

libyt: a Tool for Parallel In Situ Analysis with yt

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Abstract—In the era of exascale computing, storage and analysis of large scale data have become more important and difficult. We present *libyt*, an open source C++ library, that allows researchers to analyze and visualize data using *yt* or other Python packages in parallel during simulation runtime. We describe the code method for organizing adaptive mesh refinement grid data structure and simulation data, handling data transition between Python and simulation with minimal memory overhead, and conducting analysis with no additional time penalty using Python C API and NumPy C API. We demonstrate how it solves the problem in astrophysical simulations and increases disk usage efficiency. Finally, we conclude it with discussions about *libyt*.

Index Terms—astronomy data analysis, astronomy data visualization, in situ analysis, open source software

Introduction

In the era of exascale computing, storage and analysis of large-scale data has become a critical bottleneck. Simulations often use efficient programming language like C and C++, while many data analysis tools are written in Python, for example *yt*¹ [1]. *yt* is an open-source, permissively-licensed Python package for analyzing and visualizing volumetric data. It is a light weight tool for quantitative analysis for astrophysical data, and it has also been applied to other scientific domains. Normally, we would have to dump simulation data to hard disk first before conducting analysis using existing Python tools. This takes up lots of disk space when the simulation has high temporal and spatial resolution. This also forces us to store full datasets, even though our region of interest might contain only a small portion of simulation domain. It makes large simulation hard to analyze and manage due to the limitation of disk space. Is there a way to probe those ongoing simulation data using robust Python ecosystem? So that we don't have to re-invent data analysis tools and solve the disk usage issue at the same time.

In situ analysis, which is to analyze simulation data on-site, without intermediate step of writing data to hard disk is a promising solution. It also reduces the barrier of analyzing data by using well-developed tools instead of creating our own. We introduce in situ analysis tool *libyt*², an open source C++ library, that allows

researchers to analyze and visualize data by directly calling *yt* or any other Python packages during simulations runtime under parallel computation. Through wrapping ongoing simulation data using NumPy C API [2], constructing proper Python C-extension methods and Python objects using Python C API [3], we can reuse C++ runtime data and realize back-communication of simulation information, allowing user to define their own data generating C function, and use it to conduct analysis inside Python ecosystem. This is like using a normal Python prompt, but with direct access to simulation data. *libyt* provides another way for us to interact with simulations.

In this proceeding, we will describe the methods in Section **Code Method**, demonstrate how *libyt* solve the problem in Section **Applications**, and conlude it with Section **Discussions**.

Code Method

Overview of libyt

libyt serves as a bridge between simulation processes and Python instances as illustrated in Fig 1. It is the middle layer that handles data IO between simulations and Python instances, and between MPI processes. When launching N MPI processes, each process contains one piece of simulation and one Python interpreter. Each Python interpreter has access to simulation data. When doing in situ analysis, every simulation process pauses, and a total of N Python instances will work together to conduct Python tasks in the process space of MPI.

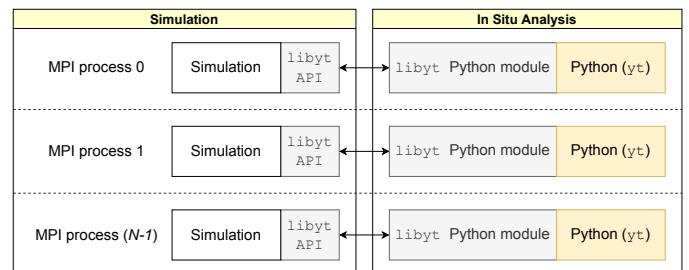


Fig. 1: This is the overall structure of *libyt*, and its relationship with simulation and Python. It provides an interface for exchanging data between simulations and Python instances, and between each process, thereby enabling in situ parallel analysis using multiple MPI processes. *libyt* can run arbitrary Python scripts and Python modules, though here we focus on using *yt* as its core analysis platform.

Simulations use *libyt API*³ to pass in data and run Python codes during runtime, and Python instances use *libyt Python*

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module to request data directly from simulations using C-extension method and access Python objects that contain simulation information. Using libyt for in situ analysis is very similar to running Python scripts in post-processing under MPI platform, except that data are stored in memory instead of hard drives. libyt is for general-purpose and can launch arbitrary Python scripts and Python modules, though here, we focus on using yt as our core analysis tool.

Connecting Python and Simulation

We can extend the functionality of Python by calling C/C++ functions, and, likewise, we can also embed Python in a C/C++ application to enhance its capability. Python and NumPy provides C API for users to connect objects in a main C/C++ program to Python.

Currently, libyt supports only adaptive mesh refinement (AMR) grid data structure.⁴ How libyt organizes simulation with AMR grid data structure is illustrated in Fig 2. It first gathers and combines local adaptive mesh refinement grid information (e.g., levels, parent id, grid edges, etc) in each process, so that every Python instance contains full information. Next, it allocates array using PyArray_SimpleNew and stores the information in a linear fashion according to global grid id. The array can be easily looked up, and we can retrieve information by libyt at C side using PyArray_GETPTR2. The operation only involves reading elements in an array. The array is accessible both in C/C++ and Python runtimes. For simulation data, libyt wraps those data pointers using NumPy C API PyArray_SimpleNewFromData. This tells Python how to interpret block of memory (e.g., shape, type, stride) and does not make a copy. libyt also marks the wrapped data as read-only⁵ to avoid Python accidentally alters it, since they are actual data used in simulation's iterative process.

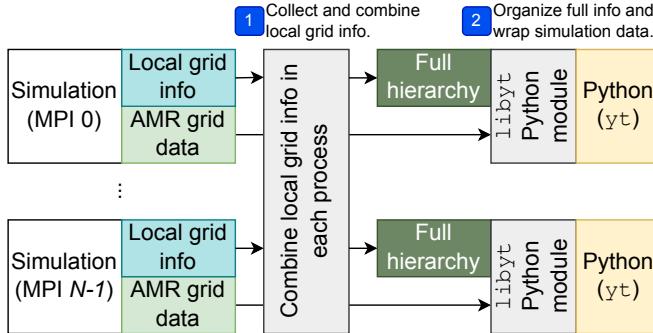


Fig. 2: This diagram shows how libyt loads and organizes simulation information and data that is based on adaptive mesh refinement (AMR) grid data structure. libyt collects local AMR grid information and combines them all, so that each Python instance contains whole information. As for simulation data, libyt wraps them using NumPy C API, which tells Python how to interpret block of memory without duplicating it.

libyt also supports back-communication of simulation information. Fig 3 shows the mechanism behind it. The process is

3. For more details, please refer to libyt documents. (<https://yt-project.github.io/libyt/libytAPI>)

4. We will support more data structures (e.g., octree, unstructured mesh grid, etc) in the future.

5. This can be done by using PyArray_CLEARFLAGS to clear writable flag NPY_ARRAY_WRITEABLE.

triggered by Python when it needs the data generated by a user-defined C function. This usually happens when the data is not part of the simulation iterative process and requires simulation to generate it, or the data isn't stored in a contiguous memory block and requires simulation to help collect it. When Python needs the data, it first calls C-extension method in libyt Python module. The C-extension method allocates a new data buffer and passes it to user-defined C function, and the function writes data in it. Finally, libyt wraps the data buffer and returns it back to Python. libyt makes the data buffer owned by Python⁶, so that the data gets freed when it is no longer needed.

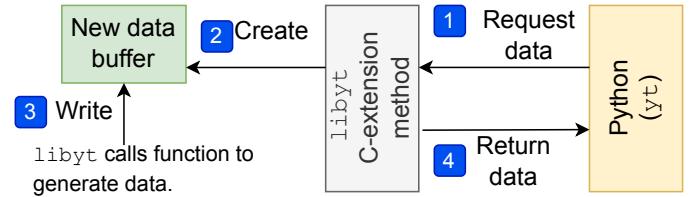


Fig. 3: This diagram describes how libyt requests simulation to generate data using user-defined C function, thus enabling back-communication of simulation information. Those generated data is freed once it is no longer used by Python.

Grid information and simulation data are properly organized in dictionaries under libyt Python module. One can import it during simulation runtime:

```
import libyt # Import libyt Python module
```

In Situ Analysis Under Parallel Computing

Each MPI process contains one simulation code and one Python instance. Each Python instance only has direct access to the data on local computing nodes, thus all Python instances must work together to make sure everything is in reach. During in situ Python analysis, workloads may be decomposed and rebalanced according to the algorithm in Python packages. It is not necessary to align with how data is distributed in simulation. Even though libyt can call arbitrary Python modules, we focus on how it uses yt and MPI to do analysis under parallel computation here.

yt supports parallelism feature⁷ using mpi4py⁸ as communication method. libyt borrows this feature and utilizes it directly. The way yt calculates and distributes jobs to each MPI process is based on data locality, but it does not always guarantee to do so⁹. In other words, in in situ analysis, the data requested by yt in each MPI process does not always locate in the same process.

Furthermore, there is no way for libyt to know what kind of communication pattern a Python script needs in advance. For a much more general case, it is difficult to schedule point-to-point communications that fit any kind of algorithms and any number of MPI processes. libyt uses one-sided communication in MPI, also known as Remote Memory Access (RMA), by which one no longer needs to explicitly specify senders and receivers. Fig 4

6. This can be done by using PyArray_ENABLEFLAGS to enable own-data flag NPY_ARRAY_OWNDATA.

7. See [Parallel Computation With yt](#) for more details.

8. mpi4py is Python bindings for MPI. (<https://mpi4py.readthedocs.io/en/stable/>)

9. yt functionalities like find_max, ProjectionPlot, create_profile, PhasePlot, etc are based on data locality, others like OffAxisProjectionPlot, SlicePlot, OffAxisSlicePlot, etc don't.

describes the data redistribution process in `libyt`. `libyt` first collects requested data in each process and asks each process to prepare it. Then `libyt` creates an epoch, for which all MPI processes will enter, and each process can fetch the data located on different processes without explicitly waiting for the remote process to respond. The caveat in data exchanging procedure in `libyt` is that it is a collective operation, and requires every MPI process to participate.

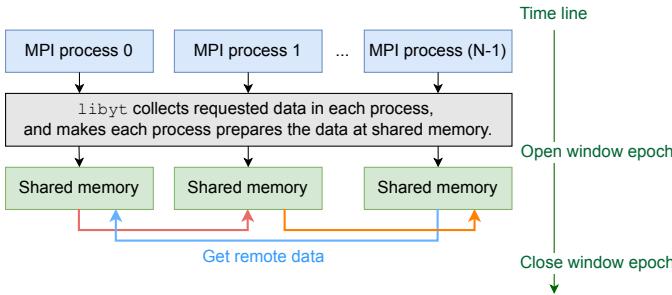


Fig. 4: This is the workflow of how `libyt` redistributes data. It is done via one-sided communication in MPI. Each process prepares the requested data from other processes, after this, every process fetches data located on different processes. This is a collective operation, and data is redistributed during this window epoch. Since the data fetched from other processes is only for analysis purpose, it gets freed once Python doesn't need it at all.

Executing Python Codes and Handling Errors

`libyt` imports user's Python script at the initialization stage. Every Python statement is executed inside the imported script's namespace using `PyRun_SimpleString`. The namespace holds Python functions and objects. Every change made will also be stored under this namespace and will be brought to the following round.

Using `libyt` for in situ analysis is just like running Python scripts in post-processing. The only difference lies in how the data is loaded. Post-processing has everything store on hard disk, while data in in situ analysis is distributed in memory space in different computing nodes. Though `libyt` can call arbitrary Python modules, here, we focus on using `yt` as the core method. This is an example of doing slice plot using `yt` function `SlicePlot` in post-processing:

```
1 import yt
2 yt.enable_parallelism()
3 def do_sliceplot(data):
4     ds = yt.load(data)
5     slc = yt.SlicePlot(ds, "z", ("gamer", "Dens"))
6     if yt.is_root():
7         slc.save()
8 if __name__ == "__main__":
9     do_sliceplot("Data000000")
```

Converting the post-processing script to inline script is a two-line change. We need to import `yt.libyt`¹⁰, which is the `yt` frontend for `libyt`. And then we change `yt.load` to `yt.libyt.libytDataset()`. That's it! Now data is loaded from `libyt` instead of loading from hard disk. The following is the inline Python script:

```
1 import yt.libyt
2 import yt
```

10. https://github.com/data-exp-lab/yt_libyt

```
3 yt.enable_parallelism()
4 def do_sliceplot_inline():
5     ds = yt.libyt.libytDataset()
6     slc = yt.SlicePlot(ds, "z", ("gamer", "Dens"))
7     if yt.is_root():
8         slc.save()
```

Simulation can call Python function using `libyt` API `yt_run_Function` and `yt_run_FunctionArguments`. For example, this calls the Python function `do_sliceplot_inline`:

```
yt_run_Function("do_sliceplot_inline");
```

Beside calling Python function, `libyt` also provides interactive prompt for user to update Python function, enter statements, and get feedbacks instantly.¹¹ This is like running Python prompt inside the ongoing simulation with access to data. Fig 5 describes the workflow. The root process takes user inputs and checks the syntax through compiling it to code object using `Py_CompilerString`. If error occurs, it parses the error to see if this is caused by input not done yet or a real error. If it is indeed caused by user hasn't done yet, for example, when using an `if` statement, the prompt will continue waiting for user inputs. Otherwise, it simply prints the error to inform the user. If the code can be compiled successfully, the root process broadcasts the code to every other MPI processes. Then they evaluate the code using `PyEval_EvalCode` inside the script's namespace simultaneously.

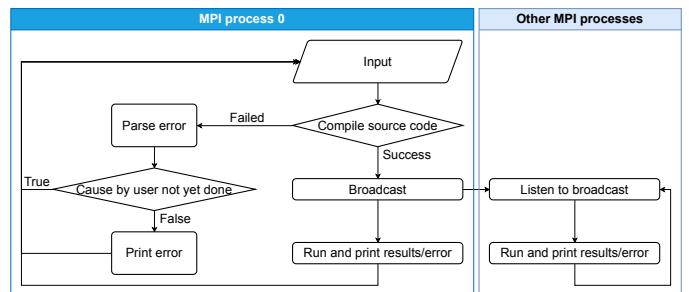


Fig. 5: The procedure shows how `libyt` supports interactive Python prompt. It takes user inputs on root process and executes Python codes across whole MPI processes. The root process handles syntax errors and distinguishes whether or not the error is caused by user hasn't done inputting yet.

Applications

`libyt` has already been implemented in `GAMER`¹² [4] and `Enzo`¹³ [5]. `GAMER` is a GPU-accelerated adaptive mesh refinement code for astrophysics. It features extremely high performance and parallel scalability and supports a rich set of physics modules. `Enzo` is a community-developed adaptive mesh refinement simulation code, designed for rich, multi-physics hydrodynamic astrophysical calculations.

Here, we demonstrate the results from `GAMER` using `libyt`, and we show how `libyt` solves the problem of limitation in disk space and improves disk usage efficiency.

¹¹. Currently, `libyt` interactive prompt only works on local machine or submitting the job to HPC platforms using interactive queue (e.g., `qsub -I` on PBS scheduler). We will support accessing through Jupyter Notebook in the future.

¹². <https://github.com/gamer-project/gamer>

¹³. <https://enzo-project.org/>

Analyzing Fuzzy Dark Matter Vortices Simulation

Fuzzy dark matter (FDM) is a promising dark matter candidate [6]. It is best described by a classical scalar field governed by the Schrödinger-Poisson equation, because of the large de Broglie wavelength compared to the mean interparticle separation. FDM halos feature a central compact solitonic core surrounded by fluctuating density granules resulting from wave function interference. Quantum vortices can form in density voids caused by fully destructive interference [7] [8]. The dynamics of these vortices in FDM halo have not been investigated thoroughly, due to the very high spatial and temporal resolution is required, which leads to tremendously huge disk space. libyt provides a promising approach for this study.

We use GAMER to simulate the evolution of an FDM halo on the Taiwania 3¹⁴. We use 560 CPU cores by launching 20 MPI processes with 28 OpenMP threads per MPI process to run the simulation. The simulation box size is 2.5×10^5 pc, covered by a 640^3 base-level grid with six refinement levels. The highest level has a maximum resolution of 6.2 pc, so that it is able to resolve the fine structure and dynamical evolution of vortices within a distance of 3200 pc from the center. To properly capture the dynamics, we aim for analyzing vortex properties with a temporal resolution of 3.5×10^{-2} Myr, corresponding to 321 analysis samples. Each simulation snapshot, including density, real part, imaginary part, gravitational potential, and AMR grid information, takes 116 GB. It will take ~ 37 TB if we do this in post-processing, which is really expensive. However, it is actually unnecessary to dump all these snapshots since our region of interest is only the vortex lines around the halo center.

We solve this by using libyt to invoke yt function covering_grid to extract a uniform-resolution grid centered at the halo center and store these grid data instead of simulation snapshots on disk. The uniform grid has dimension 1024^3 with spatial resolution 6.2 pc (i.e., the maximum resolution in the simulation), corresponding to the full extracted uniform grid width of 6300 pc. By storing only the imaginary and real parts of the wave function in single precision, each sample step now consumes only 8 GB, which is 15 times smaller than the snapshot required in post-processing.

We further analyze these uniform grids in post-processing, and do volume rendering and create animation¹⁵ using ParaView [9]. Fig 6 is the volume rendering of the result. Vortex lines and rings are manifest in the entire domain. Fig 7 shows a zoom in version of Fig 6, where reconnection of vortex lines take place. With the help of libyt, we are able to achieve a very high temporal resolution and very high spatial resolution at the same time.

Analyzing Core-Collapse Supernova Simulation

We use GAMER to simulate core-collapse supernova explosions. The simulations have been performed on a local cluster using 64 CPU cores and 4 GPUs by launching 8 MPI processes with 8 OpenMP threads per MPI process, and having two MPI processes access the same GPU. The simulations involve a rich set of physics modules, including hydrodynamics, self-gravity, a parameterized light-bulb scheme for neutrino heating and cooling with a fixed neutrino luminosity [10], a parameterized deleptonization scheme [11], an effective general relativistic potential [12], and a nuclear

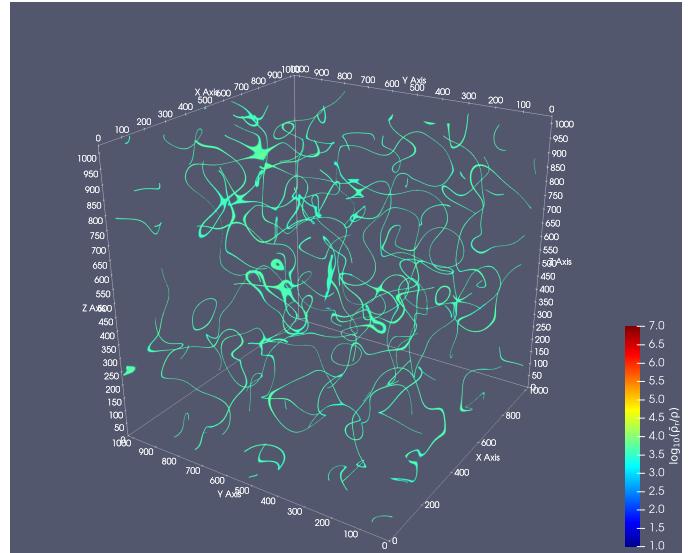


Fig. 6: Volume rendering of quantum vortices in a fuzzy dark matter halo with GAMER. Here we use libyt to extract uniform-resolution data from an AMR simulation on-the-fly, and then visualize it with ParaView in post-processing. The colormap is the logarithm of reciprocal of density averaging over radial density profile, which highlights the fluctuations and null density. Tick labels represent cell indices.

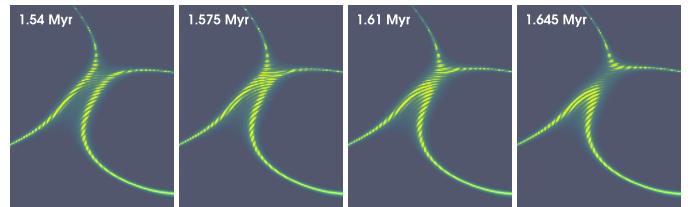


Fig. 7: Vortex reconnection process in a fuzzy dark matter halo. This is the result we get if we zoom in to one of the vortex lines in Fig 6 where reconnection of lines take place. We are able to clearly capture the dynamics, and at the same time, preserve high spatial resolution.

equation of state [13]. For the hydrodynamics scheme, we adopt the van Leer predictor-corrector integrator [14] [15], the piecewise parabolic method for spatial data reconstruction [16], and the HLLC Riemann solver [17]. The simulation box size is 2×10^4 km. The base-level grid dimension is 160^3 and there are eight refinement levels, reaching a maximum spatial resolution of ~ 0.5 km.

We use libyt to closely monitor the simulation progress during runtime, such as the grid refinement distribution, the status and location of shock wave (e.g., stalling, revival, breakout), and the evolution of the central proto-neutron star. libyt calls yt function SlicePlot to draw entropy distribution every 1.5×10^{-2} ms. Fig 8 is the output in a time step. Since entropy is not part of the variable in simulation's iterative process, these entropy data will only be generated through user-defined C function, which in turn calls the nuclear equation of state defined inside GAMER to get entropy, when they are needed by yt. libyt tries to minimize memory usage by generating relevant data only. We can combine every output figure and animate the actual simulation process¹⁶ without storing any datasets.

¹⁴. Supercomputer at the National Center for High-performance Computing in Taiwan. (<https://www.nchc.org.tw/>)

¹⁵. <https://youtu.be/6iwHzN-FsHw>

¹⁶. <https://youtu.be/6iwHzN-FsHw>

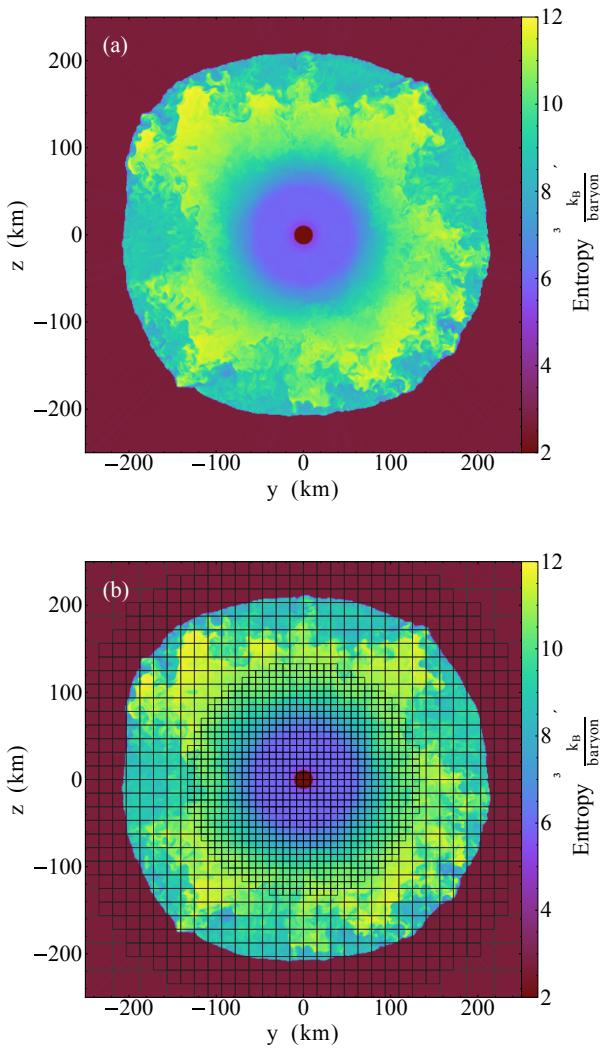


Fig. 8: Entropy distribution in a core-collapse supernova simulated by GAMER and plotted by yt function `SlicePlot` using libyt. Plot (a) shows a thin slice cut through the central proto-neutron star in the post-bounce phase. The proto-neutron star has a radius of ~ 10 km and the shock stalls at ~ 200 km. Plot (b) shows the underlying AMR grid structure, where each grid consists of 16^3 cells.

Discussions

libyt is free and open source, which does not depend on any non-free or non-open source software. Converting the post-processing script to inline script is a two-line change, which lowers the barrier of using this in situ analysis tool.

Though currently, only simulations that use AMR grid data structure are supported by libyt, we will extend to more data structure (e.g., octree, particle, unstructured mesh, etc) and hope to engage more simulations and data structures in the future.

Using libyt does not add time penalty to the analysis process, because using Python for in situ analysis and post-processing are exactly the same, except that the former one reads data from memory and the latter one reads data from disks. Fig 9 shows the strong scaling of libyt. The test compares the performance between in situ analysis with libyt and post-processing for computing 2D profiles on a GAMER dataset. The dataset contains seven adaptive mesh refinement levels with a total of 9.9×10^8

cells. libyt outperforms post-processing by $\sim 10 - 30\%$, since it avoids loading data from disk to memory. libyt and post-processing have similar deviation from the ideal scaling since libyt directly borrows the algorithm in yt. Some improvements have been made in yt, while some are still undergoing to eliminate the scaling bottleneck. But also, due to some parts cannot be parallelized, like the import of Python and the current data structure, the speed up is saturated at large number of processors and can be described by Amdahl's law.

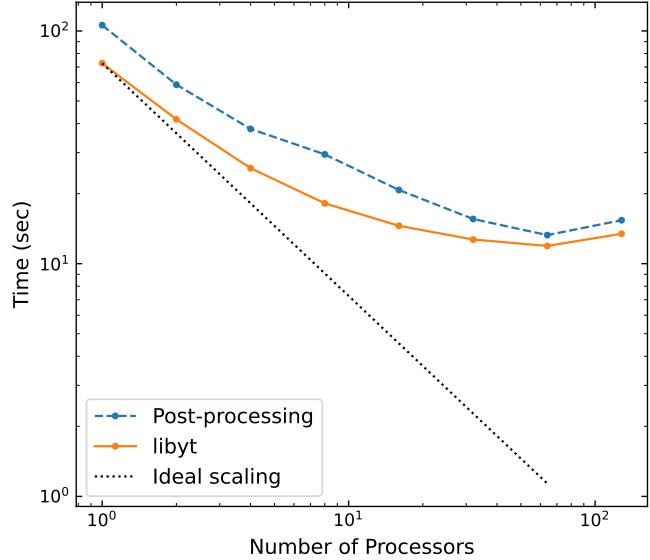


Fig. 9: Strong scaling of libyt. libyt outperforms post-processing by $\sim 10 - 30\%$ since the former avoids loading data from disk to memory. The dotted line is the ideal scaling. libyt and post-processing show a similar deviation from the ideal scaling because it directly borrows the algorithm in yt. Improvements have been made and will be made in yt to eliminate the scaling bottleneck.

libyt provides a promising solution that binds simulation to Python with minimal memory overhead and no additional time penalty. It makes analyzing large scale simulation feasible, and it can analyze the data with much higher frequency. It also reduces the barrier of heavy computational jobs written in C/C++ to use Python tools, which are normally well-developed. libyt focuses on using yt as its core analytic method, even though it can call other Python modules, and has the ability to enable back-communication of simulation information. A use case of this tool could be using yt to select data and then make it as an input source for further analysis. libyt provides us another way to interact with simulation and data.

REFERENCES

- [1] M. J. Turk, B. D. Smith, J. S. Oishi, S. Skory, S. W. Skillman, T. Abel, and M. L. Norman, "yt: A Multi-code Analysis Toolkit for Astrophysical Simulation Data," *The Astrophysical Journal Supplement Series*, vol. 192, p. 9, Jan. 2011, <https://doi.org/10.1088/0067-0049/192/1/9>.
- [2] C. R. Harris, K. J. Millman, S. J. van der Walt, R. Gommers, P. Virtanen, D. Cournapeau, E. Wieser, J. Taylor, S. Berg, N. J. Smith, R. Kern, M. Picus, S. Hoyer, M. H. van Kerwijk, M. Brett, A. Haldane, J. F. del Río, M. Wiebe, P. Peterson, P. Gérard-Marchant, K. Sheppard, T. Reddy, W. Weckesser, H. Abbasi, C. Gohlke, and T. E. Oliphant, "Array programming with NumPy," *Nature*, vol. 585, no. 7825, pp. 357–362, Sep. 2020, <https://doi.org/10.1038/s41586-020-2649-2>. [Online]. Available: <https://doi.org/10.1038/s41586-020-2649-2>

- [3] G. Van Rossum and F. L. Drake, *Python 3 Reference Manual*. Scotts Valley, CA: CreateSpace, 2009.
- [4] H.-Y. Schive, J. A. ZuHone, N. J. Goldbaum, M. J. Turk, M. Gaspari, and C.-Y. Cheng, “gamer-2: a GPU-accelerated adaptive mesh refinement code – accuracy, performance, and scalability,” *Monthly Notices of the Royal Astronomical Society*, vol. 481, no. 4, pp. 4815–4840, 09 2018, <https://doi.org/10.1093/mnras/sty2586>. [Online]. Available: <https://doi.org/10.1093/mnras/sty2586>
- [5] G. L. Bryan, M. L. Norman, B. W. O’Shea, T. Abel, J. H. Wise, M. J. Turk, D. R. Reynolds, D. C. Collins, P. Wang, S. W. Skillman, B. Smith, R. P. Harkness, J. Bordner, J.-h. Kim, M. Kuhlen, H. Xu, N. Goldbaum, C. Hummels, A. G. Krutsuk, E. Tasker, S. Skory, C. M. Simpson, O. Hahn, J. S. Oishi, G. C. So, F. Zhao, R. Cen, Y. Li, and The Enzo Collaboration, “ENZO: An Adaptive Mesh Refinement Code for Astrophysics,” *The Astrophysical Journal Supplement Series*, vol. 211, p. 19, Apr. 2014, <https://doi.org/10.1088/0067-0049/211/2/19>.
- [6] H.-Y. Schive, T. Chiueh, and T. Broadhurst, “Cosmic structure as the quantum interference of a coherent dark wave,” *Nature Physics*, vol. 10, pp. 496–499, Jul. 2014, <https://doi.org/10.1038/nphys2996>.
- [7] T. Chiueh, “Dynamical quantum chaos as fluid turbulence.” *Phys. Rev. E*, vol. 57, pp. 4150–4154, Apr 1998, <https://doi.org/10.1103/PhysRevE.57.4150>. [Online]. Available: <https://link.aps.org/doi/10.1103/PhysRevE.57.4150>
- [8] L. Hui, A. Joyce, M. J. Landry, and X. Li, “Vortices and waves in light dark matter,” *Journal of Cosmology and Astroparticle Physics*, vol. 2021, no. 1, p. 011, Jan. 2021, <https://doi.org/10.1088/1475-7516/2021/01/011>.
- [9] J. Ahrens, B. Geveci, and C. Law, *Visualization Handbook*. Burlington, MA: Elsevier Butterworth-Heinemann, 2005, ch. ParaView: An End-User Tool for Large-Data Visualization, p. 717.
- [10] S. M. Couch, “The Dependence of the Neutrino Mechanism of Core-collapse Supernovae on the Equation of State,” *The Astrophysical Journal*, vol. 765, no. 1, p. 29, Mar. 2013, <https://doi.org/10.1088/0004-637X/765/1/29>.
- [11] M. Liebendörfer, “A Simple Parameterization of the Consequences of Deleptonization for Simulations of Stellar Core Collapse,” *The Astrophysical Journal*, vol. 633, no. 2, pp. 1042–1051, Nov. 2005, <https://doi.org/10.1086/466517>.
- [12] E. P. O’Connor and S. M. Couch, “Two-dimensional Core-collapse Supernova Explosions Aided by General Relativity with Multidimensional Neutrino Transport,” *The Astrophysical Journal*, vol. 854, no. 1, p. 63, Feb. 2018, <https://doi.org/10.3847/1538-4357/aaa893>.
- [13] A. W. Steiner, M. Hempel, and T. Fischer, “Core-collapse Supernova Equations of State Based on Neutron Star Observations,” *The Astrophysical Journal*, vol. 774, no. 1, p. 17, Sep. 2013, <https://doi.org/10.1088/0004-637X/774/1/17>.
- [14] S. A. E. G. Falle, “Self-similar jets,” *Monthly Notices of the Royal Astronomical Society*, vol. 250, no. 3, pp. 581–596, 1991, <https://doi.org/10.1093/mnras/250.3.581>. [Online]. Available: [+http://dx.doi.org/10.1093/mnras/250.3.581](http://dx.doi.org/10.1093/mnras/250.3.581)
- [15] B. van Leer, “Upwind and high-resolution methods for compressible flow: From donor cell to residual-distribution schemes,” *Communications in Computational Physics*, vol. 1, pp. 192–206, 2006, <https://doi.org/10.2514/6.2003-3559>.
- [16] P. Colella and P. R. Woodward, “The piecewise parabolic method (ppm) for gas-dynamical simulations,” *Journal of Computational Physics*, vol. 54, no. 1, pp. 174–201, 1984, [https://doi.org/10.1016/0021-9991\(84\)90143-8](https://doi.org/10.1016/0021-9991(84)90143-8). [Online]. Available: <https://www.sciencedirect.com/science/article/pii/0021999184901438>
- [17] E. F. Toro, *Riemann Solvers and Numerical Methods for Fluid Dynamics. A Practical Introduction*, 3rd ed. Berlin: Springer, 2009.