Dirac dark matter, neutrino masses,



and dark baryogenesis

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Focus on

arXiv:2205.05762 [PRD]

In collaboration with

Andrés Rivera (UdeA), Walter Tangarife (Loyola University Chicago)

Dark sectors









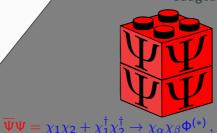


 $\mathcal{L} = -\frac{1}{4} V_{\mu\nu} V^{\mu\nu} + i \sum_{i} \chi_{i}^{\dagger} \mathcal{D} \chi_{i}$

$$\mathcal{L} = -\frac{1}{4}V_{\mu\nu}V^{\mu\nu} + i\sum_{i}\chi_{i}^{\dagger}\mathcal{D}_{\lambda}$$
$$-h(\chi_{1}\chi_{2}\Phi + \text{h.c})$$

Anomalons: SM-singlet Dirac fermion dark matter $m_{\Psi} = h\langle \Phi \rangle$

Gauged Symmetry: $\mathcal{X} \to B$: $q\overline{q} \to Z' \to \text{jets}$ Gauged Symmetry: $\mathcal{X} \to L$:



$$\alpha = 1,.$$

$$\alpha = 1, \dots N \rightarrow N > 4$$

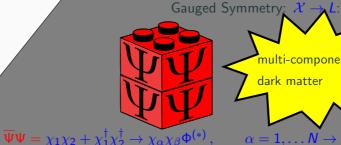


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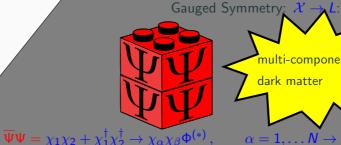


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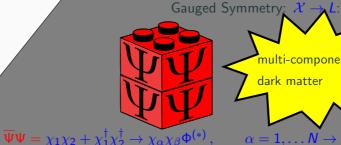
Local $U(1)_{\mathcal{X}}$ $\mathcal{L} = -\frac{1}{4}V_{\mu\nu}V^{\mu\nu} + i\sum_{i}\chi_{i}^{\dagger}\mathcal{D}\chi_{i}$

$$-y\left(\chi_1\chi_2S+\text{h.c}\right)$$

Anomalons: SM-singlet Dirac fermion

CP violation Yukawa y

Gauged Symmetry: $\mathcal{X} \to B$: $q\overline{q} \to Z' \to \text{jets}$



$$\alpha = 1, \dots N \rightarrow N > 4$$

Anomaly cancellation

Any local Abelian extension of the Standard Model can be reduced to a set of integers which must satisfy the gravitational anomaly, $[SO(1,3)]^2 U(1)_Y$, and the cubic anomaly, $[U(1)_X]^3$ conditions:

$$\sum_{\alpha=1}^{N} z_{\alpha} = 0, \qquad \sum_{\alpha=1}^{N} z_{\alpha}^{3} = 0, \qquad (1)$$

• From a list of N-2 integers, e.g., for N even

$$\mathbf{q} = [l_1, l_2, \cdots, l_n, k_1, k_2, \cdots, k_n], \qquad n = (N-2)/2.$$
 (2)

in the range [-m, m], build two vector-like solutions of N integers,

$$\mathbf{x} = [l_1, k_1, \dots k_n, -l_1, -k_1, \dots - k_n,]$$
 $\mathbf{y} = [0, 0, l_1, \dots l_n, -l_1, \dots - l_n]$ (3)

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• Obtain a (some times) non vector-like solution with $z_{\text{max}} = 2m$

$$\mathbf{z} = \mathbf{x} \oplus \mathbf{y} = \left(\sum_{i=1}^{N} x_i y_i^2\right) \mathbf{x} + \left(\sum_{i=1}^{N} x_i^2 y_i\right) \mathbf{y},$$
 (4)

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The parameter space to be explored with $z_{\rm max}=20~(m=10)$ has $96\,153$ non vector-like solutions

of
$$\mathbf{q}$$
 lists = $(2m+1)^{N-2}$ =
$$\begin{cases} 9261 \to 3 & N = 5 \\ 194841 \to 38 & N = 6 \\ \vdots & \vdots \\ 1.6 \times 10^{13} \to 65910 & N = 12 , \text{ instead } 10^{19} \end{cases}$$
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 lists = $(2m+1)^{N-2} = \begin{cases} 9261 \to 3 & N = 5 \to [1, 5, -7, -8, 9] \\ 194841 \to 38 & N = 6 \\ \vdots & \vdots \\ 1.6 \times 10^{13} \to 65910 & N = 12, \text{ instead } 10^{19} \end{cases}$ (5)

Simplest secluded model with SM-singlet massive chiral fermions from SSB

$$\mathcal{L} = i\psi_i^{\dagger} \mathcal{D} \psi_i - \frac{1}{4} V_{\mu\nu} V^{\prime\mu\nu} + \sum_{i < j} h_{ij} \psi_i \psi_j \phi^{(*)} + \text{h.c}$$
 (6)

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ightarrow 5\,196$ multi-component DM (N = 8,12) ightarrow 28 with two Dirac-fermion DM

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$$z = [1, 2, 2, 4, -5, -5, -7, 8] \rightarrow \phi = 3 \rightarrow [(1, 2), (2, -5), (-5, 8), (4, -7)]$$
 (7)

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$$\mathbf{z} = \begin{bmatrix} 1, 2, 2, 4, -5, -5, -7, 8 \end{bmatrix} \to \phi = 3 \to \begin{bmatrix} (1, 2), (2, -5), (-5, 8), (4, -7) \end{bmatrix}$$

$$1 \quad 2 \quad 2 \quad -5 \quad -5 \quad 8$$

$$1 \quad \begin{bmatrix} 0 & h_{(1,2)} & h_{(1,2)} & 0 & 0 & 0 \\ h_{(1,2)} & 0 & 0 & h_{(2,-5)} & h_{(2,-5)} & 0 \\ h_{(1,2)} & 0 & 0 & 0 & 0 & 0 \\ 0 & h_{(2,-5)} & 0 & 0 & 0 & h_{(-5,8)} \\ -5 & 0 & h_{(2,-5)} & 0 & 0 & 0 & h_{(-5,8)} \\ 0 & 0 & 0 & h_{(-5,8)} & h_{(-5,8)} & 0 \end{bmatrix} \Psi \phi^{(*)} + h_{(4,-7)} \psi_4 \psi_{-7} \phi^*$$

$$(8)$$

4

Standard model extended with $U(1)_{\mathcal{X}=L \text{ or } \mathbf{B}}$ gauge symmetry

Fields	$SU(2)_L$	$U(1)_Y$	$U(1)_{\mathcal{X}=\mathbf{B}}$ or \mathbf{L}
Q_i^{\dagger}	2	-1/6	Q
d_{Ri}	1	-1/2	d
u_{Ri}	1	+2/3	и
L_i^{\dagger}	2	+1/2	L
e_{Ri}	1	-1	e
Н	2	1/2	h = 0
χ_{α}	1	0	z_{α}
$(L'_L)^{\dagger}$	2	1/2	-x'
$L_R^{\prime\prime}$	2	-1/2	x''
e_R'	1	-1	x'
$\left(e_L^{\prime\prime} ight)^\dagger$	1	1	-x''
Ф	1	0	φ
S	1	0	5

Table 1: A minimal set of new fermion content: L = e = 0 for $\mathcal{X} = B$. Or Q = u = d = 0 for $\mathcal{X} = L$.

$$i = 1, 2, 3, \alpha = 1, 2, \dots, N'$$

Effective Dirac neutrino mass operator

$$\chi_1 \to \nu_{R1}, \cdots, \chi_{N_{\nu}} \to \nu_{RN_{\nu}}, \qquad 2 \le N_{\nu} \le 3,$$
 (9)

$$\mathcal{L}_{\mathrm{eff}} = h_{
u}^{lpha i} \; (
u_{Rlpha})^{\dagger} \; \epsilon_{\mathsf{a} \mathsf{b}} \; \mathsf{L}_{i}^{\mathsf{a}} \; \mathsf{H}^{\mathsf{b}} \left(rac{\Phi^{*}}{\Lambda}
ight)^{\delta} + \mathrm{H.c.}, \qquad \mathrm{with} \; i = 1, 2, 3 \,,$$

S is the complex singlet scalar responsible for the SSB of the anomaly-free gauge symmetry with D or X-charge

$$\phi = -(\nu + \mathbf{L})/\delta \,, \tag{10}$$

Anomaly cancellation I

The anomaly-cancellation conditions on $[SU(3)_c]^2 U(1)_X$, $[SU(2)_L]^2 U(1)_X$, $[U(1)_Y]^2 U(1)_X$, allow us to express three of the X-charges in terms of the others

$$\mathbf{u} = -\mathbf{e} - \frac{2}{3}\mathbf{L} - \frac{1}{9}(x' - x'') , \quad \mathbf{d} = \mathbf{e} + \frac{4}{3}\mathbf{L} - \frac{1}{9}(x' - x'') , \quad \mathbf{Q} = -\frac{1}{3}\mathbf{L} + \frac{1}{9}(x' - x'') ,$$
(11)

while the $[U(1)_X]^2 U(1)_Y$ anomaly condition reduces to

$$(e+L)(x'-x'')=0.$$
 (12)

- Previously: x' = x''
- We choose instead (h = 0):

$$e = -L, (13)$$

so that (*L* is still a free parameter)

$$Q = -u = -d = -\frac{1}{3}L + \frac{1}{9}(x' - x''). \tag{14}$$

If $B = 0 \rightarrow U(1)_L$

Anomaly cancellation II

The gravitational anomaly, $[SO(1,3)]^2 U(1)_Y$, and the cubic anomaly, $[U(1)_X]^3$, can be written as the following system of Diophantine equations, respectively,

$$\sum_{\alpha=1}^{N} z_{\alpha} = 0, \qquad \sum_{\alpha=1}^{N} z_{\alpha}^{3} = 0, \qquad (15)$$

where N = N' + 5 and

$$z_{N'+1} = -x',$$
 $z_{N'+2+i} = L, \quad i = 1, 2, 3$ (16)

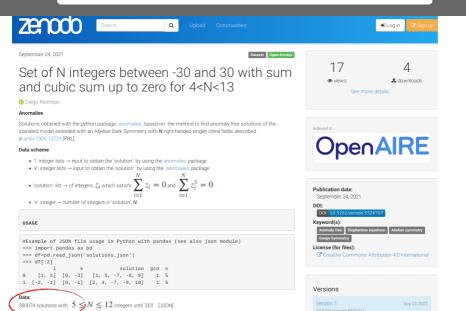
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$$9Q = -\sum_{\alpha=N'+1}^{N'+5} z_{\alpha} = -x' + x'' + L + L + L, \qquad (17)$$

$$Q = 0 \gg U(1)_L$$







$U(1)_{B}$ selection

•
$$L = 0$$

$$(5,5,-3,-2,1,-6)$$

$U(1)_{\mathbb{B}}$ selection

- L = 0
- Effective neutrino mass: $\phi = -\nu = -5$

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- Electroweak-scale vector-like fermions:

$$(L'_L)^{\dagger} L''_R \Phi^* \to x' = -1, \ x'' = 6$$

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$U(1)_{\mathbb{B}}$ selection

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• Dirac-fermionic DM: $(\chi_L)^{\dagger} \chi_R'' \Phi^* \rightarrow z_3 = -3, \ z_4 = -2$

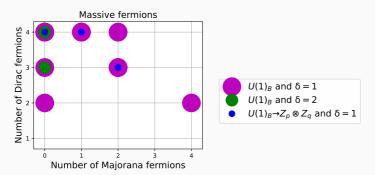
$$(5,5,-3,-2,1,-6)$$

$U(1)_{\mathbf{B}}$ selection

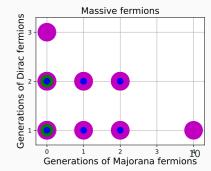
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- 959 solutions from \sim 400,000

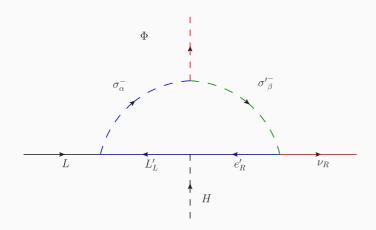


(5,5,-3,-2,1,-6)



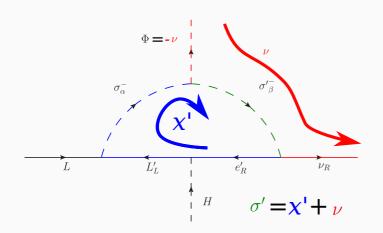
Scotogenic realization

Any realization which does not affect anomaly cancellation is allowed



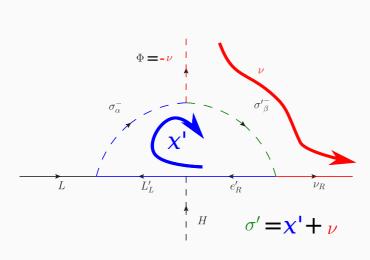
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i is allowed				
Field	$SU(2)_L$	$U(1)_Y$	$U(1)_B$	
u _{Ri}	1	2/3	u = 1/3	
d_{Ri}	1	-1/3	d = 1/3	
$(Q_i)^{\dagger}$	2	-1/6	Q = -1/3	
$(L_i)^{\dagger}$	2	1/2	L=0	
e_R	1	-1	e = 0	
$(L'_L)^{\dagger}$	2	1/2	-x' = -3/5	
e'_R	1	-1	x' = 3/5	
$L_R^{\prime\prime}$	2	-1/2	x'' = 18/5	
$\left(e_L^{\prime\prime} ight)^\dagger$	1	1	-x'' = -18/5	
$ u_{R,1}$	1	0	-3	
$\nu_{R,2}$	1	0	-3	
XR	1	0	6/5	
$(\chi_L)^{\dagger}$	1	0	9/5	
Н	2	1/2	0	
S	1	0	3	
Ф	1	0	3	
σ_{lpha}^-	1	1	3/5	
σ'_{α}^{-}	1	-1	-12/5	

Electroweak baryogenesis

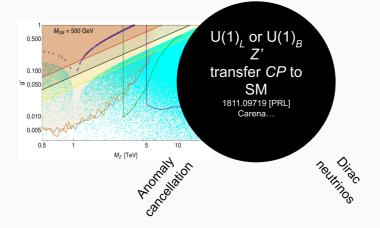
Problems

- Standard model (SM) $m_h \sim$ 40 GeV. \bigcirc
- Beyond the SM: Source of CP contains fields charged under SM
 - ightarrow too large electric dipole moments 😩

Dark sectors

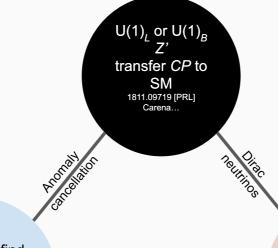
- Inert SM-singlet complex scalar field which acquires vev with temperature to have strong electroweak phase transition 😊
- CP violation (CPV) triggered in dark sectors through SM gauge singlets
 - → CPV Yukawa between SM-singlet complex scalar and SM-singlet quiral fermions \(\to\)





Anomalons:

DM



Method to find $\Sigma n=0$, $\Sigma n^3=0$ solutions 1905.13729 [PRL] Costa...

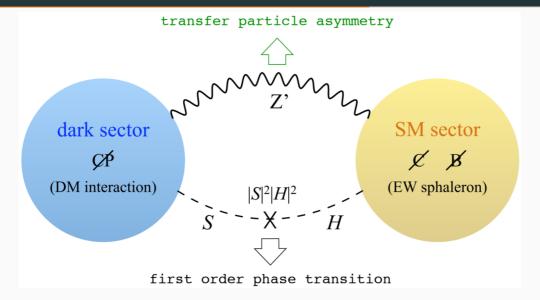
Anomalons:

Multicomponent DM

Scotogenic neutrino masses

hep-ph/0601225 [PRL→PRD] Ma

Dark sector baryogenesis



Baryogenesis

CP violation occurs in the dark sector and is transmitted to SM sector by the new Z' gauge boson.

- High scale fields: Φ , $(\langle \Phi \rangle \to L'_L, L''_R, e'_L, e''_R$: EW-scale vector-like anomalons)
- Electroweak scale (EW) fields: $Z'_{\mu}, S, \chi_L, \chi_R$
- CP-violation

$$\mathcal{L}_{\mathsf{Dirac}\;\mathsf{DM}} = h\left(\chi_{L}\right)^{\dagger} \chi_{R} \Phi^{*} + y\left(\chi_{L}\right)^{\dagger} \chi_{R} S^{*} + \mathsf{h.c}\,, \qquad y \in \mathbb{C}$$
$$\supset \left(m_{\chi} + |y| \,\mathsf{e}^{i\theta} \,|S|\right) \left(\chi_{L}\right)^{\dagger} \chi_{R} + \mathsf{h.c}\,.$$

• CP-violation Portal

$$\mathcal{L}_{\text{anomalous}} \supset g' Z'_{\mu} \left[3\bar{\chi}_L \gamma^{\mu} \chi_L - 2\bar{\chi}_R \gamma^{\mu} \chi_R + \bar{Q}_i \gamma^{\mu} Q_i + \bar{q}_{Ri} \gamma^{\mu} q_{Ri} \right]$$

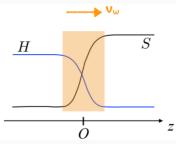
• Strong electroweak phase transition (EWPT) portal

$$\mathcal{L}_{\mathsf{first}}$$
 order EWPT $\supset -\lambda_{SH}H^{\dagger}HS^{*}S$.

$$h = H/\sqrt{2}$$
, $s = |S|$ with vevs: $v(T)$ and $w(T)$ such that $v(T_c) = w(T_c)$

$$V_T(h,s) = \frac{\lambda_H v_c^4}{4} \left(\frac{h^2}{v_c^2} + \frac{s^2}{w_c^2} - 1 \right)^2 + \frac{\lambda_H v_c^2}{m_{s,c}^2 w_{0,c}^4} h^2 s^2 + (T^2 - T_c^2) (c_h h^2 + c_s s^2), \quad (18)$$

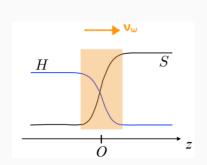
$$c_h = \frac{1}{48} \left(9g_2^2 + 3g_1^2 + 12y_t^2 + 24\lambda_H + \lambda_{HS} \right), \quad c_s = \frac{1}{12} \left(3\lambda_S + 2\lambda_{HS} \right).$$
 (19)

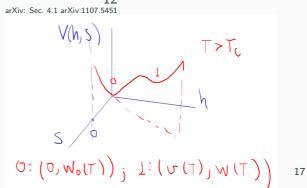


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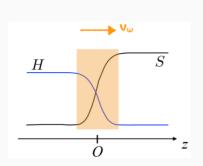


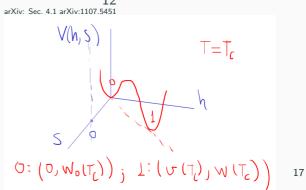


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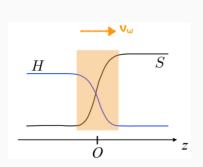


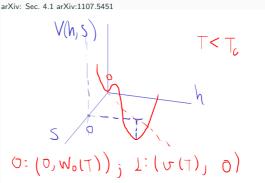


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CP assymetry generation i

Using the thin wall approximantion for the nucleation bubbles, we use the ansatz in which the space dependence of the fields is given by

$$h(z) = \frac{1}{2}v(T_n)(1-\tanh(z/L_w)), \qquad s(z) = \frac{1}{2}w_0(T_n)(1+\tanh(z/L_w)),$$

where z is the direction normal to the wall and L_w is the wall width.

The nucleation temperature, T_n , is defined by the condition

$$\exp(-S_3/T_n) = \frac{3}{4\pi} \left(\frac{H(T_n)}{T_n}\right)^4 \left(\frac{2\pi T_n}{S_3}\right)^{\frac{3}{2}},$$

where S_3 is the Euclidean action of the bubble and H(T) is the Hubble rate.

Boltzmann equation i

$$egin{align} \xi_i(z) &\equiv \mu_i(z)/T = 6\left(n_i - \overline{n}_i
ight)/T^3, \ &-D_L \xi_{\chi_L}'' - v_w \xi_{\chi_L}' + \Gamma_L (\xi_{\chi_L} - \xi_{\chi_R}) = S_{\mathcal{LP}}, \end{aligned}$$

where D_L is the diffusion constant for χ_L , which is related to the scattering rate Γ_L by

$$D_{L} = \frac{3x+2}{x^{2}+3x+2} \frac{1}{3\Gamma_{L}}, \qquad x \equiv m_{\chi}/T$$
 (20)

and

$$S_{QP} = -\frac{\lambda}{2} \frac{v_w D_L}{\frac{3x+2}{x^2+3x+2} T} \frac{(1-x)e^{-x} + x^2 E_1(x)}{4m_\chi^2 K_2(x)} \frac{m_\chi w_0(T_n) \lambda \left(-2 + \cosh\left(\frac{2z}{L_w}\right)\right) \sin\theta}{L_w^3 \cosh^4\left(\frac{z}{L_w}\right)}, \quad (21)$$

where v_w is the wall's velocity $E_1(x)$ is the error function and $K_2(x)$ is the modified Bessel function of the second kind. $\mathbf{y} = \lambda e^{i\theta - i\pi/2}$

Transfer DM assymetry to SM quarks

The chiral particle give rise to a non-zero $U(1)_B$ charge density in the proximity of the wall. This results in a Z' background that couples to the SM fields with $U(1)_B$ charge,

$$\langle Z'_0(z) \rangle = \frac{g_B (q_{\chi_L} - q_{\chi_R}) T_n^3}{6 M_{Z'}} \int_{-\infty}^{\infty} dz_1 \, \xi_{\chi_L}(z_1) \, e^{-M_{Z'}|z-z_1|} \,,$$

which generates a chemical potential for the SM quarks,

$$\mu_Q(z) = \mu_{d_R,u_R}(z) = 3 \times \frac{5}{9} \times g_B \langle Z_0'(z) \rangle.$$

This chemical potential sources a thermal-equilibrium asymmetry in the quarks, $\Delta n_O^{\rm EQ}(z) \sim T_n^2 \mu_O(z)$.

From [1]

If the Z' is sufficiently light, it mediates a long range force that extends into the region outside the bubble wall with unbroken electroweak symmetry.

Finally, the baryon-number asymmetry is then given by

$$n_B \,=\, rac{\Gamma_{
m sph}}{v_W} \int_0^\infty {
m d}\, z \, n_Q^{
m EQ}(z) \, \exp\left(-rac{\Gamma_{
m sph}}{v_W} \, z
ight) \,,$$

where Γ_{sph} is the sphaleron rate. The baryon-to-photon-number ratio is then obtained by

$$\eta_B = \frac{n_B}{s(T_n)}, \quad s(T) \equiv \frac{2\pi^2}{45} g_{*S}(T) T^3,$$

where $g_{*S}(T)$ is the effective number of relativistic degrees of freedom.

Our goal is to find what regions of the parameter space yield

$$0.82 \times 10^{-10} < \eta_B < 0.92 \times 10^{-10} \,. \tag{22}$$

https://github.com/anferivera/DarkBariogenesis

- SARAH→SPheno→MicroMegas
- η_B calculation code
- Python notebook with the scan

arXiv:1810.08055

Ten Simple Rules for Reproducible Research in Jupyter Notebook Fernando Pérez, et al

[...] In this paper, we address several questions about reproducibility [...] Combined with software repositories and open source licensing, notebooks are powerful tools for transparent, collaborative, reproducible, and reusable data analyses.

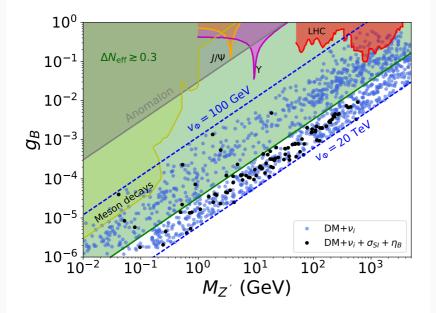
Results

We vary the typical Dirac-fermion DM parameter space and for each point that satisfy neutrino oscillation data, relic density and DM direct detection constraints. For each point we ...

Parameter	Range
θ	$(-\pi/2,\pi/2)$
$w_0(T_n)/{\rm GeV}$	100 - 500
$T_n/{ m GeV}$	100 - 200
$L_w/{ m GeV^{-1}}$	$1/T_n - 10/T_n$
V_W	0.05 - 0.5

Table 2: Scan ranges for the free parameters that are involved in the baryogenesis mechanism.

Black points: Dirac neutrinos with proper DM and baryon assymetry



Conclusions

A $U(1)_B$ is presented as an example of models where all new fermions required to cancel out the anomalies are used to solve phenomenological problems of the standard model (SM):

- EW-scale fermion vector-like doublets and iso-singlet charged singlets, in conjunction
 with right-handed neutrinos with repeated Abelian charges, participate in the generation
 of small neutrino masses through the Dirac-dark Zee mechanism
- The other SM-singlets are used to explain the dark matter in the universe, while their coupling to an inert singlet scalar is the source of the *CP* violation.

In the presence of a strong first-order electroweak phase transition, this "dark" CP violation allows for successful electroweak baryogenesis by using long range force mediated by a sufficiently light Z' which transfers the assymmetry from the Dark sector into the SM.