} + ...$ 

Substitution  $d\vec{v} = (\vec{\imath} \hat{u}_x + \vec{k} \hat{w}_x) dx$  and rearrangement<sup>f</sup>

$$\int_{x_1}^{x_2} \sigma \vec{s} \cdot d\vec{v} = \int_{x_1}^{x_2} \sigma \left( \frac{\eta_x \eta_{xt}}{s_x} + \frac{\partial}{\partial x} (s_x \hat{u}) \right) dx$$
$$= \frac{\mathrm{d}e_p^{(k)}}{\mathrm{d}t} + (\hat{u}\sigma s_x)|_{x_1}^{x_2}$$

where the last term is advection of capillary energy.

<sup>f</sup> Alternative:  $\vec{s} \cdot d\vec{v}$  can be interpreted as the extension of a material surface segment. Temporal change of L is achieved through taking the advection of arc-length  $(s_x \hat{u})$  out of  $\Omega$ .

#### Total fluxes

If everything is collected we retrieve (15) with the energy flux:

$$F = \int_{-H}^{\eta} \{ (p + \frac{1}{2}\rho\vec{v}^2 + \rho\Phi)u \} dz + \sigma u|_{z=\eta} \frac{\partial s}{\partial x} - \sigma \vec{v} \cdot \vec{s}$$
 (32)

By means of Euler's pressure equation F is simplified further

$$F = -\int_{-H}^{\eta} \phi_t \phi_x dz + \sigma u|_{z=\eta} \frac{\partial s}{\partial x} - \sigma \vec{v} \cdot \vec{s}$$
 (33)

Insertion of linear wave mode reproduces previous results.