

# Coordinate Transformations RW $k < 0$ v3

## Roberston Walker Metric

We may form a 3-space of constant curvature by embedding within a flat 4-space, just as we may embed a 2-sphere or 2 dimensional hyperbola (or also a flat plane) within 3 dimensional space. Constraining to a space of constant curvature, we have

$$\mathbf{x}^2 + z^2 = C^2. \quad (1)$$

Here  $C^2$  represents the degree and sign of curvature, with dimension of length  $C \sim [L]$ . For  $C^2$  positive, we have a bound 3-sphere, while for  $C^2 = 0$ , we have unbound Euclidean geometry, and for  $C^2 < 0$  we have an unbound hyperbolic geometry. Constructing the flat 4-space line element,

$$ds^2 = d\mathbf{x}^2 + dz^2. \quad (2)$$

Taking the differential of (1) allows us to relate  $dz$  to the three space variables  $\mathbf{x}$  via

$$dz^2 = \frac{(\mathbf{x} \cdot d\mathbf{x})^2}{C^2 - \mathbf{x}^2} \quad (3)$$

Substituting into the line element we have

$$ds^2 = d\mathbf{x}^2 + \frac{(\mathbf{x} \cdot d\mathbf{x})^2}{C^2 - \mathbf{x}^2} \quad (4)$$

Adopting polar coordinates, this becomes

$$ds^2 = \frac{dr^2}{1 - r^2/C^2} + r^2 d\Omega^2 \quad (5)$$

With the above general form for a maximally symmetric 3-space with constant curvature, we may form the invariant spacetime interval as

$$ds^2 = dt^2 - a(t)^2 \left( \frac{dr^2}{1 - r^2/C^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2 \right) \quad (6)$$

where  $a(t)$  is an arbitrary function of time to be set by dynamics. Worth noting is that if we rescale  $r' = r/|C|$ , radial distances will be dimensionless and  $a_{rescaled}(t) = a(t)/|C|$  will have dimension of  $[L]$ . Such a rescaling is necessary for the metric convention in which  $\frac{dr^2}{1 - Kr^2}$  for  $K \in [-1, 0, 1]$ . However, cosmological convention utilizes a dimensionless  $a(t)$ , thus we leave in the form of  $r^2/C^2$ .

By a coordinate transformation upon  $t$  via

$$\tau = \int \frac{dt}{a(t)}, \quad (7)$$

we may express (6) in terms of conformal time  $\tau$  as

$$ds^2 = a^2(\tau) \left( d\tau^2 - \frac{dr^2}{1 - r^2/C^2} + r^2 d\Omega^2 \right) \quad (8)$$

# RW to Conformal to Flat Form

## First Transformation

As the first step towards bringing the metric to conformal-flat form for  $C^2 < 0$ , we introduce curvature magnitude  $L^2 = -C^2$  (an inherently positive quantity) and we make coordinate transformations

$$p = \frac{\tau}{L}, \quad \sinh \chi = \frac{r}{L}, \quad (9)$$

which take the line element of (8) into

$$ds^2 = L^2 a^2(p) (dp^2 - d\chi^2 - \sinh^2 \chi d\Omega^2). \quad (10)$$

In this form, all length dimension lies within  $L^2$ .

## Second Transformation (Alternative)

To finally bring (10) to the flat form, we make coordinate substitutions

$$T = e^p \cosh \chi, \quad R = e^p \sinh \chi. \quad (11)$$

It is convenient to introduce a somewhat 'light-like' coordinate defined by

$$X^2 \equiv T^2 - R^2. \quad (12)$$

The coordinate relation for the time coordinate  $p(T, R)$  is in fact only a function of  $X^2$ , viz.

$$e^{2p} = X^2, \quad p = \frac{1}{2} \ln(X^2). \quad (13)$$

For the radial coordinate  $\chi(T, R)$  we have the relations

$$\sinh \chi = \frac{R}{X}, \quad \cosh \chi = \frac{T}{X}. \quad (14)$$

Though not as useful, we may invert (14) to find  $\chi(T, R)$  as

$$\chi = \ln \left( \frac{T + R}{X} \right) \quad (15)$$

To aid in determining the differentials, we note

$$dX = \frac{\partial X}{\partial T} dT + \frac{\partial X}{\partial R} dR = \frac{TdT - RdR}{X}. \quad (16)$$

We first determine  $dp$ :

$$dp = \frac{T}{X^2} dT - \frac{R}{X^2} dR. \quad (17)$$

To find  $d\chi$ , we differentiate  $\sinh \chi$ :

$$d(\sinh \chi) = \cosh \chi d\chi = \frac{dR}{X} - \frac{R}{X^3} (TdT - RdR) \quad (18)$$

$$\frac{T}{X} d\chi = \frac{dR}{X} - \frac{TR}{X^3} dT + \frac{R^2}{X^3} dR, \quad (19)$$

hence

$$d\chi = \frac{dR}{T} - \frac{R}{X^2} dT + \frac{R^2}{TX^2} dR. \quad (20)$$

After repeated usage of  $X^2 = T^2 - R^2$ , we find the coordinate relation between infinitesimals

$$dp^2 - d\chi^2 = \frac{1}{X^2} (dT^2 - dR^2). \quad (21)$$

Finally, with  $\sinh^2 \chi = \frac{R^2}{X^2}$ , we may write the line element in these new coordinates:

$$ds^2 = L^2 \frac{a^2(X)}{X^2} (dT^2 - dR^2 - R^2 d\Omega^2) \quad (22)$$

## Conformal Flat to RW Coordinates

### Conformal Factor

We note that the conformal factor in the flat  $T, R$  coordinates is only a function of  $X^2 = T^2 - R^2$ . The factor is simply

$$\Omega(X)^2 = L^2 \frac{a^2(X)}{X^2} \quad (23)$$

where

$$a(X) = a\left(\frac{1}{2} \ln(X^2)\right). \quad (24)$$

The relation of the conformal factor to the  $p, \chi$  geometry is simple,

$$\Omega^2(X) \equiv \Omega^2(p, \chi) = L^2 a^2(p) e^{-2p}. \quad (25)$$

Interestingly, it is a function entirely of time coordinate  $p$ . We may bring this to the comoving RW form by successive transformations

$$p = \frac{\tau}{L}, \quad \tau = \int \frac{a(t)}{dt}, \quad (26)$$

in which the conformal factor becomes

$$\Omega^2(X) \equiv \Omega^2(t) = L^2 a^2(t) \exp\left[-\frac{2}{L} \int \frac{dt}{a(t)}\right] \quad (27)$$

### Two Step Transformation

From the relations

$$T = e^p \cosh \chi, \quad R = e^p \sinh \chi \quad (28)$$

and

$$p = \frac{\tau}{L}, \quad \sinh \chi = \frac{r}{L} \quad (29)$$

we see that we could enact a coordinate transformation from conformal time ( $\tau$ ) RW geometry

$$ds^2 = a^2(\tau) \left( d\tau^2 - \frac{dr^2}{1 + r^2/L^2} + r^2 d\Omega^2 \right) \quad (30)$$

to conformal to flat (polar) geometry

$$ds^2 = L^2 \frac{a^2(X)}{X^2} (dT^2 - dR^2 - R^2 d\Omega^2) \quad (31)$$

via the effective transformation

$$T = \exp\left(\frac{\tau}{L}\right) \left(1 + \left(\frac{r}{L}\right)^2\right)^{1/2}, \quad R = \exp\left(\frac{\tau}{L}\right) \frac{r}{L}, \quad X^2 \equiv T^2 - R^2 = \exp\left(\frac{2\tau}{L}\right) \quad (32)$$

## One Step Transformation

Lastly, we may substitute the transformation of  $\tau$  viz

$$\tau = \int \frac{dt}{a(t)}, \quad (33)$$

to finally bring us to comoving coordinates. That is, via coordinate transformation

$$T = \exp\left(\frac{1}{L} \int \frac{dt}{a(t)}\right) \left(1 + \left(\frac{r}{L}\right)^2\right), \quad R = \exp\left(\frac{1}{L} \int \frac{dt}{a(t)}\right) \frac{r}{L}, \quad X^2 \equiv T^2 - R^2 = \exp\left(\frac{2}{L} \int \frac{dt}{a(t)}\right) \quad (34)$$

we may transform from comoving coordinates

$$ds^2 = dt^2 - a(t)^2 \left( \frac{dr^2}{1 + r^2/L^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2 \right) \quad (35)$$

to conformal flat (polar) coordinates

$$ds^2 = L^2 \frac{a^2(X)}{X^2} (dT^2 - dR^2 - R^2 d\Omega^2). \quad (36)$$

When  $a(t)$  is specified apriori via a dynamics, exponential factors will simplify, especially for a  $\tau$  which behaves logarithmically. For example, in the early universe radiation era, we have determined  $\tau$  as

$$\tau = L \int_0^t \frac{dt}{(d^2 + t^2)^{1/2}} = L \operatorname{arcsinh}\left(\frac{t}{d}\right). \quad (37)$$

This is equivalent to

$$\tau = L \ln \left( \frac{t}{d} + \sqrt{\left(\frac{t}{d}\right)^2 + 1} \right) \quad (38)$$

in which our exponential calculates to

$$\exp\left(\frac{1}{L} \int \frac{dt}{a(t)}\right) = \frac{t}{d} + \sqrt{\left(\frac{t}{d}\right)^2 + 1}. \quad (39)$$

In the (conformal) early universe then, the conformal factor  $\Omega(X)$  goes as

$$\Omega^2(X) = L^2 a^2(t) \exp\left[-\frac{2}{L} \int \frac{dt}{a(t)}\right] \quad (40)$$

$$= (d^2 + t^2) \left( \frac{t}{d} + \sqrt{\left(\frac{t}{d}\right)^2 + 1} \right)^{-2} \quad (41)$$

The flat space coordinate transformations  $T$  and  $R$  then are specified as

$$T = \left( \frac{t}{d} + \sqrt{\left(\frac{t}{d}\right)^2 + 1} \right) \left( 1 + \left(\frac{r}{L}\right)^2 \right)^{1/2}, \quad R = \left( \frac{t}{d} + \sqrt{\left(\frac{t}{d}\right)^2 + 1} \right) \frac{r}{L} \quad (42)$$

$$X^2 \equiv T^2 - R^2 = \left( \frac{t}{d} + \sqrt{\left(\frac{t}{d}\right)^2 + 1} \right)^2 \quad (43)$$

$$a^2(X) = \frac{d^2}{L^2} \frac{(X^2 + 1)^2}{4X^2} \quad (44)$$

$$\Omega^2(X) = L^2 \frac{a^2(X)}{X^2} = \left[ \frac{d}{2} \left( 1 + \frac{1}{X^2} \right) \right]^2 \quad (45)$$

$$\Omega(X) = \frac{d}{2} (1 + X^{-2}) \quad (46)$$

## Cartesian to Polar

### Transformation Matrices

$$\begin{pmatrix} dx \\ dy \\ dz \end{pmatrix} = \begin{pmatrix} \frac{\partial x}{\partial r} & \frac{\partial x}{\partial \theta} & \frac{\partial x}{\partial \phi} \\ \frac{\partial y}{\partial r} & \frac{\partial y}{\partial \theta} & \frac{\partial y}{\partial \phi} \\ \frac{\partial z}{\partial r} & \frac{\partial z}{\partial \theta} & \frac{\partial z}{\partial \phi} \end{pmatrix} \begin{pmatrix} dr \\ d\theta \\ d\phi \end{pmatrix} = \begin{pmatrix} \sin \theta \cos \phi & r \cos \theta \cos \phi & -r \sin \theta \sin \phi \\ \sin \theta \sin \phi & r \cos \theta \sin \phi & r \sin \theta \cos \phi \\ \cos \theta & -r \sin \theta & 0 \end{pmatrix} \begin{pmatrix} dr \\ d\theta \\ d\phi \end{pmatrix} \quad (47)$$

$$\begin{pmatrix} dr \\ d\theta \\ d\phi \end{pmatrix} = \begin{pmatrix} \sin \theta \cos \phi & \sin \theta \sin \phi & \cos \theta \\ \frac{\cos \theta \cos \phi}{r} & \frac{\cos \theta \sin \phi}{r} & -\frac{\sin \theta}{r} \\ -\frac{\sin \phi}{r \sin \theta} & \frac{\cos \phi}{r \sin \theta} & 0 \end{pmatrix} \begin{pmatrix} dx \\ dy \\ dz \end{pmatrix} \quad (48)$$

### Time-Time

$$K'_{00} = K_{00} \quad (49)$$

### Time-Space

$$K'_{0i} = \frac{\partial x^j}{\partial x'^i} K_{0j} \quad (50)$$

$$\begin{pmatrix} K'_{01} \\ K'_{02} \\ K'_{03} \end{pmatrix} = \begin{pmatrix} \frac{\partial x^1}{\partial x'^1} & \frac{\partial x^2}{\partial x'^1} & \frac{\partial x^3}{\partial x'^1} \\ \frac{\partial x^1}{\partial x'^2} & \frac{\partial x^2}{\partial x'^2} & \frac{\partial x^3}{\partial x'^2} \\ \frac{\partial x^1}{\partial x'^3} & \frac{\partial x^2}{\partial x'^3} & \frac{\partial x^3}{\partial x'^3} \end{pmatrix} \begin{pmatrix} K_{01} \\ K_{02} \\ K_{03} \end{pmatrix} \quad (51)$$

$$K'_{01} = K_{01} \sin(\theta) \cos(\phi) + K_{02} \sin(\theta) \sin(\phi) + K_{03} \cos(\theta) \quad (52)$$

$$K'_{02} = K_{01} r \cos(\theta) \cos(\phi) + K_{02} r \cos(\theta) \sin(\phi) - K_{03} r \sin(\theta) \quad (53)$$

$$K'_{03} = -K_{01} r \sin(\theta) \sin(\phi) + K_{02} r \sin(\theta) \cos(\phi) \quad (54)$$

### Space-Space

$$K'_{ij} = \frac{\partial x^k}{\partial x'^i} K_{kl} \frac{\partial x^l}{\partial x'^j} \quad (55)$$

$$\begin{pmatrix} K'_{11} & K'_{12} & K'_{13} \\ K'_{21} & K'_{22} & K'_{23} \\ K'_{31} & K'_{32} & K'_{33} \end{pmatrix} = \begin{pmatrix} \frac{\partial x^1}{\partial x'^1} & \frac{\partial x^2}{\partial x'^1} & \frac{\partial x^3}{\partial x'^1} \\ \frac{\partial x^1}{\partial x'^2} & \frac{\partial x^2}{\partial x'^2} & \frac{\partial x^3}{\partial x'^2} \\ \frac{\partial x^1}{\partial x'^3} & \frac{\partial x^2}{\partial x'^3} & \frac{\partial x^3}{\partial x'^3} \end{pmatrix} \begin{pmatrix} K_{11} & K_{12} & K_{13} \\ K_{21} & K_{22} & K_{23} \\ K_{31} & K_{32} & K_{33} \end{pmatrix} \begin{pmatrix} \frac{\partial x^1}{\partial x'^1} & \frac{\partial x^2}{\partial x'^1} & \frac{\partial x^3}{\partial x'^1} \\ \frac{\partial x^1}{\partial x'^2} & \frac{\partial x^2}{\partial x'^2} & \frac{\partial x^3}{\partial x'^2} \\ \frac{\partial x^1}{\partial x'^3} & \frac{\partial x^2}{\partial x'^3} & \frac{\partial x^3}{\partial x'^3} \end{pmatrix}^T \quad (56)$$

$$K'_{11} = K_{11} \sin^2(\theta) \cos^2(\phi) + K_{12} \sin^2(\theta) \sin(2\phi) + K_{13} \sin(2\theta) \cos(\phi) + K_{22} \sin^2(\theta) \sin^2(\phi) + K_{23} \sin(2\theta) \sin(\phi) + K_{33} \cos^2(\theta) \quad (57)$$

$$K'_{22} = K_{11} r^2 \cos^2(\theta) \cos^2(\phi) + K_{12} r^2 \cos^2(\theta) \sin(2\phi) - K_{13} r^2 \sin(2\theta) \cos(\phi) + K_{22} r^2 \cos^2(\theta) \sin^2(\phi) - K_{23} r^2 \sin(2\theta) \sin(\phi) + K_{33} r^2 \sin^2(\theta) \quad (58)$$

$$K'_{33} = K_{11} r^2 \sin^2(\theta) \sin^2(\phi) - 2K_{12} r^2 \sin^2(\theta) \sin(\phi) \cos(\phi) + K_{22} r^2 \sin^2(\theta) \cos^2(\phi) \quad (59)$$

$$K'_{12} = K_{11}r \sin(\theta) \cos(\theta) \cos^2(\phi) + K_{12}r \sin(\theta) \cos(\theta) \sin(2\phi) + K_{13}r \cos(2\theta) \cos(\phi) + K_{22}r \sin(\theta) \cos(\theta) \sin^2(\phi) + K_{23}r \cos(2\theta) \sin(\phi) - K_{33}r \sin(\theta) \cos(\theta) \quad (60)$$

$$K'_{13} = -K_{11}r \sin^2(\theta) \sin(\phi) \cos(\phi) + K_{12}r \sin^2(\theta) \cos(2\phi) - K_{13}r \sin(\theta) \cos(\theta) \sin(\phi) + K_{22}r \sin^2(\theta) \sin(\phi) \cos(\phi) + K_{23}r \sin(\theta) \cos(\theta) \cos(\phi) \quad (61)$$

$$K'_{23} = -K_{11}r^2 \sin(\theta) \cos(\theta) \sin(\phi) \cos(\phi) + K_{12}r^2 \sin(\theta) \cos(\theta) \cos(2\phi) + K_{13}r^2 \sin^2(\theta) \sin(\phi) + K_{22}r^2 \sin(\theta) \cos(\theta) \sin(\phi) \cos(\phi) - K_{23}r^2 \sin^2(\theta) \cos(\phi) \quad (62)$$

## Early Universe Setup

Given the geometry

$$ds^2 = (g_{\mu\nu} + K_{\mu\nu})dx^\mu dx^\nu = \Omega^2(\eta_{\mu\nu} + k_{\mu\nu})dx^\mu dx^\nu, \quad (63)$$

upon imposing the conformal gauge condition  $\nabla_\nu K^{\mu\nu} - \frac{1}{2}K^{\mu\nu}g_{(0)}^{\alpha\beta}\partial_\nu g_{\alpha\beta}^{(0)} = 0$ , solutions to the first order source free Bach tensor  $\delta W_{\mu\nu} = 0$  are found to obey

$$\frac{1}{2}\Omega^{-2}\square^2 k_{\mu\nu} = 0 \quad (64)$$

After performing residual gauge transformations to eliminate gauge degrees of freedom, the general momentum eigenstate solution to (64) for a given  $k$ -mode is

$$k_{\mu\nu} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & A_{11} & A_{12} & 0 \\ 0 & A_{12} & -A_{11} & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} e^{ikx} + \begin{pmatrix} 0 & B_{01} & B_{02} & 0 \\ B_{01} & B_{11} & B_{12} & 0 \\ B_{02} & B_{12} & -B_{11} & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} n_\alpha x^\alpha e^{ikx} \quad (65)$$

with timelike  $n_\alpha = (1, 0, 0, 0)$ . The full solution for  $K_{\mu\nu}$  is then given as

$$K_{\mu\nu} = \Omega^2 k_{\mu\nu}. \quad (66)$$

The  $k < 0$  R.W. line element is given in comoving coordinates as

$$ds^2 = dt^2 - a(t)^2 \left( \frac{dr^2}{1 + r^2/L^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2 \right) \quad (67)$$

where  $k = -1/L^2$  (with  $k < 0$ ). By coordinate transformation, the hyperbolic R.W. background geometry may be expressed in the form of  $g_{\mu\nu}^{(0)} = \Omega^2 \eta_{\mu\nu}$ , with the general conformal factor  $\Omega$  having time and spatial dependence in the Minkowski coordinates.

Within the early universe radiation era, the perfect fluid energy momentum tensor obeys  $\rho = 3p$ ,  $\rho = A/a^4(t)$ ,  $A > 0$ , with  $a(t)$  following the evolution equation

$$\begin{aligned} \dot{a}^2 - \frac{1}{L^2} &= \alpha a^2 - \frac{2A}{S_0^2 a^2} \\ &= -2 \frac{a^2}{S_0^2} \left( \lambda_S S_0^4 + \frac{A}{a^4} \right) \end{aligned} \quad (68)$$

With the radiation dominating over the cosmological constant in the early universe (since  $a(t)$  is small), i.e.

$$\frac{A}{a^4} \gg \lambda_S S_0^4, \quad (69)$$

the evolution equation can then be brought to the form

$$L^2 \dot{a}^2 = 1 - \frac{d^2}{L^2} \left( \frac{1}{a^2} \right), \quad (70)$$

in which the solution  $a(t)$  is

$$a^2(t) = \frac{1}{L^2} (d^2 + t^2) \quad (71)$$

where we have defined

$$d^2 \equiv \frac{2AL^4}{S_0^2}. \quad (72)$$

(With  $A \sim [L]^{-4}$  and  $S_0 \sim [L]^{-1}$  fixed early on, the relevant quantities to compare in the radiation dominated era should be the dimensionless  $a(t)$  and  $\lambda_S$ ).

## Notation

From the original form of the scale factor

$$a^2(t) = \frac{2AL^2}{S_0^2} + \frac{t^2}{L^2} \quad (73)$$

we see that for setting up a definition for large  $t$ , we should take

$$\frac{t^2}{L^2} \gg \frac{2AL^2}{S_0^2}. \quad (74)$$

This is equivalent to requiring  $t \gg d$ . If the scale behaves such that  $2AL^2/S_0^2 \ll 1$ , then  $t \gg d$  does not necessarily imply  $t \gg L$ . Noting in addition the R.W. comoving geometry distance  $r/L$ , we introduce two scales of comparison

$$u \equiv \frac{t}{d}, \quad v \equiv \frac{r}{L}. \quad (75)$$

Thus we define large  $t$  behavior as taking  $u \gg 1$ , holding  $v$  finite.

In terms of  $u$  and  $v$ , the scale factor takes the form

$$a^2(u) = \frac{d^2}{L^2} (1 + u^2) \quad (76)$$

comoving R.W. metric takes the form

$$\begin{aligned} ds^2 &= dt^2 - a(t)^2 \left( \frac{dr^2}{1 + r^2/L^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2 \right) \\ &= d^2 \left[ du^2 - (1 + u^2) \left( \frac{dv^2}{1 + v^2} + v^2 d\Omega^2 \right) \right] \end{aligned} \quad (77)$$

## Coordinate Transformations

### Cartesian to Polar

In going from the geometry of

$$ds^2 = \Omega^2 (\eta_{\mu\nu} + k_{\mu\nu}) dx^\mu dx^\nu \quad (78)$$

to

$$ds^2 = \Omega^2(dt^2 - dr^2 - r^2 d\Omega^2 + k_{\mu\nu}^{(P)} dx^\mu dx^\nu), \quad (79)$$

we must perform the appropriate coordinate transformation (given in the Appendix). Denoting the polar coordinate system as  $x^{(P)}$ , we find, after imposing the transverse and residual relations, the following:

$$\begin{aligned} k_{00}^{(P)} &= 0 \\ k_{01}^{(P)} &= k_{01} \sin(\theta) \cos(\phi) + k_{02} \sin(\theta) \sin(\phi) \\ k_{02}^{(P)} &= k_{01} r \cos(\theta) \cos(\phi) + k_{02} r \cos(\theta) \sin(\phi) \\ k_{03}^{(P)} &= -k_{01} r \sin(\theta) \sin(\phi) + k_{02} r \sin(\theta) \cos(\phi) \\ k_{11}^{(P)} &= k_{11} \sin^2(\theta) \cos(2\phi) + k_{12} \sin^2(\theta) \sin(2\phi) \\ k_{22}^{(P)} &= k_{11} r^2 \cos^2(\theta) \cos(2\phi) + k_{12} r^2 \cos^2(\theta) \sin(2\phi) \\ k_{33}^{(P)} &= -k_{11} r^2 \sin^2(\theta) \cos(2\phi) - 2k_{12} r^2 \sin^2(\theta) \sin(\phi) \cos(\phi) \\ k_{12}^{(P)} &= \frac{1}{2} k_{11} r \sin(2\theta) \cos(2\phi) + k_{12} r \sin(\theta) \cos(\theta) \sin(2\phi) \\ k_{13}^{(P)} &= -2k_{11} r \sin^2(\theta) \sin(\phi) \cos(\phi) + k_{12} r \sin^2(\theta) \cos(2\phi) \\ k_{23}^{(P)} &= -2k_{11} r^2 \sin(\theta) \cos(\theta) \sin(\phi) \cos(\phi) + k_{12} r^2 \sin(\theta) \cos(\theta) \cos(2\phi) \end{aligned} \quad (80)$$

Since the  $\square^2 k_{\mu\nu} = 0$  is only valid in a conformal to Minkowski background, upon transforming the solution for  $k_{\mu\nu}$  to polar coordinates, we must account for the factors of  $R(t, r)$  and  $r'(t, r)$  in regards to the asymptotic time behavior. As a rule, every angular index gets a power of  $r$ .

## Original Coordinates

Performing coordinate transformations

$$p' = \frac{u}{(1+u^2)^{1/2} + (1+v^2)^{1/2}}, \quad r' = \frac{v}{(1+u^2)^{1/2} + (1+v^2)^{1/2}} \quad (81)$$

transforms the comoving R.W. line element to the conformal to flat (polar)

$$ds^2 = \Omega^2(p', r')(dp'^2 - dr'^2 - r'^2 d\Omega^2) \quad (82)$$

with conformal factor

$$\Omega^2(p', r') = \frac{4L^2 a^2}{(1 - (p' + r')^2)(1 - (p' - r')^2)} = d^2(1 + u^2) \left[ (1 + u^2)^{1/2} + (1 + v^2)^{1/2} \right]^2. \quad (83)$$

We will soon make use of the coordinate relations

$$\begin{aligned} \frac{\partial p'}{\partial t} &= \frac{1}{d} \frac{\partial p'}{\partial u} = \left( \frac{1}{d} \right) \frac{1 + (1 + u^2)^{1/2}(1 + v^2)^{1/2}}{(1 + u^2)^{1/2}((1 + u^2)^{1/2} + (1 + v^2)^{1/2})^2} \\ \frac{\partial p'}{\partial r} &= \frac{1}{L} \frac{\partial p'}{\partial v} = - \left( \frac{1}{L} \right) \frac{uv}{(1 + v^2)^{1/2}((1 + u^2)^{1/2} + (1 + v^2)^{1/2})^2} \\ \frac{\partial r'}{\partial t} &= \frac{1}{d} \frac{\partial r'}{\partial u} = - \left( \frac{1}{d} \right) \frac{uv}{(1 + u^2)^{1/2}((1 + u^2)^{1/2} + (1 + v^2)^{1/2})^2} \\ \frac{\partial r'}{\partial r} &= \frac{1}{L} \frac{\partial r'}{\partial v} = \left( \frac{1}{L} \right) \frac{1 + (1 + u^2)^{1/2}(1 + v^2)^{1/2}}{(1 + v^2)^{1/2}((1 + u^2)^{1/2} + (1 + v^2)^{1/2})^2} \end{aligned} \quad (84)$$



After transforming from Minkowski to polar, it remains to transform the  $k_{\mu\nu}$  from polar to comoving coordinates. We note that angular coordinates are unaffected. In calculating the transformation (given in the appendix), we have

$$\begin{aligned}
k_{00}^{(cm)} &= 2 \frac{\partial p'}{\partial t} \frac{\partial r'}{\partial t} k_{01}^{(P)} + \left( \frac{\partial r'}{\partial t} \right)^2 k_{11}^{(P)} \\
k_{01}^{(cm)} &= \frac{\partial p'}{\partial t} \frac{\partial r'}{\partial r} k_{01}^{(P)} + \frac{\partial r'}{\partial t} \frac{\partial p'}{\partial r} k_{01}^{(P)} + \frac{\partial r'}{\partial t} \frac{\partial r'}{\partial r} k_{11}^{(P)} \\
k_{02}^{(cm)} &= \frac{\partial p'}{\partial t} k_{02}^{(P)} + \frac{\partial r'}{\partial t} k_{12}^{(P)} \\
k_{03}^{(cm)} &= \frac{\partial p'}{\partial t} k_{03}^{(P)} + \frac{\partial r'}{\partial t} k_{13}^{(P)} \\
k_{11}^{(cm)} &= 2 \frac{\partial p'}{\partial r} \frac{\partial r'}{\partial r} k_{01}^{(P)} + \left( \frac{\partial r'}{\partial r} \right)^2 k_{11}^{(P)} \\
k_{22}^{(cm)} &= k_{22}^{(P)} \\
k_{33}^{(cm)} &= k_{33}^{(P)} \\
k_{12}^{(cm)} &= \frac{\partial p'}{\partial r} k_{02}^{(P)} + \frac{\partial r'}{\partial r} k_{12}^{(P)} \\
k_{13}^{(cm)} &= \frac{\partial p'}{\partial r} k_{03}^{(P)} + \frac{\partial r'}{\partial r} k_{13}^{(P)} \\
k_{23}^{(cm)} &= k_{23}^{(P)}
\end{aligned} \tag{85}$$

### Asymptotics

The leading order solution for  $K_{\mu\nu}$  for a wave propagating along the  $z'$  axis is

$$K_{\mu\nu} = \Omega^2(p', r') [C_{\mu\nu} p' \cos(k(r' \cos \theta - p')) + D_{\mu\nu} \sin(k(r' \cos \theta - p'))] \tag{86}$$

where  $k_\mu = (-k, 0, 0, k)$ ,  $z' = r' \cos \theta$ ,  $C_{\mu\nu} = B_{\mu\nu} + B_{\mu\nu}^*$ , and  $D_{\mu\nu} = i(B_{\mu\nu} - B_{\mu\nu}^*)$ .

Up to leading order in  $u$ , we have:

$$p' \sim 1, \quad r' \sim \frac{1}{u}, \quad \Omega^2(p', r') \sim d^2 u^4. \tag{87}$$

$$\frac{\partial p'}{\partial t} \sim \frac{1}{d} \left( \frac{1}{u^2} \right), \quad \frac{\partial p'}{\partial r} \sim -\frac{1}{L} \left( \frac{1}{u} \right), \quad \frac{\partial r'}{\partial t} \sim -\frac{1}{d} \left( \frac{1}{u^2} \right), \quad \frac{\partial r'}{\partial r} \sim \frac{1}{L} \left( \frac{1}{u} \right). \tag{88}$$

For the plane wave  $\sin(k(z' - p'))$ , the phase equates to

$$z' - p' = \frac{v \cos \theta - u}{(1 + u^2)^{1/2} + (1 + v^2)^{1/2}}. \tag{89}$$

For  $u \rightarrow \infty$ , the above converges and has asymptotic expansion

$$z' - p' \approx -1 + \frac{1}{u} (v + (1 + v^2)^{1/2}) \cos \theta - \frac{1}{u^2} \left( \frac{1}{2} + v^2 + v(1 + v^2)^{1/2} \cos \theta \right) + O\left(\frac{1}{u^3}\right). \tag{90}$$

Hence, to second leading order, the  $(p', z')$  plane wave behave asymptotically as

$$\begin{aligned}
\sin(k(z' - p')) &\approx -\sin(k) + \frac{k \cos(k)}{u} (v \cos \theta + (1 + v^2)^{1/2}) \\
\cos(k(z' - p')) &\approx \cos(k) + \frac{k \sin(k)}{u} (v \cos \theta + (1 + v^2)^{1/2})
\end{aligned} \tag{91}$$

For the tensor transformation behavior, recalling that each angular index goes as  $\sim r'$ , the leading large  $u$  behavior of  $B_{\mu\nu}^{(cm)}$  is calculated as:

$$\begin{aligned} B_{00}^{(cm)} &\sim \frac{1}{d^2} \left( \frac{1}{u^4} \right), & B_{01}^{(cm)} &\sim \frac{1}{dL} \left( \frac{1}{u^3} \right), & B_{02}^{(cm)} &\sim \frac{1}{d} \left( \frac{1}{u^3} \right), & B_{03}^{(cm)} &\sim \frac{1}{d} \left( \frac{1}{u^3} \right) \\ B_{11}^{(cm)} &\sim \frac{1}{L^2} \left( \frac{1}{u^2} \right), & B_{22}^{(cm)} &\sim \frac{1}{u^2}, & B_{33}^{(cm)} &\sim \frac{1}{u^2}, & B_{12}^{(cm)} &\sim \frac{1}{L} \left( \frac{1}{u^2} \right), & B_{13}^{(cm)} &\sim \frac{1}{L} \left( \frac{1}{u^2} \right), & B_{23}^{(cm)} &\sim \frac{1}{u^2} \end{aligned} \quad (92)$$

Finally, we calculate the leading  $u = t/d$  behavior for the comoving  $K_{\mu\nu}^{(cm)}$ , which follows

$$K_{\mu\nu}^{(cm)} = \Omega^2(p', r') B_{\mu\nu}^{(cm)} r' \sin(k(z' - p')) \sim d^2 u^4 B_{\mu\nu}^{(cm)}. \quad (93)$$

$$\begin{aligned} K_{00}^{(cm)} &\sim 1 \\ K_{01}^{(cm)} &\sim \frac{d}{L} u \\ K_{02}^{(cm)} &\sim \frac{1}{d} u \\ K_{03}^{(cm)} &\sim \frac{1}{d} u \\ K_{11}^{(cm)} &\sim \frac{d^2}{L^2} (u^2) \\ K_{22}^{(cm)} &\sim d^2 (u^2) \\ K_{33}^{(cm)} &\sim d^2 (u^2) \\ K_{12}^{(cm)} &\sim \frac{d^2}{L} (u^2) \\ K_{13}^{(cm)} &\sim \frac{d^2}{L} (u^2) \\ K_{23}^{(cm)} &\sim d^2 (u^2) \end{aligned} \quad (94)$$

## New Coordinates

Performing coordinate transformations

$$T = \left[ u + (1 + u^2)^{1/2} \right] (1 + v^2)^{1/2}, \quad R = \left[ u + (1 + u^2)^{1/2} \right] v, \quad X^2 = T^2 - R^2, \quad (95)$$

transforms the comoving R.W. line element to the conformal to flat (polar)

$$ds^2 = \Omega^2(T, R) (dT^2 - dR^2 - R^2 d\Omega^2) \quad (96)$$

with conformal factor

$$\Omega^2(T, R) = \frac{L^2 a^2}{T^2 - R^2} = d^2 (1 + u^2) ((1 + u^2)^{1/2} - u)^2. \quad (97)$$

We will soon make use of the coordinate relations

$$\begin{aligned} \frac{\partial T}{\partial t} &= \frac{1}{d} \frac{\partial T}{\partial u} = \left( \frac{1}{d} \right) \frac{(u + (1 + u^2)^{1/2})(1 + v^2)^{1/2}}{(1 + u^2)^{1/2}} \\ \frac{\partial T}{\partial r} &= \frac{1}{L} \frac{\partial T}{\partial v} = \left( \frac{1}{L} \right) \frac{(u + (1 + u^2)^{1/2})v}{(1 + v^2)^{1/2}} \\ \frac{\partial R}{\partial t} &= \frac{1}{d} \frac{\partial R}{\partial u} = \left( \frac{1}{d} \right) \frac{(u + (1 + u^2)^{1/2})v}{(1 + u^2)^{1/2}} \\ \frac{\partial R}{\partial r} &= \frac{1}{L} \frac{\partial R}{\partial v} = \left( \frac{1}{L} \right) (u + (1 + u^2)^{1/2}) \end{aligned} \quad (98)$$

After transforming from Minkowski to polar, it remains to transform the  $k_{\mu\nu}$  from polar to comoving coordinates. We note that angular coordinates are unaffected. In calculating the transformation (given in the appendix), we have

$$\begin{aligned}
k_{00}^{(cm)} &= 2 \frac{\partial T}{\partial t} \frac{\partial R}{\partial t} k_{01}^{(P)} + \left( \frac{\partial R}{\partial t} \right)^2 k_{11}^{(P)} \\
k_{01}^{(cm)} &= \frac{\partial T}{\partial t} \frac{\partial R}{\partial r} k_{01}^{(P)} + \frac{\partial R}{\partial t} \frac{\partial T}{\partial r} k_{01}^{(P)} + \frac{\partial R}{\partial t} \frac{\partial R}{\partial r} k_{11}^{(P)} \\
k_{02}^{(cm)} &= \frac{\partial T}{\partial t} k_{02}^{(P)} + \frac{\partial R}{\partial t} k_{12}^{(P)} \\
k_{03}^{(cm)} &= \frac{\partial T}{\partial t} k_{03}^{(P)} + \frac{\partial R}{\partial t} k_{13}^{(P)} \\
k_{11}^{(cm)} &= 2 \frac{\partial T}{\partial r} \frac{\partial R}{\partial r} k_{01}^{(P)} + \left( \frac{\partial R}{\partial r} \right)^2 k_{11}^{(P)} \\
k_{22}^{(cm)} &= k_{22}^{(P)} \\
k_{33}^{(cm)} &= k_{33}^{(P)} \\
k_{12}^{(cm)} &= \frac{\partial T}{\partial r} k_{02}^{(P)} + \frac{\partial R}{\partial r} k_{12}^{(P)} \\
k_{13}^{(cm)} &= \frac{\partial T}{\partial r} k_{03}^{(P)} + \frac{\partial R}{\partial r} k_{13}^{(P)} \\
k_{23}^{(cm)} &= k_{23}^{(P)}
\end{aligned} \tag{99}$$

### Asymptotics

The leading order solution for  $K_{\mu\nu}$  for a wave propagating along the  $Z$  axis is

$$K_{\mu\nu} = \Omega^2(T, R) [C_{\mu\nu} T \cos(k(R \cos \theta - T)) + D_{\mu\nu} \sin(k(R \cos \theta - T))] \tag{100}$$

where  $k_\mu = (-k, 0, 0, k)$ ,  $Z = R \cos \theta$ ,  $C_{\mu\nu} = B_{\mu\nu} + B_{\mu\nu}^*$ , and  $D_{\mu\nu} = i(B_{\mu\nu} - B_{\mu\nu}^*)$ .

Up to leading order in  $u$ , we have:

$$T \sim u, \quad R \sim u, \quad \Omega^2(T, R) \sim d^2 \tag{101}$$

$$\frac{\partial T}{\partial t} \sim \frac{1}{d}, \quad \frac{\partial T}{\partial r} \sim \frac{u}{L}, \quad \frac{\partial R}{\partial t} \sim \frac{1}{d}, \quad \frac{\partial R}{\partial r} \sim \frac{u}{L} \tag{102}$$

For the plane wave  $\sin(k(Z - T))$ , the phase equates to

$$Z - T = \left[ u + (1 + u^2)^{1/2} \right] \left[ v \cos \theta - (1 + v^2)^{1/2} \right] \tag{103}$$

For  $u \rightarrow \infty$ , the above diverges and has asymptotic expansion

$$Z - T \approx 2u \left( v \cos \theta - (1 + v^2)^{1/2} \right) + \frac{1}{2u} \left( v \cos \theta - (1 + v^2)^{1/2} \right) + O\left(\frac{1}{u^3}\right) \tag{104}$$

Hence, in the  $(T, Z)$  coordinate system, plane waves remain at least periodic with asymptotic form

$$\begin{aligned}
\sin(k(Z - T)) &\approx \sin \left[ 2ku \left( v \cos \theta - (1 + v^2)^{1/2} \right) \right] \\
\cos(k(Z - T)) &\approx \cos \left[ 2ku \left( v \cos \theta - (1 + v^2)^{1/2} \right) \right]
\end{aligned} \tag{105}$$

For the tensor transformation behavior, recalling that each angular index goes as  $\sim R$ , the leading large  $u$  behavior of  $B_{\mu\nu}^{(cm)}$  is calculated as:

$$\begin{aligned}
B_{00}^{(cm)} &\sim \frac{1}{d^2}, & B_{01}^{(cm)} &\sim \frac{1}{d^2}, & B_{02}^{(cm)} &\sim \frac{u}{d}, & B_{03}^{(cm)} &\sim \frac{u}{d}, & B_{11}^{(cm)} &\sim \frac{u^2}{L^2} \\
B_{22}^{(cm)} &\sim u^2, & B_{33}^{(cm)} &\sim u^2, & B_{12}^{(cm)} &\sim \frac{u^2}{L}, & B_{13}^{(cm)} &\sim \frac{u^2}{L}, & B_{23}^{(cm)} &\sim u^2
\end{aligned} \tag{106}$$

Finally, we calculate the leading  $u = t/d$  behavior for the comoving  $K_{\mu\nu}^{(cm)}$ , which follows

$$K_{\mu\nu}^{(cm)} = \Omega^2(T, R) B_{\mu\nu}^{(cm)} T \sin(k(Z - T)) \sim d^2 u B_{\mu\nu}^{(cm)}. \quad (107)$$

$$\begin{aligned} K_{00}^{(cm)} &\sim u \\ K_{01}^{(cm)} &\sim u \\ K_{02}^{(cm)} &\sim d(u^2) \\ K_{03}^{(cm)} &\sim d(u^2) \\ K_{11}^{(cm)} &\sim \frac{d^2}{L^2}(u^3) \\ K_{22}^{(cm)} &\sim d^2(u^3) \\ K_{33}^{(cm)} &\sim d^2(u^3) \\ K_{12}^{(cm)} &\sim \frac{d^2}{L}(u^3) \\ K_{13}^{(cm)} &\sim \frac{d^2}{L}(u^3) \\ K_{23}^{(cm)} &\sim d^2(u^3) \end{aligned} \quad (108)$$

## Conformal Minkowski to Polar RW Comoving

In going from the geometry of

$$ds^2 = \Omega^2(dT^2 - dx^2 - dy^2 - dz^2) \quad (109)$$

to

$$ds^2 = \Omega^2(dT^2 - dR^2 - R^2 d\Omega^2), \quad (110)$$

we utilize the Cartesian to polar conversions given in the Appendix. Denoting the polar coordinate system as  $x^{(P)}$ , we find, after imposing the transverse and residual relations, the following:

$$\begin{aligned} k_{00}^{(P)} &= 0 \\ k_{01}^{(P)} &= k_{01} \sin(\theta) \cos(\phi) + k_{02} \sin(\theta) \sin(\phi) \\ k_{02}^{(P)} &= k_{01} r \cos(\theta) \cos(\phi) + k_{02} r \cos(\theta) \sin(\phi) \\ k_{03}^{(P)} &= -k_{01} r \sin(\theta) \sin(\phi) + k_{02} r \sin(\theta) \cos(\phi) \\ k_{11}^{(P)} &= k_{11} \sin^2(\theta) \cos(2\phi) + k_{12} \sin^2(\theta) \sin(2\phi) \\ k_{22}^{(P)} &= k_{11} r^2 \cos^2(\theta) \cos(2\phi) + k_{12} r^2 \cos^2(\theta) \sin(2\phi) \\ k_{33}^{(P)} &= -k_{11} r^2 \sin^2(\theta) \cos(2\phi) - 2k_{12} r^2 \sin^2(\theta) \sin(\phi) \cos(\phi) \\ k_{12}^{(P)} &= \frac{1}{2} k_{11} r \sin(2\theta) \cos(2\phi) + k_{12} r \sin(\theta) \cos(\theta) \sin(2\phi) \\ k_{13}^{(P)} &= -2k_{11} r \sin^2(\theta) \sin(\phi) \cos(\phi) + k_{12} r \sin^2(\theta) \cos(2\phi) \\ k_{23}^{(P)} &= -2k_{11} r^2 \sin(\theta) \cos(\theta) \sin(\phi) \cos(\phi) + k_{12} r^2 \sin(\theta) \cos(\theta) \cos(2\phi) \end{aligned} \quad (111)$$

$$K'_{\mu\nu}(t, r, \theta, \phi) = \frac{\partial x^\alpha}{\partial x'^\mu} \frac{\partial x^\beta}{\partial x'^\nu} k_{\alpha\beta}(T, R, \theta, \phi) \quad (112)$$

$$J_{\mu\nu} = \frac{\partial x^\nu}{\partial x'^\mu}, \quad \text{where } x(T, R, \theta, \phi) = x'(t, r, \theta, \phi) \quad (113)$$

$$J_{\mu\nu} = \begin{pmatrix} \frac{\partial T}{\partial t} & \frac{\partial R}{\partial t} & 0 & 0 \\ \frac{\partial T}{\partial r} & \frac{\partial R}{\partial r} & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \quad (114)$$

$$k_{\mu\nu}^{(cm)} = \frac{\partial x_{(P)}^k}{\partial x_{(cm)}^i} k_{kl}^{(P)} \frac{\partial x_{(P)}^l}{\partial x_{(cm)}^j} \quad (115)$$

$$\begin{pmatrix} k_{00}^{(cm)} & k_{01}^{(cm)} & k_{02}^{(cm)} & k_{03}^{(cm)} \\ k_{10}^{(cm)} & k_{11}^{(cm)} & k_{12}^{(cm)} & k_{13}^{(cm)} \\ k_{20}^{(cm)} & k_{21}^{(cm)} & k_{22}^{(cm)} & k_{23}^{(cm)} \\ k_{30}^{(cm)} & k_{31}^{(cm)} & k_{32}^{(cm)} & k_{33}^{(cm)} \end{pmatrix} = \begin{pmatrix} \frac{\partial T}{\partial t} & \frac{\partial R}{\partial t} & 0 & 0 \\ \frac{\partial T}{\partial r} & \frac{\partial R}{\partial r} & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} k_{00}^{(P)} & k_{01}^{(P)} & k_{02}^{(P)} & k_{03}^{(P)} \\ k_{10}^{(P)} & k_{11}^{(P)} & k_{12}^{(P)} & k_{13}^{(P)} \\ k_{20}^{(P)} & k_{21}^{(P)} & k_{22}^{(P)} & k_{23}^{(P)} \\ k_{30}^{(P)} & k_{31}^{(P)} & k_{32}^{(P)} & k_{33}^{(P)} \end{pmatrix} \begin{pmatrix} \frac{\partial T}{\partial t} & \frac{\partial R}{\partial t} & 0 & 0 \\ \frac{\partial T}{\partial r} & \frac{\partial R}{\partial r} & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}^T \quad (116)$$

$$\begin{aligned} k_{00}^{(cm)} &= 2 \frac{\partial T}{\partial t} \frac{\partial R}{\partial t} k_{01}^{(P)} + \left( \frac{\partial R}{\partial t} \right)^2 k_{11}^{(P)} \\ k_{01}^{(cm)} &= \frac{\partial T}{\partial t} \frac{\partial R}{\partial r} k_{01}^{(P)} + \frac{\partial R}{\partial t} \frac{\partial T}{\partial r} k_{01}^{(P)} + \frac{\partial R}{\partial t} \frac{\partial R}{\partial r} k_{11}^{(P)} \\ k_{02}^{(cm)} &= \frac{\partial T}{\partial t} k_{02}^{(P)} + \frac{\partial R}{\partial t} k_{12}^{(P)} \\ k_{03}^{(cm)} &= \frac{\partial T}{\partial t} k_{03}^{(P)} + \frac{\partial R}{\partial t} k_{13}^{(P)} \\ k_{11}^{(cm)} &= 2 \frac{\partial T}{\partial r} \frac{\partial R}{\partial r} k_{01}^{(P)} + \left( \frac{\partial R}{\partial r} \right)^2 k_{11}^{(P)} \\ k_{22}^{(cm)} &= k_{22}^{(P)} \\ k_{33}^{(cm)} &= k_{33}^{(P)} \\ k_{12}^{(cm)} &= \frac{\partial T}{\partial r} k_{02}^{(P)} + \frac{\partial R}{\partial r} k_{12}^{(P)} \\ k_{13}^{(cm)} &= \frac{\partial T}{\partial r} k_{03}^{(P)} + \frac{\partial R}{\partial r} k_{13}^{(P)} \\ k_{23}^{(cm)} &= k_{23}^{(P)} \end{aligned} \quad (117)$$

$$T = \left( \frac{t}{d} + \sqrt{1 + \left( \frac{t}{d} \right)^2} \right) \sqrt{1 + \left( \frac{r}{L} \right)^2}, \quad R = \left( \frac{t}{d} + \sqrt{1 + \left( \frac{t}{d} \right)^2} \right) \frac{r}{L} \quad (118)$$

$$\frac{\partial T}{\partial t} = \frac{1}{d} \left( 1 + \frac{\frac{t}{d}}{\sqrt{1 + \left( \frac{t}{d} \right)^2}} \right) \sqrt{1 + \left( \frac{r}{L} \right)^2} \quad (119)$$

$$\frac{\partial R}{\partial t} = \frac{1}{d} \left( 1 + \frac{\frac{t}{d}}{\sqrt{1 + \left( \frac{t}{d} \right)^2}} \right) \frac{r}{L} \quad (120)$$

$$\frac{\partial T}{\partial r} = \frac{1}{L} \left( \frac{t}{d} + \sqrt{1 + \frac{t^2}{d^2}} \right) \left( \frac{\frac{r}{L}}{\sqrt{1 + \left( \frac{r}{L} \right)^2}} \right) \quad (121)$$

$$\frac{\partial R}{\partial r} = \frac{1}{L} \left( \frac{t}{d} + \sqrt{1 + \frac{t^2}{d^2}} \right) \quad (122)$$

For late times such that  $t \gg d$ , the large time behavior goes as:

$$T \sim \frac{t}{d}, \quad R \sim \frac{t}{d}, \quad \frac{\partial T}{\partial t} \sim \frac{1}{d} \left( \frac{t}{d} \right)^0, \quad \frac{\partial R}{\partial t} \sim \frac{1}{d} \left( \frac{t}{d} \right)^0, \quad \frac{\partial T}{\partial r} \sim \frac{1}{L} \left( \frac{t}{d} \right), \quad \frac{\partial R}{\partial r} \sim \frac{1}{L} \left( \frac{t}{d} \right) \quad (123)$$

We note that in converting from Cartesian to polar, there reside factors of  $R$  in the Jacobian of transformation. Thus, for  $k_{\mu\nu}^{(P)}$  we have, to leading order

$$\begin{aligned} k_{00}^{(P)} &= 0 & k_{01}^{(P)} &\sim \left( \frac{t}{d} \right) & k_{02}^{(P)} &\sim \left( \frac{t}{d} \right)^2 & k_{03}^{(P)} &\sim \left( \frac{t}{d} \right)^2 \\ k_{11}^{(P)} &\sim \left( \frac{t}{d} \right) & k_{22}^{(P)} &\sim \left( \frac{t}{d} \right)^3 & k_{33}^{(P)} &\sim \left( \frac{t}{d} \right)^3 \\ k_{12}^{(P)} &\sim \left( \frac{t}{d} \right)^2 & k_{13}^{(P)} &\sim \left( \frac{t}{d} \right)^2 & k_{23}^{(P)} &\sim \left( \frac{t}{d} \right)^3 \end{aligned} \quad (124)$$

Next, we use (82-83) to determine the late time behavior in comoving coordinates:

$$\begin{aligned} k_{00}^{(cm)} &\sim \frac{1}{d^2} \left( \frac{t}{d} \right) & k_{01}^{(cm)} &\sim \frac{1}{Ld} \left( \frac{t}{d} \right)^2 & k_{02}^{(cm)} &\sim \frac{1}{d} \left( \frac{t}{d} \right)^2 & k_{03}^{(cm)} &\sim \frac{1}{d} \left( \frac{t}{d} \right)^2 \\ k_{11}^{(cm)} &\sim \frac{1}{L^2} \left( \frac{t}{d} \right)^3 & k_{22}^{(cm)} &\sim \left( \frac{t}{d} \right)^3 & k_{33}^{(cm)} &\sim \left( \frac{t}{d} \right)^3 \\ k_{12}^{(cm)} &\sim \frac{1}{L} \left( \frac{t}{d} \right)^3 & k_{13}^{(cm)} &\sim \frac{1}{L} \left( \frac{t}{d} \right)^3 & k_{23}^{(cm)} &\sim \left( \frac{t}{d} \right)^3 \end{aligned} \quad (125)$$

Finally, with the conformal factor late time dependence behaving as

$$\Omega^2(X^2) = \frac{L^2 a^2(X^2)}{X^2} = \frac{d^2 \left( 1 + \frac{t^2}{d^2} \right)}{\left( \frac{t}{d} + \sqrt{1 + \left( \frac{t}{d} \right)^2} \right)^2} \sim d^2 \quad (126)$$

we construct the comoving  $K_{\mu\nu}^{(cm)}$  as

$$K_{\mu\nu}^{(cm)} = \Omega^2 k_{\mu\nu}^{(cm)} \quad (127)$$

to thus have

$$\begin{aligned} K_{00}^{(cm)} &\sim \left( \frac{t}{d} \right) & K_{01}^{(cm)} &\sim \frac{d}{L} \left( \frac{t}{d} \right)^2 & K_{02}^{(cm)} &\sim d \left( \frac{t}{d} \right)^2 & K_{03}^{(cm)} &\sim d \left( \frac{t}{d} \right)^2 \\ K_{11}^{(cm)} &\sim \frac{d^2}{L^2} \left( \frac{t}{d} \right)^3 & K_{22}^{(cm)} &\sim d^2 \left( \frac{t}{d} \right)^3 & K_{33}^{(cm)} &\sim d^2 \left( \frac{t}{d} \right)^3 \\ K_{12}^{(cm)} &\sim \frac{d^2}{L} \left( \frac{t}{d} \right)^3 & K_{13}^{(cm)} &\sim \frac{d^2}{L} \left( \frac{t}{d} \right)^3 & K_{23}^{(cm)} &\sim d^2 \left( \frac{t}{d} \right)^3 \end{aligned} \quad (128)$$

$$\begin{aligned}
k_{00}^{(cm)} &= 2 \frac{\partial p'}{\partial t} \frac{\partial r'}{\partial t} k_{01}^{(P)} + \left( \frac{\partial r'}{\partial t} \right)^2 k_{11}^{(P)} \\
k_{01}^{(cm)} &= \frac{\partial p'}{\partial t} \frac{\partial r'}{\partial r} k_{01}^{(P)} + \frac{\partial r'}{\partial t} \frac{\partial p'}{\partial r} k_{01}^{(P)} + \frac{\partial r'}{\partial t} \frac{\partial r'}{\partial r} k_{11}^{(P)} \\
k_{02}^{(cm)} &= \frac{\partial p'}{\partial t} k_{02}^{(P)} + \frac{\partial r'}{\partial t} k_{12}^{(P)} \\
k_{03}^{(cm)} &= \frac{\partial p'}{\partial t} k_{03}^{(P)} + \frac{\partial r'}{\partial t} k_{13}^{(P)} \\
k_{11}^{(cm)} &= 2 \frac{\partial p'}{\partial r} \frac{\partial r'}{\partial r} k_{01}^{(P)} + \left( \frac{\partial r'}{\partial r} \right)^2 k_{11}^{(P)} \\
k_{22}^{(cm)} &= k_{22}^{(P)} \\
k_{33}^{(cm)} &= k_{33}^{(P)} \\
k_{12}^{(cm)} &= \frac{\partial p'}{\partial r} k_{02}^{(P)} + \frac{\partial r'}{\partial r} k_{12}^{(P)} \\
k_{13}^{(cm)} &= \frac{\partial p'}{\partial r} k_{03}^{(P)} + \frac{\partial r'}{\partial r} k_{13}^{(P)} \\
k_{23}^{(cm)} &= k_{23}^{(P)}
\end{aligned} \tag{129}$$

## Original Transformation

$$p' = \frac{u}{(1+u^2)^{1/2} + (1+v^2)^{1/2}}, \quad r' = \frac{v}{(1+u^2)^{1/2} + (1+v^2)^{1/2}} \quad (130)$$

The above transformation takes us from

$$ds^2 = \Omega^2(p', r')(dp'^2 - dr'^2 - r'^2 d\Omega^2) \quad (131)$$

to

$$ds^2 = dt^2 - a(t)^2 \left( \frac{dr^2}{1+r^2/L^2} + r^2 d\Omega^2 \right) \quad (132)$$

$$\Omega^2(p', r') \sim d^2 \frac{u^4}{4} = \frac{t^4}{4} \quad (133)$$

$$p' \sim u^0, \quad r' \sim \frac{v}{u} \quad (134)$$

## New Transformation

$$T = \left[ u + (1+u^2)^{1/2} \right] (1+v^2)^{1/2}, \quad R = \left[ u + (1+u^2)^{1/2} \right] v \quad (135)$$

The above transformation takes us from

$$ds^2 = \Omega^2(X^2)(dT^2 - dR^2 - R^2 d\Omega^2) \quad (136)$$

to

$$ds^2 = dt^2 - a(t)^2 \left( \frac{dr^2}{1+r^2/L^2} + r^2 d\Omega^2 \right). \quad (137)$$

$$\Omega^2(X^2) \sim \frac{d^2}{4} \quad (138)$$

$$T \sim 2u(1+v^2)^{1/2}, \quad R \sim 2uv \quad (139)$$

## Plane Waves

### Original Coordinates

$$\sin(k(z' - p')) \quad (140)$$

$$p' = \frac{u}{(1+u^2)^{1/2} + (1+v^2)^{1/2}}, \quad r' = \frac{v}{(1+u^2)^{1/2} + (1+v^2)^{1/2}} \quad (141)$$

$$z' = r' \cos \theta = \frac{v \cos \theta}{(1+u^2)^{1/2} + (1+v^2)^{1/2}} \quad (142)$$

$$z' - p' = \frac{v \cos \theta - u}{(1+u^2)^{1/2} + (1+v^2)^{1/2}} \quad (143)$$

For  $u \rightarrow \infty$ , the above converges and has asymptotic expansion

$$z' - p' \approx -1 + \frac{1}{u}(v + (1+v^2)^{1/2}) \cos \theta - \frac{1}{u^2} \left( \frac{1}{2} + v^2 + v(1+v^2)^{1/2} \cos \theta \right) + O\left(\frac{1}{u^3}\right). \quad (144)$$

Hence, to second leading order, the  $(p', z')$  plane waves behave asymptotically as

$$\sin(k(z' - p')) \approx -\sin(k) + \cos(k) \left[ \frac{1}{u}(v + (1+v^2)^{1/2}) \cos \theta \right] \quad (145)$$



### New Coordinates

$$\sin(k(Z - T)) \tag{146}$$

$$T = \left[ u + (1 + u^2)^{1/2} \right] (1 + v^2)^{1/2}, \quad R = \left[ u + (1 + u^2)^{1/2} \right] v \tag{147}$$

$$Z = R \cos \theta = \left[ u + (1 + u^2)^{1/2} \right] v \cos \theta \tag{148}$$

$$Z - T = \left[ u + (1 + u^2)^{1/2} \right] \left[ v \cos \theta - (1 + v^2)^{1/2} \right] \tag{149}$$

For  $u \rightarrow \infty$ , the above diverges and has asymptotic expansion

$$Z - T \approx 2u \left( v \cos \theta - (1 + v^2)^{1/2} \right) + \frac{1}{2u} \left( v \cos \theta - (1 + v^2)^{1/2} \right) + O \left( \frac{1}{u^3} \right) \tag{150}$$

Hence, in the  $(T, Z)$  coordinate system, plane waves remain at least periodic with asymptotic form

$$\sin(k(Z - T)) \approx \sin \left[ 2ku \left( v \cos \theta - (1 + v^2)^{1/2} \right) \right]. \tag{151}$$