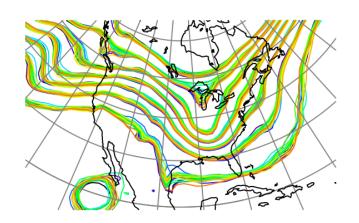


DART_LAB Tutorial Section 5: Adaptive Inflation

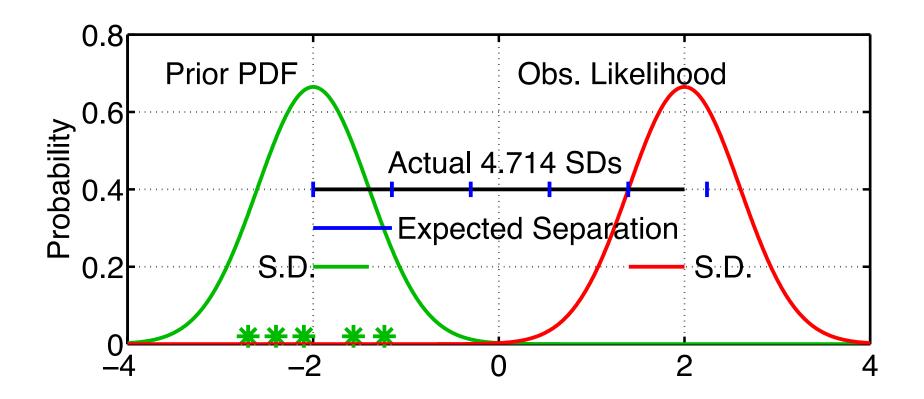




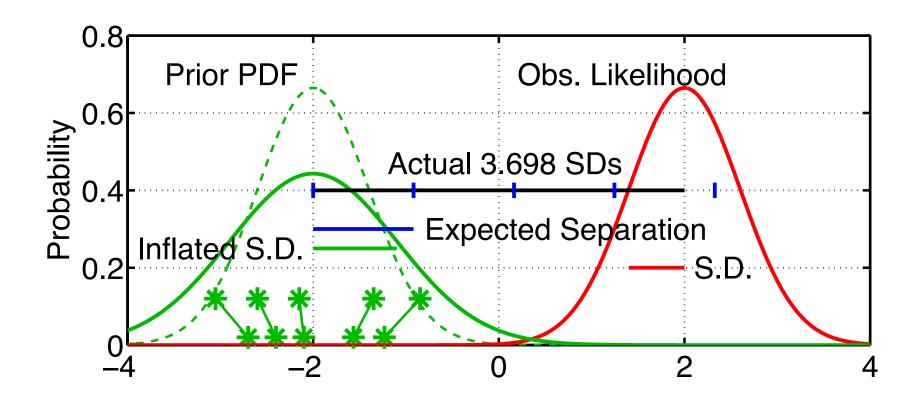
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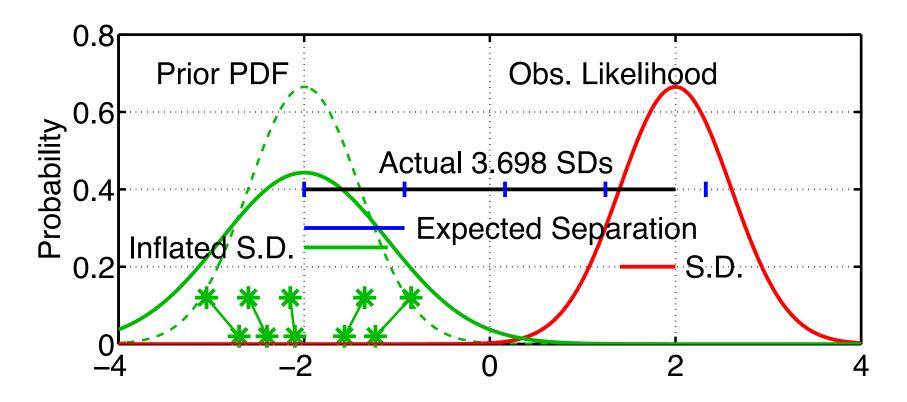




- 1. For observed variable, have estimate of prior-observed inconsistency.
- 2. Expected (prior_mean observation) = $\sqrt{\sigma_{prior}^2 + \sigma_{obs}^2}$ Assumes that prior and observation are supposed to be unbiased. Is it model error or random chance?



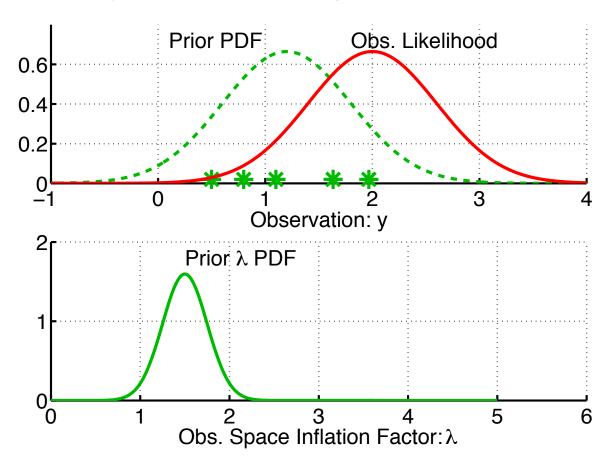
- 1. For observed variable, have estimate of prior-observed inconsistency.
- 2. Expected (prior_mean observation) = $\sqrt{\sigma_{prior}^2 + \sigma_{obs}^2}$
- 3. Inflating increases expected separation Increases 'apparent' consistency between prior and observation.



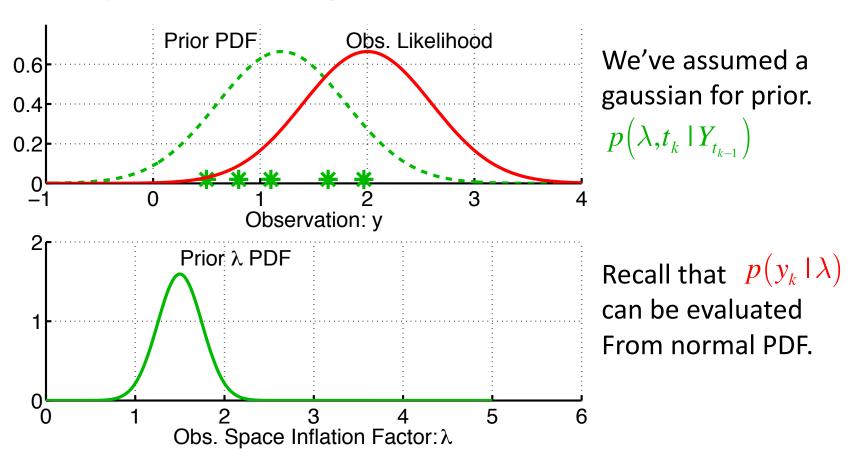
Distance D from prior mean y to obs is $N\left(0,\sqrt{\lambda\sigma_{prior}^2+\sigma_{obs}^2}\right)=N\left(0,\theta\right)$

Prob y₀ is observed given λ : $p(y_o \mid \lambda) = (2\pi\theta^2)^{-1/2} \exp(-D^2/2\theta^2)$

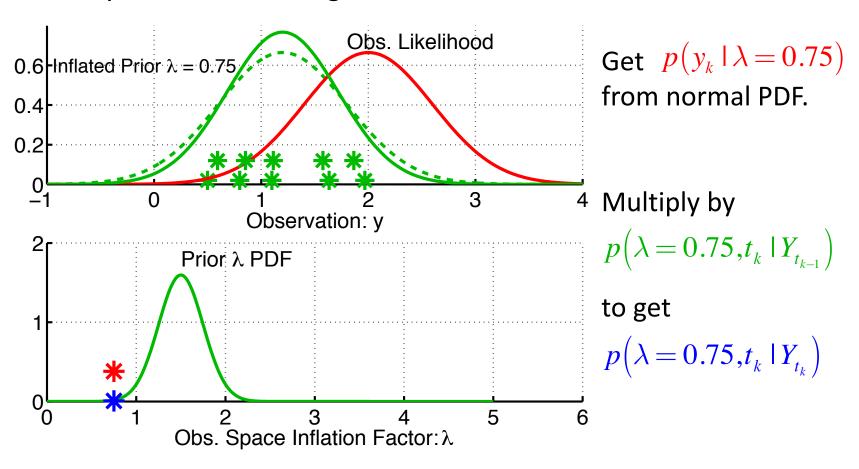
Use Bayesian statistics to get estimate of inflation factor λ .



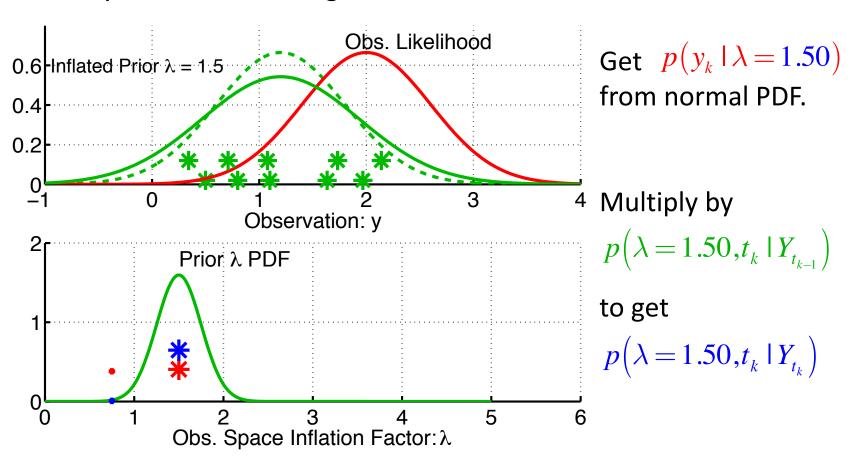
Assume prior is gaussian: $p(\lambda, t_k \mid Y_{t_{k-1}}) = N(\overline{\lambda}_p, \sigma_{\lambda, p}^2)$



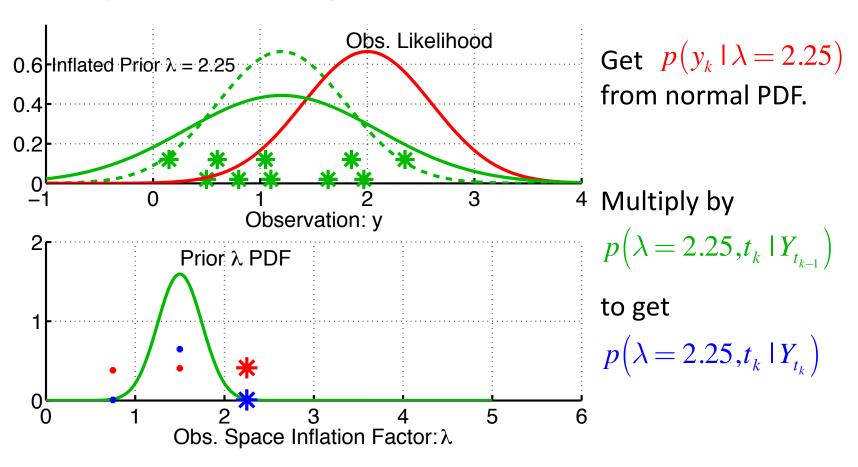
$$p(\lambda, t_k | Y_{t_k}) = p(y_k | \lambda) p(\lambda, t_k | Y_{t_{k-1}}) / normalization$$



$$p(\lambda,t_k \mid Y_{t_k}) = p(y_k \mid \lambda) p(\lambda,t_k \mid Y_{t_{k-1}}) / normalization$$

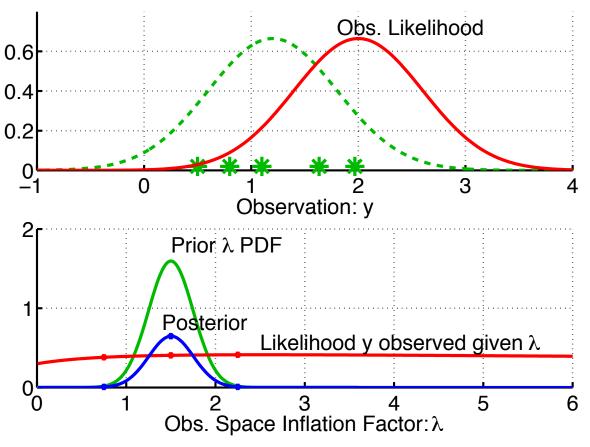


$$p(\lambda,t_k \mid Y_{t_k}) = p(y_k \mid \lambda) p(\lambda,t_k \mid Y_{t_{k-1}}) / normalization$$



$$p(\lambda, t_k | Y_{t_k}) = p(y_k | \lambda) p(\lambda, t_k | Y_{t_{k-1}}) / normalization$$

Use Bayesian statistics to get estimate of inflation factor λ .

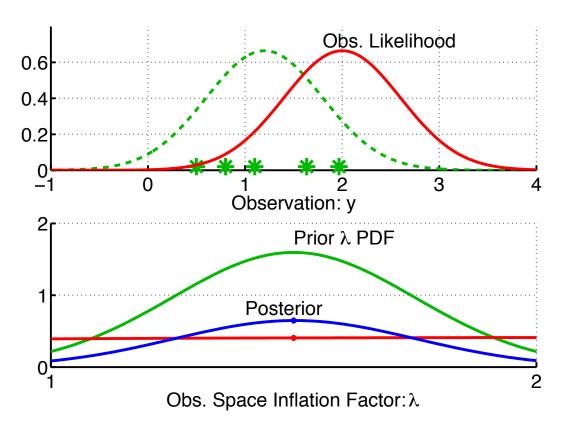


Repeat for a range of values of λ .

Now must get posterior in same form as prior (gaussian).

$$p(\lambda,t_k \mid Y_{t_k}) = p(y_k \mid \lambda) p(\lambda,t_k \mid Y_{t_{k-1}}) / normalization$$

Use Bayesian statistics to get estimate of inflation factor λ .



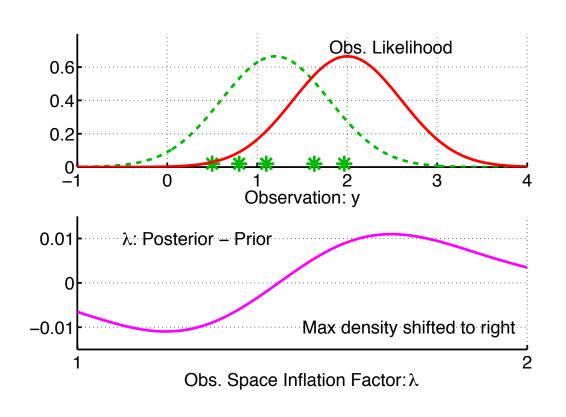
Very little information about λ in a single observation.

Posterior and prior are very similar.

Normalized posterior indistinguishable from prior.

$$p(\lambda, t_k | Y_{t_k}) = p(y_k | \lambda) p(\lambda, t_k | Y_{t_{k-1}}) / normalization$$

Use Bayesian statistics to get estimate of inflation factor λ .



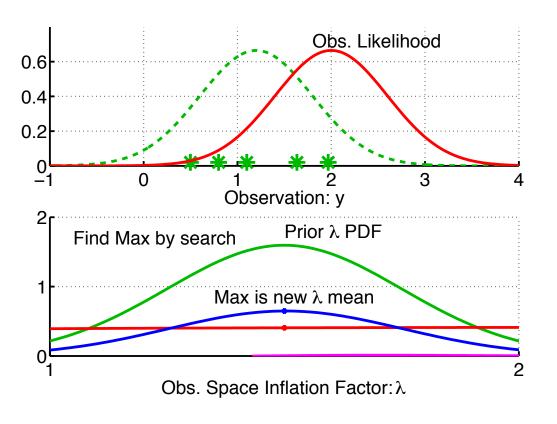
Very little information about λ in a single observation.

Posterior and prior are very similar.

Difference shows slight shift to larger values of λ .

$$p(\lambda,t_k \mid Y_{t_k}) = p(y_k \mid \lambda) p(\lambda,t_k \mid Y_{t_{k-1}}) / normalization$$

Use Bayesian statistics to get estimate of inflation factor λ .



One option is to use Gaussian prior for λ .

Select max (mode) of posterior as mean of updated Gaussian.

Do a fit for updated standard deviation.

$$p(\lambda,t_k \mid Y_{t_k}) = p(y_k \mid \lambda) p(\lambda,t_k \mid Y_{t_{k-1}}) / normalization$$

A. Computing updated inflation mean, $\overline{\lambda}_{u}$.

Mode of $p(y_k | \lambda) p(\lambda, t_k | Y_{t_{k-1}})$ can be found analytically!

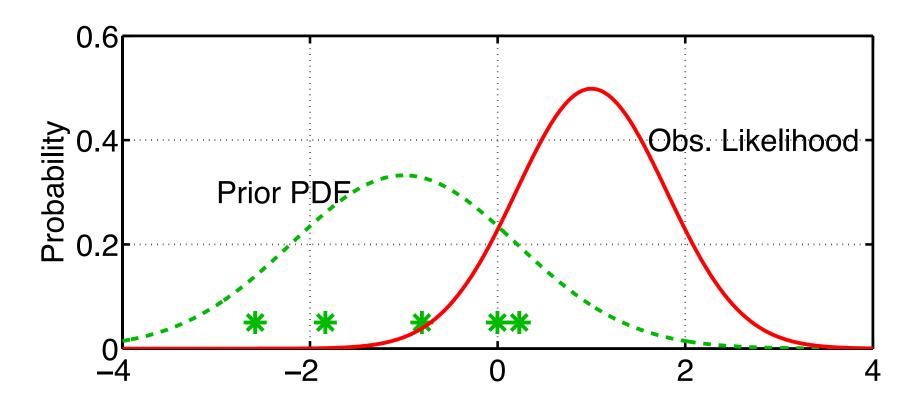
Solving
$$\partial \left[p(y_k \mid \lambda) p(\lambda, t_k \mid Y_{t_{k-1}}) \right] / \partial y = 0$$
 leads to 6^{th} order poly in θ .

This can be reduced to a cubic equation and solved to give mode.

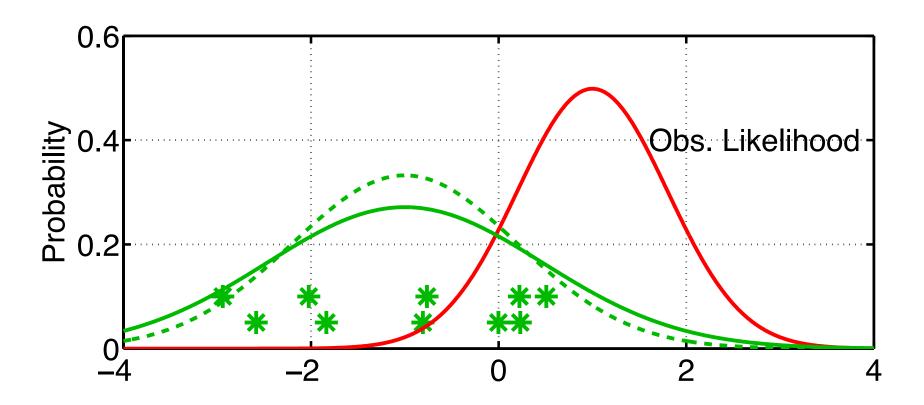
New $\overline{\lambda}_u$ is set to the mode.

This is relatively cheap compared to computing regressions.

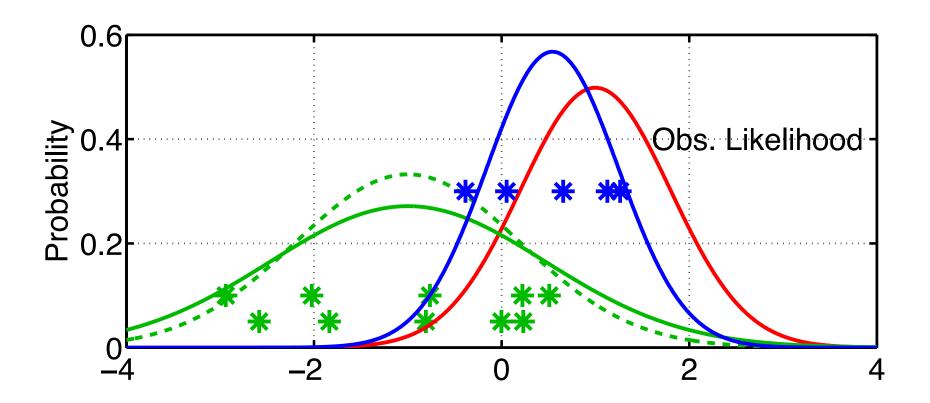
- B. Computing updated inflation variance, $\sigma_{\lambda,u}^2$.
 - 1. Evaluate numerator at mean $\overline{\lambda}_{\!\scriptscriptstyle u}$ and second point, e.g. $\overline{\lambda}_{\!\scriptscriptstyle u}+\sigma_{\lambda,p}$
 - 2. Find $\sigma_{\lambda,u}^2$ so $N(\overline{\lambda}_u,\sigma_{\lambda,u}^2)$ goes through $p(\overline{\lambda}_u)$ and $p(\overline{\lambda}_u+\sigma_{\lambda,p})$
 - 3. Compute as $\sigma_{\lambda,u}^2 = -\sigma_{\lambda,p}^2 / 2 \ln r$ where $r = p(\overline{\lambda}_u + \sigma_{\lambda,p}) / p(\overline{\lambda}_u)$



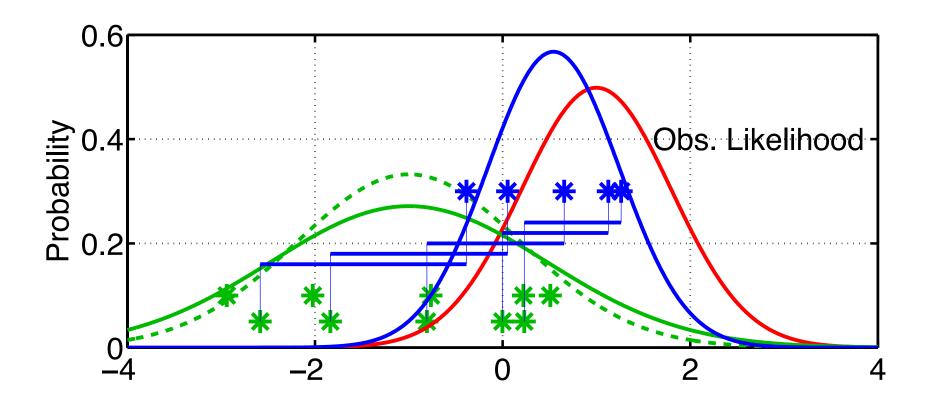
1. Compute updated inflation distribution, $p(\lambda, t_k | Y_{t_k})$.



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- 2. Inflate ensemble using mean of updated $\hat{\lambda}$ distribution.



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- 2. Inflate ensemble using mean of updated $\hat{\lambda}$ distribution.
- 3. Compute posterior for y using inflated prior.

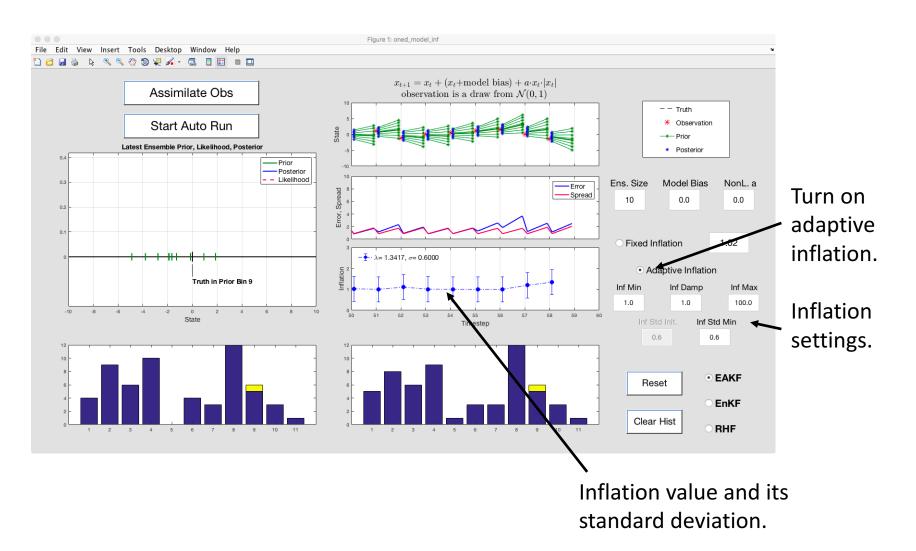


- 1. Compute updated inflation distribution $p(\lambda, t_k | Y_{t_k})$
- 2. Inflate ensemble using mean of updated λ distribution.
- 3. Compute posterior for y using inflated prior.
- 4. Compute increments from ORIGINAL prior ensemble.

Adaptive inflation can be tested with matlab script *oned_model_inf.m*

Can explore 5 different values that control adaptive inflation:

- Minimum value of inflation, often set to 1 (no deflation).
- Inflation damping, more on this later. Value of 1.0 turns it off.
- Maximum value of inflation.
- The initial value of the inflation standard deviation.
- A lower bound on inflation standard deviation (it will asymptote to zero if allowed).



Try adaptive inflation.

Pick a lower value for standard deviation. Initial lower bound on inflation of 1.0, upper bound large (100).

Try introducing a model bias.

What happens if lower bound is less than 1?

Inflation Damping

Inflation mean damped towards 1 every assimilation time.

• *inf_damping* 0.9: 90% of the inflation difference from 1.0 is retained.

Can be useful in models with heterogeneous observations in time. For instance, a well-observed hurricane crosses a model domain. Adaptive inflation increases along hurricane trace. After hurricane, fewer observations, no longer need so much inflation.

For large earth system models, following values may work:

```
inf_sd_initial = 0.6,
inf_damping = 0.9,
inf_sd_lower_bound = 0.6.
```

Adaptive Multivariate Inflation Algorithm

Suppose we want a global multivariate inflation, λ_s , instead.

Make same least squares assumption that is used in ensemble filter.

Inflation of λ_s for state variables inflates obs. priors by same amount.

Get same likelihood as before: $p(y_o \mid \lambda) = (2\pi\theta^2)^{-1/2} \exp(-D^2/2\theta^2)$

$$\theta = \sqrt{\lambda_s \sigma_{prior}^2 + \sigma_{obs}^2}$$

Compute updated distribution for λ_s exactly as for single observed variable.

Implementation of Adaptive Multivariate Inflation Algorithm

- 1. Apply inflation to state variables with mean of λ_s distribution.
- 2. Do following for observations at given time sequentially:
 - a. Compute forward operator to get prior ensemble.
 - b. Compute updated estimate for λ_s mean and variance.
 - c. Compute increments for prior ensemble.
 - d. Regress increments onto state variables.

Spatially varying adaptive inflation algorithm

Have a distribution for λ for each state variable, $\lambda_{s,i}$.

Use prior correlation from ensemble to determine impact of $\lambda_{s,i}$ on prior variance for given observation.

If γ is correlation between state variable i and observation then

$$\theta = \sqrt{\left[1 + \gamma\left(\sqrt{\lambda_{s,i}} - 1\right)\right]^2 \sigma_{prior}^2 + \sigma_{obs}^2}$$

Equation for finding mode of posterior is now full 12th order:

Analytic solution appears unlikely.

Can do Taylor expansion of θ around $\lambda_{s,i}$.

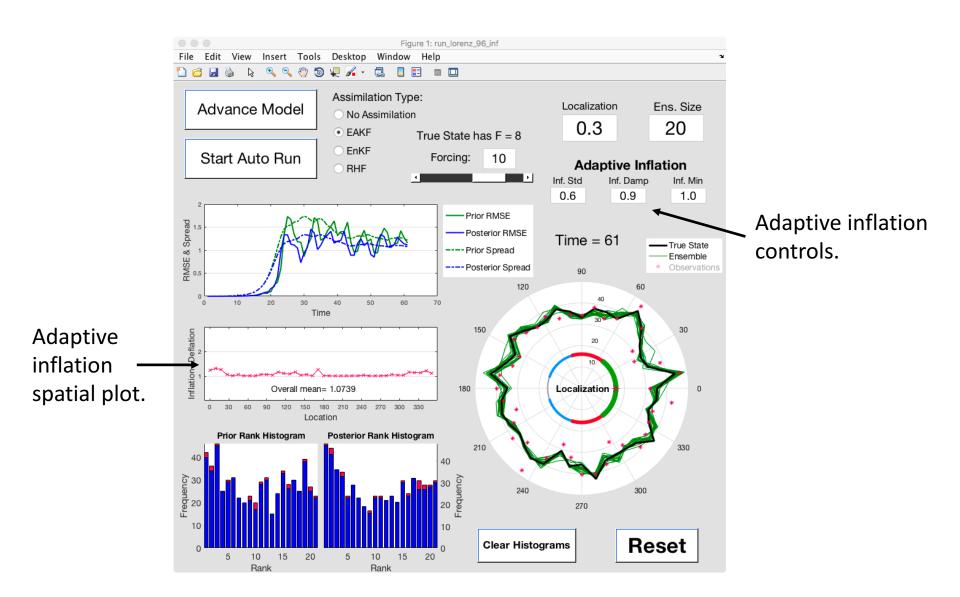
Retaining linear term is normally quite accurate.

There is an analytic solution to find mode of product in this case!

Spatially Varying Adaptive inflation can be tested with matlab script run_lorenz96_inf.m

Can explore 3 different values that control adaptive inflation:

- Minimum value of inflation, often set to 1 (no deflation).
- Inflation damping. Value of 1.0 turns it off.
- The value of the inflation standard deviation.
 - Lower bound on standard deviation is set to same value.
 - In this case, standard deviation just stays fixed at selected value.



Explore some of the following:

How does adaptive inflation change as localization is changed? How does adaptive inflation change for different values of inflation standard deviation? If the lower bound is smaller than 1, does deflation (inflation < 1) happen?

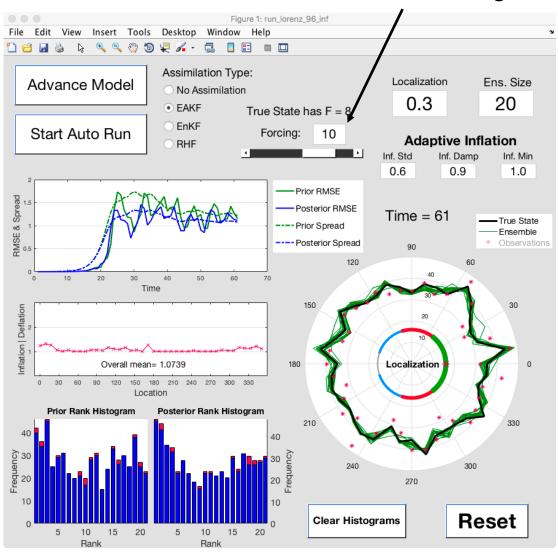
Simulating Model Error in 40-Variable Lorenz-96 Model

Inflation can deal with all sorts of errors, including model error.

Can simulate model error in Lorenz96 by changing forcing. Synthetic observations are from model with forcing = 8.0.

Both run_lorenz_96 and run_lorenz_96_inf allow model error.

Model error: Change forcing for assimilating model here.



Spatially Varying Adaptive Inflation with Model Error

Explore some of the following:

Change the model forcing to a larger or smaller value (say 6 or 10).

How does adaptive inflation respond to model error?

Do good values of localization change as model error increases?