

## Multicomponent fermionic dark matter and dark baryogenesis

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Local Abelian extensions of the SM

#### Anomaly cancellation of a gauge $U(1)_x$ extension

Any *universal* local Abelian extension of the Standard Model can be reduced to a set of integers which must satisfy the gravitational anomaly,  $[SO(1,3)]^2 U(1)_Y$ , and the cubic anomaly,  $[U(1)_X]^3$  conditions:

$$\sum_{\alpha=1}^{N} z_{\alpha} = 0, \qquad \sum_{\alpha=1}^{N} z_{\alpha}^{3} = 0, \qquad (1)$$

• From a list of N-2 integers, e.g., for N even

$$\mathbf{q} = [l_1, l_2, \cdots, l_n, k_1, k_2, \cdots, k_n], \qquad n = (N-2)/2.$$
 (2)

in the range [-m, m], build two vector-like solutions of N integers,

$$\mathbf{x} = [l_1, k_1, \dots k_n, -l_1, -k_1, \dots - k_n,]$$
  $\mathbf{y} = [0, 0, l_1, \dots l_n, -l_1, \dots - l_n]$  (3)

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• Obtain a (some times) non vector-like solution with  $z_{max} = 2m$ 

$$\mathbf{z} = \mathbf{x} \oplus \mathbf{y} = \left(\sum_{i=1}^{N} x_i y_i^2\right) \mathbf{x} + \left(\sum_{i=1}^{N} x_i^2 y_i\right) \mathbf{y}, \tag{4}$$

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in the range [-m, m], build two vector-like solutions of N integers,

$$\mathbf{x} = [l_1, \mathbf{k_1}, \dots \mathbf{k_n}, -l_1, -\mathbf{k_1}, \dots - \mathbf{k_n},]$$
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The parameter space to be explored with  $z_{\text{max}}=20~(m=10)$  has  $96\,153$  non vector-like solutions

# of 
$$\mathbf{q}$$
 lists =  $(2m+1)^{N-2} = \begin{cases} 9261 \to 3 & N=5 \\ 194841 \to 38 & N=6 \\ \vdots & \vdots \\ 1.6 \times 10^{13} \to 65910 & N=12 \end{cases}$ , instead  $10^{19}$ 

• From a list of N-2 integers, e.g., for N even

$$\mathbf{q} = [l_1 = 2, l_2 = 3, k_1 = -1, k_2 = -3], \qquad n = 2.$$
 (2)

in the range [-3,3], build two vector-like solutions of 6 integers,

$$x = [2, -1, -3, -2, 1, 3,]$$
  $y = [0, 0, 2, \dots 3, -2, \dots -3]$  (3)

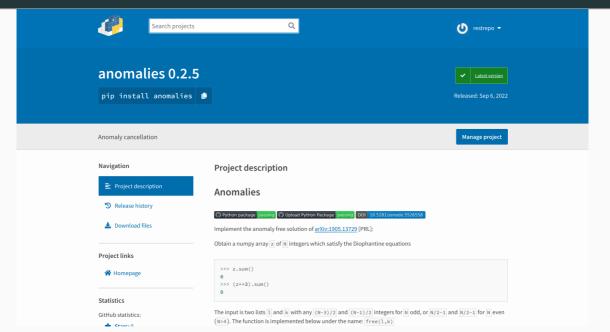
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# https://pypi.org/project/anomalies/







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# Minimal secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_{D}$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (6)

 $96\,153\rightarrow5\,196$ 

multi-component DM  $({\it N}=8,12) 
ightarrow 142$  with three Dirac-fermion DM

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96 153 ightarrow 5 196 multi-component DM ( $\mathit{N}=8,12$ ) ightarrow 142 with three Dirac-fermion DM

$$z = [1, -2, -2, 4, 5, -7, -7, 8] \rightarrow \phi = 9 \rightarrow [(1, 8), (-2, -7), (4, 5)]$$
 (7)

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 (7)

$$\mathcal{L} \subset h_{(1,8)} \psi_1 \psi_8 \phi^* \phi^{(*)} + \sum_{a,b=1}^2 h_{(-2a,-7b)} \psi_2 \psi_{-7} \phi + h_{(4,5)} \psi_4 \psi_5 \phi^* \phi^{(*)} + \text{h.c.}$$
 (8)

# Simplest secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_{D}$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
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$$z = [1, 2, 2, 4, -5, -5, -7, 8] \rightarrow \phi = 3 \rightarrow [(1, 2), (2, -5), (-5, 8), (4, -7)]$$
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# Simplest secluded model with SM-singlet massive chiral fermions from SSB: $U(1)_{D}$

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (9)

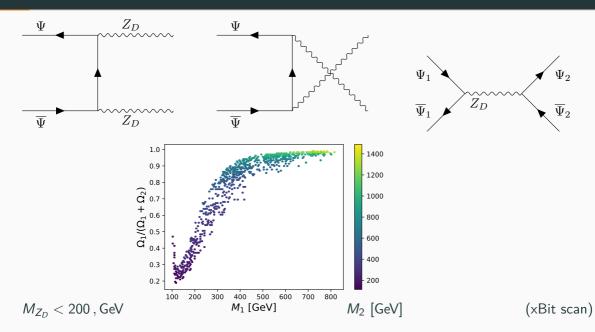
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$$\mathbf{z} = [1, 2, 2, 4, -5, -5, -7, 8] \to \phi = 3 \to [(1, 2), (2, -5), (-5, 8), (4, -7)] \tag{10}$$

$$1 \qquad 2 \qquad 2 \qquad -5 \qquad -5 \qquad 8$$

$$1 \qquad \begin{bmatrix}
0 & h_{(1,2)} & h'_{(1,2)} & 0 & 0 & 0 \\
h_{(1,2)} & 0 & 0 & h_{(2,-5)} & h_{(2,-5)} & 0 \\
h'_{(1,2)} & 0 & 0 & 0 & 0 & 0 \\
0 & h_{(2,-5)} & 0 & 0 & 0 & h_{(-5,8)} \\
-5 & 0 & h_{(2,-5)} & 0 & 0 & 0 & h'_{(-5,8)} \\
0 & 0 & 0 & h_{(-5,8)} & h'_{(-5,8)} & 0
\end{bmatrix}$$

$$\Psi \phi^{(*)} + h_{(4,-7)} \psi_4 \psi_{-7} \phi^*$$

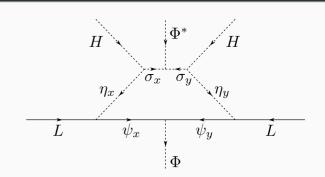


# Majorana neutrino masses and mixings

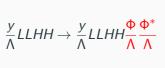
$$\frac{y}{\Lambda}$$
LLHH

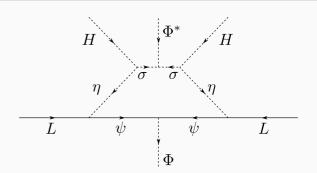
## Scotogenic Majorana neutrino masses and mixings

$$\frac{y}{\Lambda}LLHH \to \frac{y}{\Lambda}LLHH\frac{\Phi}{\Lambda}\frac{\Phi^*}{\Lambda}$$



## Scotogenic Majorana neutrino masses and mixings





Already found by Chi-Fong Wong in arXiv:2008.08573 (subset with  $N \leq 9$  and  $z_{\text{max}} \leq 10$ )

$$z = [\underbrace{1,1}_{y_0}, 2, 3, -4, -4, -5, 6] \rightarrow \phi = 2 \rightarrow [(1,1)_a, (2,-4), (4,-6), (4,-7)]$$
 (11)

Additional conditions to reduce

multiplicity

Decrease the number of charges to be assigned to dark matter particles,  $\psi_i$  below

$$[\chi_1,\chi_2,\cdots,\psi_1,\psi_2,\cdots,\psi_{N'}]$$

#### Secluded case:

$$[\nu, \nu, (\nu), \psi_1, \psi_2, \cdots, \psi_{N'}]$$

$$\chi_1 \to \nu_{R1}, \cdots, \chi_{N_{\nu}} \to \nu_{RN_{\nu}}, \qquad 2 \le N_{\nu} \le 3,$$

(12)

$$\mathcal{L}_{\mathrm{eff}} = h_{\nu}^{\alpha i} (\nu_{R\alpha})^{\dagger} \epsilon_{ab} L_{i}^{a} H^{b} \left( \frac{\Phi^{*}}{\Lambda} \right)^{\delta} + \mathrm{H.c.}, \qquad \mathrm{with} \ i = 1, 2, 3 \,,$$

 $\Phi$  is the complex singlet scalar responsible for the SSB of the anomaly-free gauge symmetry and give mass to all  $\psi_i$ 

$$\phi = -\frac{\nu}{\delta} \,, \tag{13}$$

(12)

Decrease the number of charges to be assigned to dark matter particles,  $\psi_i$  below

$$[\chi_1,\chi_2,\cdots,\psi_1,\psi_2,\cdots,\psi_{N'}]$$

Secluded case:

Active case:

$$[\nu, \nu, (\nu), \psi_1, \psi_2, \cdots, \psi_{N'}]$$

$$[\nu, \nu, (\nu), m, m, m, \psi_1, \psi_2, \cdots, \psi_{N'}]$$

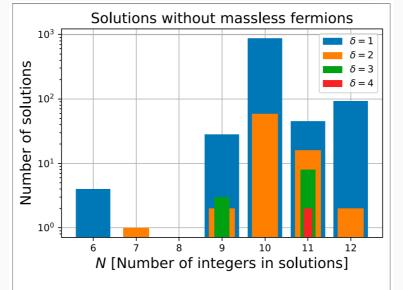
$$\chi_1 \to \nu_{R1}, \cdots, \chi_{N_{\nu}} \to \nu_{RN_{\nu}}, \qquad 2$$

$$\chi_1 \rightarrow \nu_{R1}, \cdots, \chi_{N_{\nu}} \rightarrow \nu_{RN_{\nu}}, \qquad 2 \leq N_{\nu} \leq 3, \quad X(L_i) = -L, \quad X(H) = h \qquad \rightarrow m = L - h$$

$$\mathcal{L}_{\mathrm{eff}} = h_{\nu}^{\alpha i} \left( \nu_{R\alpha} \right)^{\dagger} \epsilon_{ab} \, L_{i}^{a} \, H^{b} \left( rac{\Phi^{*}}{\Lambda} 
ight)^{\delta} + \mathrm{H.c.}, \qquad \mathrm{with} \, \, i = 1, 2, 3 \, ,$$

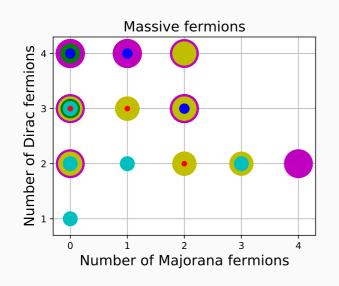
Φ is the complex singlet scalar responsible for the SSB of the anomaly-free gauge symmetry and give mass to all  $\psi_i$   $\rightarrow$  [-4, -4, 1, 1, 1, 5]

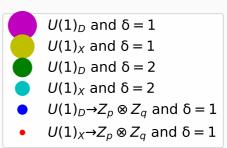
$$\phi = -\frac{(\nu + m)}{\delta},\tag{13}$$



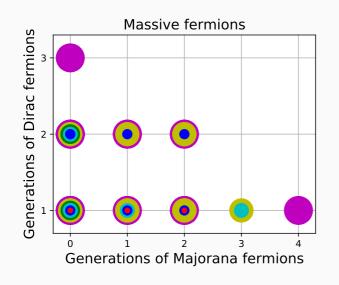
**FIGURE 1** Distribution of solutions with N integers to the Diophantine **Eq. 1** which allow the effective Dirac neutrino mass operator at  $d = (4 + \delta)$  for at least two right-handed neutrinos and have non-vanishing Dirac o Majorana masses for the other SM-singlet chiral fermions in the solution.

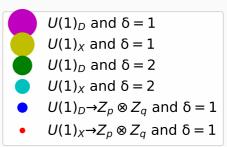
#### Multi-component dark matter





#### Multi-flavor dark matter





## $U(1)_X$ selection with Dirac-fermionic DM

• Active symmetry m = 3

$$(-5, -5, 3, 3, 3, -7, 8)$$

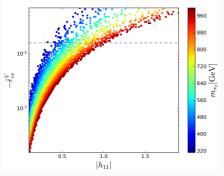
#### $U(1)_X$ selection with Dirac-fermionic DM

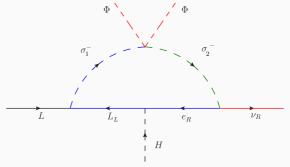
- Active symmetry m = 3
- Effective neutrino mass  $\delta=2 \rightarrow \nu=-5$ :

$$(-5, -5, 3, 3, 3, -7, 8)$$

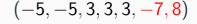
- Active symmetry m = 3
- Effective neutrino mass  $\delta = 2 \rightarrow \nu = -5$ :
- Active symmetry:  $m=3 \rightarrow \phi = -(\nu + m)/\delta = 1$

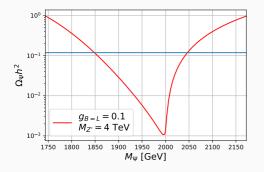
(-5, -5, 3, 3, 3, -7, 8)

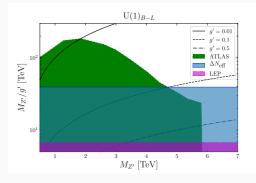




- Active symmetry m = 3
- Effective neutrino mass  $\delta = 2 \rightarrow \nu = -5$ :
- Active symmetry:  $m=3 \rightarrow \phi = -(\nu+m)/\delta = 1$
- Dirac-fermionic DM:  $(\psi_L)^{\dagger} \psi_R'' \Phi^* \rightarrow z_6 = -7, z_7 = 8$

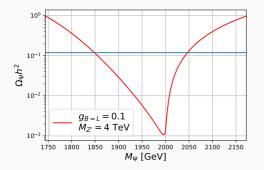


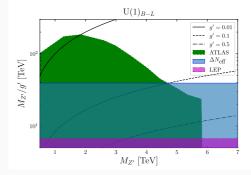




(-5, -5, 3, 3, 3, -7, 8)

- Active symmetry m = 3
- Effective neutrino mass  $\delta = 2 \rightarrow \nu = -5$ :
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# Beyond SM-fermion singlets

# Standard model extended with $U(1)_{\mathcal{X}=X \text{ or } D}$ gauge symmetry

Fields	$SU(2)_L$	$U(1)_Y$	$U(1)_{\mathcal{X}=D \text{ or } X}$
$Q_i^{\dagger}$	2	-1/6	Q
$d_{Ri}$	1	-1/2	d
$u_{Ri}$	1	+2/3	и
$L_i^{\dagger}$	2	+1/2	L
$e_{Ri}$	1	-1	e
Н	2	1/2	h
$\chi_{\alpha}$	1	0	$z_{\alpha}$

Ф	1	0	$\phi$

Table 1: LHC: hadronic production and dileptonic decay

$$i = 1, 2, 3, \ \alpha = 1, 2, \dots, N'$$

# Standard model extended with $U(1)_{\mathcal{X}=L \text{ or } \mathbf{B}}$ gauge symmetry

Fields	$SU(2)_L$	$U(1)_Y$	$U(1)_{\mathcal{X}=\mathbf{B} \text{ or } \mathbf{L}}$
$Q_i^{\dagger}$	2	-1/6	Q
$d_{Ri}$	1	-1/2	d
$u_{Ri}$	1	+2/3	и
$L_i^{\dagger}$	2	+1/2	L
$e_{Ri}$	1	-1	e
Н	2	1/2	h = 0
$\chi_{\alpha}$	1	0	$z_{\alpha}$
$(L'_L)^{\dagger}$	2	1/2	-x'
$L_R^{\prime\prime}$	2	-1/2	x''
$e_R'$	1	-1	×
$(e_L^{\prime\prime})^\dagger$	1	1	-x''
Ф	1	0	$\phi$
5	1	0	S

**Table 1:** minimal set of new fermion content: L = e = 0 for  $\mathcal{X} = B$ . Or Q = u = d = 0 for  $\mathcal{X} = L$ .  $i = 1, 2, 3, \alpha = 1, 2, \dots, N'$ 

# Anomaly cancellation: $\mathcal{X} = L$ or B: beyond SM-singlet fermions

The anomaly-cancellation conditions on  $[SU(3)_c]^2 U(1)_X$ ,  $[SU(2)_L]^2 U(1)_X$ ,  $[U(1)_Y]^2 U(1)_X$ , allow us to express three of the X-charges in terms of the others

$$\mathbf{u} = -\mathbf{e} - \frac{2}{3}\mathbf{L} - \frac{1}{9}(x' - x''), \quad \mathbf{d} = \mathbf{e} + \frac{4}{3}\mathbf{L} - \frac{1}{9}(x' - x''), \quad \mathbf{Q} = -\frac{1}{3}\mathbf{L} + \frac{1}{9}(x' - x''), \quad (14)$$

while the  $[U(1)_X]^2 U(1)_Y$  anomaly condition reduces to

$$(e+L)(x'-x'')=0.$$
 (15)

- Previously: x' = x''
- We choose instead (h = 0):

$$e = -L, (16)$$

so that (L is still a free parameter)

$$Q = -u = -d = -\frac{1}{3}L + \frac{1}{9}(x' - x''). \tag{17}$$

#### Anomaly cancellation: $\mathcal{X} = L$ or **B**

The gravitational anomaly,  $[SO(1,3)]^2 U(1)_Y$ , and the cubic anomaly,  $[U(1)_X]^3$ , can be written as the following system of Diophantine equations, respectively,

$$\sum_{\alpha=1}^{N} z_{\alpha} = 0, \qquad \sum_{\alpha=1}^{N} z_{\alpha}^{3} = 0, \qquad (18)$$

where

$$z_1 = -x',$$
  $z_2 = x'',$   $z_{2+i} = L, \quad i = 1, 2, 3$  (19)

 $\rightarrow$ 

$$9Q = -\sum_{\alpha=1}^{5} z_{\alpha} = -x' + x'' + L + L + L, \qquad (20)$$

$$L = 0 \rightarrow U(1)_B$$
 but  $Q = 0 \rightarrow U(1)_L$ 

# $U(1)_B$ selection: Neutrinos, dark matter and baryogenesis

• 
$$L = 0$$

$$(5,5,-3,-2,1,-6)$$

# $U(1)_B$ selection: Neutrinos, dark matter and baryogenesis

- L=0
- Effective Dirac neutrino masses:  $\phi = -\nu = -5$

$$(5, 5, -3, -2, 1, -6)$$

# $U(1)_B$ selection: Neutrinos, dark matter and baryogenesis

- L = 0
- Effective Dirac neutrino masses:  $\phi = -\nu = -5$
- Electroweak-scale vector-like fermions:

$$(L'_L)^{\dagger} L''_R \Phi^* \to x' = -1, \ x'' = 6$$

$$(5,5,-3,-2,1,-6)$$

## $U(1)_{B}$ selection: Neutrinos, dark matter and baryogenesis

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# $U(1)_{\it B}$ selection: Neutrinos, dark matter and baryogenesis

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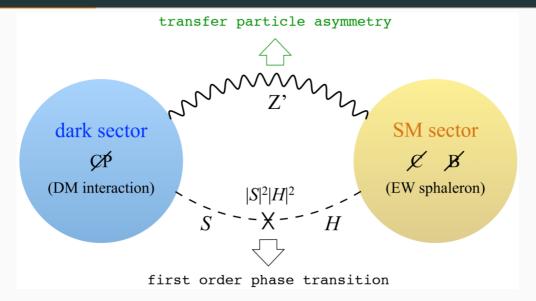
$$(L'_L)^{\dagger} L''_R \Phi^* \to x' = -1, \ x'' = 6$$

• Dirac-fermionic DM:  $(\chi_L)^{\dagger} \chi_R'' \Phi^* \rightarrow z_3 = -3, \ z_4 = -2$ 

959 solutions

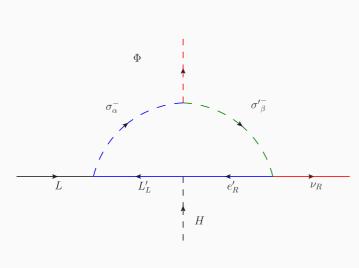
$$(5,5,-3,-2,1,-6)$$

### Dark sector baryogenesis



### Gauge Baryon number scotogenic realization: arXiv:2205.05762 [PRD]

with Andrés Rivera (UdeA) and Walter Tangarife (Loyola U.)

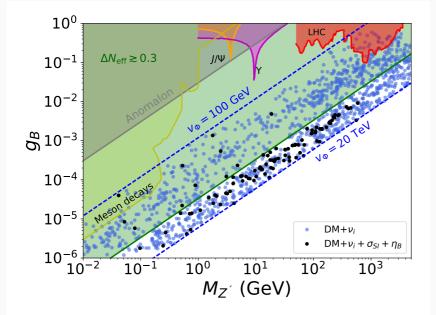


0.)				
	Field	$SU(2)_L$	$U(1)_Y$	$U(1)_B$
	u <sub>Ri</sub>	1	2/3	u = 1/3
	$d_{Ri}$	1	-1/3	d = 1/3
	$(Q_i)^{\dagger}$	2	-1/6	Q = -1/3
	$(L_i)^{\dagger}$	2	1/2	L=0
	$e_R$	1	-1	e = 0
	$(L'_L)^{\dagger}$	2	1/2	-x' = -3/5
	$e'_R$	1	-1	x' = 3/5
	$L_R^{\prime\prime}$	2	-1/2	x'' = 18/5
	$\left(e_L^{\prime\prime} ight)^\dagger$	1	1	-x'' = -18/5
	$ u_{R,1}$	1	0	-3
	$\nu_{R,2}$	1	0	-3
	$\chi_R$	1	0	6/5
	$(\chi_L)^{\dagger}$	1	0	9/5
	Н	2	1/2	0
	S	1	0	3
	Φ	1	0	3
	$\sigma_{lpha}^-$	1	1	3/5
	$\sigma'_{\alpha}^{-}$	1	-1	-12/5

arXiv:2205.05762 [PRD] https://github.com/anferivera/DarkBariogenesis

- $SARAH \rightarrow SPheno \rightarrow MicroMegas$
- $\eta_B$  calculation code
- Python notebook with the scan

## Black points: Dirac neutrinos with proper DM and baryon assymetry



#### **Conclusions**

A methodology was designed to find all the *universal* gauge Abelian extensions of the standard model:

All of the extensions can be reformulated as the solution of

$$\sum_{\alpha=1}^{N} z_{\alpha} = 0, \qquad \sum_{\alpha=1}^{N} z_{\alpha}^{3} = 0,$$

which we thoroughly scan in an efficient way until N=12 and  $|z_{\rm max}|=20$ 

Once the physical conditions are stablished, the full set of self-consistent models are found from a simple data-analysis procedure, providing enough freedom to solve several phenomenological problems.

#### **Baryogenesis**

CP violation occurs in the dark sector and is transmitted to SM sector by the new Z' gauge boson.

- High scale fields:  $\Phi$ ,  $(\langle \Phi \rangle \to L'_L, L''_R, e'_L, e''_R$ : EW-scale vector-like anomalons)
- Electroweak scale (EW) fields:  $Z'_{\mu}, S, \chi_L, \chi_R$
- CP-violation

$$\mathcal{L}_{\mathsf{Dirac}\;\mathsf{DM}} = h(\chi_L)^{\dagger} \chi_R \Phi^* + y(\chi_L)^{\dagger} \chi_R S^* + \mathsf{h.c.}, \qquad y \in \mathbb{C}$$
$$\supset \left( m_{\chi} + |y| \, \mathrm{e}^{\mathrm{i}\theta} \, |S| \right) (\chi_L)^{\dagger} \chi_R + \mathsf{h.c.}.$$

CP-violation Portal

$$\mathcal{L}_{\text{anomalous}} \supset g' Z'_{\mu} \left[ 3\bar{\chi}_{L} \gamma^{\mu} \chi_{L} - 2\bar{\chi}_{R} \gamma^{\mu} \chi_{R} + \bar{Q}_{i} \gamma^{\mu} Q_{i} + \bar{q}_{Ri} \gamma^{\mu} q_{Ri} \right]$$

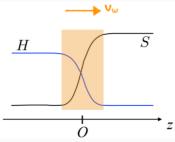
Strong electroweak phase transition (EWPT) portal

$$\mathcal{L}_{\mathsf{first}\ \mathsf{order}\ \mathsf{EWPT}} \supset -\lambda_{\mathsf{SH}} H^\dagger H S^* S$$
 .

$$h = H/\sqrt{2}$$
,  $s = |S|$  with vevs:  $v(T)$  and  $w(T)$  such that  $v(T_c) = w(T_c)$ 

$$V_T(h,s) = \frac{\lambda_H v_c^4}{4} \left( \frac{h^2}{v_c^2} + \frac{s^2}{w_c^2} - 1 \right)^2 + \frac{\lambda_H v_c^2}{m_{s,c}^2 w_{0,c}^4} h^2 s^2 + (T^2 - T_c^2) (c_h h^2 + c_s s^2),$$
 (21)

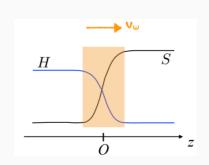
$$c_h = \frac{1}{48} \left( 9g_2^2 + 3g_1^2 + 12y_t^2 + 24\lambda_H + \lambda_{HS} \right) , \quad c_s = \frac{1}{12} \left( 3\lambda_S + 2\lambda_{HS} \right) .$$
 (22)

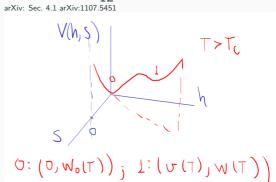


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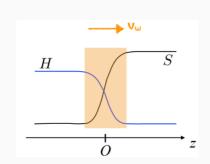


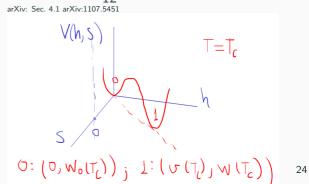


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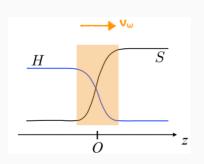


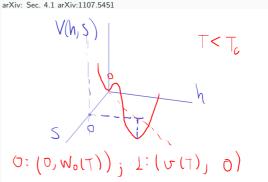


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#### CP assymetry generation i

Using the thin wall approximantion for the nucleation bubbles, we use the ansatz in which the space dependence of the fields is given by

$$h(z) = \frac{1}{2}v(T_n)(1-\tanh(z/L_w)), \qquad s(z) = \frac{1}{2}w_0(T_n)(1+\tanh(z/L_w)),$$

where z is the direction normal to the wall and  $L_w$  is the wall width.

The nucleation temperature,  $T_n$ , is defined by the condition

$$\exp(-S_3/T_n) = \frac{3}{4\pi} \left(\frac{H(T_n)}{T_n}\right)^4 \left(\frac{2\pi T_n}{S_3}\right)^{\frac{3}{2}},$$

where  $S_3$  is the Euclidean action of the bubble and H(T) is the Hubble rate.

#### Boltzmann equation i

$$egin{align} \xi_i(z) &\equiv \mu_i(z)/\mathcal{T} = \left.6\left(n_i - \overline{n}_i\right)/\mathcal{T}^3, 
ight. \ &\left. -D_L \xi_{\chi_L}'' - v_w \xi_{\chi_L}' + \Gamma_L (\xi_{\chi_L} - \xi_{\chi_R}) \right. = S_{\mathcal{R}}, \end{aligned}$$

where  $D_L$  is the diffusion constant for  $\chi_L$ , which is related to the scattering rate  $\Gamma_L$  by

$$D_{L} = \frac{3x+2}{x^{2}+3x+2} \frac{1}{3\Gamma_{L}}, \qquad x \equiv m_{\chi}/T$$
 (23)

and

$$S_{\mathcal{CP}} = -\frac{\lambda}{2} \frac{v_w D_L}{\frac{3x+2}{x^2+3x+2} T} \frac{(1-x)e^{-x} + x^2 E_1(x)}{4m_\chi^2 K_2(x)} \frac{m_\chi w_0(T_n) \lambda \left(-2 + \cosh\left(\frac{2z}{L_w}\right)\right) \sin\theta}{L_w^3 \cosh^4\left(\frac{z}{L_w}\right)}, \qquad (24)$$

where  $v_w$  is the wall's velocity  $E_1(x)$  is the error function and  $K_2(x)$  is the modified Bessel function of the second kind.  $\mathbf{y} = \lambda e^{i\theta - i\pi/2}$ 

### Transfer DM assymetry to SM quarks

The chiral particle give rise to a non-zero  $U(1)_B$  charge density in the proximity of the wall. This results in a Z' background that couples to the SM fields with  $U(1)_B$  charge,

$$\langle Z_0'(z) \rangle = \frac{g_B (q_{\chi_L} - q_{\chi_R}) T_n^3}{6 M_{Z'}} \int_{-\infty}^{\infty} dz_1 \, \xi_{\chi_L}(z_1) \, e^{-M_{Z'}|z-z_1|} \,,$$

which generates a chemical potential for the SM quarks,

$$\mu_Q(z) = \mu_{d_R,u_R}(z) = 3 \times \frac{5}{9} \times g_B \langle Z'_0(z) \rangle.$$

This chemical potential sources a thermal-equilibrium asymmetry in the quarks,

$$\Delta n_Q^{\text{EQ}}(z) \sim T_n^2 \mu_Q(z).$$

From [1]

If the Z' is sufficiently light, it mediates a long range force that extends into the region outside the bubble wall with unbroken electroweak symmetry.

### Finally, the baryon-number asymmetry is then given by

$$n_B \,=\, rac{\Gamma_{
m sph}}{v_w} \int_0^\infty {
m d}\,z\, n_Q^{
m EQ}(z) \, \exp\left(-rac{\Gamma_{
m sph}}{v_w}\,z
ight) \,,$$

where  $\Gamma_{\rm sph}$  is the sphaleron rate. The baryon-to-photon-number ratio is then obtained by

$$\eta_B = \frac{n_B}{s(T_n)}, \quad s(T) \equiv \frac{2\pi^2}{45} g_{*S}(T) T^3,$$

where  $g_{*S}(T)$  is the effective number of relativistic degrees of freedom.

Our goal is to find what regions of the parameter space yield

$$0.82 \times 10^{-10} < \eta_B < 0.92 \times 10^{-10} \,. \tag{25}$$