

Multicomponent fermionic dark matter and dark baryogenesis

Diego Restrepo

Instituto de Física Universidad de Antioquia Phenomenology Group http://gfif.udea.edu.co



Local Abelian extensions of the SM

Anomaly cancellation of a gauge $U(1)_{\times}$ extension

Any *universal* local Abelian extension of the Standard Model can be reduced to a set of integers which must satisfy the gravitational anomaly, $[SO(1,3)]^2 U(1)_Y$, and the cubic anomaly, $[U(1)_X]^3$ conditions:

$$\sum_{\alpha=1}^{N} z_{\alpha} = 0, \qquad \sum_{\alpha=1}^{N} z_{\alpha}^{3} = 0, \qquad (1)$$

• From a list of N-2 integers, e.g., for N even

$$\mathbf{q} = [l_1, l_2, \cdots, l_n, k_1, k_2, \cdots, k_n], \qquad n = (N-2)/2.$$
 (2)

in the range [-m, m], build two vector-like solutions of N integers,

$$\mathbf{x} = [l_1, k_1, \dots k_n, -l_1, -k_1, \dots - k_n,]$$
 $\mathbf{y} = [0, 0, l_1, \dots l_n, -l_1, \dots - l_n]$ (3)

• From a list of N-2 integers, e.g., for N even

$$\mathbf{q} = [l_1, l_2, \cdots, l_n, k_1, k_2, \cdots, k_n], \qquad n = (N-2)/2.$$
 (2)

in the range [-m, m], build two vector-like solutions of N integers,

$$\mathbf{x} = [l_1, k_1, \dots k_n, -l_1, -k_1, \dots - k_n,]$$
 $\mathbf{y} = [0, 0, l_1, \dots l_n, -l_1, \dots - l_n]$ (3)

• Obtain a (some times) non vector-like solution with $z_{max} = 2m$

$$\mathbf{z} = \mathbf{x} \oplus \mathbf{y} = \left(\sum_{i=1}^{N} x_i y_i^2\right) \mathbf{x} + \left(\sum_{i=1}^{N} x_i^2 y_i\right) \mathbf{y},$$
(4)

• From a list of N-2 integers, e.g., for N even

$$\mathbf{q} = [l_1, l_2, \cdots, l_n, k_1, k_2, \cdots, k_n], \qquad n = (N-2)/2.$$
 (2)

in the range [-m, m], build two vector-like solutions of N integers,

$$\mathbf{x} = [l_1, k_1, \dots k_n, -l_1, -k_1, \dots - k_n,]$$
 $\mathbf{y} = [0, 0, l_1, \dots l_n, -l_1, \dots - l_n]$ (3)

• Obtain a (some times) non vector-like solution with $z_{max} = 2m$

$$\mathbf{z} = \mathbf{x} \oplus \mathbf{y} = \left(\sum_{i=1}^{N} x_i y_i^2\right) \mathbf{x} + \left(\sum_{i=1}^{N} x_i^2 y_i\right) \mathbf{y},$$
 (4)

The parameter space to be explored with $z_{\rm max}=20~(m=10)$ has $96\,153$ non vector-like solutions

of
$$\mathbf{q}$$
 lists = $(2m+1)^{N-2}$ =
$$\begin{cases} 9261 \to 3 & N=5 \\ 194841 \to 38 & N=6 \\ \vdots & \vdots \\ 1.6 \times 10^{13} \to 65910 & N=12 \end{cases}$$
 instead 10^{19}

• From a list of N-2 integers, e.g., for N even

$$\mathbf{q} = [l_1 = 2, l_2 = 3, k_1 = -1, k_2 = -3], \qquad n = 2.$$
 (2)

in the range [-3,3], build two vector-like solutions of 6 integers,

$$\mathbf{x} = [2, -1, -3, -2, 1, 3,]$$
 $\mathbf{y} = [0, 0, 2, \dots, 3, -2, \dots, -3]$ (3)

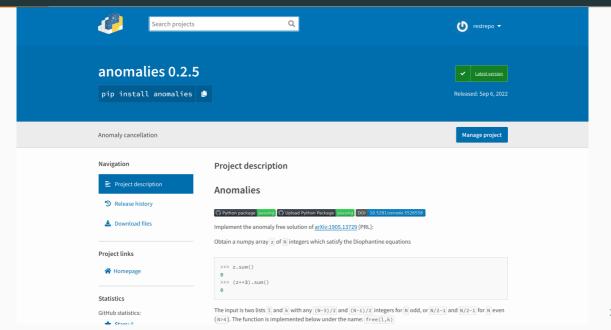
• Obtain a (some times) non vector-like solution with $z_{\text{max}} = 2 \times 3 = 6$

$$\mathbf{z} = \mathbf{x} \oplus \mathbf{y} = \left(\sum_{i=1}^{N} x_i y_i^2\right) \mathbf{x} + \left(\sum_{i=1}^{N} x_i^2 y_i\right) \mathbf{y},$$
(4)

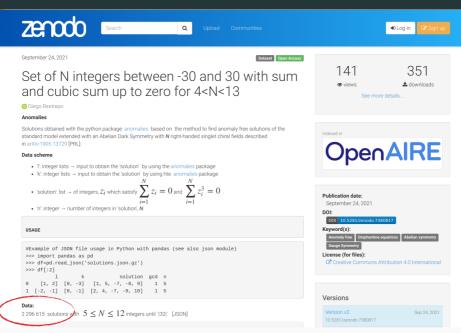
The parameter space to be explored with $z_{\text{max}}=20~(m=10)$ has $96\,153$ non vector-like solutions

of **q** lists =
$$(2m+1)^{N-2}$$
 =
$$\begin{cases} 9261 \to 3 & N = 5 \to [1, -2, -3, 5, 5, -6] \\ 194841 \to 38 & N = 6 \\ \vdots & \vdots \\ 1.6 \times 10^{13} \to 65910 & N = 12 , \text{ instead } 10^{19} \end{cases}$$
 (5)

https://pypi.org/project/anomalies/



https://doi.org/10.5281/zenodo.7380817



Minimal secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_D$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (6)

 $96\,153 \rightarrow 5\,196$

multi-component DM (N=8,12)
ightarrow 142 with three Dirac-fermion DM

Minimal secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_D$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (6)

 $96\,153
ightarrow 5\,196$ multi-component DM (N=8,12) ightarrow 142 with three Dirac-fermion DM

$$z = [1, -2, -2, 4, 5, -7, -7, 8] \rightarrow \phi = 9 \rightarrow [(1, 8), (-2, -7), (4, 5)]$$
 (7)

Minimal secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_D$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (6)

96 153 ightarrow 5 196 multi-component DM ($\mathit{N}=8,12$) ightarrow 142 with three Dirac-fermion DM

$$z = [1, -2, -2, 4, 5, -7, -7, 8] \rightarrow \phi = 9 \rightarrow [(1, 8), (-2, -7), (4, 5)]$$
 (7)

$$\mathcal{L} \subset h_{(1,8)} \psi_1 \psi_8 \phi^* \phi^{(*)} + \sum_{a,b=1}^2 h_{(-2a,-7b)} \psi_4 \psi_{-7} \phi + h_{(4,5)} \psi_4 \psi_5 \phi^* \phi^{(*)} + \text{h.c.}$$
(8)

5

Simplest secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_{D}$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (9)

 $96\,153\rightarrow5\,196$

multi-component DM (N=8,12)
ightarrow 41 with two Dirac-fermion DM

Simplest secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_D$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (9)

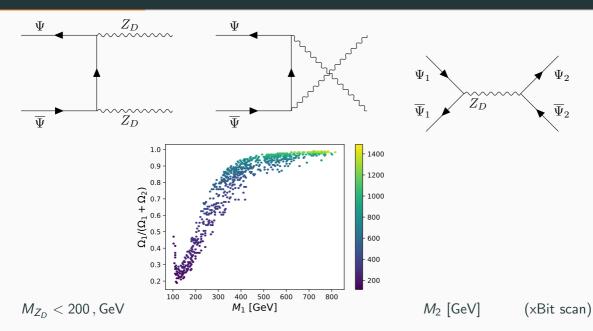
 $96\,153
ightarrow 5\,196$ multi-component DM (N=8,12)
ightarrow 41 with two Dirac-fermion DM

$$z = [1, 2, 2, 4, -5, -5, -7, 8] \rightarrow \phi = 3 \rightarrow [(1, 2), (2, -5), (-5, 8), (4, -7)]$$
 (10)

Simplest secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_{D}$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (9)

 $96\,153
ightarrow 5\,196$ multi-component DM (N=8,12) ightarrow 41 with two Dirac-fermion DM

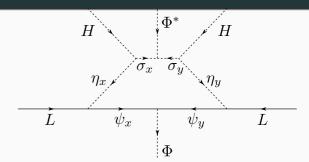


Majorana neutrino masses and mixings

$$\frac{y}{\Lambda}$$
LLHH

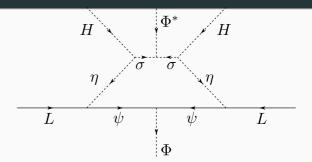
Scotogenic Majorana neutrino masses and mixings

$$\frac{y}{\Lambda}LLHH \to \frac{y}{\Lambda}LLHH\frac{\Phi}{\Lambda}\frac{\Phi^*}{\Lambda}$$



Scotogenic Majorana neutrino masses and mixings

$$\frac{y}{\Lambda} LLHH \rightarrow \frac{y}{\Lambda} LLHH \frac{\Phi}{\Lambda} \frac{\Phi^*}{\Lambda}$$



Already found by Chi-Fong Wong in arXiv:2008.08573 (subset with $N \leq 9$ and $z_{\text{max}} \leq 10$)

$$z = [\underbrace{1,1}_{\psi}, 2, 3, -4, -4, -5, 6] \rightarrow \phi = 2 \rightarrow [(1,1)_a, (2,-4), (4,-6), (4,-7)]$$
 (11)

Additional conditions to reduce

multiplicity

Decrease the number of charges to be assigned to dark matter particles, ψ_i below

$$[\chi_1,\chi_2,\cdots,\psi_1,\psi_2,\cdots,\psi_{N'}]$$

Secluded case:

$$[\nu, \nu, (\nu), \psi_1, \psi_2, \cdots, \psi_{N'}]$$

$$\chi_1 \to \nu_{R1}, \cdots, \chi_{N_\nu} \to \nu_{R\,N_\nu}, \qquad 2 \le N_\nu \le 3\,,$$

(12)

$$\mathcal{L}_{\mathrm{eff}} = h_{\nu}^{\alpha i} \left(\nu_{R\alpha} \right)^{\dagger} \epsilon_{ab} \, L_{i}^{a} \, H^{b} \left(rac{\Phi^{*}}{\Lambda}
ight)^{\delta} + \mathrm{H.c.}, \qquad \mathrm{with} \, \, i = 1, 2, 3 \, ,$$

 Φ is the complex singlet scalar responsible for the SSB of the anomaly-free gauge symmetry and give mass to all ψ_i

$$\phi = -\frac{\nu}{\delta} \,, \tag{13}$$

Decrease the number of charges to be assigned to dark matter particles, ψ_i below

$$[\chi_1,\chi_2,\cdots,\psi_1,\psi_2,\cdots,\psi_{N'}]$$

Secluded case:

Active case:

$$[\nu, \nu, (\nu), \psi_1, \psi_2, \cdots, \psi_{N'}]$$

$$[\nu, \nu, (\nu), m, m, m, \psi_1, \psi_2, \cdots, \psi_{N'}]$$

$$\chi_1 \to \nu_{R1}, \cdots, \chi_{N_{\nu}} \to \nu_{RN_{\nu}}, \qquad 2 \le N_{\nu} \le 3, \quad X(L_i) = -L, \quad X(H) = h \qquad \to m = L - h$$
(12)

$$\mathcal{L}_{\mathrm{eff}} = h_{
u}^{lpha i} \left(
u_{Rlpha}
ight)^{\dagger} \epsilon_{ab} \, L_{i}^{a} \, H^{b} \left(rac{\Phi^{*}}{\Lambda}
ight)^{\delta} + \mathrm{H.c.}, \qquad \mathrm{with} \, \, i = 1, 2, 3 \, ,$$

 Φ is the complex singlet scalar responsible for the SSB of the anomaly-free gauge symmetry and give mass to all ψ_i \to [-4, -4, 1, 1, 5]

$$\phi = -\frac{(\nu + m)}{\delta} \,, \tag{13}$$

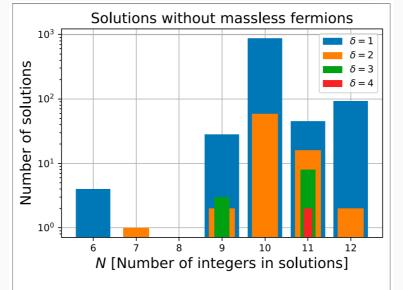
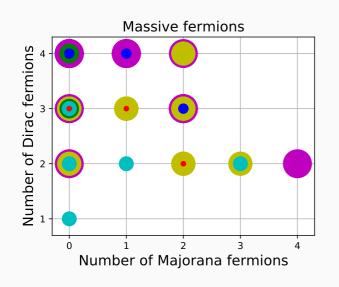
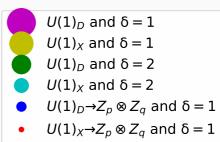


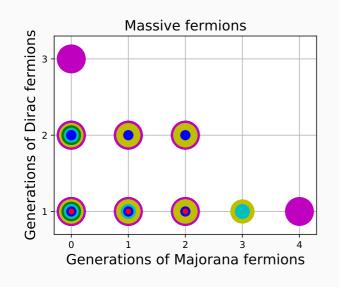
FIGURE 1 Distribution of solutions with N integers to the Diophantine **Eq. 1** which allow the effective Dirac neutrino mass operator at $d = (4 + \delta)$ for at least two right-handed neutrinos and have non-vanishing Dirac o Majorana masses for the other SM-singlet chiral fermions in the solution.

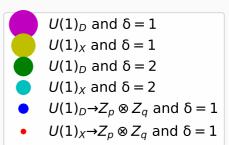
Multi-component dark matter





Multi-flavor dark matter





$U(1)_X$ selection with Dirac-fermionic DM

• Active symmetry m = 3

$$(-5, -5, 3, 3, 3, -7, 8)$$

$U(1)_X$ selection with Dirac-fermionic DM

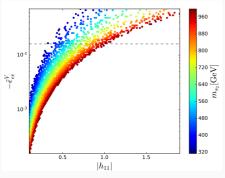
- Active symmetry m = 3
- Effective neutrino mass $\delta = 2 \rightarrow \nu = -5$:

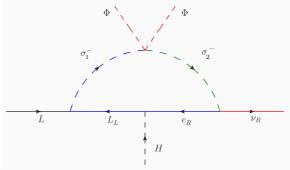
$$(-5, -5, 3, 3, 3, -7, 8)$$

$U(1)_X$ selection with Dirac-fermionic DM

- Active symmetry m = 3
- Effective neutrino mass $\delta = 2 \rightarrow \nu = -5$:
- Active symmetry: $m=3 \rightarrow \phi = -(\nu+m)/\delta = 1$

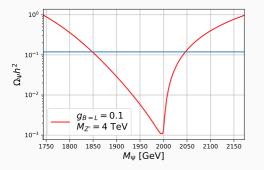
(-5, -5, 3, 3, 3, -7, 8)

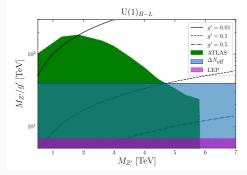




(-5, -5, 3, 3, 3, -7, 8)

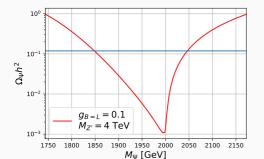
- Active symmetry m = 3
- Effective neutrino mass $\delta = 2 \rightarrow \nu = -5$:
- Active symmetry: $m=3 \rightarrow \phi = -(\nu + m)/\delta = 1$
- Dirac-fermionic DM: $(\psi_L)^{\dagger} \psi_R'' \Phi^* \rightarrow z_6 = -7, z_7 = 8$

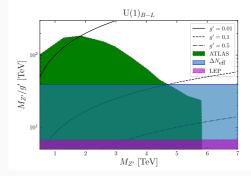




(-5, -5, 3, 3, 3, -7, 8)

- Active symmetry m = 3
- Effective neutrino mass $\delta = 2 \rightarrow \nu = -5$:
- Active symmetry: $m=3 \rightarrow \phi = -(\nu + m)/\delta = 1$
- Dirac-fermionic DM: $(\psi_L)^{\dagger} \psi_R'' \Phi^* \rightarrow z_6 = -7, \ z_7 = 8$





Beyond SM-fermion singlets

Standard model extended with $U(1)_{\mathcal{X}=X \text{ or } \mathbf{D}}$ gauge symmetry

Fields	$SU(2)_L$	$U(1)_Y$	$U(1)_{\mathcal{X}=\mathbf{D} \text{ or } X}$
Q_i^{\dagger}	2	-1/6	Q
d_{Ri}	1	-1/2	d
u_{Ri}	1	+2/3	и
L_i^{\dagger}	2	+1/2	L
e_{Ri}	1	-1	e
Н	2	1/2	h
χ_{α}	1	0	z_{α}

Φ	1	0	ϕ

Table 1: LHC: hadronic production and dileptonic decay

$$i = 1, 2, 3, \ \alpha = 1, 2, \dots, N'$$

Standard model extended with $U(1)_{\mathcal{X}=L \text{ or } \mathbf{B}}$ gauge symmetry

Fields	$SU(2)_L$	$U(1)_Y$	$U(1)_{\mathcal{X}=\mathbf{B}}$ or L
Q_i^{\dagger}	2	-1/6	Q
d_{Ri}	1	-1/2	d
u_{Ri}	1	+2/3	и
L_i^{\dagger}	2	+1/2	L
e_{Ri}	1	-1	e
Н	2	1/2	h = 0
χ_{α}	1	0	z_{lpha}
$(L'_L)^{\dagger}$	2	1/2	-x'
$L_R^{\prime\prime}$	2	-1/2	x''
e_R'	1	-1	x'
$(e_L^{\prime\prime})^\dagger$	1	1	-x"
Ф	1	0	ϕ
S	1	0	S

Table 1: minimal set of new fermion content: L = e = 0 for $\mathcal{X} = B$. Or Q = u = d = 0 for $\mathcal{X} = L$. $i = 1, 2, 3, \alpha = 1, 2, \dots, N'$

14

Anomaly cancellation: $\mathcal{X} = L$ or **B**: beyond SM-singlet fermions

The anomaly-cancellation conditions on $[SU(3)_c]^2 U(1)_X$, $[SU(2)_L]^2 U(1)_X$, $[U(1)_Y]^2 U(1)_X$, allow us to express three of the X-charges in terms of the others

$$\mathbf{u} = -\mathbf{e} - \frac{2}{3}\mathbf{L} - \frac{1}{9}(x' - x''), \quad \mathbf{d} = \mathbf{e} + \frac{4}{3}\mathbf{L} - \frac{1}{9}(x' - x''), \quad \mathbf{Q} = -\frac{1}{3}\mathbf{L} + \frac{1}{9}(x' - x''),$$
(14)

while the $[U(1)_X]^2 U(1)_Y$ anomaly condition reduces to

$$(e+L)(x'-x'')=0.$$
 (15)

- Previously: x' = x''
- We choose instead (h = 0):

$$e = -L, (16)$$

so that (L is still a free parameter)

$$Q = -u = -d = -\frac{1}{3}L + \frac{1}{9}(x' - x''). \tag{17}$$

Anomaly cancellation: $\mathcal{X} = L$ or **B**

The gravitational anomaly, $[SO(1,3)]^2 U(1)_Y$, and the cubic anomaly, $[U(1)_X]^3$, can be written as the following system of Diophantine equations, respectively,

$$\sum_{\alpha=1}^{N} z_{\alpha} = 0, \qquad \sum_{\alpha=1}^{N} z_{\alpha}^{3} = 0, \qquad (18)$$

where

$$z_1 = -x',$$
 $z_2 = x'',$ $z_{2+i} = L, \quad i = 1, 2, 3$ (19)

 \rightarrow

$$9Q = -\sum_{\alpha=1}^{5} z_{\alpha} = -x' + x'' + L + L + L, \qquad (20)$$

$$L = 0 \rightarrow U(1)_B$$
 but $Q = 0 \rightarrow U(1)_L$

$\overline{U(1)_B}$ selection: Neutrinos, dark matter and baryogenesis

•
$$L = 0$$

$$(5,5,-3,-2,1,-6)$$

$U(1)_{B}$ selection: Neutrinos, dark matter and baryogenesis

- L = 0
- Effective Dirac neutrino masses: $\phi = -\nu = -5$

$$(5,5,-3,-2,1,-6)$$

$U(1)_{B}$ selection: Neutrinos, dark matter and baryogenesis

- L = 0
- Effective Dirac neutrino masses: $\phi = -\nu = -5$
- Electroweak-scale vector-like fermions:

$$(L'_L)^{\dagger} L''_R \Phi^* \to x' = -1, \ x'' = 6$$

$$(5,5,-3,-2,\frac{1}{1},-6)$$

$U(1)_{\it B}$ selection: Neutrinos, dark matter and baryogenesis

- L = 0
- Effective Dirac neutrino masses: $\phi = -\nu = -5$
- Electroweak-scale vector-like fermions:

$$(L'_L)^{\dagger} L''_R \Phi^* \to x' = -1, \ x'' = 6$$

• Dirac-fermionic DM:
$$(\chi_L)^{\dagger} \chi_R'' \Phi^* \rightarrow z_3 = -3, \ z_4 = -2$$

$$(5, 5, -3, -2, 1, -6)$$

$U(1)_{\it B}$ selection: Neutrinos, dark matter and baryogenesis

- L = 0
- Effective Dirac neutrino masses: $\phi = -\nu = -5$
- Electroweak-scale vector-like fermions:

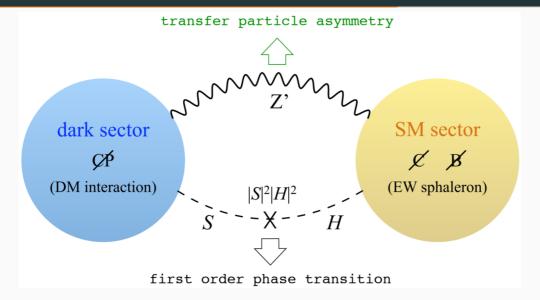
$$(L'_L)^{\dagger} L''_R \Phi^* \to x' = -1, \ x'' = 6$$

• Dirac-fermionic DM: $(\chi_L)^{\dagger} \chi_R'' \Phi^* \rightarrow z_3 = -3, \ z_4 = -2$

959 solutions

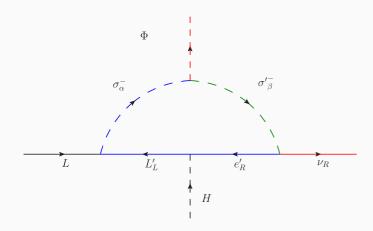
$$(5,5,-3,-2,1,-6)$$

Dark sector baryogenesis



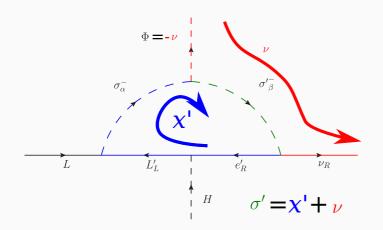
Gauge Baryon number scotogenic realization: arXiv:2205.05762 [PRD]

with Andrés Rivera (UdeA) and Walter Tangarife (Loyola U.)



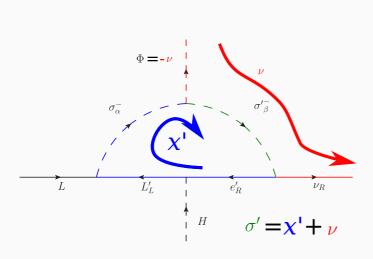
Gauge Baryon number scotogenic realization: arXiv:2205.05762 [PRD]

with Andrés Rivera (UdeA) and Walter Tangarife (Loyola U.)



Gauge Baryon number scotogenic realization: arXiv:2205.05762 [PRD]

with Andrés Rivera (UdeA) and Walter Tangarife (Loyola U.)



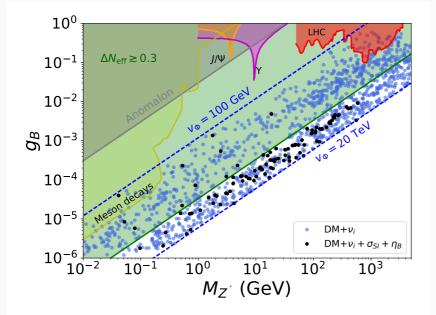
- ')			
Field	$SU(2)_L$	$U(1)_Y$	$U(1)_B$
u_{Ri}	1	2/3	u = 1/3
d_{Ri}	1	-1/3	d = 1/3
$(Q_i)^{\dagger}$	2	-1/6	Q = -1/3
$(L_i)^{\dagger}$	2	1/2	L=0
e_R	1	-1	e = 0
$(L'_L)^{\dagger}$	2	1/2	-x' = -3/5
e'_R	1	-1	x' = 3/5
$L_R^{\prime\prime}$	2	-1/2	x'' = 18/5
$\left(e_L^{\prime\prime} ight)^\dagger$	1	1	-x'' = -18/5
$\nu_{R,1}$	1	0	-3
$\nu_{R,2}$	1	0	-3
χ_R	1	0	6/5
$(\chi_L)^{\dagger}$	1	0	9/5
Н	2	1/2	0
S	1	0	3
Ф	1	0	3
σ_{lpha}^-	1	1	3/5
σ'_{α}^{-}	1	-1	-12/5

19

arXiv:2205.05762 [PRD] https://github.com/anferivera/DarkBariogenesis

- SARAH \rightarrow SPheno \rightarrow MicroMegas
- η_B calculation code
- Python notebook with the scan

Black points: Dirac neutrinos with proper DM and baryon assymetry



Conclusions

A methodology to find all the *universal* gauge Abelian extensions of the standard model is designed.

All of the extensions can be reformulated as the solution of

$$\sum_{\alpha=1}^{N} z_{\alpha} = 0, \qquad \qquad \sum_{\alpha=1}^{N} z_{\alpha}^{3} = 0,$$

which we fully scan until N = 12 and $|z_{max}| = 20$

Once the physical conditions are stablished, the full set of self-consistent models are found from a simple data-analysis procedure. With enough freedom to solve several phenomenological problems.

Baryogenesis

CP violation occurs in the dark sector and is transmitted to SM sector by the new Z' gauge boson.

- High scale fields: Φ , $(\langle \Phi \rangle \to L'_L, L''_R, e'_L, e''_R$: EW-scale vector-like anomalons)
- Electroweak scale (EW) fields: $Z'_{\mu}, S, \chi_L, \chi_R$
- CP-violation

$$\mathcal{L}_{\mathsf{Dirac}\;\mathsf{DM}} = h\left(\chi_{L}\right)^{\dagger} \chi_{R} \Phi^{*} + y\left(\chi_{L}\right)^{\dagger} \chi_{R} S^{*} + \mathsf{h.c}\,, \qquad y \in \mathbb{C}$$
$$\supset \left(m_{\chi} + |y| \,\mathsf{e}^{i\theta} \,|S|\right) \left(\chi_{L}\right)^{\dagger} \chi_{R} + \mathsf{h.c}\,.$$

• CP-violation Portal

$$\mathcal{L}_{\text{anomalous}} \supset g' Z'_{\mu} \left[3\bar{\chi}_L \gamma^{\mu} \chi_L - 2\bar{\chi}_R \gamma^{\mu} \chi_R + \bar{Q}_i \gamma^{\mu} Q_i + \bar{q}_{Ri} \gamma^{\mu} q_{Ri} \right]$$

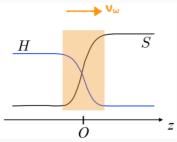
• Strong electroweak phase transition (EWPT) portal

$$\mathcal{L}_{\mathsf{first}}$$
 order EWPT $\supset -\lambda_{SH}H^{\dagger}HS^{*}S$.

$$h = H/\sqrt{2}$$
, $s = |S|$ with vevs: $v(T)$ and $w(T)$ such that $v(T_c) = w(T_c)$

$$V_T(h,s) = \frac{\lambda_H v_c^4}{4} \left(\frac{h^2}{v_c^2} + \frac{s^2}{w_c^2} - 1 \right)^2 + \frac{\lambda_H v_c^2}{m_{s,c}^2 w_{0,c}^4} h^2 s^2 + (T^2 - T_c^2) (c_h h^2 + c_s s^2), \quad (21)$$

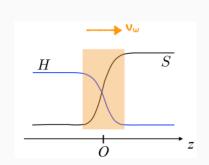
$$c_h = \frac{1}{48} \left(9g_2^2 + 3g_1^2 + 12y_t^2 + 24\lambda_H + \lambda_{HS} \right) , \quad c_s = \frac{1}{12} \left(3\lambda_S + 2\lambda_{HS} \right) .$$
 (22)

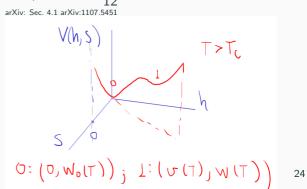


$$h = H/\sqrt{2}$$
, $s = |S|$ with vevs: $v(T)$ and $w(T)$ such that $v(T_c) = w(T_c)$

$$V_T(h,s) = \frac{\lambda_H v_c^4}{4} \left(\frac{h^2}{v_c^2} + \frac{s^2}{w_c^2} - 1 \right)^2 + \frac{\lambda_H v_c^2}{m_{s,c}^2 w_{0,c}^4} h^2 s^2 + (T^2 - T_c^2) (c_h h^2 + c_s s^2), \tag{21}$$

$$c_h = \frac{1}{48} \left(9g_2^2 + 3g_1^2 + 12y_t^2 + 24\lambda_H + \lambda_{HS} \right) , \quad c_s = \frac{1}{12} \left(3\lambda_S + 2\lambda_{HS} \right) . \tag{22}$$

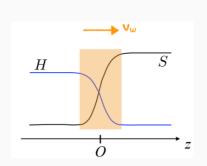


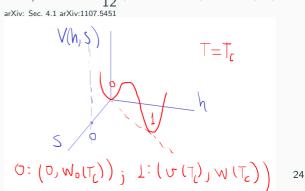


$$h = H/\sqrt{2}$$
, $s = |S|$ with vevs: $v(T)$ and $w(T)$ such that $v(T_c) = w(T_c)$

$$V_T(h,s) = \frac{\lambda_H v_c^4}{4} \left(\frac{h^2}{v_c^2} + \frac{s^2}{w_c^2} - 1 \right)^2 + \frac{\lambda_H v_c^2}{m_{s,c}^2 w_{0,c}^4} h^2 s^2 + (T^2 - T_c^2) (c_h h^2 + c_s s^2), \tag{21}$$

$$c_h = \frac{1}{48} \left(9g_2^2 + 3g_1^2 + 12y_t^2 + 24\lambda_H + \lambda_{HS} \right), \quad c_s = \frac{1}{12} \left(3\lambda_S + 2\lambda_{HS} \right).$$
 (22)

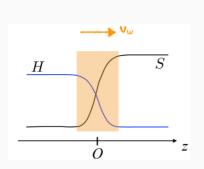


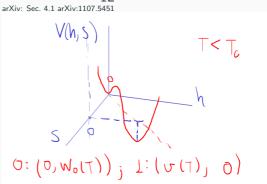


$$h = H/\sqrt{2}$$
, $s = |S|$ with vevs: $v(T)$ and $w(T)$ such that $v(T_c) = w(T_c)$

$$V_T(h,s) = \frac{\lambda_H v_c^4}{4} \left(\frac{h^2}{v_c^2} + \frac{s^2}{w_c^2} - 1 \right)^2 + \frac{\lambda_H v_c^2}{m_{s,c}^2 w_{0,c}^4} h^2 s^2 + (T^2 - T_c^2) (c_h h^2 + c_s s^2), \quad (21)$$

$$c_h = \frac{1}{48} \left(9g_2^2 + 3g_1^2 + 12y_t^2 + 24\lambda_H + \lambda_{HS} \right) , \quad c_s = \frac{1}{12} \left(3\lambda_S + 2\lambda_{HS} \right) . \tag{22}$$





CP assymetry generation i

Using the thin wall approximantion for the nucleation bubbles, we use the ansatz in which the space dependence of the fields is given by

$$h(z) = \frac{1}{2}v(T_n)(1-\tanh(z/L_w)), \qquad s(z) = \frac{1}{2}w_0(T_n)(1+\tanh(z/L_w)),$$

where z is the direction normal to the wall and L_w is the wall width.

The nucleation temperature, T_n , is defined by the condition

$$\exp(-S_3/T_n) = \frac{3}{4\pi} \left(\frac{H(T_n)}{T_n}\right)^4 \left(\frac{2\pi T_n}{S_3}\right)^{\frac{3}{2}},$$

where S_3 is the Euclidean action of the bubble and H(T) is the Hubble rate.

Boltzmann equation i

$$egin{align} \xi_i(z) &\equiv \mu_i(z)/T = 6\left(n_i - \overline{n}_i
ight)/T^3, \ &-D_L \xi_{\chi_L}'' - v_w \xi_{\chi_L}' + \Gamma_L (\xi_{\chi_L} - \xi_{\chi_R}) = S_{\mathcal{CP}}, \end{aligned}$$

where D_L is the diffusion constant for χ_L , which is related to the scattering rate Γ_L by

$$D_{L} = \frac{3x+2}{x^{2}+3x+2} \frac{1}{3\Gamma_{L}}, \qquad x \equiv m_{\chi}/T$$
 (23)

and

$$S_{QP} = -\frac{\lambda}{2} \frac{v_w D_L}{\frac{3x+2}{x^2+3x+2} T} \frac{(1-x)e^{-x} + x^2 E_1(x)}{4m_\chi^2 K_2(x)} \frac{m_\chi w_0(T_n) \lambda \left(-2 + \cosh\left(\frac{2z}{L_w}\right)\right) \sin\theta}{L_w^3 \cosh^4\left(\frac{z}{L_w}\right)}, \quad (24)$$

where v_w is the wall's velocity $E_1(x)$ is the error function and $K_2(x)$ is the modified Bessel function of the second kind. $\mathbf{y} = \lambda e^{i\theta - i\pi/2}$

Transfer DM assymetry to SM quarks

The chiral particle give rise to a non-zero $U(1)_B$ charge density in the proximity of the wall. This results in a Z' background that couples to the SM fields with $U(1)_B$ charge,

$$\langle Z'_0(z) \rangle = \frac{g_B (q_{\chi_L} - q_{\chi_R}) T_n^3}{6 M_{Z'}} \int_{-\infty}^{\infty} dz_1 \, \xi_{\chi_L}(z_1) \, e^{-M_{Z'}|z-z_1|} \,,$$

which generates a chemical potential for the SM quarks,

$$\mu_Q(z) = \mu_{d_R,u_R}(z) = 3 \times \frac{5}{9} \times g_B \langle Z_0'(z) \rangle.$$

This chemical potential sources a thermal-equilibrium asymmetry in the quarks, $\Delta n_O^{\rm EQ}(z) \sim T_n^2 \mu_Q(z)$.

From [1]

If the Z' is sufficiently light, it mediates a long range force that extends into the region outside the bubble wall with unbroken electroweak symmetry.

Finally, the baryon-number asymmetry is then given by

$$n_B \,=\, rac{\Gamma_{
m sph}}{v_W} \int_0^\infty {
m d}\, z \, n_Q^{
m EQ}(z) \, \exp\left(-rac{\Gamma_{
m sph}}{v_W} \, z
ight) \,,$$

where $\Gamma_{\rm sph}$ is the sphaleron rate. The baryon-to-photon-number ratio is then obtained by

$$\eta_B = \frac{n_B}{s(T_n)}, \quad s(T) \equiv \frac{2\pi^2}{45} g_{*S}(T) T^3,$$

where $g_{*S}(T)$ is the effective number of relativistic degrees of freedom.

Our goal is to find what regions of the parameter space yield

$$0.82 \times 10^{-10} < \eta_B < 0.92 \times 10^{-10} \,. \tag{25}$$