Dark matter from SM gauge extensions



with neutrino masses

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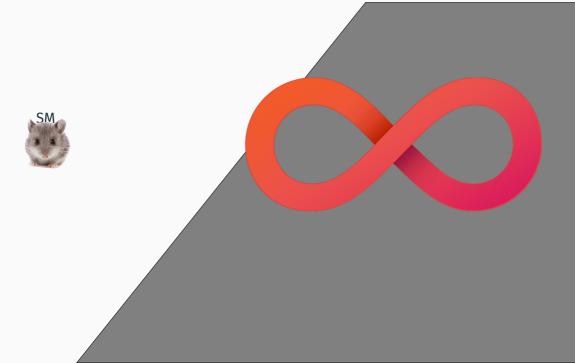


Focus on

In collaboration wit

M. Hirsch (IFIC), C. Álvarez (UTFSM), A. Flórez (UniAndes), B. Dutta(Texas A& M), C. Yaguna (UPTC), J. Calle, O. Zapata, A. Rivera (UdeA), W. Tangarife (Loyola University Chicago)

Hidden sectors







$$m_{\text{Majorana}}^{\nu} = \frac{1}{\Lambda} L \cdot H L \cdot H$$
 (1-loop)

arXiv:1308.3655 [JHEP] with C. Yaguna and Ó. Zapata

1



 $m_{
m Majorana}^{
u}=rac{1}{\Lambda}{
m L}\cdot{
m HL}\cdot{
m H}$ $m_{
m Dirac}^{
u}=rac{1}{\Lambda}\left(
u_{
m R}
ight)^{\dagger}{
m L}\cdot{
m HS}$

-

Local $U(1)_X \rightarrow Z_N$



$$m_{
m Majorana}^{
u} = rac{1}{\Lambda} L \cdot HL \cdot H$$

 $m_{
m Dirac}^{
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u_R)^{\dagger} L \cdot HS$

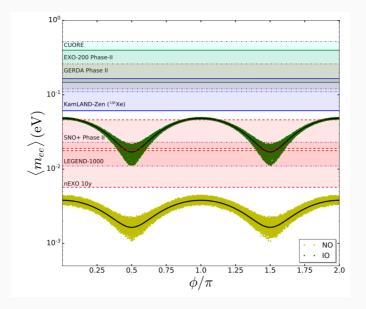
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Neutrino masses

Lepton number

- Lepton number (*L*) is an accidental discret or Abelian symmetry of the standard model (SM).
- · Without neutrino masses L_e , L_μ , L_τ are also conserved.
- The processes which violates individual *L* are called Lepton flavor violation (LFV) processes.
- · All the neutrino mass models predict, to some extent, LFV processes
- Only models with Majorana neutrinos predict processes with total $L = L_e + L_\mu + L_\tau$ violation, like neutrino less doublet beta decay (NLDBD).
- NLDBD is experimentally challenging, specially if there is a massless neutrino in the spectrum.

NLDBD prospects for a model with a massless neutrino (arXiv:1806.09977 [PLB] with Reig, Valle and Zapata)



Total lepton number: $L = L_e + L_\mu + L_{\tau_1}$

Majorana U(1)

Field
$$Z_2 (\omega^2 = 1)$$

SM 1
 $L \qquad \omega$
 $(e_R)^{\dagger} \qquad \omega$
 $(\nu_R)^{\dagger} \qquad \omega$

$$\mathcal{L}_{\nu} = h_D (\nu_R)^{\dagger} L \cdot H + M_R \nu_R \nu_R + \text{h.c.}$$

$$h_D \sim \mathcal{O}(1)$$

Dirac $U(1)_L$

Field
$$Z_3$$
 ($\omega^3 = 1$)

SM 1

L ω
 $(e_R)^{\dagger}$ ω^2
 $(\nu_R)^{\dagger}$ ω^2

$$\mathcal{L}_{\nu} = h_D (\nu_R)^{\dagger} L \cdot H + \text{h.c.}$$

$$h_D \sim \mathcal{O}(10^{-11})$$

Total lepton number: $L=L_e+L_\mu+L_ au$

Majorana U(1)[

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Dirac $U(1)_{B-L}$

Field
$$Z_3$$
 ($\omega^3 = 1$)
SM 1
 L ω
 $(e_R)^{\dagger}$ ω^2
 $(\nu_R)^{\dagger}$ ω^2

$$\mathcal{L}_{\nu} = h_D (\nu_R)^{\dagger} L \cdot H + \text{h.c.}$$

$$h_D \sim \mathcal{O}(10^{-11})$$

Explain smallness ala Peccei-Quinn: $U(1)_{B-L} \xrightarrow{(S)} Z_N$, $N \ge 3$.

To explain the smallness of Dirac neutrino masses choose $U(1)_{B-L}$ which:

• Forbids tree-level mass (TL) term (Y(H) = +1/2)

$$\mathcal{L}_{T.L} = h_D \epsilon_{ab} (\nu_R)^{\dagger} L^a H^b + \text{h.c}$$
$$= h_D (\nu_R)^{\dagger} L \cdot H + \text{h.c}$$

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• Forbids Majorana term: $u_{R}
u_{R}$

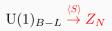
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- Forbids Majorana term: $\nu_R \nu_R$
- Realizes of the 5-dimension operator which conserves lepton number in $SU(3)_c \times SU(2)_L \times U(1)_Y \times U(1)_{B-L}$:

$$\mathcal{L}_{5-D} = \frac{h_{\nu}}{\Lambda} (\nu_R)^{\dagger} L \cdot HS + \text{h.c}$$





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H V_R

 $U(1)_{B-L} \stackrel{\langle S \rangle}{\to} Z_N$

• Enhancement to the effective number of degrees of freedom in the early Universe $\Delta N_{\rm eff} = N_{\rm eff}^{\rm SM}$ (see arXiv:1211.0186)

See E. Ma, Rahul Srivastava: arXiv:1411.5042 [PLB] for tree-level realization



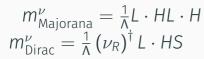
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Dark matter and unification

Unification: SO(10)

 $\Rightarrow \mathcal{L}_{\text{SM}} \supset h\, \textbf{16}_{\text{F}} \times \textbf{16}_{\text{F}} \times \textbf{10}_{\text{S}} + \text{h.c}$



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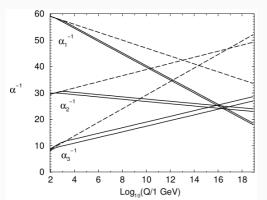
SO(10) breakings

$$SO(10) \rightarrow \begin{cases} SU(5) \times U(1)_X & \text{with } \begin{cases} Z_X \text{ at GUT: } U(1)_X \rightarrow Z_N \\ Z_X \text{ at EW} \end{cases} \\ SU(3)_C \times SU(2)_R \times SU(2)_L \times U(1)_Y \times U(1)_{B-L} \end{cases}$$

Majorana neutrinos case

Standard Model: Z ₂ -even	New Z ₂ -odd particles
Fermions: 16 _F	$10_F, 45_F, \cdots$
Scalars: 10 _{<i>H</i>} , 45 _{<i>H</i>} · · ·	16 _{<i>H</i>} , ⋅ ⋅ ⋅

Lightest Odd Particle (LOP) may be a suitable dark matter candidate, and can improve gauge coupling unification



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	fermions	scalars
$SU(2)_L \times U(1)_Y$	even $SO(10)$	odd $SO(10)$
representation	representations	representations
1_0	45, 54, 126, 210	16, 144
$2_{\pm 1/2}$	10, 120, 126, 210, 210'	16, 144
3_0	45, 54, 210	144

 $SU(3)_{C}: 3(T), 6, 8(\Lambda)$

$$m_{
m 3_0} = 2.7 \; {
m TeV}, \qquad m_{\Lambda} \sim 10^{10} \; {
m TeV}, \qquad m_{
m GUT} \sim 10^{16} \; {
m GeV} \, .$$

arXiv:0912.1545 [PRD] (Frigerio-Hambye)

Standard Model: Z ₂ -even	New Z ₂ -odd particles
Fermions: 16 _F	$10_F, 45_F, \cdots$
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$\overline{(3_0)}$	45, 54, 210	144

 $SU(3)_{C}: 3(T), 6, 8(\Lambda)$

Split-SUSY like

arXiv:1509.06313 [PRD] with C. Arbelaez, R. Longas, and O. Zapata.

Standard Model: Z_2 -evenNew Z_2 -odd particlesFermions: $\mathbf{16}_F$ $\mathbf{10}_F, \mathbf{45}_F, \cdots$ Scalars: $\mathbf{10}_H, \mathbf{45}_H \cdots$ $\mathbf{16}_H, \cdots$

Lightest Odd Particle (LOP) may be a suitable dark matter candidate, and can improve gauge coupling unification

		fermions	scalars	
	$SU(2)_L \times U(1)_Y$	even $SO(10)$	odd $SO(10)$	
	representation	representations	representations	
	1_0	45, 54, 126, 210	16 144	
	$(2_{\pm 1/2})(2_{1/2}^{S})$	10, 120, 126, 210, 210'	16, 144	
	(3_0)	45, 54, 210	144	
-	$SU(3)_C: 3(T), 6, 8(7)$	Radiative h	ybrid seesaw (Parid	a 1106.4137) or 1509.06313

Partial Split-SUSY-like spectrum: bino-higgsino-wino

$$\mathbf{10}_{H}'$$
 with fermion DM or, $\mathbf{16}_{H}, \cdots$ with scalar DM

Standard Model: Z_2 -evenNew Z_2 -odd particlesFermions: $\mathbf{16}_F$ $\mathbf{10}_F, \mathbf{45}_F, \cdots$ Scalars: $\mathbf{10}_H, \mathbf{45}_H \cdots$ $\mathbf{16}_H, \cdots$

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CU(2) $G(T)$	1509.06313	

 $SU(3)_C : (3 (T)) 6, (8 (\Lambda))$

SUSY-like spectrum: bino-higgsino-wino

 $10'_H$ with fermion DM or, 16_H , ... with scalar DM

Singlet-Doublet-Triplet fermion dark-matter

The most general SO(10) invariant Lagrangian contains the following Yukawa terms

$$-\mathcal{L} \supset Y10_F45_F10_H + M_{45_F}45_F45_F + M_{10_F}10_F10_F$$

Basis
$$\psi^0 = \left(N, \Sigma^0, \psi_L^0, (\psi_R^0)^\dagger\right)^T$$

$$\mathcal{M}_{\psi^0} = \begin{pmatrix} M_N & 0 & -yc_\beta v/\sqrt{2} & ys_\beta v/\sqrt{2} \\ 0 & M_\Sigma & fc_\beta v/\sqrt{2} & -fs_\beta v/\sqrt{2} \\ -yc_\beta v/\sqrt{2} & fc_\beta v/\sqrt{2} & 0 & -M_D \\ ys_\beta v/\sqrt{2} & -fs_\beta v/\sqrt{2} & -M_D & 0 \end{pmatrix},$$

$$\mathbf{10}_F \to \psi_L, (\psi_R)^\dagger$$

$$\mathbf{45}_F \to \Sigma, \Lambda$$

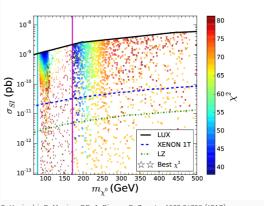
$$\mathbf{45}_F' \to N$$

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$$\begin{aligned} \text{Basis } \boldsymbol{\psi}^0 &= \begin{pmatrix} N, \Sigma^0, \boldsymbol{\psi}_L^0, \left(\boldsymbol{\psi}_R^0\right)^\dagger \end{pmatrix}^T \\ \boldsymbol{\mathcal{M}}_{\boldsymbol{\psi}^0} &= \\ \begin{pmatrix} M_N & 0 & -\mathbf{y} c_\beta \mathbf{v}/\sqrt{2} & \mathbf{y} s_\beta \mathbf{v}/\sqrt{2} \\ 0 & M_\Sigma & f c_\beta, \mathbf{v}/\sqrt{2} & -f s_\beta, \mathbf{v}/\sqrt{2} \\ -\mathbf{y} c_\beta \mathbf{v}/\sqrt{2} & -f s_\beta, \mathbf{v}/\sqrt{2} & 0 & -M_D \\ \mathbf{y} s_\beta \mathbf{v}/\sqrt{2} & -f s_\beta, \mathbf{v}/\sqrt{2} & -M_D & 0 \end{pmatrix}, \\ \mathbf{10}_F &\to \boldsymbol{\psi}_L, \left(\boldsymbol{\psi}_R\right)^\dagger \\ \mathbf{45}_F &\to \boldsymbol{\Sigma}, \boldsymbol{\Lambda} \\ \mathbf{45}_F' &\to \boldsymbol{N} \end{aligned}$$



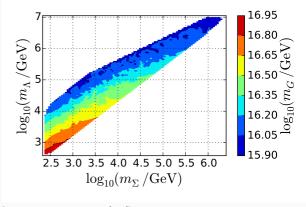
S. Horiuchi, O. Macias, DR, A. Rivera, O. Zapata, 1602.04788 (JCAP)

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$$-\mathcal{L} \supset Y10_F45_F10_H + M_{45_F}45_F45_F + M_{10_F}10_F10_F + \mathcal{L}(10_{\Phi})$$

$$\begin{aligned} \text{Basis } \boldsymbol{\psi}^0 &= \begin{pmatrix} \mathsf{N}, \boldsymbol{\Sigma}^0, \boldsymbol{\psi}_\mathsf{L}^0, \begin{pmatrix} \boldsymbol{\psi}_\mathsf{R}^0 \end{pmatrix}^\dagger \end{pmatrix}^\mathsf{T} \\ \boldsymbol{\mathcal{M}}_{\boldsymbol{\psi}^0} &= \\ \begin{pmatrix} \mathsf{M}_\mathsf{N} & \mathbf{0} & -\mathsf{y} c_\mathsf{B} \mathsf{v}/\sqrt{2} & \mathsf{y} s_\mathsf{B} \mathsf{v}/\sqrt{2} \\ \mathbf{0} & \mathsf{M}_{\boldsymbol{\Sigma}} & f c_{\mathsf{B}}, \mathsf{v}/\sqrt{2} & -f s_{\mathsf{B}}, \mathsf{v}/\sqrt{2} \\ -\mathsf{y} c_\mathsf{B} \mathsf{v}/\sqrt{2} & -f s_\mathsf{B}, \mathsf{v}/\sqrt{2} & \mathbf{0} & -\mathsf{M}_\mathsf{D} \\ \mathbf{y} s_\mathsf{B} \mathsf{v}/\sqrt{2} & -f s_\mathsf{B}, \mathsf{v}/\sqrt{2} & -\mathsf{M}_\mathsf{D} & \mathbf{0} \end{pmatrix}, \\ \mathbf{10}_F &\to \boldsymbol{\psi}_\mathsf{L}, (\boldsymbol{\psi}_\mathsf{R})^\dagger \\ \mathbf{45}_F &\to \boldsymbol{\Sigma}, \boldsymbol{\Lambda} \\ \mathbf{45}_\mathsf{L}^c &\to \boldsymbol{N} \end{aligned}$$



(See previous arXiv:1509.06313 [PRD]): Split-SUSY: like $M_{\Phi} = 2 \text{ TeV}$

Not-susy SO(10) \rightarrow SU(3)_c \times SU(2)_R \times SU(2)_L \times U(1)_Y \times U(1)_{B-L}

Fermionic triplet dark matter in an SO(10)-inspired left right model, with C. Arbeláez and M. Hirsch, arXiv:1703.08148 [PRD]

Minimal Left-Right Symmetric Standard Model

Field	Multiplicity	$3_c 2_L 2_R 1_{B-L}$	Spin	SO(10) origin
Q	3	$(3,2,1,+\frac{1}{3})$	1/2	16
Q ^c	3	$(\bar{3},1,2,-\frac{1}{3})$	1/2	16
L	3	(1,2,1,-1)	1/2	16
_C	3	(1,1,2,+1)	1/2	16
Ф	1	(1, 2, 2, 0)	0	10
Δ_R	1	(1,1,3,-2)	0	126

Left-singlet right-triplet DM

Field	Multiplicity	$3_c 2_L 2_R 1_{B-L}$	Spin	SO(10) origin
Q	3	$(3,2,1,+\frac{1}{3})$	1/2	16
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	ĺ	'		

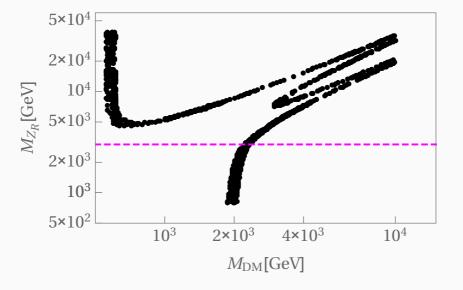


Figure 1: Proper relic density scan: $0.5 < v_R/\text{TeV} < 50$

Mixed Left-singlet right-triplet DM

Field	Multiplicity	$3_c 2_L 2_R 1_{B-L}$	Spin	SO(10) origin
Q	3	$(3,2,1,+\frac{1}{3})$	1/2	16
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Δ_R	1	(1,1,3,-2)	0	126
Ψ ₁₁₃₀	1	(1, 1, 3, 0)	1/2	45
Ψ_{1132}	1	(1,1,3,2)	1/2	126
Ψ_{113-2}	1	(1,1,3,-2)	1/2	126

$$\Psi_{1132} = \begin{pmatrix} \Psi^{+}/\sqrt{2} & \Psi^{++} \\ \Psi^{0} & -\Psi^{+}/\sqrt{2} \end{pmatrix}, \qquad \bar{\Psi}_{113-2} = \begin{pmatrix} \Psi^{-}/\sqrt{2} & \overline{\Psi}^{0} \\ \Psi^{--} & -\Psi^{-}/\sqrt{2} \end{pmatrix}. \tag{1}$$

$$L \supset M_{11} \operatorname{Tr}(\Psi_{1130}\Psi_{1130}) + M_{23} \operatorname{Tr}(\Psi_{1132}\bar{\Psi}_{113-2}) + \lambda_{13} \operatorname{Tr}(\Delta_R\bar{\Psi}_{113-2}\Psi_{1130}) + \lambda_{12} \operatorname{Tr}(\Delta_R^{\dagger}\Psi_{1132}\Psi_{1130}),$$
 (2)

$$\tan \gamma = \frac{\lambda_{13}}{\lambda_{12}}, \qquad \lambda = \sqrt{\lambda_{12}^2 + \lambda_{13}^2}. \tag{3}$$

Blind spot at

$$M_{23}\sin 2\gamma - M_{\rm DM} = 0 \tag{4}$$

Proper relic density scan

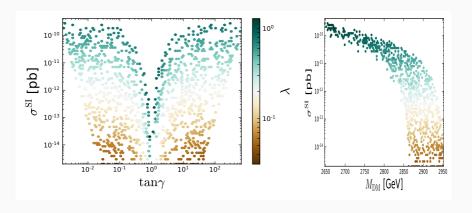


Figure 2: $M_{11} = 50 \text{ TeV } 2.7 < M_{23}/\text{TeV} < 3.1$ (Right: $\tan \gamma > 5$)

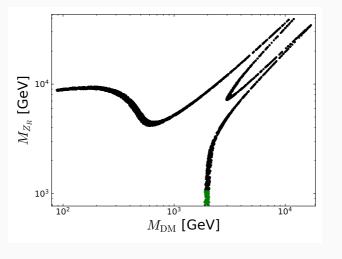


Figure 3:

Direct detection cross section

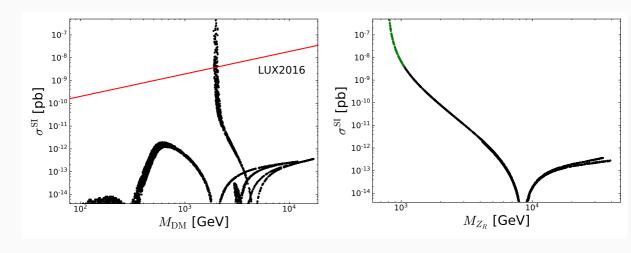


Figure 4: v_R : [2,50] TeV, M_{23} : [0.2,50] TeV, M_{11} : 50 TeV, $\tan \gamma = -1$ and $\lambda = 0.14$.

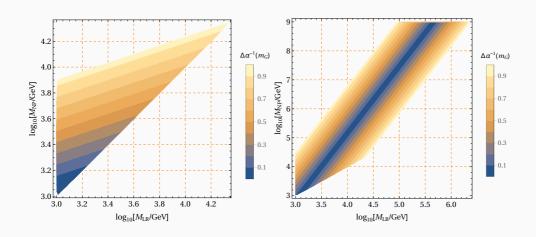
Unification

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Ψ_{1132}	1	(1, 1, 3, 2)	1/2	126
Ψ_{113-2}	1	(1,1,3,-2)	1/2	126
Ψ_{1310}	1	(1, 3, 1, 0)	1/2	45
Ψ_{8110}	1	(1, 1, 8, 0)	1/2	45
$\Psi_{321\frac{1}{3}}$	1	(3, 2, 1, 1/3)	1/2	16
$\Psi_{321-\frac{1}{3}}$	1	(1,2,3,-1/3)	1/2	16

Unification quality



Remarks

In addition to accommodate usual simplified dark matter models, Left-right symmetric standard models have additional DM portals:

New Δ_R portal for direct detection of left-singlet right-triplet mixed dark matter, in companion with left-singlets charged and doubly charged fermions.

Next: Search for them in compressed spectra scenarios at the LHC

Dirac neutrinos case:

SO(10) \rightarrow SU(5) \times U(1)_X

From 1210.6350 and 1805.02025: $\Delta N_{\text{eff}} = 3 (T_{\nu_R}/T_{\nu_L})^4$

$$\Gamma_{\nu_R}(T) = n_{\nu_R}(T) \sum_f \langle \sigma_f(\nu_R \bar{\nu}_R \to f\bar{f}) v \rangle$$

$$= \sum_f \frac{g_{\nu_R}^2}{n_{\nu_R}} \int \frac{d^3p}{(2\pi)^3} \frac{d^3q}{(2\pi)^3} f_{\nu_R}(p) f_{\nu_R}(q) \sigma_f(s) (1 - \cos\theta),$$

$$s = 2pq(1 - \cos \theta), f_{\nu_R}(k) = 1/(e^{k/T} + 1)$$

$$n_{\nu_R}(T) = g_{\nu_R} \int \frac{d^3k}{(2\pi)^3} f_{\nu_R}(k), with g_{\nu_R} = 2$$

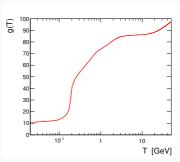
$$\sigma_f(s) \simeq \frac{N_C^f (Q_{BL}^f)^2 Q^2 s}{12\pi} \left(\frac{g'}{M_{el}}\right)^4, In the limit M_{Z'}^2 \gg s.$$

with three right-handed neutrinos, the Hubble parameter is

$$H(T) = \sqrt{\frac{4\pi^3 G_N \left[g(T) + 21/4\right]}{45}} T^2.$$

The right-handed neutrinos decouple when

$$\Gamma_{\nu_R}(T_{\text{dec}}^{\nu_R}) = H(T_{\text{dec}}^{\nu_R}).$$



A. Solaguren-Beascoa, M. C. Gonzalez-Garcia: arXiv:1210.6350 [PLB]

From 1210.6350 and 1805.02025: $\Delta N_{\text{eff}} = 3 \left(T_{\nu_R} / T_{\nu_L} \right)^4$

$$\begin{split} \Gamma_{\nu_R}(T) &= n_{\nu_R}(T) \sum_f \langle \sigma_f(\nu_R \bar{\nu}_R \to f\bar{f}) v \rangle \\ &= \sum_f \frac{g_{\nu_R}^2}{n_{\nu_R}} \int \frac{d^3p}{(2\pi)^3} \frac{d^3q}{(2\pi)^3} f_{\nu_R}(p) f_{\nu_R}(q) \sigma_f(s) (1 - \cos\theta), \end{split}$$

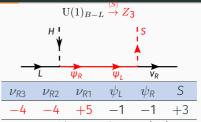
$$\begin{split} s = &2pq(1-\cos\theta), & f_{\nu_R}(k) = &1/(e^{k/T}+1) \\ n_{\nu_R}(T) = &g_{\nu_R} \int \frac{d^3k}{(2\pi)^3} f_{\nu_R}(k), & \text{with } g_{\nu_R} = &2 \\ \sigma_f(s) \simeq &\frac{N_C^f(Q_{BL}^f)^2 Q^2 s}{12\pi} \left(\frac{g'}{M_{Pl}}\right)^4, & \text{In the limit } M_{Z'}^2 \gg s. \end{split}$$

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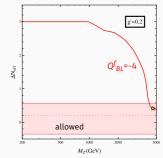
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E. Ma, R. Srivastava: arXiv:1411.5042 [PLB]

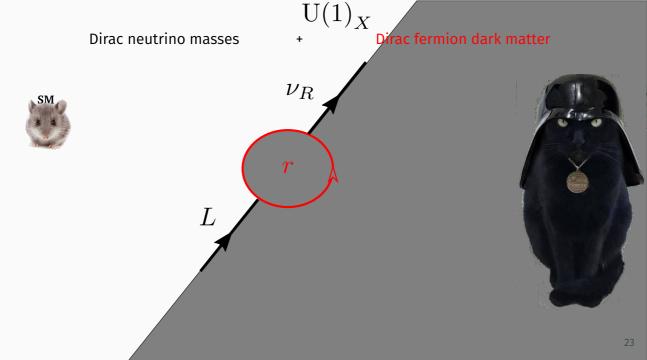


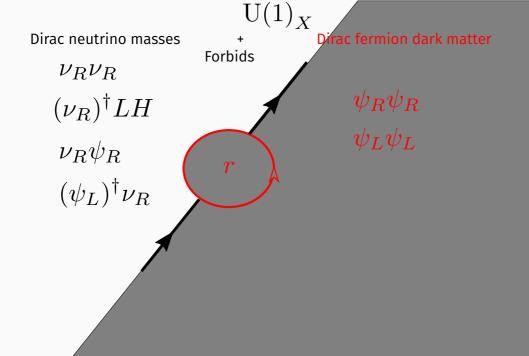
Z.-L. Han, W. Wang: arXiv:1805.02025 [EJPC]

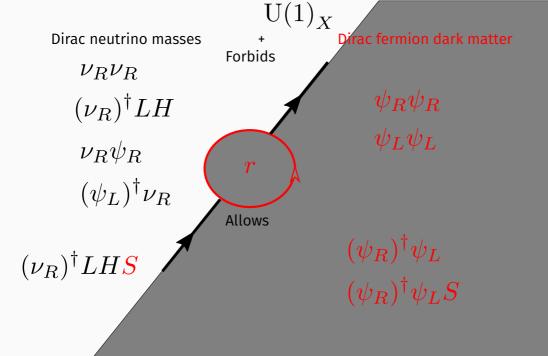
(also: Planck 1807.06209, Riess et al 1903.07603)

One-loop realization of \mathcal{L}_{5-D} with

total L







Dirac neutrino masses
$$\nu_R\nu_R$$

$$(\nu_R)^\dagger LH$$

$$\nu_R\psi_R$$

$$(\nu_R)^\dagger \nu_R$$

$$(\nu_R)^\dagger \nu_R$$

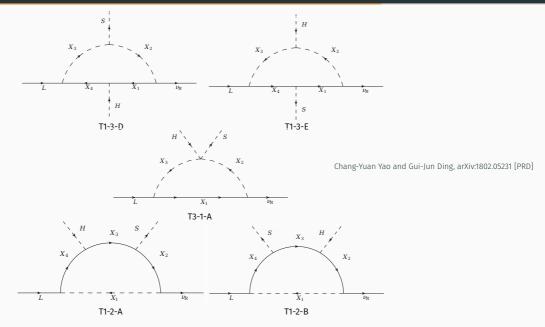
$$(\nu_R)^\dagger \nu_R$$

$$(\nu_R)^\dagger \nu_R$$
 Allows
$$(\nu_R)^\dagger \psi_L$$

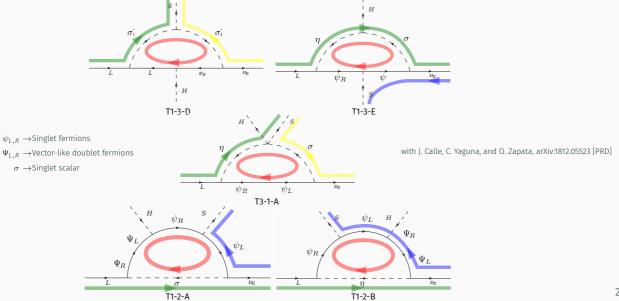
$$(\nu_R)^\dagger \psi_L$$

$$(\nu_R)^\dagger \psi_L S$$

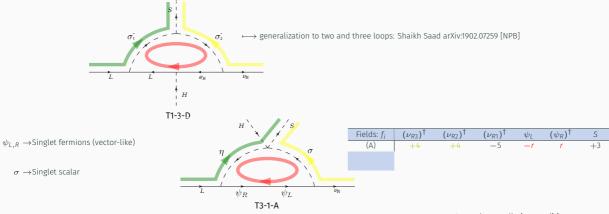
One loop topologies $U(1)_{B-L} \oplus Z_2 \oplus Z_2$



One loop topologies $U(1)_{B-L}$ only!



One loop topologies $U(1)_{B-L}$ only! with J. Calle, C. Yaguna, and O. Zapata, arXiv:1812.05523 [PRD]

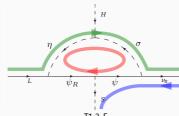


Anomaly cancellation conditions

$$\sum_{i} f_{i} = 3$$

$$\sum_{i} f_{i}^{3} = 3$$

One loop topologies $U(1)_{B-L}$ only! with J. Calle, C. Yaguna, and O. Zapata, arXiv:1812.05523 [PRD]



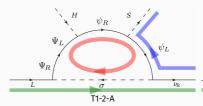
T1-3-E

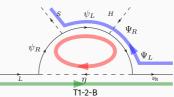
Fields	$: f_i \mid (\nu_{R3})^{\dagger}$	$(\nu_{R2})^{\dagger}$	$(\nu_{R1})^{\dagger}$	ψ_{L}	$(\psi_R)^{\dagger}$	S
(A)	+4	+4	-5	-r	r	+3
(B)	$+\frac{8}{5}$	$+\frac{8}{5}$	$+\frac{2}{5}$	7 _ 5	$-\frac{10}{5}$	+ - 5

 $\psi_{L,R} \rightarrow \text{Singlet fermions (quiral)}$

 $\Psi_{L,R}
ightarrow$ Vector-like doublet fermions

 $\sigma \rightarrow Singlet scalar$





Anomaly cancellation conditions

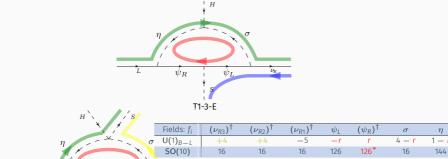
$$\sum_{i} f_{i} = \sum_{i} f_{i}^{3} = \sum_{i} f_{i}^{3$$

$$f_i^3 = 3$$

Scotogenic Dirac from SO(10) E. Ma arXiv:1901.09091 [PLB]

 ψ_R

T3-1-A



 ν_R

 $\psi_{L,R} o$ Singlet fermions (quiral)

 $\sigma
ightarrow ext{Singlet scalar}$

 $\eta \to \text{Doublet scalar}$

Anomaly cancellation conditions

$$\sum_{i} f_{i} = 3$$

$$\sum_{i} f_{i}^{3} = 3$$

+3

672









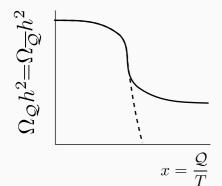


Colored dark matter: De Luca, Mitridate, Redi, Smirnov & Strumia, arXiv:1801.01135 [PRD]

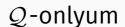
(Switch to Dirac fermions) Because Q is a Dirac fermion, QQ is also stable

 $QQ \rightarrow g$,

 $\overline{QQ} \rightarrow g$.

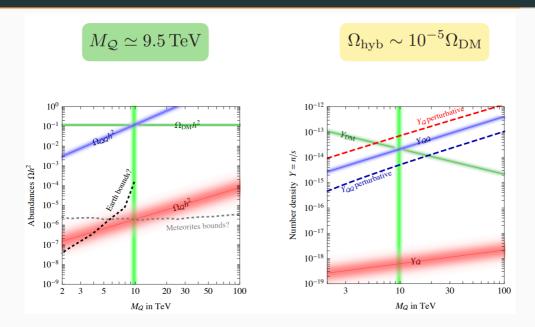


Step one

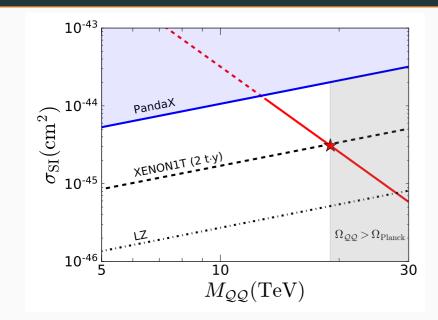




Step two



Direct detection

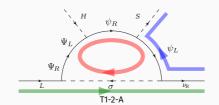


$SD^3M+\sigma_i~(i=1,2)$ with J. Calle, C. Yaguna, and O. Zapata, arXiv:1812.05523 [PRD]

 $\psi_{L,R} o$ Singlet fermions (quiral)

 $\Psi_{L,R} \rightarrow$ Vector-like doublet fermions : 10/5

 $\sigma \to \text{Singlet scalar}: 15/5$



Fields: fi	$(\nu_{R3})^{\dagger}$	$(\nu_{R2})^{\dagger}$	$(\nu_{R1})^{\dagger}$	ψ_{L}	$(\psi_R)^{\dagger}$	S	Ψ_L	$\widetilde{(\Psi_R)}$	σ
$U(1)_{B-L}$	$+\frac{8}{5}$	$+\frac{8}{5}$	$+\frac{2}{5}$	7 - 5	$-\frac{10}{5}$	+ = 5	10 5	$-\frac{10}{5}$	7 - 5
SO(10)	16	16	16	45	45	126	10	10	16

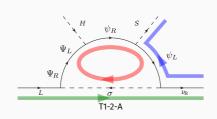
Second Scotogenic Dirac from **SO(10)** E. Ma arXiv:1901.09091 [PLB]

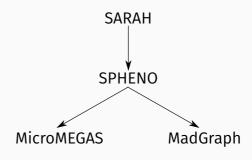
$SD^{3}M+\sigma_{i}$ (i=1,2)

$$M_{\psi} = h_1 \langle S \rangle$$
, $y_2 = 0$:

$$\mathcal{L} = \mathcal{L}_{\text{SD}^{3}\text{M}} + h_{3}^{ia}\widetilde{(\Psi_{R})} \cdot L_{i} \sigma_{a} + h_{2}^{\beta a} (\nu_{R\beta})^{\dagger} \psi_{L} \sigma_{a}^{*} - V(\sigma_{a}, S, H).$$

with A.F Rivera, W. Tangarife, arXiv:1906.09685





Singlet-Doublet Dirac Dark Matter (SD³M) By Carlos E. Yaguna. arXiv:1510.06151 [PRD].

The model extends the standard model (SM) particle content with Dirac Fermions: from SU(2) doublets of Weyl fermions: $\Psi_L = (\Psi_L^0, \Psi_L^-)^\mathsf{T}, \widetilde{(\Psi_R)} = ((\Psi_R^-)^\dagger, -(\Psi_R^0)^\dagger)^\mathsf{T}$ and singlet Weyl fermions ψ_{LR} that interact among themselves and with the SM fields

$$\mathcal{L} \supset M_{\psi} (\psi_{R})^{\dagger} \psi_{L} + M_{\psi} (\widetilde{\Psi_{R}}) \cdot \Psi_{L} + y_{1} (\psi_{R})^{\dagger} \Psi_{L} \cdot H + y_{2} (\widetilde{\Psi_{R}}) \cdot \widetilde{H} \psi_{L} + \text{h.c}$$
 (5)

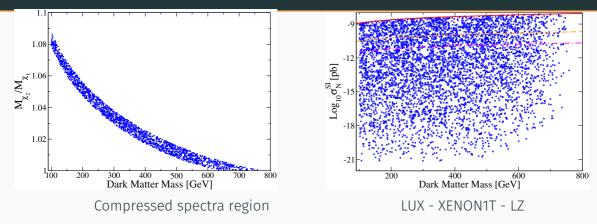
Four free parameters:

$$M_{\psi}, M_{\Psi} < 2 \text{ GeV},$$
 $y_1, y_2 > 10^{-6}$ (6)

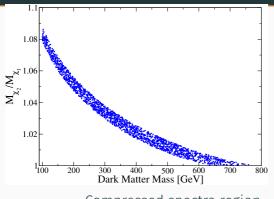
Two neutral Dirac fermion eigenstates:

$$M = \begin{pmatrix} M_{\psi} & y_2 v / \sqrt{2} \\ y_1 v / \sqrt{2} & M_D \end{pmatrix}, \qquad M_{\text{diag}} = \begin{pmatrix} M_{\chi_1} & 0 \\ 0 & M_{\chi_2} \end{pmatrix} = U_L^{\dagger} M U_R$$
 (7)

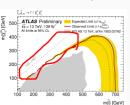
SD³M By Carlos E. Yaguna. arXiv:1510.06151 [PRD].

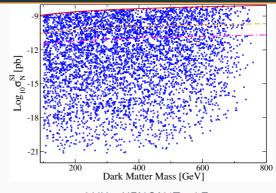


SD³M By Carlos E. Yaguna. arXiv:1510.06151 [PRD].



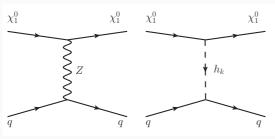
Compressed spectra region



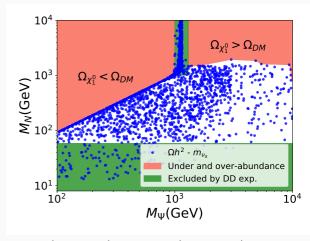


LUX - XENON1T - LZ

Spin independent (SI) direct detection cross section



Decoupled Z' limit



Vector SI (blue points) and scalar SI (green points)

Conclusions

A single U(1) symmetry to explain both the smallnes of neutrino masses and the stability of fermion dark matter

Conclusions

A single U(1) symmetry to explain both the smallnes of neutrino masses and the stability of fermion dark matter

Neutrino masses and DM

- Spontaneously broken $U(1)_X$ generates a radiative neutrino masses
- A remnant symmetry makes the lightest field circulating the loop stable and good dark matter candidate.
- For T1-2-A: Either Singet Doublet (Dirac) Dark Matter or Singlet Scalar Dark Matter with extra scalar and vector portal
- · With relaxed direct detection constraints

