Dark matter from SM gauge extensions



with neutrino masses

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Focus on

In collaboration wit

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Hidden sectors



 $\frac{1}{\Lambda}L \cdot HL \cdot H$ $\frac{1}{\Lambda}(\nu_R)^{\dagger}L \cdot HS$









 $\frac{1}{\Lambda}L \cdot HL \cdot H$ $\frac{1}{\Lambda} (\nu_R)^{\dagger} L \cdot HS$



















Dark matter and unification

Unification: SO(10)

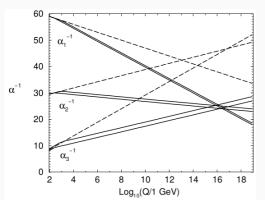
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u_L
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$$\Rightarrow \mathcal{L}_{SM} \supset h \, \mathbf{16}_F \times \mathbf{16}_F \times \mathbf{10}_S + \text{h.c}$$



Standard Model: Z ₂ -even	New Z ₂ -odd particles
Fermions: 16 _F	$10_F, 45_F, \cdots$
Scalars: 10 _{<i>H</i>} , 45 _{<i>H</i>} · · ·	16 _{<i>H</i>} , ⋅ ⋅ ⋅

Lightest Odd Particle (LOP) may be a suitable dark matter candidate, and can improve gauge coupling unification



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	fermions	scalars
$SU(2)_L \times U(1)_Y$	even $SO(10)$	odd $SO(10)$
representation	representations	representations
1_0	45, 54, 126, 210	16, 144
$2_{\pm 1/2}$	10, 120, 126, 210, 210'	16, 144
(3_0)	45, 54, 210	144

 $SU(3)_C: 3(T), 6, 8(\Lambda)$

$$m_{3_0} = 2.7 \text{ TeV}, \qquad m_{\Lambda} \sim 10^{10} \text{ TeV}, \qquad m_{\text{GUT}} \sim 10^{16} \text{ GeV} \,.$$

arXiv:0912.1545 (Frigerio-Hambye)

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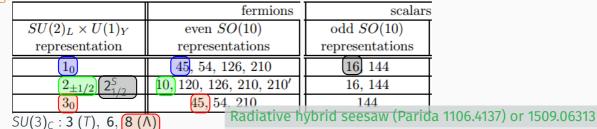
 $SU(3)_{C}: 3(T), 6, 8(\Lambda)$

Split-SUSY like

arXiv:1509.06313 (C. Arbelaez, R. Longas, D.R, O. Zapata)

Standard Model: Z_2 -evenNew Z_2 -odd particlesFermions: $\mathbf{16}_F$ $\mathbf{10}_F, \mathbf{45}_F, \cdots$ Scalars: $\mathbf{10}_H, \mathbf{45}_H \cdots$ $\mathbf{16}_H, \cdots$

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Partial Split-SUSY-like spectrum: bino-higgsino-wino

rartial spiit-sost-like spectrum. bino-mggsmo-wii

 $10'_H$ with fermion DM or, $16_H, \cdots$ with scalar DM

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CU(2) $O(T)$ $O(C)$	1509.06313	

 $SU(3)_C: \overline{3(T)} \ 6, \overline{8(\Lambda)}$

SUSY-like spectrum: bino-higgsino-wino

 $10'_H$ with fermion DM or, $16_H, \cdots$ with scalar DM

Left-Right symmetric realization

Singlet-doublet fermion dark matter

Field	Multiplicity	$3_c 2_L 2_R 1_{B-L}$	Spin	SO(10) origin
Φ	1	(1, 2, 2, 0)	0	10
χ , χ^c	1	(1, 2, 2, 0)	1/2	10
N	1	(1, 1, 1, 0)	1/2	45

Table 1: The relevant part of the field content. Note that, the two fermion doublets χ and χ^c come from an only fermionic LR bidoublet. In the third column the relevant fields are characterized by their $SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$ quantum numbers while their SO(10) origin is specified in the fourth column.

Unification

m_{LR} (GeV)	$3_c 2_L 2_R 1_{B-L}$	m_G (GeV)
2×10^{3}	$\Phi_{1,2,2,0} + 2\Phi_{1,1,3,-2} + \Psi_{1,1,3,0} + \Phi_{1,1,3,0} + \Phi_{8,1,1,0}$	1.65×10^{16}
÷	:	÷

Table 2: $\Delta_{L,R} = 2\Phi_{1,1,3,-2}$. m_{LR} and m_G are given in GeV.

Triplets

Minimal Left-Right Symmetric Standard Model

Field	Multiplicity	$3_c 2_L 2_R 1_{B-L}$	Spin	SO(10) origin
Q	3	$(3,2,1,+\frac{1}{3})$	1/2	16
Q ^c	3	$(\overline{3},1,2,-\frac{1}{3})$	1/2	16
L	3	(1, 2, 1, -1)	1/2	16
_C	3	(1,1,2,+1)	1/2	16
Ф	1	(1, 2, 2, 0)	0	10
Δ_R	1	(1,1,3,-2)	0	126

Left-singlet right-triplet DM

Field	Multiplicity	$3_c 2_L 2_R 1_{B-L}$	Spin	SO(10) origin
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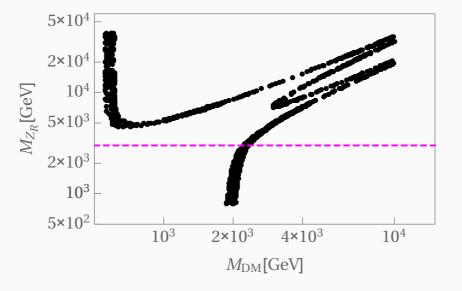


Figure 1: Proper relic density scan: $0.5 < v_R/\text{TeV} < 50$

Mixed Left-singlet right-triplet DM

Field	Multiplicity	$3_c 2_L 2_R 1_{B-L}$	Spin	SO(10) origin
Q	3	$(3,2,1,+\frac{1}{3})$	1/2	16
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Ψ ₁₁₃₀	1	(1,1,3,0)	1/2	45
Ψ_{1132}	1	(1,1,3,2)	1/2	126
Ψ_{113-2}	1	(1,1,3,-2)	1/2	126

$$\Psi_{1132} = \begin{pmatrix} \Psi^{+}/\sqrt{2} & \Psi^{++} \\ \Psi^{0} & -\Psi^{+}/\sqrt{2} \end{pmatrix}, \qquad \bar{\Psi}_{113-2} = \begin{pmatrix} \Psi^{-}/\sqrt{2} & \overline{\Psi}^{0} \\ \Psi^{--} & -\Psi^{-}/\sqrt{2} \end{pmatrix}. \tag{1}$$

$$L \supset M_{11} \operatorname{Tr}(\Psi_{1130}\Psi_{1130}) + M_{23} \operatorname{Tr}(\Psi_{1132}\bar{\Psi}_{113-2}) + \lambda_{13} \operatorname{Tr}(\Delta_R\bar{\Psi}_{113-2}\Psi_{1130}) + \lambda_{12} \operatorname{Tr}(\Delta_R^{\dagger}\Psi_{1132}\Psi_{1130}),$$
 (2)

$$\tan \gamma = \frac{\lambda_{13}}{\lambda_{12}}, \qquad \lambda = \sqrt{\lambda_{12}^2 + \lambda_{13}^2}. \tag{3}$$

Blind spot at

$$M_{23}\sin 2\gamma - M_{\rm DM} = 0 \tag{4}$$

Proper relic density scan

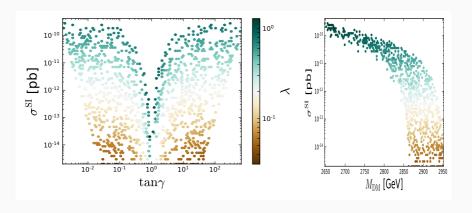


Figure 2: $M_{11} = 50 \text{ TeV } 2.7 < M_{23}/\text{TeV} < 3.1$ (Right: $\tan \gamma > 5$)

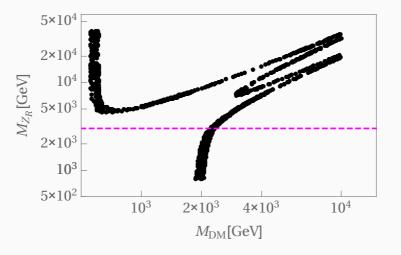


Figure 3:

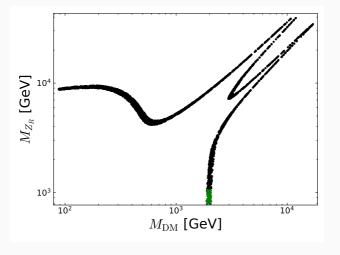


Figure 3: Proper relic density scan: v_R : [2,50] TeV, M_{23} : [0.2,50] TeV, M_{11} : 50 TeV, $\tan \gamma = -1$ and $\lambda = 0.14$.

Direct detection cross section

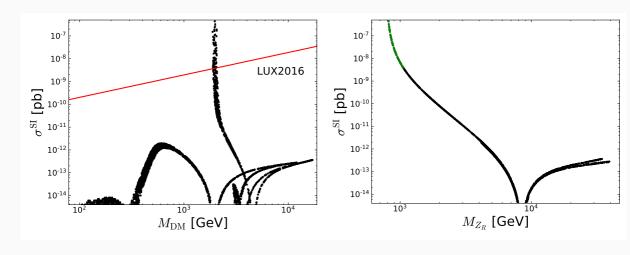


Figure 4: v_R : [2,50] TeV, M_{23} : [0.2,50] TeV, M_{11} : 50 TeV, $\tan \gamma = -1$ and $\lambda = 0.14$.

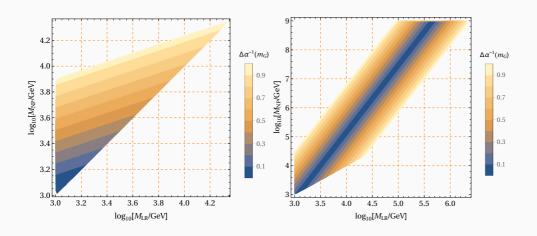
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Ψ_{1132}	1	(1, 1, 3, 2)	1/2	126
Ψ_{113-2}	1	(1,1,3,-2)	1/2	126
Ψ_{1310}	1	(1, 3, 1, 0)	1/2	45
Ψ_{8110}	1	(1, 1, 8, 0)	1/2	45
$\Psi_{321\frac{1}{3}}$	1	(3, 2, 1, 1/3)	1/2	16
$\Psi_{321-\frac{1}{3}}$	1	(1,2,3,-1/3)	1/2	16

Unification quality



Conclusions

In addition to accommodate usual simplified dark matter models, Left-right symmetric standard models have additional DM portals:

New Δ_R portal for direct detection of left-singlet right-triplet mixed dark matter, in companion with left-singlets charged and doubly charged fermions.

Next: Search for them in compressed spectra scenarios at the LHC

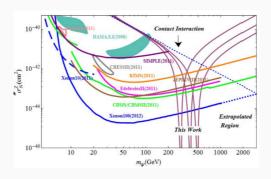
Dirac fermion dark matter

Isosinglet dark matter candidate

 ψ as a isosinglet Dirac dark matter fermion charged under a local U(1)_X (SM) cuples to a SM-singlet vector mediator X as

$$\mathcal{L}_{\text{int}} = -g_{\psi} \, \overline{\psi} \gamma^{\mu} \psi X_{\mu} - \sum_{f} g_{f} \bar{f} \gamma^{\mu} f X_{\mu} \,,$$

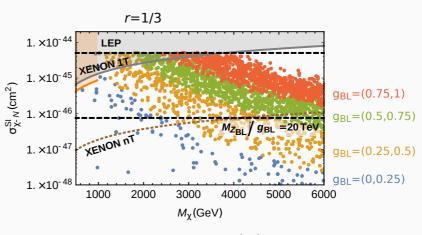
where f are the Standard Model fermions



Isosinglet Dirac fermion dark matter model and Seesaw scale

Left Field	$U(1)_{B-L}$
$(\nu_{R_1})^{\dagger}$	+1
$(u_{R_2})^\dagger$	+1
$(u_{R_2})^\dagger$	+1
ψ_L	-r
$(\psi_{R})^\dagger$	r
ϕ	2

$$\chi = \begin{pmatrix} \psi_{\mathsf{L}} \\ \psi_{\mathsf{R}} \end{pmatrix}$$



Duerr et al: 1803.07462 [PRD]

Singlet-Doublet Dirac Dark matter

Model (SD³M)

Singlet-Doublet Dirac Dark Matter (SD³M) By Carlos E. Yaguna. arXiv:1510.06151 [PRD].

The model extends the standard model (SM) particle content with Dirac Fermions: from SU(2) doublets of Weyl fermions: $\Psi_L = (\Psi_L^0, \Psi_L^-)^\mathsf{T}, \widetilde{(\Psi_R)} = ((\Psi_R^-)^\dagger, -(\Psi_R^0)^\dagger)^\mathsf{T}$ and singlet Weyl fermions ψ_{LR} that interact among themselves and with the SM fields

$$\mathcal{L} \supset M_{\psi} (\psi_R)^{\dagger} \psi_L + M_{\psi} (\widetilde{\Psi}_R) \cdot \Psi_L + y_1 (\psi_R)^{\dagger} \Psi_L \cdot H + y_2 (\widetilde{\Psi}_R) \cdot \widetilde{H} \psi_L + \text{h.c}$$
 (5)

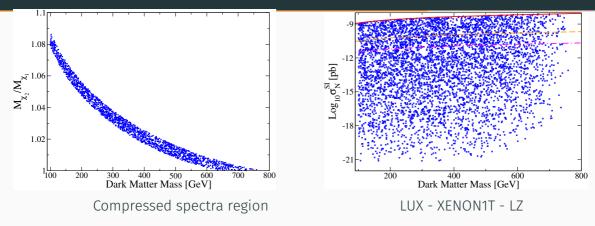
Four free parameters:

$$M_{\psi}, M_{\Psi} < 2 \text{ GeV},$$
 $y_1, y_2 > 10^{-6}$ (6)

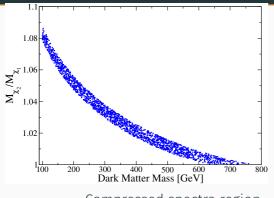
Two neutral Dirac fermion eigenstates:

$$M = \begin{pmatrix} M_{\psi} & y_2 v / \sqrt{2} \\ y_1 v / \sqrt{2} & M_D \end{pmatrix}, \qquad M_{\text{diag}} = \begin{pmatrix} M_{\chi_1} & 0 \\ 0 & M_{\chi_2} \end{pmatrix} = U_L^{\dagger} M U_R$$
 (7)

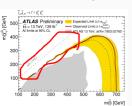
SD³M By Carlos E. Yaguna. arXiv:1510.06151 [PRD].

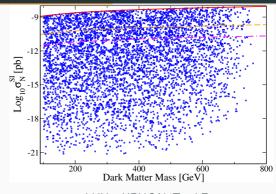


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Compressed spectra region





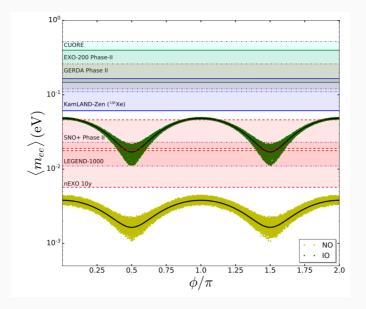
LUX - XENON1T - LZ

Neutrino masses

Lepton number

- Lepton number (*L*) is an accidental discret or Abelian symmetry of the standard model (SM).
- · Without neutrino masses L_e , L_μ , L_τ are also conserved.
- The processes which violates individual *L* are called Lepton flavor violation (LFV) processes.
- · All the neutrino mass models predict, to some extent, LFV processes
- Only models with Majorana neutrinos predict processes with total $L = L_e + L_\mu + L_\tau$ violation, like neutrino less doublet beta decay (NLDBD).
- NLDBD is experimentally challenging, specially if there is a massless neutrino in the spectrum.

NLDBD prospects for a model with a massless neutrino (arXiv:1806.09977 [PLB] with Reig, Valle and Zapata)



Total lepton number: $L = L_e + L_\mu + L_\tau$

Majorana U(1)[

Field
$$Z_2 (\omega^2 = 1)$$

SM 1
 $L \qquad \omega$
 $(e_R)^{\dagger} \qquad \omega$
 $(\nu_R)^{\dagger} \qquad \omega$

$$\mathcal{L}_{\nu} = h_D (\nu_R)^{\dagger} L \cdot H + M_R \nu_R \nu_R + \text{h.c.}$$

$$h_D \sim \mathcal{O}(1)$$

Dirac $U(1)_L$

Field
$$Z_3$$
 ($\omega^3 = 1$)

SM 1

 L ω
 $(e_R)^{\dagger}$ ω^2
 $(\nu_R)^{\dagger}$ ω^2

$$\mathcal{L}_{\nu} = h_D (\nu_R)^{\dagger} L \cdot H + \text{h.c.}$$

$$h_D \sim \mathcal{O}(10^{-11})$$

Total lepton number: $L=L_e+L_\mu+L_ au$

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$$Z_2$$
 ($\omega^2 = 1$)

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Dirac $U(1)_{B-L}$

Field
$$Z_3$$
 ($\omega^3 = 1$)
SM 1
 L ω
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$$\mathcal{L}_{\nu} = h_D (\nu_R)^{\dagger} L \cdot H + \text{h.c.}$$

$$h_D \sim \mathcal{O}(10^{-11})$$

Explain smallness ala Peccei-Quinn: $U(1)_{B-L} \xrightarrow{(S)} Z_N$, $N \ge 3$.

To explain the smallness of Dirac neutrino masses choose $U(1)_{B-L}$ which:

• Forbids tree-level mass (TL) term (Y(H) = +1/2)

$$\mathcal{L}_{T.L} = h_D \epsilon_{ab} (\nu_R)^{\dagger} L^a H^b + \text{h.c}$$
$$= h_D (\nu_R)^{\dagger} L \cdot H + \text{h.c}$$

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• Forbids Majorana term: $u_{R}
u_{R}$

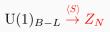
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- Forbids Majorana term: $\nu_R \nu_R$
- Realizes of the 5-dimension operator which conserves lepton number in $SU(3)_c \times SU(2)_L \times U(1)_Y \times U(1)_{B-L}$:

$$\mathcal{L}_{5-D} = \frac{h_{\nu}}{\Lambda} (\nu_R)^{\dagger} L \cdot HS + \text{h.c}$$





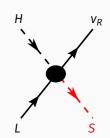
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$$\mathcal{L}_{5-D} = \frac{h_{\nu}}{\Lambda} (\nu_R)^{\dagger} L \cdot HS + \text{h.c.}$$



 $U(1)_{B-L} \stackrel{\langle S \rangle}{\to} Z_N$

• Enhancement to the effective number of degrees of freedom in the early Universe $\Delta N_{\rm eff} = N_{\rm eff} - N_{\rm eff}^{\rm SM}$ (see arXiv:1211.0186)

See E. Ma, Rahul Srivastava: arXiv:1411.5042 [PLB] for tree-level realization

From 1210.6350 and 1805.02025: $\Delta N_{\text{eff}} = 3 (T_{\nu_R}/T_{\nu_L})^4$

$$\Gamma_{\nu_R}(T) = n_{\nu_R}(T) \sum_f \langle \sigma_f(\nu_R \bar{\nu}_R \to f\bar{f}) v \rangle$$

$$= \sum_f \frac{g_{\nu_R}^2}{n_{\nu_R}} \int \frac{d^3p}{(2\pi)^3} \frac{d^3q}{(2\pi)^3} f_{\nu_R}(p) f_{\nu_R}(q) \sigma_f(s) (1 - \cos\theta),$$

$$s = 2pq(1 - \cos \theta), f_{\nu_R}(k) = 1/(e^{k/T} + 1)$$

$$n_{\nu_R}(T) = g_{\nu_R} \int \frac{d^3k}{(2\pi)^3} f_{\nu_R}(k), with g_{\nu_R} = 2$$

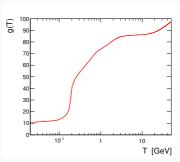
$$\sigma_f(s) \simeq \frac{N_C^f (Q_{BL}^f)^2 Q^2 s}{12\pi} \left(\frac{g'}{M_{el}}\right)^4, In the limit M_{Z'}^2 \gg s.$$

with three right-handed neutrinos, the Hubble parameter is

$$H(T) = \sqrt{\frac{4\pi^3 G_N \left[g(T) + 21/4\right]}{45}} T^2.$$

The right-handed neutrinos decouple when

$$\Gamma_{\nu_R}(T_{\text{dec}}^{\nu_R}) = H(T_{\text{dec}}^{\nu_R}).$$



A. Solaguren-Beascoa, M. C. Gonzalez-Garcia: arXiv:1210.6350 [PLB]

From 1210.6350 and 1805.02025: $\Delta N_{\text{eff}} = 3 \left(T_{\nu_R} / T_{\nu_L} \right)^4$

$$\begin{split} \Gamma_{\nu_R}(T) &= n_{\nu_R}(T) \sum_f \langle \sigma_f(\nu_R \bar{\nu}_R \to f\bar{f}) v \rangle \\ &= \sum_f \frac{g_{\nu_R}^2}{n_{\nu_R}} \int \frac{d^3p}{(2\pi)^3} \frac{d^3q}{(2\pi)^3} f_{\nu_R}(p) f_{\nu_R}(q) \sigma_f(s) (1 - \cos\theta), \end{split}$$

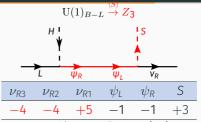
$$\begin{split} s = &2pq(1-\cos\theta), & f_{\nu_R}(k) = &1/(e^{k/T}+1) \\ n_{\nu_R}(T) = &g_{\nu_R} \int \frac{d^3k}{(2\pi)^3} f_{\nu_R}(k), & \text{with } g_{\nu_R} = &2 \\ \sigma_f(s) \simeq &\frac{N_C^f(Q_{BL}^f)^2 Q^2 s}{12\pi} \left(\frac{g'}{M_{Pl}}\right)^4, & \text{In the limit } M_{Z'}^2 \gg s. \end{split}$$

with three right-handed neutrinos, the Hubble parameter is

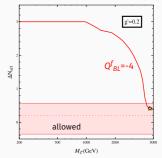
$$H(T) = \sqrt{\frac{4\pi^3 G_N \left[g(T) + 21/4\right]}{45}} T^2.$$

The right-handed neutrinos decouple when

$$\Gamma_{\nu_R}(T_{\text{dec}}^{\nu_R}) = H(T_{\text{dec}}^{\nu_R}).$$



E. Ma, R. Srivastava: arXiv:1411.5042 [PLB]

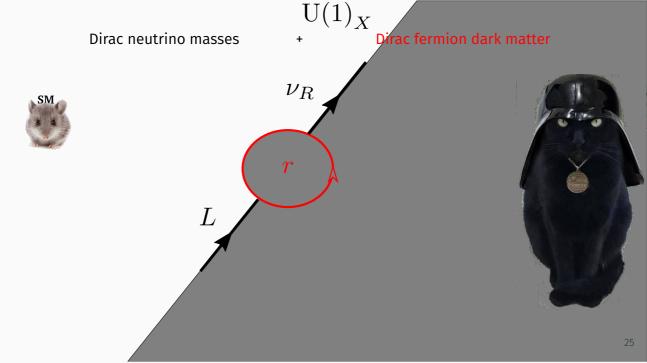


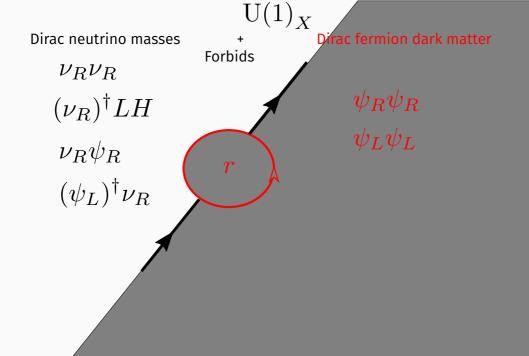
Z.-L. Han, W. Wang: arXiv:1805.02025 [EJPC]

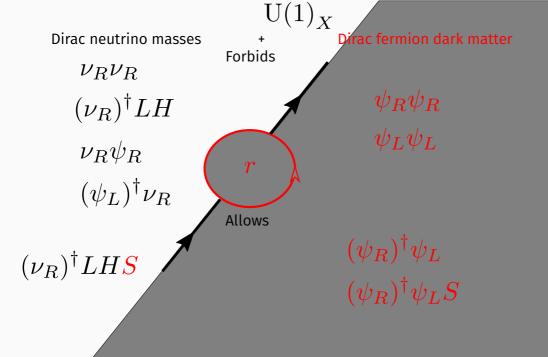
(also: Planck 1807.06209, Riess et al 1903.07603)

One-loop realization of \mathcal{L}_{5-D} with

total L







Dirac neutrino masses
$$\nu_R\nu_R$$

$$(\nu_R)^\dagger LH$$

$$\nu_R\psi_R$$

$$(\nu_R)^\dagger \nu_R$$

$$(\nu_R)^\dagger \nu_R$$

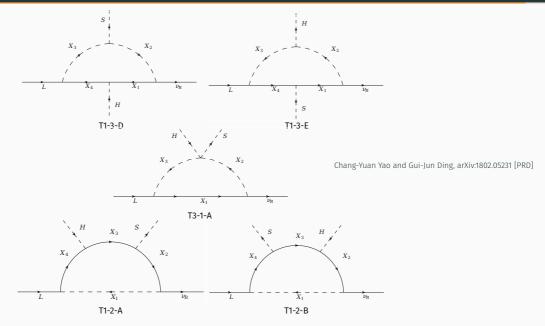
$$(\nu_R)^\dagger \nu_R$$

$$(\nu_R)^\dagger \nu_R$$
 Allows
$$(\nu_R)^\dagger \psi_L$$

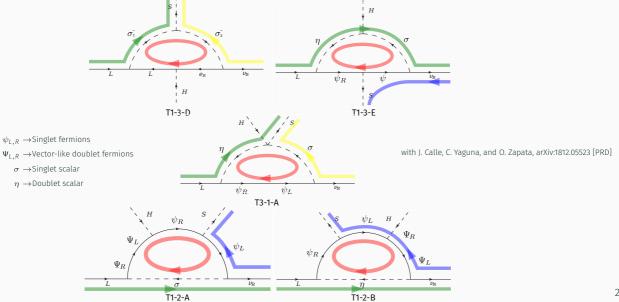
$$(\nu_R)^\dagger \psi_L$$

$$(\nu_R)^\dagger \psi_L S$$

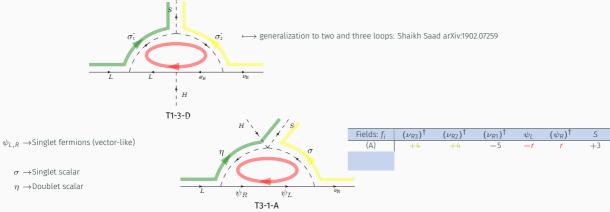
One loop topologies $U(1)_{B-L} \oplus Z_2 \oplus Z_2$



One loop topologies $U(1)_{B-L}$ only!



One loop topologies $U(1)_{B-L}$ only! with J. Calle, C. Yaguna, and O. Zapata, arXiv:1812.05523 [PRD]

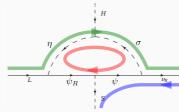


Anomaly cancellation conditions

$$\sum_{i} f_{i} = 3$$

$$\sum_{i} f_{i}^{3} = 3$$

One loop topologies $U(1)_{B-L}$ only! with J. Calle, C. Yaguna, and O. Zapata, arXiv:1812.05523 [PRD]



T1-3-E

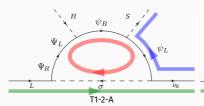
Fields: fi	$(\nu_{R3})^{\dagger}$	$(\nu_{R2})^{\dagger}$	$(\nu_{R1})^{\dagger}$	ψ_{L}	$(\psi_R)^\dagger$	S
(A)	+4	+4	-5	-r	r	+3
(B)	+ = 5	+ = 5	+ 2/5	7 _ 5	$-\frac{10}{5}$	+ - 5

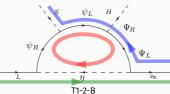
 $\psi_{L,R} \rightarrow \text{Singlet fermions (quiral)}$

 $\Psi_{L,R}
ightarrow$ Vector-like doublet fermions

 $\sigma \to Singlet scalar$

 $\eta \rightarrow Doublet scalar$





Anomaly cancellation conditions

$$\sum_{i} f_{i} = 3$$

$$\sum_{i} f_{i}^{3} = 3$$

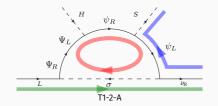
$$\int_{i}^{3} =3$$

$SD^3M+\sigma_i~(i=1,2)$ with J. Calle, C. Yaguna, and O. Zapata, arXiv:1812.05523 [PRD]

 $\psi_{L,R} \to \text{Singlet fermions (quiral)}$

 $\Psi_{L,R} o$ Vector-like doublet fermions : 10/5

 $\sigma \to \text{Singlet scalar}: 15/5$



Fields: fi	$(\nu_{R3})^{\dagger}$	$(\nu_{R2})^{\dagger}$	$(\nu_{R1})^{\dagger}$	ψ_{L}	$(\psi_R)^\dagger$	S
(A)	+4	+4	-5	-r	r	+3
(B)	+ 8/5	+ 8 - 5	$+\frac{2}{5}$	7 _ 5	$-\frac{10}{5}$	$+\frac{3}{5}$

Anomaly cancellation conditions

$$\sum_{i} f_{i} = 3$$

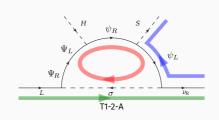
$$\sum_{i} f_{i}^{3} = 3$$

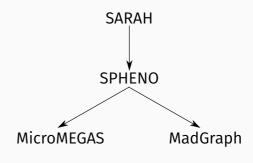
$SD^{3}M+\sigma_{i}$ (*i* = 1, 2)

$$M_{\psi} = h_1 \langle S \rangle$$
, $y_2 = 0$:

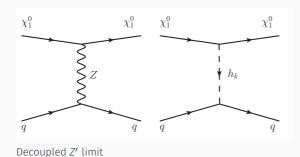
$$\mathcal{L} = \mathcal{L}_{\text{SD}^{3}\text{M}} + h_{3}^{ia}\widetilde{(\Psi_{R})} \cdot L_{i} \, \sigma_{a} + h_{2}^{\beta a} \left(\nu_{R\beta}\right)^{\dagger} \psi_{L} \, \sigma_{a}^{*} - V(\sigma_{a}, S, H) \, . \label{eq:loss_loss}$$

with A.F Rivera, W. Tangarife, arXiv:19nn.nnnnn





Spin independent (SI) direct detection cross section



 10^{-4} 10-8 $\begin{array}{c}
10^{-12} \\
\text{qd} \\
10^{-16}
\end{array}$ $\begin{array}{c}
(IS) \\
10^{-20}
\end{array}$ Vector S 10-24 Scalar SI XENON1T 10-28 PandaX 10-32 10^{1} 10³ $m_{\chi_1^0}(\text{GeV})$

Vector SI (blue points) and scalar SI (green points)

Conclusions

A single U(1) symmetry to explain both the smallnes of Dirac neutrino masses and the stability of Dirac fermion dark matter

Conclusions'

A single U(1) symmetry to explain both the smallnes of Dirac neutrino masses and the stability of Dirac fermion dark matter

Dirac neutrino masses and DM

- Spontaneously broken $U(1)_{B-L}$ generates a radiative Dirac neutrino masses
- A remnant symmetry makes the lightest field circulating the loop stable and good dark matter candidate.
- For T1-2-A: Either Singet Doublet Dirac Dark Matter or Singlet Scalar Dark Matter with extra scalar and vector portal
- · With relaxed direct detection constraints

