# Solar Atmosphere Wave Dynamics Generated by Solar Global Oscillating Eigenmodes

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#### Abstract

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#### 1. Introduction

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Figure 1: This figure shows .....

of your manuscript.

## 2. An equation

$$\sum_{i=0}^{\infty} A^n \int dx \frac{F_n(x)}{A_n + B_n} = B^n C^n \int dx \int dy \frac{G_n(x, y)}{A_n x + B_n y}$$
(1)

## 3. Figures

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#### References

Elson, R.A.W., Santiago, B.X., & Gilmore, G.F. 1996, Halo stars, starbursts, and distant globular clusters: A survey of unresolved objects in the Hubble Deep Field, NewA, 1, 1-16.

Table 1: This is the caption of this table

Parameter	Value
Parameter 1	$526.849 \pm 0.003 \text{ s}$
Parameter 2	$5268.49 \pm 0.03 \text{ s}$
Parameter 3	$52684.9 \pm 0.3 \text{ s}$

- Governato, F., Moore, B., Cen, R., Stadel, J., Lake, G., & Quinn, T. 1997, The Local Group as a test of cosmological models, NewA 2, 91-106.
- Lamport, L., Lamport, L., Lamport, Addison-Wesley (Reading, Massachusetts, 1994).
- Wettig, T., & Brown, G.E. 1996, The evolution of relativistic binary pulsars, NewA, 1, 17-34.