

planetMagFields: A Python package for analyzing and plotting planetary magnetic field data

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Summary

Long term observations and space missions have generated a wealth of data on the magnetic fields of the Earth and other solar system planets (Alken et al., 2021; Anderson et al., 2012; Cao et al., 2020; Connerney et al., 1987, 1991, 2022; Kivelson et al., 2002). planetMagFields is a Python package designed to have all the planetary magnetic field data currently available in one place and to provide an easy interface to access the data. planetMagFields focuses on planetary bodies that generate their own magnetic field, namely Mercury, Earth, Jupiter, Saturn, Uranus, Neptune and Ganymede. planetMagFields provides functions to compute as well as plot the magnetic field on the planetary surface or at a distance above or under the surface. It also provides functions to filter out the field to large or small scales as well as to produce .vts files to visualize the field in 3D using Paraview (Ahrens et al., 2005; Ayachit, 2015), VisIt (Childs et al., 2012) or similar rendering software. Lastly, the planetMagFields repository also provides a Jupyter notebook for easy interactive visualizations.

Statement of need

Planetary scientists studying the magnetic field of planets need to constantly access, visualize, analyze and extrapolate magnetic field data. In addition, with technological advancements in space exploration and planetary missions, we are constantly getting new data for planetary magnetic fields and hence, better field models. Though reviews of these field models are often written (Schubert & Soderlund, 2011; Stanley, 2014), there is very little software available that provides easy access to these models with a high level language and a way to easily visualize and analyze them. To the knowledge of the authors, there are a few publicly available repositories that are capable of providing access to planetary magnetic field data and tools to analyze them such as JupiterMag (James et al., 2024; Wilson et al., 2023), KMAG (Khurana, 2020), ChaosMagPy (Kloss, 2024), SHTools (Wieczorek & Meschede, 2018), PlanetMag (Styczinski & Cochrane, 2024) and libinternalfield (<https://github.com/mattkjames7/libinternalfield>). Out of these, only libinternalfield provides data and software to analyze and access magnetic fields of all planets. However, it is a C++ library which needs to be interfaced with something at a higher level to enable fast analyses and visualization. Thus, a software package that has different magnetic field models for all different planets of the solar system in one place, as well as provides a high level API to access, analyze and visualize them is not available. planetMagFields is intended not only to currently fill this gap, but also to provide a central repository, to be constantly updated, as more magnetic field models become available.

In addition to the research aspect of our software, the interactive Jupyter notebook serves as a valuable educational resource, fostering a deeper appreciation for the complexities of planetary magnetic environments.

Mathematics

Magnetic fields in planets are generated by electric currents in a fluid region inside them through a process called dynamo action (Jones, 2011; Schubert & Soderlund, 2011; Stanley, 2014). Outside this region, in the absence of current sources, the magnetic field \vec{B} can be written as the gradient of a scalar potential, $\vec{B} = -\nabla V$. The potential V is usually written as an expansion in orthogonal functions in spherical coordinates (r, θ, ϕ) ,

$$V = R_p \sum_{l,m} \left(\frac{R_p}{r} \right)^{l+1} [g_l^m \cos(m\phi) + h_l^m \sin(m\phi)] P_l^m(\cos \theta), \quad (1)$$

where, g_l^m and h_l^m are called the Gauss coefficients. R_p represents the radius of the planet and P_l^m are associated Legendre functions of order l and degree m , where l and m are integers. The above equation can be recast in terms of spherical harmonics, which is what the code uses.

The raw data obtained from satellites or space missions are usually inverted to obtain these Gauss coefficients which are the key to describing the surface magnetic field of a planet as well as how that field looks at a certain altitude from the surface. The magnetic energy content on the surface in a certain degree l is given by the Lowes spectrum:

$$R_l = (l+1) \sum_m ((g_l^m)^2 + (h_l^m)^2),$$

l plays the role of a wavenumber. Low degrees represent large spatial features in the field while high degrees represent small scale features. The maximum available degree l_{max} of data for a particular planet depends on the quality of observations.

Benchmarking

We benchmarked our software against two publicly available repositories : JupiterMag (James et al., 2024; Wilson et al., 2023) for Jupiter and the CHAOS-7 (Finlay et al., 2020; Kloss, 2024) for Earth. For Jupiter, we compare the field at a depth of 85% of planetary radius, thus testing our extrapolation capability while for Earth, we compare the field on the surface in 2016, testing our implementation of taking into account changes in the Earth's field in a linear fashion (as is done for the IGRF model, Alken et al., 2021) The comparison for Jupiter is shown in Figure 1. We also use these cases in our unit testing.

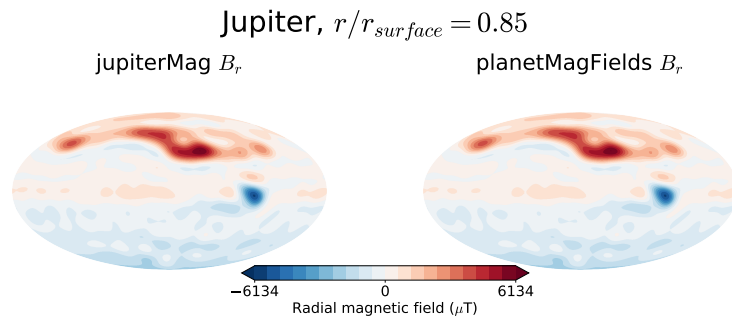


Figure 1: Benchmarking the code against publicly available repositories.

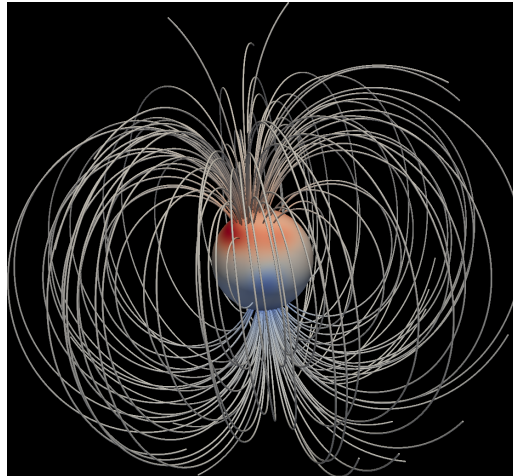


Figure 3: 3D rendering of Jupiter's magnetic field using Paraview, using a vts file produced by planet-Magfields.

Magnetic field models used

planetMagfields currently supports the following magnetic field models:

- *Mercury* : Anderson et al. (2012) , Thébault et al. (2018), Wardinski et al. (2019)
- *Earth* : The International Geomagnetic Reference Field (IGRF) (Alken et al., 2021)
- *Jupiter* : The VIP-4 model (Connerney et al., 1998), JRM09 (Connerney et al., 2018), JRM33 (Connerney et al., 2022)
- *Saturn* : Cassini Saturn orbit insertion (SOI) (Burton et al., 2009), Cassini11 (Dougherty et al., 2018), Cassini11+ (Cao et al., 2020)
- *Uranus* : Connerney et al. (1987)
- *Neptune* : Connerney et al. (1991)
- *Ganymede* : Kivelson et al. (2002)

When new magnetic field models become available, either through newly available data or through reanalysis of existing observations, we will add them to the current repository, either ourselves or through a community effort of pull requests.

Jupyter frontend

We provide a Jupyter notebook, along with a binder link (<https://mybinder.org/v2/gh/AnkitBarik/planetMagFields/HEAD?labpath=%2FExploreFieldsInteractively.ipynb>) that gives interactive access for visualizing the radial magnetic fields and the corresponding Lowes spectra at various depths, with different background color options.

Documentation

The software has been documented using Sphinx (<https://www.sphinx-doc.org/>) and the documentation is available here: <https://ankitbarik.github.io/planetMagFields/>.

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