

Modern middleware for the data acquisition of the Cherenkov Telescope Array

Etienne Lyard^a, Roland Walter^a, Karl Kosack^b, Jean Jacquemier^c, Igor Oya^d, Peter Wegner^d, Matthias Fuessling^d and Xin Wu^{*a}

^aObservatory of Geneva, University of Geneva, Chemin d'Ecogia 19, CH-1290 Versoix,
Switzerland; ^bCEA Saclay, DSM/IRFU/SAp, Bat 709, F-91191 Gif-Sur-Yvette, France;
^cLaboratoire d'Annecy-le-Vieux de Physique des Particules, Université de Savoie, CNRS/IN2P3,
F-74941 Annecy-le-Vieux, France; ^dDESY Zeuthen, Platanenallee 6, D-15738 Zeuthen,
Germany

E-mail: etienne.lyard@unige.ch, roland.walter@unige.ch

The data acquisition system (DAQ) of the future Cherenkov Telescope Array (CTA) must be efficient, modular and robust to be able to cope with the very large data rate of up to 550 Gbps coming from many telescopes with different characteristics. The use of modern middleware, namely ZeroMQ and Protocol Buffers, can help to achieve these goals while keeping the development effort to a reasonable level. Protocol Buffers are used as an on-line data format, while ZeroMQ is employed to communicate between processes. The DAQ will be controlled and monitored by the Alma Common Software (ACS).

Protocol Buffers from Google are a way to define high-level data structures through an interface description language (IDL) and a meta-compiler. ZeroMQ is a middleware that augments the capabilities of TCP/IP sockets. It does not implement very high-level features like those found in CORBA for example, but makes use of sockets easier, more robust and almost as effective as raw TCP. The use of these two middlewares enabled us to rapidly develop a robust prototype of the DAQ including data persistence to compressed FITS files.

The 34th International Cosmic Ray Conference, 30 July- 6 August, 2015 The Hague, The Netherlands

*Speaker.

1. Introduction

The Cherenkov Telescope Array [1] will generate large amounts of raw data. It will consist of a heterogeneous system composed of telescopes of different sizes, designs and cameras. The total throughput from the telescopes is thought to be as high as 550 Gbps while this data is to be filtered so that less than 50 Gbps are stored on-site. Additionally, ways to further reduce the amount of raw data by another factor 10 are under investigation [2]. To handle such large data rates along with the multiple filtering processes it is desirable that the Array Control software [3] and more specifically the data acquisition system (DAQ) is efficient, modular and robust. An illustration of a possible pipeline is illustrated in Fig. 1. This architecture will be deployed on top of the Alma Common Software framework (ACS) [4] so that the DAQ can be controlled and monitored like any other component of the array. We propose a modular architecture that brings several improvements compared to a monolithic approach. Complex processing can be decomposed into simple tasks, thus easing the development and validation procedures. It allows for easier load-balancing across several physical nodes and helps to achieve the high availability requirement by allowing the reconfiguration of the system at any time. In addition to air-shower observations, it is proposed that CTA will also operate in optical interferometry mode [5]. This extra functionality would continuously record the observed value of a single pixel, sampled at 100 MHz or more. Therefore the DAQ is designed to have the necessary level of modularity, which allows for observing mode reconfiguration on a short timescale.

The DAQ will make use of three main external libraries. The Protocol Buffers library from Google [6], is used for serializing structured data. It maintains backward and forward compatibility and provides reflection. ZeroMQ [7] is a middleware that simplifies the usage of streams. Although designed for ALMA [8], it is proposed that the ACS framework will be used to deploy, control and monitor all software components of CTA.

2. Data Volume Reduction Scenarios

The general concept for camera data capture involves transmission from each camera to a camera server process running at a central location. In most cases a stream of all available raw data (i.e. locally triggered events) can be continuously delivered, without any event pre-selection (or array-level triggering). This approach minimises the complexity on the camera side and delivers maximum flexibility for data collection. Subsystem-level hardware triggers are envisaged internally by the camera groups of the large-sized and Schwarzschild-Couder telescopes. Under these assumptions, the volume of data delivered by cameras must be reduced by a large factor (\approx 100) so that it can be written to permanent storage. Reducing this large data set should not discard scientifically useful data, thus each new data trimming process should be carefully evaluated and tested off-line [9] before it is introduced in the real-time analysis pipeline [2]. The proposed modular architecture achieves this by storing the full dataset during commissioning and gradually introducing new data trimming algorithms in the real-time pipeline once they are well understood.

Various data reduction strategies are under consideration, including obvious choices such as waveform reduction and zero suppression, but also more ambitious ones such as background suppression. The latter aims at discarding proton and helium nuclei initiated cascades, from which

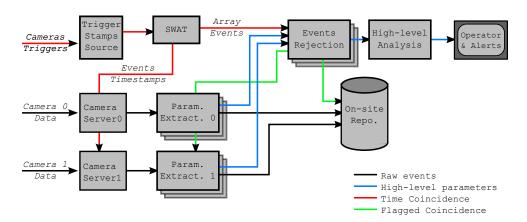


Figure 1: Data flow diagram of the proposed real-time data reduction configuration. Single-camera triggers are time-stamped before being passed to the software array trigger (SWAT). Timestamps passing the SWAT criteria are sent to the camera servers, which in turn send the associated event data to the parameter extraction nodes. There the data are calibrated, a first analysis performed and image parameters extracted. From there, event parameters are delivered to the event rejection module (blue lines) where they are assembled into array-events and further filtered. The result of this analysis is fed back to the parameter extraction nodes (green lines) and forward to the high-level analysis. Events that remain of interest are then compressed and written to the repository (black lines), while the result of the high-level analysis is presented to the operator. The parameter extraction nodes must provide enough buffering capabilities to wait until the analysis is complete before deciding whether to discard an event or not. All boxes and connecting lines in this diagram are logical entities. More than one box can run on a single physical server, and a given box can be loaddistributed across several servers. The lines also represent logical connections, which will be implemented via the ZeroMQ middleware for the data transfer and ACS remote calls for the configuration. In order to improve fault tolerance, information from the event rejection could be routed in parallel to an algorithm that detects and corrects faults in the information used to associate events. The results of that algorithm would then be provided to the off-line analysis.

most of the background Cherenkov light is produced. Tagging such events in a reliable way is a computationally intensive process, which will be challenging to implement in real-time. Further details on these approaches, as well as the implication on the data pipelines and calibration are provided in [10].

3. Data Streams

All control and monitoring will be implemented using ACS, while the bulk data transfer will be implemented using <code>ZeroMQ</code>, as seen in figure 2. All connection procedures and the messages concept will be implemented with <code>ZeroMQ</code>. Because <code>ZeroMQ</code> is also capable of connecting different processes or threads running on a single machine, it is very easy to build single large components from several smaller ones if the communication performance requirements are not met by network protocols. Connection paradigms will be chosen individually for each connection item. The possible cases are request-reply, push-pull and publish-subscribe. <code>ZeroMQ</code> implement all connection procedures along with the concept of messages. This way the user does not need to worry about buffers or protocols. <code>ZeroMQ</code> delivers incoming messages as they arrive, depending on the selected distribution paradigm.

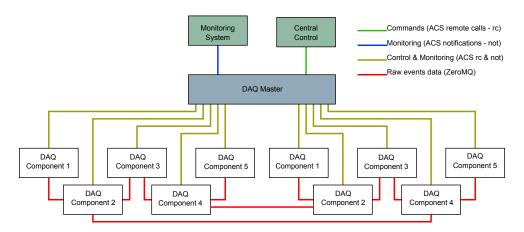


Figure 2: Conceptual view of the DAQ data flow with an arbitrary number of low-level components interconnection. Commands are provided by the central control, directly to the DAQ master. The master then takes appropriate action to reconfigure the relevant components accordingly, for instance by connecting two low-level components with a raw events stream. Each low-level component is monitored by the DAQ master which in turn forwards summary information to the monitoring system. If necessary, such as for debugging purposes, it is possible that the monitoring system directly connects to the low-level components notification channels. The vertical data flow is entirely handled by ACS (either through remote procedure calls or notification channels) while the horizontal, high throughput data flow is entirely handled by <code>ZeroMQ</code> streams. This architecture reduces the complexity for the central control, and lets the DAQ master handle errors.

4. Structured Data Format

The internal DAQ format will rely on Protocol Buffers developed and widely used by Google [6]. Protocol Buffers provide a standardised way to handle the serialization problem. Structures are defined via a high-level interface description language. A compiler then generates the run-time code. The serialized version of the structure is highly optimized, with values being packed as efficiently as possible to reduce the network throughput or left untouched to minimize the serialization overload. Moreover, the built-in forward and backward compatibility of the format allows modification of the structure content without breaking the pipeline. As an option that can be selected, the built-in reflection of the structures allows reading and writing of any data structure to/from persistent storage without having to write a single line of code specific to that data structure.

Last but not least, Protocol Buffers are available in many programming languages. This will allow us to mix different languages to build the real-time analysis pipeline and thus to reuse offline algorithms that may have been written in languages other than C/C++.

One weakness of Protocol Buffers is that they do not include a 16-bit integer type. We overcame this problem by implementing our own array class that operates on top of the standard arrays in protocol buffers. In addition, the serialisation is sometimes slower than expected. To remedy this, we had to use the fixed32 type instead of the int32 type. This prevented the occurrence of time-consuming on-the-fly compression.

5. Prototyping

As a proof of concept, a prototype involving the components described below was developed: **Data Buffers** store events and/or event parameters in memory for some time until a decision is taken regarding their fate. They provide flexibility in the pipeline by allowing the array event builders to retrieve event parameters only when the time coincidence is known.

Repository Writers are responsible for committing incoming data to the on-site repository. If the data is provided raw then it is compressed on-the-fly. Each repository writer deals with one single stream and opens new files once the size or event number limit is reached or after a new run is started. The target storage node will remain the same at least during a single run. Changing the storage node from run to run would allow for the next-day analysis to start while data taking is still ongoing, possibly providing more accurate results earlier.

Processing Nodes are meant to be all-purpose skeleton components so that any kind of computation can easily be introduced to the pipeline. Their primary goal is to embed real-time processing into the DAQ pipeline. For instance, extracting the image parameters from the raw event data is desirable early in the pipeline to help reduce the total network throughput. The data reduction algorithms will be provided by another workpackage (DATA). We developed a simple image cleaner and muon extractor based on the Hough transform [11] to use within this demonstrator.

We also developed a monitoring tool to be used in place of ACS so that the required third party software is reduced to a minimum. A display of the running pipeline can be seen in figure 3. Due to time constraints, this prototype only focuses on load-balancing and automatic failure recovery. The maximum throughput was evaluated between two nodes only and reached ≈ 9 Gbps. This suggests that the chosen middleware makes the pipeline perform $\approx 10\%$ slower compared to when using raw C-structs and tcp sockets in low-latency environments. This should not be an issue as more than one stream can easily be created to accommodate higher throughput.

Integration tests to verify if this approach is suitable to readout the data from the camera servers is yet to be implemented. In addition, these tests will highlight if any optimizations and/or specific constraints are required. The small-sized telescopes should not present a problem, as the data rates remain low. For the larger-sized telescopes the high throughput from the cameras will most likely prevent the DAQ from performing on-the-fly data conversion. Thus delaying the unification of the data format to a later stage in the acquisition pipeline.

The prototype pipeline allows for nodes to be discarded or new ones to be introduced at any time. For example, if a server fails and all data sitting in its memory is lost but all other data manages to buffer (until a replacement becomes available).

Using the Protocol Buffers allowed us to make our data model evolve in a flexible way. This evolution was triggered by the introduction of new concepts that required new data structures or the modification of existing ones.

6. Conclusion

The use of the proposed middlewares allowed us to quickly develop a proof-of-concept toy model. This would not have been possible using traditional sockets and structures as many low-level features would have had to be reimplemented. Using these middlewares to achieve greater

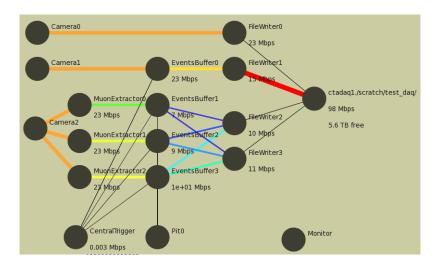


Figure 3: Screenshot of the monitor display from the DAQ demonstrator. Each circle represents a DAQ component while the lines connecting them are actual <code>ZeroMQ</code> connections. The colour and width of the lines reflect the average throughput for each connection. Due to the use of <code>ZeroMQ</code> and <code>ProtocolBuffers</code>, new components can be added or removed without stopping operations. This particular view shows the data taking process from 3 cameras. The first one has all its events written to disk, the second camera pipeline discards events based on their stereo coincidence while the third camera pipeline has an image cleaner and muon extractor which analyses the data in real-time. The EventsBuffer nodes retain the events until their stereo coincidence is known, while the FileWriter nodes compress and write the events to disk in compressed FITS format. The node called <code>PitO</code> is a proxy that only receives events classified as muons for monitoring purposes

flexibility comes at a computing power and throughput cost of $\approx 10\%$. We strongly believe that this approach is the best one to follow because high-level implementations can be replaced by lower-level code that is optimised toward the chosen features and architecture of the system, and which is unlikely to change over time.

References

- [1] B.S. Acharya et al., Introducing the CTA concept, Astroparticle Physics vol. 43 No. 3, 2013
- [2] A. Bulgarelli, V. Fioretti, *The On-Site Analysis of the Cherenkov Telescope Array*, In these proceedings, 2015
- [3] I. Oya et al., Status and Plans for the Array Control and Data Acquisition system of the Cherenkov Telescope Array In these proceedings, 2015
- [4] G. Chiozzi et al., *The ALMA common software: a developer-friendly CORBA-based framework.* in Advanced Software, Control, and Communication Systems for Astronomy, Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series 5496, 205-218, 2004
- [5] D. Dravins et al., *Optical intensity interferometry with the Cherenkov Telescope Array. Astroparticle Physics* vol. 43 No. 0, 2013
- [6] Google Inc., Protocol Buffers, http://developers.google.com/protocol-buffers, 2014
- [7] Hintjens, P., ZeroMQ: Code Connected, http://zeromq.org, 2014

- [8] A. Wootten, *The Atacama large millimeter/submillimeter array.*, in proceedings of the IEEE 97.8: 1463-1471, 2009
- [9] K. Kosack et al., Cherenkov Telescope Array Data Processing Pipeline, In these proceedings, 2015
- [10] J. Hinton et al., *Data Volume Reduction in CTA*, in the CTA documentation, ref. MAN-PO/130828, 2014
- [11] P. E. Hart, *How the Hough Transform Was Invented*, in IEEE Signal Processing Magazine, 26(6):18 22, 2009