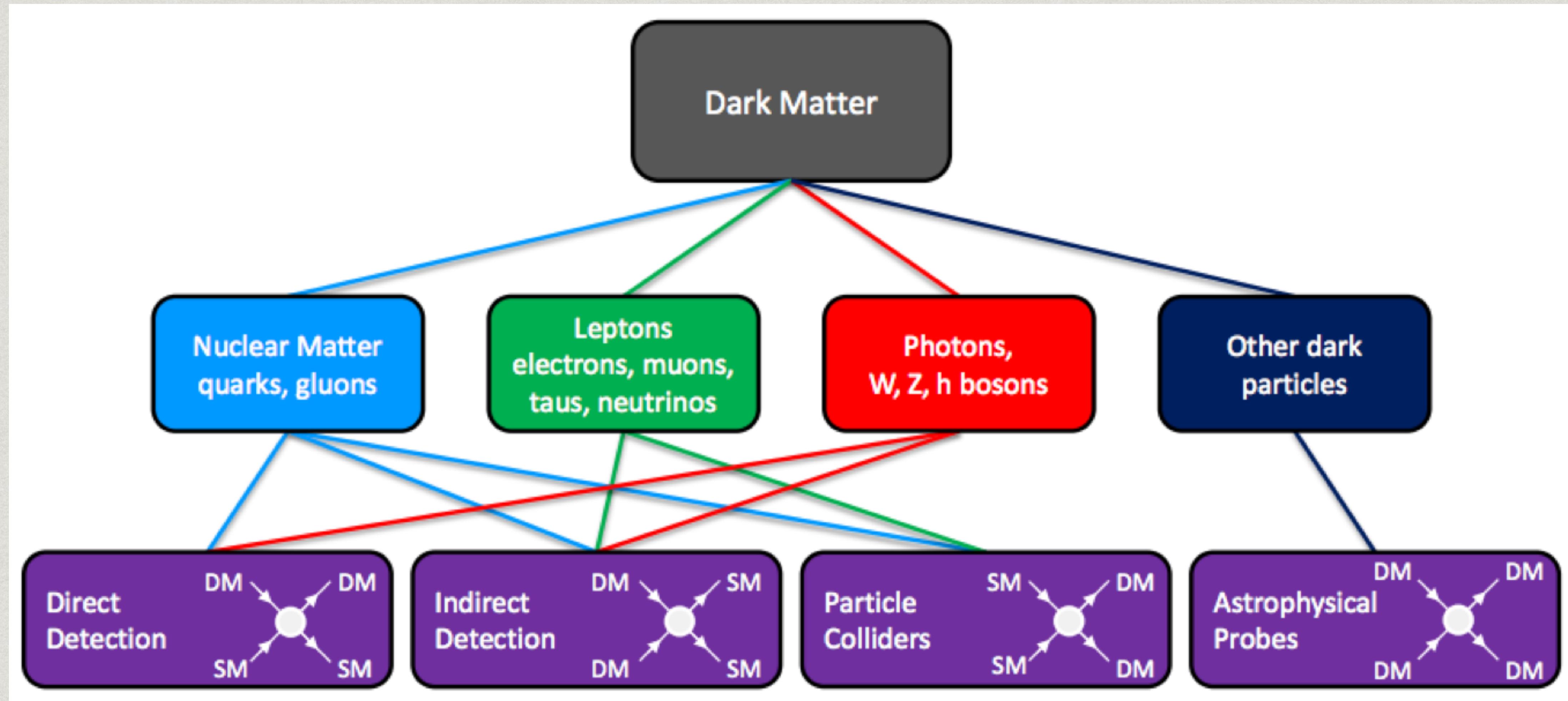


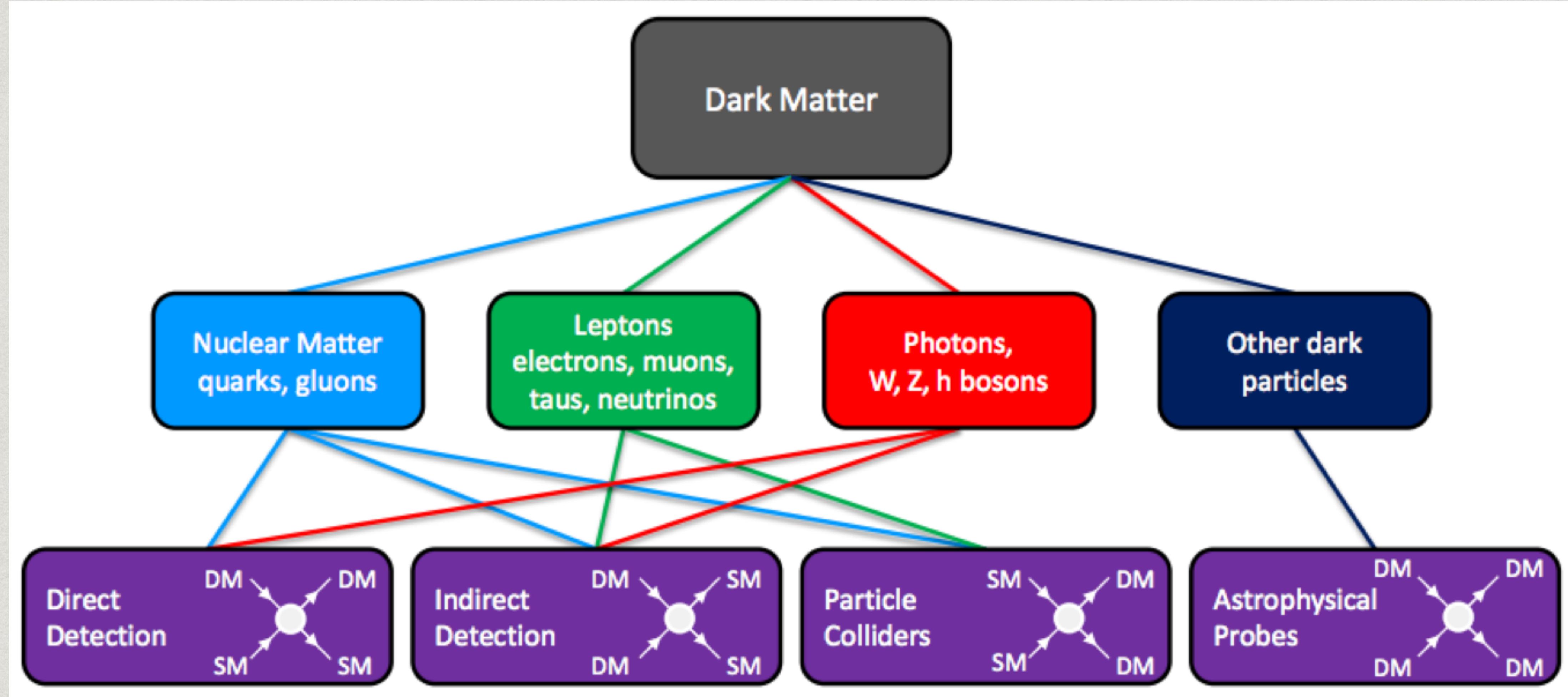
ASTRONOMICAL INSIGHTS INTO DARK MATTER PARTICLE CONSTRAINTS

WILL DAWSON (LLNL)

Probes of Dark Matter



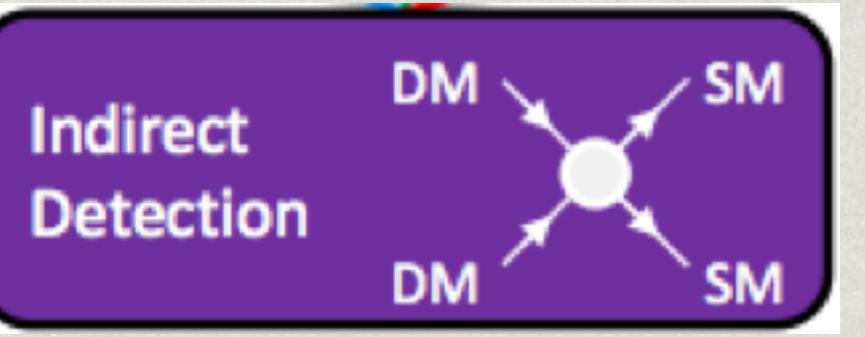
Astrophysical probes are leading the way on two of the avenues to studying dark matter



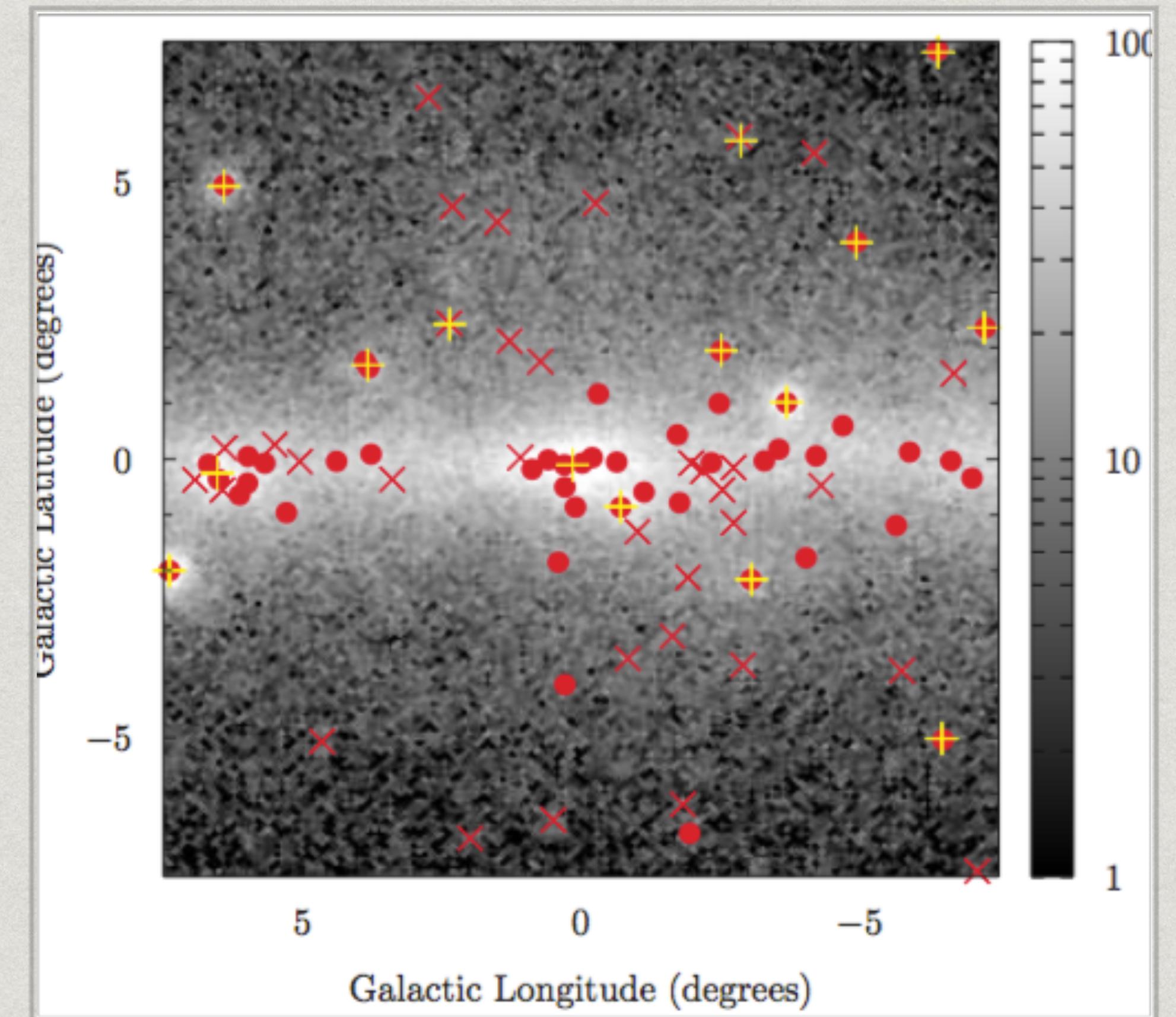
Observational Dark Matter Probes

Astrophysical DM Probe	Scattering DM	Annihilating DM	Warm DM	Observatory/Survey/Method
Stellar Streams	x		x	Gaia, DES, HSC, LSST, Euclid, WFIRST
Gamma-ray Emission (MW, dwarf, clusters, merging clusters)		x		Fermi, Cherenkov telescopes
Number of dwarf satellites			x	DES, HSC, LSST, Euclid, WFIRST
Ly-alpha forest			x	Optical spectroscopy, etc.?
CMB	x	x	x	Planck, ACT, SPT, etc.?
Dark matter halo profiles (dwarf to cluster)	x			Optical spectroscopy, strong lensing, weak lensing
Galaxy strong lensing			x	DES, HSC, LSST, Euclid, WFIRST, HST, Ground-based AO
Cluster-cluster mergers	x	x		Strong and weak lensing, optical spectroscopy, X-ray
Galaxy-cluster mergers	x			Strong lensing, optical spectroscopy
Neutrino emission (from sun etc.)	x	x		Icecube , etc.
Velocity distribution of stars in MW	x		x	Gaia, 4MOST, MKSE, WEAVE

Gamma-ray Emission



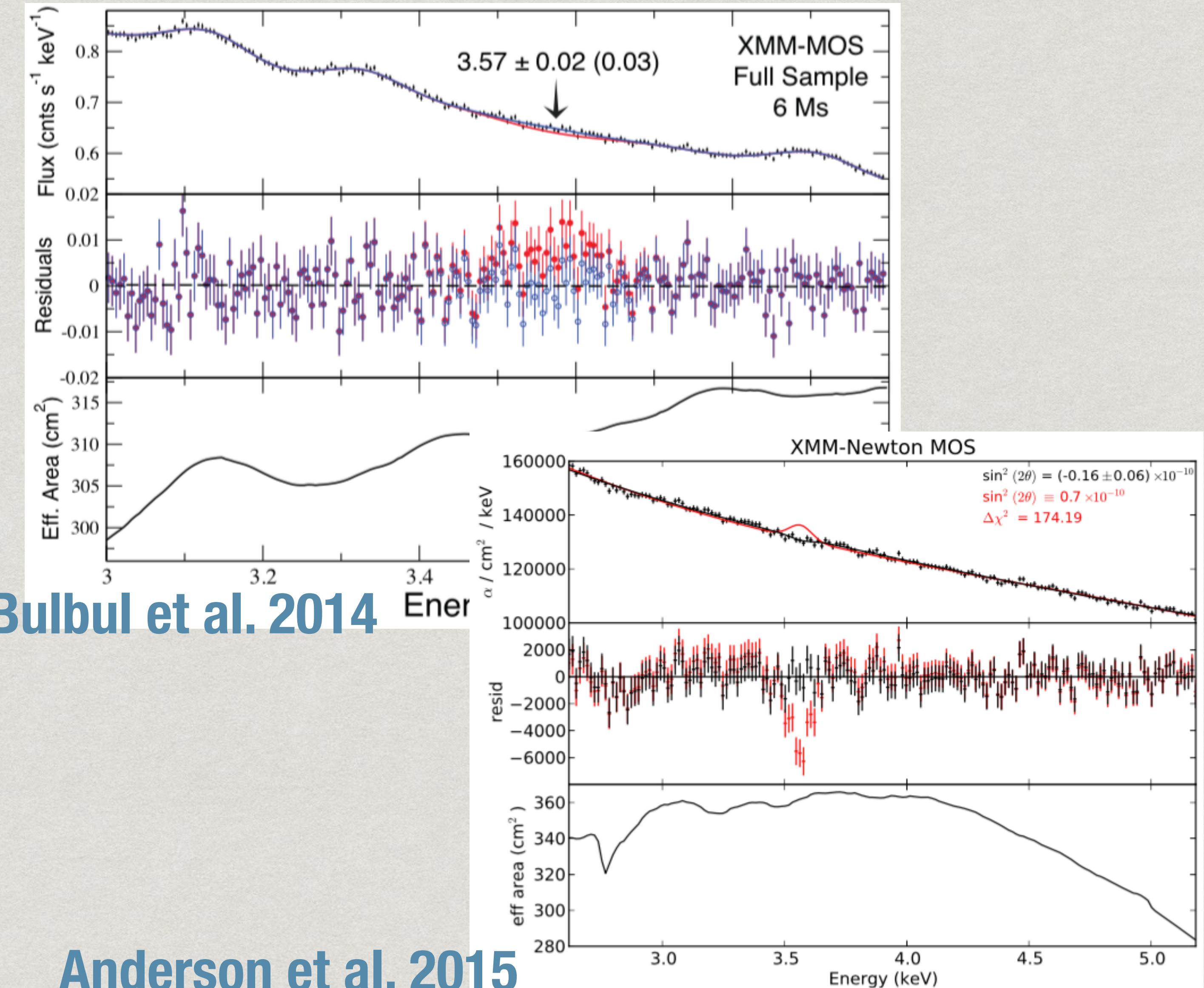
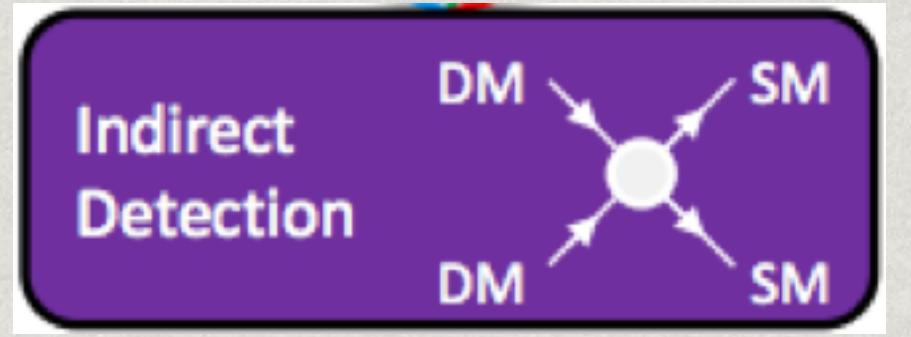
- * Milky Way Center
- * Stacked Clusters
- * Stacked Dwarfs
- * (Others?)
- * Possibly how current studies could be improved?



Fermi-LAT Collaboration 2015

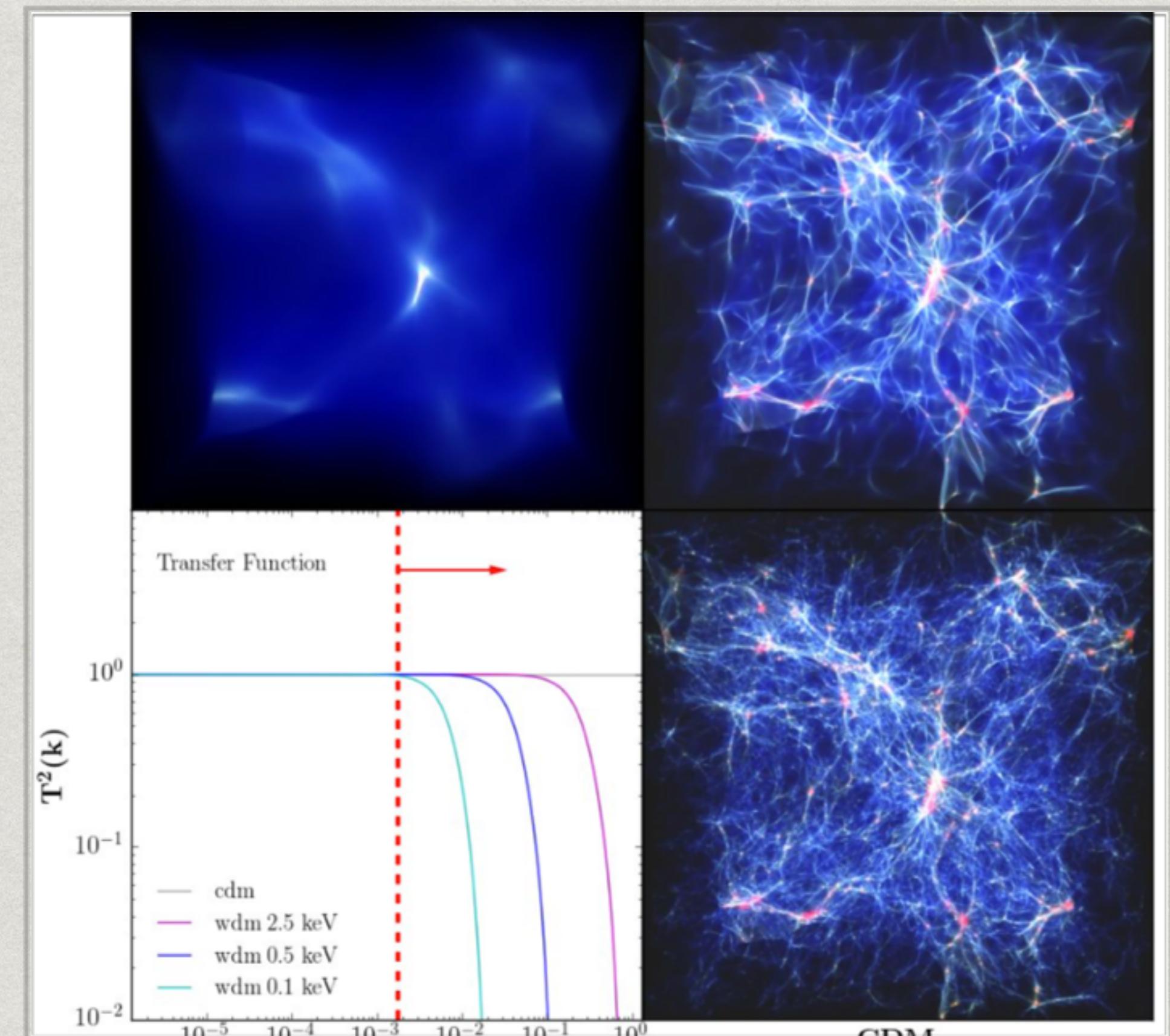
X-ray Emission

- * Stacked galaxy clusters (merging galaxy clusters)
- * Stacked galaxies
- * Others?



Lyman-alpha forest

- * constraint on dark matter temperature (there-by mass)
- * Tightest constraint on warm dark matter



Neutrinos from possible dark matter decay in center of sun

indirect searches for DM candidates

DM-induced neutrinos:

$$\text{KK, XX, SS} \rightarrow \left\{ \begin{array}{l} \text{qq} \\ \text{l+l-} \\ \text{W, Z, H} \end{array} \right\} \rightarrow \nu$$

Kaluza-Klein modes an additional useful channel:

$$\text{KK} \rightarrow \nu\nu$$

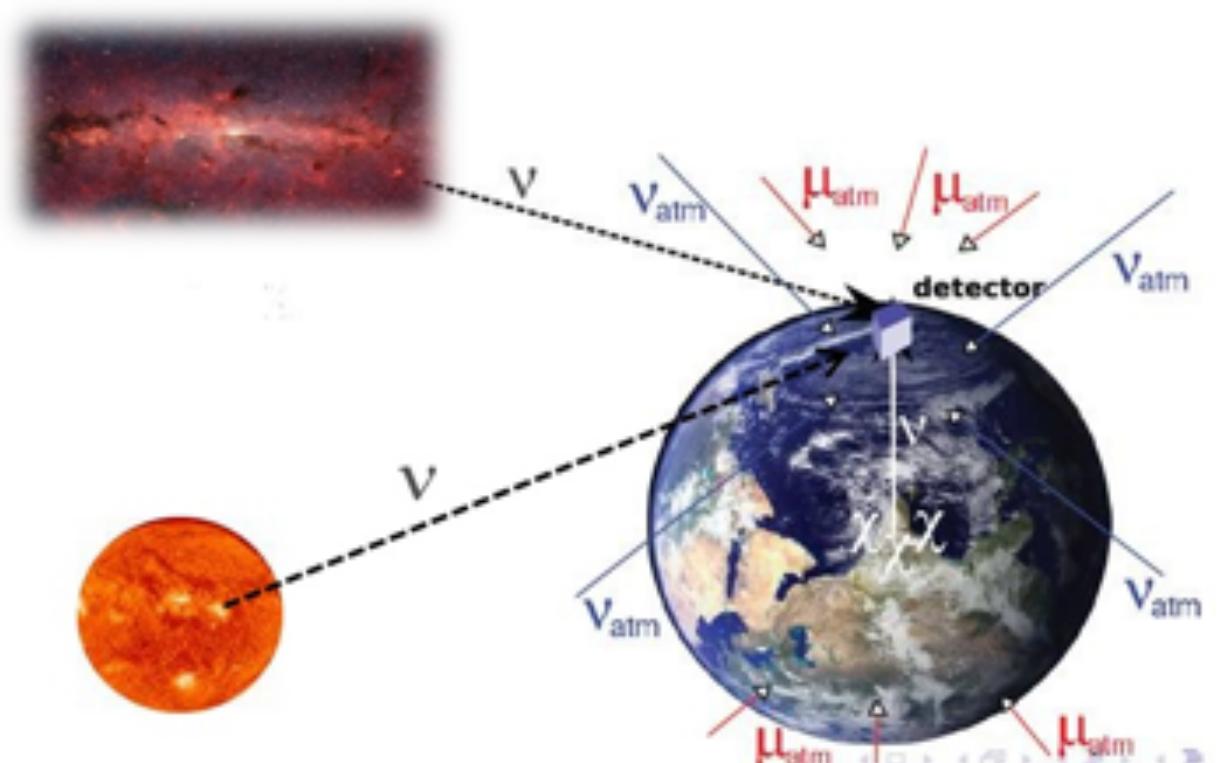
signature: ν excess over background from Sun/Earth/Galactic Halo direction

A lot of physics uncertainties involved:

- relic density calculations
- DM distribution in the halo
- velocity distribution
- χ, K, S properties (MSSM/UED...)
- interaction of χ, K, S with matter (capture)
- self interaction (annihilation)

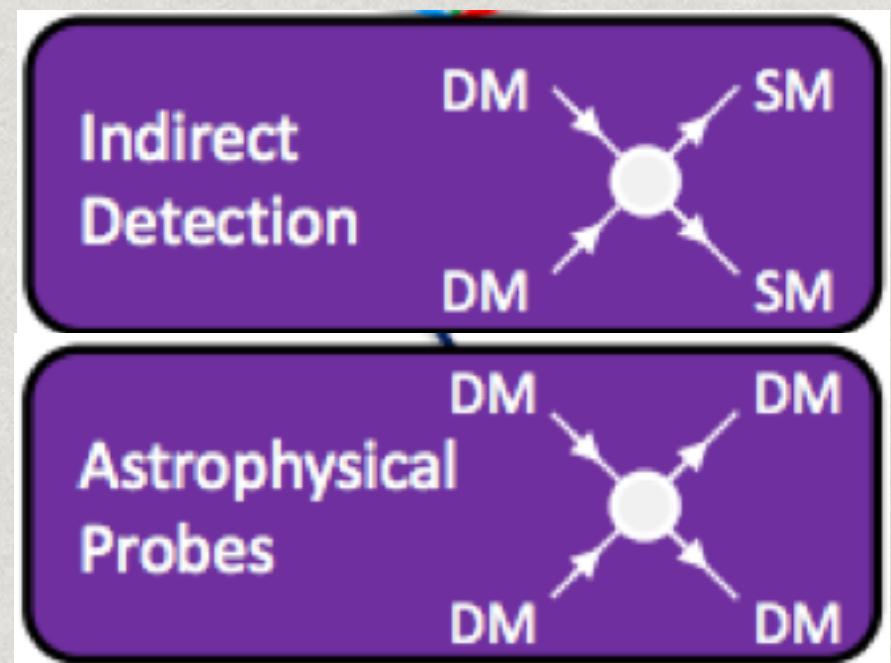
look at objects where the WIMP can be gravitationaly trapped and annihilate:

Sun, Earth and Galactic Halo



Atmospheric muons $\sim O(10^9)$ events/year (downwards)

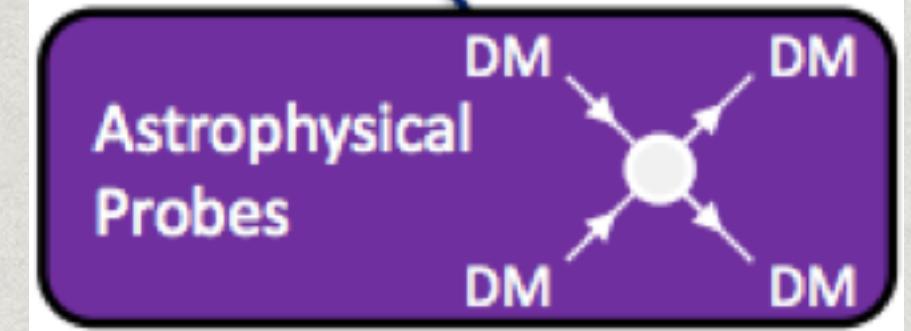
Atmospheric neutrinos $\sim O(10^3)$ events/year (all directions)



- * using flux limits for annihilation products of DM that are at the center of the sun due to DM-DM interactions

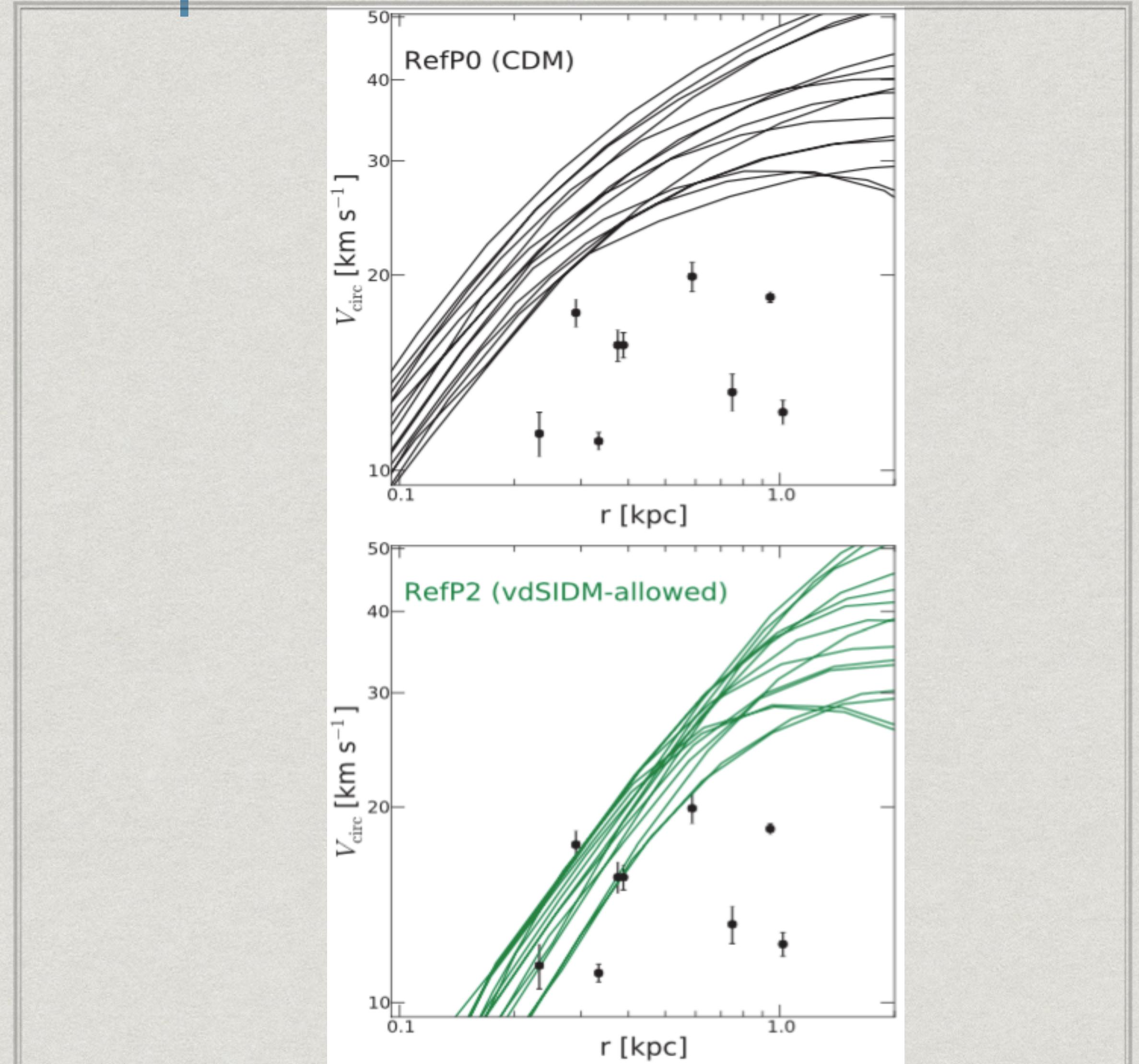
Slide from Carlos de los Heros

Dwarf galaxies: mass profile

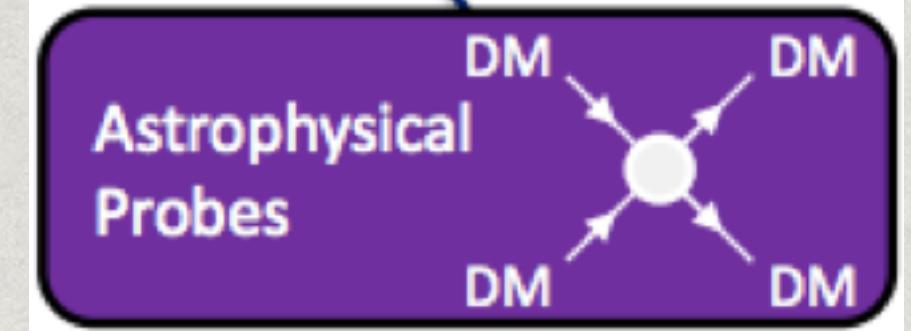


- * cusp-core problem
- * too big to fail
- * how they could be explained by standard model effects

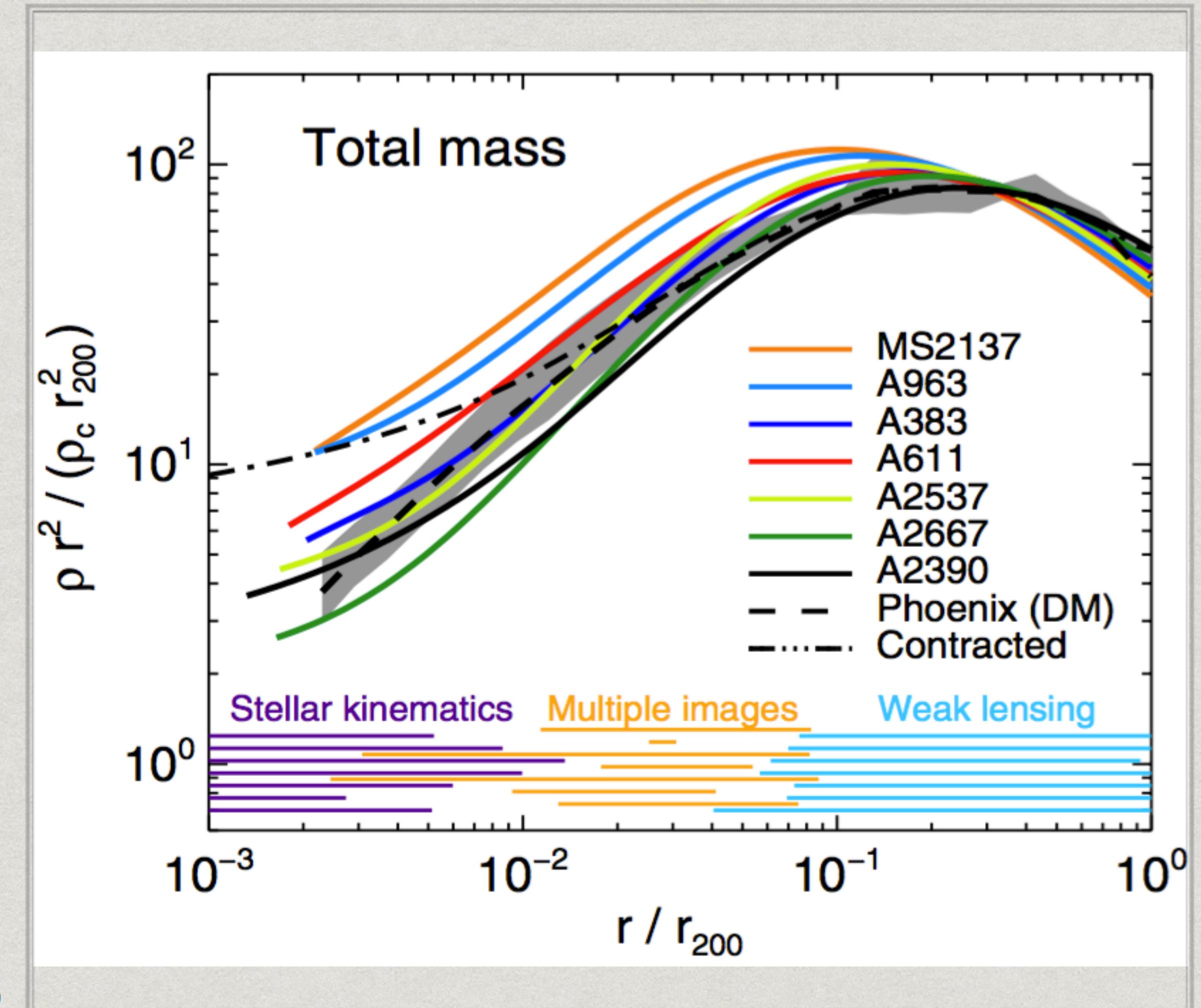
Vogelsberger et al. 2012



Galaxy clusters: mass profile



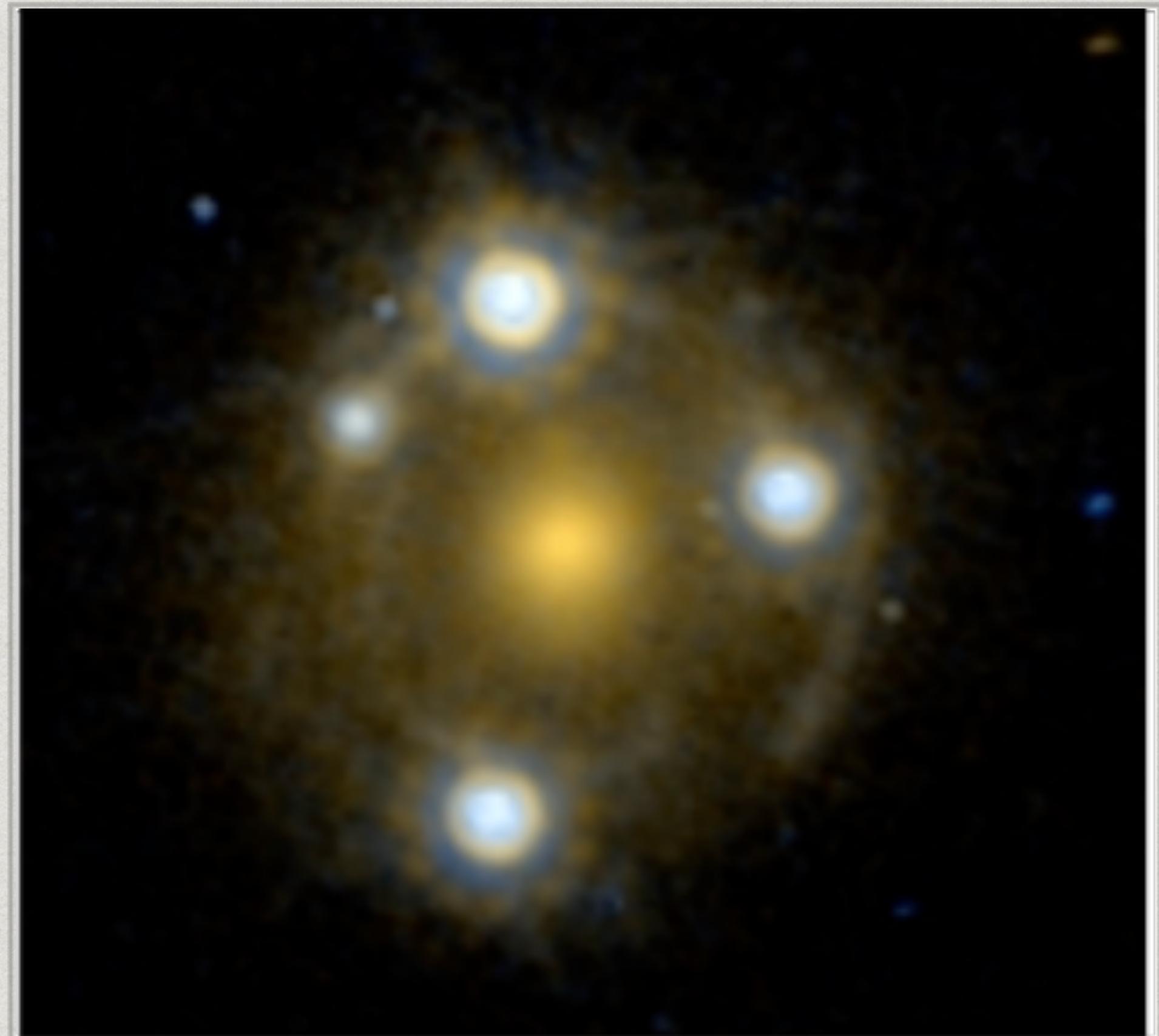
- * Inner density profile
- * degeneracy with baryonic effects



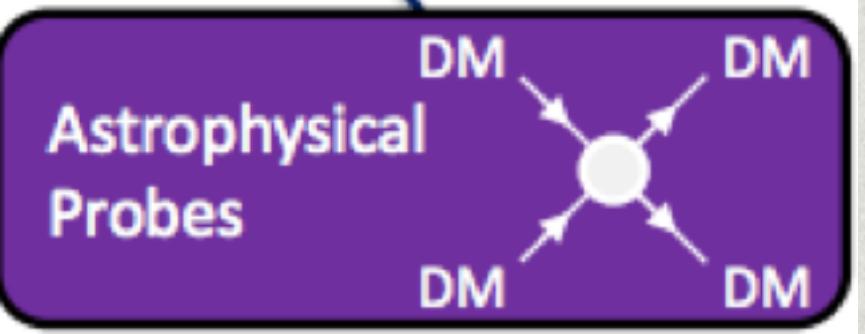
Newman et al. 2012

Galaxy-galaxy strong lensing

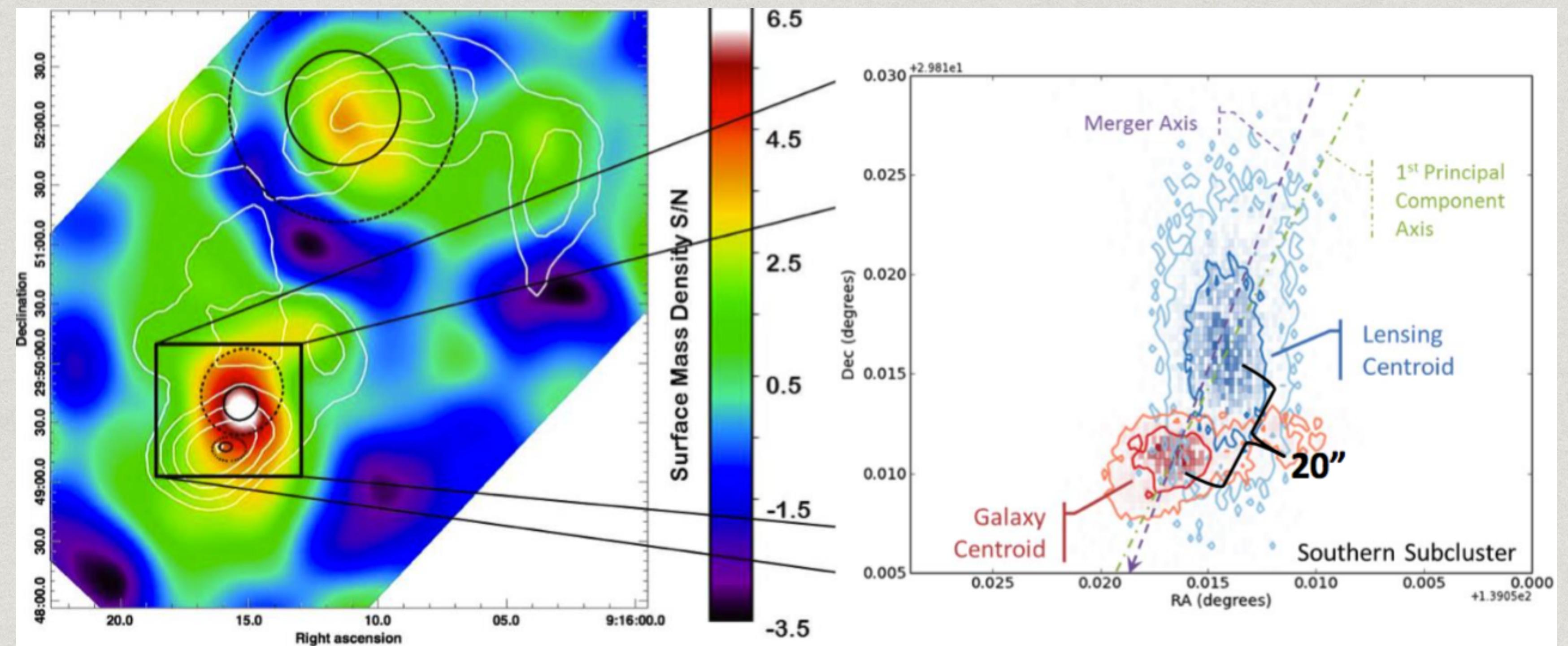
- * Probe of amount of substructure



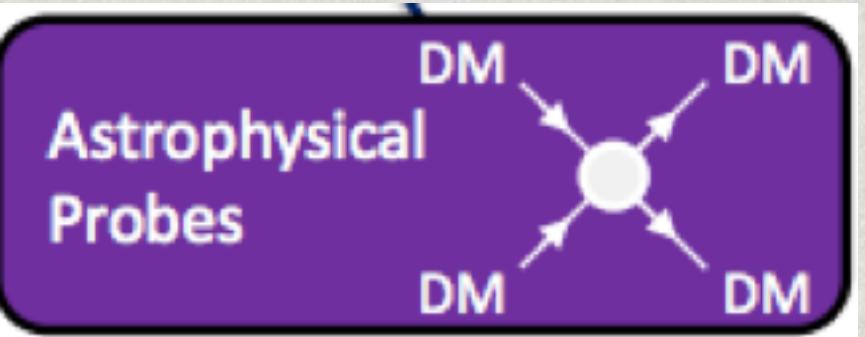
Merging Galaxy Clusters



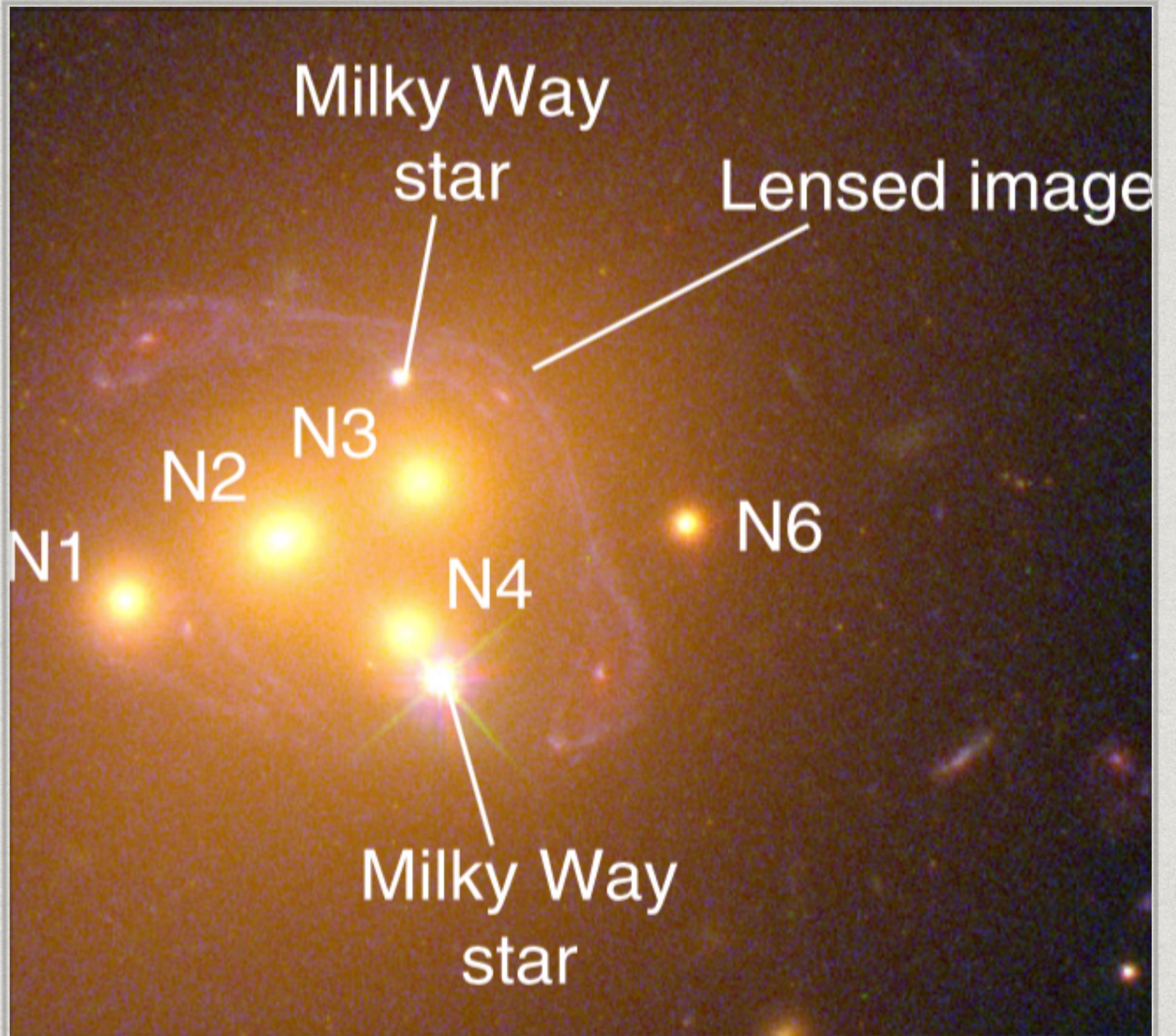
- ✳ Offset between galaxies and dark matter
 - * either offset of peak
 - * or rain drop distribution



In-falling galaxies

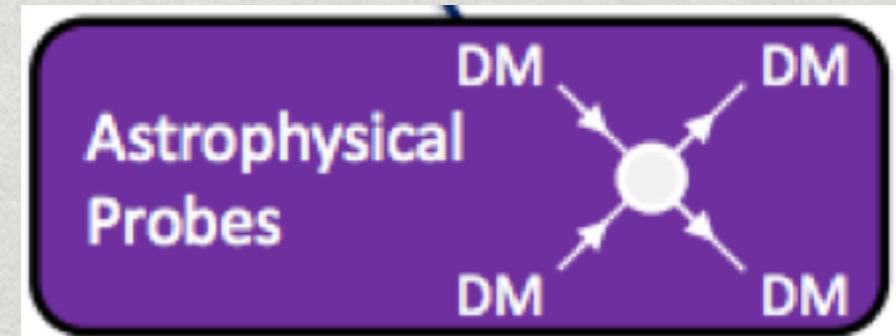


- * Galaxy offset from its dark matter sub-halo
- * requires fortuitous strong lensing

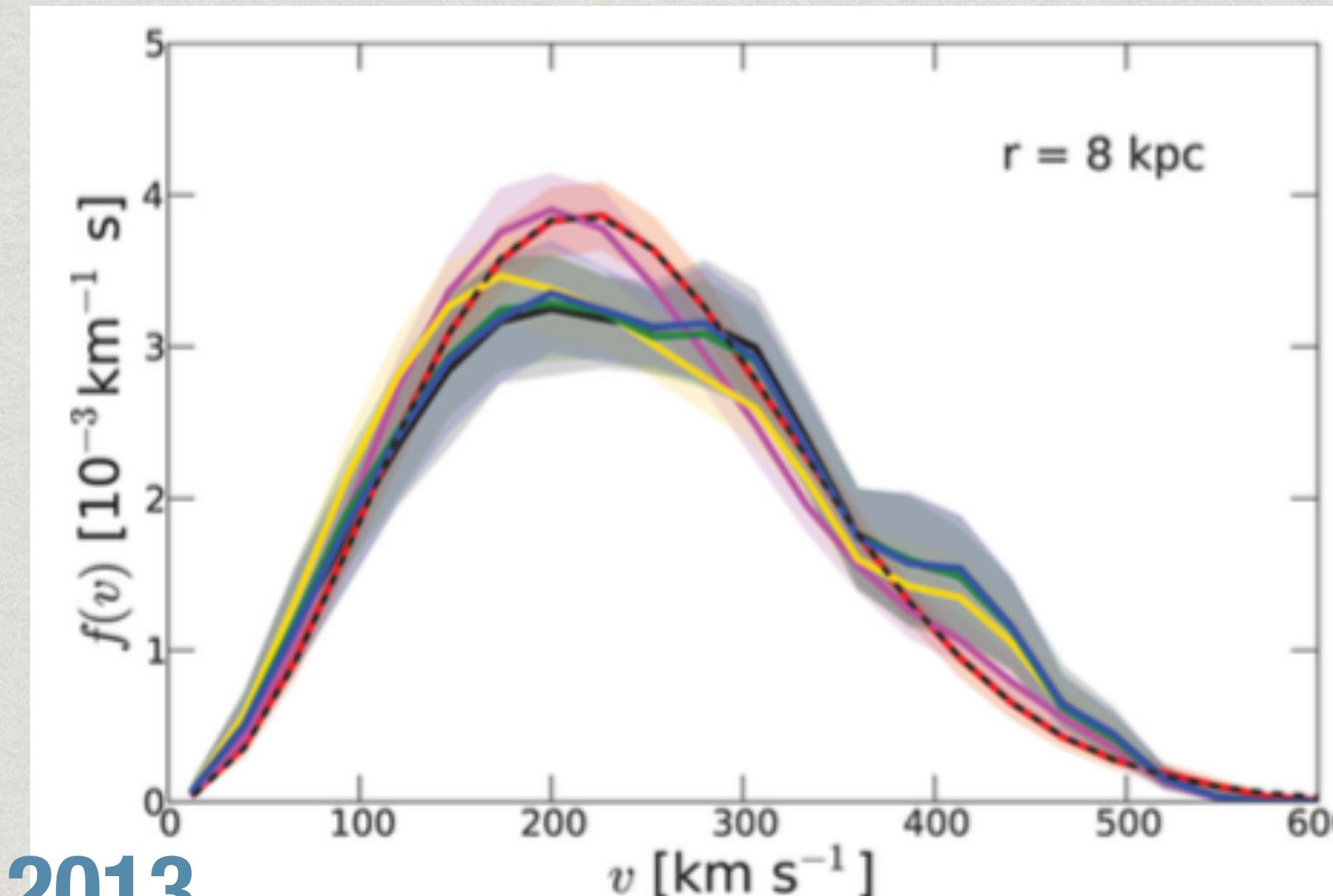


Massey et al. 2015

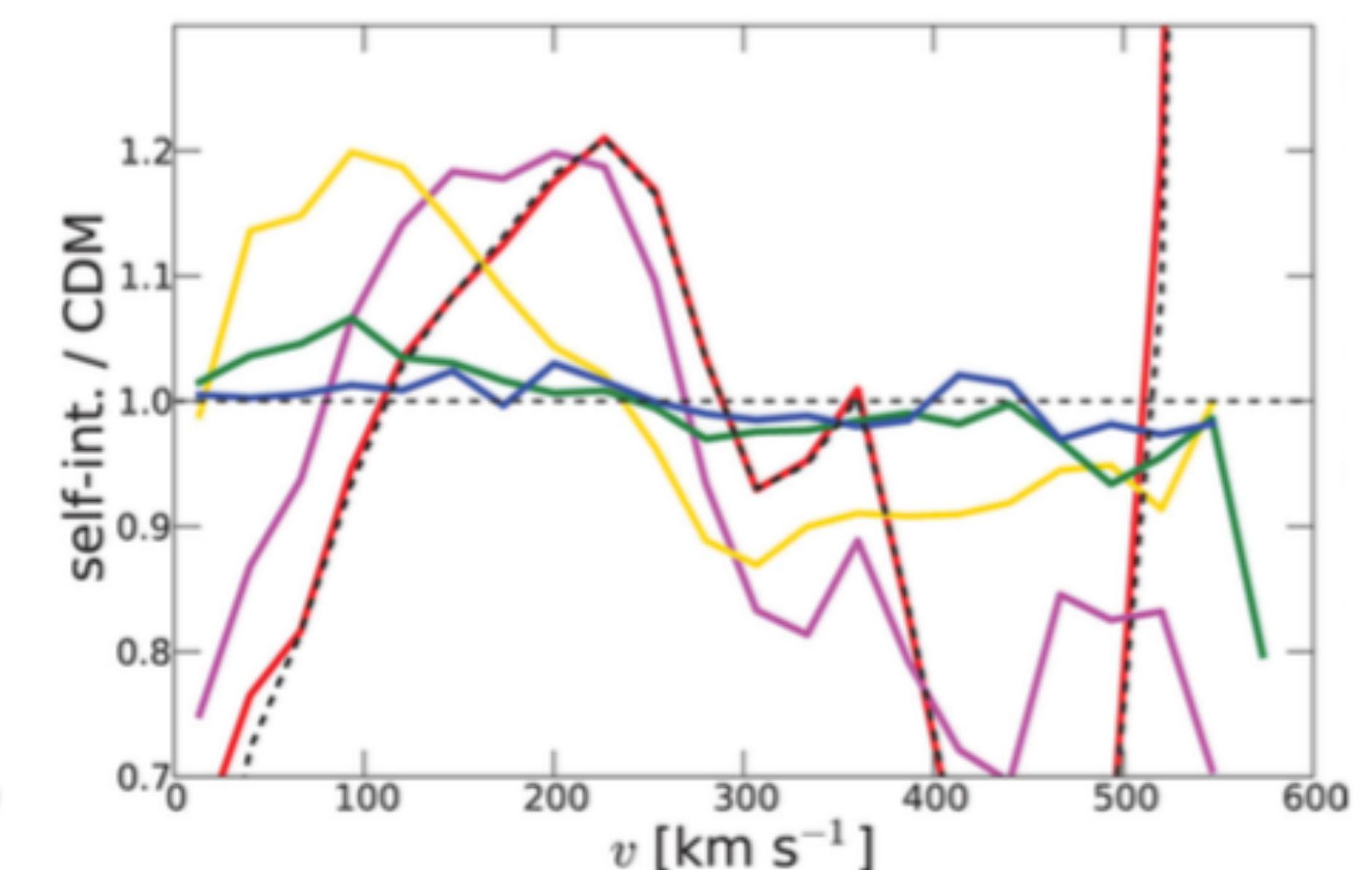
Stellar streams & velocity distribution in MW



- * Velocity distribution varies as a function of cross-section
- * Gaia



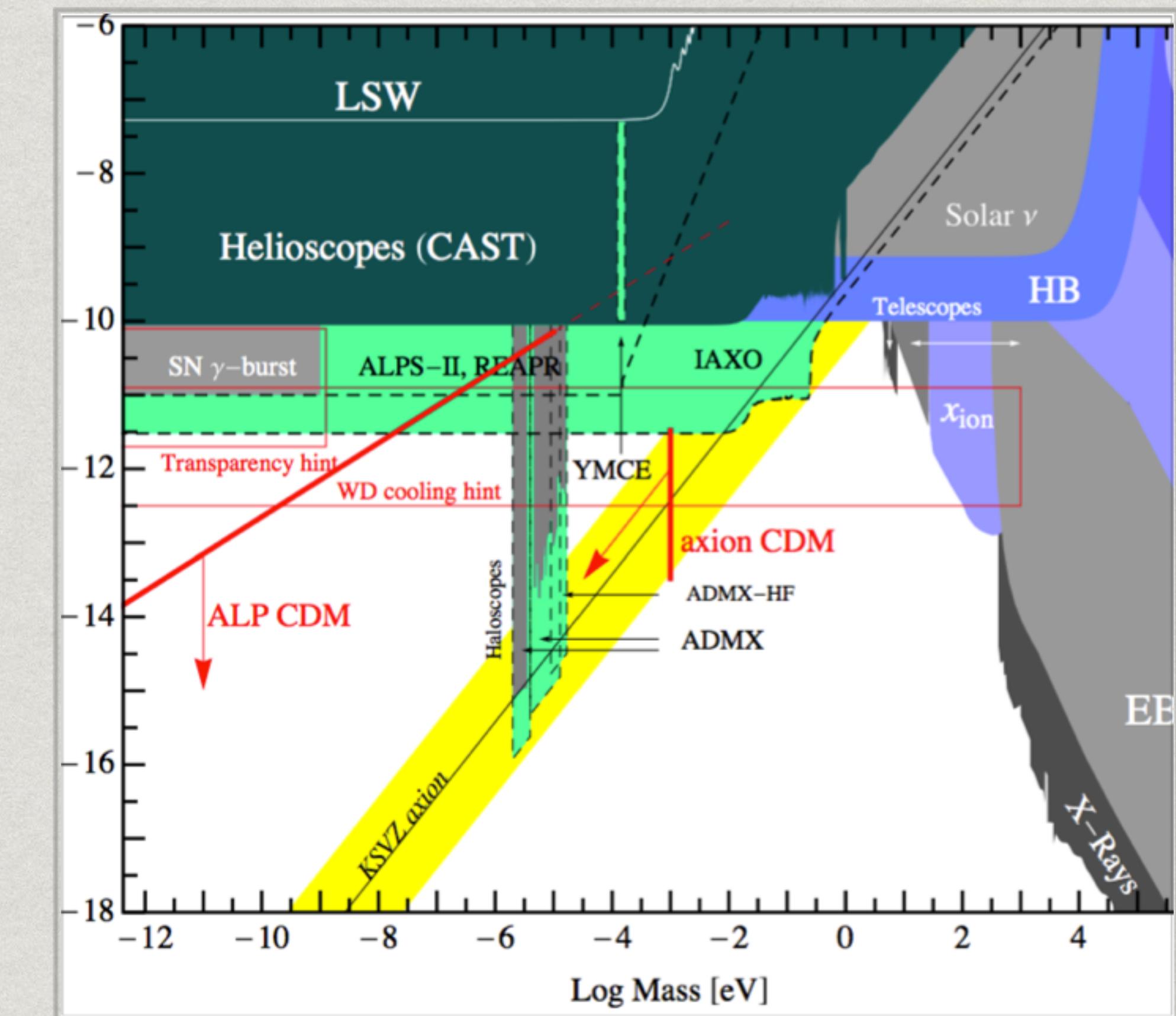
Vogelsberger et al. 2013



CMB BAO

Astrophysical constraints on axions

- * Need something about the astrophysical constraints of axion dark matter.
- * perhaps mention Bose-Einstein condensate constraints
- * See e.g. Roswald 2012 (blue and gray regions in figure to right) <http://adsabs.harvard.edu/abs/2012PDU.....1..116R>



Challenge to combine all of these constraints