

SURROGATE MODELS FOR GRAVITATIONAL WAVE ASTRONOMY

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GRAVITATIONAL WAVE TEMPLATE BANKS

- **SEOBNRv4** (Bohe' et al. 2017)

- Spin-aligned

Intrinsic parameters are

$$\chi_i = \pm \frac{|S_i|}{m_i^2}$$

$$\lambda = (q, \chi_1, \chi_2)$$

where the mass ratio is

$$q = m_1/m_2$$

- **Effective One Body:** binary motion is replaced by the effective one-body problem of a test particle in a deformed Schwarzschild metric, with the symmetric mass ratio as the deformation parameter:

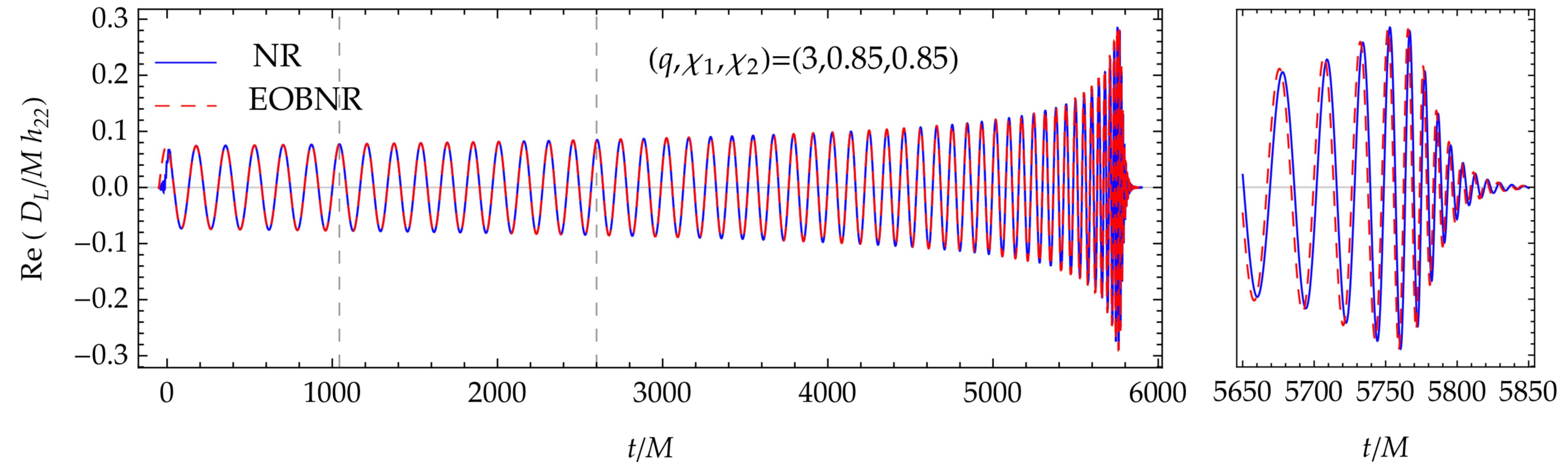
$$\eta = \frac{m_1 m_2}{M^2} = \frac{q}{(1+q)^2}$$

Includes all **4PN terms** in the EOB radial potential + higher-order PN terms that are calibrated with numerical relativity simulations.

Describes dominant (2,2) mode including the **merger-ringdown phase**.

SEOBNRv4 TEMPLATE BANK

- Numerical-Relativity Calibrated to 141 simulations

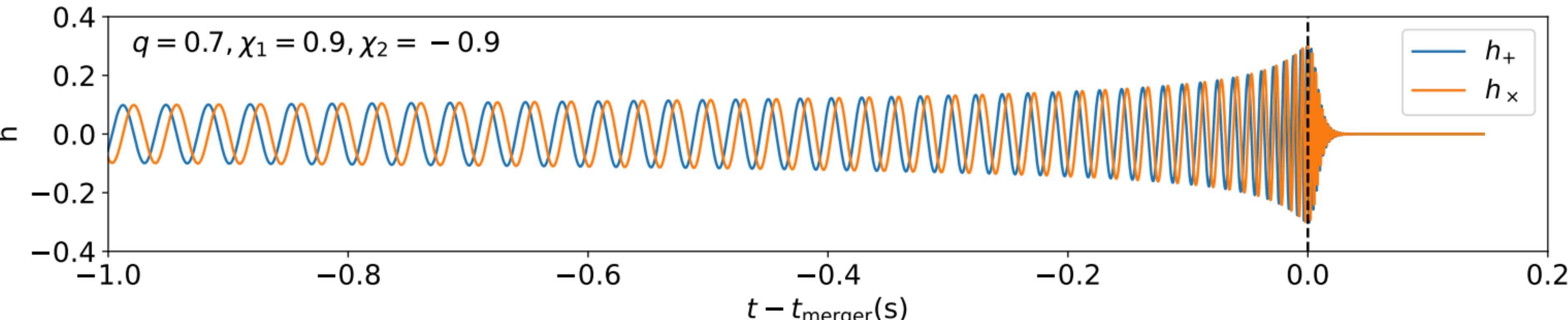


- $O(300,000)$ waveforms are used in detection and parameter estimation of BBH mergers using matched filtering -> very costly computations!
- Extension to random spins impractical -> construct fast surrogate models + use machine learning for rapid detection and parameter estimation!

SAMPLE SEOBNRv4 WAVEFORM

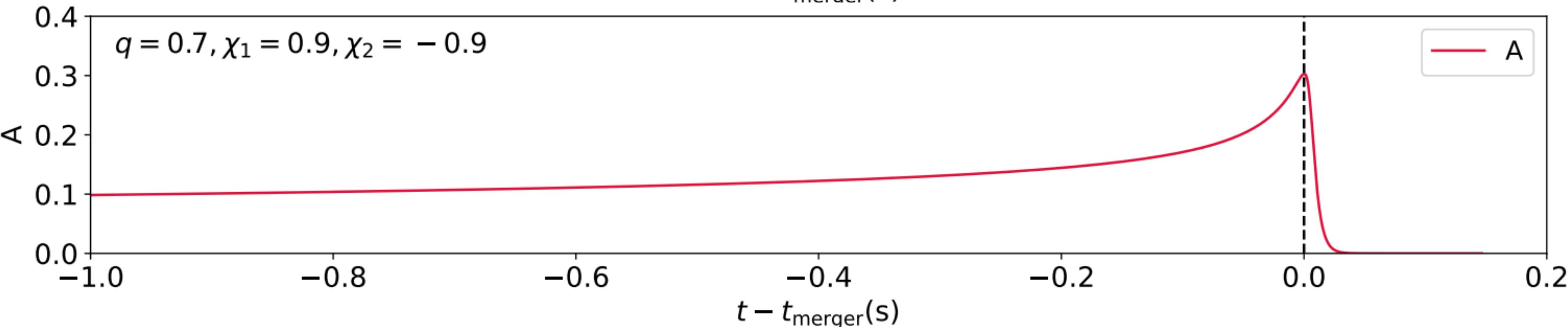
- Strain

$$h = h_+ - i h_\times$$



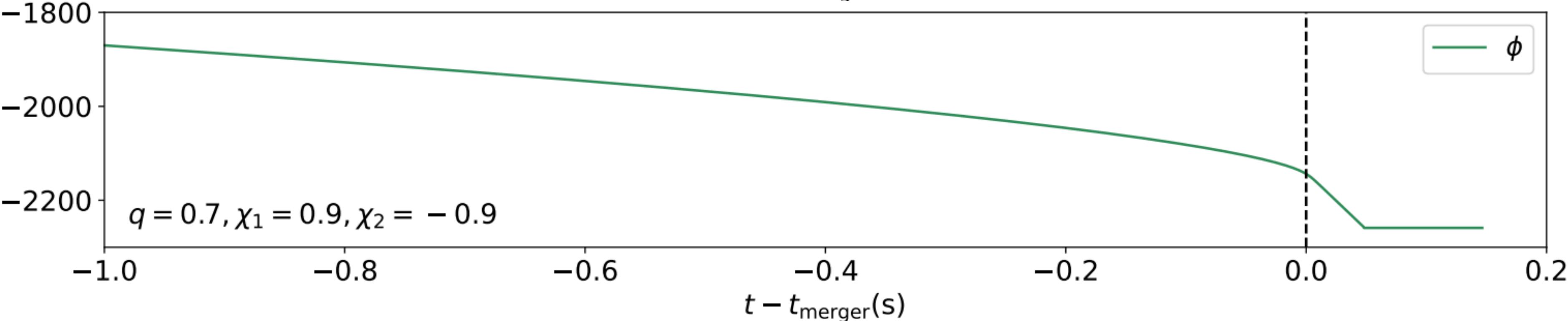
- Amplitude

$$A = \sqrt{h_+^2 + h_\times^2}$$



- Phase

$$\phi = \tan^{-1} \left(-\frac{h_\times}{h_+} \right)$$



REDUCED BASIS

- Complex waveform:

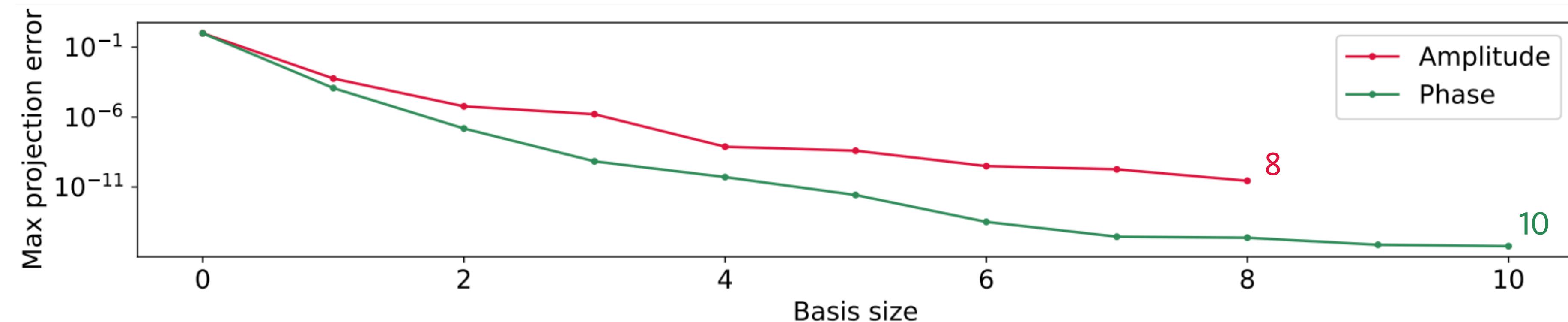
$$h(t; \boldsymbol{\lambda}) = h_+(t; \boldsymbol{\lambda}) - i h_\times(t; \boldsymbol{\lambda})$$

where $\boldsymbol{\lambda} = (q, \chi_1, \chi_2)$ = **intrinsic parameters**

- Use N waveforms, randomly chosen in range $0 \leq q \leq 8, -1 \leq \chi_{1,2} \leq 1$ with mass $M = 60M_\odot$
- Construct **reduced basis** $e_j(t)$ of size $n < N$, using an **iterative greedy algorithm** and Gram-Schmidt orthonormalization

$$h(t; \boldsymbol{\lambda}_i) \approx \sum_{j=1}^n c_j(\boldsymbol{\lambda}_i) e_j(t) \quad i = 1, \dots, N$$

- Example: nonspinning $N = 40$ & tolerance = 10^{-12} , then **8** for amplitude and **10** for phase.



EMPIRICAL INTERPOLATION METHOD

- Construct Empirical Interpolation Method (EIM) basis $B_k(t)$ of size n
- The coefficients $\alpha_k(\boldsymbol{\lambda}_j)$ coincide with the waveform at the empirical time nodes $\{T_k\}_{k=1}^n$, i.e.

$$\alpha_k(\boldsymbol{\lambda}_j) = h(T_k; \boldsymbol{\lambda}_j)$$

- Then, for the greedy points:

$$h(t; \boldsymbol{\lambda}_j) = \sum_{k=1}^n \alpha_k(\boldsymbol{\lambda}_j) B_k(t) \quad j = 1, \dots, n$$

- and for any other waveform in the training set:

$$\alpha_k(\boldsymbol{\lambda}_i) = h(T_k; \boldsymbol{\lambda}_i)$$

(no costly projection is required!)

- Calculations with Chad Galley's [Rompy](#) package.

MISMATCH BETWEEN TWO WAVEFORMS

- Inner product

$$\langle h(\cdot; \boldsymbol{\lambda}_1), h(\cdot; \boldsymbol{\lambda}_2) \rangle = 4\Re \int_{f_{min}}^{f_{max}} \frac{\tilde{h}(f; \boldsymbol{\lambda}_1)\tilde{h}^*(f; \boldsymbol{\lambda}_2)}{S_n(f)} df.$$

- Scaled waveform spectrum

$$\hat{h}(f; \boldsymbol{\lambda}) = \frac{\tilde{h}(f; \boldsymbol{\lambda})}{\langle h(\cdot; \boldsymbol{\lambda}), h(\cdot; \boldsymbol{\lambda}) \rangle}$$

power spectral
density (PSD)

- Overlap

$$\mathcal{O}(\hat{h}(\cdot; \boldsymbol{\lambda}_1), \hat{h}(\cdot; \boldsymbol{\lambda}_2)) = \max_{t_0, \phi_0} \langle h(\cdot; \boldsymbol{\lambda}_1), h(\cdot; \boldsymbol{\lambda}_2) \rangle$$

- Mismatch

$$\mathcal{M}(\hat{h}(\cdot; \boldsymbol{\lambda}_1), \hat{h}(\cdot; \boldsymbol{\lambda}_2)) = 1 - \mathcal{O}(\hat{h}(\cdot; \boldsymbol{\lambda}_1), \hat{h}(\cdot; \boldsymbol{\lambda}_2))$$

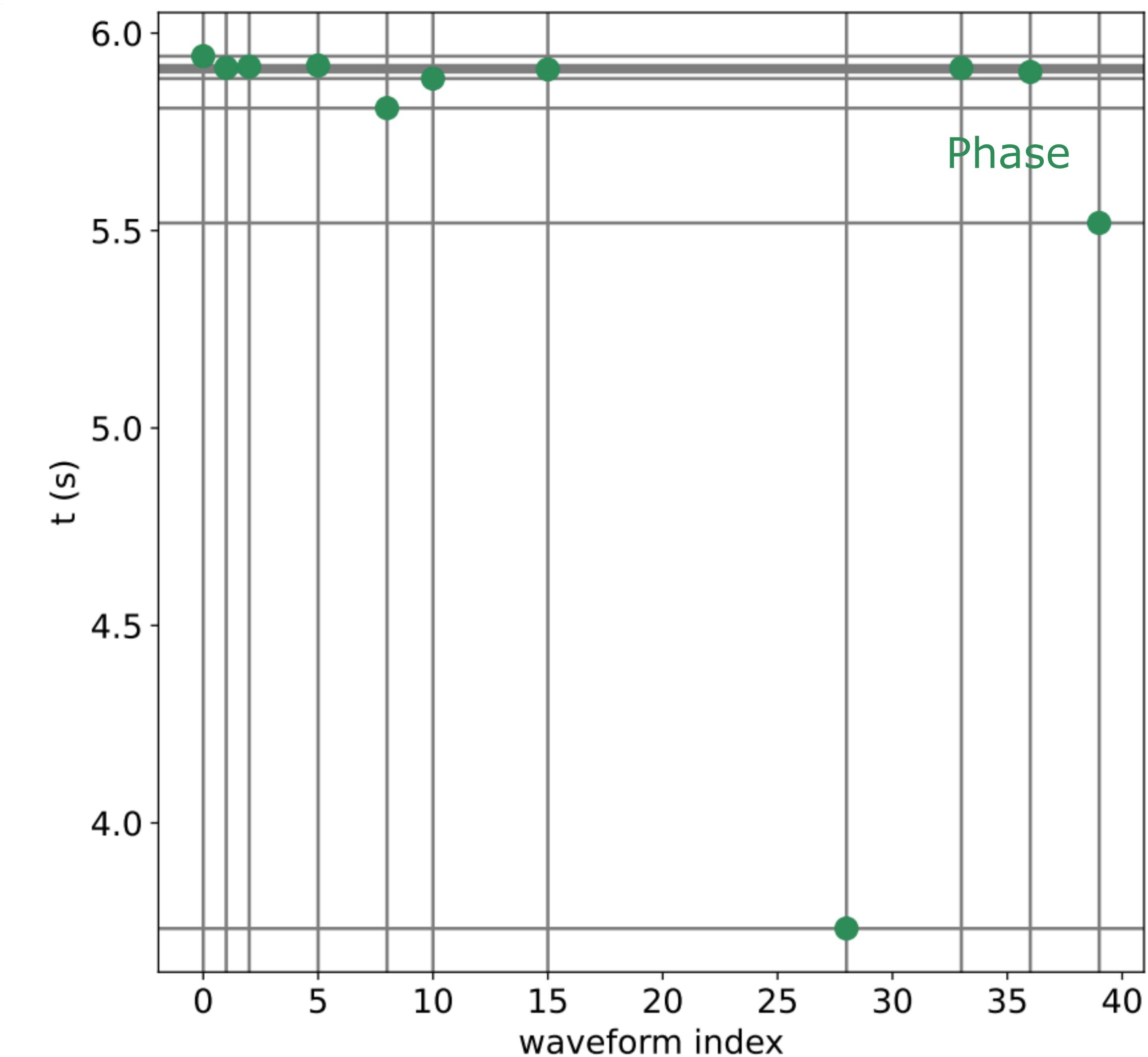
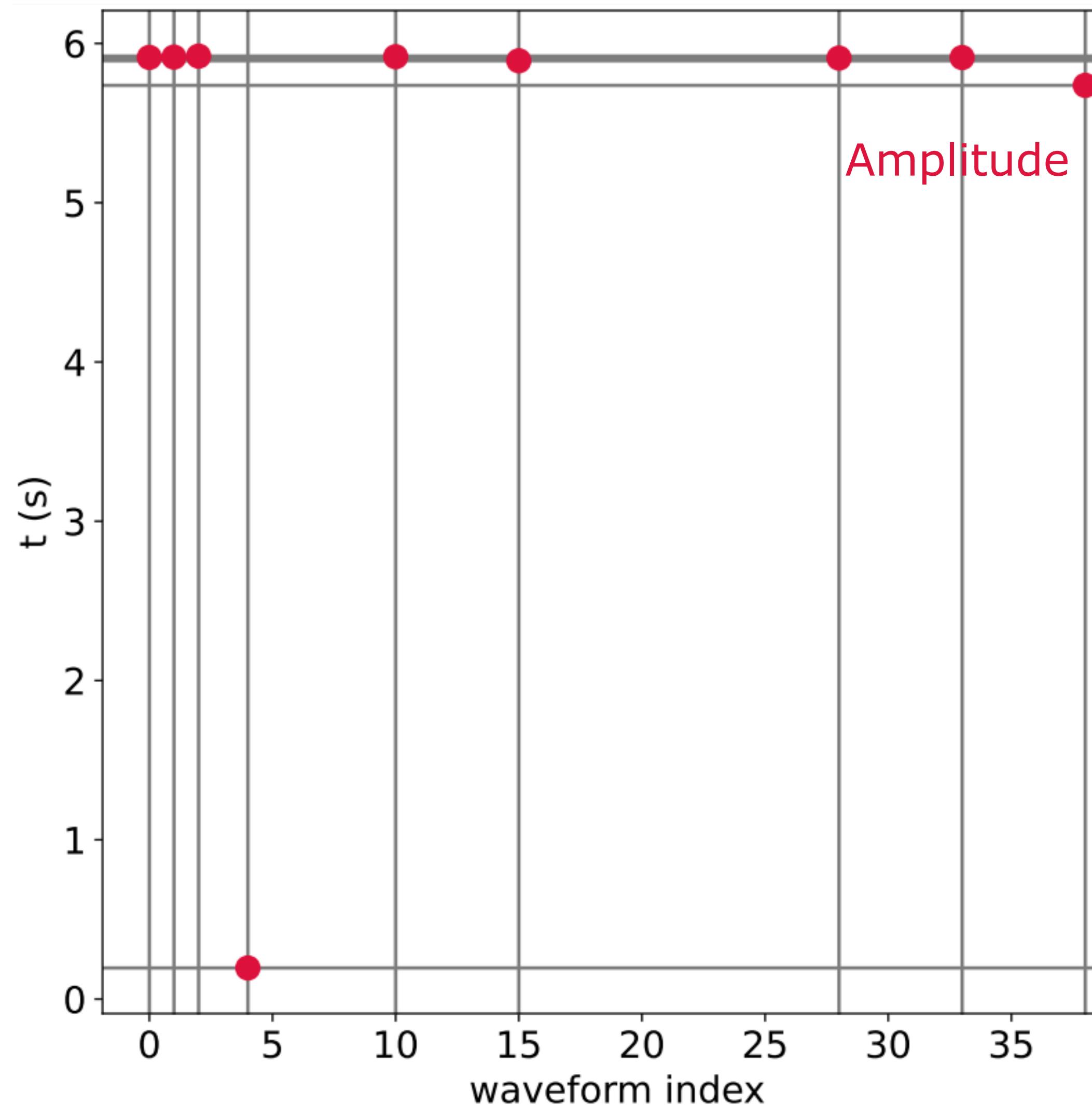
- Use $N = 2 \times 10^5$ waveforms in range $0 \leq q \leq 8, -1 \leq \chi_{1,2} \leq 1$

Greedy Tolerance	n (amplitude)	n (phase)	mismatch \mathcal{M} (max)	mismatch \mathcal{M} (median)	mismatch \mathcal{M} (95 th percentile)
10^{-6}	8	4	8.44×10^{-3}	5.47×10^{-4}	1.81×10^{-3}
10^{-8}	13	4	8.44×10^{-3}	5.45×10^{-4}	1.80×10^{-3}
10^{-10}	18	8	4.95×10^{-4}	1.30×10^{-5}	8.22×10^{-5}
10^{-12}	41	12	2.07×10^{-6}	7.45×10^{-8}	2.83×10^{-7}
10^{-14}	84	32	1.34×10^{-8}	5.64×10^{-10}	3.95×10^{-9}
10^{-16}	93	48	6.60×10^{-9}	4.59×10^{-10}	3.02×10^{-9}

baseline for
comparisons

WAVEFORMS AND EMPIRICAL TIMES OF THE EIM BASIS

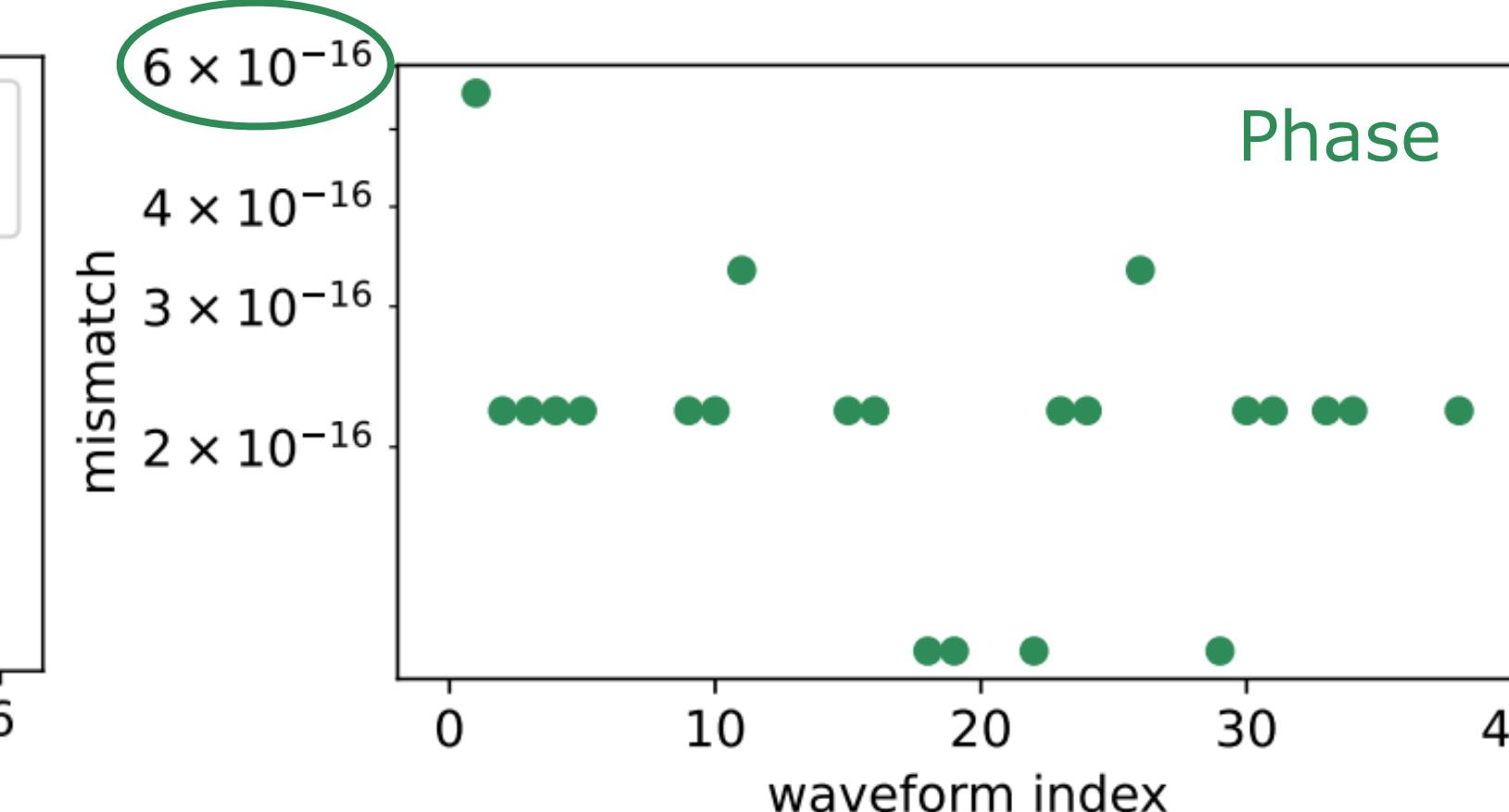
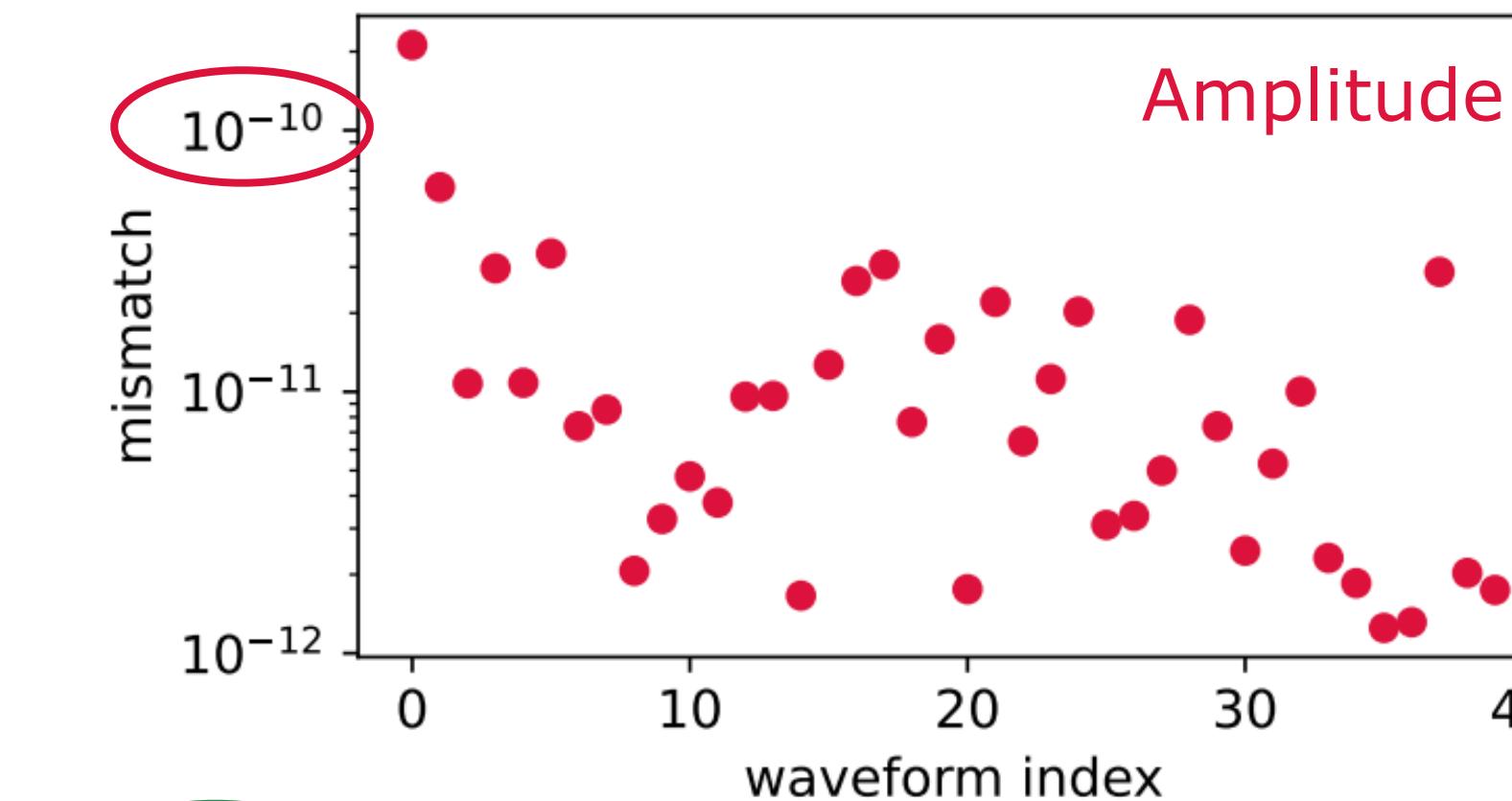
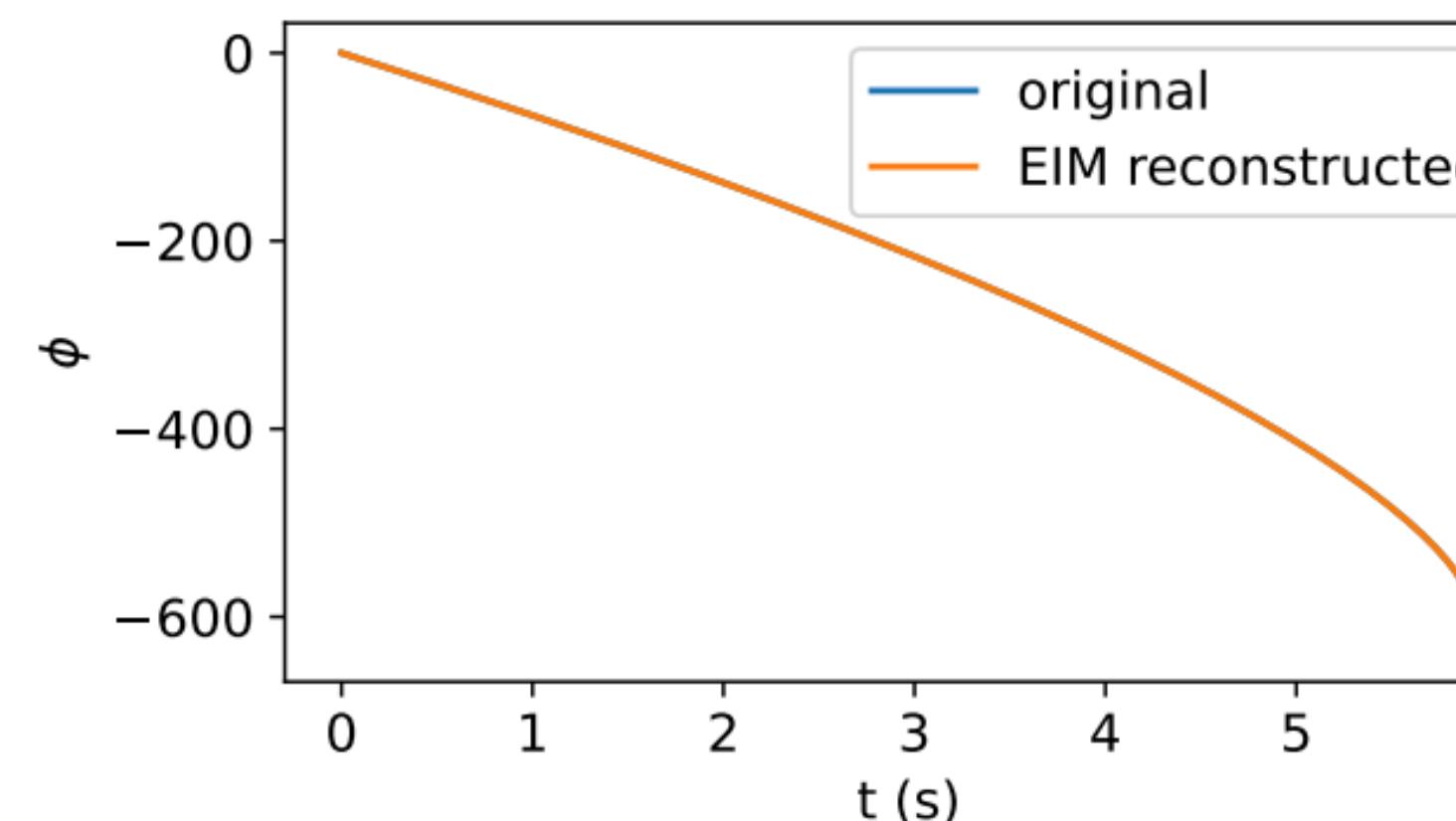
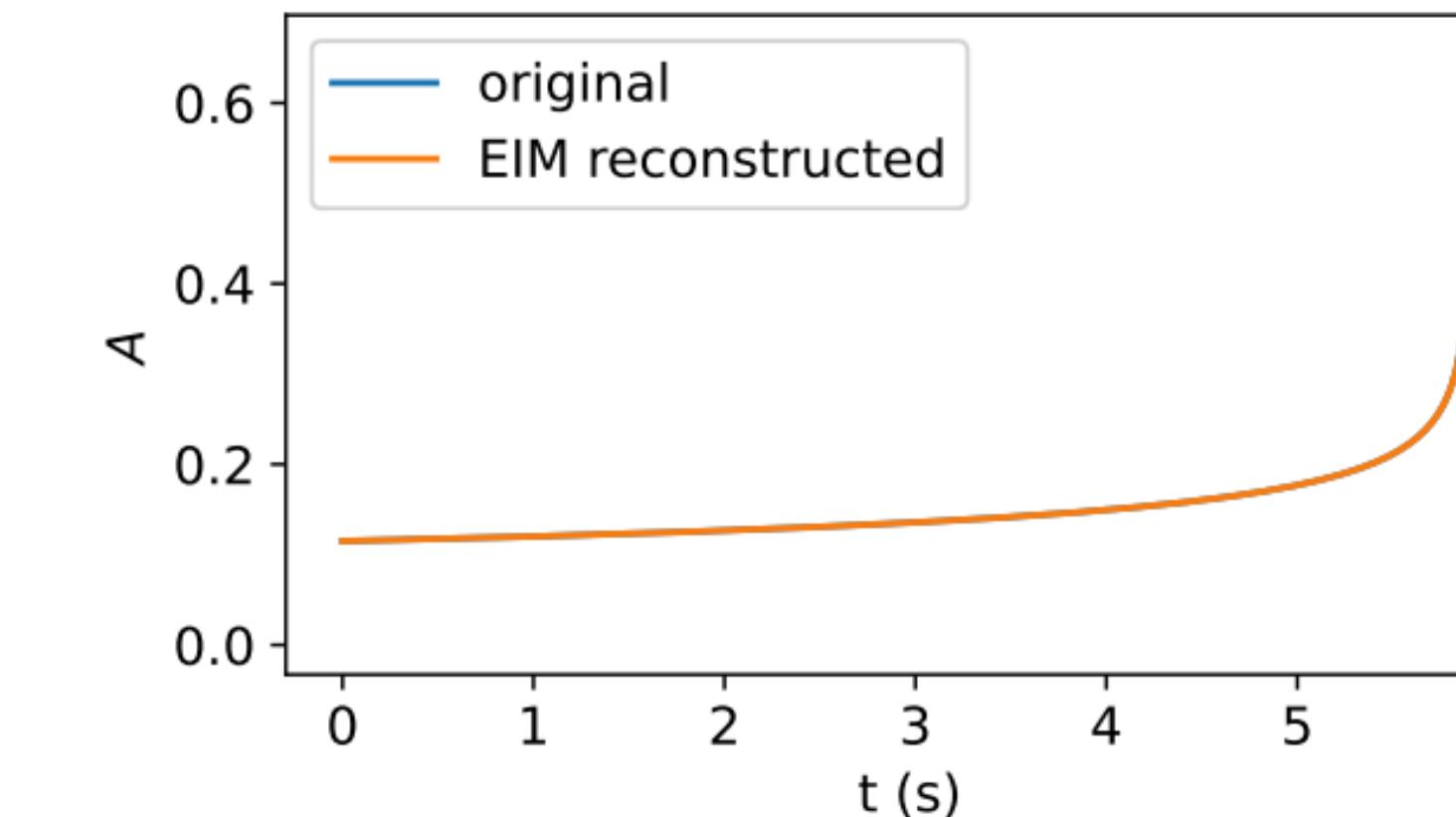
- Example: nonspinning $N = 40$ & tolerance = 10^{-12} , then $n = 8$ for amplitude and $n= 10$ for phase.



SURROGATE MODEL

- Interpolating the EIM coefficients $\alpha_k(\lambda_i)$ of the training set for arbitrary values λ of the intrinsic parameters, we obtain the interpolated EIM coefficients $\hat{\alpha}_k(\lambda)$
- A surrogate model of the waveform model is then
- Example reconstruction:

$$h(t; \lambda) \approx \sum_{k=1}^n \hat{\alpha}_k(\lambda) B_k(t)$$



Mismatch for random validation set of $N=40$ waveforms.

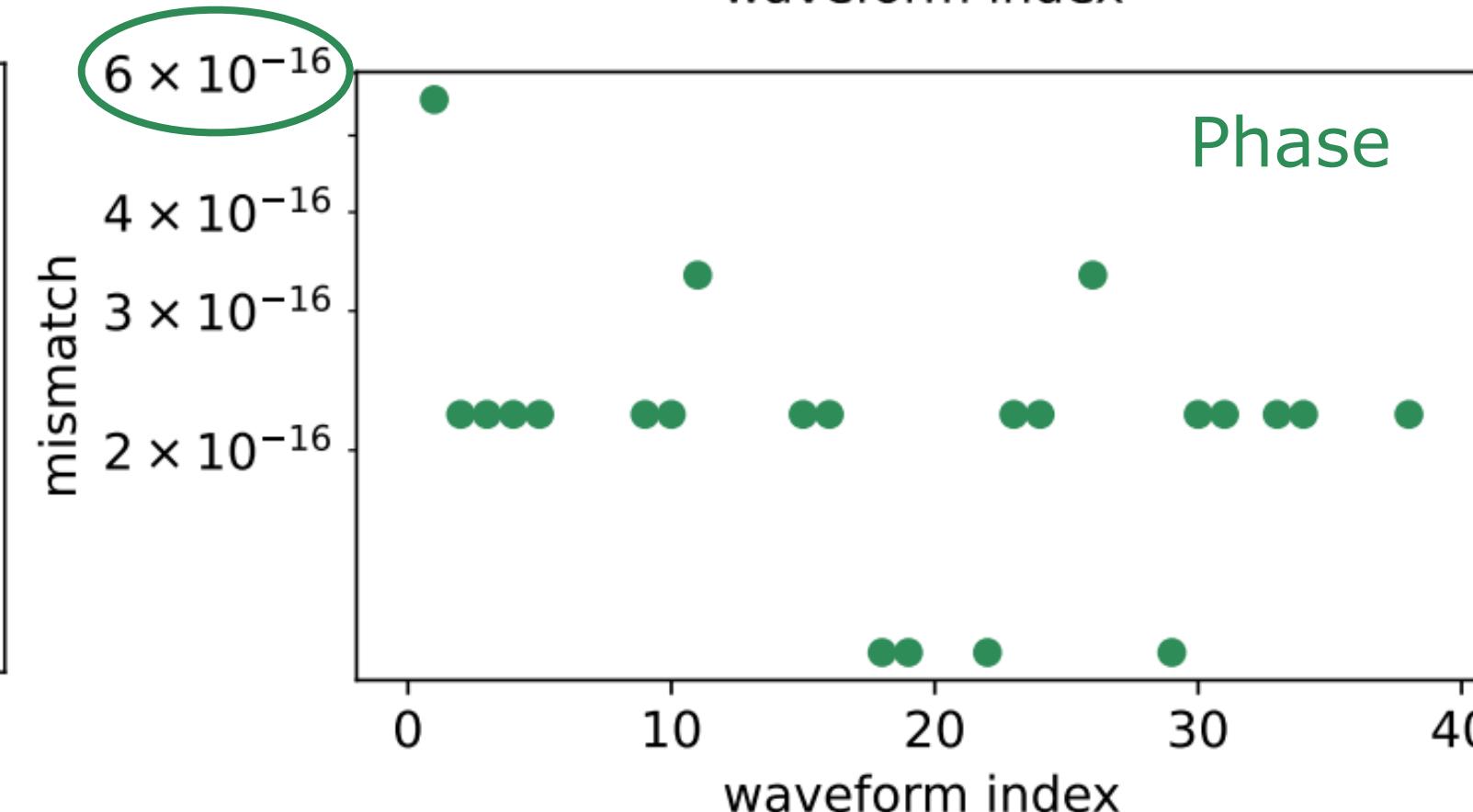
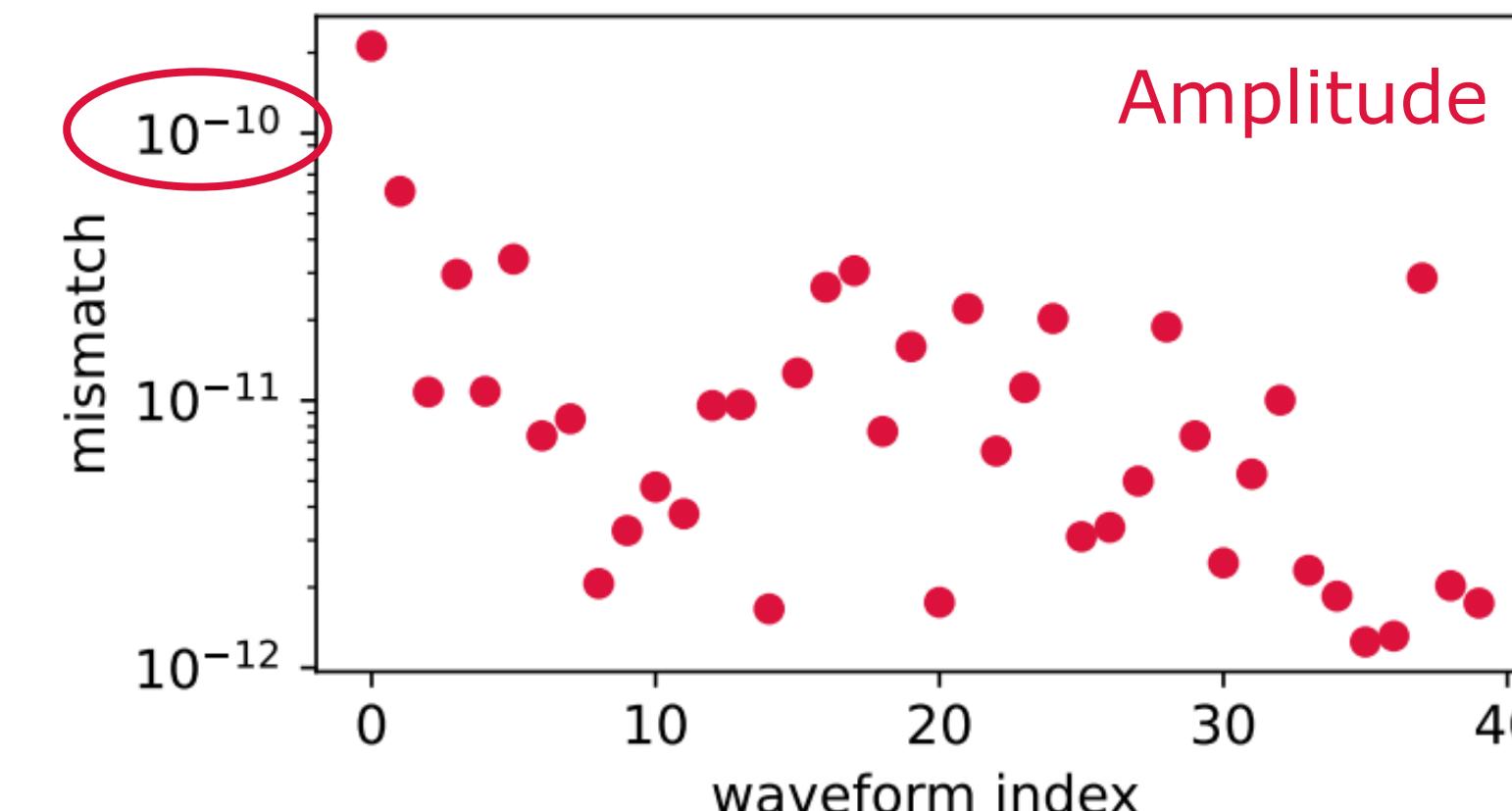
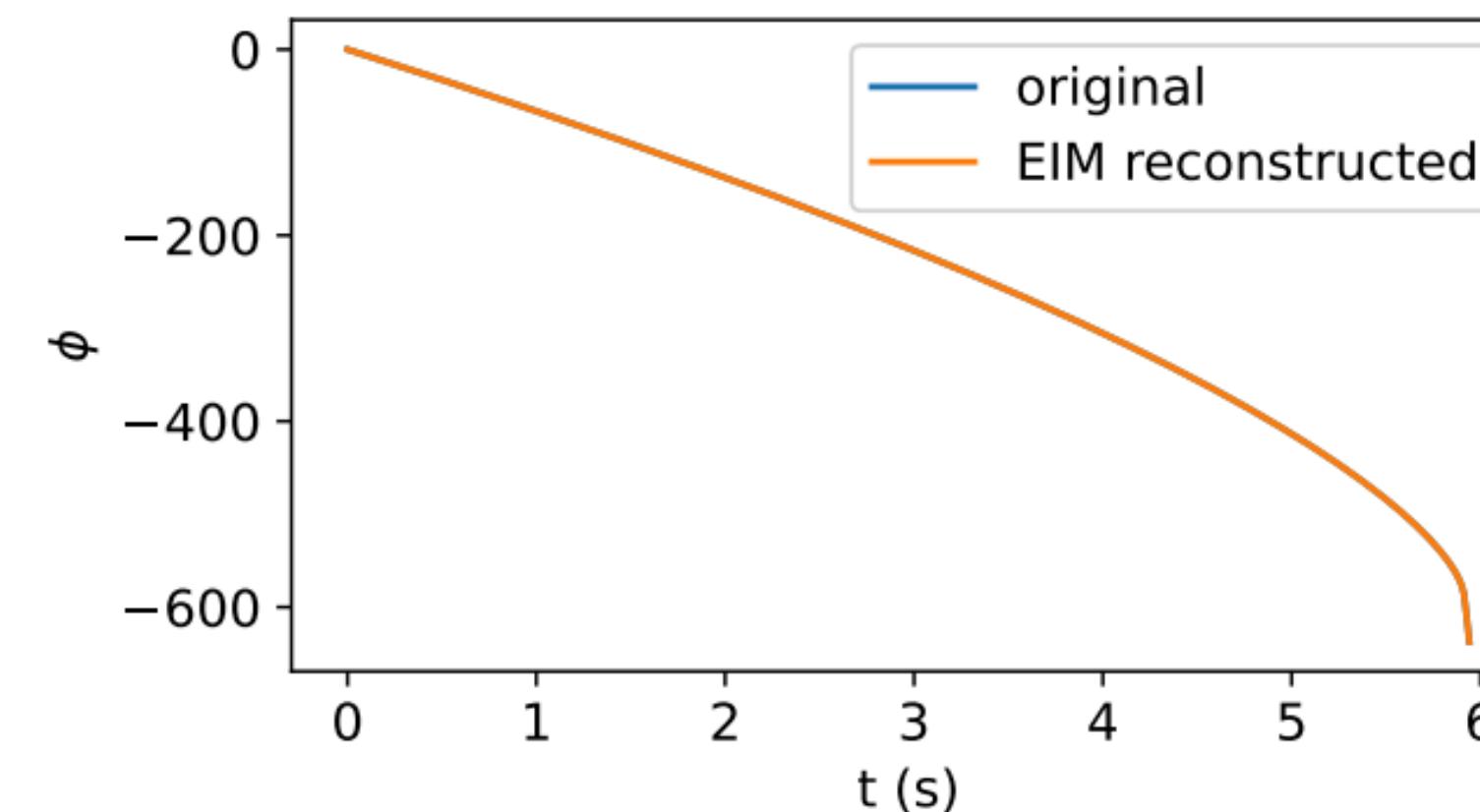
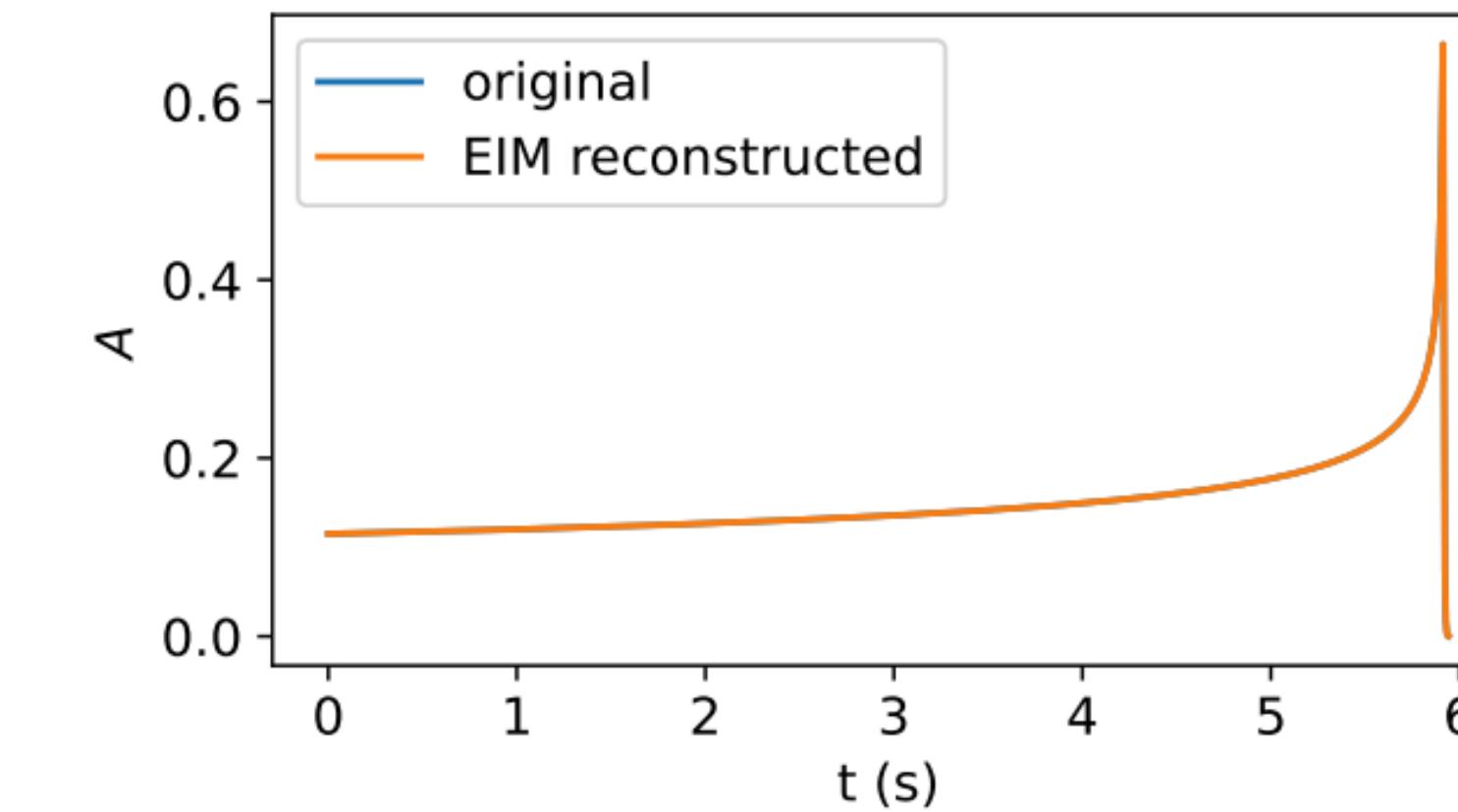
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Artificial
Neural
Network

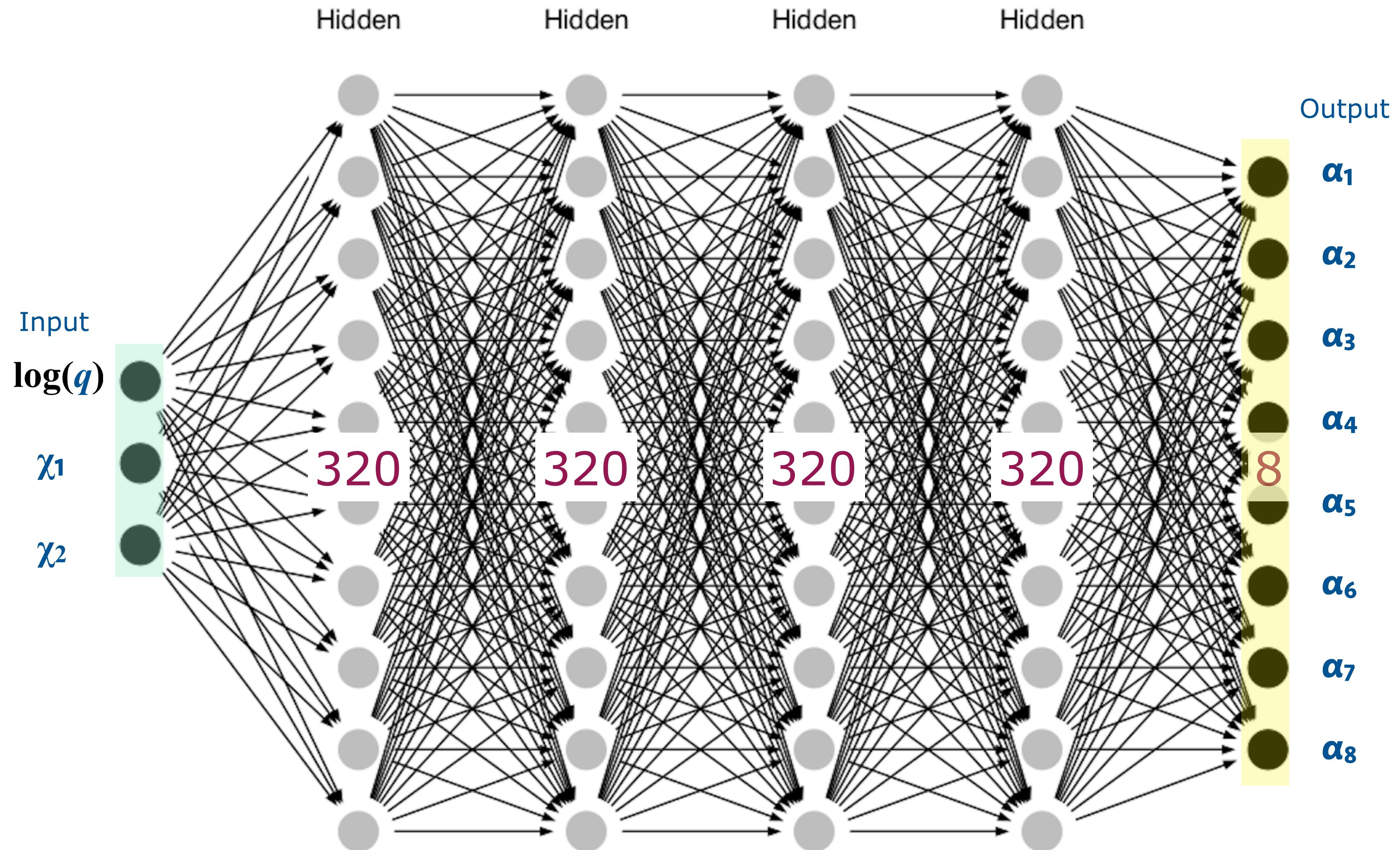
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ARTIFICIAL NEURAL NETWORK



ARTIFICIAL NEURAL NETWORK

- Mini-batch size = 1000
- Epochs = 1000
- For amplitude:

Input: standard scaler

Output: raw

Adam optimizer, initial learning rate= 10^{-3}

Activation function = RELU

- For Phase:

Input: standard scaler

Output: minmax scaler

Adamax optimizer, initial learning rate = 10^{-2}

Activation function = softplus

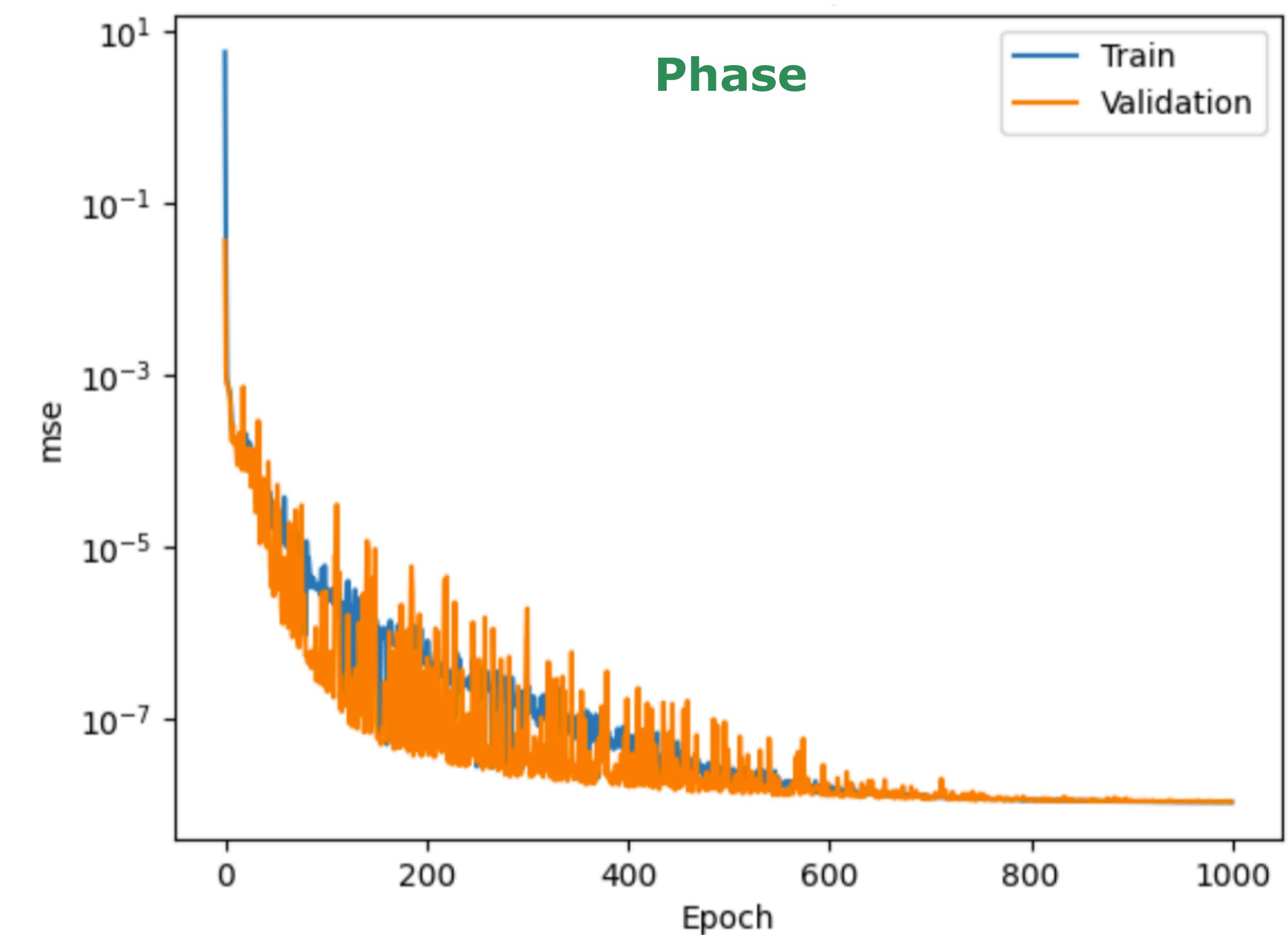
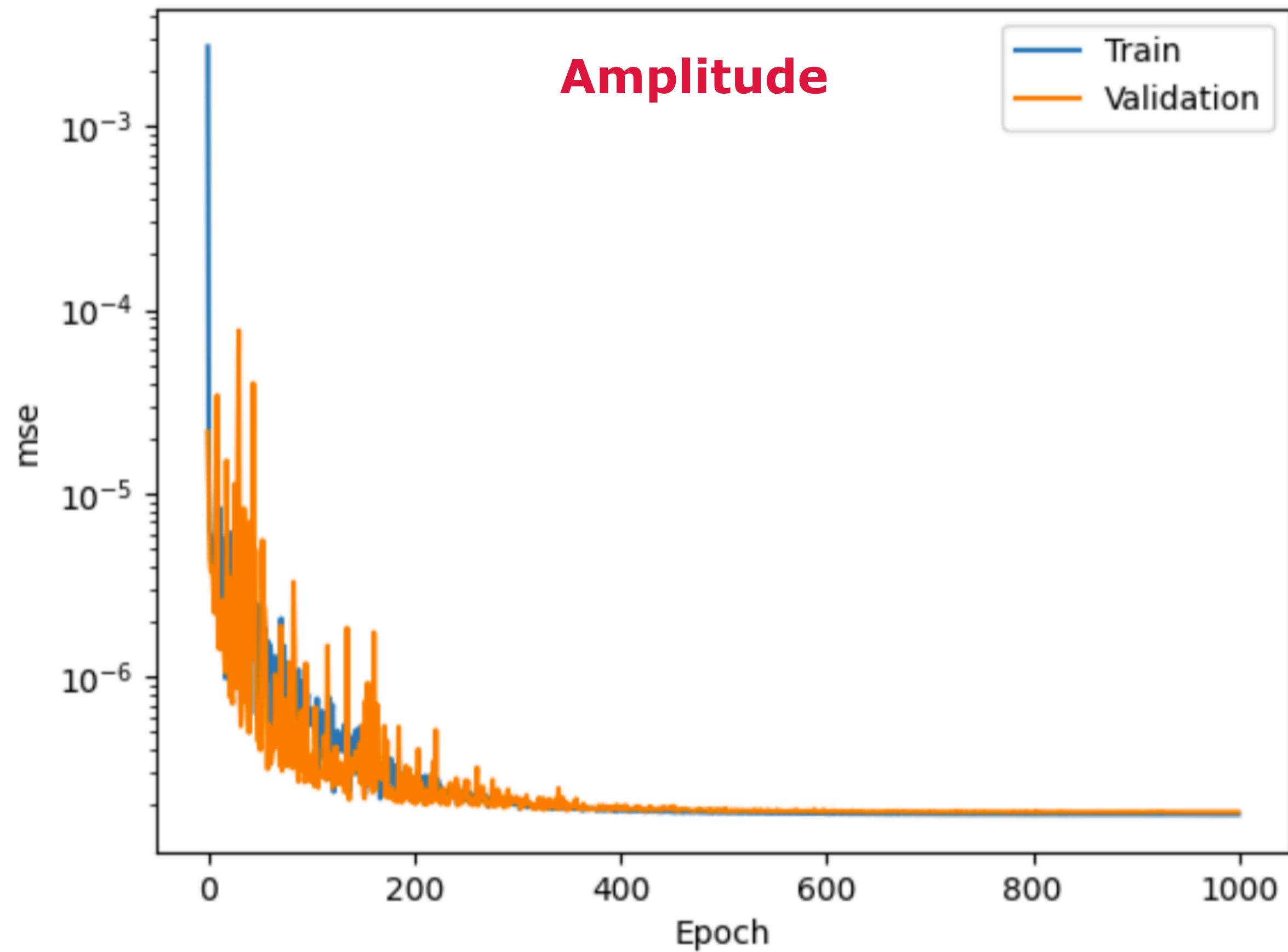
Surrogate with ANN **~5000 faster** than original SEOBNRv4 waveform generation.

EOBNRv2 WAVEFORM MODEL

- Training: minimization of Mean Square Error

$$MSE = \frac{1}{N} \sum_{i=1}^N \|\hat{\mathbf{y}}_i - \mathbf{y}_i\|_2^2$$

- Validation set: 3×10^4 waveforms used to confirm training of model



RESIDUAL NETWORK

- Define the **residual** (for each waveform) of the first network as

$$\mathbf{e}_i \equiv \mathbf{y}(\boldsymbol{\lambda}_i) - \hat{\mathbf{y}}(\boldsymbol{\lambda}_i)$$

- A **second network**, with *same architecture* as the first one, is trained on these residuals.

The output scaling was MinMaxScaler

- The MSE is

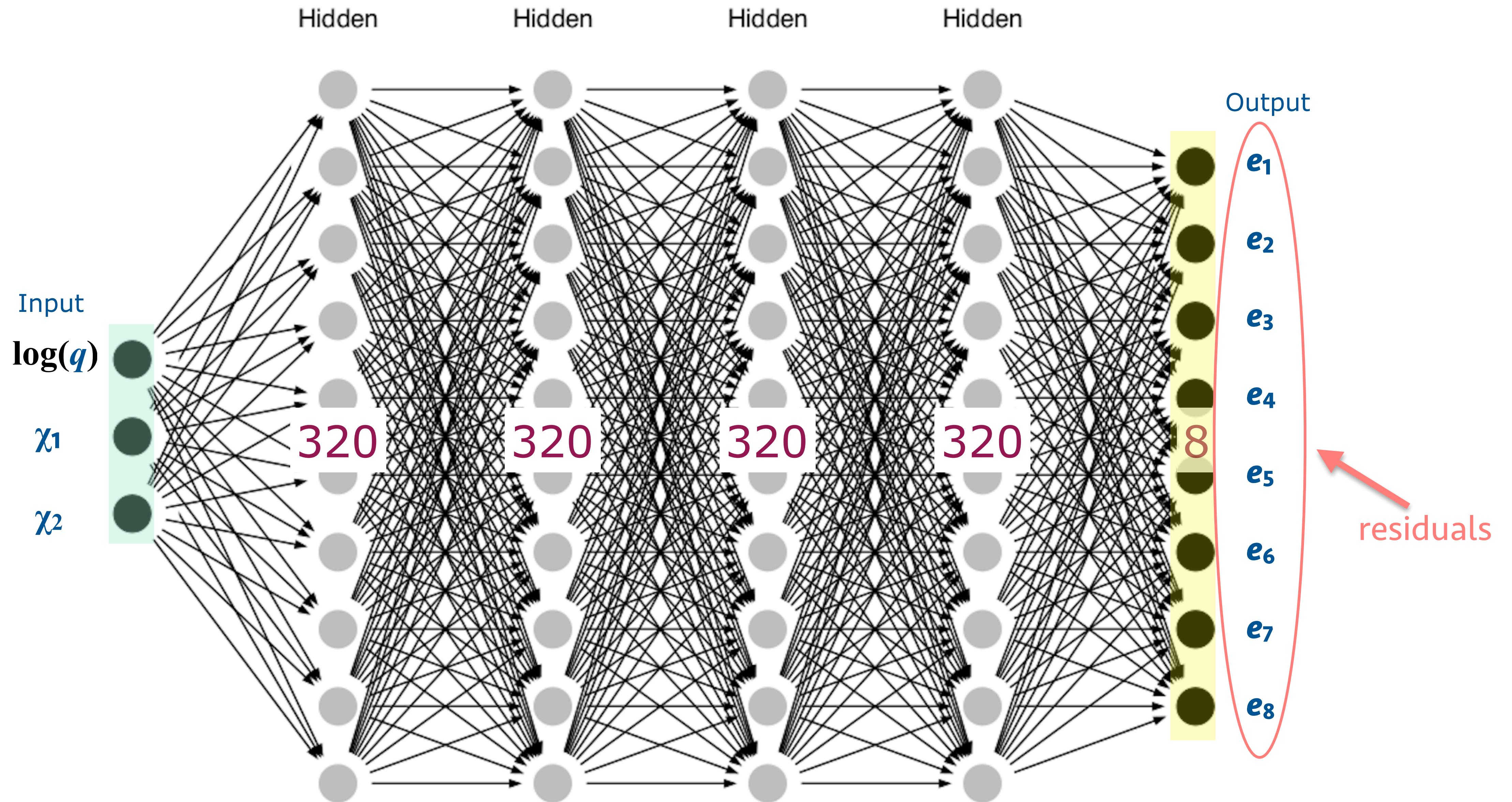
$$MSE = \frac{1}{N} \sum_{i=1}^N \|\hat{\mathbf{e}}_i - \mathbf{e}_i\|_2^2$$

- The prediction $\hat{\mathbf{e}}$ of the second network is then added to the prediction $\hat{\mathbf{y}}$ of the first network

- Improved prediction:

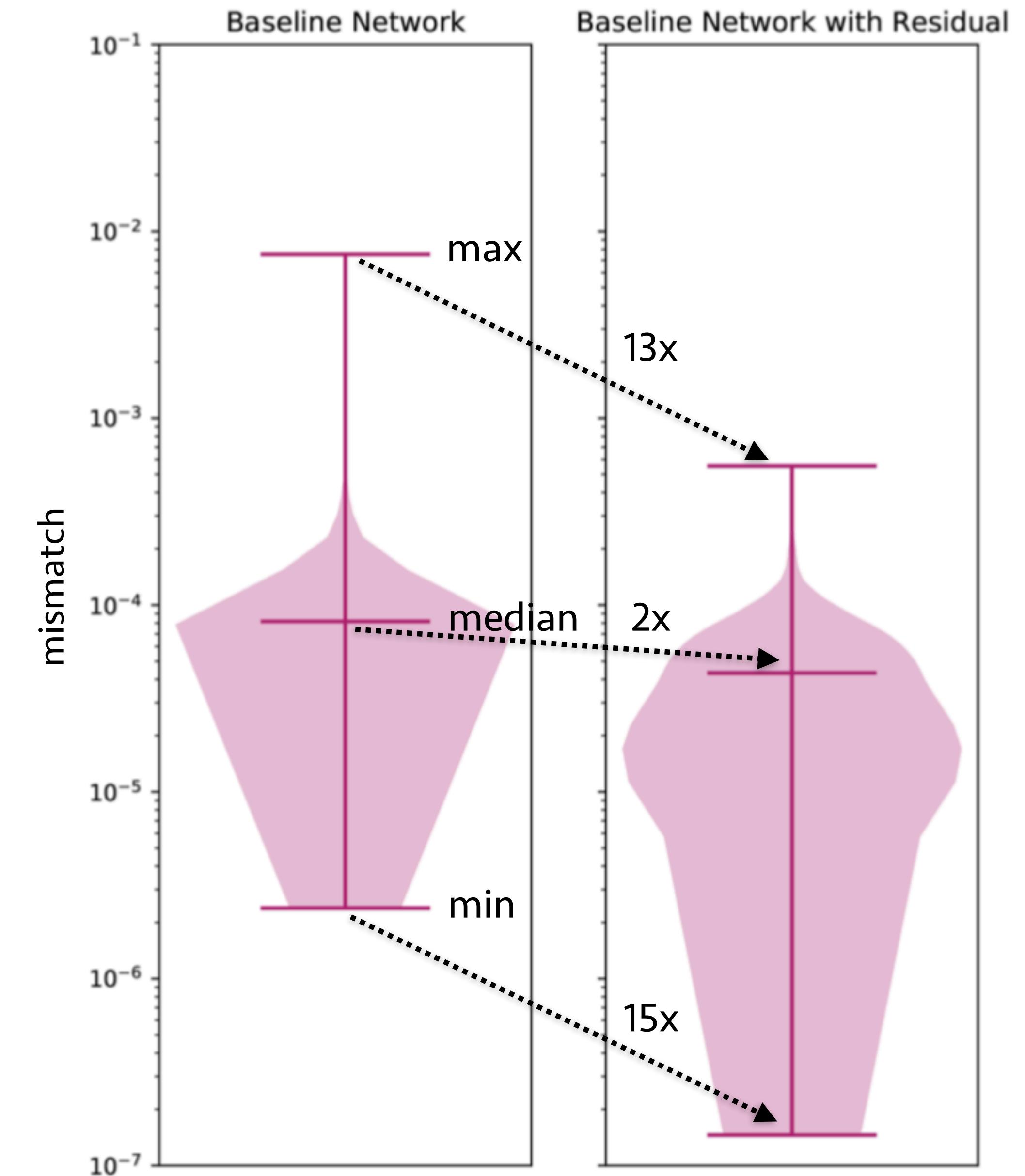
$$\tilde{\mathbf{y}} \equiv \hat{\mathbf{y}} + \hat{\mathbf{e}}$$

ARTIFICIAL NEURAL NETWORK



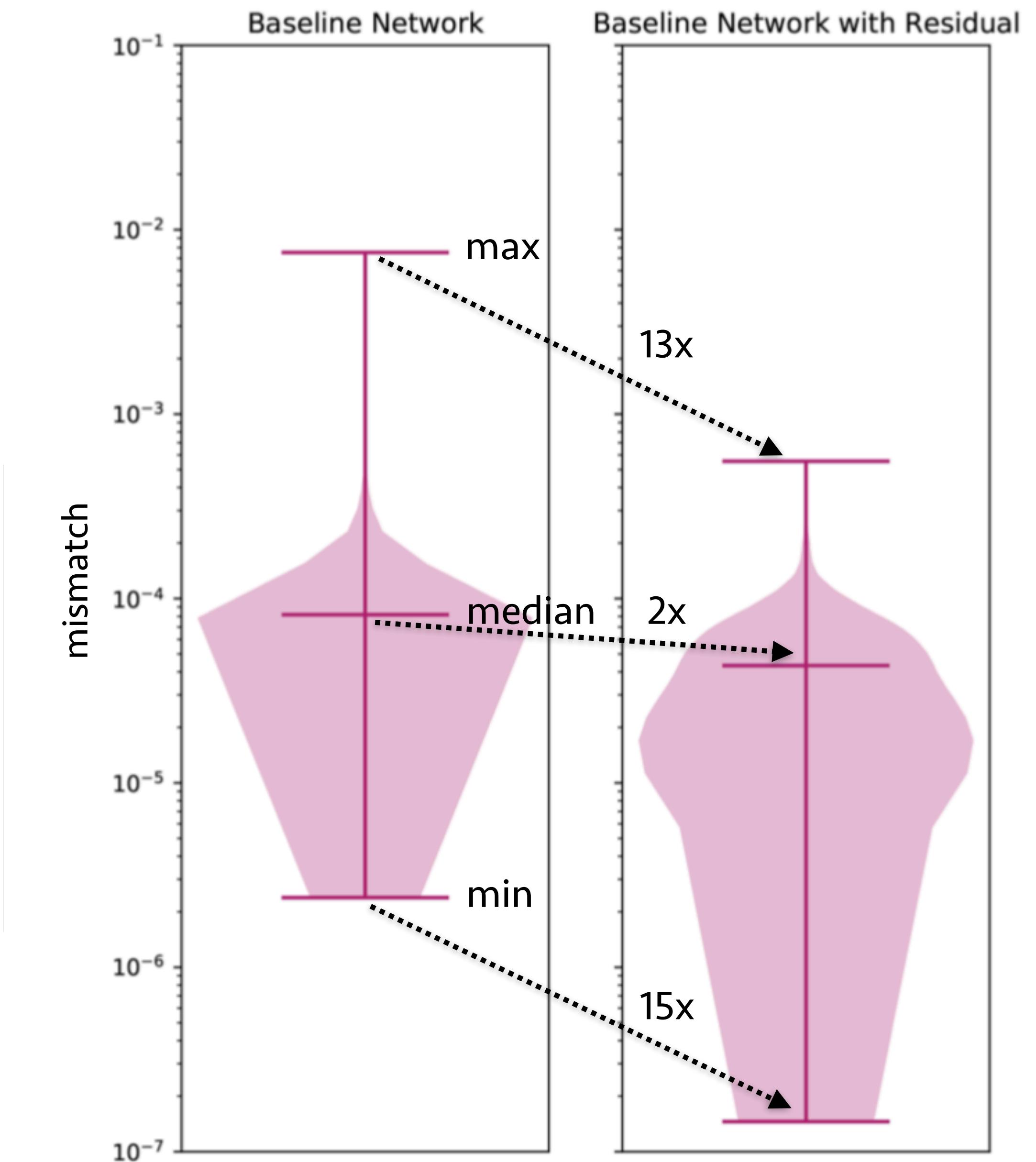
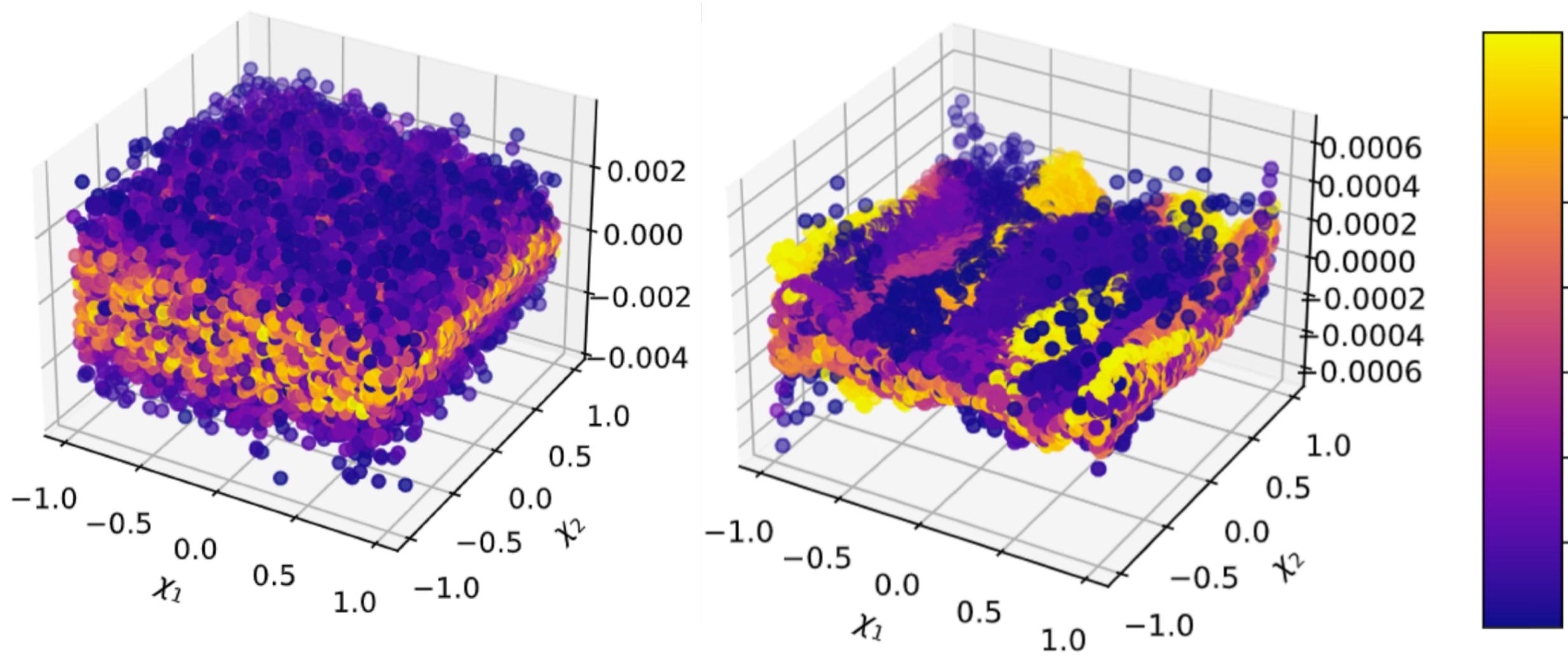
IMPACT OF RESIDUAL NETWORK

- Adding the second network trained on the residuals of the first, improves the maximum, median and minimum mismatches by factors of 13, 2 and 15, correspondingly.



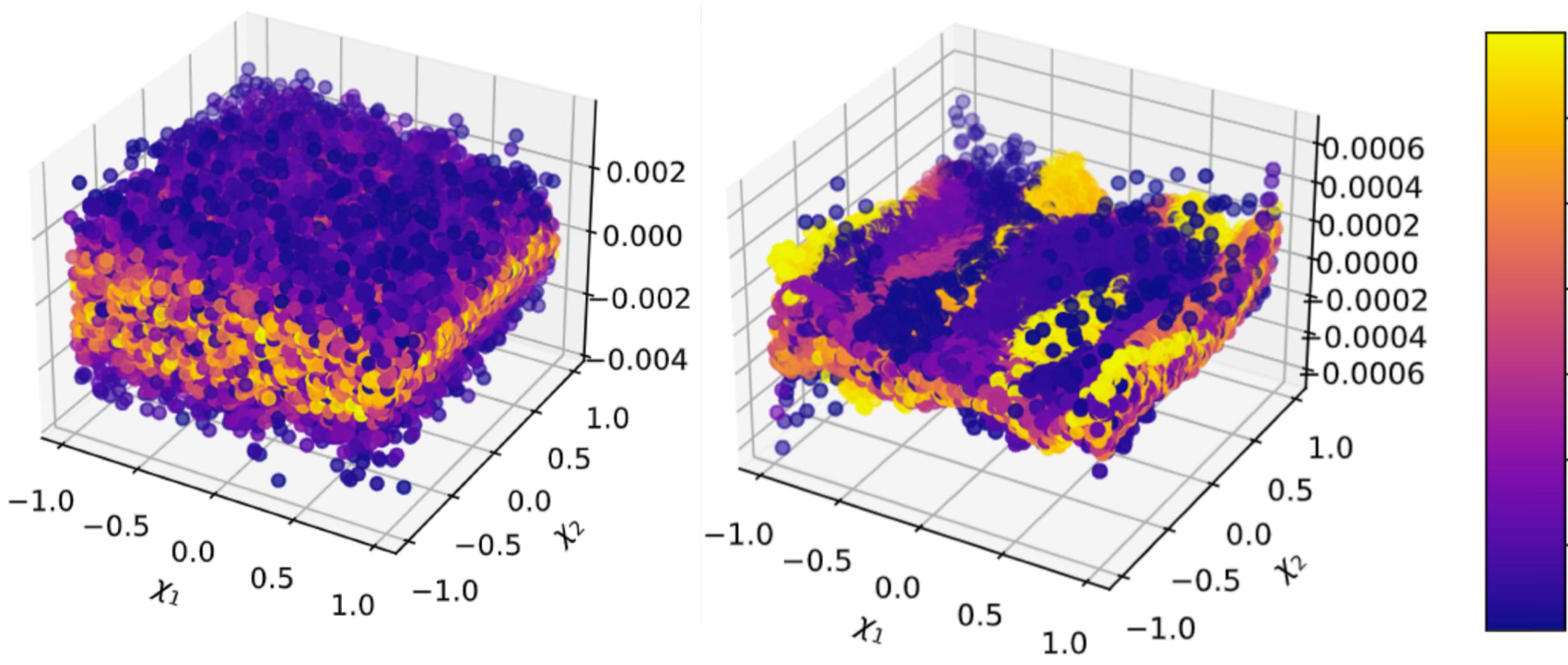
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- Explanation: residual errors of some coefficients $\hat{\alpha}_k$ obtained by the first network have **structure**. Examples:

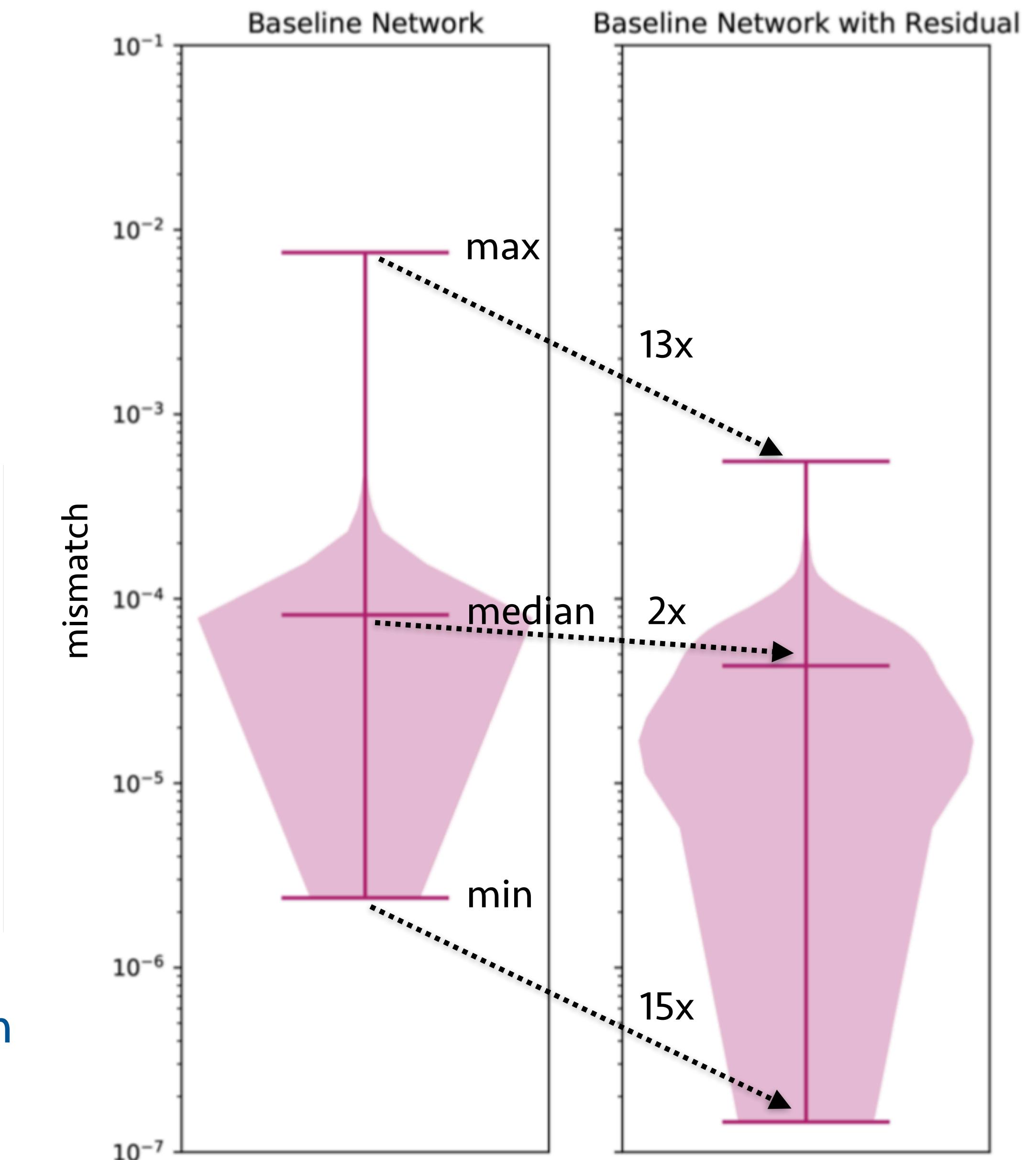


IMPROVEMENTS THROUGH RESIDUAL NETWORK

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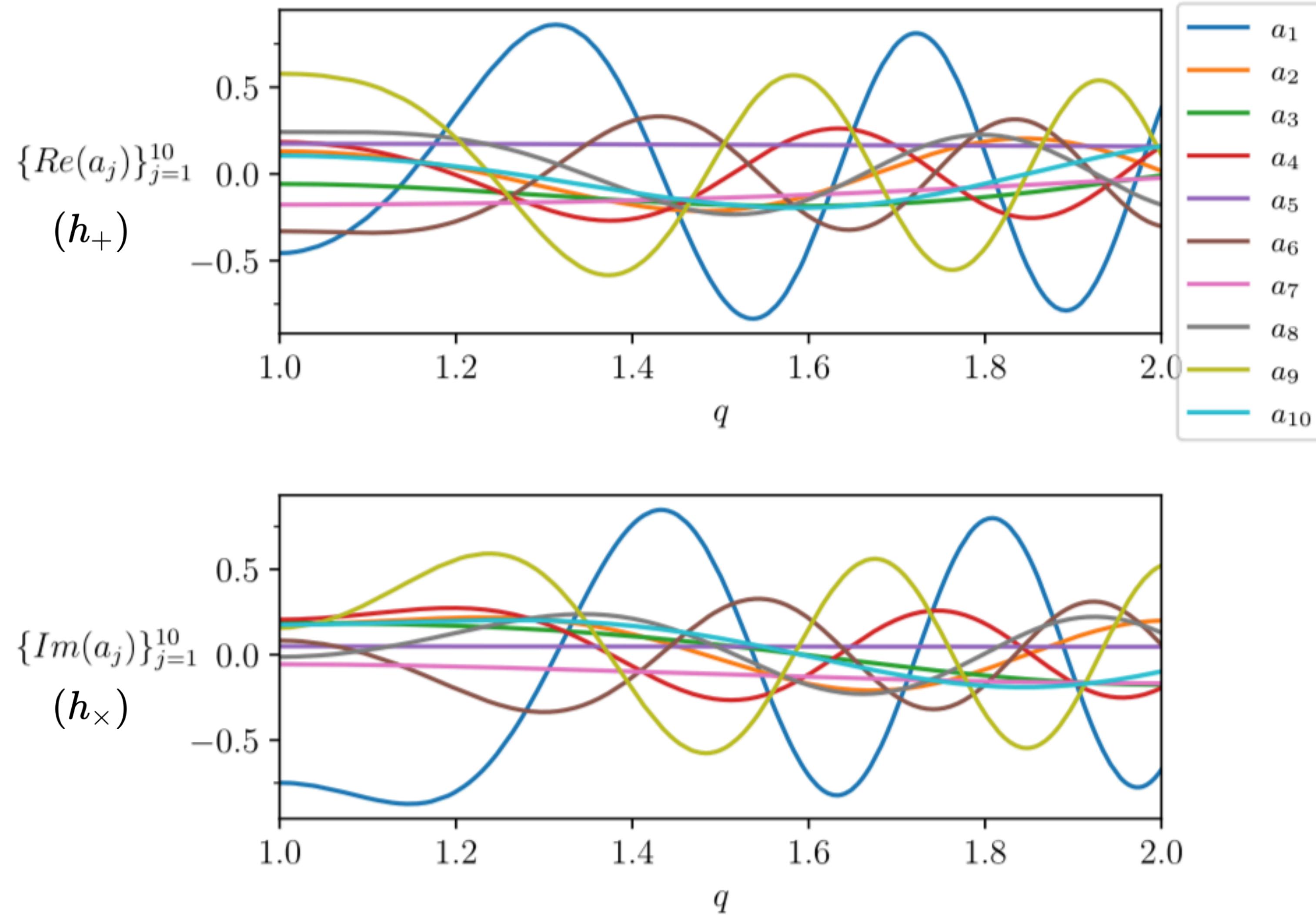


- Outlook: generalize to precessing BBH and BNS waveform models.



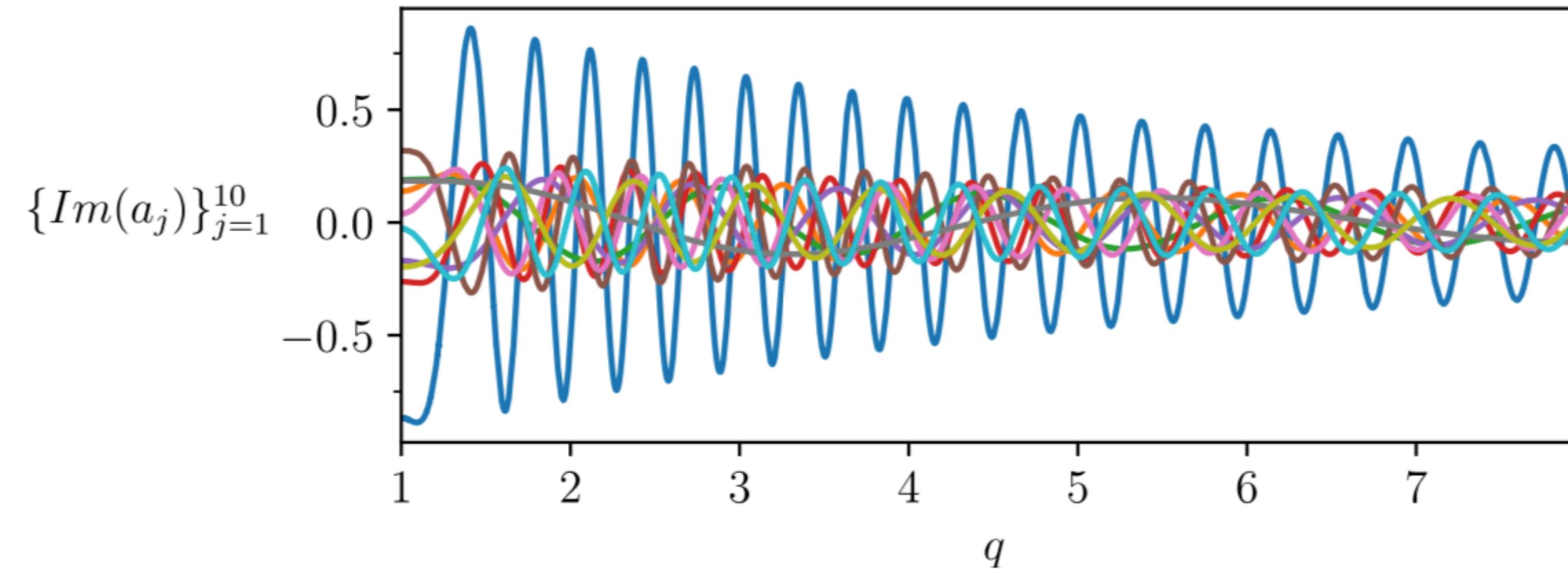
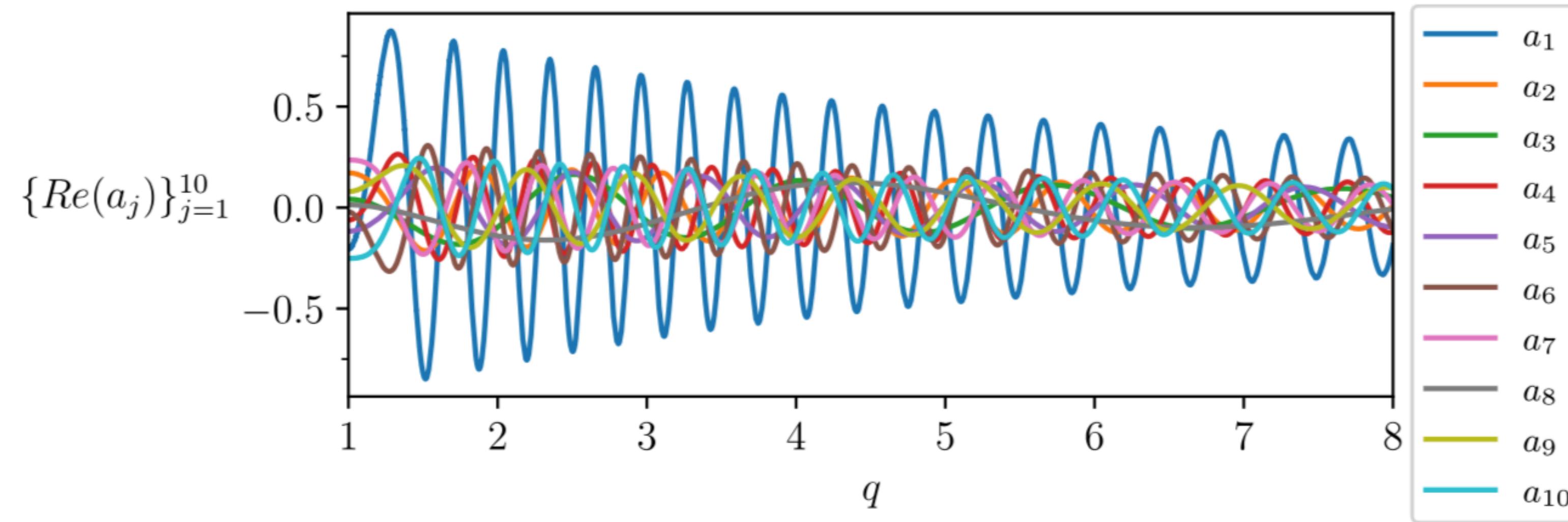
AUTOENCODERS

- Nonspinning case
 - Construct a surrogate model for the two polarizations h_+, h_\times of nonspinning EOBNRv2 waveforms.
- $$h(t; q) = h_+(t; q) - i h_\times(t; q)$$
- Initial range of mass ratio $0 \leq q \leq 2$
 - $N = 1000$ waveforms
 - With tolerance $= 10^{-10}$ the reduced basis has a size of $n = 11$.
 - The first ten coefficients a_j are shown as a function of q .



EXTENSION TO LARGER MASS RATIOS

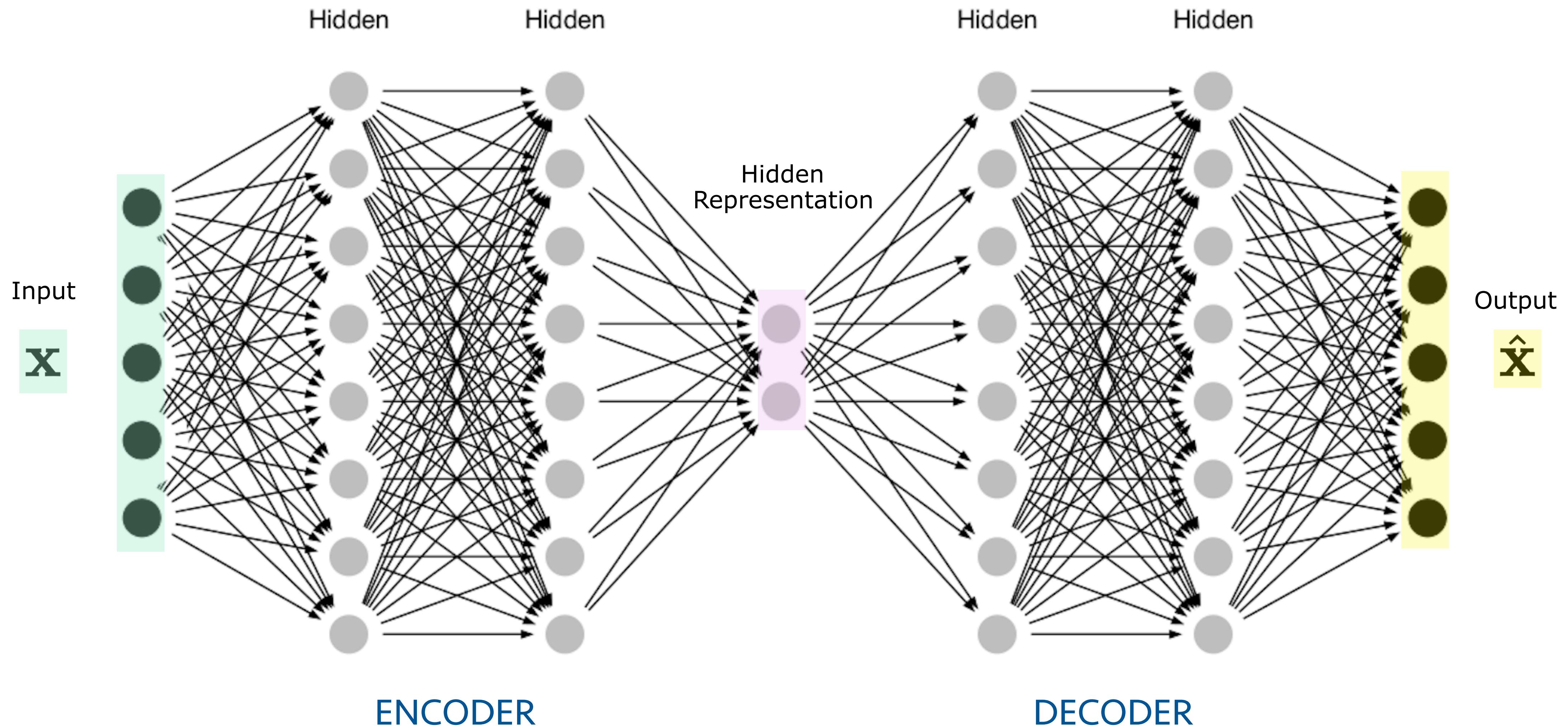
Reconstructed EIM coefficients for $0 \leq q \leq 8$



AUTOENCODER

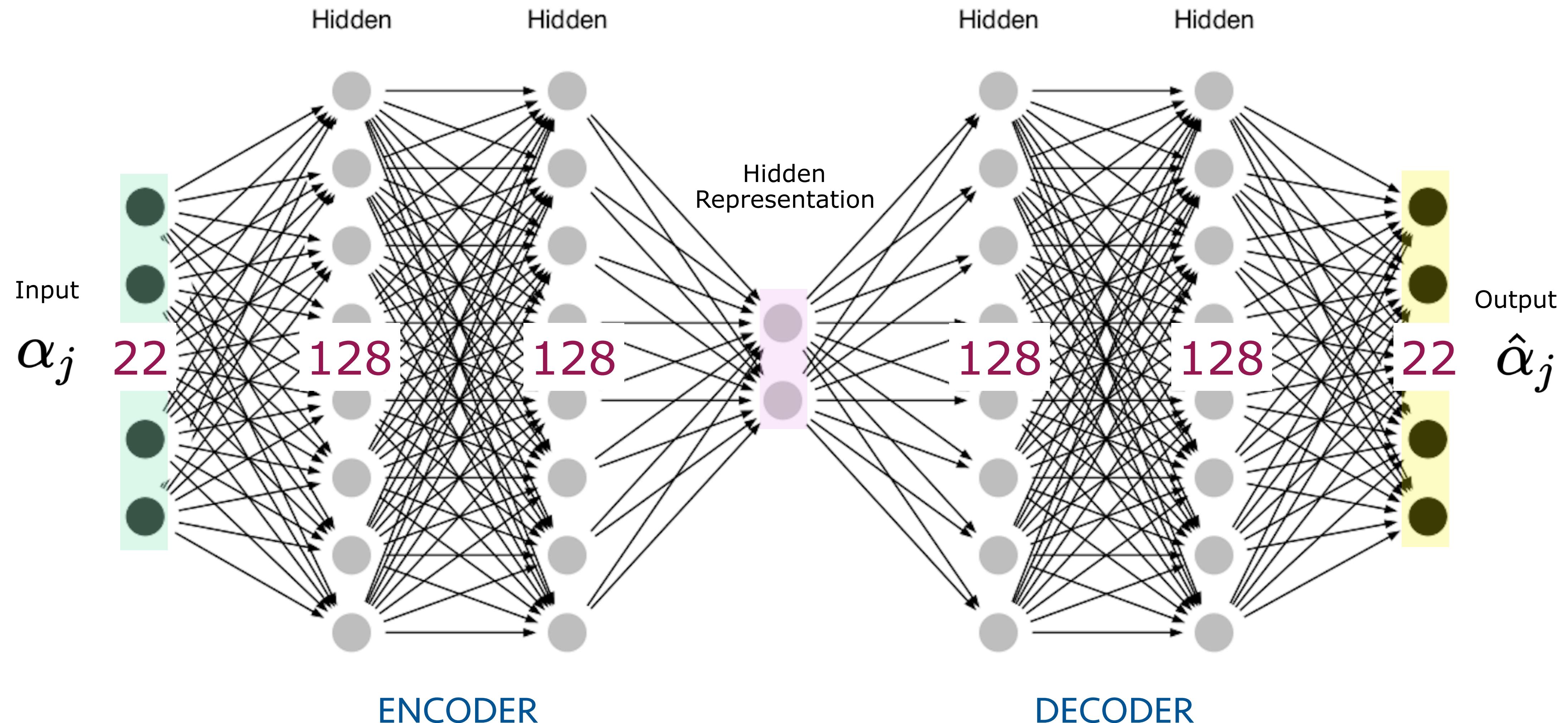
Unsupervised - encodes the input \mathbf{x} into a lower-dimensional representation $g(\mathbf{x})$ and then decodes it at the output as $\hat{\mathbf{x}} = f(g(\mathbf{x}))$

$$\text{MSE} = \frac{1}{N} \sum_{i=1}^N \|\hat{\mathbf{x}}_i - \mathbf{x}_i\|_2^2$$



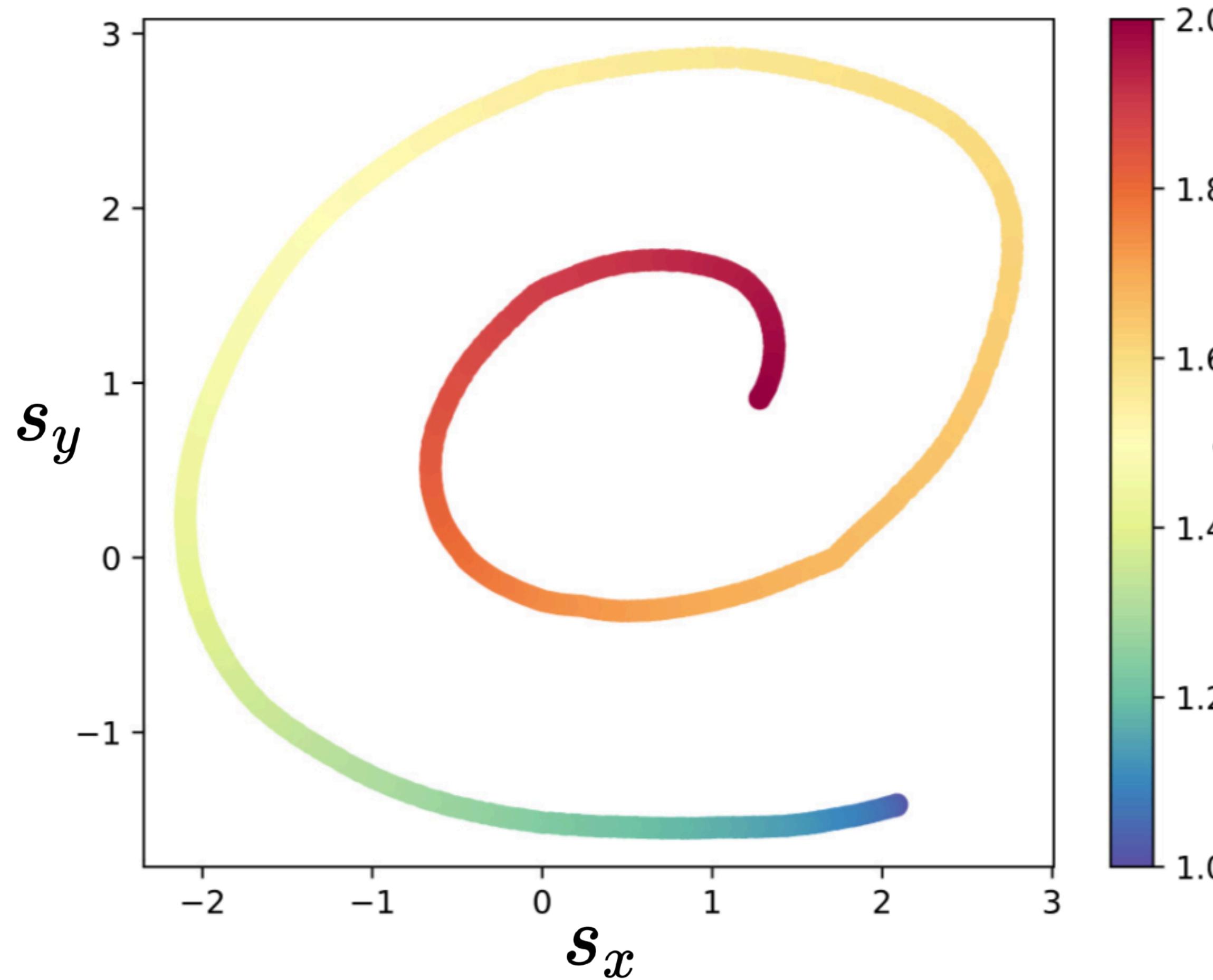
AUTOENCODERS

Training: 100 epochs. Batch size: 32. Initial learning rate = 1e-3. Learning schedule x0.9 every 15 epochs.
Activation function PreLU (parametrized RELU), Gradient descent. Achieved MSE = 6.8e-5.



SPIRAL STRUCTURE OF HIDDEN REPRESENTATION

Values s_x, s_y of hidden representation form a **spiral structure!**



We use the knowledge of q from the training set, to write down a simple relation:

$$s_x = (\alpha + \beta \cdot \theta) \cdot \cos \theta$$

$$s_y = (\alpha + \beta \cdot \theta) \cdot \sin \theta$$

where

$$\theta = w \cdot q + b$$

depends on the mass ratio q .

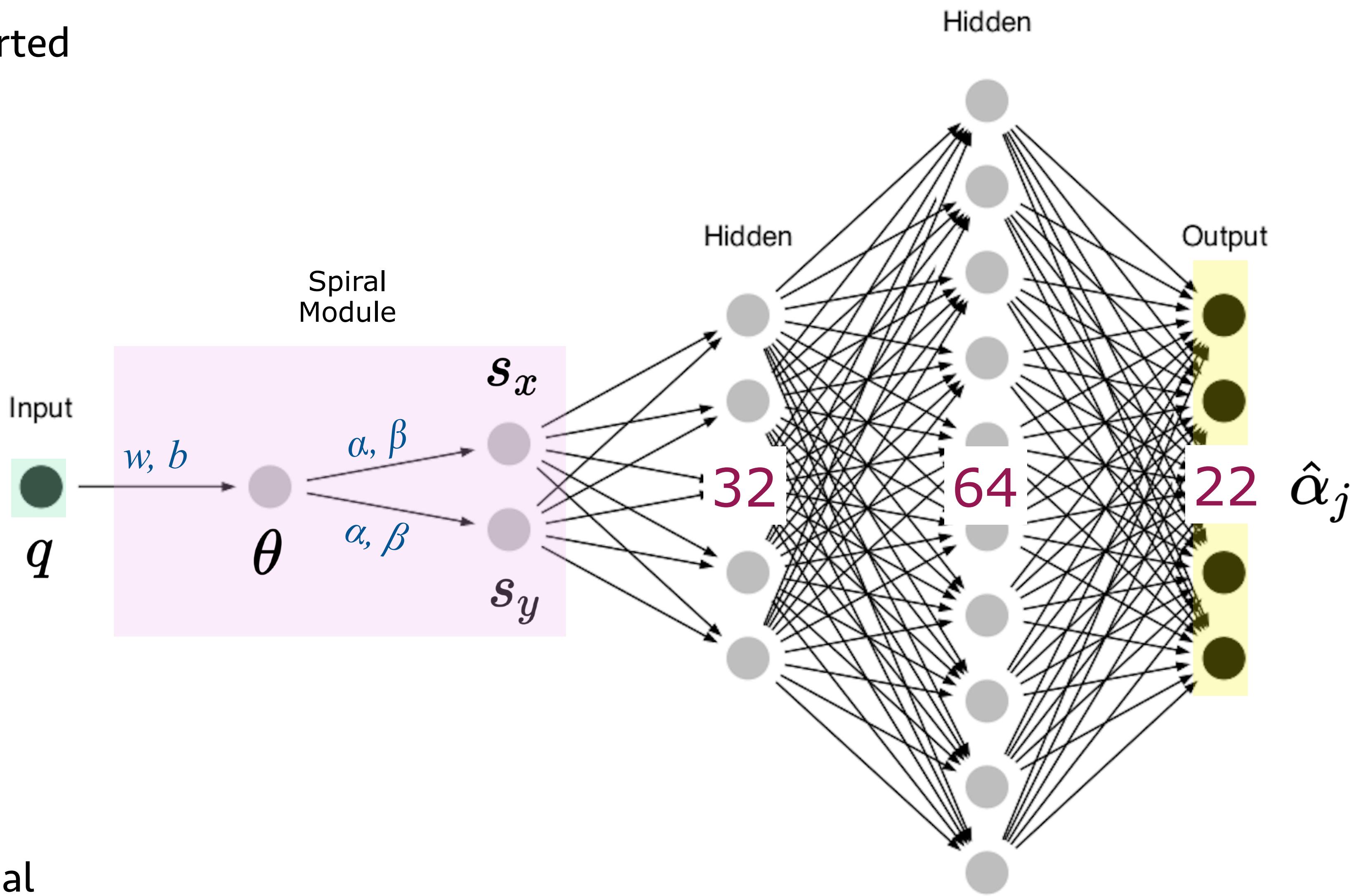
The **coefficients a, β, w, b** of the spiral are **learnable**.

We thus **insert a spiral module** that maps

$$q \rightarrow \theta \rightarrow s_x, s_y$$

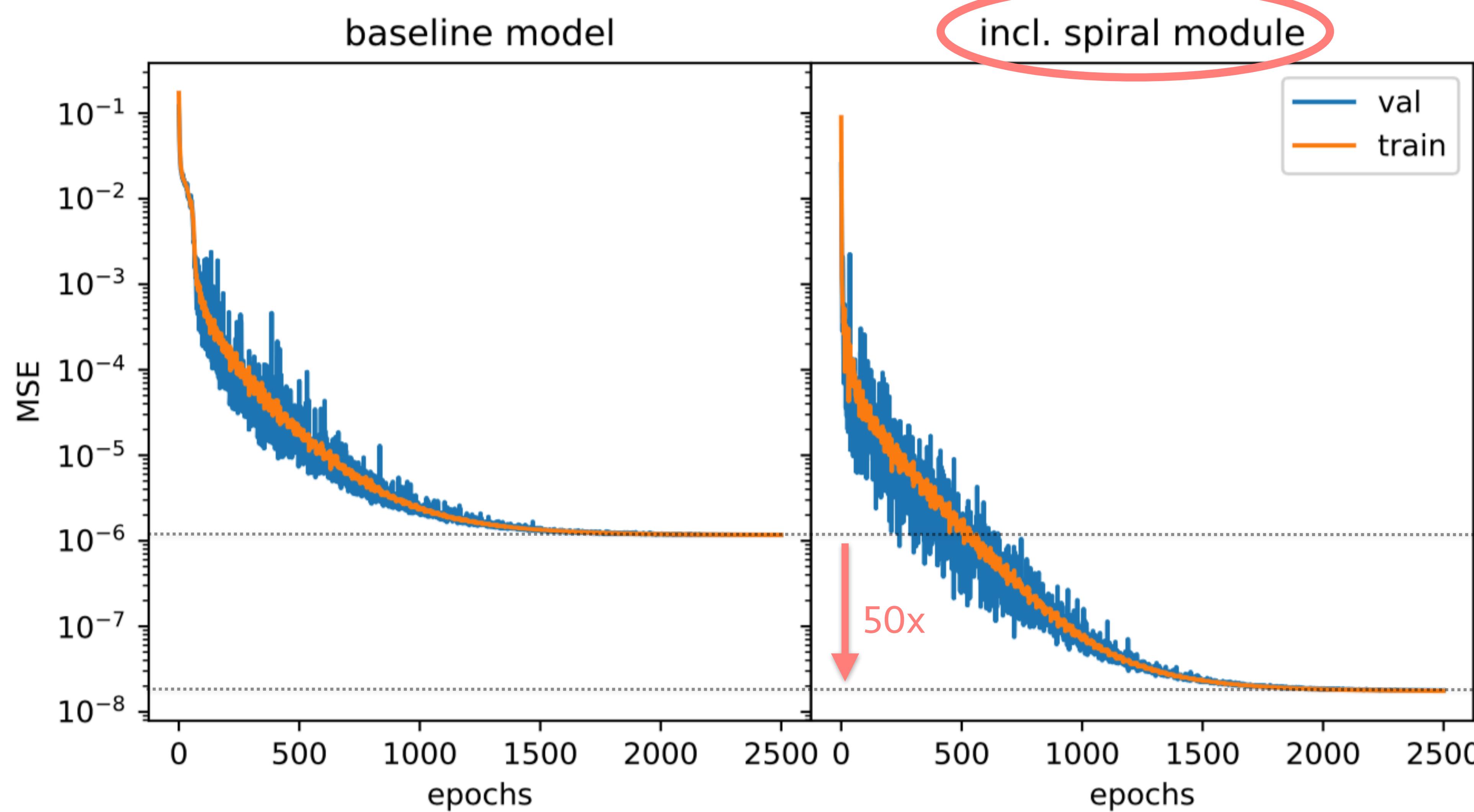
NEURAL NETWORK WTH SPIRAL MODULE

- Fully connected ANN with inserted spiral module.
- Training set: $N = 10,000$
Validation set: $N = 2,000$
- Training for 2500 epochs
- Batch size = 16
- Adam optimizer
- Initial learning rate = 10^{-3}
- Learning schedule:
x0.9 every 150 epochs
- 400 times faster than traditional spline interpolation



LOSS PER EPOCH

Training and validation loss per epoch for a 32-64-128-64 network:



COMPARISON OF ERRORS WITH AND WITHOUT SPIRAL MODULE

Same MSE is achieved with simpler networks \rightarrow faster execution time

network	max \mathcal{M}	median \mathcal{M}	$p = 95\% \mathcal{M}$	network simplicity (max batch size)
16-64	4.36×10^{-1}	1.44×10^{-1}	3.47×10^{-1}	8 m
\mathcal{S} -16-64	1.33×10^{-3}	9.14×10^{-5}	2.25×10^{-4}	8 m
32-64	4.43×10^{-1}	5.85×10^{-3}	2.84×10^{-1}	7.3 m
\mathcal{S} -32-64	4.39×10^{-3}	1.12×10^{-5}	2.56×10^{-5}	7.3 m
32-32-64	1.48×10^{-3}	4.97×10^{-5}	1.87×10^{-3}	6.1 m
\mathcal{S} -32-32-64	2.99×10^{-5}	9.04×10^{-7}	1.99×10^{-6}	6.1 m
32-64-128	9.34×10^{-5}	8.00×10^{-6}	5.83×10^{-5}	4.2 m
\mathcal{S} -32-64-128	1.80×10^{-6}	1.66×10^{-7}	3.63×10^{-7}	4.2 m
32-64-128-64	7.79×10^{-5}	1.13×10^{-6}	4.62×10^{-6}	3.4 m
\mathcal{S} -32-64-128-64	2.55×10^{-6}	4.16×10^{-8}	1.07×10^{-7}	3.4 m
64-128-256-128	2.46×10^{-5}	1.50×10^{-7}	2.23×10^{-6}	1.9 m
\mathcal{S} -64-128-256-128	5.25×10^{-7}	1.69×10^{-8}	3.93×10^{-8}	1.9 m

COMPUTATIONAL GAINS

2x the batch size in a single forward pass when the spiral module is included (for same MSE)!

