

14.22 (a) I have tried to follow the procedure outlined in Prob. 14.15, but the elliptic orbit makes the calculation impossible. Instead, the following comes Landau & Lifshitz, Classical Theory of Fields, section 70. We are going to use the dipole radiation results, Prob. 9.9 (a) of Jackson. For the n -th harmonic, $\vec{p}(t) = \vec{p}_n e^{-in\omega_0 t}$. Taking into account of $-n$ component, in Gauss units, we have

$$P_n = 2 \cdot \frac{2}{3c^3} (n\omega_0)^4 |\vec{p}_n|^2 = \frac{4e^2}{3c^3} (n\omega_0)^4 |\vec{r}_n|^2$$

where $\vec{r}(t) = \vec{r}_n e^{-in\omega_0 t}$, or $\dot{\vec{r}}_n = \frac{\omega_0}{2\pi} \int_0^{2\pi/\omega_0} \dot{\vec{r}}(t) e^{in\omega_0 t} dt$. However, \vec{r}_n is still hard to compute. Notice that $\dot{\vec{v}}(t) = -in\omega_0 \dot{\vec{r}}_n e^{-in\omega_0 t} = \ddot{\vec{r}}_n e^{-in\omega_0 t}$, or equivalently, $\ddot{\vec{r}}_n = \frac{i}{n\omega_0} \ddot{\vec{v}}_n$.

On the other hand,

$$\ddot{\vec{v}}_n = \frac{\omega_0}{2\pi} \int_0^{2\pi/\omega_0} e^{in\omega_0 t} \ddot{\vec{r}}(t) dt = \frac{\omega_0}{2\pi} \int_0^{2\pi} e^{in(u-\varepsilon \sin u)} d\vec{r} = \frac{\omega_0}{2\pi} \int_0^{2\pi} e^{in(u-\varepsilon \sin u)} \left\{ \frac{-a \sin u}{a\sqrt{1-\varepsilon^2} \cos u} \right\} du$$

$$\text{Therefore, } \ddot{\vec{r}}_n = \frac{i}{2\pi n} \int_0^{2\pi} e^{in(u-\varepsilon \sin u)} \left\{ \frac{-a \sin u}{a\sqrt{1-\varepsilon^2} \cos u} \right\} du.$$

Using the identity $e^{in \cos \phi} = \sum_{m=-\infty}^{+\infty} i^m J_m(x) e^{im\phi}$, we can write the exponential term in the

integrand as

$$e^{in(u-\varepsilon \sin u)} = e^{inu} \exp \left\{ -in\varepsilon \cos \left(u + \frac{\pi}{2} \right) \right\} = e^{inu} \sum_{m=-\infty}^{+\infty} J_m(n\varepsilon) (-i)^m \exp \left\{ -im \left(u + \frac{\pi}{2} \right) \right\}$$

$$\text{Then, } \ddot{r}_{nx} = -\frac{ia}{2\pi n} \int_0^{2\pi} e^{in(u-\varepsilon \sin u)} \sin u du = -\frac{a}{4\pi n} \int_0^{2\pi} e^{inu} \sum_{m=-\infty}^{+\infty} J_m(n\varepsilon) (-i)^m e^{-im(u+\pi/2)} (e^{iu} - e^{-iu}) du$$

$$= -\frac{a}{2n} \left\{ J_{n+1}(n\varepsilon) (-i)^{n+1} e^{-i(n+1)\pi/2} - J_{n-1}(n\varepsilon) (-i)^{n-1} e^{-i(n-1)\pi/2} \right\}$$

$$= -\frac{a}{2n} (-i)^{n+1} \left\{ J_{n+1}(n\varepsilon) - J_{n-1}(n\varepsilon) \right\} = (-i)^n \frac{a}{n} J'_n(n\varepsilon).$$

$$\text{and } \ddot{r}_{ny} = \frac{ia\sqrt{1-\varepsilon^2}}{2n} \left\{ J_{n+1}(n\varepsilon) (-i)^{n+1} e^{-i(n+1)\pi/2} + J_{n-1}(n\varepsilon) (-i)^{n-1} e^{-i(n-1)\pi/2} \right\}$$

$$= (-i)^{n+1} \frac{ia\sqrt{1-\varepsilon^2}}{2n} \left\{ J_{n+1}(n\varepsilon) + J_{n-1}(n\varepsilon) \right\} = (-i)^{n+1} \frac{ia\sqrt{1-\varepsilon^2}}{n\varepsilon} J_n(n\varepsilon).$$

We finally arrive at the result for the power radiated in the n -th multiple of ω_0 as

$$P_n = \frac{4e^2}{3c^3} (n\omega_0)^2 a^2 \cdot \left\{ \frac{1}{n^2} \left[(J'_n(n\varepsilon))^2 + \frac{1-\varepsilon^2}{\varepsilon^2} J_n^2(n\varepsilon) \right] \right\}$$

(b) For circular orbit, $\epsilon \rightarrow 0$. In this limit, $J_n(n\epsilon) \rightarrow \left(\frac{n\epsilon}{2}\right)^n$, $J'_n(n\epsilon) \rightarrow \frac{n}{2} \left(\frac{n\epsilon}{2}\right)^{n-1}$,

and only $n=1$ term will contribute. Therefore,

$$P_1 = \frac{4e^2}{3c^3} \omega_0^4 a^2 \left(\frac{1}{4} + \frac{1}{4} \right) = \frac{2e^2}{3c^3} \omega_0^4 a^2.$$

For hydrogen-like atoms, its radius for n -th eigenstate is $a = n^2 a_0 / Z = \frac{n^2 \hbar^2}{Z m e^2}$. The fundamental frequency can be determined as

$$m \omega_0^2 a = \frac{Z e^2}{a^2}, \text{ or } \omega_0 = \frac{Z e^2}{m a^3} = \frac{Z e^2}{m} \left(\frac{Z m e^2}{n^2 \hbar^2} \right)^3 = \frac{Z^4 m^2 e^8}{n^6 \hbar^6}$$

Thus, $\omega_0 = \frac{Z^2 m e^4}{n^3 \hbar^3}$. The reciprocal mean life time is then given by

$$\begin{aligned} \frac{1}{\tau} &= P_1 / \hbar \omega_0 = \frac{2e^2}{3c^3} \frac{1}{\hbar} \omega_0^3 a^2 = \frac{2e^2}{3c^3} \frac{1}{\hbar} \left(\frac{Z^2 m e^4}{n^3 \hbar^3} \right)^3 \left(\frac{n^2 \hbar^2}{Z m e^2} \right)^2 \\ &= \frac{2e^2}{3c^3} \frac{1}{\hbar} \frac{Z^6 m^3 e^{12}}{n^9 \hbar^9} \frac{n^4 \hbar^4}{Z^2 m^2 e^4} = \frac{2}{3} \frac{Z^4 m e^{10}}{n^5 \hbar^6 c^3}, \end{aligned}$$

which agrees with Prob. 14.21 (a).