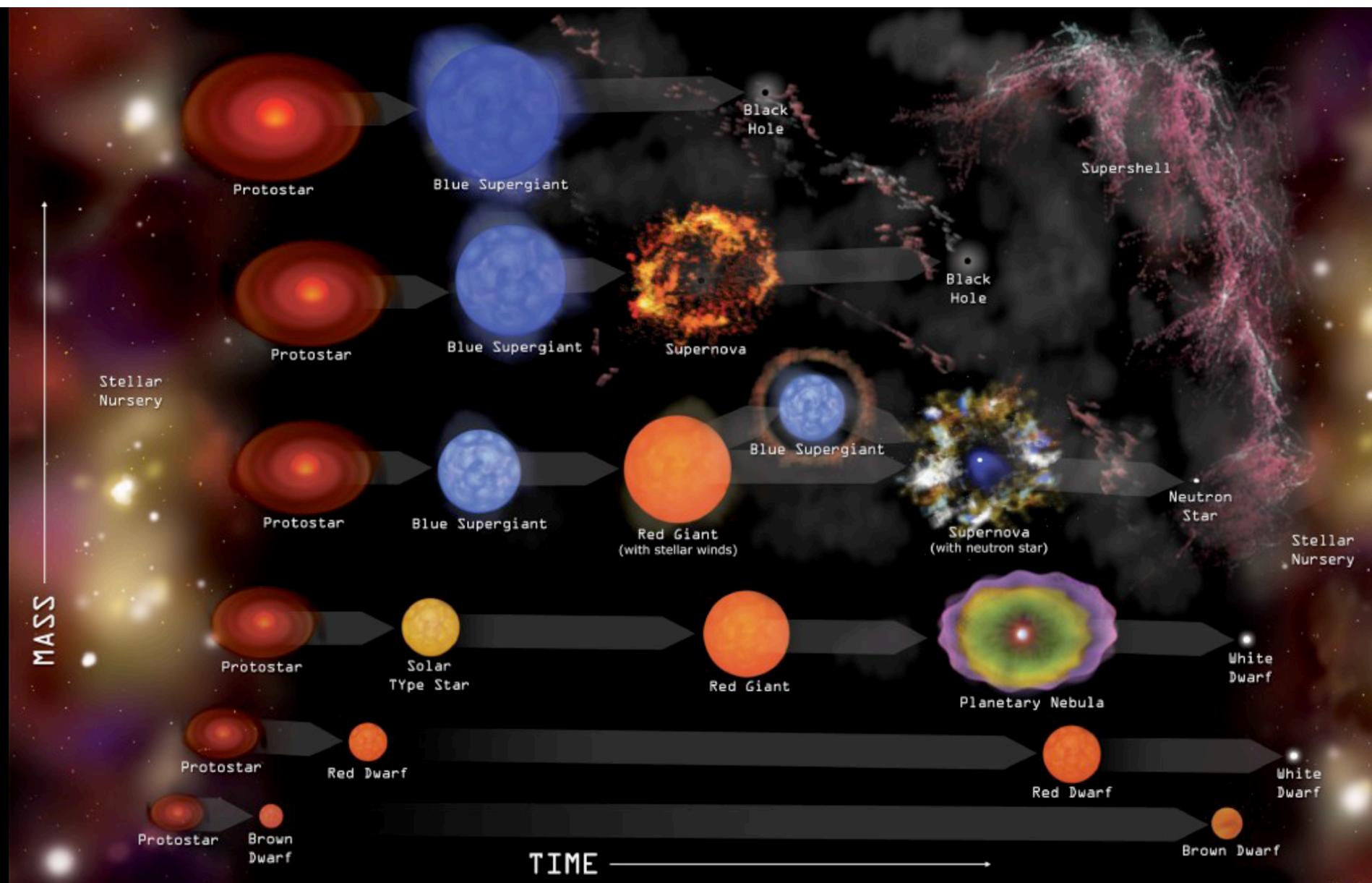


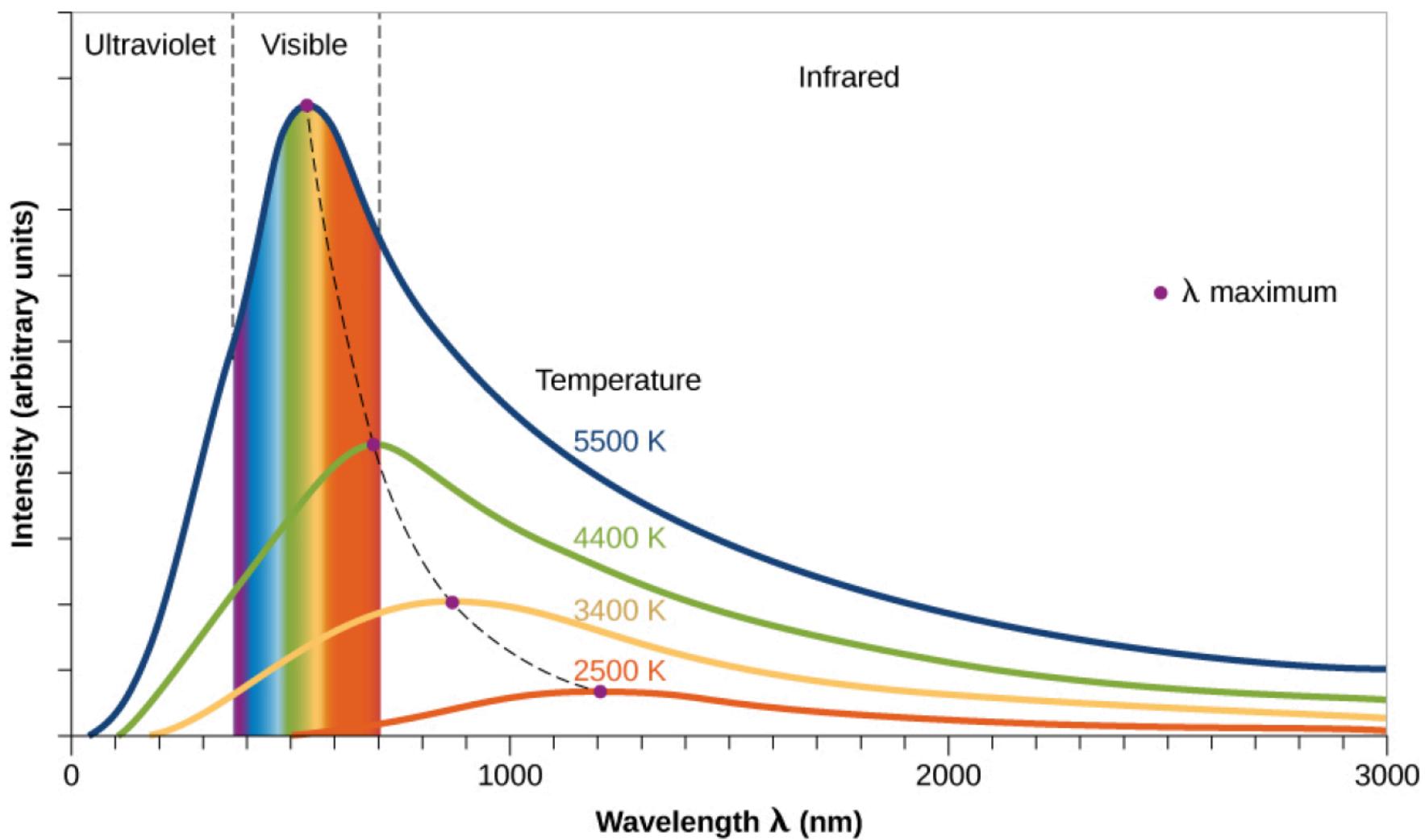


Stars: Star Formation and the Stellar Graveyard

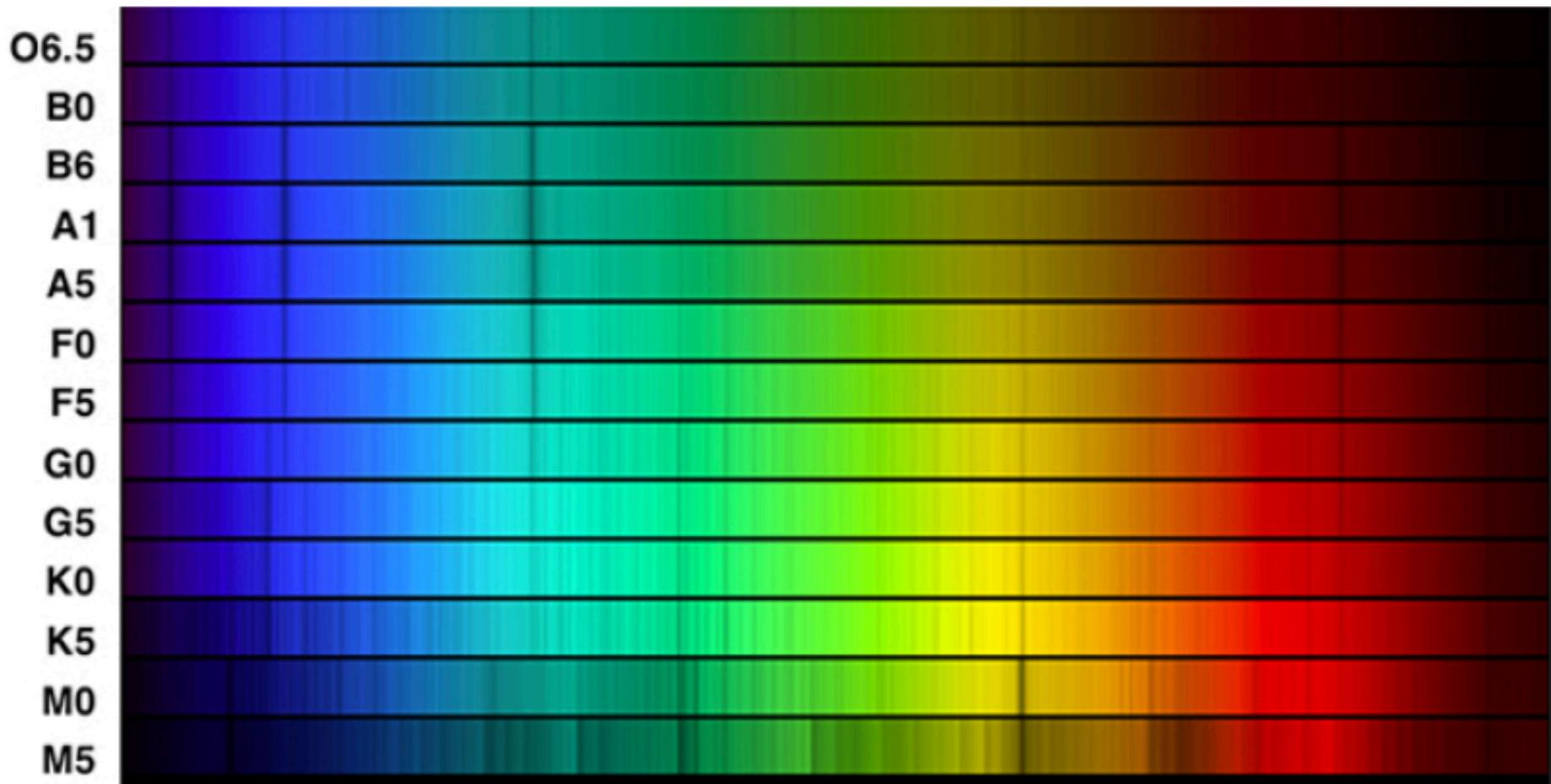


Blackbody emission: hotter things emit at higher energies (=shorter wavelengths)

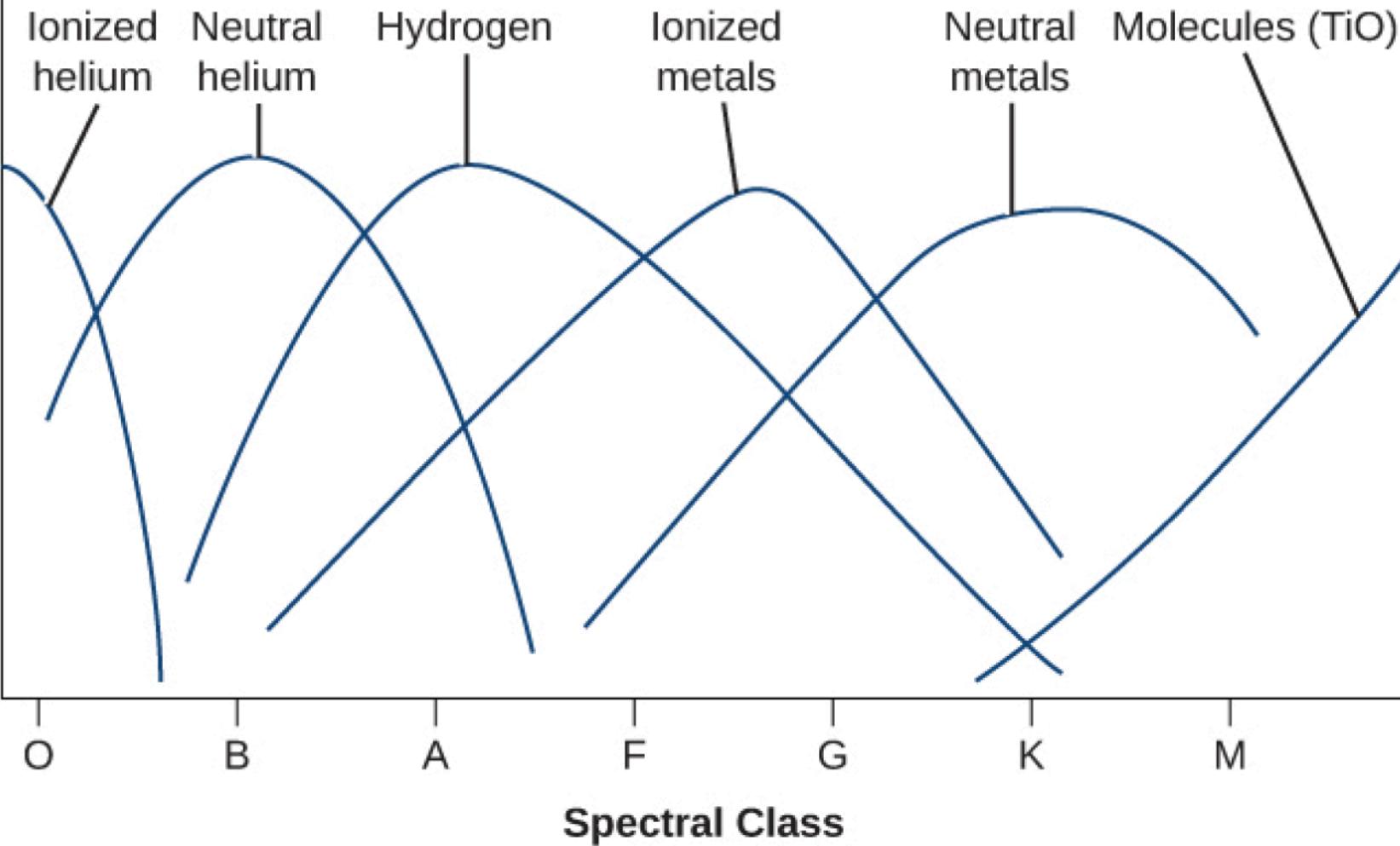
Peak of blackbody: $\lambda_{\max} \cdot T = 0.288 \text{ cm} \cdot \text{K}$



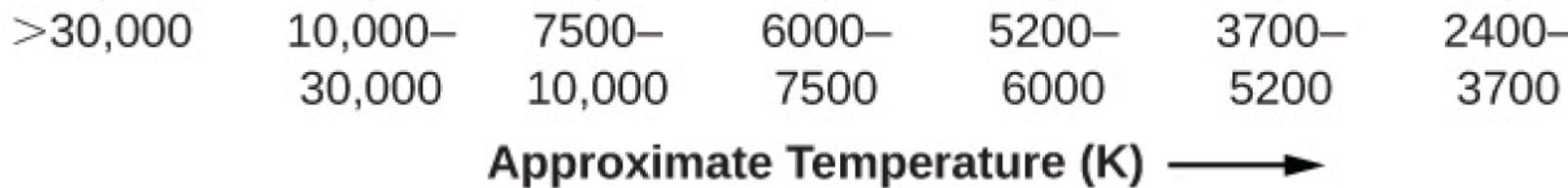
Spectral type = temperature sequence



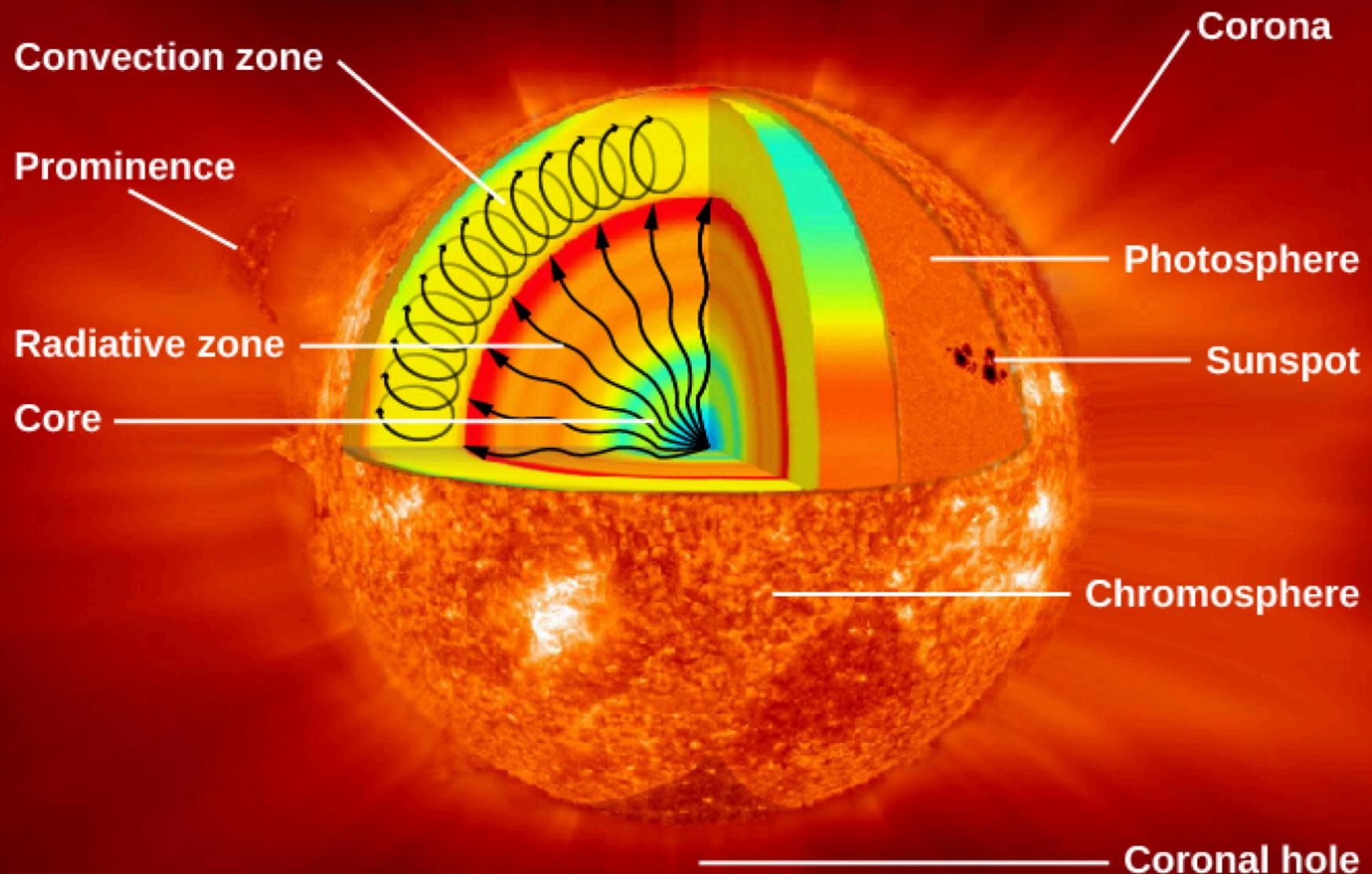
Relative Strengths of
Absorption Lines



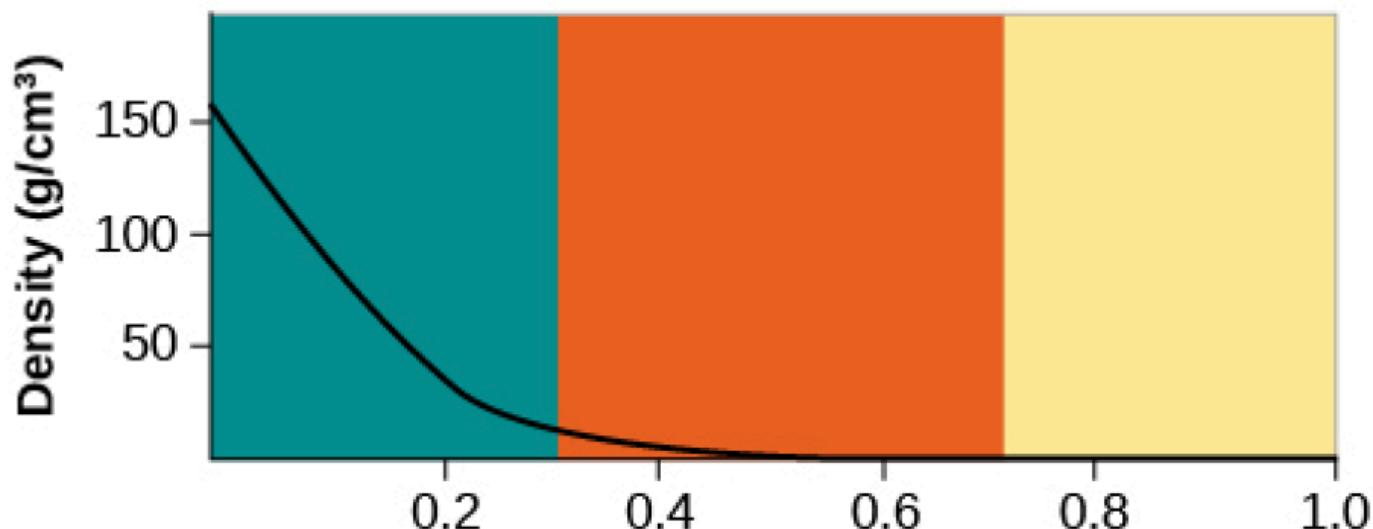
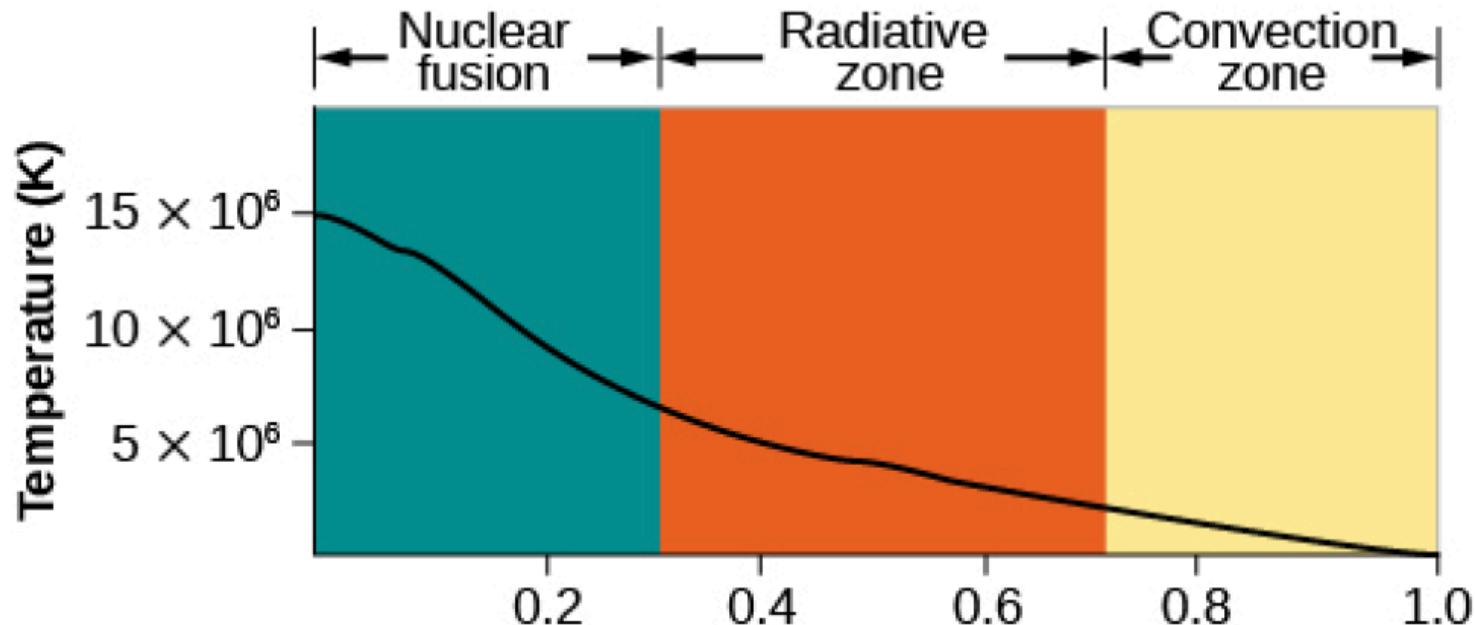
Spectral Class



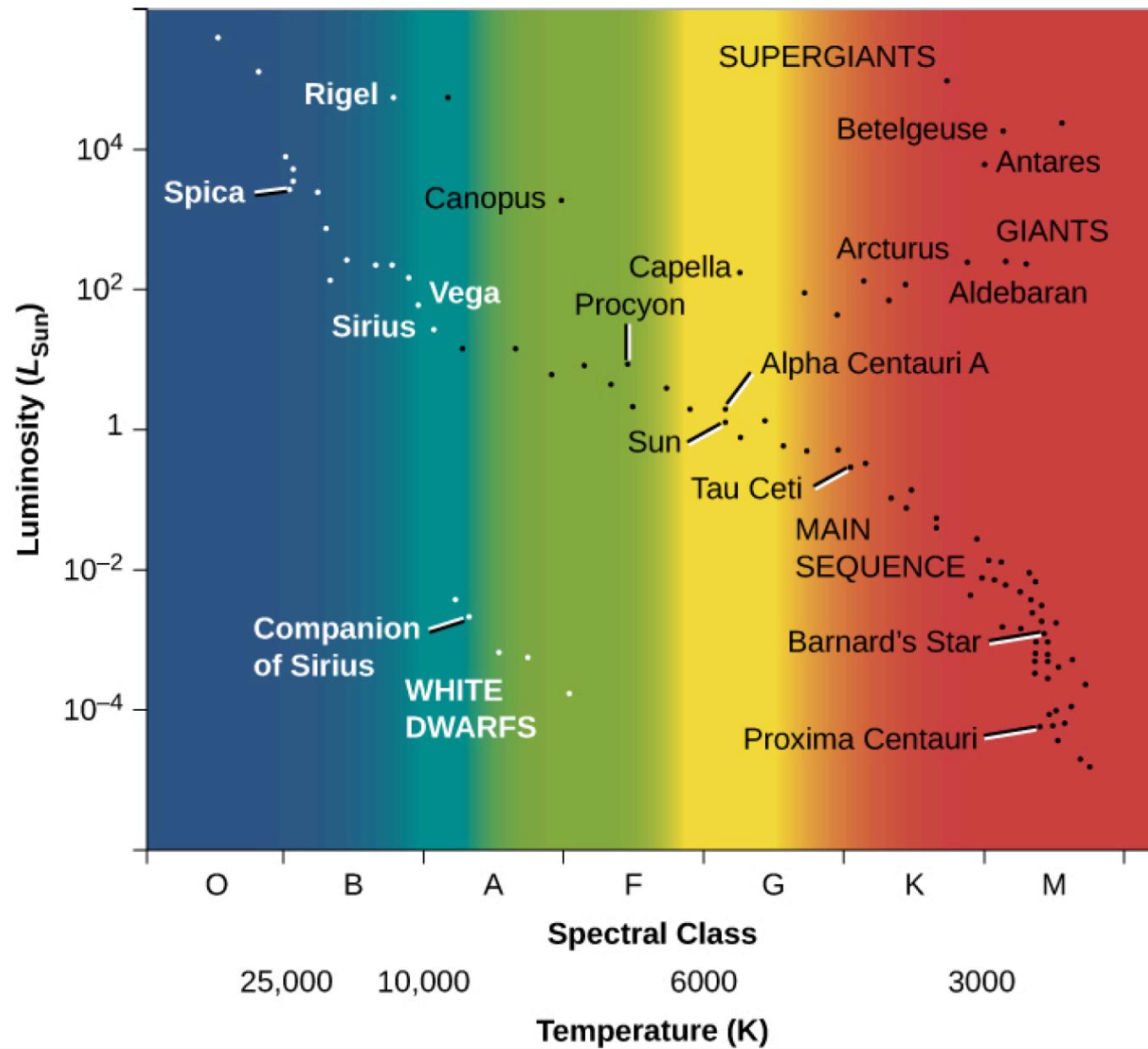
Approximate Temperature (K)



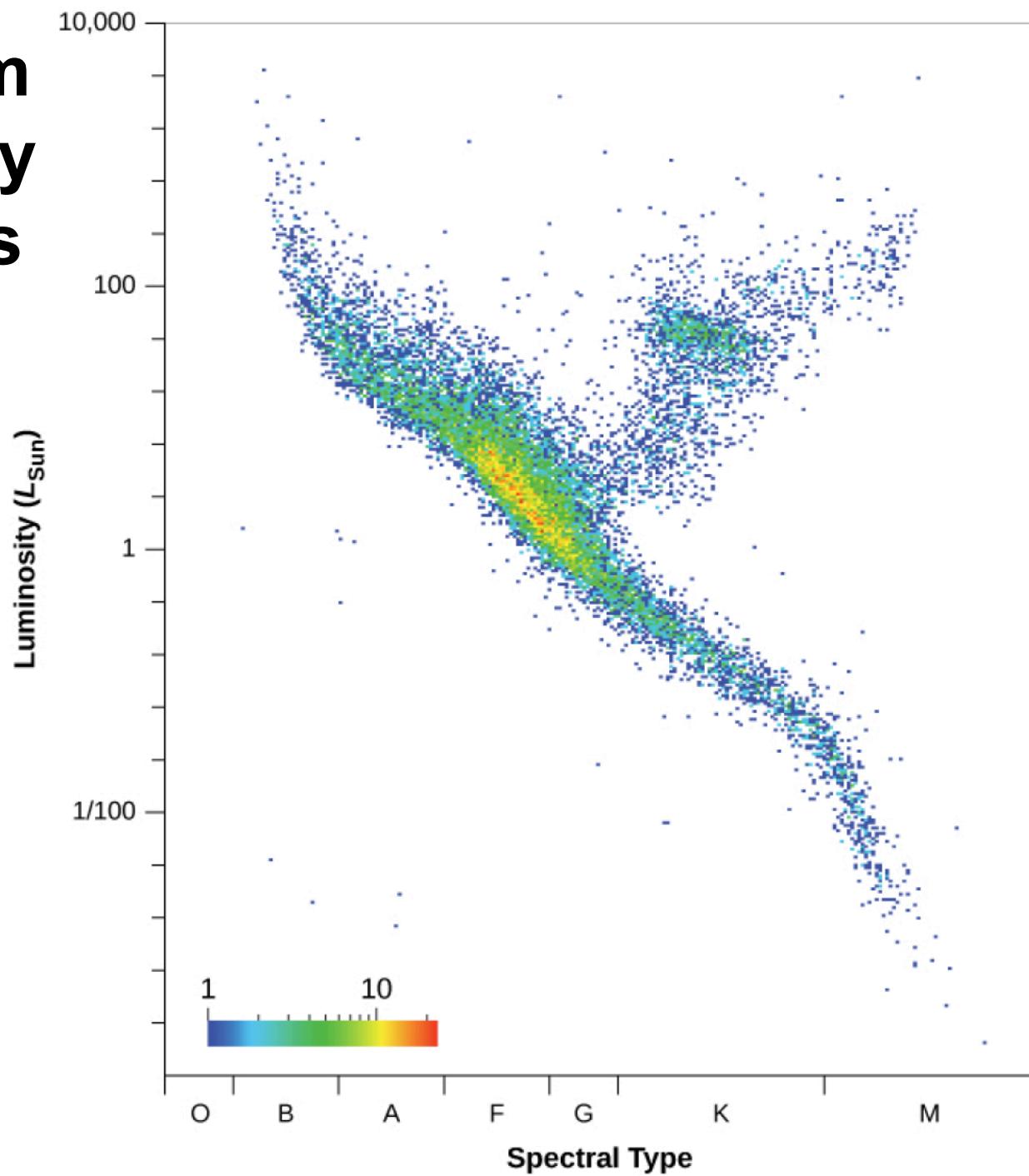
Core of star: very dense, 15 million K



HR Diagram



HR Diagram of all nearby bright stars



Interstellar medium

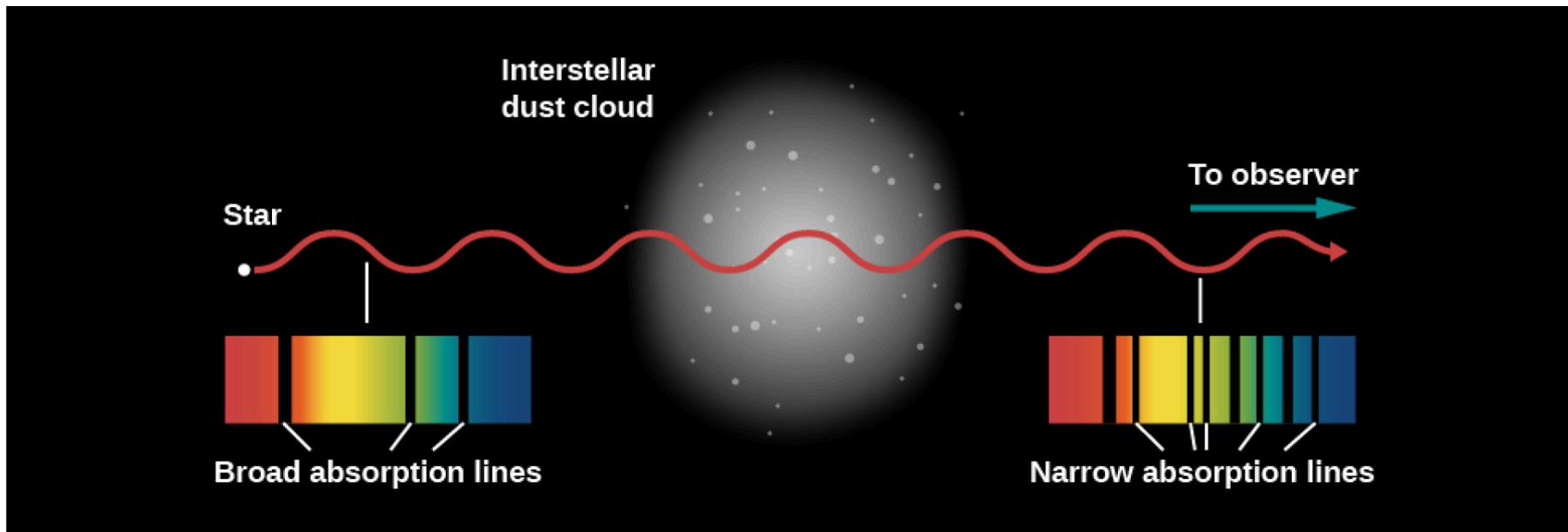
- Space is not quite empty
 - Hot interstellar medium: 10^{-4} ions per cm^3
 - In this room: 10^{19} molecules/ cm^3
 - Best vacuum in lab: 10^{10} molecules/ cm^3
- Some places are denser and colder
 - Molecular clouds, where stars form
 - Densities of 10^2 - 10^6 molecules/ cm^3



Intestellar medium: how to detect?

Absorption of photons by gas

Emission from gas/dust



Orion Nebula



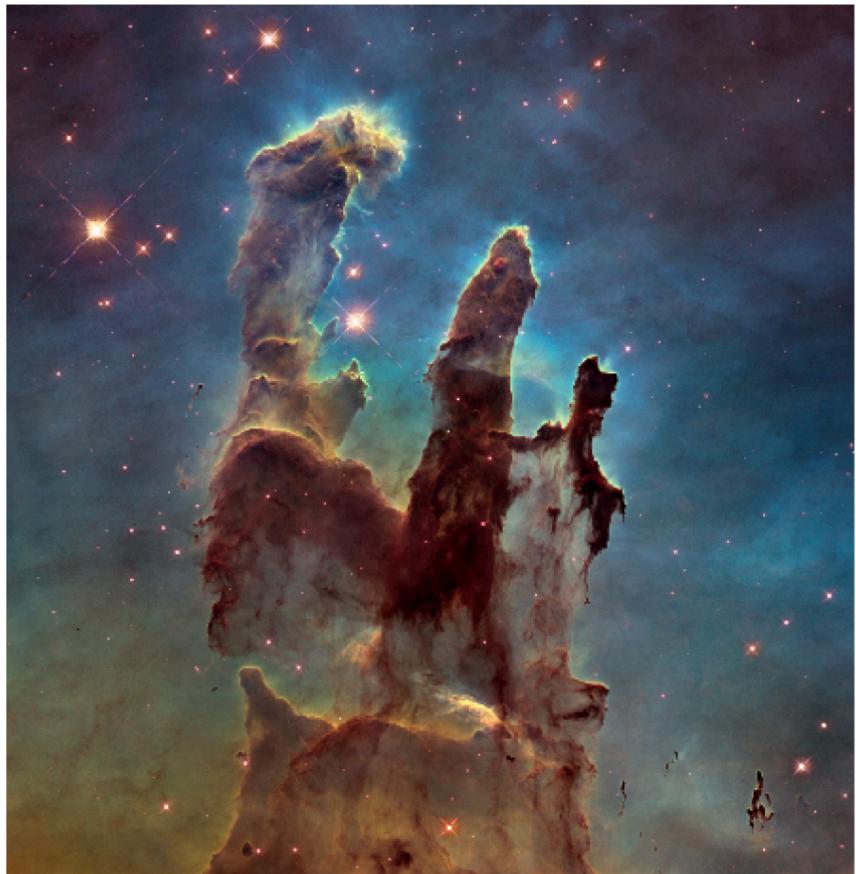
Eta Carina Cluster, Hubble Space Telescope Much larger than Orion Nebula



"Mystic Mountain" A Pillar of Gas and Dust in the Carina Nebula



HUBBLESITE.org





Hubble's "Pillars of Creation"
[shown to scale]



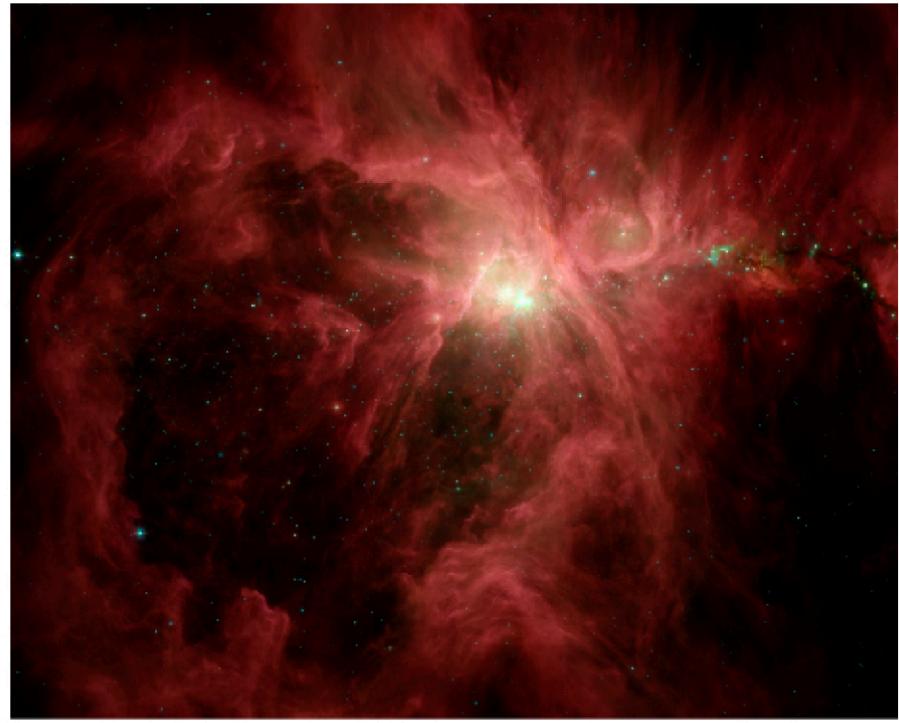
"Pillars" and "Mountains" of Star Formation

NASA / JPL-Caltech / L. Allen (Harvard-Smithsonian CfA)

Spitzer Space Telescope • IRAC

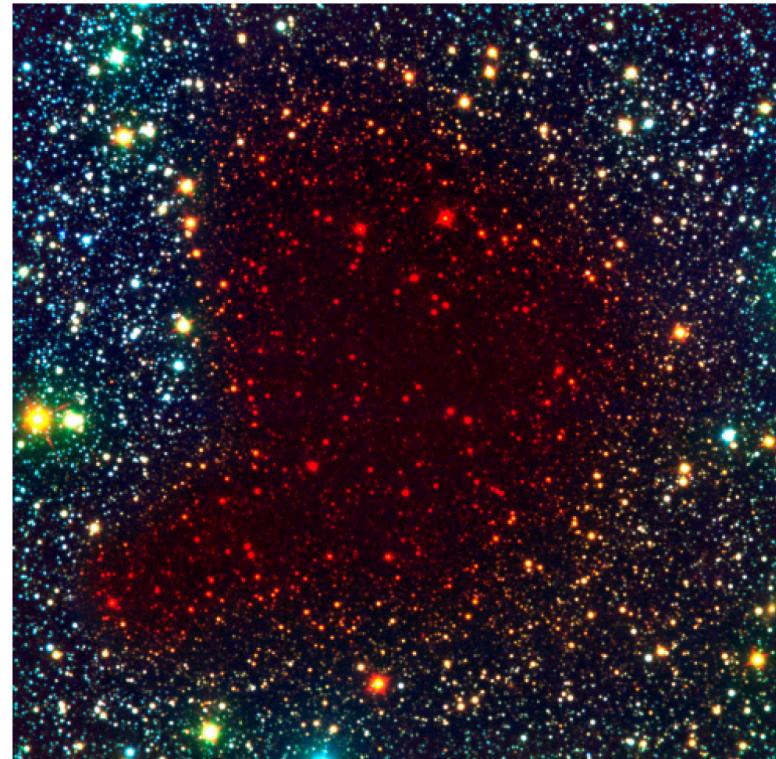
Inset: Hubble Space Telescope

ssc2005-23b



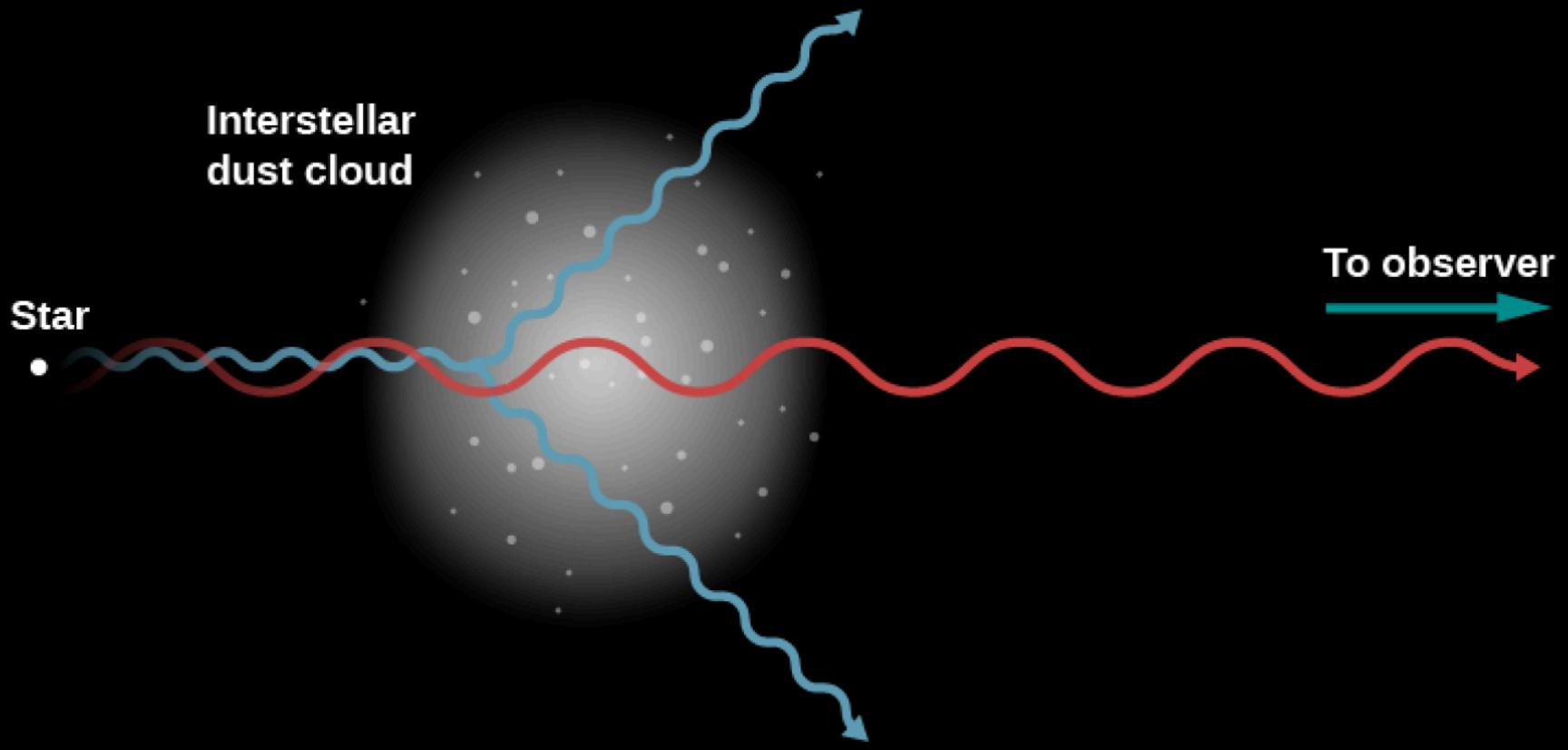


Optical



Near-infrared

Barnard 68: very dusty.
Longer wavelengths look through dust



Blue wavelengths: absorbed/scattered by dust
Red wavelengths: pass through dust



Rosette Nebula
Far-infrared: dust in emission