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space_time_splits.wxm
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An application document for Geometric Algebra using wxMaxima
Ref: The Survey, paragraph 4.1.2
Use of G(1,3) to split spacetime!
Initialization
(%i35) ext:["wxm"]$
        file_type_maxima:append(ext,file_type_maxima)$
        batchload("initialize_fns")$
the pseudoscalar and its inverse
the lowest useable dimension pseudoscalar should be \{e1,e2\} i.e. Plen = 2
e.g. for four dimensions edit Pseudos: {e1,e2,e3}$ to Pseudos: {e1,e2,e3,e4}$
(%i1) Pseudos:{e1,e2,e3,e4}$
       Pvar:listofvars(Pseudos)$
       Plen:length(Pvar)$
       I:Pseudos$
       ni:(Plen-1)*Plen/2$
       Ii:(-1)^ni*I$
       kill(ni)$
       ldisplay(Pvar)$
  (\%t8) Pvar = [e1, e2, e3, e4]
(%i9) batchload("initialize_lsts")$
  (%t9) | lstb||ds = [[{e1},{e2},{e3},{e4}],[{e1,e2},{e1,e3},{e1,e4},{e2,e3},{e2,e4},{
e3,e4}],[{e1,e2,e3},{e1,e2,e4},{e1,e3,e4},{e2,e3,e4}],[{e1,e2,e3,e4}]]
e4},{e1,e2,e3},{e1,e2,e4},{e1,e3,e4},{e2,e3,e4},{e1,e2,e3,e4}]
(\%t11) invblds = [\{e1\}, \{e2\}, \{e3\}, \{e4\}, -\{e1, e2\}, -\{e1, e3\}, -\{e1, e4\}, -\{e2, e3\}, -\{e2, e4\}\}]
,-{e3,e4},-{e1,e2,e3},-{e1,e2,e4},-{e1,e3,e4},-{e2,e3,e4},{e1,e2,e3,e4}]
end of Initialization
floating point print (display) precision
(%i12) fpprintprec:6$
        ratprint:false$
        ldisplay(fpprintprec,fpprec,ratprint)$
 (%t14) fpprintprec=6
 (\%t15) \text{ fpprec} = 16
 (\%t16) ratprint = false
The Survey, para.4.1.2
show the spacetime gammas required for the imitation of G(1,3), where, to avoid the use
of gamma_zero, we have used the fourth axis, e4, for the time axis and the intrinsic
maxima imaginary, %i, for the space axes
(%i17) g1:%i*{e1}$
        g2:%i*{e2}$
        g3:%i*{e3}$
        g4:{e4}$
the spacetime coordinate vector using the gammas
(\%i21) x:x1*g1+x2*g2+x3*g3+t*g4;
(\%o21)\%i*{e3}*x3+\%i*{e2}*x2+\%i*{e1}*x1+{e4}*t
for a small vector, find the spacetime interval, delta x squared
(%i22) dx2:x&*x$
        ldisplay(dx2)$
 (\%t23)/R/dx2 = -x3^2-x2^2-x1^2+t^2
The Survey, paragraph 4.1.2, actually suggests a spacetime split using
the geometric product; x&*gamma_zero, thus, using our g's as the gammas,
that is x&*g4; we could call this the split spacetime coordinate, splx
(%i24) splx:x&*g4;
(\%o24)/R/\{e3,e4\}*\%i*x3+\{e2,e4\}*\%i*x2+\{e1,e4\}*\%i*x1+t
for the imitation of G(1,3), the actual split into time and space is...
(%i25) realpart(splx);
        imagpart(splx);
(%o25) t
(\%026) \{e3,e4\} * \times 3 + \{e2,e4\} * \times 2 + \{e1,e4\} * \times 1
Created with wxMaxima.
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