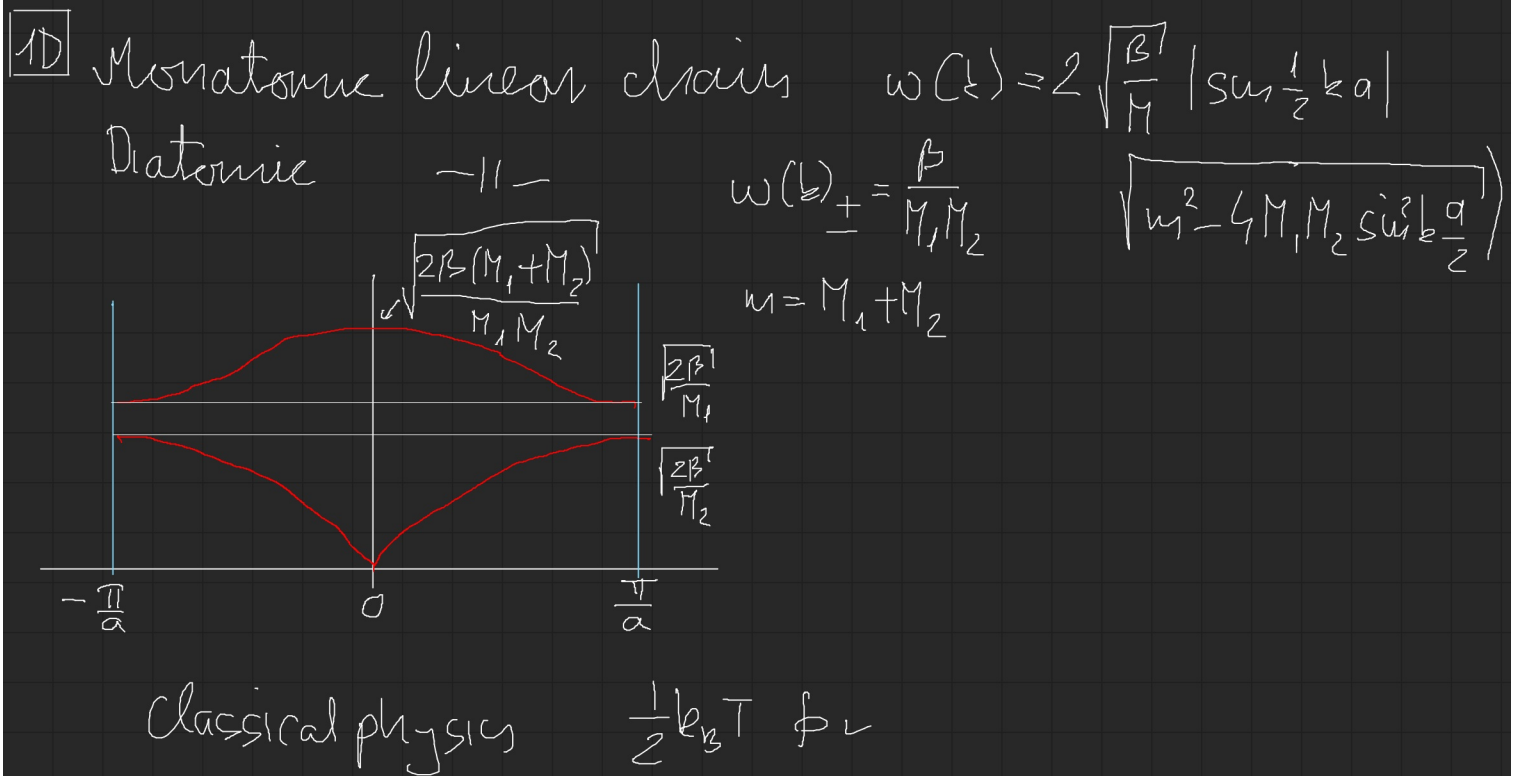


Lecture #16

Specific heat of solids



Classical physics: $\epsilon = \frac{N}{2} k_B T \cdot f$ f - degree of freedom

N molecule

examples:

$= \frac{N}{2 L_A} L_A k_B T \cdot f = \frac{f \cdot M}{2} R T$

$R = 8.31 \frac{J}{mol \cdot K}$
 ideal gas unit

monatomic
diatomic

$$\epsilon = \frac{1}{2} m (\underbrace{v_x^2 + v_y^2 + v_z^2}_{f=3}) \rightarrow U = \frac{3}{2} MRT$$

$$\epsilon = \frac{1}{2} m (\underbrace{v_x^2 + v_y^2 + v_z^2}_3) + \frac{1}{2} \Theta_1 \omega_1^2 + \frac{1}{2} \Theta_2 \omega_2^2$$

$f=5$

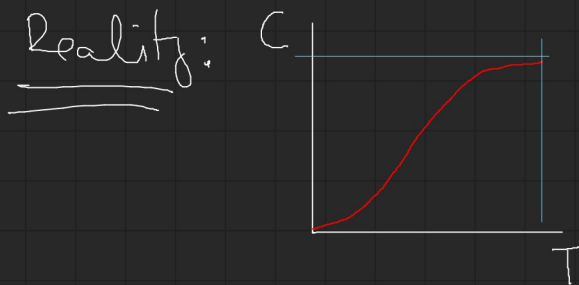
$$U = \frac{5}{2} MRT$$

$$C_V = \left(\frac{\partial U}{\partial T} \right)_{V=\text{const}}$$

heat capacity per mole

$$C = \frac{1}{N} \left(\frac{\partial U}{\partial T} \right) = \begin{cases} \frac{3}{2} R \\ \frac{5}{2} R \end{cases} = \frac{f}{2} R$$

What happens when $T \rightarrow 0$? Nothing...



of phonon states

$$dn_s(\nu, \nu + d\nu) = \frac{4\pi V}{v_s^3(\nu)} \nu^2 d\nu \quad v_s = \begin{cases} v_{\text{longitudinal}} \\ v_{\text{transversal}} \end{cases}$$

$$dn(\nu, \nu + d\nu) = 4\pi V \nu^2 \underbrace{\left(\frac{1}{v_l^3(\nu)} + \frac{2}{v_t^3(\nu)} \right)}_{g_{\text{total}}(\nu)} d\nu$$

$$N \text{ atoms} \Rightarrow f = 3N \equiv (\text{tot. \# of states})$$

$$\int_0^\infty dn = \int_0^\infty g_{\text{tot}}(\nu) d\nu$$

$$4\pi V \int_0^\infty \left(\frac{1}{v_l^3(\nu)} + \frac{2}{v_t^3(\nu)} \right) \nu^2 d\nu = 3N$$

can't calculate!

Debye

assumptions

v_s

0

U_e, U_t

2D

$$4\pi V \left(\frac{1}{v_e^3} + \frac{2}{v_t^3} \right) \int_0^{v_D} v^2 dv = 3N$$

$$\left[\frac{v^3}{3} \right]_0^{v_D} = \frac{v_D^3}{3}$$

$$\frac{1}{v_e^3} + \frac{2}{v_t^3} = \frac{3N}{4\pi V v_D^3}$$

$$g(v) = \frac{3N}{v_D^3} v^2$$

of phonons? $du = g(v) f_{BE}(v) dv$

$$f_{BE} = \frac{v^2 dv}{e^{h\nu/k_B T} - 1}$$

$$U = \int_0^{v_D} E(v, T) g(v) f_{BE}(v, T) dv + \text{const.}$$

$$U = \frac{3Nh}{v_D^3} \int_0^{v_D} \frac{v^3}{e^{h\nu/k_B T} - 1} dv + \text{const.}$$

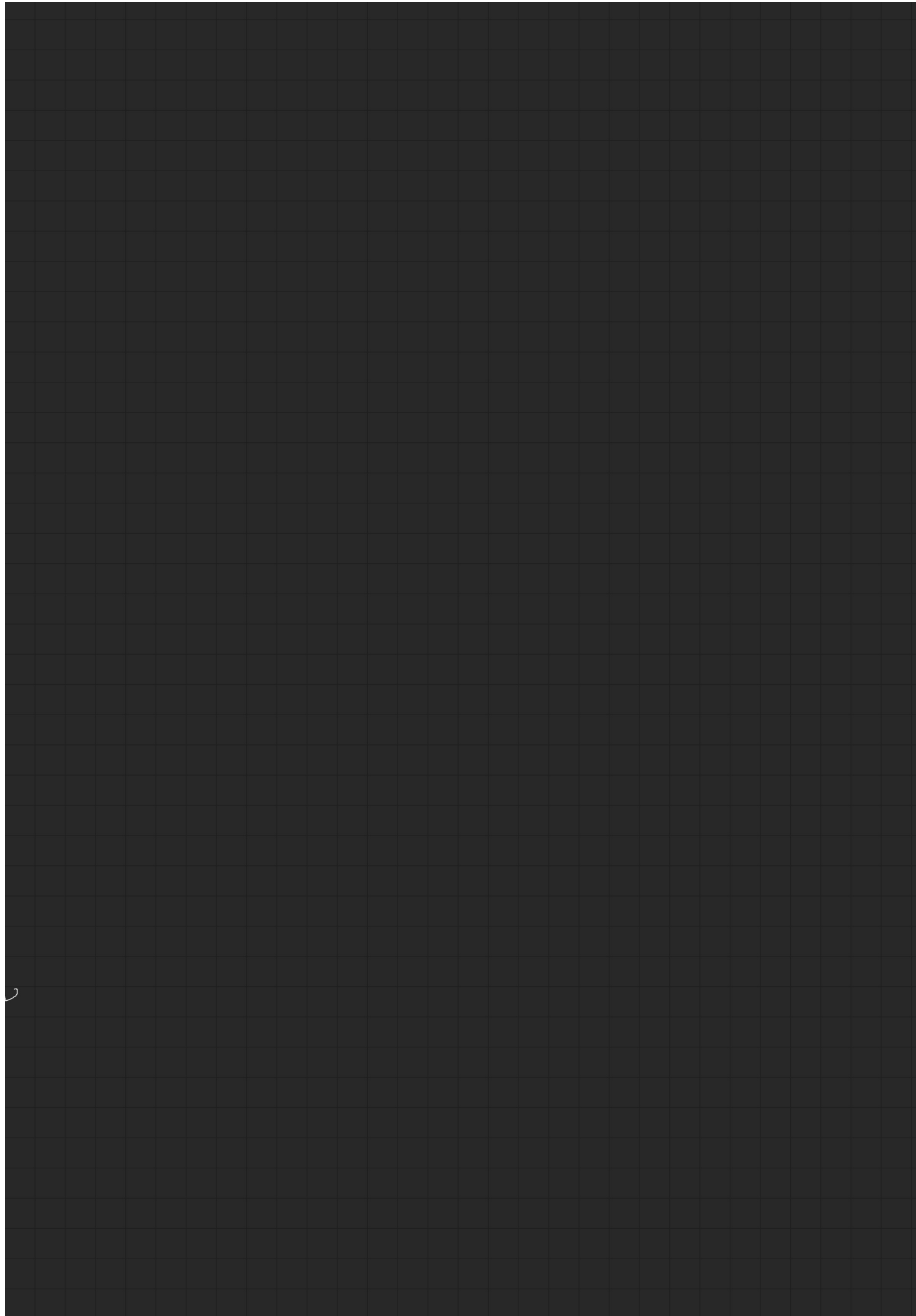
$$C_V = \left(\frac{\partial U}{\partial T} \right)_{V=\text{const}}$$

$$\frac{\partial}{\partial T} \int (\dots) dv = \int \frac{\partial}{\partial T} (\dots) dv$$

$$\frac{\partial}{\partial T} \frac{v^3}{e^{h\nu/k_B T} - 1} = \frac{v^4 e^{h\nu/k_B T}}{(e^{h\nu/k_B T} - 1)^2}$$

$$\Theta_D \equiv \frac{h\nu_D}{k_B}$$

$$C_V = 9R \left(\frac{T}{\Theta_D} \right)^3 \int_0^{\Theta_D/k_B T} \frac{y^4 e^y}{(e^y - 1)^2} dy$$



$$C_V \sim T^3$$

$$C_V \sim 3R$$

$$T \ll \Theta_D$$

$$T \gg \Theta_D \Leftarrow$$

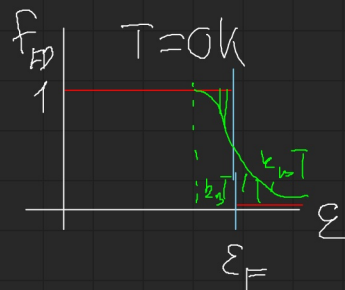
Classical physics:

$$f = 2.3$$

$$C_V = \frac{6}{2} R$$

$$\left(\frac{1}{2} m \vec{v}^2 + \frac{1}{2} D x^2 \right)_{x,y,z}$$

electrons?



of electrons

$$U \sim (k_B T)^2$$

$$C_V^{(e)} \sim T$$

$$\frac{C_V^{(e)}}{C_V} \approx$$

