

## Function GET\_MW – single-thread version

Calling syntax:

*For IDL users:*

```
res = call_external(libname, 'GET_MW', Lparms, Rparms, Parms, $  
                    E_arr, mu_arr, f_arr, RL)
```

*For Python users:* firstly, you need to load the libraries:

```
import GScores
```

```
GET_MW = GScores.initGET_MW(libname)
```

Then, when necessary, the function is called as:

```
res = GET_MW(Lparms, Rparms, Parms, E_arr, mu_arr, f_arr, RL)
```

In both cases, libname is the name of the appropriate executable library (\*.dll or \*.so).

*Notes for Python users:*

- All function parameters should be of numpy.ndarray type. For multi-dimensional arrays, the Fortran-like row-column ordering should be used, e.g.:  
RL = np.zeros((7, Nf), dtype = 'double', order = 'F').
- In the below descriptions, the IDL conventions are used; e.g., “\*” symbols should be replaced everywhere by “:”, etc.

Function parameters:

0. Lparms – 11-element long (32-bit) integer array of dimensions and global (for all voxels) integer parameters (see below).
1. Rparms – 5-element double array of global (for all voxels) real parameters (see below).
2. Parms – array of LOS parameters,  $24 \times N_z$  elements, double. Parms[:, i] represents the parameters for  $i$ th voxel (see below).
3. E\_arr – array of energies  $E_i$  where the electron distribution function is specified, NE elements, double, in MeV. The values must be monotonically increasing.
4. mu\_arr – array of pitch-angle cosines  $\mu_j = \cos \alpha_j$  where the electron distribution function is specified, Nmu elements, double. The values must be monotonically increasing and should cover the entire range of possible values from -1 to +1.
5. f\_arr – array of electron distribution functions  $f_{ijk} = f_{ij}^{(k)}(E_i, \mu_j)$ ,  $NE \times Nmu \times N_z$  elements, double, in  $\text{cm}^{-3} \text{MeV}^{-1}$ . f\_arr[:, \*, k] represents the distribution function for  $k$ th voxel. The distribution function in each voxel is assumed to satisfy the normalization condition

$$2\pi \int_{E_{\min}}^{E_{\max}} dE \int_{\mu_{\min}}^{\mu_{\max}} f(E, \mu) d\mu = n_b,$$

where  $n_b$  is the local concentration of energetic electrons (in  $\text{cm}^{-3}$ ), and the energy  $E$  is in MeV.

6. RL – input/output array,  $7 \times N_f$  elements, double. RL[\* , i] corresponds to  $i$ th frequency (see below).

Array of dimensions and global integer parameters Lparms:

Lparms = [Nz, Nf, NE, Nmu, Nnodes, \$  
match\_key, Qopt\_key, arr\_key, log\_key, PK\_key, spline\_key]

0. Lparms[0] = Nz – number of voxels along the LOS.
1. Lparms[1] = Nf – number of frequencies in the spectrum.
2. Lparms[2] = NE – number of energies in the E\_arr array; must be  $\geq 3$  – otherwise the array-defined electron distribution function is ignored.
3. Lparms[3] = Nmu – number of pitch-angle nodes in the mu\_arr array; must be  $\geq 3$  – otherwise the array-defined electron distribution function is ignored.
4. Lparms[4] = Nnodes – number of energy nodes used for integration over energy in the continuous gyrosynchrotron code.
  - a. Minimum value: 16; if  $0 \leq Nnodes < 16$ , 16 nodes are used instead.
  - b. If  $Nnodes < 0$ , an adaptive integration grid with the target relative accuracy of  $10^{-5}$  is used.
5. Lparms[5] = match\_key – controls the behaviour of the hybrid gyrosynchrotron code at the boundary frequencies  $f^c$  and  $f^{WH}$  (see Fleishman & Kuznetsov 2010):
  - a. 0: additional re-normalization of the spectrum is performed to remove possible jumps at the boundary frequencies;
  - b.  $\neq 0$ : re-normalization is not performed.
6. Lparms[6] = Qopt\_key – controls the  $Q$ -optimization of the continuous gyrosynchrotron code (see Fleishman & Kuznetsov 2010):
  - a. 0:  $Q$ -optimization is on, which improves accuracy;
  - b.  $\neq 0$ :  $Q$ -optimization is off, which improves speed.
7. Lparms[7] = arr\_key – global key specifying which electron distribution functions (analytical or/and array-defined) are used to compute the gyrosynchrotron emission:
  - a. 0 (*default*): contributions of both the analytical and array-defined electron distribution functions are included (*this choice can be overridden in some voxels, depending on the local keys, see below; the array-defined distribution requires also  $NE \geq 3$  and  $Nmu \geq 3$* );
  - b. 1: the array-defined electron distribution function is disabled for all voxels, regardless on the local on/off keys;
  - c. 2: the analytical electron distribution function is disabled for all voxels (*equivalent to using the “free-free only” analytical model*), regardless on the local on/off keys.

*These flags can be combined: Lparms[7] = 3 disables both the analytical and array-defined distributions.*

8. Lparms[8] = log\_key – controls the assumptions about the energy grid for the array-defined electron distribution:
  - a. 0: the nodes are assumed to be logarithmically-spaced ( $E_{i+1}/E_i = \text{const}$ );
  - b.  $\neq 0$ : the nodes are assumed to be equidistant ( $E_{i+1} - E_i = \text{const}$ ).

*Note: if neither of above is applied to your energy grid, choose the option that fits the actual energy spacing better – this can improve the calculation accuracy greatly.*
9. Lparms[9] = PK\_key – specifies how the pitch-angle dependences of the electron distribution functions are treated:
  - a. 0: the exact (possibly anisotropic) electron distributions are used (*default option*);
  - b. 1: for the analytical distribution function, the pitch-angle distribution (specified by Parm[14]) is switched to an isotropic one; for the array-defined distribution function, the electron distribution at each energy is replaced by an isotropic (pitch-angle-averaged) one;
  - c. 2: same as 1, and the continuous gyrosynchrotron code uses the fast approximation by Petrosian (1981) and Klein (1987).
10. Lparms[10] = spline\_key – controls the 2D interpolation method for the array-defined electron distribution function:
  - a. 0: spline interpolation is used (*usually provides higher speed and accuracy*);
  - b.  $\neq 0$ : local linear-quadratic interpolation over 2-3 adjacent nodes is used (*sometimes works better for the distributions with very sharp gradients*).

Array of global real parameters Rparms:

$$\text{Rparms} = [S, f_0, \Delta f, f^c, f^{WH}]$$

0. Rparms[0] =  $S$  – visible source area, in  $\text{cm}^2$ .
1. Rparms[1] =  $f_0$  – starting frequency of the spectrum, in Hz:
  - a. is used, only if  $f_0 > 0$ ;
  - b. if  $f_0 \leq 0$ , the frequencies are taken from the RL[0, \*] array.
2. Rparms[2] =  $\Delta f$  – logarithmic frequency step used to produce the spectrum,  $f_{i+1}/f_i = 10^{\Delta f}$  (is used only if  $f_0 > 0$ ).
3. Rparms[3] =  $f^c$  – boundary frequency of the hybrid gyrosynchrotron code (Fleishman & Kuznetsov 2010), expressed in units of the local electron gyrofrequency.
  - If the emission frequency  $f < f^c$ , the exact code with summation over cyclotron harmonics is used.
  - If  $f > f^c$ , the continuous code is used.
  - If  $f^c < 0$ , the code is purely continuous with additional re-normalization using the exact parameters computed at  $f = f^{WH}$ .

4.  $\text{Rparms}[4] = f^{\text{WH}}$  – boundary frequency for the exact/approximated expressions for the Bessel functions in the exact gyrosynchrotron code, expressed in units of the local electron gyrofrequency.

- If  $f < f^{\text{WH}}$ , the exact gyrosynchrotron code (at  $f < f^{\text{C}}$ ) uses the exact expressions for the Bessel functions.
- If  $f > f^{\text{WH}}$ , the exact gyrosynchrotron code (at  $f < f^{\text{C}}$ ) uses the approximate expressions for the Bessel functions by Wild & Hill (1971).

Array of parameters  $\text{Parms}$  (for a single voxel, 24 parameters):

0.  $\text{Parms}[0] = \Delta z$  – voxel length, in cm.
1.  $\text{Parms}[1] = T_0$  – plasma temperature, in K.
2.  $\text{Parms}[2] = n_0$  – either thermal electron concentration or total atomic concentration (depending on other parameters, see the separate diagram [Diagram.pdf](#)), in  $\text{cm}^{-3}$ .
3.  $\text{Parms}[3] = B$  – magnetic field strength, in G.
4.  $\text{Parms}[4] = \theta$  – viewing angle, in degrees.
5.  $\text{Parms}[5]$  – emission mechanism flag (rounded down to the nearest integer):

- a. 0: all emission mechanisms (gyrosynchrotron + e-ions + e-neutrals) are included;
- b. 1: gyrosynchrotron is off;
- c. 2: e-ions is off;
- d. 4: e-neutrals is off.

*Several flags can be combined by usual or bitwise summation: e.g.,  $\text{Parms}[5] = 2 + 4$  turns off both e-ions and e-neutrals, etc.*

6.  $\text{Parms}[6]$  – specifies the chosen analytical electron distribution over energy (index of the model distribution function, see the separate document [AnalyticalDistributions.pdf](#)); non-integer values are rounded down to the nearest integer. *Default option: 0.*

*Note: if the kappa-distribution ( $\text{Parms}[6] = 6$ ) is selected, the e-ions contribution is also computed using the formulae for the kappa-distribution (Fleishman & Kuznetsov 2014); in all other cases, the Maxwellian thermal distribution is assumed.*

7.  $\text{Parms}[7] = n_b$  – concentration of nonthermal electrons in the analytical electron distributions, in  $\text{cm}^{-3}$ .

8.  $\text{Parms}[8] = \varepsilon$  or  $\kappa$  – either the matching parameter  $\varepsilon$  in the thermal/nonthermal electron distributions or the parameter  $\kappa$  in the kappa-distribution.

9.  $\text{Parms}[9] = E_{\text{min}}$  – the low-energy cutoff in the analytical electron distributions (when relevant), in MeV.

10.  $\text{Parms}[10] = E_{\text{max}}$  – the high-energy cutoff in the analytical electron distributions (when relevant), in MeV.

11. Parms[11] =  $E_{\text{break}}$  – the break energy in the double-power-law analytical electron distributions, in MeV.

12. Parms[12] =  $\delta_1$  – the power-law index in the single-power-law analytical electron distributions or the low-energy power-law index in the double-power-law analytical electron distributions.

13. Parms[13] =  $\delta_2$  – the high-energy power-law index in the double-power-law analytical electron distributions.

14. Parms[14] – specifies the chosen analytical electron distribution over pitch-angle (index of the model distribution function, see the separate document AnalyticalDistributions.pdf); non-integer values are rounded down to the nearest integer. *Default option: 0.*

15. Parms[15] =  $\alpha_c$  or  $\alpha_0$  – either the loss-cone boundary  $\alpha_c$  in the loss-cone analytical electron distributions or the beam direction  $\alpha_0$  in the beam-like analytical electron distributions, in degrees.

16. Parms[16] =  $\Delta\mu$  – either the loss-cone boundary width or the beam angular width in the loss-cone or beam-like analytical electron distributions, respectively.

17. Parms[17] =  $a_4$  – the coefficient  $a_4$  in the supergaussian beam-like analytical electron distribution.

18. Parms[18] =  $n_p$  – proton concentration, in  $\text{cm}^{-3}$ ; is used only as a switch (see the separate diagram Diagram.pdf).

19. Parms[19] =  $n_{\text{HI}}$  – neutral hydrogen concentration, in  $\text{cm}^{-3}$  (see the separate diagram Diagram.pdf).

20. Parms[20] =  $n_{\text{HeI}}$  – neutral helium concentration, in  $\text{cm}^{-3}$  (see the separate diagram Diagram.pdf).

21. Parms[21] = arr\_key\_local – local key (rounded down to the nearest integer) specifying which electron distribution functions (analytical or/and array-defined) are used to compute the gyrosynchrotron emission in this voxel:

a. 0 (*default*): contributions of both the analytical and array-defined electron distribution functions are included (*provided that they are enabled by the global key; the array-defined distribution requires also  $NE \geq 3$  and  $Nmu \geq 3$* );

b. 1: the array-defined electron distribution function in this voxel is ignored even if it is specified.

c. 2: the analytical electron distribution function in this voxel is ignored (*equivalent to using the “free-free only” analytical model*).

*These flags can be combined: Parms[21] = 3 disables both the analytical and array-defined distributions.*

22. Parms[22] – element abundance model (used to compute the e-ions contribution):

a. -1: “classical” formulae from Dulk (1985) are used;

b. 0: solar coronal abundance (by Feldman 1992) is used (*default option*);

- c. 1: solar photospheric abundance (by Scott et al. 2015) is used.
23. Parms[23] – currently unused.

Input/output array RL:

0. First row (RL[0, \*]) – emission frequencies, in GHz. On input, this array is used if  $f_0 = \text{Parms}[1] \leq 0$  (the specified frequency values must be monotonically increasing). Otherwise, the frequencies are computed using the  $f_0$  and  $\Delta f$  parameters:  $f_1 = f_0 10^{\Delta f}$ ,  $f_2 = f_1 10^{\Delta f}$ , etc. On output, this array contains the computed or pre-defined emission frequencies.

Other rows – emission intensities, as observed from the Earth, in sfu:

1. RL[1, \*] – left polarization, weak mode coupling;
2. RL[2, \*] – right polarization, weak mode coupling;
3. RL[3, \*] – left polarization, strong mode coupling;
4. RL[4, \*] – right polarization, strong mode coupling;
5. RL[5, \*] – left polarization, exact mode coupling.
6. RL[6, \*] – right polarization, exact mode coupling.

On input, these arrays specify the emission intensities at the start of the line-of-sight; on output, they contain the emission intensities at the end of the line-of-sight.

Return value:

- 0: no errors;
- -1: error (insufficient number of parameters);
- 1: error (incorrect parameters of the analytical electron distribution function);
- 2: error (incorrect parameters of the array-defined electron distribution function).

In case of any errors, the input/output array RL remains unchanged. *Note: the parameter checking has not been fully implemented yet, so that some invalid parameter combinations can pass without notice.*

## Function GET\_MW\_SLICE – multi-thread version

Calling syntax:

*For IDL users:*

```
res = call_external(libname, 'GET_MW_SLICE', $  
                    Lparms_M, Rparms_M, Parms_M, $  
                    E_arr, mu_arr, f_arr_M, RL_M)
```

*For Python users:* firstly, you need to load the libraries:

```
import GScodes
```

```
GET_MW_SLICE = GScodes.initGET_MW_SLICE(libname)
```

Then, when necessary, the function is called as:

```
res = GET_MW_SLICE(Lparms_M, Rparms_M, Parms_M,  
                  E_arr, mu_arr, f_arr_M, RL_M)
```

Function parameters:

0. Lparms\_M – 12-element long (32-bit) integer array of dimensions and global (for all voxels and LOSs) integer parameters (see below).
1. Rparms\_M – array of real parameters common for all voxels within each LOS,  $5 \times \text{Npix}$  elements, double (see below).
2. Parms\_M – array of voxel parameters,  $24 \times \text{Nz} \times \text{Npix}$  elements, double (see below).
3. E\_arr – array of energies where the electron distribution function is specified, NE elements, double, in MeV. This parameter is the same as in the GET\_MW function.
4. mu\_arr – array of pitch-angle cosines where the electron distribution function is specified, Nmu elements, double. This parameter is the same as in the GET\_MW function.
5. f\_arr\_M – array of electron distribution functions,  $\text{NE} \times \text{Nmu} \times \text{Nz} \times \text{Npix}$  elements, double, in  $\text{cm}^{-3} \text{MeV}^{-1}$  (see below).
6. RL\_M – input/output array,  $7 \times \text{Nf} \times \text{Npix}$  elements, double (see below).

Array of dimensions and global integer parameters Lparms\_M:

```
Lparms_M = [Npix, Nz, Nf, NE, Nmu, Nnodes, $  
            match_key, Qopt_key, arr_key, log_key, PK_key, spline_key]
```

0. Lparms\_M[0] = Npix – number of LOSs.

Other elements (1<sup>st</sup> to 11<sup>th</sup>) are respectively the same as the 0<sup>th</sup> to 10<sup>th</sup> elements of the Lparms array in the GET\_MW function. In particular:

- all LOSs have the same number of voxels Nz;
- the number of frequencies Nf is the same for all LOSs (although the frequency grids can be different);
- the energy and pitch-angle grids (including their dimensions NE and Nmu) are the same in all voxels of all LOSs;

- all other global parameters and keys (Nnodes, match\_key, Qopt\_key, arr\_key, log\_key, PK\_key, spline\_key) are applied to all voxels of all LOSs.

Other parameters: sub-arrays Rparms\_M[\* , i], Parms\_M[\* , \*, i], f\_arr\_M[\* , \*, \*, i] and RL\_M[\* , \*, i] correspond respectively to the parameters Rparms, Parms, f\_arr and RL of the single-thread GET\_MW function, for *i*th LOS.

Return value:

- 0: no errors;
- -1: error (insufficient number of parameters); the input/output array RL\_M remains unchanged;
- 1: error (incorrect parameters of an electron distribution function in, at least, one of the LOSs); the elements of the input/output array RL\_M corresponding to those incorrect LOSs remain unchanged.