

# The Doodle Verse

## 1. Introduction

## 2. Overview

## 3. Warmup

### a. Mollweide Projection

One issue is that our extracted feature points will have Cartesian coordinates and our star data set is in spherical coordinates. Here is a readout of the first 10 stars of our data set:

RA	Dec	Mag
float64	float64	float64
0.02662528	-77.06529438	4.78
0.03039927	-3.02747891	5.13
0.03266433	-6.01397169	4.37
0.03886504	-29.72044805	5.04
0.06232565	-17.33597002	4.55
0.07503398	-10.50949443	4.99
0.08892938	-5.70783255	4.61
0.13976888	29.09082805	2.07
0.15280269	59.15021814	2.28
0.15583908	-27.9879039	5.42

The "RA" or measures the distance from a central meridian and ranges from  $0^\circ$  to  $360^\circ$ . The "Dec" or "declination" measures the distances above or below the equator and ranges from  $-90^\circ$  to  $90^\circ$ .

To convert from spherical coordinates to Cartesian coordinates, we need to pick a center point and project the stars in the neighborhood of this point onto a plane. If we attempt to project all or most of the stars, this will distort the stars far from our center point, which would be undesirable considering we are trying to match shapes. This restricts our search to small neighborhoods at a time.

The Mollweide projection of a star with spherical coordinates  $(\lambda, \phi)$  centered around  $(\lambda_c, \phi_c)$  can be obtained as follows:

$$x = R \frac{2\sqrt{2}}{\pi} (\lambda - \lambda_c) \cos(\theta)$$

$$y = R \sqrt{2} \sin(\theta)$$

Where  $\theta$  is the angle defined by:  $2\theta + \sin(2\theta) = \pi \sin(\phi - \phi_c)$ . Since  $\theta$  is implicitly defined, we cannot solve for it directly. But the following iteration will converge to its value quickly (it converges slow for points far from our center point, but that is not a concern for us).

$$\theta_0 = \phi - \phi_c$$

$$\theta_{k+1} = \theta_k + \frac{2\theta_k + \sin(2\theta_k) - \pi(\sin\phi - \phi_c)}{2 + 2\cos(2\theta_k)}$$

Right a script that does this?

## b. Procrustes

Here we address the problem of finding the ideal transformation  $T$  between our set of feature points  $F$  and a predetermined subset of stars  $S \subset \mathbb{S}$ . Here we are assuming  $F$  and  $S$  are the same size ( $k \times 2$ ) and represent their respective collections of points in Cartesian coordinates. We want to preserve the shape of  $F$ , and for now let's assume we also don't want to change the scale of  $F$ : our transformation will preserve distances and angles between points (and thus the shape). This simplifies the problem because it constrains  $T$  to be orthogonal. Precisely, we are looking for the orthogonal  $T$  that minimizes  $\|FT - S\|_F$ .

## 4. Lab