# Astronomical Image Processing with Hadoop

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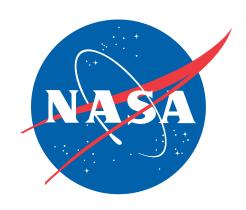


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### **Acknowledgments**

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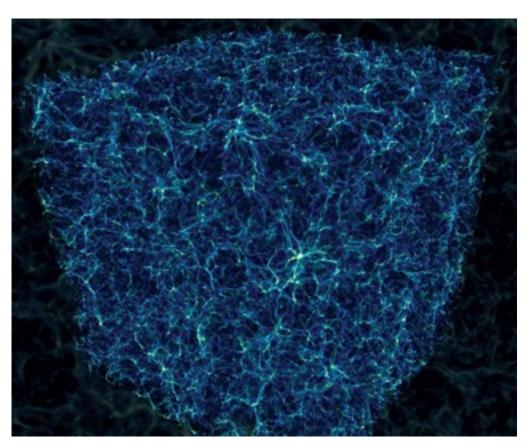


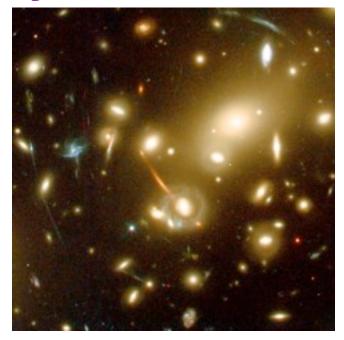
### **Session Agenda**

- Astronomical Survey Science
- Image Coaddition
- Implementing Coaddition within MapReduce
- Optimizing the Coaddition Process
- Conclusions
- Future Work

## **Astronomical Topics of Study**

- Dark energy
- Large scale structure of universe
- Gravitational lensing
- Asteroid detection/tracking







### What is Astronomical Survey Science?

- Dedicated sky surveys, usually from a single calibrated telescope/camera pair.
- Run for years at a time.
- Gather millions of images and TBs of storage\*.
- Require high-throughput data reduction pipelines.
- Require sophisticated off-line data analysis tools.

\* Next generation surveys will gather PBs of image data.

### **Sky Surveys: Today and Tomorrow**

- *SDSS*\* (1999-2005)
- Founded in part by UW
- 1/4 of the sky
- 80TBs total data



\* Sloan Digital Sky Survey

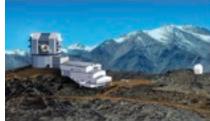
<sup>†</sup> Large Synoptic Survey Telescope

- *LSST*<sup>†</sup> (2015-2025)
- 8.4m mirror, 3.2 gigapixel camera
- Half sky every three nights
- 30TB per night......one SDSS every three nights
- 60PBs total (nonstop ten years)
- 1000s of exposures of each location







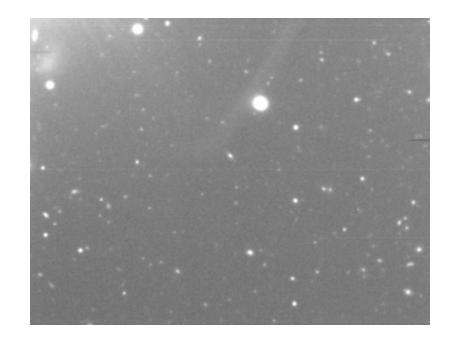


## FITS (Flexible Image Transport System)

- Common astronomical image representation file format
- Metadata tags (like EXIF):
  - Most importantly: Precise astrometry\*
  - Other:
    - Geolocation (telescope location)
    - Sky conditions, image quality, etc.

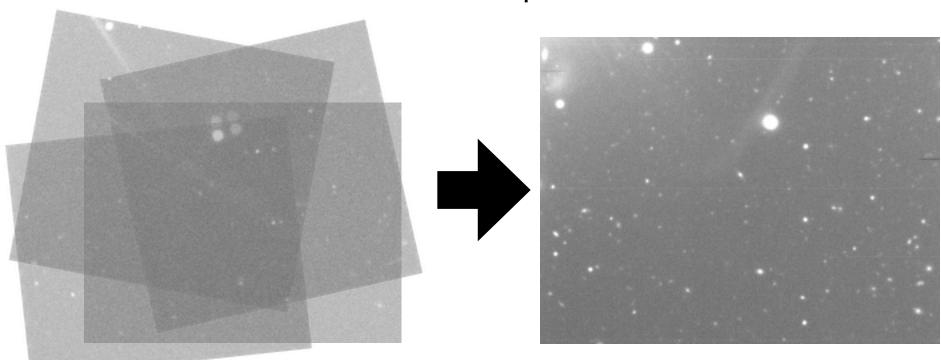
#### Bottom line:

 An image format that knows where it is looking.



## **Image Coaddition**

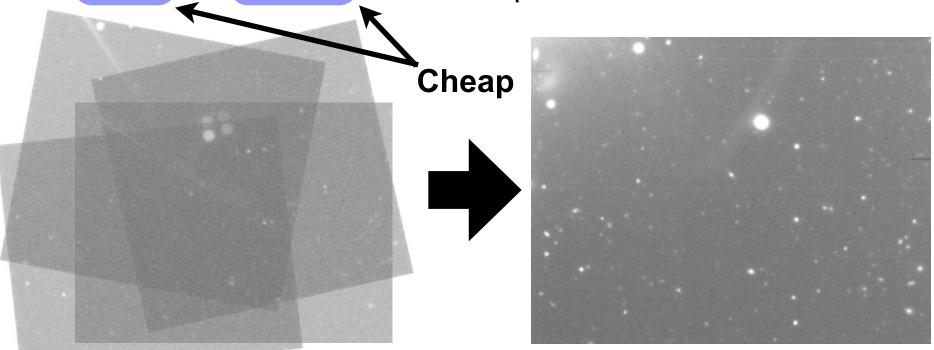
- Give multiple partially overlapping images and a *query* (color and sky bounds):
  - > Find images' intersections with the query bounds.
  - > Project bitmaps to the bounds.
  - > Stack and mosaic into a final product.



## **Image Coaddition**

### **Expensive**

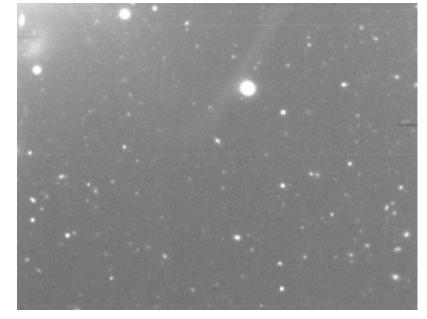
- Give multiple partially overlapping images and a *query* (color and sky bounds):
  - > Find images intersections with the query bounds.
  - > **Project** bitmaps to the bounds.
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## Image Stacking (Signal Averaging)



- Stacking improves SNR:
  - Makes fainter objects visible.
- Example (SDSS, Stripe 82):
  - Top: Single image, R-band
  - Bottom: 79-deep stack
    - (~9x SNR improvement)
    - Numerous additional detections



Variable conditions (e.g., atmosphere, PSF, haze) mean stacking algorithm complexity can exceed a mere sum.

### **Existing Image Coaddition Systems**

- SWarp
  - Multi-threaded parallelism (single machine only).
- SDSS coadds of Stripe 82 (Fermilab)
  - Same dataset used in our work.
  - One-off project not a general-purpose tool.
- Montage (run on TeraGrid)
  - Most similar to our work.
  - > MPI (complicated), TeraGrid (dedicated, expensive).
- MapReduce (our work, this talk)
  - High-level, potentially simpler to program.
  - Scalable on cheap commodity hardware (the cloud).

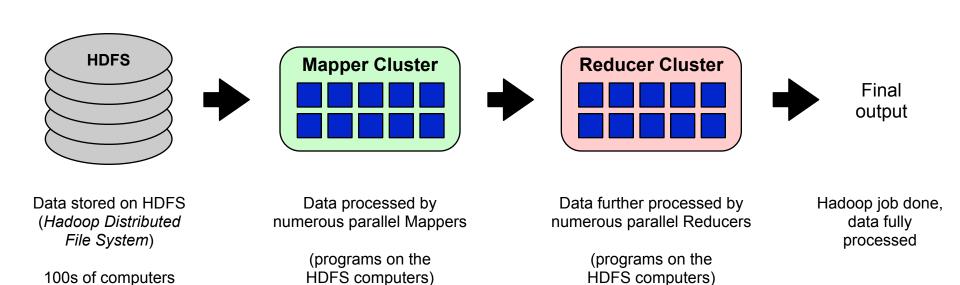
## **Advantages of MapReduce (Hadoop)**

- High-level problem description. No effort spent on internode communication, message-passing, etc.
- Programmed in Java (accessible to most science researchers, not just computer scientists and engineers).
- Runs on cheap commodity hardware, potentially in the cloud, e.g., Amazon's EC2.
- Scalable: 1000s of nodes can be added to the cluster with no modification to the researcher's software.
- Large community of users/support.

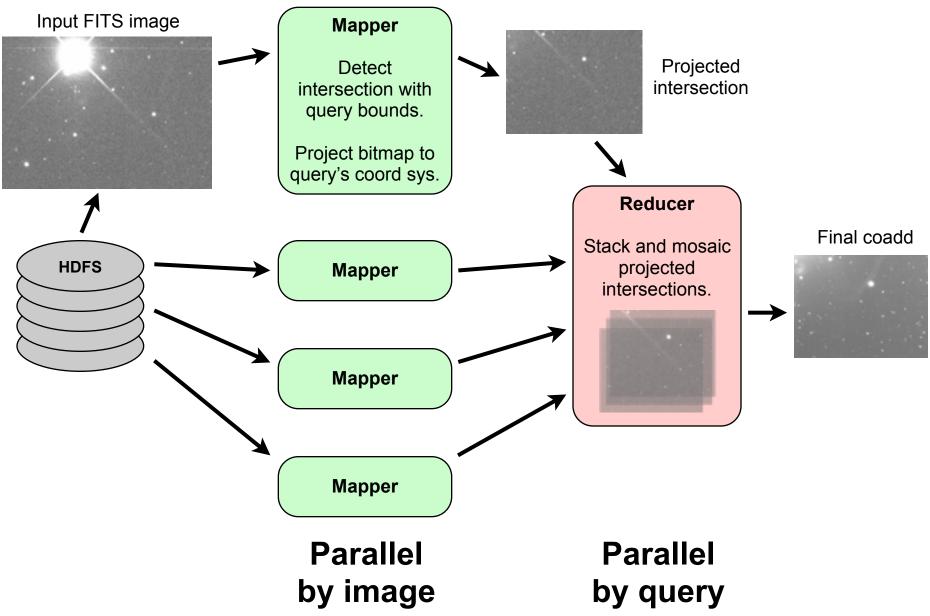
### Hadoop

- A massively parallel database-processing system:
  - In one sense: a parallel computing system (a cluster)
  - In another sense: a parallel database
  - It's both!

in a cluster

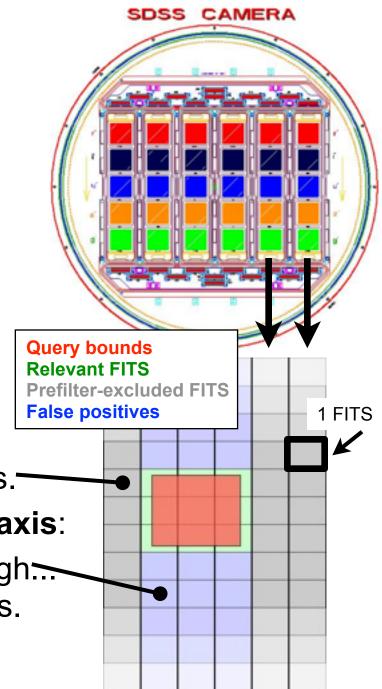


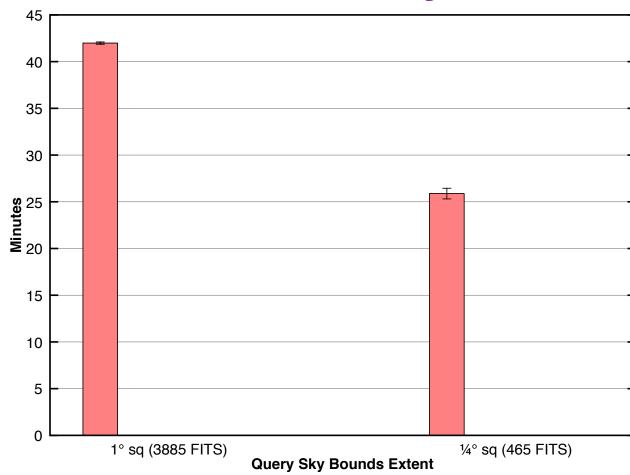
### **Coaddition in Hadoop**



## **Driver Prefiltering**

- To assist the process we prefilter the FITS files in the driver.
- SDSS camera has 30 CCDs:
  - > 5 colors
  - > 6 abutting strips of sky
- Prefilter (path glob) by color and sky coverage (single axis):
  - Exclude many irrelevant FITS files.
  - > Sky coverage filter is only single axis:
    - Thus, false positives slip through...
       ...to be discarded in the mappers.





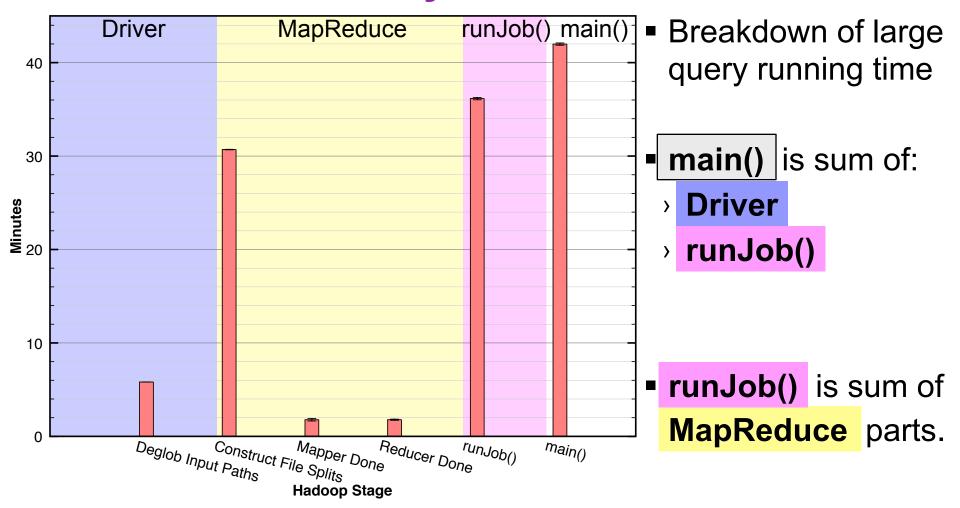
Error bars show 95% confidence intervals — Outliers removed via Chauvenet

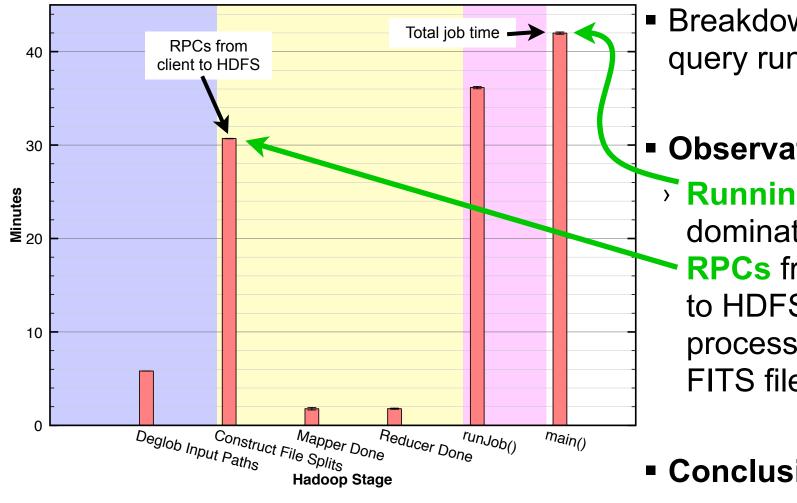
### Running time:

- > 2 query sizes
- Run against1/10th of SDSS(100,058 FITS)

#### Conclusion:

- Considering the small dataset, this is too slow!
- Remember 42 minutes for the next slide.





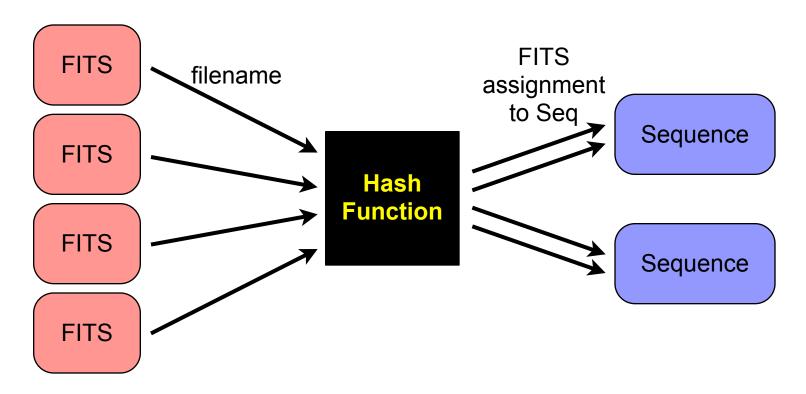
- Breakdown of large query running time
- Observation:
  - > Running time dominated by **RPCs** from client to HDFS to process 1000s of FITS file paths.

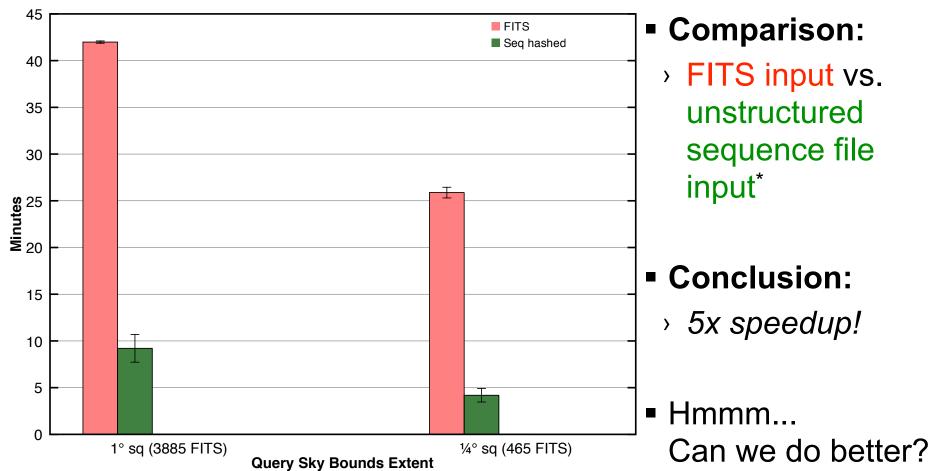
#### Conclusion:

Need to reduce number of files.

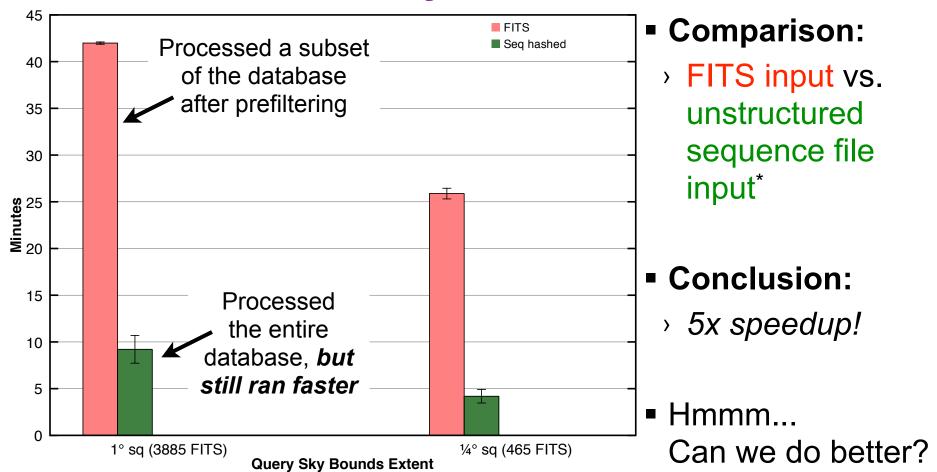
### **Sequence Files**

- Sequence files group many small files into a few large files.
- Just what we need!
- Real-time images may not be amenable to logical grouping.
  - > Therefore, sequence files filled in an arbitrary manner:





<sup>\* 360</sup> seq files in hashed seq DB.

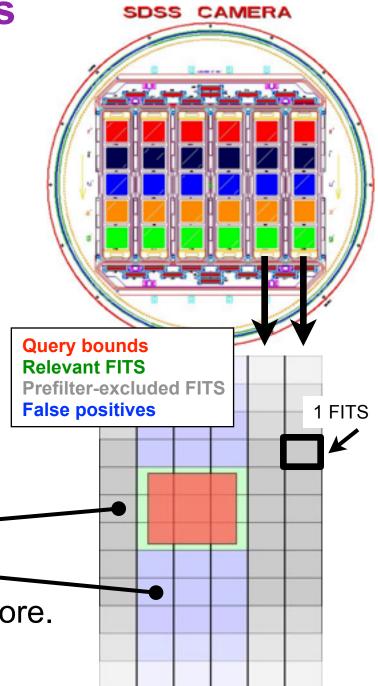


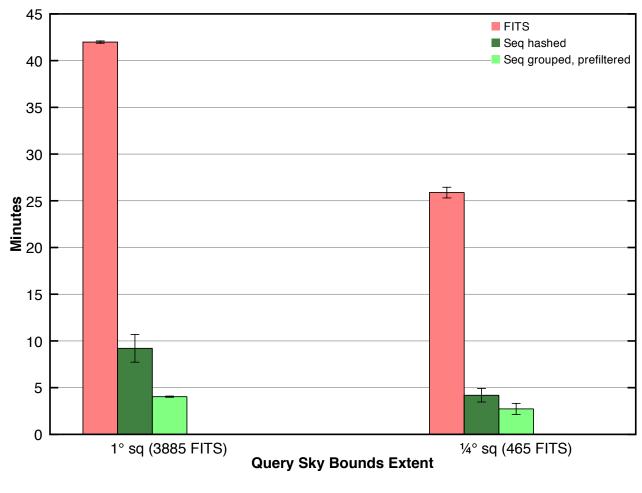
<sup>\* 360</sup> seq files in hashed seq DB.

**Structured Sequence Files** 

 Similar to the way we prefiltered FITS files...

- SDSS camera has 30 CCDs:
  - > 5 colors
  - > 6 abutting strips of sky
  - > Thus, 30 sequence file types
- Prefilter by color and sky coverage (single axis):
  - Exclude irrelevant sequence files:
  - Still have false positives.-
  - Catch them in the mappers as before.





### Comparison:

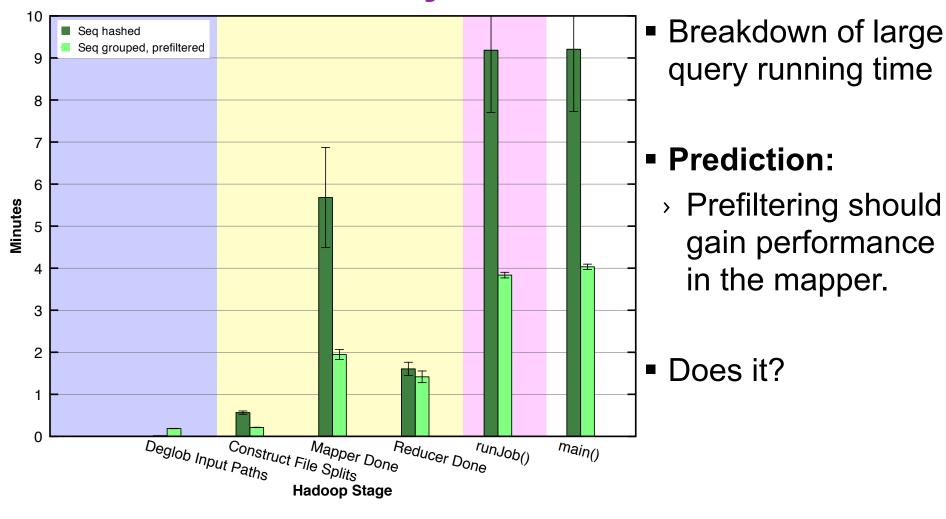
FITS vs.
 unstructured
 sequence\* vs.
 structured
 sequence files\*

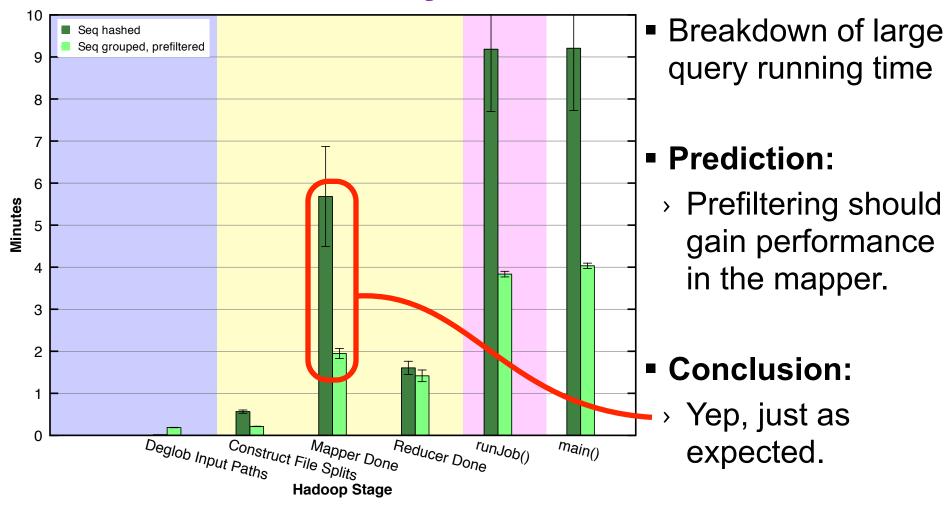
#### Conclusion:

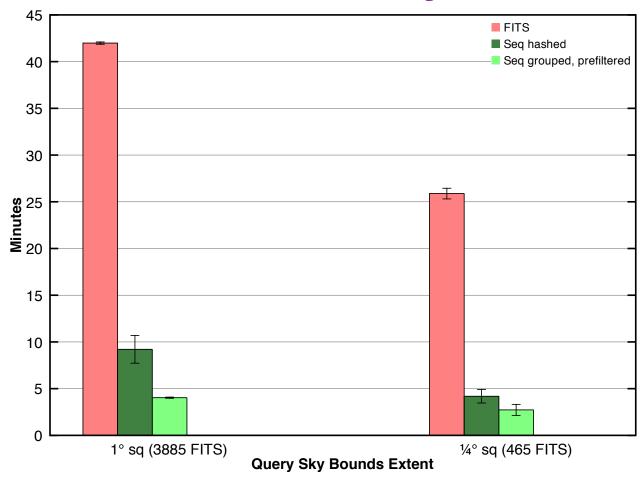
Another 2x
 speedup for the large query, 1.5x
 speedup for the small query.

<sup>\* 360</sup> seq files in hashed seq DB.

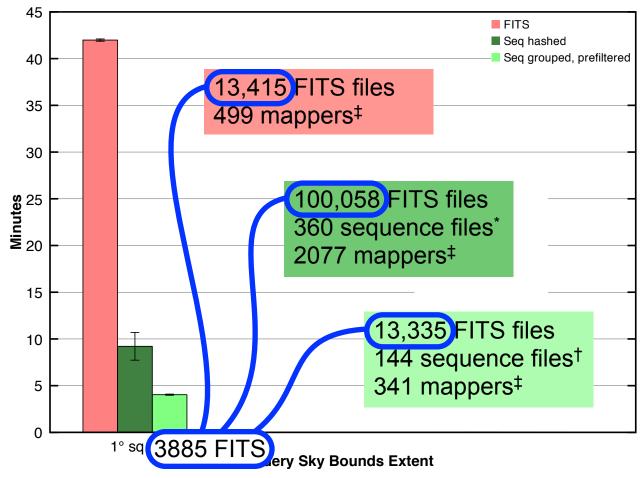
<sup>† 1080</sup> seq files in structured DB.







- Experiments were performed on a 100,058 FITS database (1/10th SDSS).
- How much of this database is Hadoop churning through?



#### \* 360 seq files in hashed seq DB.

### Comparison:

Number of FITS considered in mappers vs. number contributing to coadd

#### Conclusion:

 Mappers must discard many FITS files due to nonoverlap of query bounds.

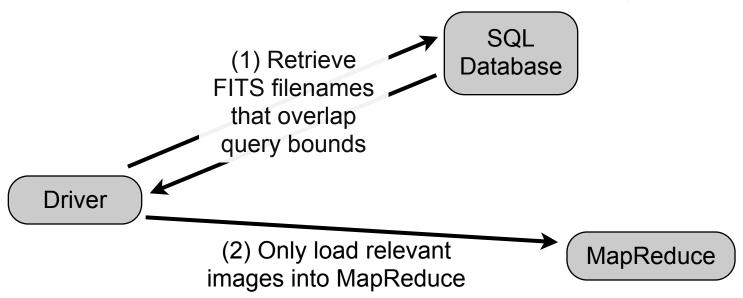
<sup>&</sup>lt;sup>†</sup> 1080 seq files in structured DB.

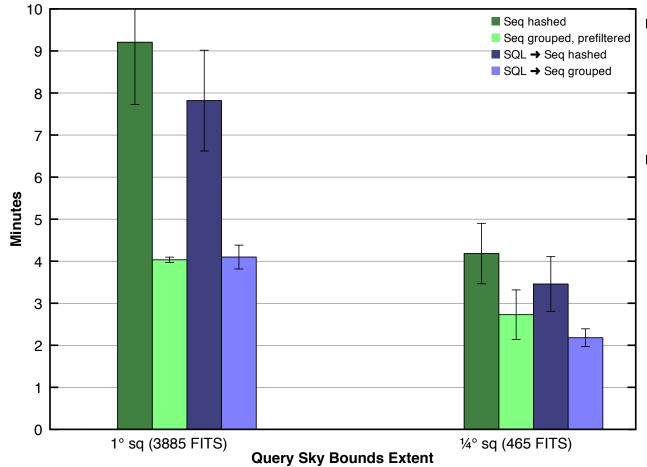
<sup>&</sup>lt;sup>‡</sup> 800 mapper slots on cluster.

### **Using SQL to Find Intersections**

- Store all image colors and sky bounds in a database:
  - First, query color and intersections via SQL.
  - Second, send only relevant images to MapReduce.
- Consequence:

All images processed by mappers contribute to coadd. No time wasted considering irrelevant images.



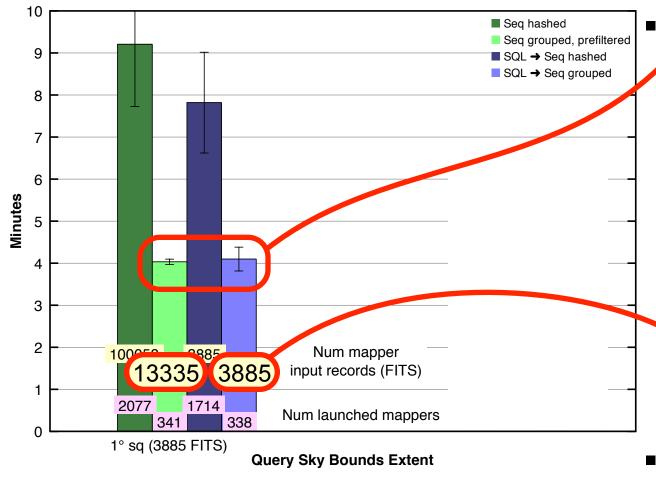


### Comparison:

> nonSQL vs. SQL

#### Conclusion:

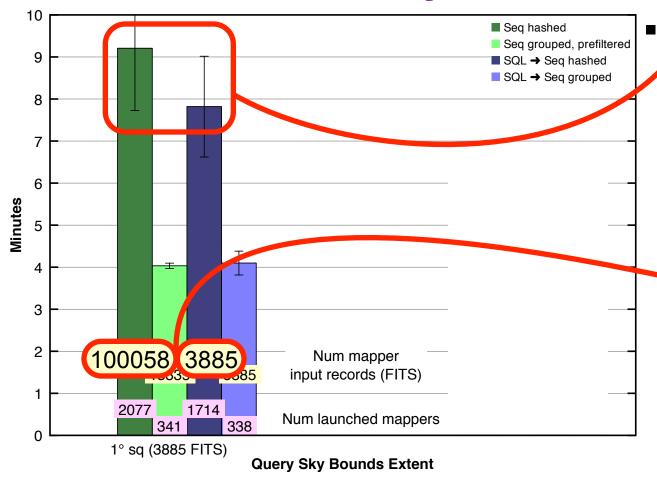
Sigh, no major improvement
 (SQL is not remarkably superior to nonSQL for given pairs of bars).



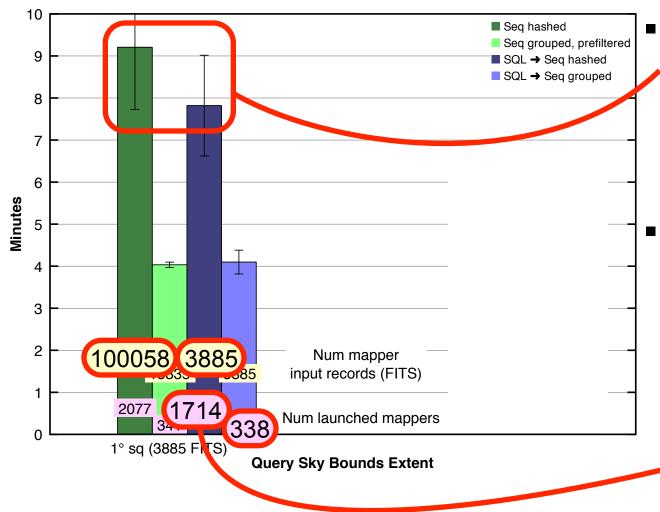
- Comparable performance here makes sense:
  - In essence,
     prefiltering and
     SQL performed
     similar tasks,
     albeit with 3.5x
     different mapper inputs (FITS).

#### Conclusion:

 Cost of discarding many images in the nonSQL case was negligible.



- Low improvement for SQL in the hashed case is surprising at first
  - ...especially considering 26x
     different mapper inputs (FITS)!



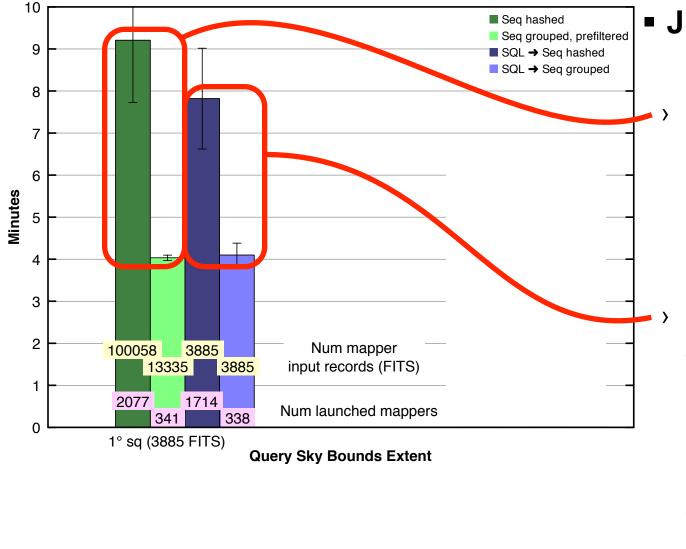
Low improvement
 for SQL in the
 hashed case is
 surprising at first

### Theory:

Scattered
 distribution of
 relevant FITS
 prevented
 efficient mapper
 reuse.

Starting each mapper is expensive. This overhead hurt overall performance.

### Results



Just to be clear:

Prefiltering improved due to reduction of mapper load.

SQL improved due to data locality and more efficient mapper allocation – the required work was unaffected (3885 FITS).

### **Utility of SQL Method**

- Despite our results (which show SQL to be equivalent to prefiltering)...
- ...we predict that SQL should outperform prefiltering on larger databases.
- Why?
  - Prefiltering would contend with an increasing number of false positives in the mappers\*.
  - SQL would incur little additional overhead.
- No experiments on this yet.

<sup>\*</sup> A spacing-filling curve for grouping the data might help.

### **Conclusions**

- Packing many small files into a few large files is essential.
- Structured packing and associated prefiltering offers significant gains (reduces mapper load).
- SQL prefiltering of unstructured sequence files yields little improvement (failure to combine scattered HDFS file-splits leads to mapper bloat).
- SQL prefiltering of structured sequence files performs comparably to driver prefiltering, but we anticipate superior performance on larger databases.
- On a shared cluster (e.g. the cloud), performance variance is high – doesn't bode well for online applications. Also makes precise performance profiling difficult.

### **Future Work**

- Parallelize the reducer.
- Less conservative CombineFileSplit builder.
- Conversion to C++, usage of existing C++ libraries.
- Query by time-range.
- Increase complexity of projection/interpolation:
  - > PSF matching
- Increase complexity of stacking algorithm:
  - Convert straight sum to weighted sum by image quality.
- Work toward the larger science goals:
  - Study the evolution of galaxies.
  - Look for moving objects (asteroids, comets).
  - Implement fast parallel machine learning algorithms for detection/classification of anomalies.

### **Questions?**

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