

# Unit 1. Antennas and Free Space Propagation

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ECE-GY 6023. INTRODUCTION TO WIRELESS COMMUNICATIONS

PROF. SUNDEEP RANGAN

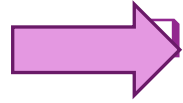
# Learning Objectives

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- ❑ Mathematically describe an EM wave:
  - Direction of motion, wavenumber, frequency, polarization, ...
- ❑ Identify radio spectrum and power levels used in common commercial wireless products
- ❑ Perform basic power calculations in dB scale
- ❑ Perform basic mathematical operations in polar coordinates
  - Conversions to cartesian coordinates, rotations, integrals, averages, ...
- ❑ Use tools from MATLAB to compute and plot key antenna parameters
  - Directivity, gain, efficiency, ...
- ❑ Compute received power in an angular region using the radiation density and intensity.
- ❑ Compute the free-space path loss using Friis Law
- ❑ Derive Friis Law

# Outline

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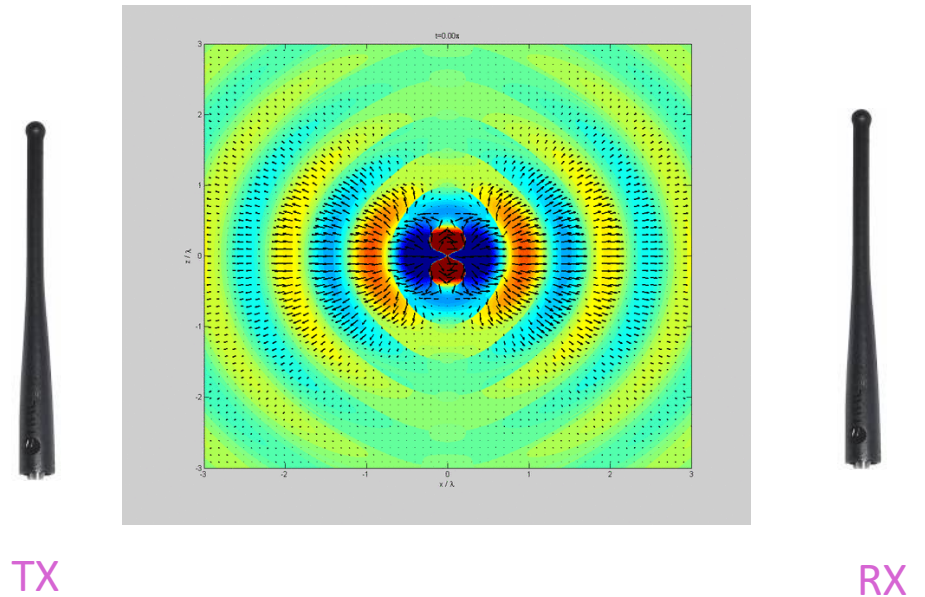
## Basics of Electromagnetic Waves

- ☐ Power and Bandwidth of Signals
- ☐ Basics of Antennas
- ☐ Free Space Propagation
- ☐ Frames of Reference and Rotations



# Electric and Magnetic Forces

- ❑ Two closely related forces:
  - **Electric**: Forces between charged particles
  - **Magnetic**: Forces between moving charged particles
- ❑ EM forces operate at a distance
  - Current in one location  $\Rightarrow$  current in another location
- ❑ Enables **communication**
  - Modulate current at a TX
  - Currents create EM fields in space
  - Detect modulation at a RX



# Electric and Magnetic Vector Fields

- E and M forces represented by a **vector field**
  - Changes with position  $\mathbf{r} = (x, y, z)$  and time  $t$
  - Force strength has a direction and magnitude

- **Electric Field:**  $\mathbf{E}(\mathbf{r}, t)$ 
  - Units: N/C (force / unit charge)

- **Magnetic field:**  $\mathbf{B}(\mathbf{r}, t)$ 
  - Units: in N/(Am) = Teslas (force / unit charge / velocity)

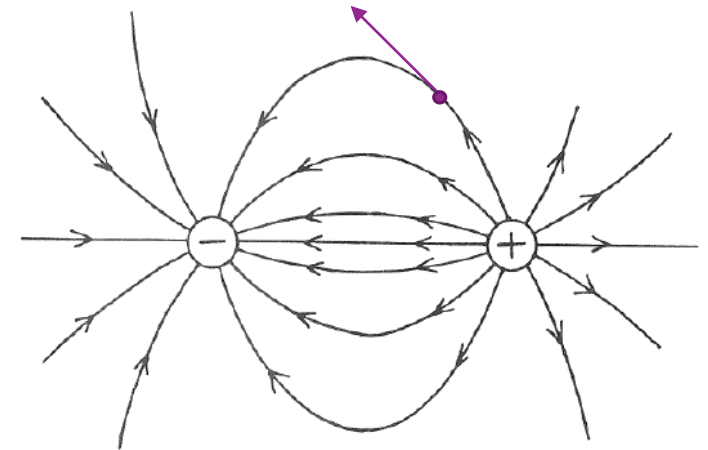
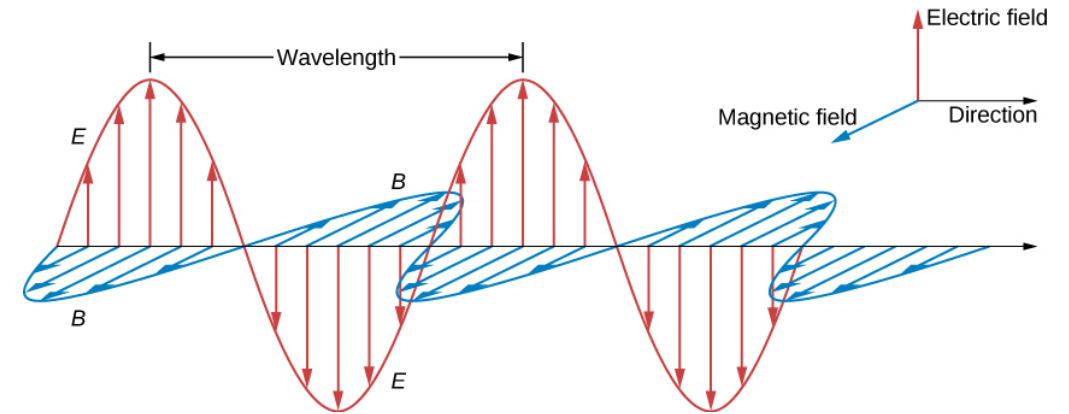


FIGURE 2.4 Electric field lines begin and end on charges.

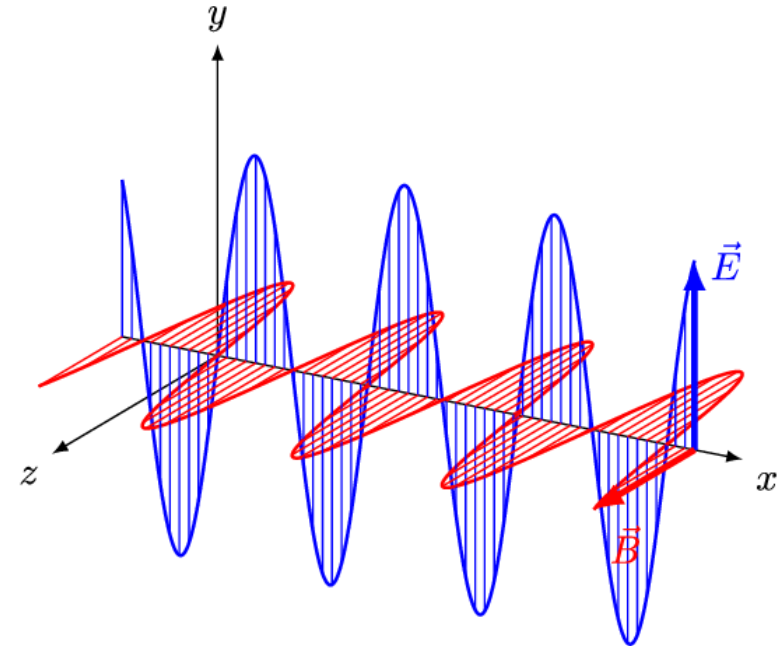
# Plane Waves

- EM field governed by Maxwell's equations
- In free space, all solutions can be decomposed into plane waves
- EM plane wave at position  $\mathbf{r} = (x, y, z)$ 
  - $\mathbf{E}(\mathbf{r}, t) = E_0 \mathbf{e}_y \cos(2\pi(ft + \lambda^{-1}x) + \phi)$
  - $\mathbf{B}(\mathbf{r}, t) = B_0 \mathbf{e}_z \cos(2\pi(ft + \lambda^{-1}x) + \phi)$
- Key constraints:
  - $\mathbf{E}(\mathbf{r}, t)$  is always perpendicular to  $\mathbf{B}(\mathbf{r}, t)$
  - $B_0 = (1/c)E_0$
  - $c = \lambda f = \text{speed of light}$



# Visualizing Plane Waves

- EM plane wave at position  $\mathbf{r} = (x, y, z)$ 
  - $\mathbf{E}(\mathbf{r}, t) = E_0 \mathbf{e}_y \cos(2\pi(ft + \lambda^{-1}z) + \phi)$
  - $\mathbf{B}(\mathbf{r}, t) = B_0 \mathbf{e}_z \cos(2\pi(ft + \lambda^{-1}z) + \phi)$
- At any given position  $\mathbf{r}$ :
  - $E$  and  $B$  fields vary sinusoidally with frequency  $f$
  - Maximum amplitudes  $E_0$  and  $B_0$
  - Phase  $\phi$
- For a fixed time  $t$ , along direction  $x$ 
  - $E$  and  $B$  fields vary sinusoidally with wavelength  $\lambda = \frac{c}{f}$
- Can be viewed as traveling along direction



# Plane Wave Direction of Motion

□ Direction of motion = direction of “energy flux”

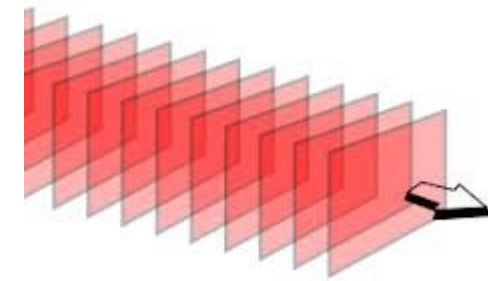
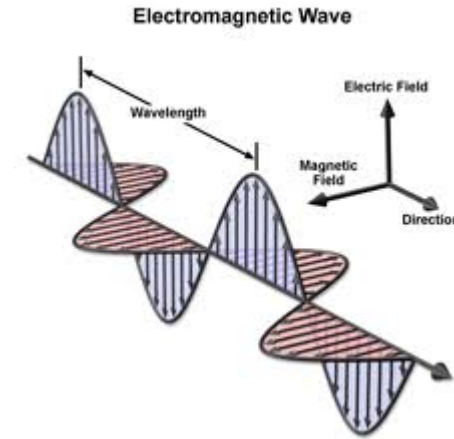
□ “Poynting” vector:

$$\mathbf{S} = \frac{1}{\mu} \mathbf{E} \times \mathbf{B} = \frac{|E_0|^2}{c\mu} \cos^2(2\pi(ft - \lambda^{-1}z)) \mathbf{e}_x$$

- $\mathbf{e}_x$  = direction of motion
- $\eta = c\mu$  = characteristic impedance
- Vacuum:  $\eta = \eta_0 \approx 377\Omega$

□ Represents energy flux

- Surface integral of  $\mathbf{S}$  = energy transferred into region
- Energy consumed =  $\nabla \cdot \mathbf{S}$
- Units =  $W/m^2$





# Polarization

❑ **Polarization**: Orientation of E-field relative to direction of motion

❑ **Linearly polarized**: Constant orientation

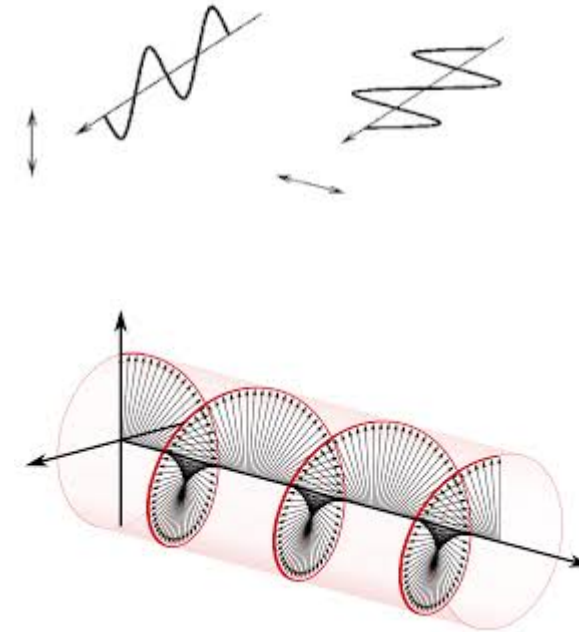
- Vertical:  $\mathbf{E}(\mathbf{r}, t) = E_0 \mathbf{e}_x \cos(\omega t + kz)$
- Horizontal:  $\mathbf{E}(\mathbf{r}, t) = E_0 \mathbf{e}_y \cos(\omega t + kz)$
- Angular frequency  $\omega = 2\pi f$  and wave number  $k = \frac{2\pi}{\lambda}$

❑ **Two degrees of freedom**:

- Consider any plane wave in some direction
- Can be decomposed as V + H

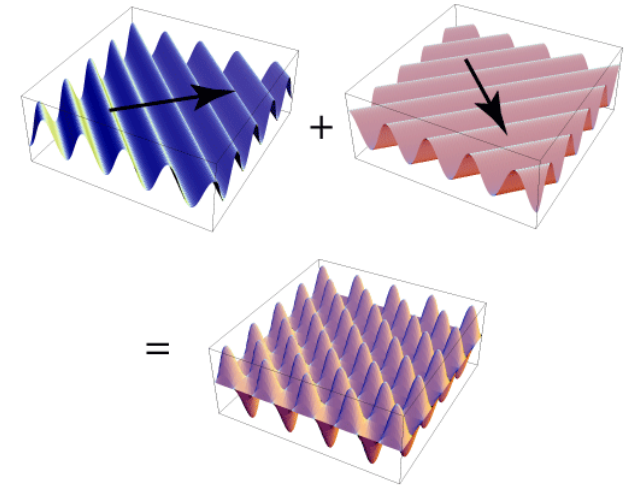
❑ **Also, circularly polarized**

- Sum of V and H that are out of phase
- $E_0[\mathbf{e}_x \cos(\omega t + kz) \pm \mathbf{e}_y \sin(\omega t + kz)]$
- Called left hand and right hand



# Plane Wave Decomposition

- ❑ Every electric field is a linear combination of plane waves
- ❑ Each plane wave in the decomposition has:
  - Frequency
  - Direction of motion
  - Gain, Phase
  - One of two polarization
- ❑ Decomposition can be found from a 4D Fourier transform
  - $\mathbf{E}(x, y, z, t) \Rightarrow \hat{E}_V(k_x, k_y, k_z, f)$  and  $\hat{E}_H(k_x, k_y, k_z, f)$
  - Converts time + space  $\Rightarrow$  wavenumber and frequency
  - Note that there are two polarization components
- ❑ This decomposition is used in many EM solvers
  - And your EM class if you take it



# In-Class Problem

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## Problem 1

Suppose that an EM-plane wave:

- Power flux density is  $1 \text{ nW/m}^2$
- Freq = 2.3 GHz

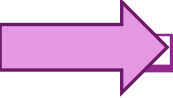
Print the:

- Maximum E-field value
- Wavelength. You may use the `physconst('Lightspeed')` command to get the speed of light.

Make sure you print the units.

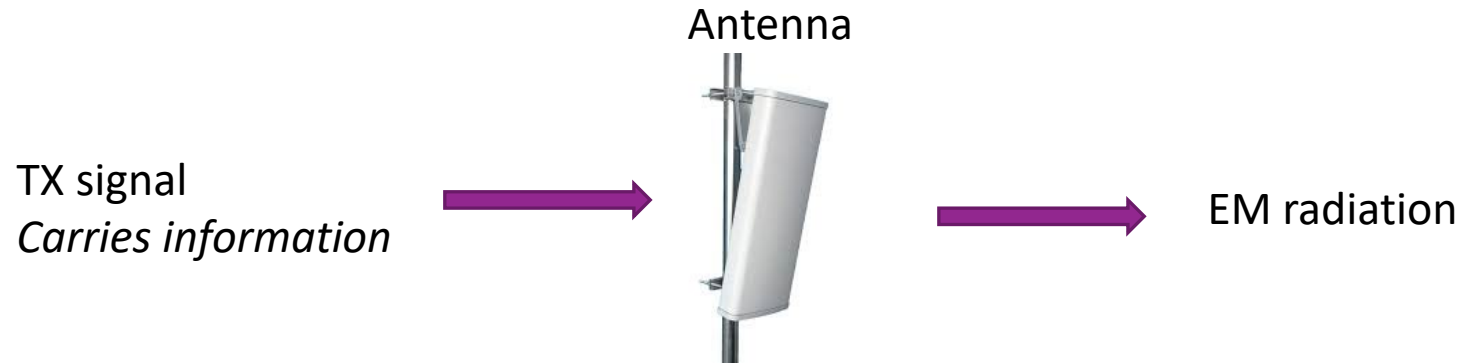
# Outline

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- ☐ Basics of Electromagnetic Waves
-  ☐ Power and Bandwidth of Signals
- ☐ Basics of Antennas
- ☐ Free Space Propagation
- ☐ Frames of Reference and Rotations

# Signals for Communication

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- ❑ **Signal:** Any quantity that varies in time
  - Continuous, discrete, complex, real, ...
- ❑ **Signals for wireless communications:**
  - Modulate an information bearing signal to a signal in the EM radiation
- ❑ **Three key characteristics of the signal:** power, bandwidth, center frequency

# Energy and Power of Signals

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- ❑ Consider a scalar-valued, continuous-time signal  $x(t)$
- ❑ Define **instantaneous power**:  $|x(t)|^2$
- ❑ Typically  $|x(t)|^2$  this is proportional to the actual power
  - Ex 1: For a voltage, power =  $\frac{|V(t)|^2}{R}$
  - Ex 2: For an EM plane wave , power flux =  $\frac{|E(t)|^2}{\eta}$
- ❑ **Energy**:
  - $E_x = \int_{-\infty}^{\infty} |x(t)|^2 dt$
  - Signal is called an “energy signal” if  $E_x < \infty$
- ❑ **Power**:
  - $P_x = \lim_{T \rightarrow \infty} \frac{1}{T} \int_{-T/2}^{T/2} |x(t)|^2 dt$
  - Energy per unit time
  - Signal is called a “power signal” if limit  $P_x$  exists and is finite

# Power: Linear and Decibel scale

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## □ Linear scale units

- Power measured in Watts (W) or mW
- Power values in W or mW called *linear scale*
- Energy measured in Joules (J) or mJ

## □ Power often measured in dB scale:

- $P_{\text{dBW}} = 10\log_{10}(P_{\text{W}} / 1\text{W})$
- $P_{\text{dBm}} = 10\log_{10}(P_{\text{mW}} / 1\text{mW})$
- $E_{\text{dBmJ}} = 10\log_{10}(E_{\text{mJ}} / 1\text{mJ})$
- dB scale is preferred since wireless signals have very large range

## □ Example: $P = 250 \text{ mW}$ (typical max mobile transmit power)

- In dBW:  $P = 10\log_{10}(0.25\text{W} / 1\text{W}) = -6 \text{ dBW}$
- In dBm:  $P = 10\log_{10}(250\text{mW} / 1\text{mW}) = 24 \text{ dBm}$

# Some important dB values

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□ Some conversions don't need a calculator:

- $10\log_{10}(2) = 3$  [Most important: Doubling power = 3dB]
- $10\log_{10}(3) = 4.7 \sim 5$
- $10\log_{10}(10) = 10$

□ You can cascade these.

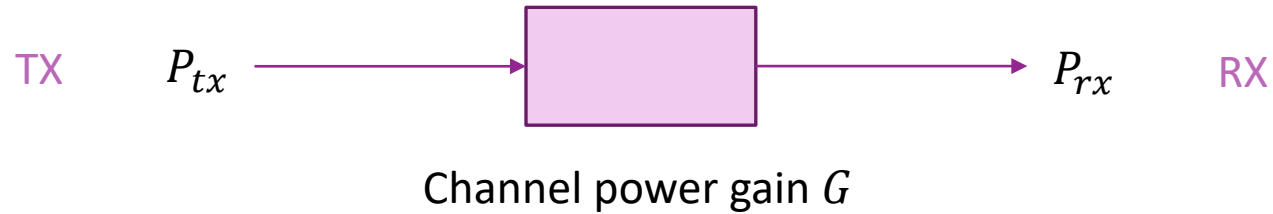
□ Ex: What is 50 mW in dBm?

□ Ans:

$$\begin{aligned} 10\log_{10}(50) &= 10\log_{10}(10^2/2) \\ &= (2)10\log_{10}(10) - 10\log_{10}(2) = 2(10) - 3 = 17 \text{ dBm} \end{aligned}$$



# Gain and Loss in dB



**Linear scale:** Gain is multiplication

$$P_{rx} = G P_{tx}$$

- $P_{tx}$ ,  $P_{rx}$ : TX and RX power in W or mW
- $G$ : Power gain (dimensionless)
- $G > 1$ : Gain (e.g. amplifier)
- $G < 1$ : Loss (e.g. propagation, attenuator, ...)

**dB scale:** Gain is addition

$$P_{rx} = G + P_{tx}$$

- $P_{tx}$ ,  $P_{rx}$ : TX and RX power in dBW or dBm
- $G$ : Power gain in dB
- $G > 0$ : Gain
- $G < 0$ : Loss

# Typical Wireless Power Transmit Levels

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- ❑ 100 kW = 80 dBm: Typical FM radio transmission with 50 km radius
- ❑ 1 kW = 60 dBm: Microwave oven element (most of this doesn't escape)
- ❑ ~300 W = 55 dBm: Geostationary satellite
- ❑ 250 mW = 24 dBm: Cellular phone maximum power (class 2)
- ❑ 200 mW = 23 dBm: WiFi access point
- ❑ 32 mW = 15 dBm: WiFi transmitter in a laptop
- ❑ 4 mW = 6 dBm: Bluetooth 10 m range
- ❑ 1 mW = 0 dBm: Bluetooth, 1 m range

# Example: Power and Time Calculation

- ❑ **Problem:** Suppose that: TX power= 17 dBm, path loss= 80 dB
- ❑ What is the RX power in dBm and mW?
  - dB scale:  $P_{rx} = P_{tx} - PL = 17 - 80 = -63$  dBm
  - Linear scale:  $-63 = -60 - 3 \Rightarrow P_{rx} = (0.5)10^{-6}$  mW = 0.5 nW
- ❑ What is the energy received in  $T = 4$  us (Symbol time for an 802.11g OFDM system):
  - Linear scale:  $E_{rx} = P_{rx}T = (0.5)10^{-6}4(10)^{-6} = 2(10)^{-12}$  mJ
  - dB Scale: Converting  $2(10)^{-12}$  mJ to dB:  $E_{rx} = -120 + 3 = -117$  dBmJ
- ❑ Note unit: dBmJ = Energy relative to 1 mJ
- ❑ Can also compute the energy directly without converting to linear scale:
  - $E_{rx} = P_{rx} + 10 \log_{10}(T) = P_{tx} - PL + 10 \log_{10}(T) = 17 - 80 + 6 - 60 = -117$  dBmJ

# Bandwidth and Carrier Frequency

## ❑ Power density spectrum:

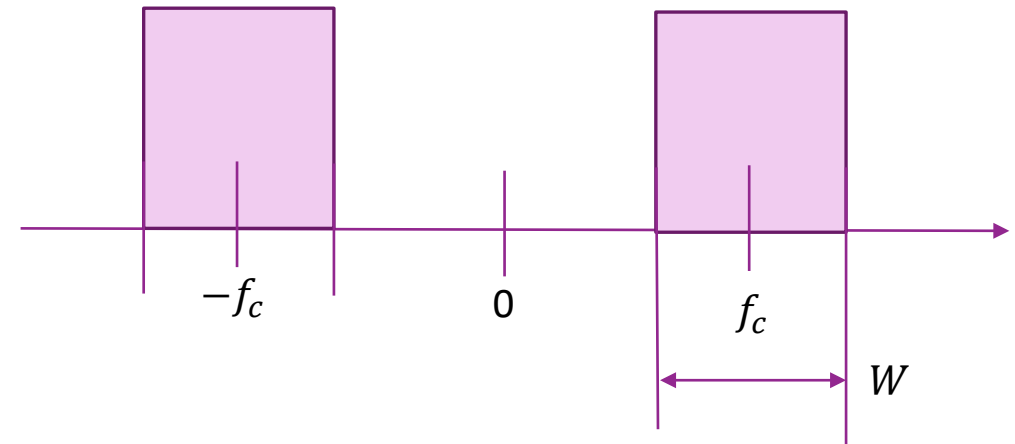
- PSD measures the power per unit frequency
- Indicates range of frequencies of the corresponding EM wave
- Measured by a spectrum analyzer



## ❑ Two key parameters for RF signals:

- Carrier or center frequency,  $f = f_c$
- Bandwidth  $W$

## ❑ Note for a real-valued signal: Always two images



# Example: PSD Calculation

❑ **Problem:** Suppose that: TX power= 17 dBm, path loss= 80 dB, bandwidth = 16.25 MHz

- Assume power is transmitted uniformly over the bandwidth
- Bandwidth is the occupied BW for an 802.11g signal

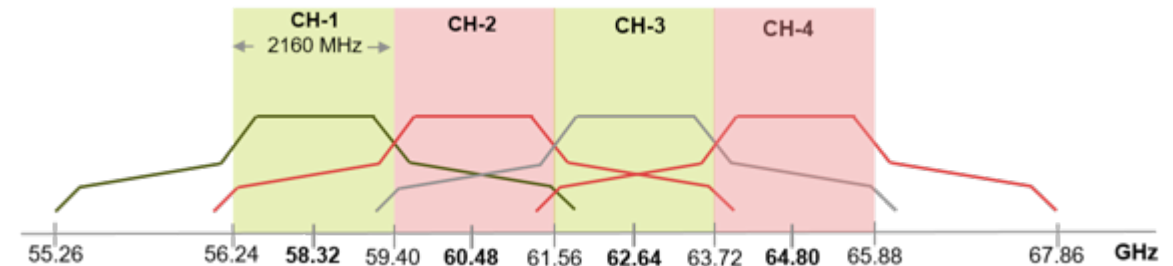
❑ What is the RX power in dBm in a 5 MHz bandwidth:

- In linear scale, RX PSD =  $S = \frac{P_{rx}}{W_{tot}}$ ,  $W_{tot} = 16.25$  MHz
- RX power in  $W_0 = 5$  MHz is  $P_0 = SW_0 = \frac{P_{rx}W_0}{W_{tot}}$
- In dB scale:

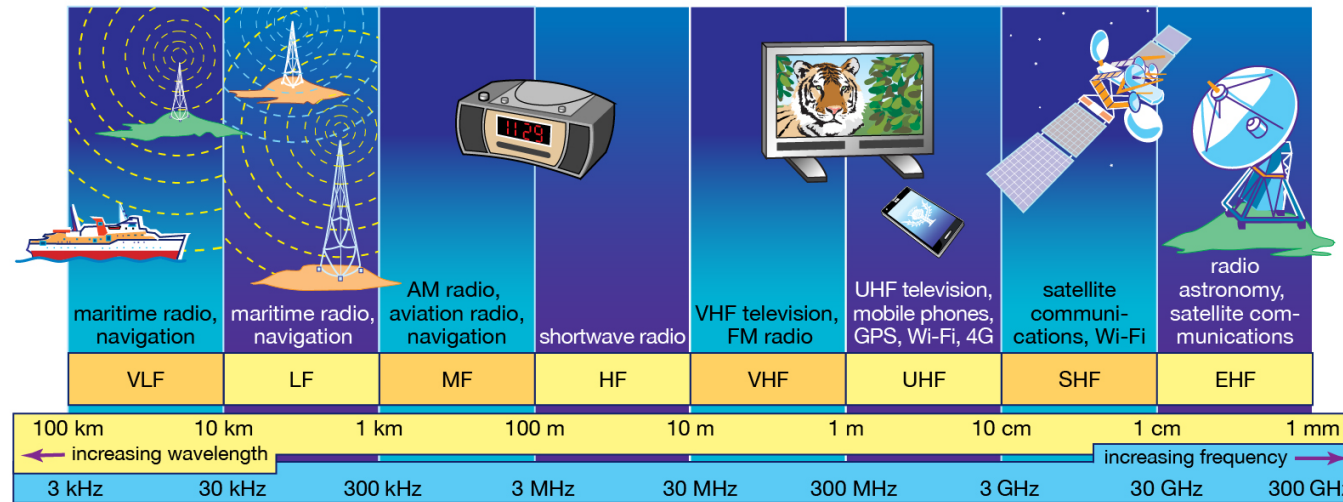
$$\begin{aligned} P_0 &= P_{rx} + 10 \log_{10} \left( \frac{W_0}{W_{tot}} \right) = P_{tx} - PL + 10 \log_{10} \left( \frac{W_0}{W_{tot}} \right) \\ &= 17 - 80 + 10 \log_{10} \left( \frac{5}{16.25} \right) = -68.1 \text{ dBm} \end{aligned}$$

# Importance of Bandwidth

- ❑ Data rate generally scales linearly in bandwidth
  - If the transmit power and bandwidth increase by  $N \Rightarrow$  the communication rate increase by  $N$
  - We will see this in detail later
- ❑ Ex: Compare GSM (2G) and LTE (4G)
  - Single channel of GSM system = 200 kHz
  - Single channel of LTE = 20 MHz
  - If power scales sufficiently, LTE would in general have 100x data rate
  - LTE, in fact, can have even more capacity due to other improvements
- ❑ Figure to the right: 802.11ad channels
  - The channels are > 2 GHz

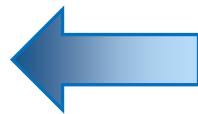


# Radio Spectrum



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Less bandwidth  
Greater Range



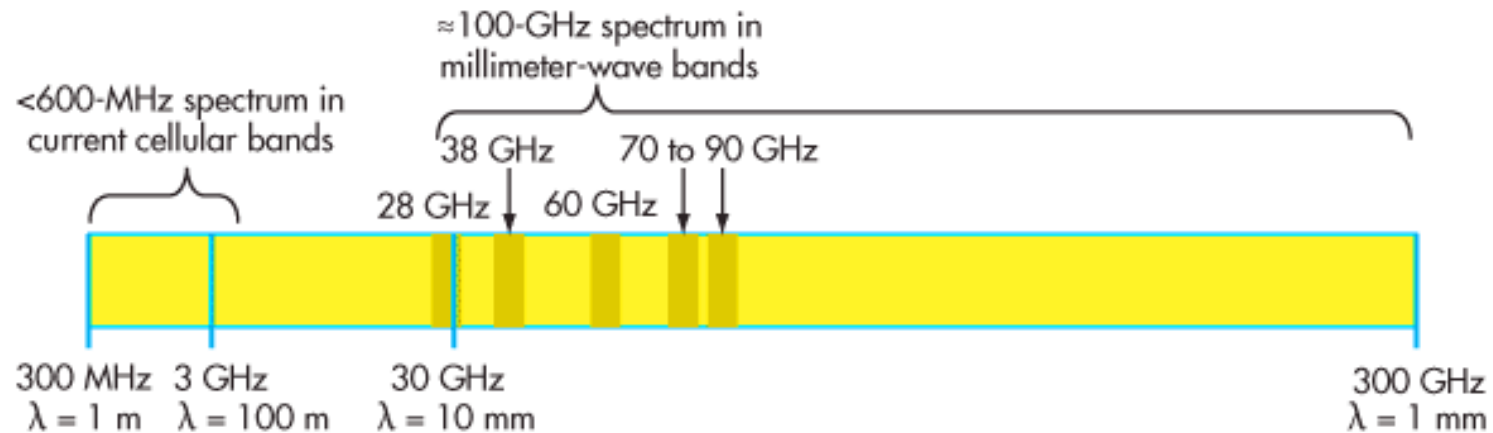
More bandwidth  
Less Range



Image: Britanica, <https://www.britannica.com/science/radio-frequency-spectrum>

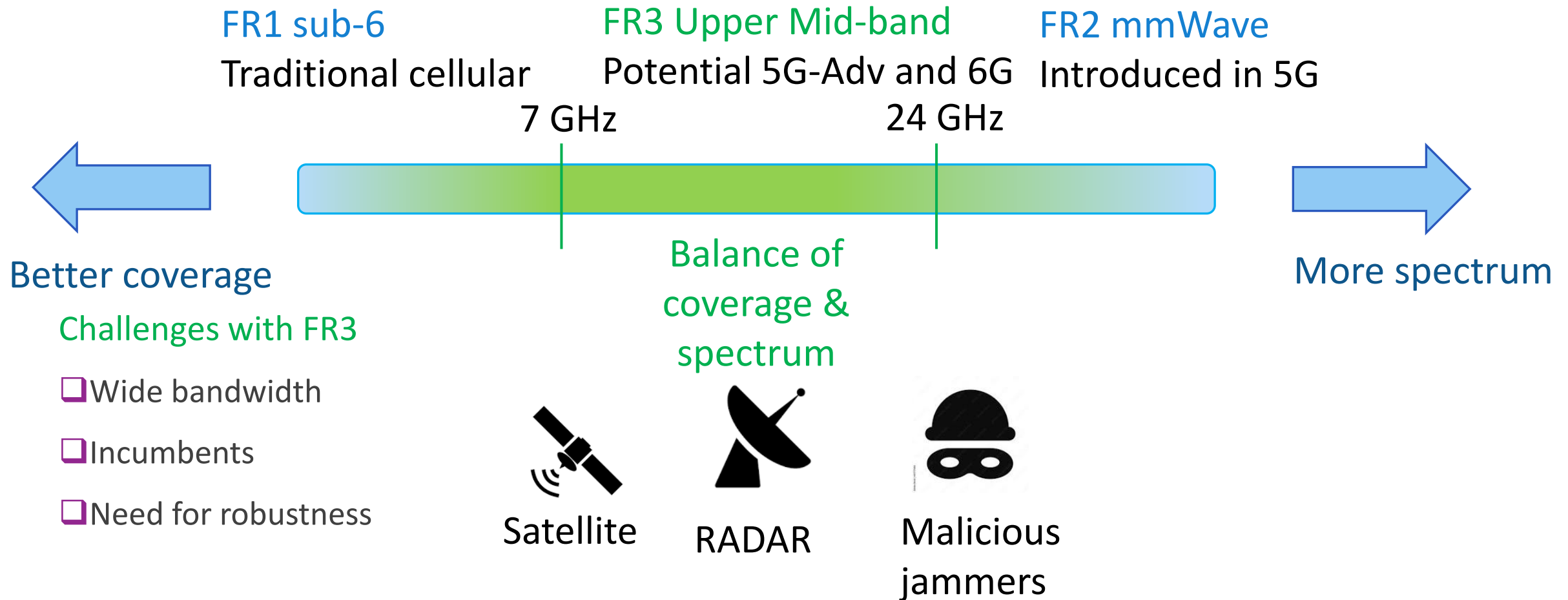
# Millimeter Wave Bands

- ❑ New bands for 5G
  - 100x more bandwidth than conventional bands below 6 GHz
  - Bands at 28 GHz and 38 GHz opened up by FCC
  - 5G systems operating have just started deployments





# Next Frontier: Upper Mid-Band



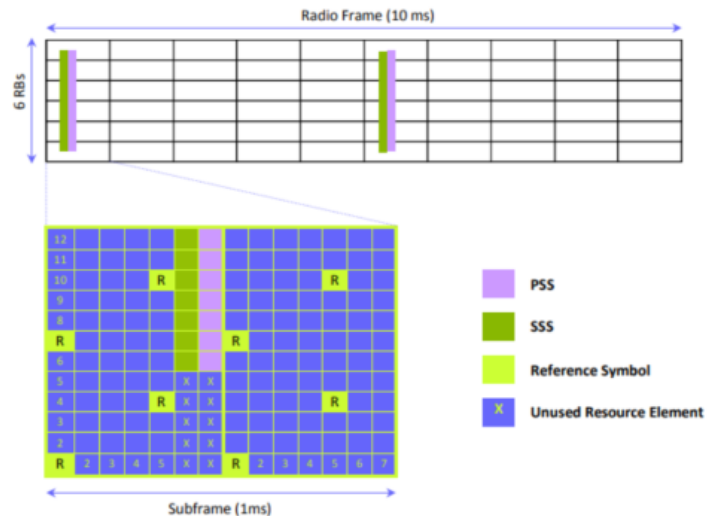
# In-Class Exercise

## Problem 2: Computing Power in a LTE PSS

In an LTE system, each base station (called eNB in 3GPP terminology) periodically transmits a Primary Synchronization Signal (PSS) so that mobiles can detect a mobile (called a UE or user equipment) can detect the base station. The PSS occupies:

- In frequency: 72 sub-carriers at 15 kHz per sub-carrier
- In time: One OFDM symbol = 2048 samples at 30.72 Ms/s


The following diagram shows the transmission of the OFDM:



Suppose the eNB has a total transmit power of 43 dBm uniformly over 20 MHz. The path loss between the eNB and UE is 100 dB. What is the received energy per PSS? Print your answer in dBmJ.

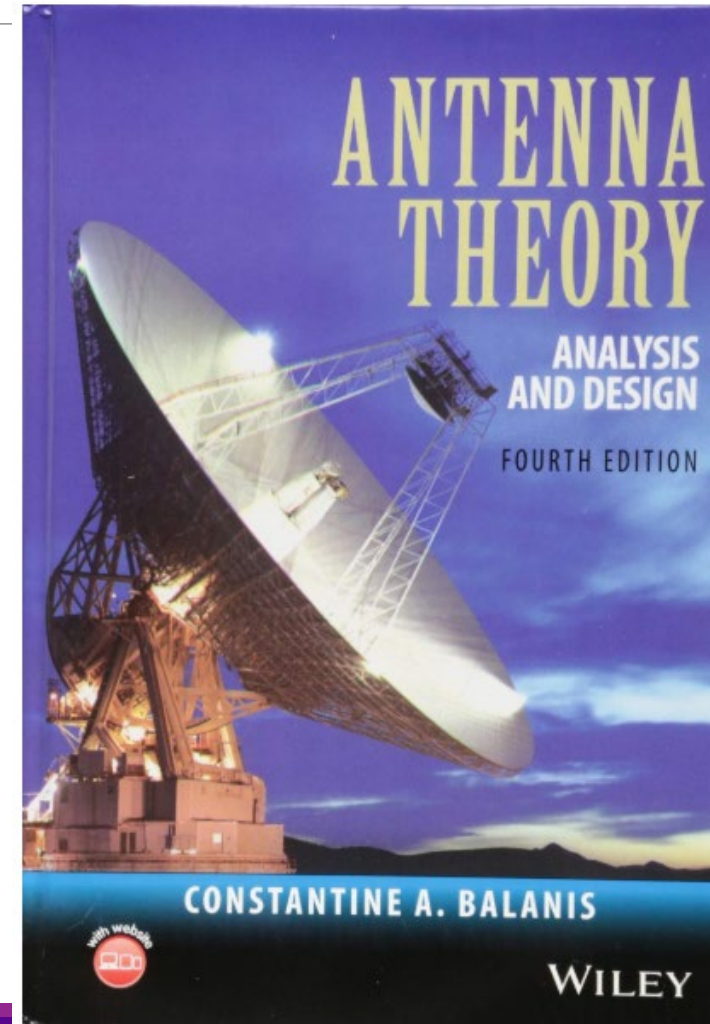
# Outline

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- ☐ Basics of Electromagnetic Waves
- ☐ Power and Bandwidth of Signals
-  Basics of Antennas
- ☐ Free Space Propagation
- ☐ Frames of Reference and Rotations

# Excellent Text for Antennas

- ❑ This lecture is based on classic text
  - Balanis, “Antenna Theory”
  - Most of the figures here are from this text
- ❑ If you want to learn more, study the text:
  - Provides full EM theory view
  - Many excellent problems and examples
  - Designed for RF engineers
- ❑ We will use only a small portion here
- ❑ Take an EM class for more!

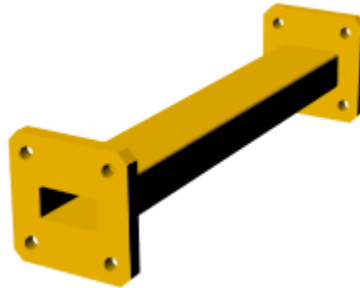


# Waveguides and Transmission Lines

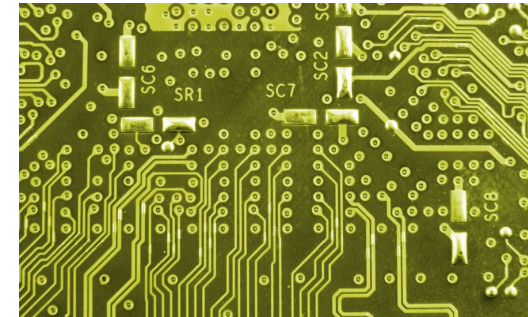
- ❑ **Transmission lines** and **waveguides**: Any structure to guide waves with minimal loss
- ❑ Some texts:
  - Transmission lines refer to conductors and waveguides to hollow structures
- ❑ Many examples



Coaxial cable



Waveguide



PCB traces

Microstrip: External layer

Stripline: Internal layer

# Antenna

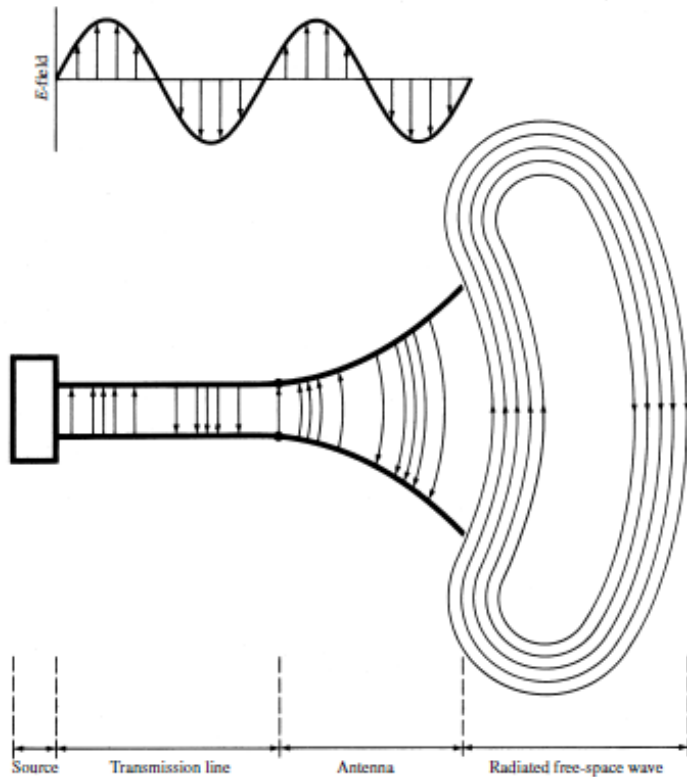


Figure 1.1 Antenna as a transition device.

- ❑ Transmit antenna: Radiates electromagnetic waves
- ❑ Converts signals:
  - From guided signals in transmission lines to
  - To radiation in free space
- ❑ Receive antenna: Collects EM wave



USRP with four vertical antennas

# Spherical Coordinates

❑ Radiation patterns are often given in spherical coordinates

❑ **Polar coordinates:**  $(\varphi, \theta, r)$

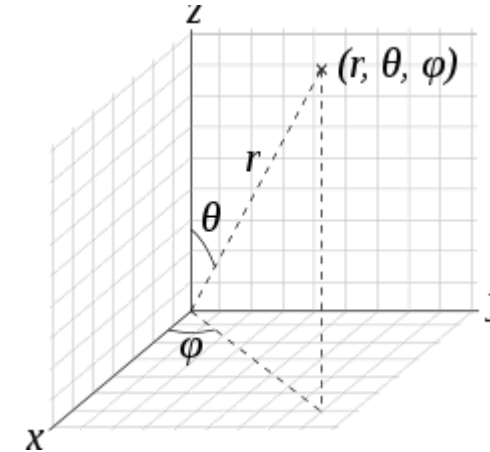
- $\varphi \in [-\pi, \pi]$ : **Azimuth**, counter-clockwise angle in xy plane
- $\theta = \theta_{inc} \in [0, \pi]$ : **Inclination** angle from z axis
- $r \geq 0$ : **Radius** from origin

❑ Wireless sometimes uses **elevation form:**  $(\varphi, \theta_{el}, r)$

- Use  $\theta_{el} = \frac{\pi}{2} - \theta_{inc} \in [\frac{\pi}{2}, \frac{\pi}{2}]$
- Measures angle from xy-plane
- Most antenna and math texts use polar form
- But MATLAB antenna toolbox uses elevation form

❑ Remember right hand rule!

Polar coordinates



**Spherical (polar form)  $\Leftrightarrow$  Cartesian**

$$r = \sqrt{x^2 + y^2 + z^2},$$

$$\varphi = \arctan \frac{y}{x},$$

$$\theta = \arccos \frac{z}{\sqrt{x^2 + y^2 + z^2}},$$

$$x = r \sin \theta \cos \varphi,$$

$$y = r \sin \theta \sin \varphi,$$

$$z = r \cos \theta.$$

# Spherical Vector Fields

□ A **vector field**: vector function of space  $\mathbf{p} = (x, y, z)$

$$\mathbf{A}(\mathbf{p}) = A_x(\mathbf{p})\mathbf{e}_x + A_y(\mathbf{p})\mathbf{e}_y + A_z(\mathbf{p})\mathbf{e}_z$$

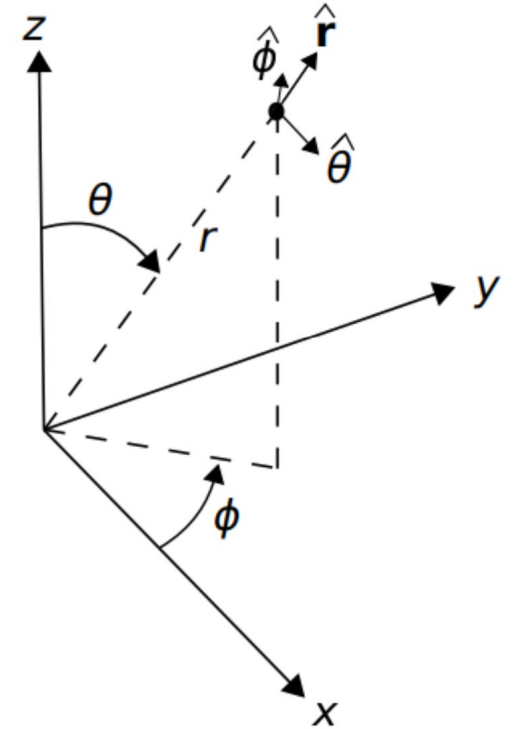
□ In **spherical coordinates**:  $\mathbf{q} = (r, \theta, \phi)$

$$\mathbf{A}(\mathbf{q}) = A_r(\mathbf{q})\mathbf{e}_r + A_\theta(\mathbf{q})\mathbf{e}_\theta + A_\phi(\mathbf{q})\mathbf{e}_\phi$$

◦ May also write  $(\hat{\mathbf{r}}, \hat{\boldsymbol{\theta}}, \hat{\boldsymbol{\phi}})$  for the unit vectors

□ Orthogonal transformation:

$$\begin{bmatrix} \hat{\mathbf{r}} \\ \hat{\boldsymbol{\theta}} \\ \hat{\boldsymbol{\phi}} \end{bmatrix} = \begin{bmatrix} \sin \theta \cos \phi & \sin \theta \sin \phi & \cos \theta \\ \cos \theta \cos \phi & \cos \theta \sin \phi & -\sin \theta \\ -\sin \phi & \cos \phi & 0 \end{bmatrix} \begin{bmatrix} \hat{\mathbf{x}} \\ \hat{\mathbf{y}} \\ \hat{\mathbf{z}} \end{bmatrix}$$





# Spherical Coordinates in MATLAB

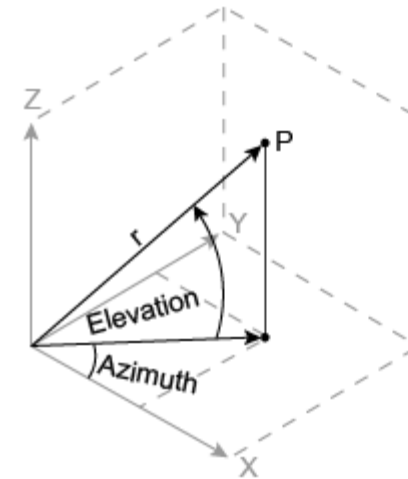
## ❑ Conversion between spherical and cartesian

```
% Generate four random points in 3D
X = randn(3,4);

% Compute spherical coordinates of a matrix of points
% Note these are in radians!
[az, el, rad] = cart2sph(X(1,:), X(2,:), X(3,:));

% Convert back
[x,y,z] = sph2cart(az,el,rad);
Xhat = [x; y; z];
```

```
x = r .* cos(elevation) .* cos(azimuth)
y = r .* cos(elevation) .* sin(azimuth)
z = r .* sin(elevation)
```



## ❑ Conversion to a coordinate system

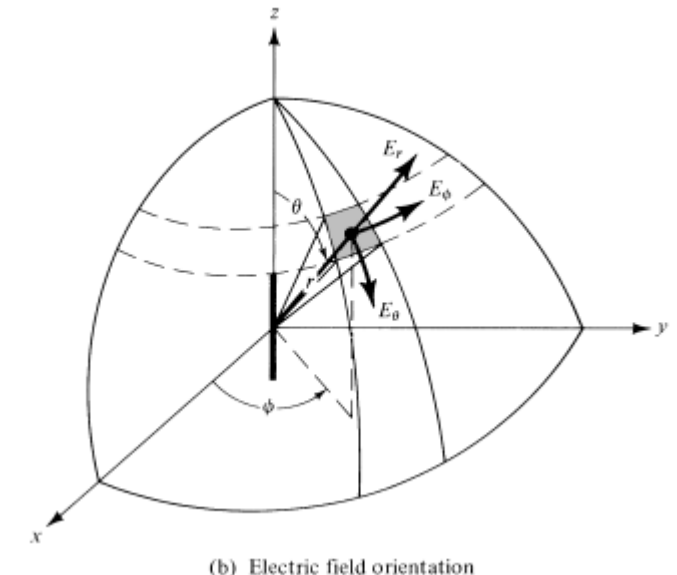
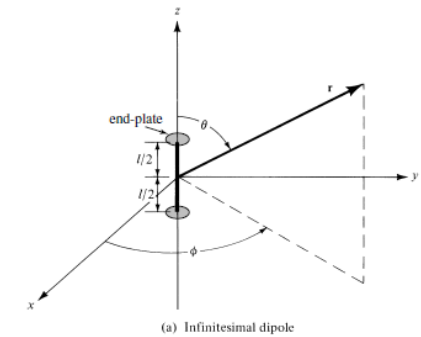
```
%% Conversion to a new frame of reference

% Angles of new frame of reference
% Note these are in degrees!
az1 = 0;
el1 = 45;

% Rotate to the new frame of reference
% This takes row vectors!
X1 = cart2sphvec(X,az1,el1);
```

# Fields from an Infinitesimal Current Source

- ❑ Consider infinitesimal dipole of length  $\ell$ 
  - Current is  $\mathbf{I} = I_0 \mathbf{e}_z$  at frequency  $\omega$
- ❑ It can be shown that in the far-field (see Balanis):
  - Magnetic field  $\mathbf{H} = \frac{jI_0 k \ell \sin \theta}{4\pi r} e^{-kr} \mathbf{e}_\phi$  (direction  $\mathbf{I} \times \mathbf{e}_r$ )
  - Electric field  $\mathbf{E} = \frac{j\eta I_0 k \ell \sin \theta}{4\pi r} e^{-kr} \mathbf{e}_\theta$ ,  $k = \frac{2\pi}{\lambda}$  (direction  $\mathbf{H} \times \mathbf{e}_r$ )
  - Poynting vector:  $\mathbf{S} = \frac{\eta}{8} \left| \frac{I_0 \ell}{\lambda} \right|^2 \frac{\sin^2 \theta}{r^2} \mathbf{e}_r$
- ❑ Notes:
  - $\theta$  is inclination angle in above formulae
  - $E$  and  $H$  decay as  $1/r$  and are tangent to the sphere
  - $S$  decays at  $1/r^2$  and is outward normal to the sphere
- ❑ The fields for any antenna:
  - Can be computed by integrating the current density



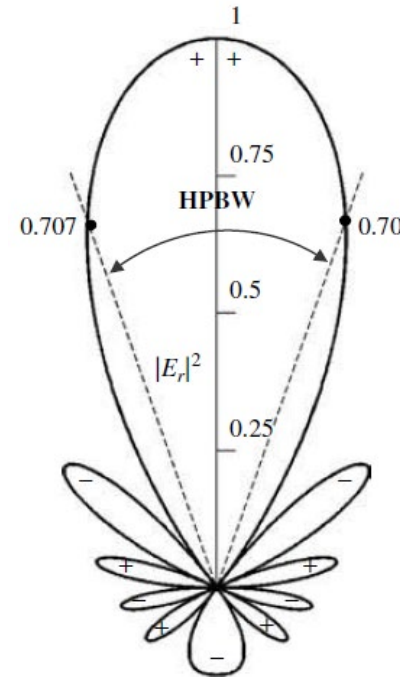
# Radiation Patterns

□ Antenna radiation typically shown via a **pattern**

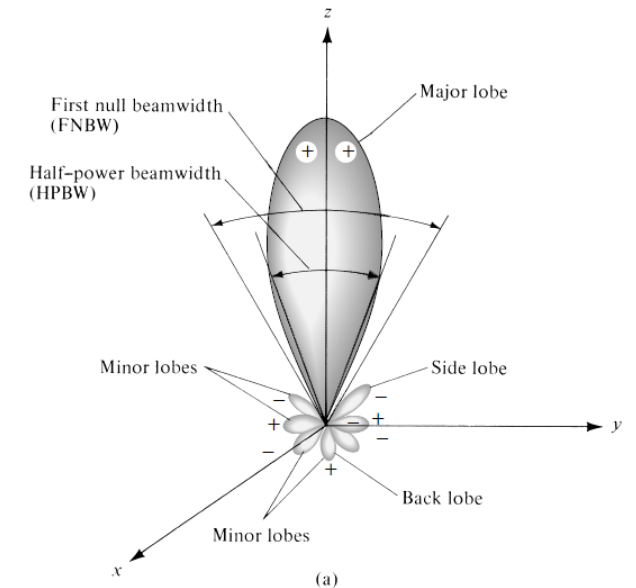
- Value of **scalar** as a function of **position**
- Antenna usually at origin
- Orientation of the antenna is important

□ Many possible quantities:

- Power, electric field, ...
- Normalized or un-normalized
- Can be 2D or 3D



2D



3D

# Radiation Density

- Each infinitesimal current source contributes Poynting vector:

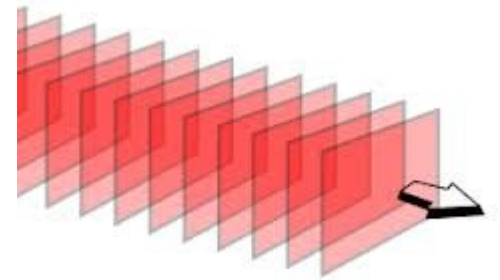
$$\mathbf{S} = \frac{\eta}{8} \left| \frac{I_0 \ell}{\lambda} \right|^2 \frac{\sin^2 \theta}{r^2} \mathbf{e}_r = \frac{1}{2\eta} \|\mathbf{E}\|^2 \mathbf{e}_r$$

- Integrating the currents across the antenna we have (in far-field):

$$\mathbf{S} = W \mathbf{e}_r$$

- Radiation density:**  $W = W(r, \theta, \phi) = \frac{1}{2\eta} |E(r, \theta, \phi)|^2 = \text{radiation density}$

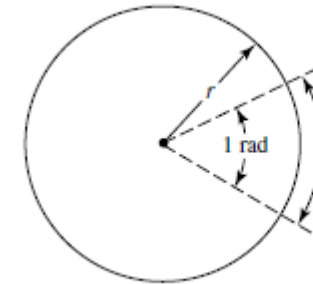
- Power radiating outward from the sphere
- Units  $W/m^2$
- This is a function of position  $W(r, \theta, \phi)$



# Radians and Steradians

## □ Radian:

- Circle of radius one
- Angle for unit length on circumference
- $2\pi$  radians in the circle



(a) Radian

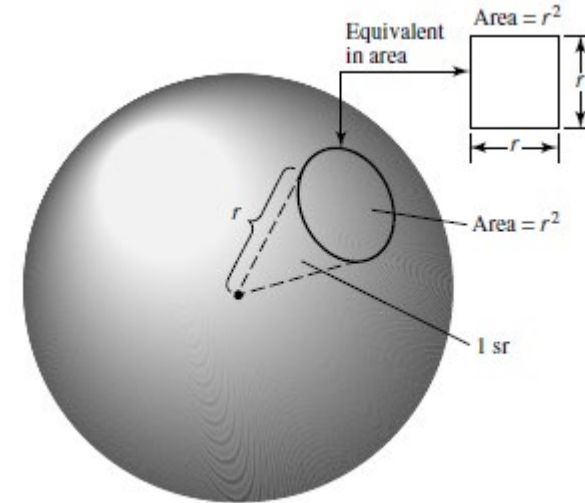
## □ Steradian

- Defined on sphere of radius one
- Angles corresponding to unit area on surface
- $4\pi$  sr in the sphere

## □ Infinitesimal area and solid angle:

$$dA = r^2 \sin \theta d\theta d\phi \quad (\text{m}^2) \quad d\Omega = \frac{dA}{r^2} = \sin \theta d\theta d\phi \quad (\text{sr})$$

- Note:  $\theta$  is the inclination angle not elevation



(b) Steradian

# Radiation Intensity

□ From previous slide: **Radiation density**:  $W = W(r, \theta, \phi) = \frac{1}{2\eta} |\mathbf{E}(r, \theta, \phi)|^2$

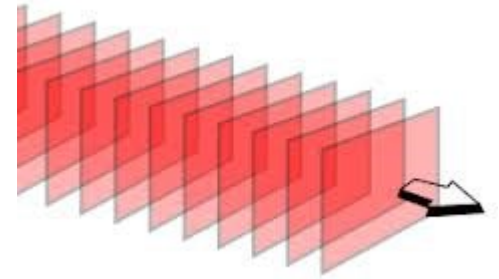
- Units  $\frac{W}{m^2}$

□ Also define **radiation intensity**:  $U = r^2 W = \frac{r^2}{2\eta} |\mathbf{E}(r, \theta, \phi)|^2$

- Watts per solid angle:  $\frac{W}{sr}$

□ In far field, radiation pattern typically decays as:

- $\mathbf{E}(r, \theta, \phi) \approx \frac{1}{r} \mathbf{E}_0(\theta, \phi)$
- In this case,  $U(r, \theta, \phi) = r^2 W(r, \theta, \phi) = \frac{r^2}{2\eta} |\mathbf{E}(r, \theta, \phi)|^2 \approx \frac{1}{2\eta} |\mathbf{E}_0(\theta, \phi)|^2$
- Only depends on angular position  $U(r, \theta, \phi) = U(\theta, \phi)$
- Does not depend on distance  $r$



# Total Radiated Power

---

□ Total radiated power:

$$P_{rad} = \iint U d\Omega = \int_{-\pi/2}^{\pi/2} \int_{-\pi}^{\pi} U(\theta, \phi) \cos \theta d\phi d\theta$$

- Units is Watts
- Note  $\cos \theta$  term! Angle here is elevation angle not polar angle

□ Typically measured in dBm or dBW:

- $P_{rad}[\text{dBm}] = 10 \log_{10} \left[ \frac{P_{rad}}{1 \text{ mW}} \right], P_{rad}[\text{dBW}] = 10 \log_{10} \left[ \frac{P_{rad}}{1 \text{ W}} \right]$
- Power relative to mW or W

# Example Problem

❑ **Problem:** Total radiated power:  $P_{rad} = 30$  dBm

- Assume power is uniformly radiated
- At distance of 2 km, what is the power that strikes a 1 cm x 2 cm surface?

❑ **Solution:**

- Since power is uniform, radiation density is  $W = \frac{P_{tx}}{4\pi d^2}$
- RX power in small area is  $P_{rx} = AW = \frac{AP_{tx}}{4\pi d^2}$
- Area in m<sup>2</sup>,  $A = 2(10)^{-4} \text{ m}^2$
- In dB scale:

$$P_{rx} = P_{tx} + 10 \log_{10} \left( \frac{A}{4\pi d^2} \right) = 30 + 10 \log_{10} \left( \frac{2(10)^{-4}}{4\pi(2000)^2} \right) = -84 \text{ dBm}$$





# Isotropic Antenna

❑ **Isotropic antenna:** Radiates uniformly in all directions

❑ Radiation density and intensity are uniform

- Radiation density:  $W(\theta, \phi, r^2) = \frac{P_{rad}}{4\pi r^2}$
- Radiation intensity:  $U(\theta, \phi) = \frac{P_{rad}}{4\pi}$
- Do not depend on angles  $\theta, \phi$

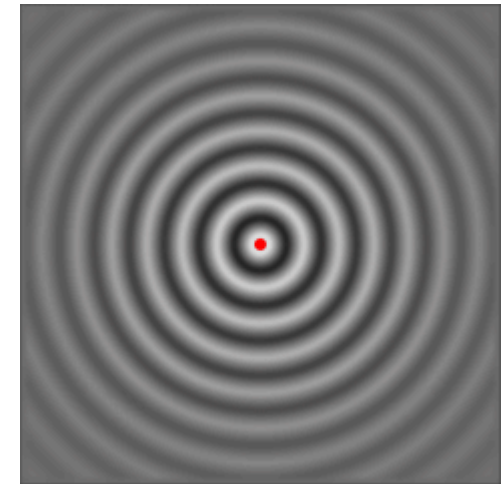
❑ Mostly theoretical construct:

- Most real antennas have some “directivity”

❑ In fact, there can be no coherent (linearly polarized) isotropic radiator

- E-field would be always tangent to sphere
- Such an E-field would have to go to zero in at least one point (“Hairy Ball Theorem”)

Theoretical isotropic pattern



# Antenna Directivity

❑ Most real antennas concentrate power in certain angles

- They are non-isotropic

❑ **Antenna directivity:**

- $D(\theta, \phi) = \frac{4\pi U(\theta, \phi)}{P_{rad}}$  [dimensionless]
- Measures power at an angle relative to average
- Average in linear domain is one
- For isotropic antenna,  $D(\theta, \phi) = 1$

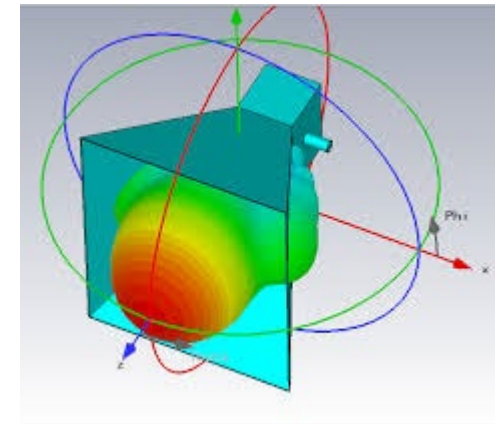
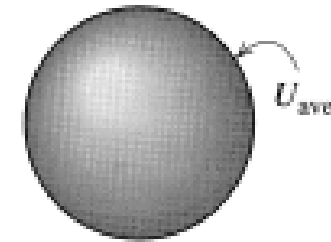
❑ **Max directivity:**  $D_{max} = \max D(\theta, \phi)$

- Directivity in direction with maximum power

❑ Typically measured in dBi

- dB relative to isotropic
- $D(\theta, \phi) [dBi] = 10 \log \left[ \frac{4\pi U(\theta, \phi)}{P_{rad}} \right]$

Theoretical isotropic antenna



Horn antenna with directivity

# Antenna Gain and Efficiency

❑ Most antennas have losses

❑ Define **efficiency**:

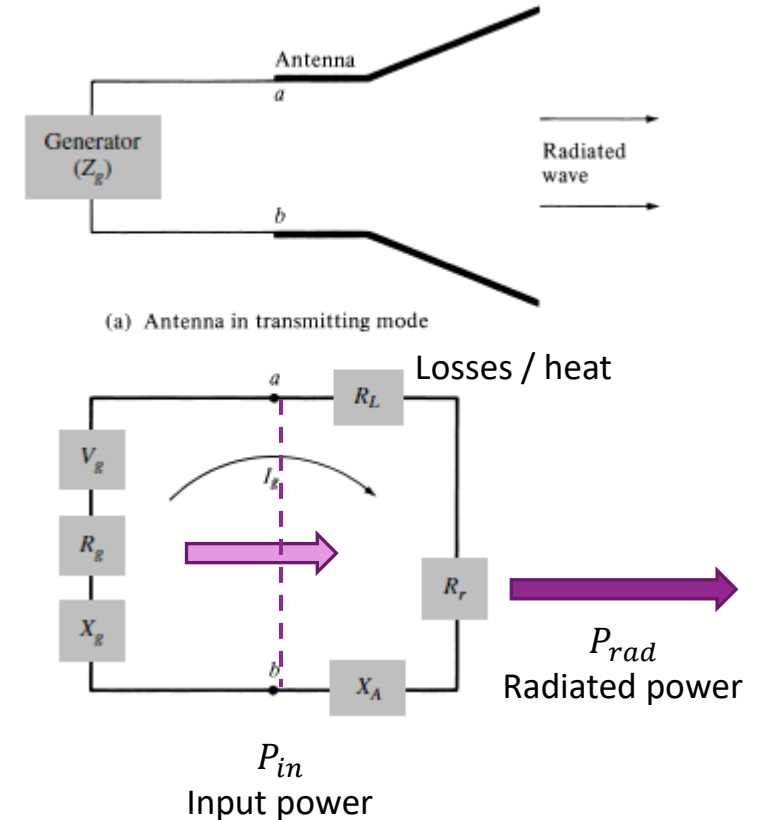
$$\epsilon = \frac{P_{rad}}{P_{in}} \in [0,1]$$

- Radiated to input power in TX mode
- Remaining power is lost in heat in the antenna
- Losses in the conductor and dielectric

❑ Lossless antenna:  $\epsilon = 1$

❑ **Antenna gain**:

- $G(\theta, \phi) = \epsilon D(\theta, \phi) = \frac{4\pi U(\theta, \phi)}{P_{in}}$
- Radiation intensity per unit input power
- For losses antennas, gain = directivity



# Antenna Toolbox in MATLAB

- ❑ Powerful routines for:
  - Design and analysis of antennas
- ❑ Benefits:
  - Supports many antennas
  - Accurate EM modeling
  - Nice visualization tools
  - Simple to use
- ❑ Also, free to NYU students
  - Just download it with MATLAB
- ❑ But...very slow for complex antennas

## Antenna Toolbox

Design, analyze, and visualize antenna elements and antenna arrays

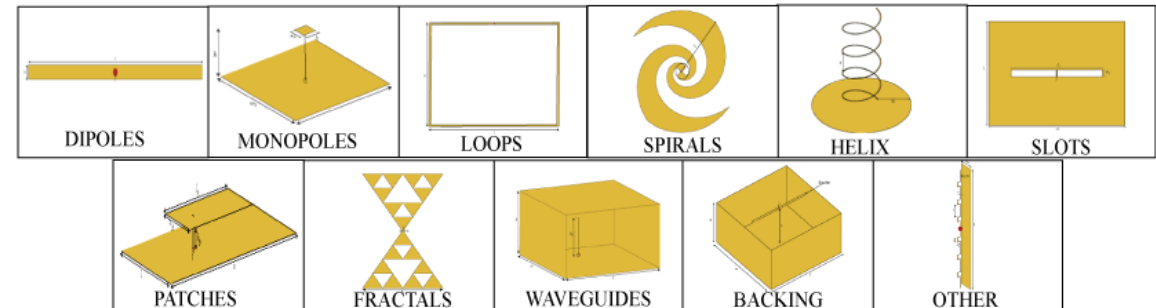
Antenna Toolbox™ provides functions and apps for the design, analysis, and visualization of antennas using either predefined elements with parameterized geometry or arbitrary shapes.

Antenna Toolbox uses the method of moments (MoM) to compute port properties such as the near-field and far-field radiation pattern. You can visualize antenna geometry and radiation patterns.

You can integrate antennas and arrays into wireless systems and use impedance matching, beam forming and beam steering algorithms. Gerber files can be generated from large platforms such as cars or airplanes and analyze the effects of the structure using a variety of propagation models.

## Get Started

Learn the basics of Antenna Toolbox

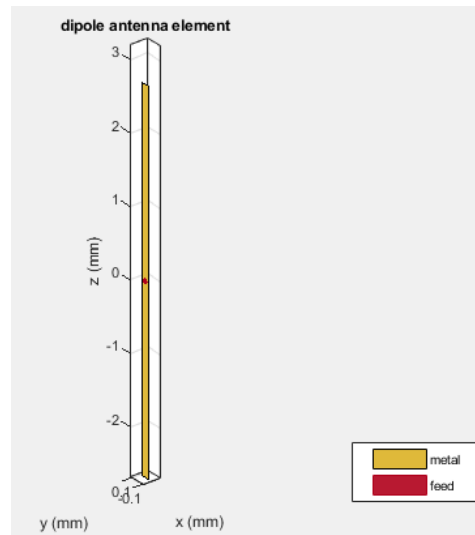


# Patterns in MATLAB: Dipole Example

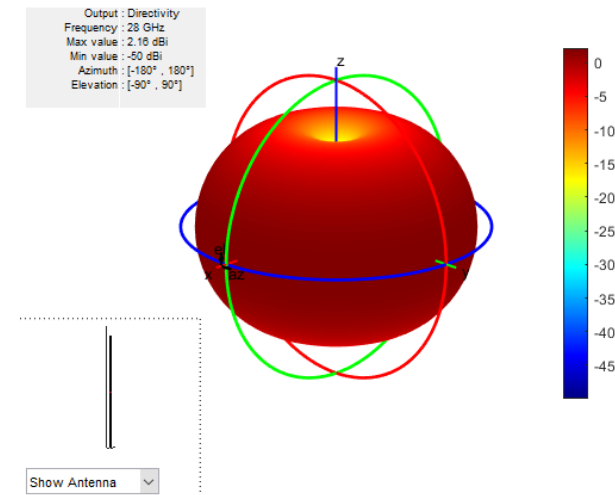
- ❑ MATLAB has powerful tools for calculating antenna patterns

```
%% Simulation constants
fc = 28e9;
vp = physconst('lightspeed');
lambda = vp/fc;

%% Dipole antenna
% Construct the antenna object
ant = dipole(...
    'Length', lambda/2,...
    'Width', 0.01*lambda );
```



```
ant.show();
```



```
ant.pattern(fc)
```

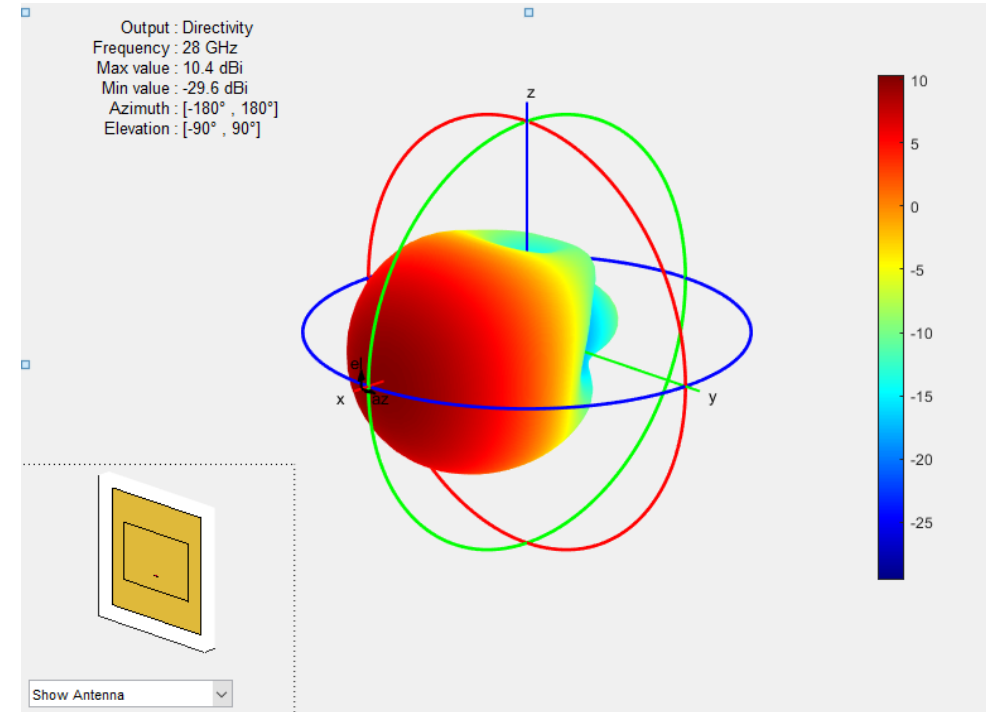
# Microstrip Patch Example

- ❑ A more complex antenna
- ❑ Many other parameters
  - Substrate selection (e.g. FR4, Rogers)
  - Shapes, notches, ...

```
%% Create a patch element
len = 0.49*lambda;
groundPlaneLen = lambda;
ant2 = patchMicrostrip(...
    'Length', len, 'Width', 1.5*len, ...
    'GroundPlaneLength', groundPlaneLen, ...
    'GroundPlaneWidth', groundPlaneLen, ...
    'Height', 0.01*lambda, ...
    'FeedOffset', [0.25*len 0]);

%%
% Tilt the element so that the maximum energy is in the x-axis
ant2.Tilt = 90;
ant2.TiltAxis = [0 1 0];

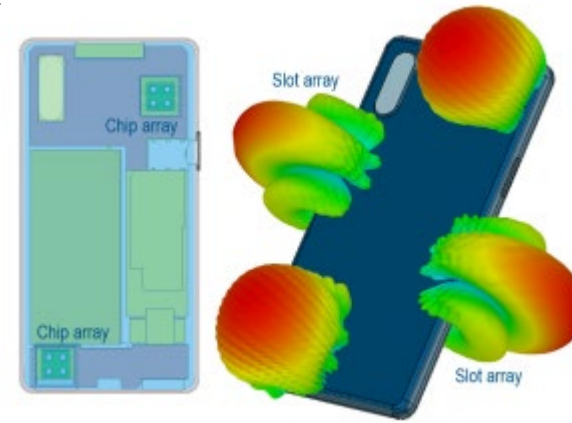
% Display the antenna pattern after rotation
ant2.pattern(fc);
```



# More Complex Antennas

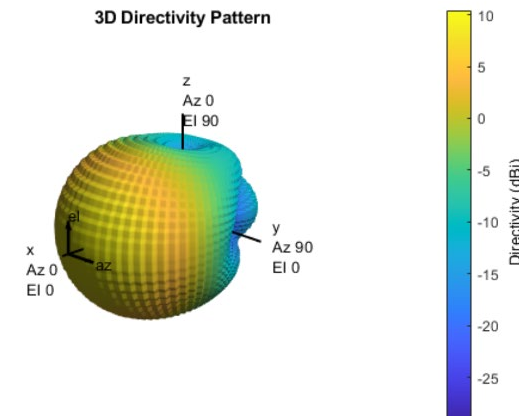
- ❑ For complex antennas:
  - MATLAB antenna toolbox is often too slow
  - Cannot handle packaging, covers, obstacles, ...
  - Need other tools (e.g. Ansoft HFSS and CST)
- ❑ Use MATLAB custom antenna object
  - Store offline computed pattern

```
phasePattern = zeros(size(dir));  
ant3 = phased.CustomAntennaElement(...  
    'AzimuthAngles', az, 'ElevationAngles', el, ...  
    'MagnitudePattern', dir, ...  
    'PhasePattern', phasePattern);  
  
% Plot the antenna pattern.  
% Note the format is slightly different since we are using  
% the pattern routine from the phased array toolbox  
ant3.pattern(fc);
```



CST simulation of 28 GHz array on a handset with cover

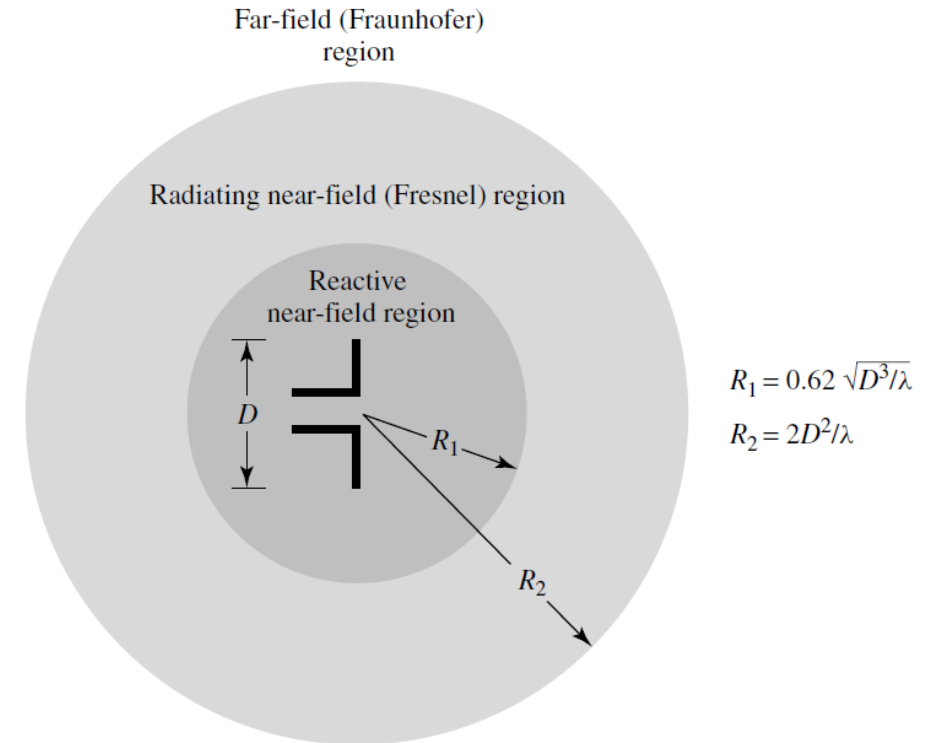
<https://blogs.3ds.com/simulia/5g-antenna-design-mobile-phones/>



Demo of custom antenna element in MATLAB

# Field Regions

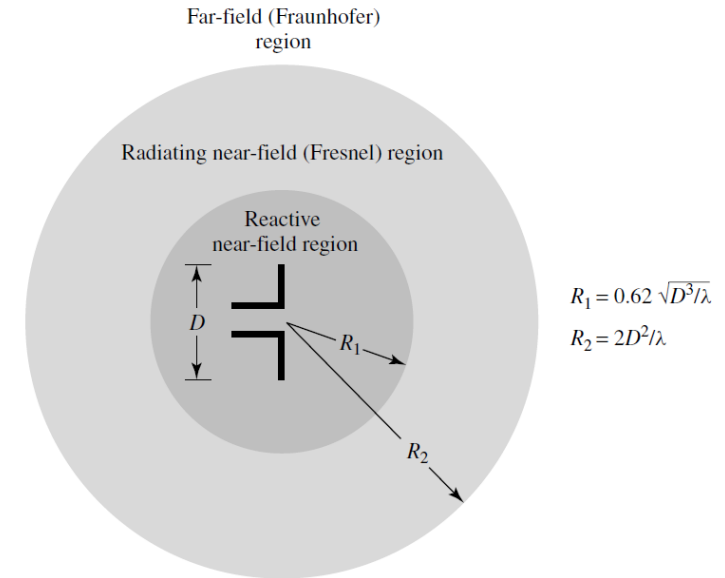
- ❑ Antenna patterns depend on the region
- ❑ **Reactive near field:**
  - Reactive pattern dominates
- ❑ **Radiating near field** or Fresnel region:
  - Angular pattern depends on distance
- ❑ **Far field** or Fraunhofer region:
  - Angular pattern independent of distance
  - Radiation is approximately plane waves
- ❑ Can be approximately calculated using:
  - $D$ : Maximum antenna dimension
  - $\lambda$ : Wavelength





# Rayleigh Distance

- Distance  $R_2$  to far-field = Rayleigh distance
- Most cellular / WLAN systems operate in far field
- Ex 1: Half wavelength dipole antenna
  - $f_c = 2.3$  GHz:
  - $D = \frac{\lambda}{2}$ ,  $R_2 = \frac{2D^2}{\lambda} = \frac{\lambda}{2} = 6.5$  cm
- Ex 2: Large cellular base station
  - $D \approx 7$  m,  $f_c = 2.3$  GHz
  - $R_2 = 751$  m
- Ex 3: MmWave wide aperture antenna
  - $D \approx 40$  cm,  $f_c = 140$  GHz
  - $R_2 = 149$  m



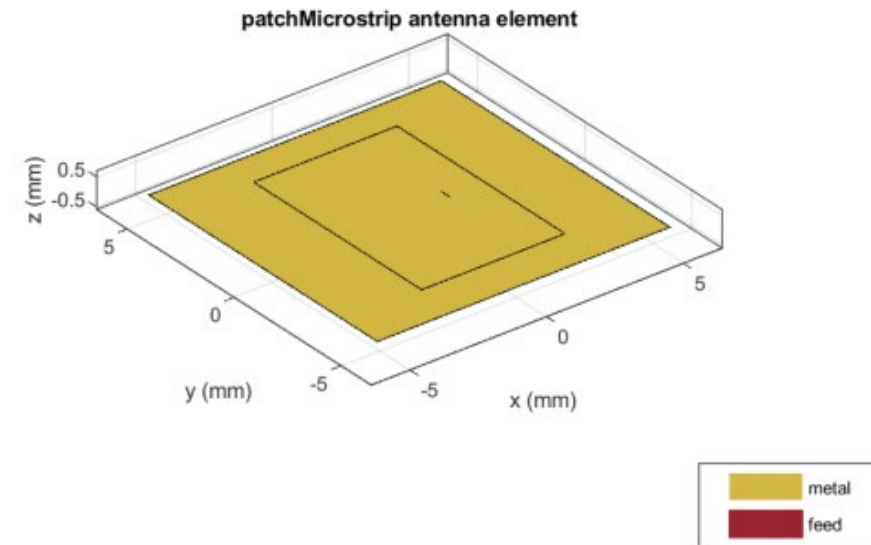
# In-Class Exercise

## Problem 3: Creating and Displaying a Patch Antenna

Create the patch antenna as follows:

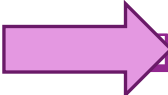
```
% Compute the wavelength
fc = 28e9;
c=physconst('Lightspeed');
lambda = c/fc;

% Create the patch antenna
len = 0.49*lambda;
groundPlaneLen = lambda;
ant = patchMicrostrip(...
    'Length', len, 'Width', 1.5*len, ...
    'GroundPlaneLength', groundPlaneLen, ...
    'GroundPlaneWidth', groundPlaneLen, ...
    'Height', 0.01*lambda, ...
    'FeedOffset', [0.25*len 0]);
```



# Outline

---

- ☐ Basics of Electromagnetic Waves
- ☐ Power and Bandwidth of Signals
- ☐ Basics of Antennas
-  ☐ Free Space Propagation
- ☐ Frames of Reference and Rotations

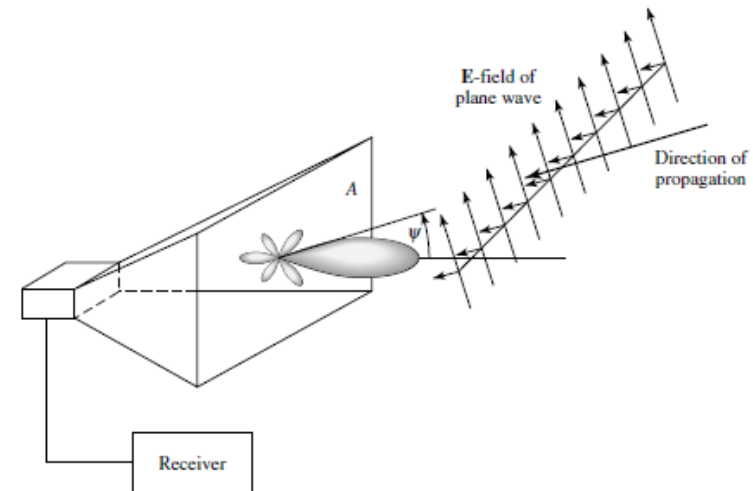
# Antenna Effective Aperture

- ❑ Suppose RX antenna sees incident plane wave
  - Assume polarization aligned to the antenna

- ❑ The effective antenna aperture (or area):

$$A_e(\theta, \phi) = \frac{P_L}{W(\theta, \phi)} \quad [m^2]$$

- $W$  = Power density of incident wave [ $W / m^2$ ]
  - $P_L$  = Power delivered to load at the receiver [ $W$ ]
- ❑ The effective area that the antenna collects
  - We will see this is different than the physical aperture
- ❑  $A_e$  will depend on the direction of arrival



# Aperture and Directivity

❑ From previous slide, effective aperture is:  $A_e(\theta, \phi) = \frac{P_L}{W(\theta, \phi)} [m^2]$

- Ratio of received power to incident radiation density

❑ Aperture-directivity relation:

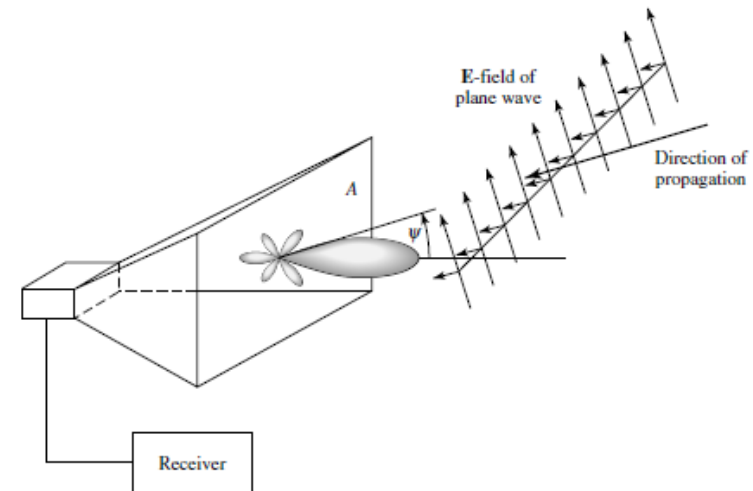
$$A_e(\theta, \phi) = D(\theta, \phi) \frac{\lambda^2}{4\pi}$$

- True for all lossless antennas
- Proof: next slide

❑ Consequence: Average aperture is always  $\frac{\lambda^2}{4\pi}$

- Why?  $\frac{1}{4\pi} \iint A_e(\theta, \phi) \cos \theta d\theta d\phi = \frac{\lambda^2}{(4\pi)^2} \iint D(\theta, \phi) \cos \theta d\theta d\phi = \frac{\lambda^2}{4\pi}$

❑ Independent of the physical size of the antenna!



# Reciprocity of Antennas

❑ To prove aperture-directivity, we need **reciprocity**

❑ Loosely stated:

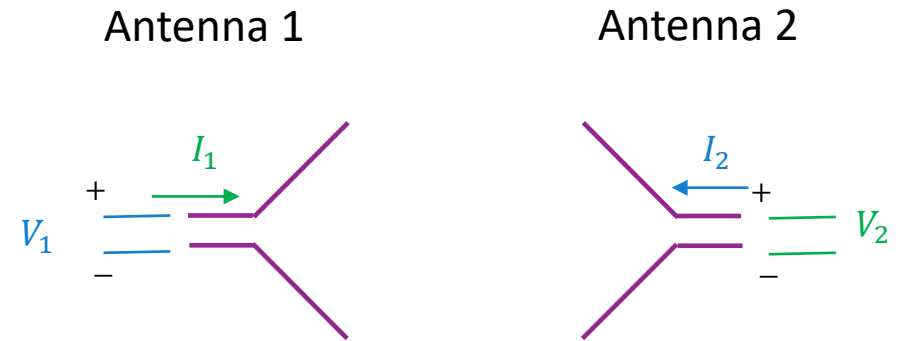
*Channel between antennas in both directions are equal*

❑ Mathematically:

- If Ant 1 Transmits: RX voltage / TX current =  $\frac{V_2^{oc}}{I_1}$
- If Ant 2 Transmits: RX voltage / TX current =  $\frac{V_1^{oc}}{I_2}$
- Reciprocity:  $\frac{V_2^{oc}}{I_1} = \frac{V_1^{oc}}{I_2}$

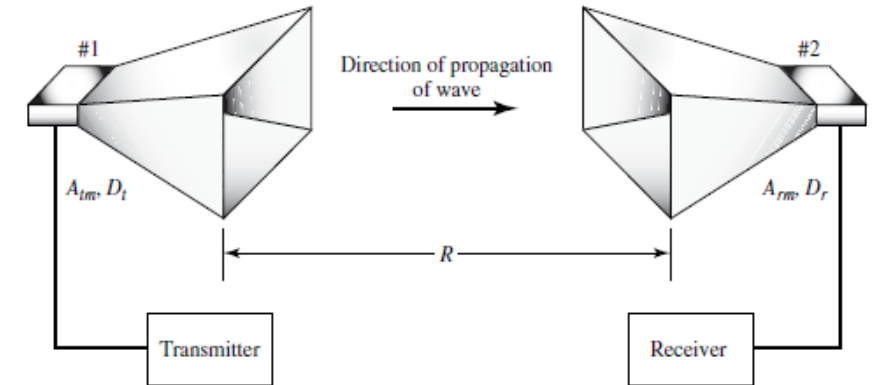
❑ Can show that with the same loads:

- Power transfer in both directions are equal:  $\frac{P_{r2}}{P_{t1}} = \frac{P_{t2}}{P_{r1}}$



# Proof of the Aperture-Directivity Relation

- Suppose Ant 1 transmits power  $P_t$
- Radiation density is:  $W = \frac{D_1 P_t}{4\pi R^2}$
- Received power at Ant 2:  $P_r = A_2 W = \frac{A_2 D_1 P_t}{4\pi R^2} \Rightarrow \frac{P_r}{P_t} = \frac{A_2 D_1}{4\pi R^2}$
- TX from Ant 2, the gain must be the same:  $\frac{P_r}{P_t} = \frac{A_1 D_2}{4\pi R^2}$ 
  - This is a consequence of **reciprocity**
- Hence, for *any* two antennas:  $\frac{D_1}{A_1} = \frac{D_2}{A_2}$
- From simple antenna calculations for a short dipole:
  - $D_2 = \frac{3}{2}$ ,  $A_2 = \frac{3\lambda^2}{8\pi} \Rightarrow \frac{D_2}{A_2} = \frac{4\pi}{\lambda^2}$  (Needs basic EM theory)



# Friis' Law

□ Consider two lossless antennas in **free space**

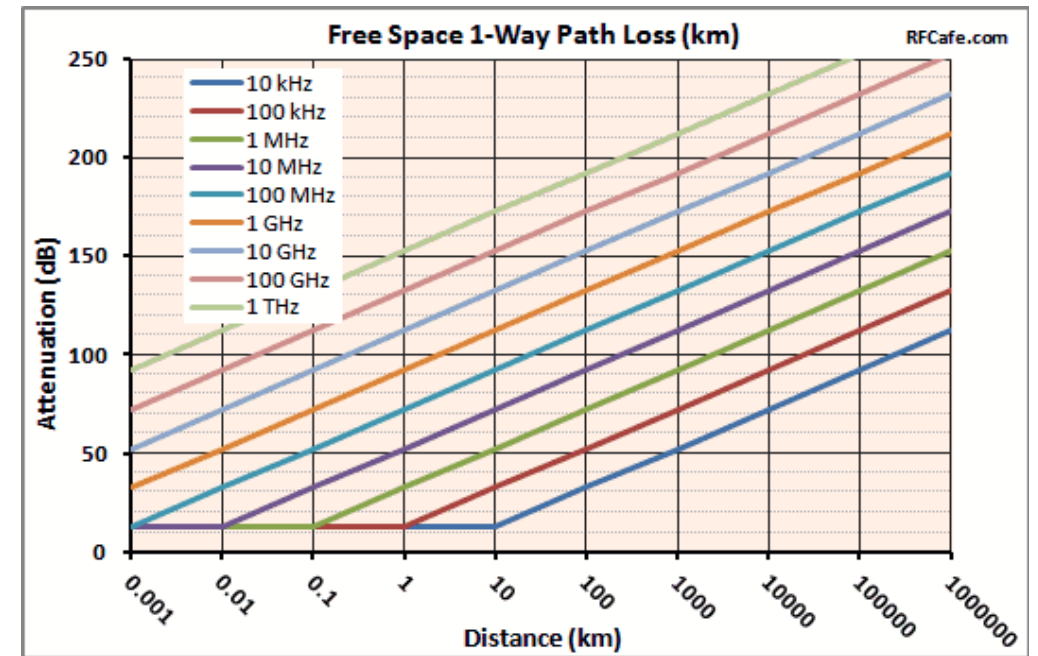
□ From previous slide:  $\frac{P_r}{P_t} = \frac{A_1 D_2}{4\pi R^2}$

□ From aperture-directivity relation:  $A_1 = D_1 \frac{\lambda^2}{4\pi}$

□ This leads to **Friis' Law** (for lossless antennas):

$$\frac{P_r}{P_t} = D_1 D_2 \left( \frac{\lambda}{4\pi R} \right)^2$$

- **Path loss** is proportional to  $R^2$
- Path loss Inversely proportional to  $\lambda^2 \Rightarrow$  proportional to  $f_c^2$





# Example: Calculating Path Loss

❑ Suppose  $f_c = 2.3$  GHz,  $d = 500$  m, what is the **omni directional path loss**?

- Omni-Directional path loss is path loss without the antenna gain

❑ This is easily done in MATLAB:

```
fc = 2.3e9;      % Carrier frequency
vp = physconst('lightspeed'); % speed of light
lambda = vp/fc;  % wavelength

d = 500; % distance in meters

% We can compute the FSPL manually from Friis' law
% Note the minus sign
plOmni1 = -20*log10(lambda/4/pi/d);

% Or, we can use MATLAB's built in function:
plOmni2 = fspl(d, lambda);

fprintf(1,'Omni PL - manual: %7.2f\n', plOmni1);
fprintf(1,'Omni PL - MATLAB: %7.2f\n', plOmni2);
```

Omni PL - manual: 93.66

Omni PL - MATLAB: 93.66

# Polarization

□ We know E-field in far-field behaves as:

- $\mathbf{E}(r, \theta, \phi) \approx \frac{1}{r} [F_V(\theta, \phi) \mathbf{e}_\theta + F_H(\theta, \phi) \mathbf{e}_\phi]$
- Since E-field decays at  $\frac{1}{r}$  and is tangent to the sphere

□ We call:

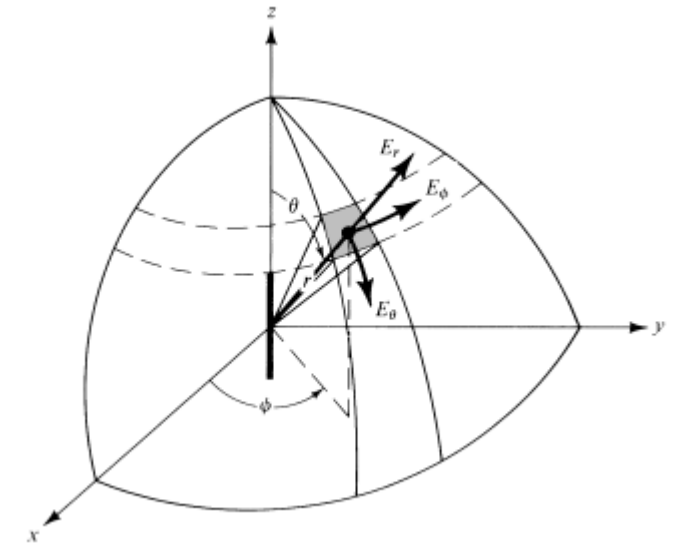
- $F_V(\theta, \phi)$  = vertical E-field pattern
- $F_H(\theta, \phi)$  = horizontal E-field pattern

□ Define **complex gain patterns**:

- $C_V = \sqrt{\frac{4\pi}{2\eta P_{rad}}} F_V$  and  $C_H = \sqrt{\frac{4\pi}{2\eta P_{rad}}} F_H$

□ Define **polarization vector**:  $\mathbf{C}(\theta, \phi) = C_V(\theta, \phi) \mathbf{e}_\theta + C_H(\theta, \phi) \mathbf{e}_\phi$

- Vector is in direction of E-field
- Then directivity is:  $G(\theta, \phi) = |C_V(\theta, \phi)|^2 + |C_H(\theta, \phi)|^2 = \|\mathbf{C}(\theta, \phi)\|^2$
- This convention is used in Nvidia Sionna



(b) Electric field orientation

# Complex Channel Gain

## □ Friis' Law:

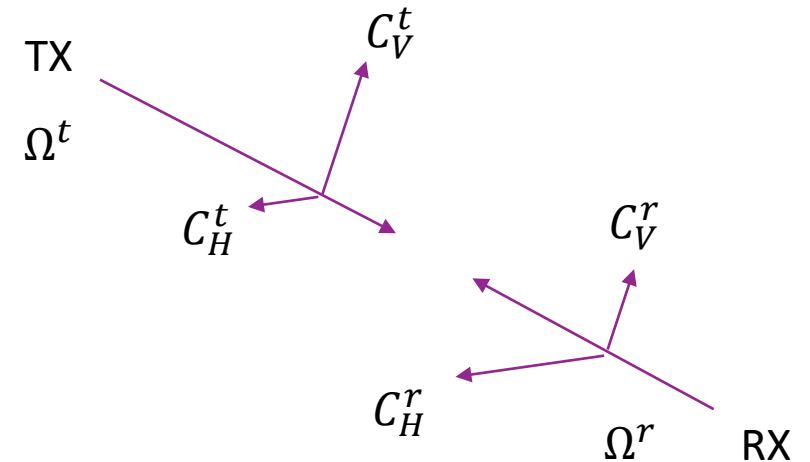
- Provides power gain, not complex channel
- Assumes aligned polarization

## □ To compute complex channel gain for arbitrary polarization

- Find angle of departure at TX:  $\Omega^t = (\theta^t, \phi^t)$
- Find angle of arrival at RX:  $\Omega^r = (\theta^r, \phi^r)$
- Find TX complex gain patterns:  $\mathbf{C}^t(\Omega^t) = C_V^t(\Omega^t)\mathbf{e}_\theta + C_H^t(\Omega^t)\mathbf{e}_\phi$
- Find RX complex gain patterns:  $\mathbf{C}^r(\Omega^r) = C_V^r(\Omega^r)\mathbf{e}_\theta + C_H^r(\Omega^r)\mathbf{e}_\phi$

## □ Complex channel gain will be:

$$H = \underbrace{\mathbf{C}^t(\Omega^t) \cdot \mathbf{C}^r(\Omega^r)}_{\text{Inner product}} e^{-kr} \frac{\lambda}{4\pi r}$$



# Polarization Loss

□ From previous slide, complex channel gain is:  $H = \mathbf{C}^t(\Omega^t) \cdot \mathbf{C}^r(\Omega^r) e^{-kr} \frac{\lambda}{4\pi r}$

□ Power gain is:  $\frac{P_r}{P_t} = |H|^2 = |\mathbf{C}^t(\Omega^t) \cdot \mathbf{C}^r(\Omega^r)|^2 \left(\frac{\lambda}{4\pi r}\right)^2$

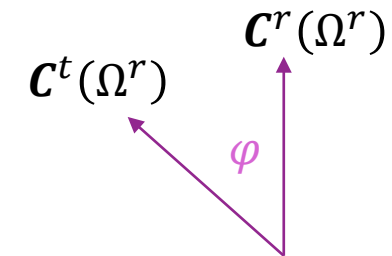
□ Define polarization unit vectors:  $\mathbf{u}^t = \frac{\mathbf{C}^t}{\|\mathbf{C}^t\|}, \mathbf{u}^r = \frac{\mathbf{C}^r}{\|\mathbf{C}^r\|}$

□ Since  $|\mathbf{C}^t(\Omega^t)|^2 = D_t, |\mathbf{C}^r(\Omega^r)|^2 = D_r$ :

$$\frac{P_r}{P_t} = |H|^2 = D_t D_r \left(\frac{\lambda}{4\pi r}\right)^2 \cos^2 \varphi$$

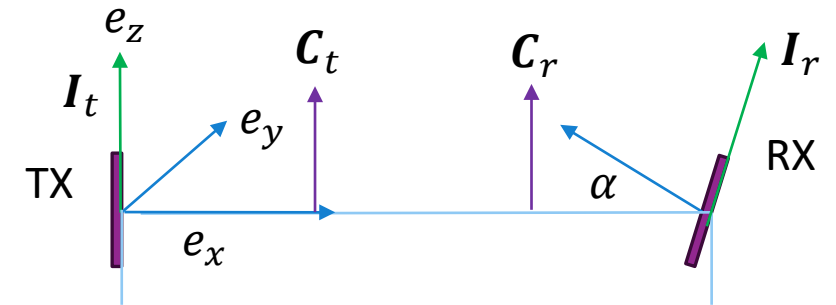
□ Polarization loss:  $\mathbf{u}^t \cdot \mathbf{u}^r = \cos^2 \varphi$  loss due to misalignment in polarization

- Often quoted in dB:  $F = 20 \log_{10} \cos \varphi$



# Ex: Rotation Orthogonal to Path Direction

- Consider two antennas:
  - TX at  $(0,0,0)$ , RX at  $(r, 0,0)$ . Both vertically polarized
- TX: AoD  $\theta^t = \phi^t = 0$ .
  - H-field is in direction  $\mathbf{I}_t \times \mathbf{r} \propto \mathbf{e}_z \times \mathbf{e}_x = \mathbf{e}_y$
  - E-field & TX pol vector  $\mathbf{C}_t \propto \mathbf{H} \times \mathbf{r} \propto \mathbf{e}_y \times \mathbf{e}_x = \mathbf{e}_z$
- RX antenna is rotated around y axis by  $\alpha$ 
  - AoA relative to antenna:  $\theta^r = -\alpha, \phi^r = 0$
  - H-field is in direction  $\mathbf{I}_r \times \mathbf{r} \propto \mathbf{I}_r \times \mathbf{e}_x \propto \mathbf{e}_y$
  - E-field & RX pol vector  $\mathbf{C}_r \propto \mathbf{H} \times \mathbf{r} \propto \mathbf{e}_y \times \mathbf{e}_x = \mathbf{e}_z$
  - Path loss =  $D_t(0,0)D_r(-\alpha, 0) \left(\frac{\lambda}{4\pi r}\right)^2$

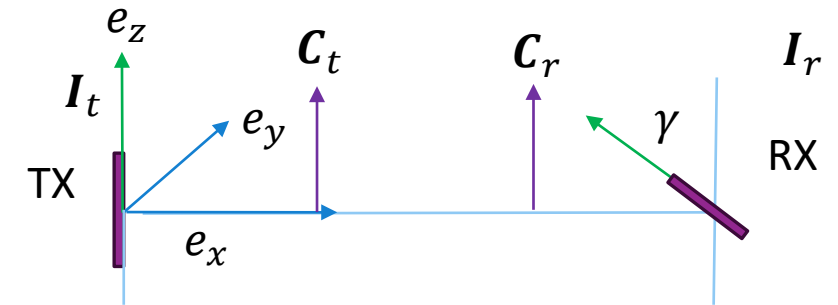


Polarization vectors aligned  
Boresight not aligned

Rotation along axis orthogonal to path direction causes no polarization loss

# Ex: Rotation Around Path Direction

- ❑ Consider two antennas:
  - TX at (0,0,0), RX at (r, 0,0). Both vertically polarized
- ❑ TX: AoD  $\theta^t = \phi^t = 0$ .
  - H-field is in direction  $\mathbf{I}_t \times \mathbf{r} \propto \mathbf{e}_z \times \mathbf{e}_x = \mathbf{e}_y$
  - E-field & TX pol vector  $\mathbf{C}_t \propto \mathbf{H} \times \mathbf{r} \propto \mathbf{e}_y \times \mathbf{e}_x = \mathbf{e}_z$
- ❑ RX antenna is rotated around x axis by  $\gamma$ 
  - AoA relative to antenna:  $\theta^r = 0, \phi^r = 0$
  - $\mathbf{I}_r \propto (0, \sin \gamma, \cos \gamma)$
  - H-field is in direction  $\mathbf{I}_r \times \mathbf{r} \propto \mathbf{I}_r \times \mathbf{e}_x \propto (0, \cos \gamma, \sin \gamma)$
  - E-field & RX pol vector  $\mathbf{C}_r \propto \mathbf{H} \times \mathbf{r} \propto (0, \sin \gamma, \cos \gamma)$
  - Path loss =  $D_t(0,0)D_r(0,0) \left(\frac{\lambda}{4\pi r}\right)^2 \cos^2 \gamma$



Polarization vectors not aligned  
Boresight aligned

Rotation around axis of path direction causes polarization loss

# Antenna Impedance and Matching

❑ Not all power from radio may be delivered to antenna

❑ Some is reflected back

❑ Described by reflection coefficient  $\Gamma$

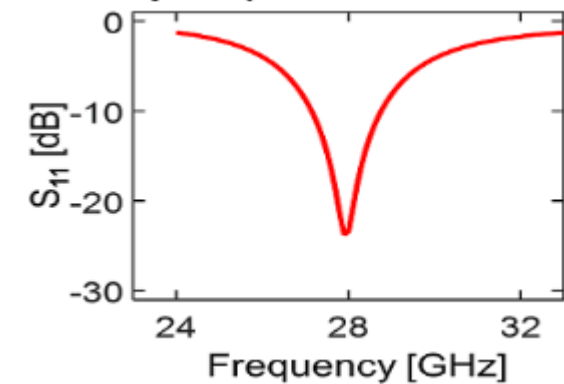
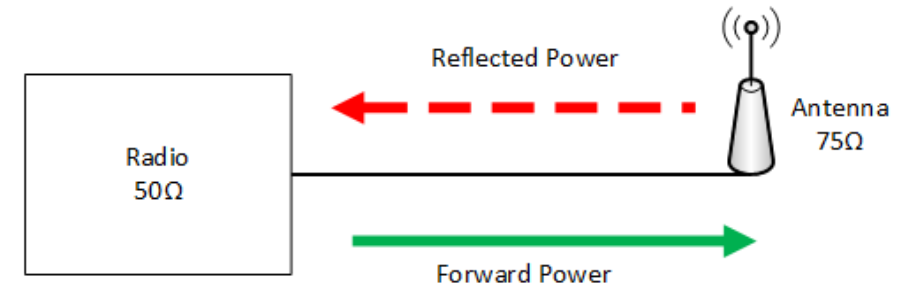
- Also referred to as  $S_{11}$
- Complex ratio of forward to reverse wave

❑ Also described by impedance mismatch:

- $\Gamma = \frac{Z_{in} - Z_0}{Z_{in} + Z_0}$

❑ Fraction of power transferred:  $1 - |\Gamma|^2$

❑ Also given as voltage standing wave ratio (VSWR) =  $\frac{1 + |\Gamma|}{1 - |\Gamma|}$



Ali et al, Small Form Factor PIFA Antenna Design at 28 GHz for 5G Applications, 2019

# Friis' Law with Losses

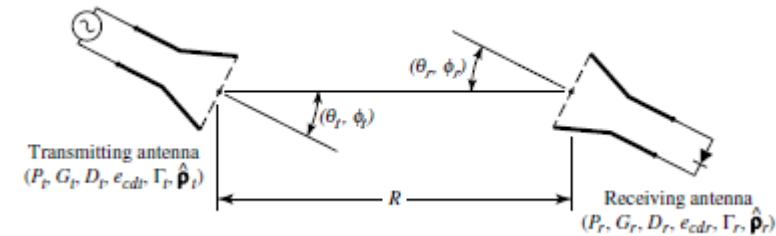
## Three losses in practice:

- Polarization loss
- Conductive / dielectric loss
- Impedance mismatch

## Friis' Law with lossy antennas:

$$\frac{P_r}{P_t} = \epsilon_1 \epsilon_2 (1 - |\Gamma_1|^2)(1 - |\Gamma_2|^2) D_1 D_2 \left( \frac{\lambda}{4\pi R} \right)^2 \cos^2 \theta_{POL}$$

- $\epsilon_i$ : Efficiency of antenna
- $\theta_{POL}$ : Angle between the polarization vectors
- Note that gain is:  $G_i = \epsilon_i D_i$



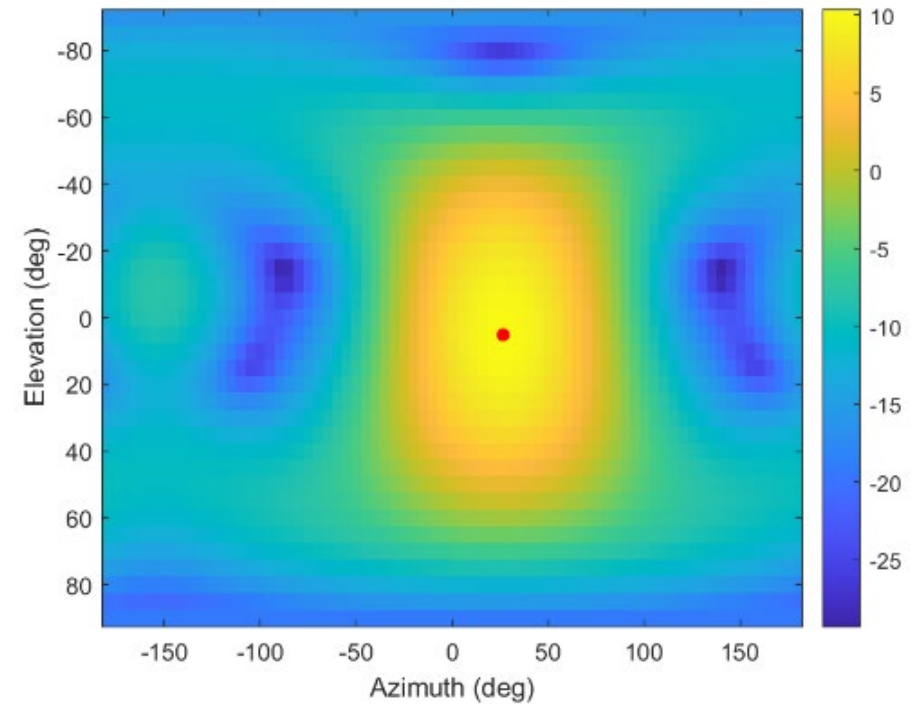


# In-Class Exercise

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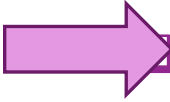
## Problem 4: Plotting the Far Field Radiation Pattern

Use the `ant.pattern` command to get the pattern of the rotated antenna.



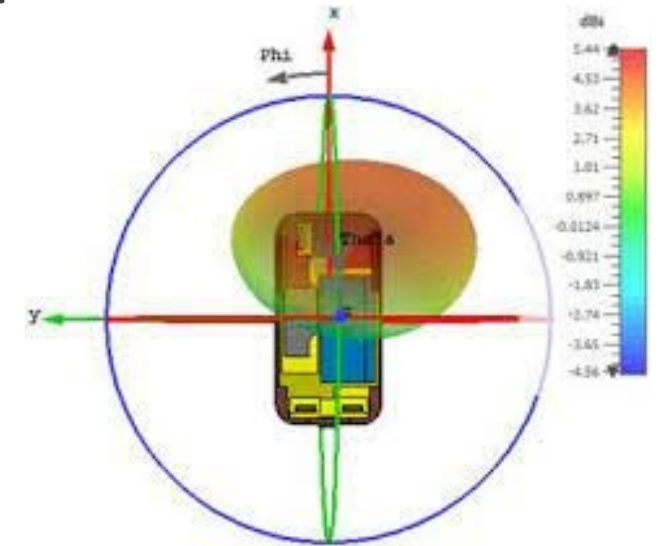
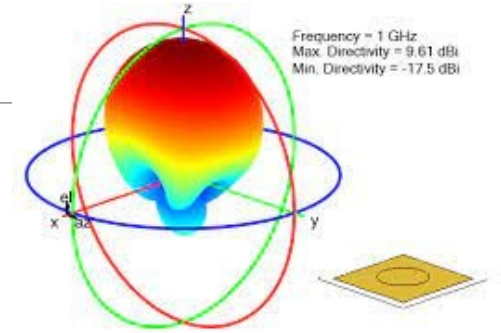
# Outline

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- ☐ Basics of Electromagnetic Waves
- ☐ Power and Bandwidth of Signals
- ☐ Basics of Antennas
- ☐ Free Space Propagation
-  ☐ Frames of Reference and Rotations

# Frames of Reference

- ❑ Antenna patterns are usually given in a “local” frame of reference
  - A coordinate system in a fixed relation with the antenna structure
  - Ex: Coordinate aligned to the normal of a patch antenna
- ❑ But antenna may have arbitrary alignment to the rest of environment
  - Ex, Base station antenna tilt and orientation
- ❑ Alignment may move over time
  - Ex: Rotation or translation of a handset
- ❑ Environment typically specified in “global” frame of reference
- ❑ This section:
  - How to represent local and global coordinate systems
  - How to translate between the two

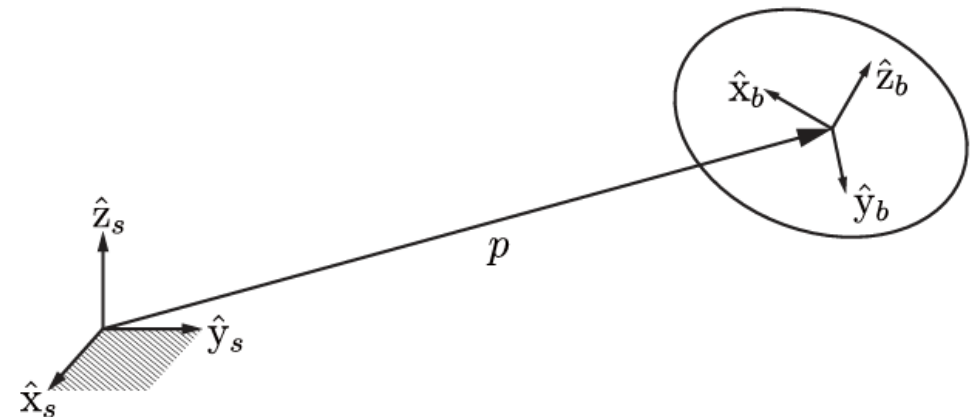


# Rigid Body

- ❑ **Rigid body**: Any structure with fixed relative distances
  - Body may be in motion
  - Ex: Handset, base station antenna, antenna mounted on a car, ...
- ❑ Rigid body's configuration described by two properties
- ❑ **Position** of some reference point in the body:
  - $\mathbf{p} = (p_x, p_y, p_z) \in \mathbb{R}^3$
- ❑ **Orientation** around that point:
  - Described by orthonormal vectors  $\{\hat{x}_b, \hat{y}_b, \hat{z}_b\}$
- ❑ Motion: Change of position and orientation over time

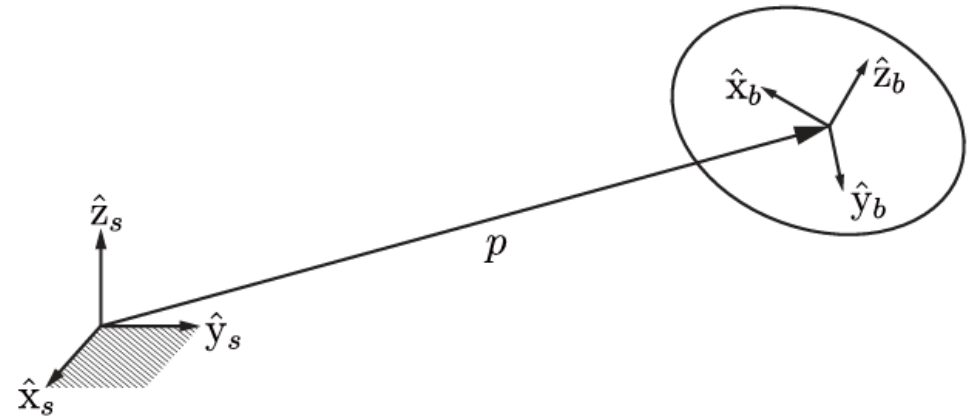


VectorStock® VectorStock.com/11147287



# Frames of Reference for Antennas

- ❑ **Global frame of reference:** An arbitrary coordinate system, typically fixed
  - Assume aligned at the origin with basis  $\hat{x}_s, \hat{y}_s, \hat{z}_s$
- ❑ **Body or local frame of reference:**
  - A coordinate system for a rigid body possibly in motion
  - Translation  $\mathbf{p}$  and basis  $\hat{x}_b, \hat{y}_b, \hat{z}_b$  in the global basis
- ❑ **Antenna systems**
  - Typ. Given antenna patterns in local frame of reference
  - Signal paths given in global frame of reference
  - Need to translate global to local



# Euler Angles

□ Rotation matrices in  $\mathbb{R}^3$  are typically described by three Euler angles

□ Robotics naming:

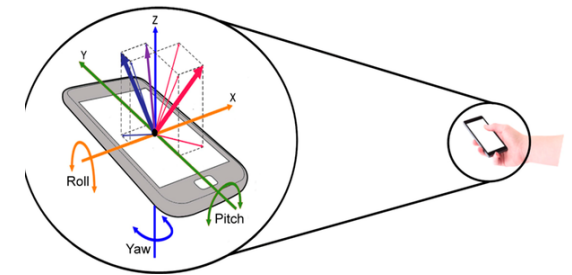
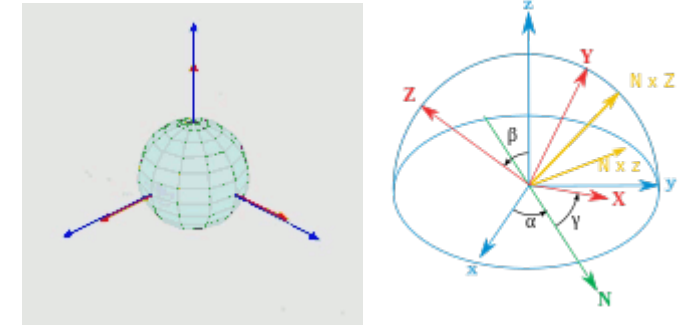
- **Yaw**  $\alpha$ : Rotation round  $z$  axis, corresponds to azimuth angle
- **Pitch**  $\beta$ : Rotation round  $y$  axis, corresponds to inclination angle
- **Roll**  $\gamma$ : Rotation around  $x$  axis

□ Rotation matrix given by a product

- Note order matters. Many different conventions. This product is XYZ:

$$R = R_z(\alpha) R_y(\beta) R_x(\gamma) = \begin{bmatrix} \cos \alpha & -\sin \alpha & 0 \\ \sin \alpha & \cos \alpha & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \cos \beta & 0 & \sin \beta \\ 0 & 1 & 0 \\ -\sin \beta & 0 & \cos \beta \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \gamma & -\sin \gamma \\ 0 & \sin \gamma & \cos \gamma \end{bmatrix}$$

$$= \begin{bmatrix} \cos \alpha \cos \beta & \cos \alpha \sin \beta \sin \gamma - \sin \alpha \cos \gamma & \cos \alpha \sin \beta \cos \gamma + \sin \alpha \sin \gamma \\ \sin \alpha \cos \beta & \sin \alpha \sin \beta \sin \gamma + \cos \alpha \cos \gamma & \sin \alpha \sin \beta \cos \gamma - \cos \alpha \sin \gamma \\ -\sin \beta & \cos \beta \sin \gamma & \cos \beta \cos \gamma \end{bmatrix}$$



[https://en.wikipedia.org/wiki/Euler\\_angles](https://en.wikipedia.org/wiki/Euler_angles)

[https://en.wikipedia.org/wiki/Rotation\\_matrix](https://en.wikipedia.org/wiki/Rotation_matrix)

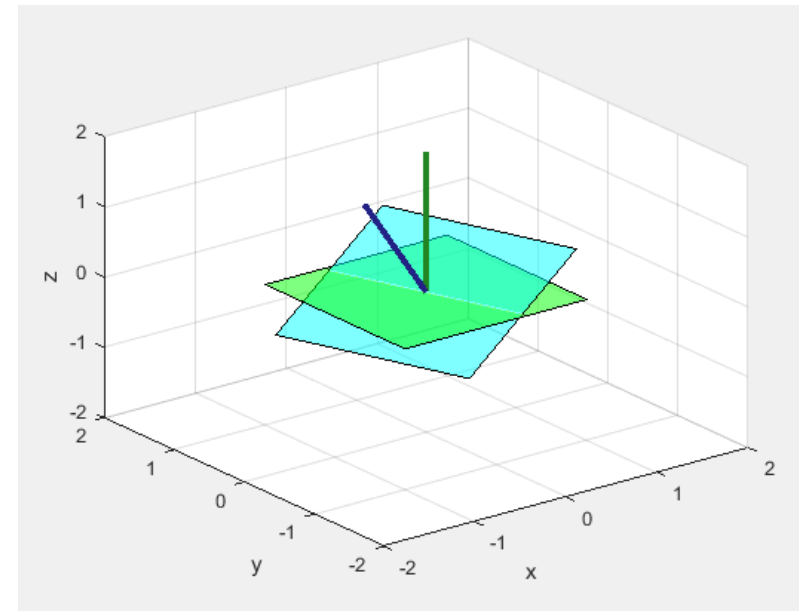
# Computing Rotation Matrices in MATLAB

- ❑ MATLAB has many tools to compute rotation matrices
- ❑ Eul2rotm:
  - Give angles in radians
  - Give order, e.g. 'ZYX'
- ❑ See demo in github

```
% Set the rotation angles
yaw = 20  ;
pitch = -30  ;
roll = 0  ;

% Create the rotation matrix
R = eul2rotm(deg2rad([yaw pitch roll]), 'ZYX');

vertRot = vert*R'; % Rotate the vertices
nvecRot = nvec*R'; % Rotate the normal vector
```



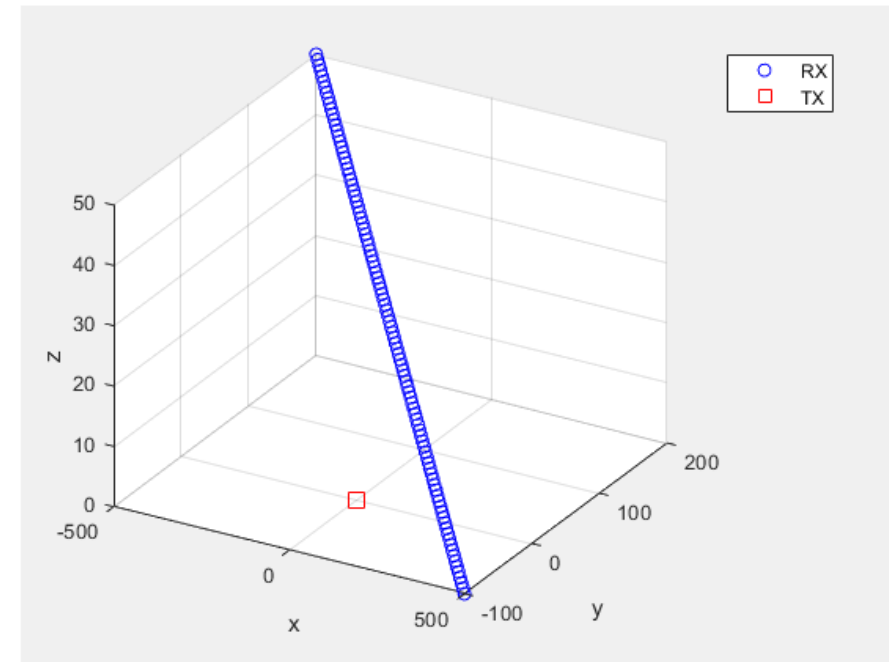
# Example: Frame of Reference of a Moving Object

- ❑ See MATLAB demo
- ❑ Object moving in 3D space (e.g., UAV)
- ❑ Antenna pointed in direction of motion

```
% Get direction of motion
v = xend-xstart;

% Compute the angle of direction of motion
[azDir, elDir, ~] = cart2sph(v(1),v(2),v(3));

% Find a rotation matrix aligned to direction of motion
yaw = azDir;
pitch = -elDir; % Note the negative sign since
roll = 0;
R = eul2rotm([yaw pitch roll], 'ZYX');
```





# Find Angles of Arrival along Path

## Find the angles in RX frame of reference

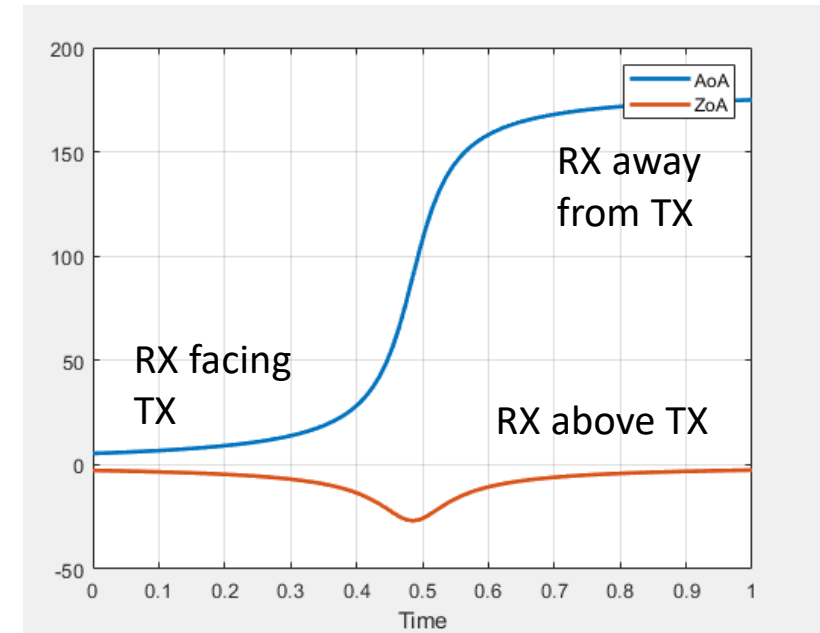
Find the angles of arrival from the TX to the RX in the RX frame of reference

```
% Create the vector from the local antenna to remote signal source
Zpath = -X;

% Rotate to the RX frame of reference
% Note: No transpose since we are multiplying on the right
Zrot = Zpath*R;

% Compute angles in local frame of reference
[azpath, elpath, dist] = cart2sph(Zrot(:,1), Zrot(:,2), Zrot(:,3));

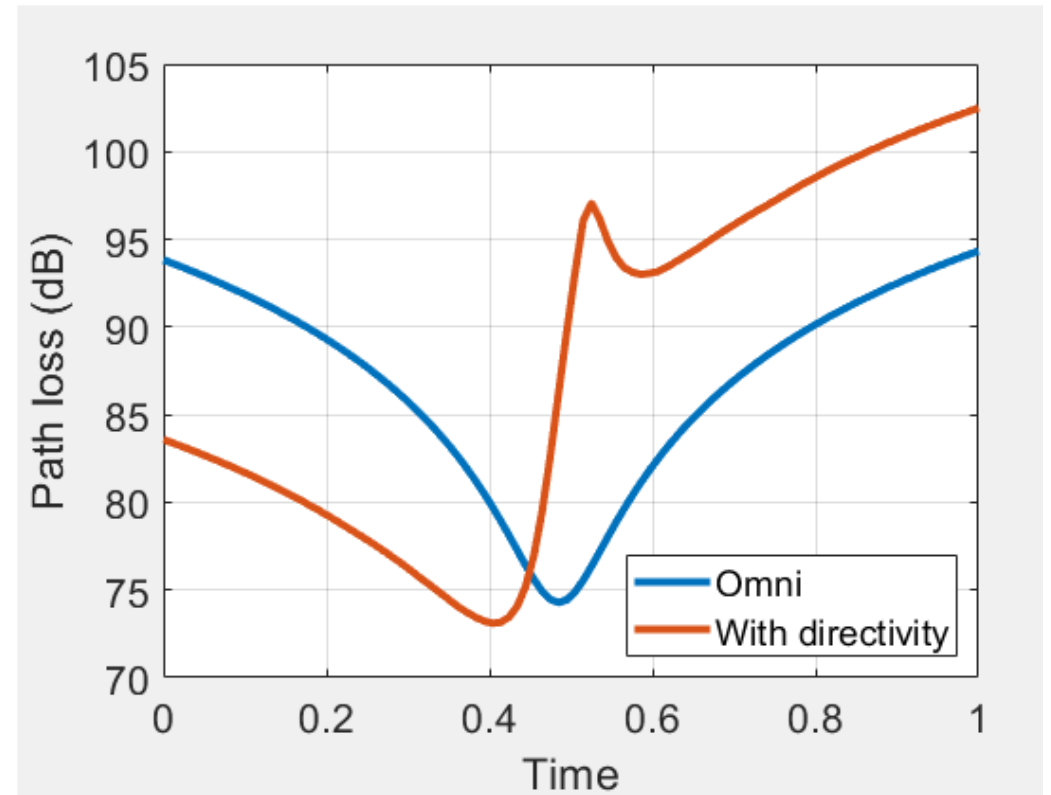
% Convert to degrees
azpath = rad2deg(azpath);
elpath = rad2deg(elpath);
```



# Find Gain along Path

- ❑ Compute omni-directional path loss and gain along path

```
% Compute the free space path loss along the path without  
% the antenna gain. We can use MATLAB's built-in function  
plOmni = fspl(dist, lambda);  
  
% Compute the directivity using interpolation of the pattern.  
% We can use the ant3.resp method for this purpose, but the  
% interpolation is not smooth. So, we will do this using  
% MATLAB's interpolation objects. First, we create the  
% interpolation object.  
F = griddedInterpolant({el,az},dir);  
  
% Then, we compute the directivity using the object  
dirPath = F(elpath,azpath);  
  
% Compute the total path loss including the directivity  
plDir = plOmni - dirPath;
```



# Computing the Polarization Loss

❑ Compute the E-field H and V patterns

```
% First we compute the electric field at the RX in both polarizations
[Ev,azv,elv] = ant.pattern(fc, Type='efield', Polarization='v');
[Eh,azh,elh] = ant.pattern(fc, Type='efield', Polarization='h');
```

❑ Then perform interpolation on path

- Code for interpolation in demo
- Not shown on slide due to space

❑ Assume TX is V pol

❑ RX polarization vector:

- Unit vector in direction of E-field

❑ Use MATLAB built-in polloss function

- Computes angle between TX and RX pol vectors
- Accounts for rotation of TX and RX

```
for i = 1:npts

    % Assume the TX polarization is vertically polarized
    fvTx = [0;1];

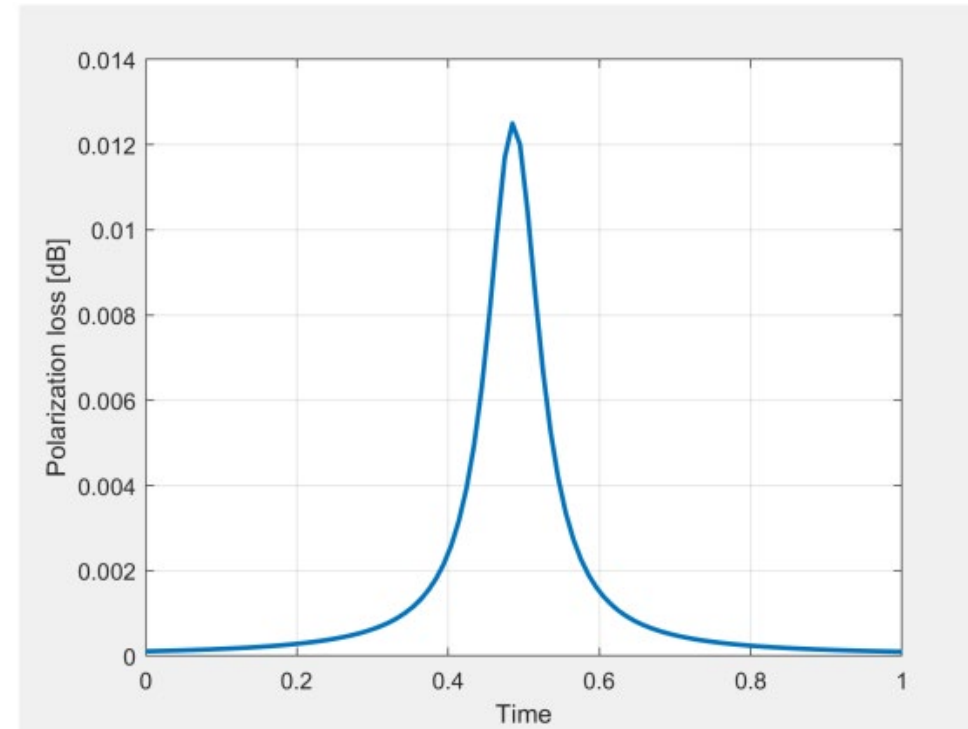
    % The RX polarization is proportional to the E-fields
    fvRx = [Ehpath(i); Evpath(i)];
    fvRx = fvRx / norm(fvRx);

    % Compute the polarization loss
    posRx=X(i,:)' ;
    axesRx=R;
    axesTx=eye(3);
    posTx=zeros(3,1);
    pollosspath(i) = polloss(fvTx,fvRx,posRx,axesRx,...
                             posTx,axesTx);

end
```

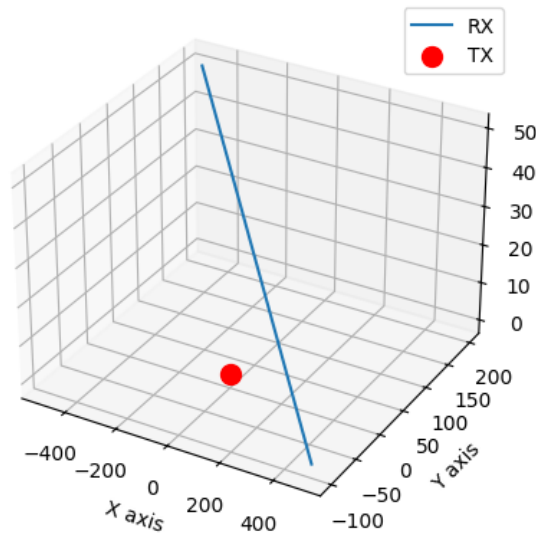
# Polarization Loss

- In this case, loss is minimal
- Loss is  $\sim 0.01$  dB
- Minimal rotation around path direction



# Nvidia Sionna Demo

- Demo in github for calculations with Nvidia Sionna
- Performs all geometric calculations
- Easy tensorflow interface



## Free Space Propagation with Nvidia's Sionna

*Sionna* is a fully open-source Tensorflow-based package developed by Nvidia to support physical layer of functionality as MATLAB's toolboxes. However, Sionna supports:

- Full python interface
- Integration with Tensorflow for machine learning
- Integration with ray tracing

```
# Configure antenna array for all transmitters
scene.tx_array = sn.rt.PlanarArray(num_rows=1,
                                   num_cols=1,
                                   vertical_spacing=0.5,
                                   horizontal_spacing=0.5,
                                   pattern="dipole",
                                   polarization="V")

# Configure antenna array for all receivers
scene.rx_array = sn.rt.PlanarArray(num_rows=1,
                                   num_cols=1,
                                   vertical_spacing=0.5,
                                   horizontal_spacing=0.5,
                                   pattern="tr38901",
                                   polarization="V")
```